

Euclid mission Overview

Sandrine Pires
CEA Paris-Saclay



on behalf of the Euclid Consortium



www.euclid-ec.org

Credits

All the Euclid material shown here is on behalf of and approved by the Euclid consortium and ESA

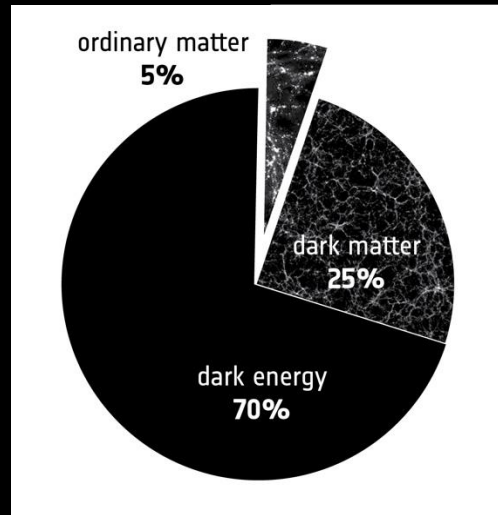
For more information and proper credit to national space agencies and funding organisations : <https://www.euclid-ec.org>

Pictures and movie:

https://www.esa.int/Science_Exploration/Space_Science/Euclid

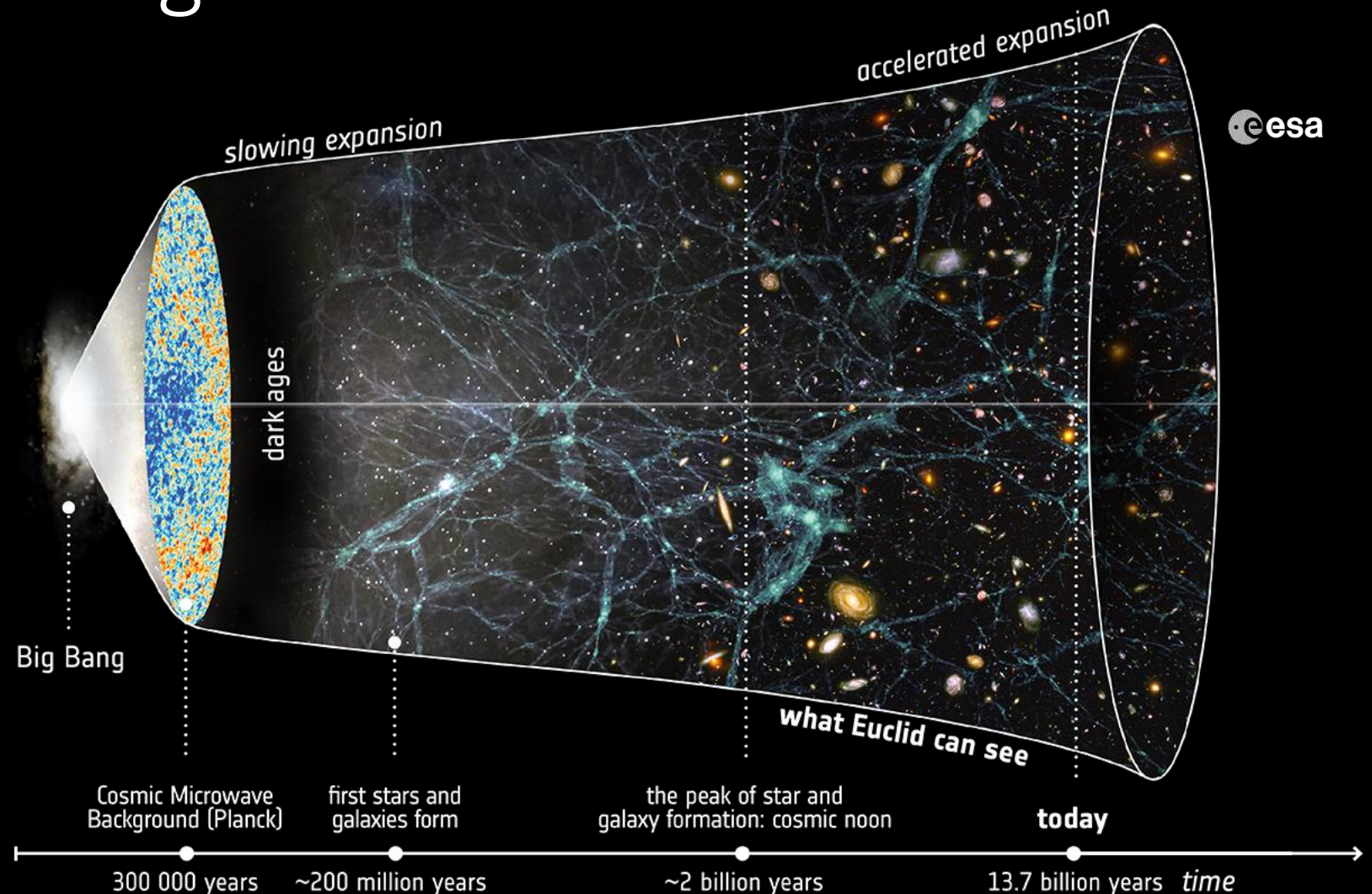


Standard cosmological model

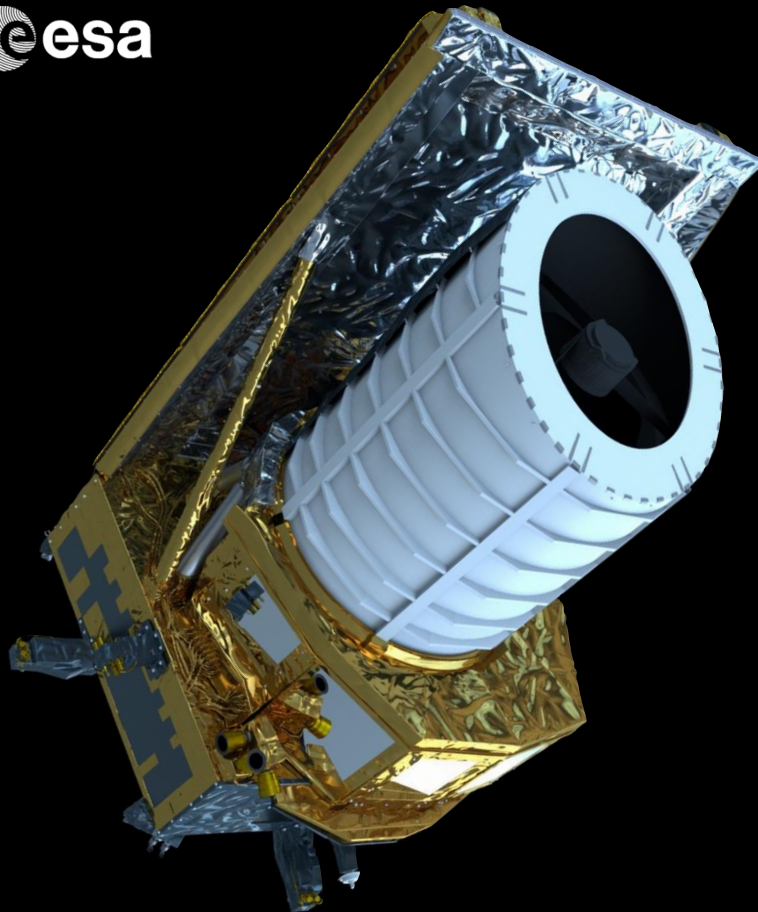


Cosmological parameters :

- H_0 - current expansion rate
- Ω_M - matter density
- Ω_b - baryon density
- Ω_Λ - dark energy density
- σ_8 - matter density fluctuations
- n_s - scale index of initial fluctuations

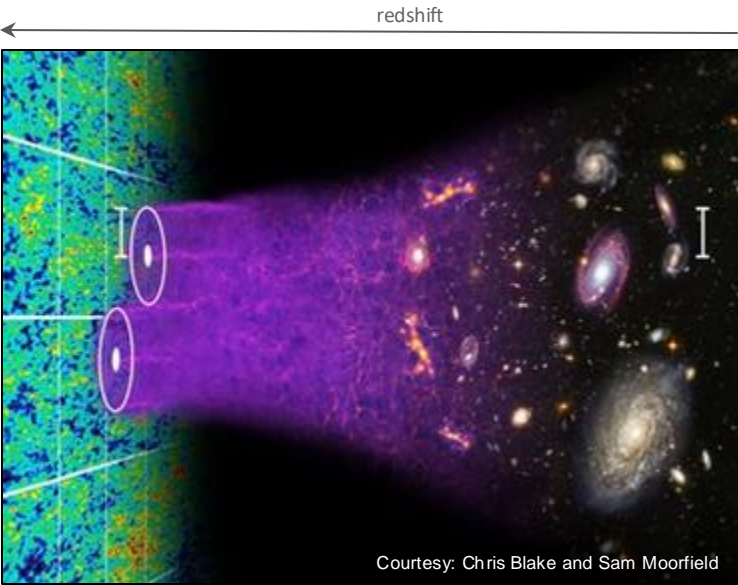
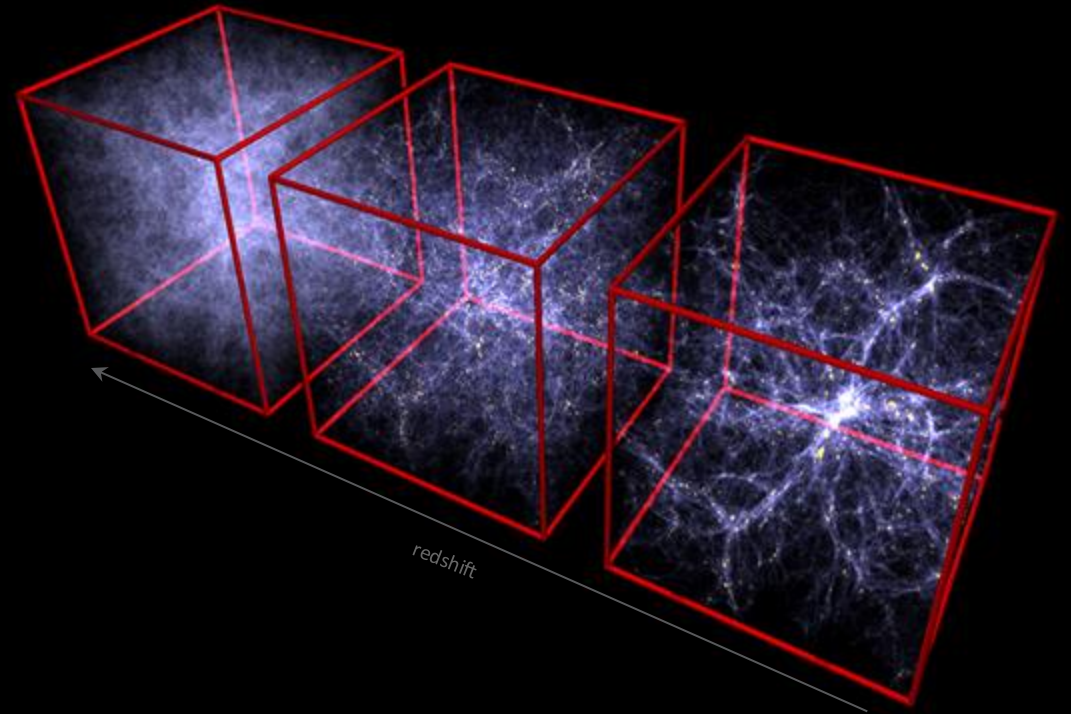


Mapping the dark Universe with Euclid



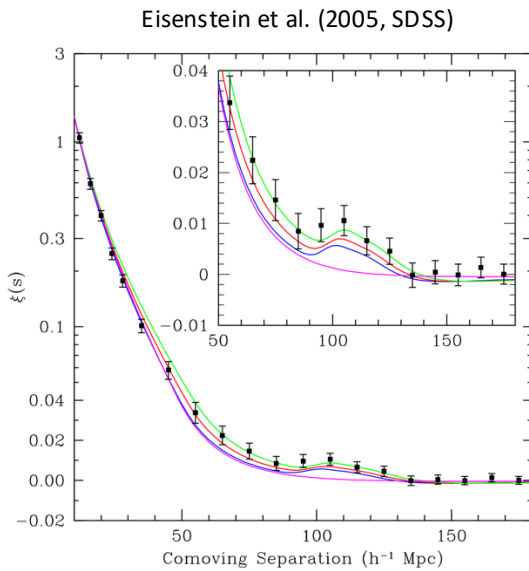
- ESA Cosmic Vision medium-class mission program
two original proposals (2007):
 - DUNE (imaging, Refregier et al.)
 - SPACE (spectroscopy, Cimatti et al.)
- Both accepted and merged into **Euclid**
 - simultaneously map the visible and dark matter distribution over one-third of the sky,
 - using galaxy redshifts and weak gravitational lensing
- The mission was selected in 2011 and adopted in 2012

3D Galaxy Clustering

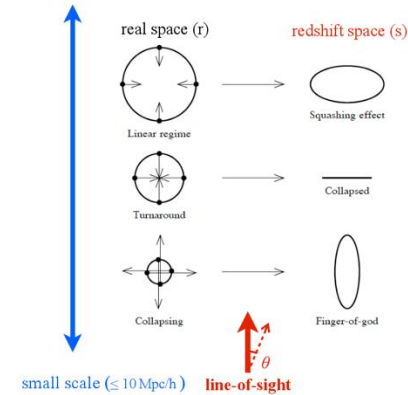


Courtesy: Chris Blake and Sam Moorfield

Baryonic Acoustic Oscillation (BAO) provides a cosmic standard ruler and is sensitive to the expansion history



large scale (~100 Mpc/h)

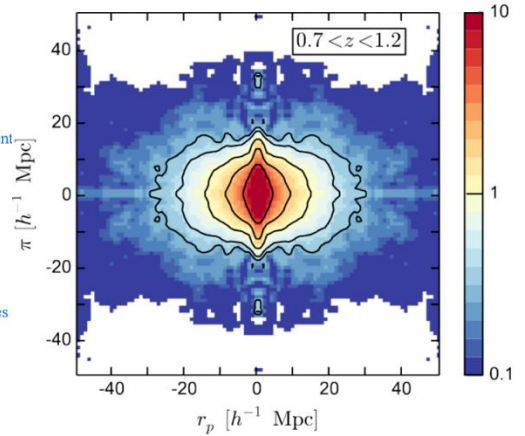


Kaiser effect:
Directional squashing,
due to central gravitational potential

Finger of God effect (FoG):
Elongation along line-of-sight
by virialized motions of galaxies

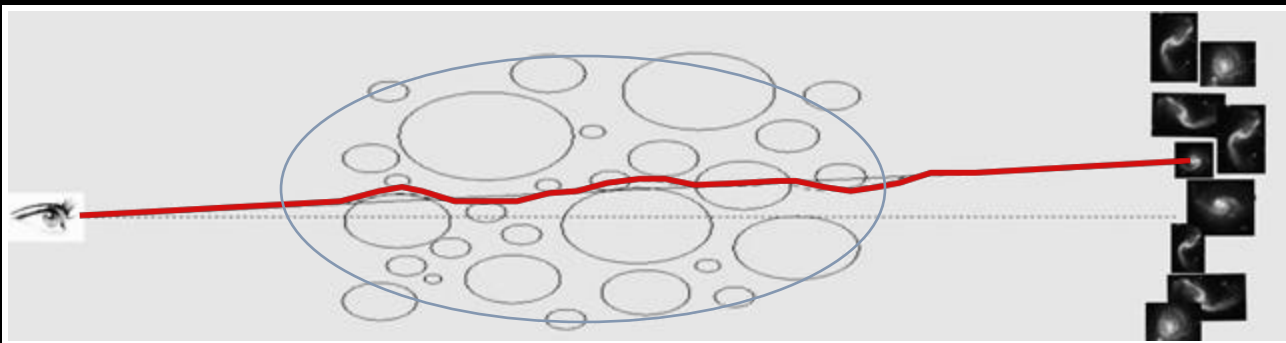
small scale (≤ 10 Mpc/h) **line-of-sight**

Pezzotta et al. (2017, VIPERS)

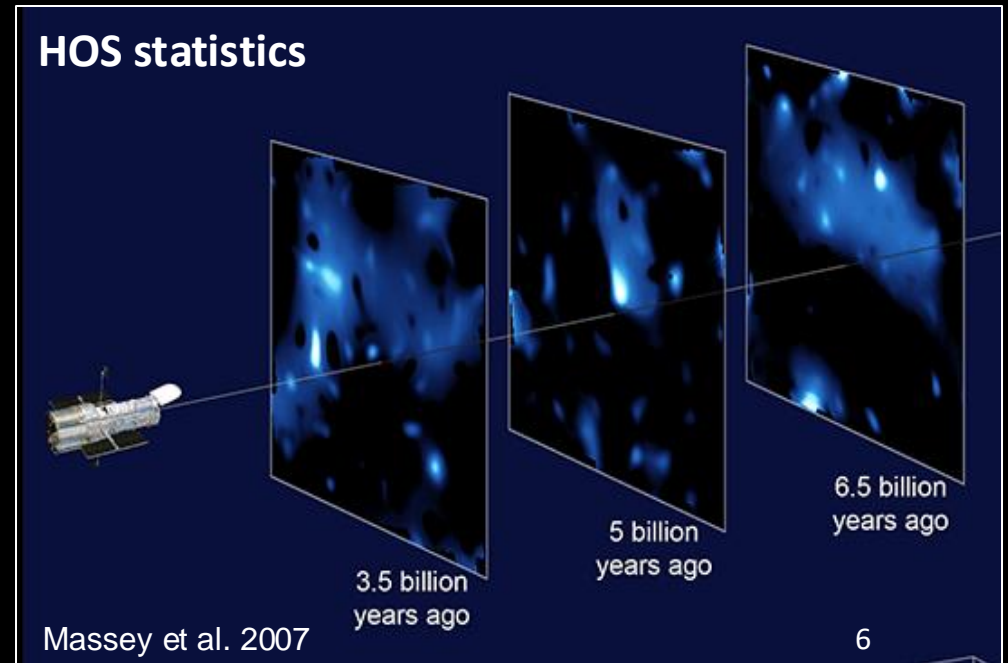
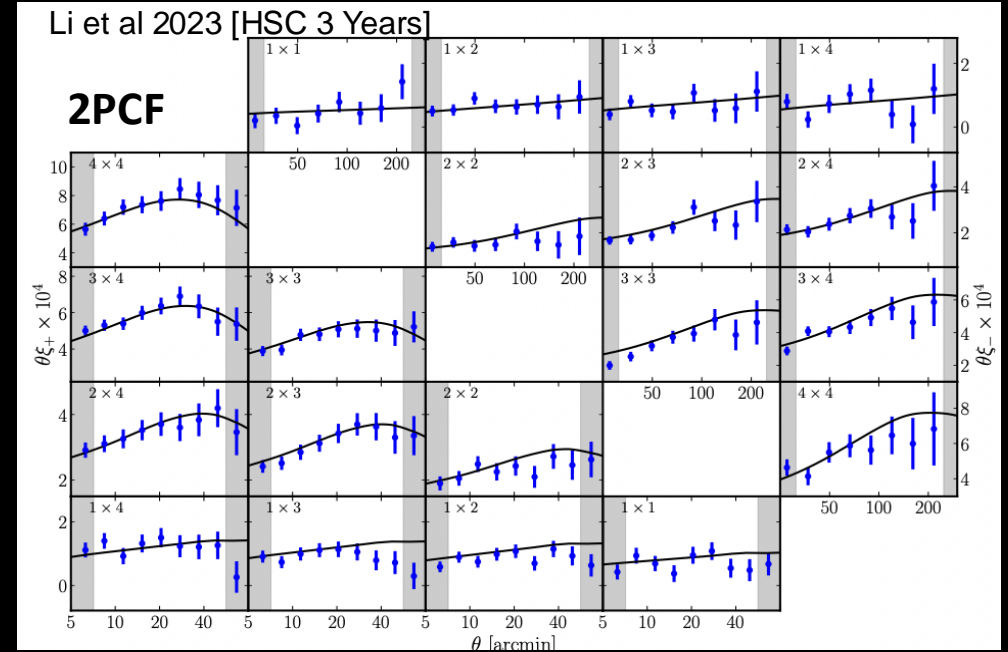


Redshift Space Distortion (RSD) provides constraints on the growth of structure

Weak Gravitational Lensing



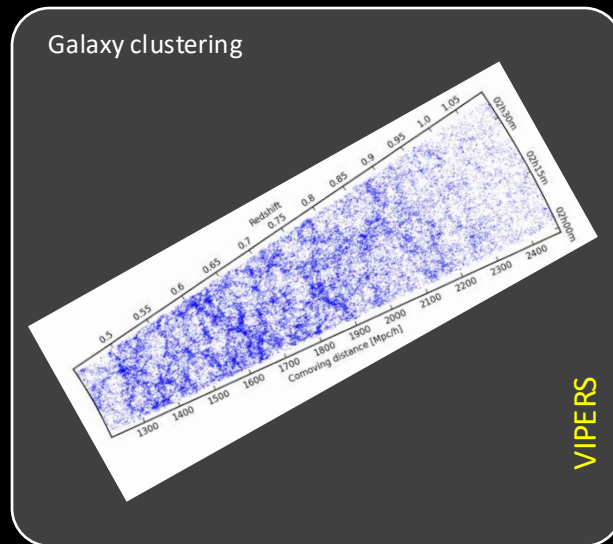
Gravitational lens



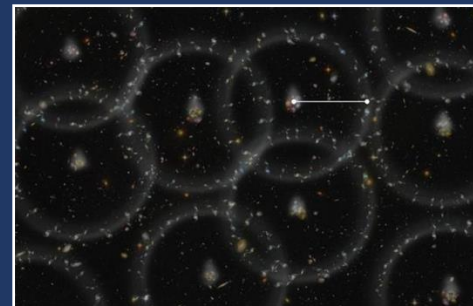
Euclid : Unveiling gravity and dark energy

 Φ

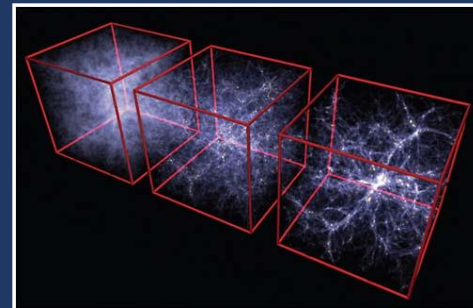
Governs the dynamic of non-relativistic objects



Expansion history

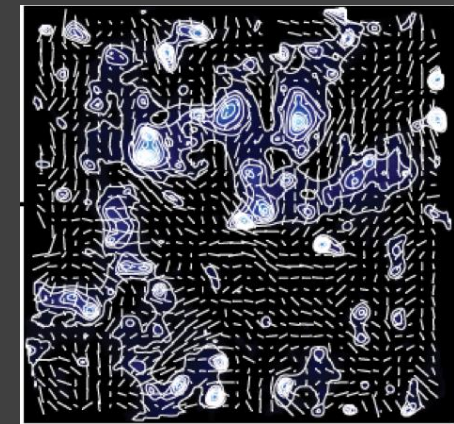


Growth of structure

 Ψ_w

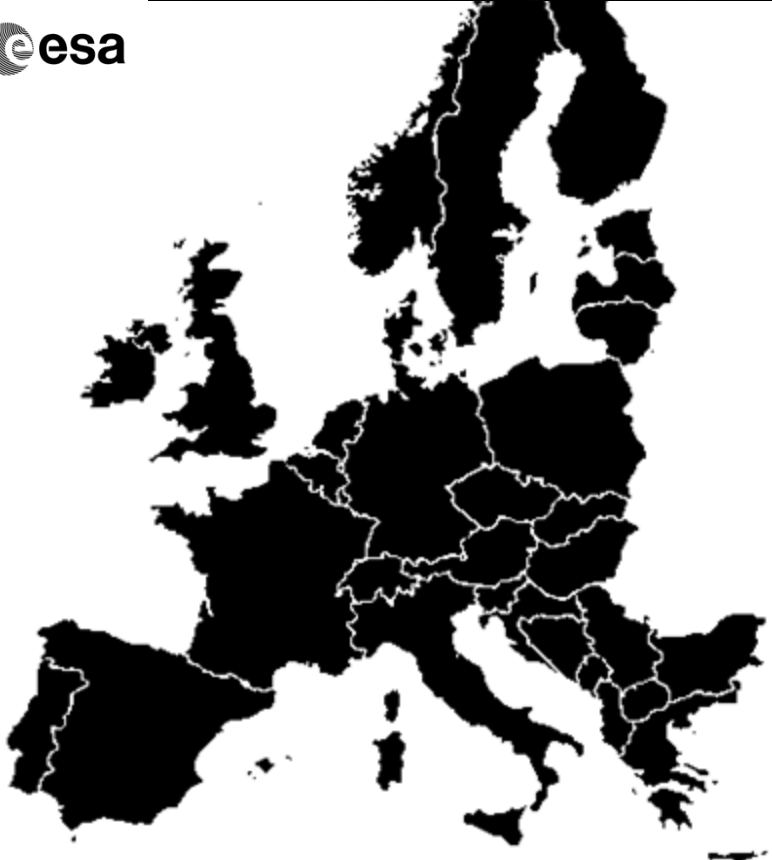
Governs the light propagation

Weak lensing





esa



- More than 2700 registered scientists;
EC Lead Y. Mellier
- 15 European countries + USA + Canada +
Japan
- Responsible for the two Euclid instruments
and the reduction and analysis of the data
(Science Ground Segment)

Spacecraft

- Satellite: Thales-Alenia Space
- Payload (telescope): Airbus Defence and Space

- Launch mass: 1988 kg
- Propellant: 137 kg hydrazine, lasting 14 years
- Data downlink: 4h / day, 820 Gbit

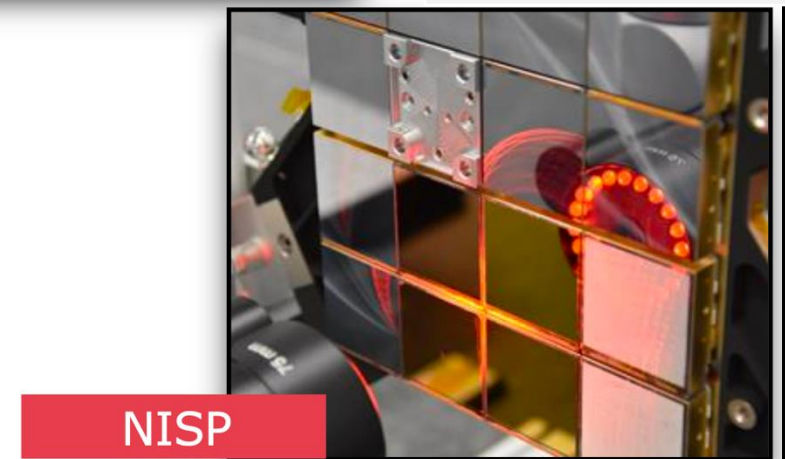
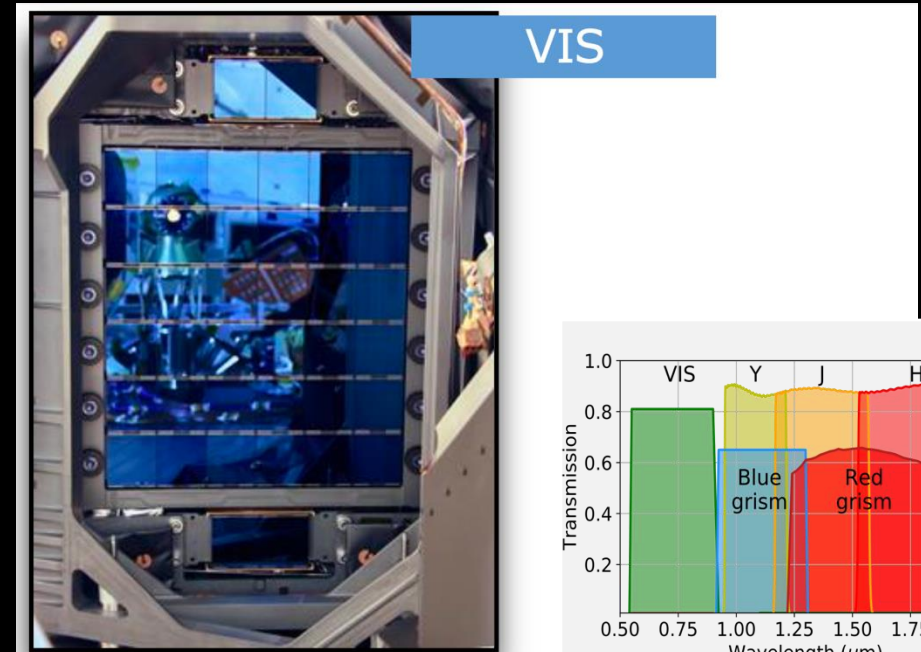
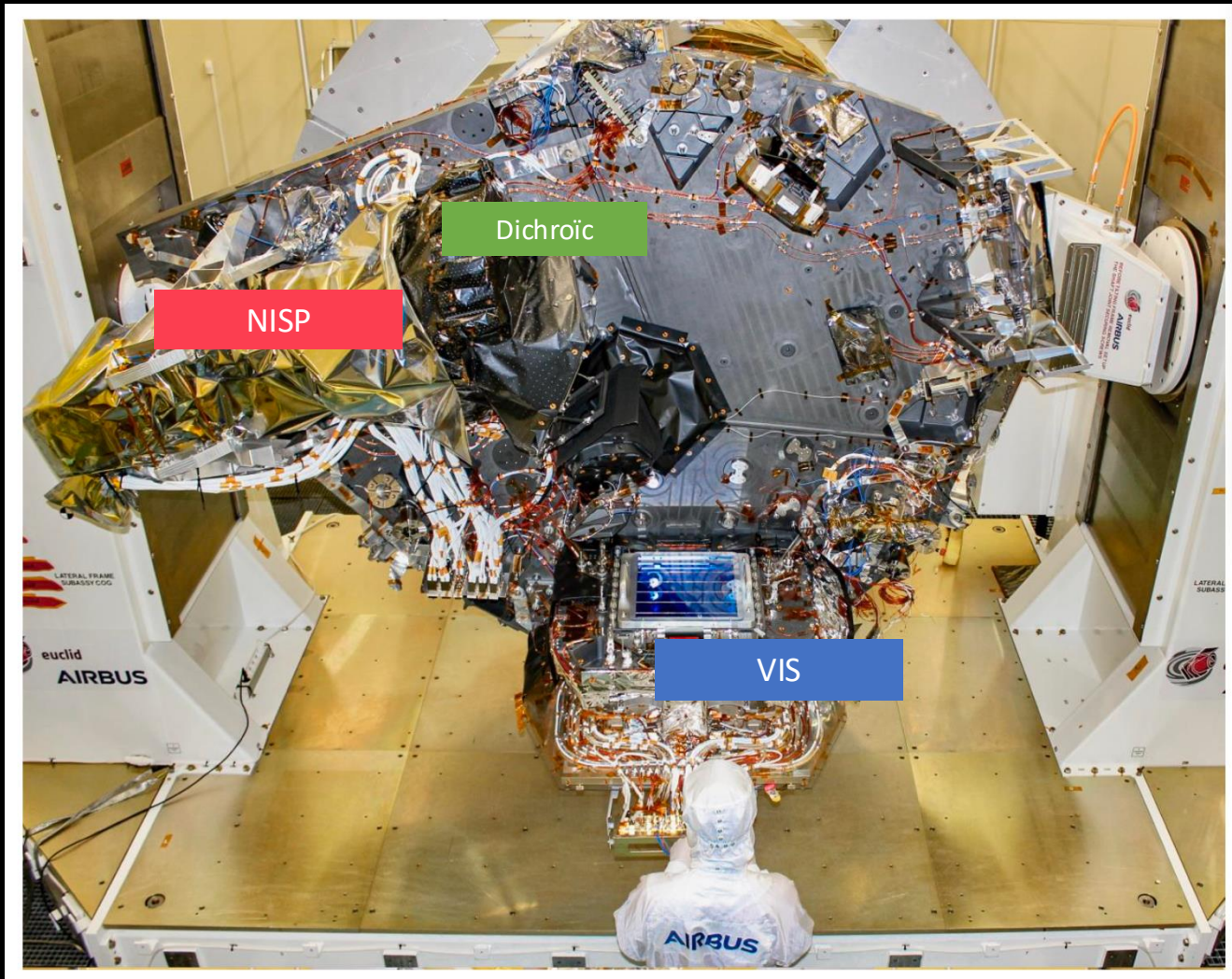
- 1.2m Korsch telescope in off-axis configuration
- Field of view 0.7 x 0.7 degrees

- Simultaneous observations (dichroic beam splitter)
- **VIS** (optical imaging)
- **NISP** (NIR imaging and slitless spectroscopy)

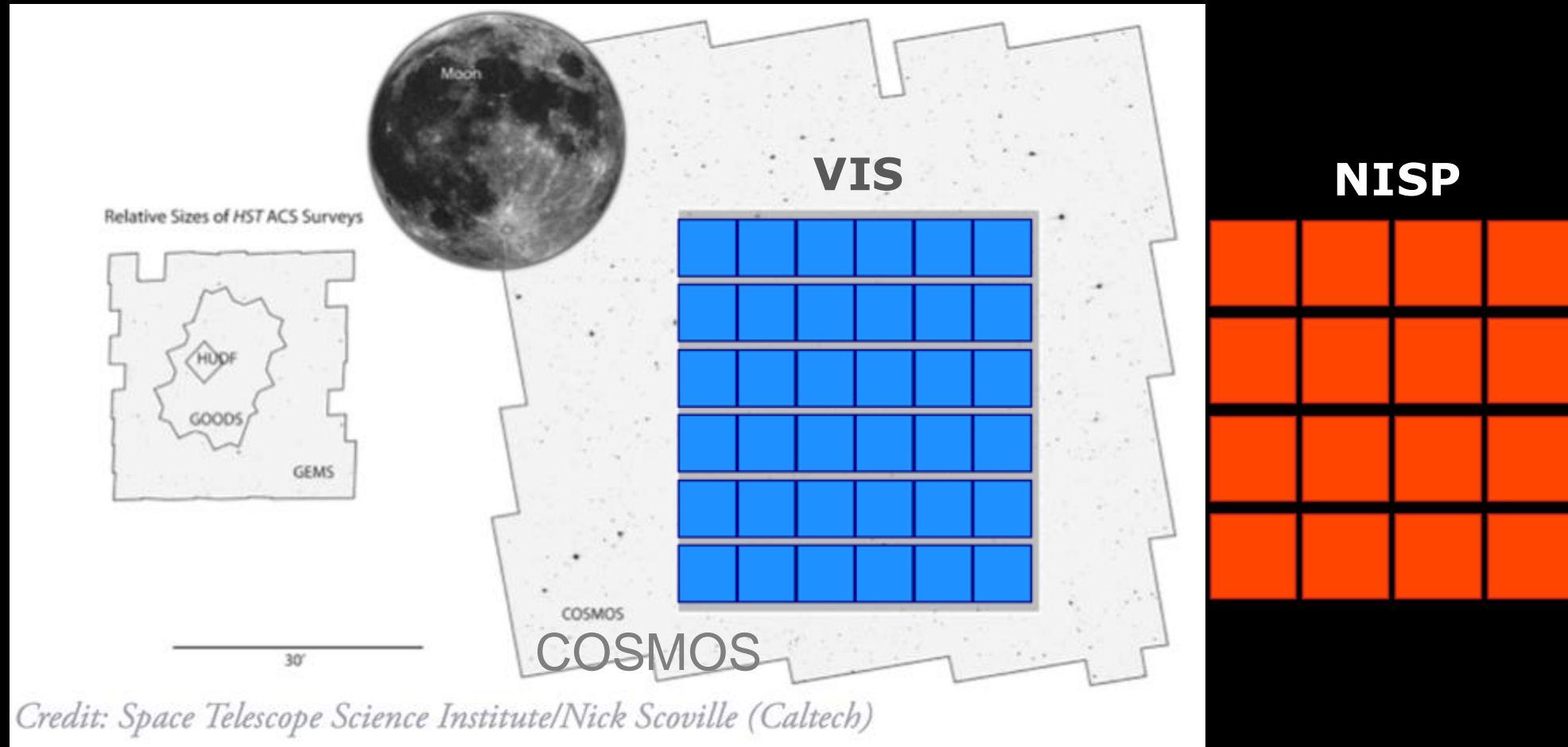


Euclid in Cannes / France, February 2023
Credit: ESA / M. Pedoussaut

Instruments



Twin wide-field imagers and NIR spectrograph

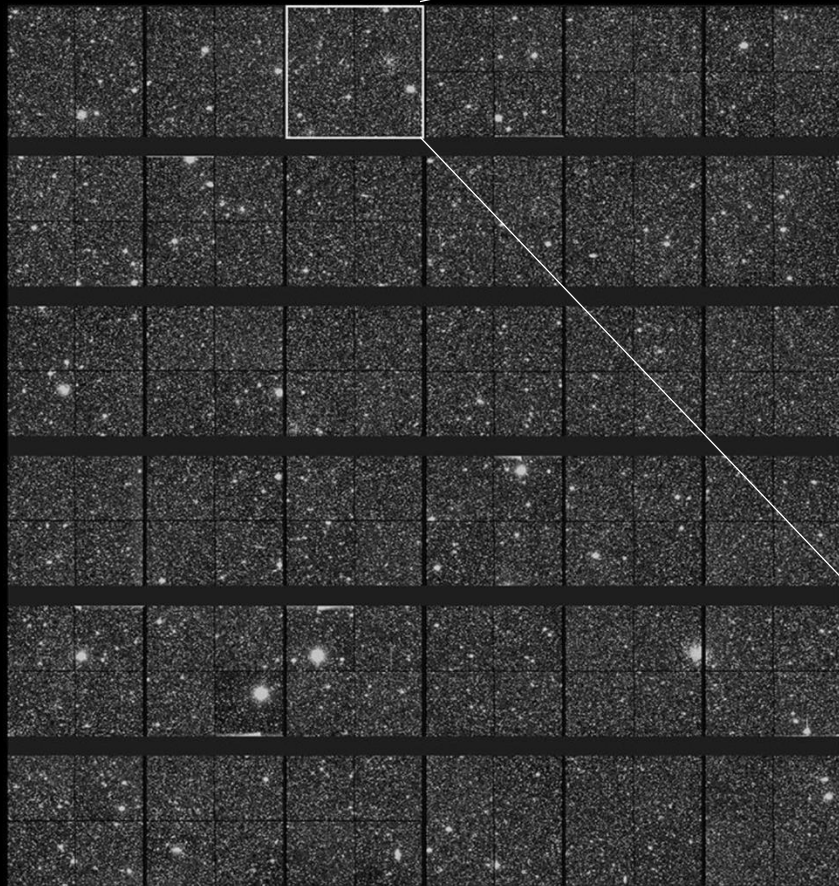


Euclid is the first panoramic space telescope ever: $10 \text{ deg}^2 / \text{day}$ in the Wide survey
Euclid field = 0.54 deg^2

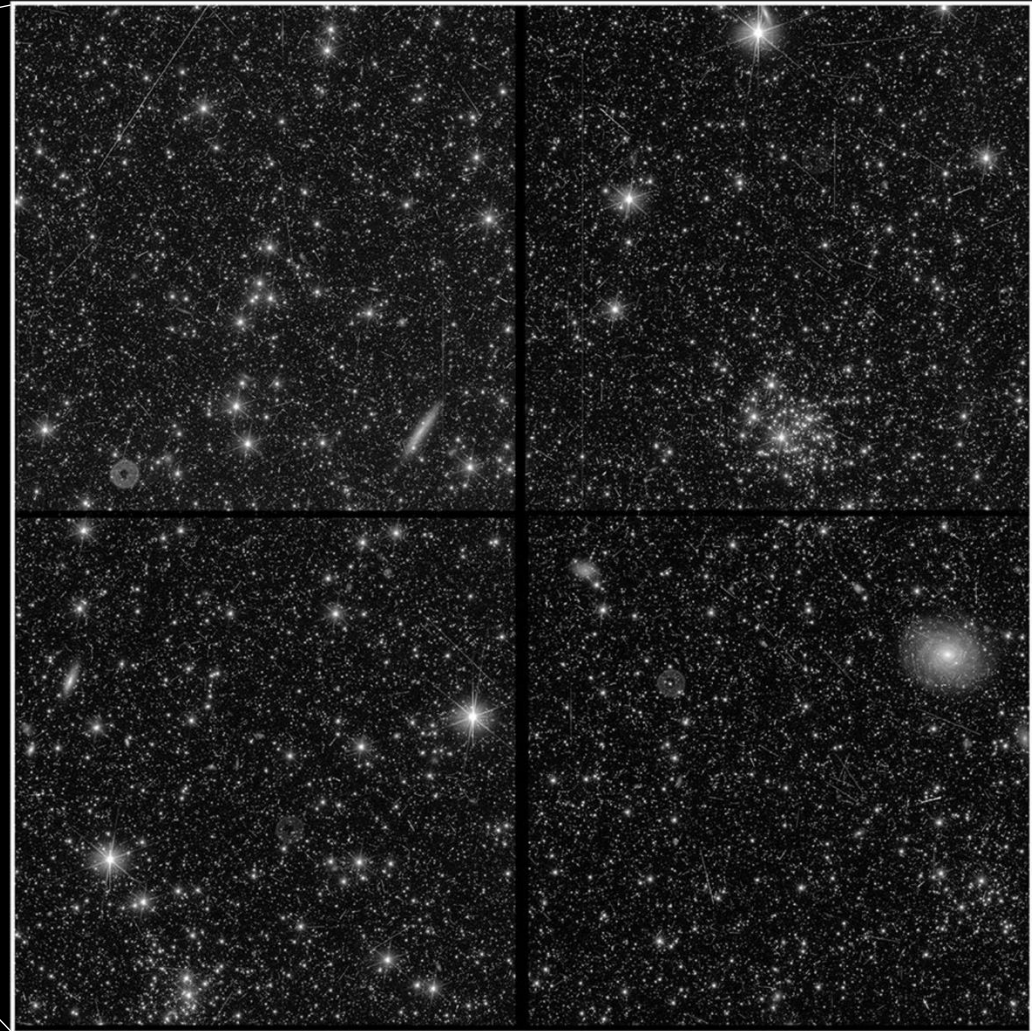
VIS instrument

Cropper et al. 2024

EARLY COMMISSIONING TEST IMAGE, VIS INSTRUMENT



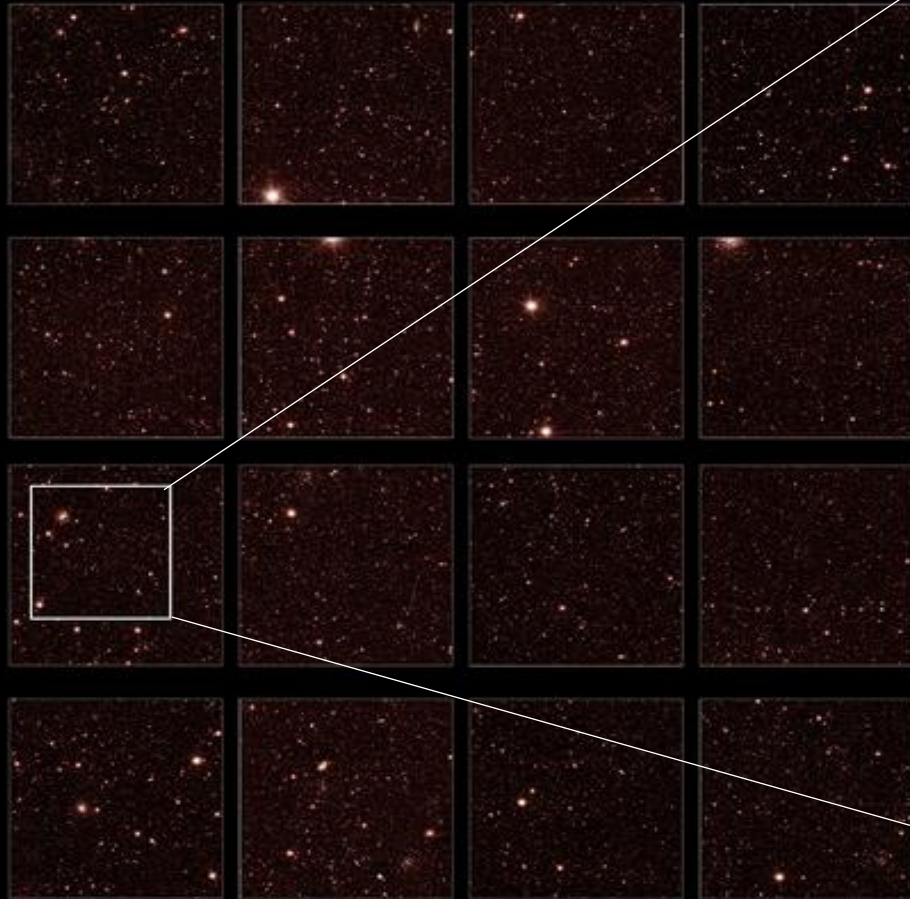
Credit : EC VIS team



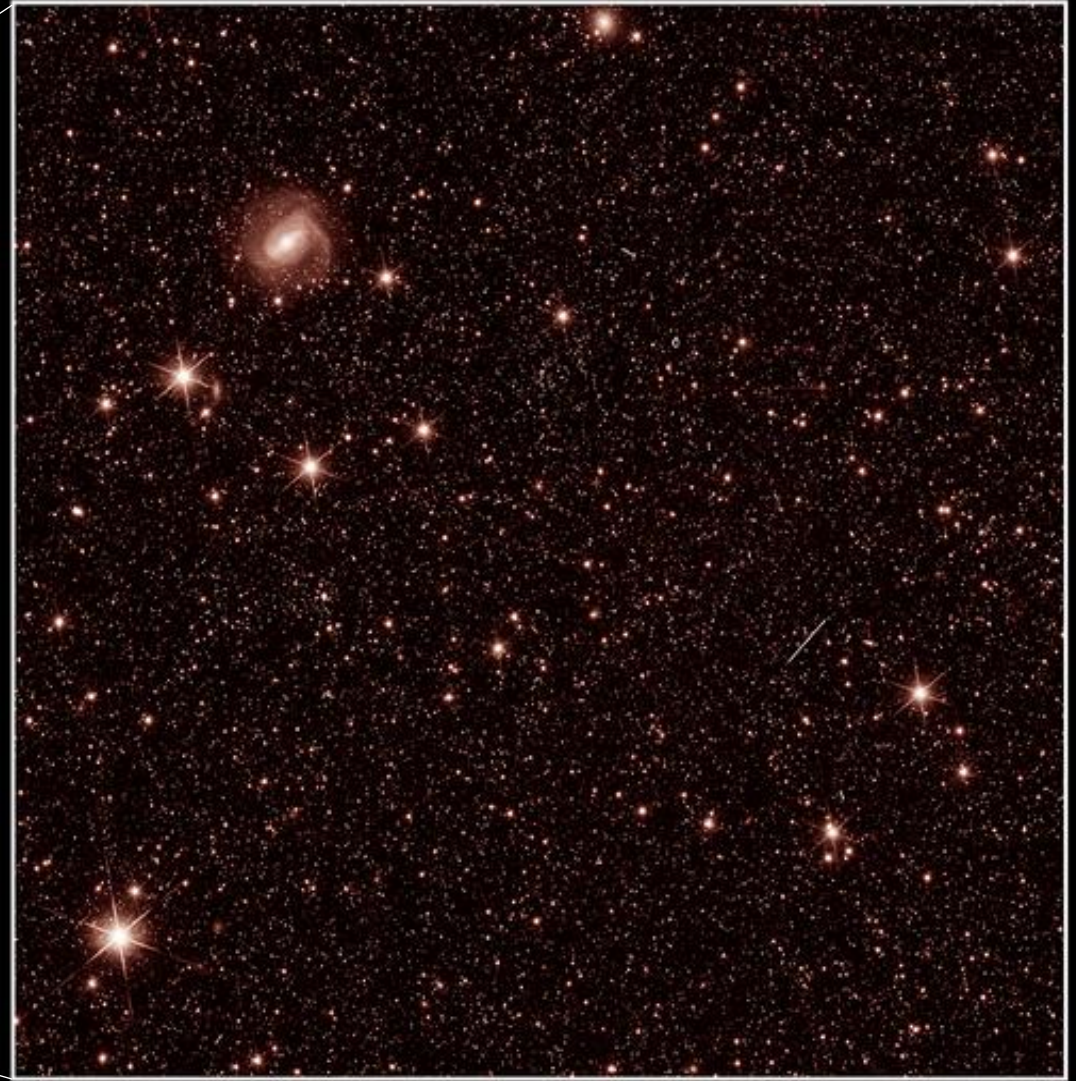
NISP Instrument

Jahnke et al. 2024

EARLY COMMISSIONING TEST IMAGE, NISP INSTRUMENT



Credit : EC NISP team

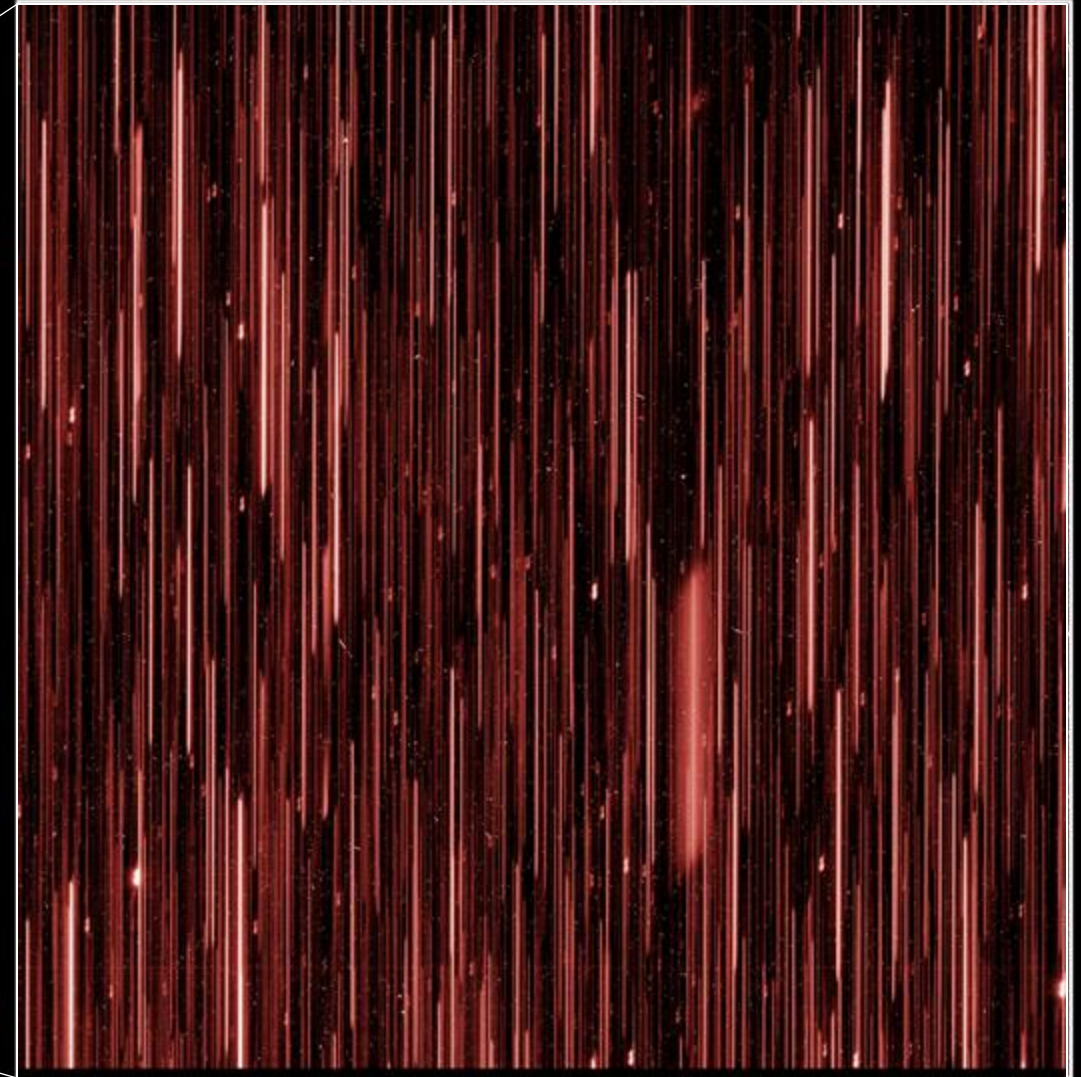
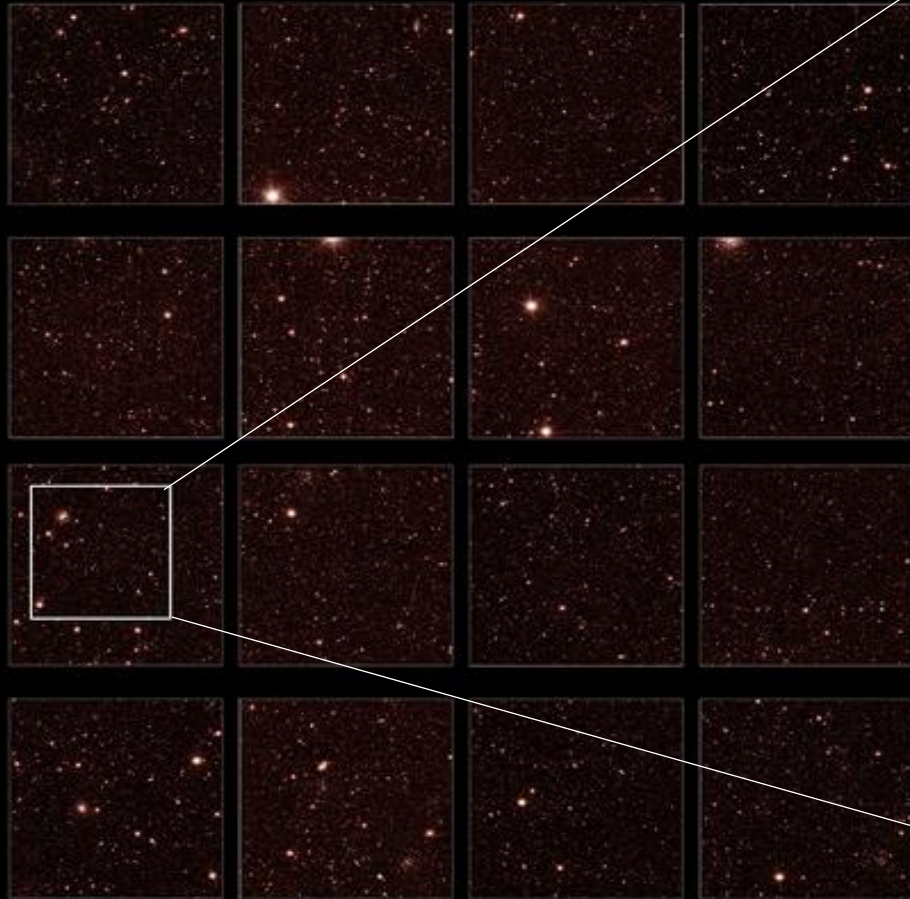


NISP Instrument

Jahnke et al. 2024

EARLY COMMISSIONING TEST IMAGE, NISP INSTRUMENT

Credit : EC NISP team



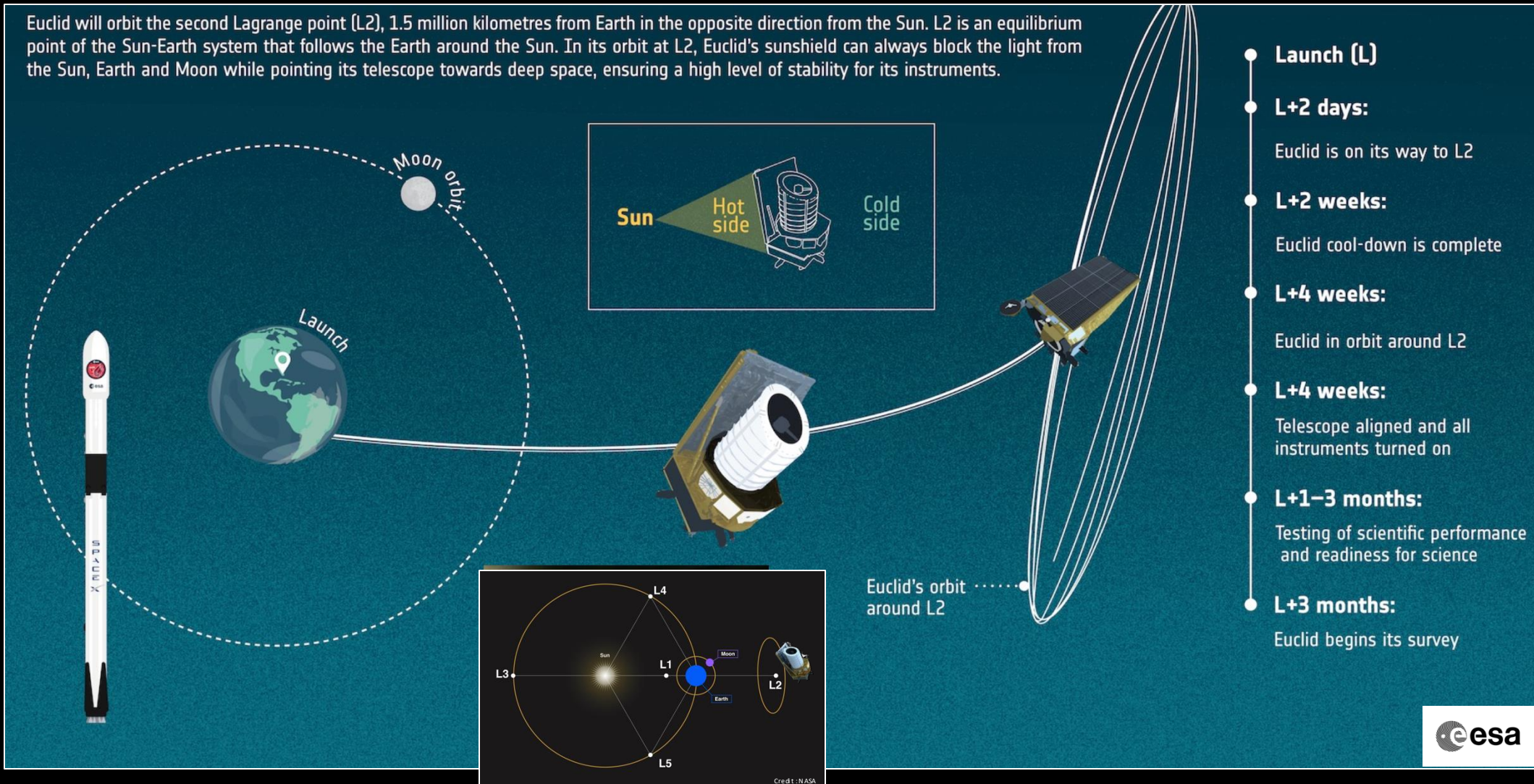
Euclid launch: 1st July 2023



Euclid launched on a Falcon 9 on 1st July 2023, from Cape Canaveral. (Credits: ESA, NASA & Space-X)

Euclid journey to L2

Euclid will orbit the second Lagrange point (L2), 1.5 million kilometres from Earth in the opposite direction from the Sun. L2 is an equilibrium point of the Sun-Earth system that follows the Earth around the Sun. In its orbit at L2, Euclid's sunshield can always block the light from the Sun, Earth and Moon while pointing its telescope towards deep space, ensuring a high level of stability for its instruments.



Early Release Observations (ERO)



Euclid on sky

Euclid consortium “On Sky” paper release:

www.euclid-ec.org/science/publications/

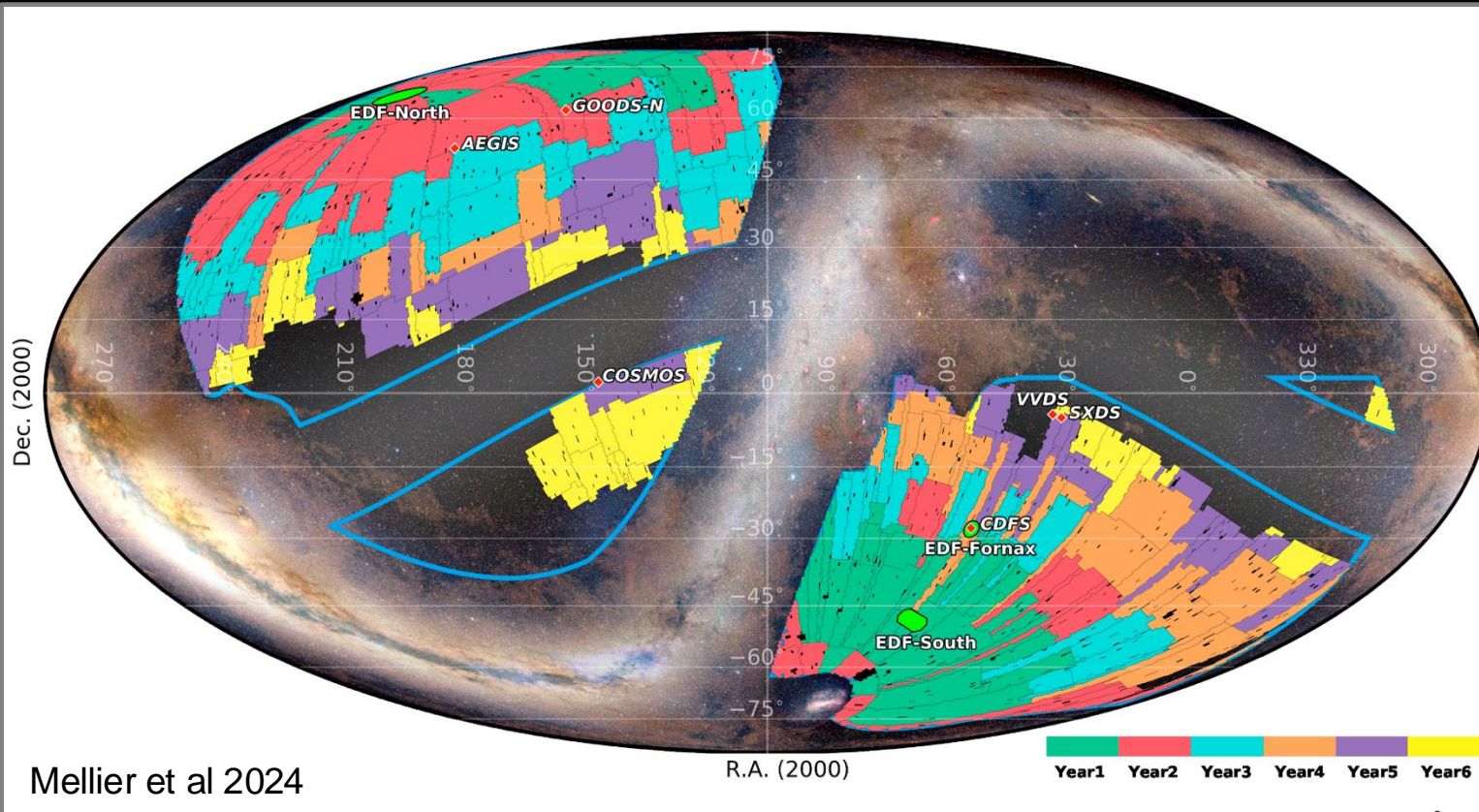
- 5 survey and instrument reference papers
- 10 papers on early release observations (ERO) results



Coordinated Release "Euclid on Sky," May 2024

- Euclid Collaboration: Mellier et al., 2024, "*Euclid. I. Overview of the Euclid mission*"
- Euclid Collaboration: Cropper et al., 2024, "*Euclid. II. The VIS Instrument*"
- Euclid Collaboration: Jahnke et al., 2024, "*Euclid. III. The NISP Instrument*"
- Euclid Collaboration: Hormuth et al., 2024, "*Euclid. IV. The NISP calibration unit*"
- Euclid Collaboration: Castander et al., 2024, "*Euclid. V. The Flagship galaxy mock catalogue: a comprehensive simulation for the Euclid mission*"
- Cuillandre et al., 2024, "*Euclid: Early Release Observations - Programme overview and pipeline for compact- and diffuse-emission photometry*"
- Martin et al., 2024, "*Euclid: Early Release Observations -- A glance at free-floating new-born planets in the σ Orionis cluster*"
- Massari et al., 2024, "*Euclid: Early Release Observations -- Unveiling the morphology of two Milky Way globular clusters out to their periphery*"
- Hunt et al., 2024, "*Euclid: Early Release Observations -- Deep anatomy of nearby galaxies*"
- Saifollahi et al., 2024, "*Euclid: Early Release Observations -- Globular clusters in the Fornax galaxy cluster, from dwarf galaxies to the intracluster field*"
- Cuillandre et al., 2024, "*Euclid: Early Release Observations - Overview of the Perseus cluster and analysis of its luminosity and stellar mass functions*"
- Marleau et al., 2024, "*Euclid: Early Release Observations -- Dwarf galaxies in the Perseus galaxy cluster*"
- Kluge et al., 2024, "*Euclid: Early Release Observations -- The intracluster light and intracluster globular clusters of the Perseus cluster*"
- Atek et al., 2024, "*Euclid: Early Release Observations -- A preview of the Euclid era through a magnifying lens*"
- Weaver et al., 2024, "*Euclid: Early Release Observations -- NISP-only sources and the search for luminous $z=6-8$ galaxies*"

Euclid reference survey



Euclid Wide Survey

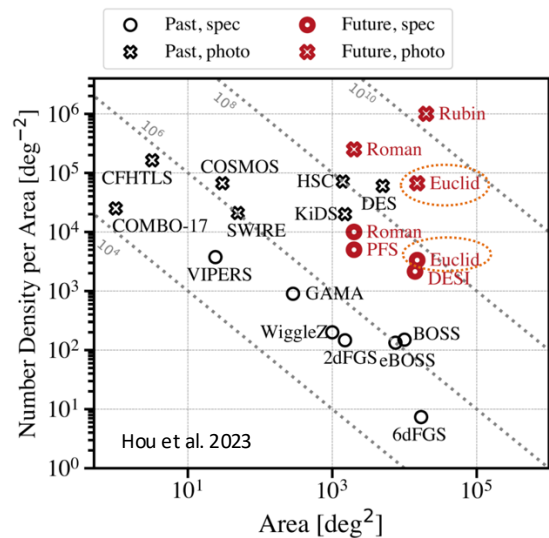
- 14 000 deg²
- 1.5 billion galaxies with photometric redshift and shape measurement ($n_g = 30$ gal/arcmin²)
- 30 million of galaxies with spectroscopic redshift

Euclid Deep Survey

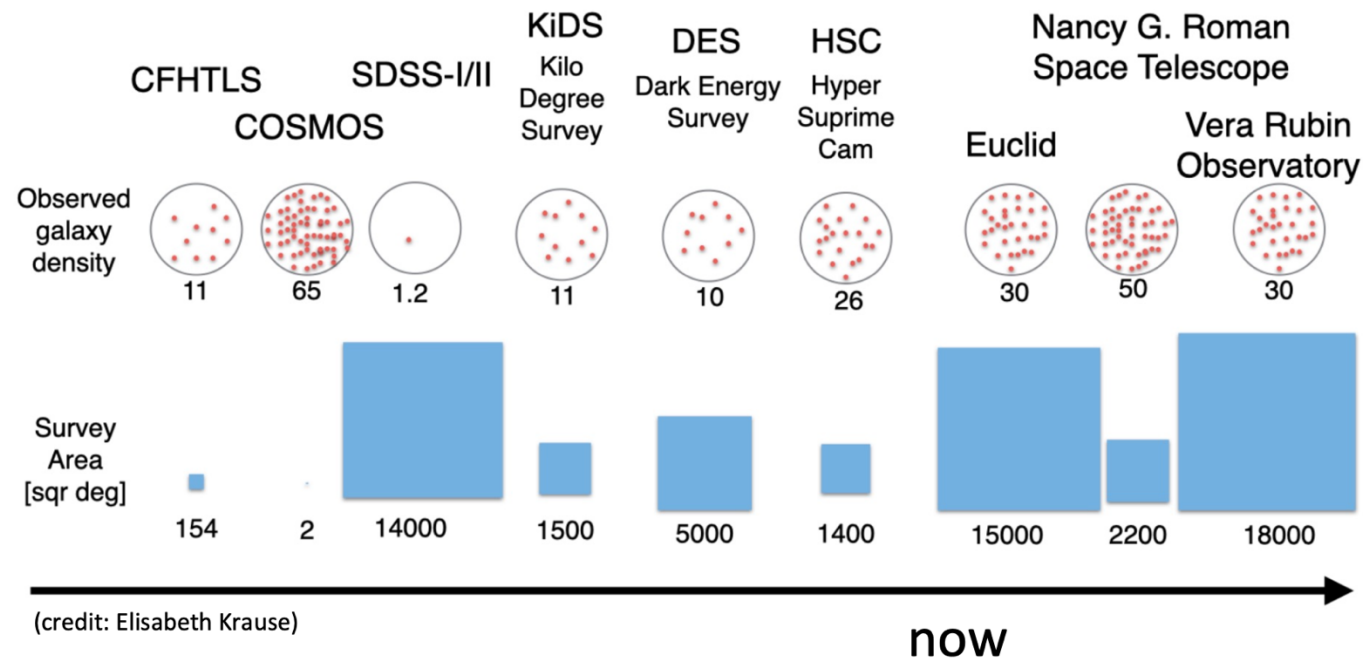
- Expected galaxy density: $n_g = 50$ gal/arcmin²
- 3 fields
 - EDF-N : 20 deg²
 - EDF-S : 23 deg²
 - EDF-F : 10 deg²

- Duration: 6 years survey
- Euclid will observe 10 deg²/day of the Euclid Wide Survey
- 12% of Euclid time is spent on Euclid Deep Survey

Eculid Wide Survey in context

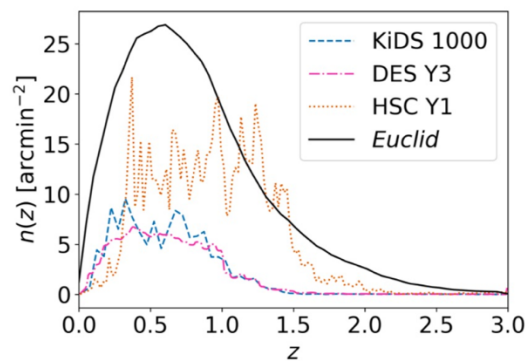


Euclid WL vs. Stage III & IV experiments



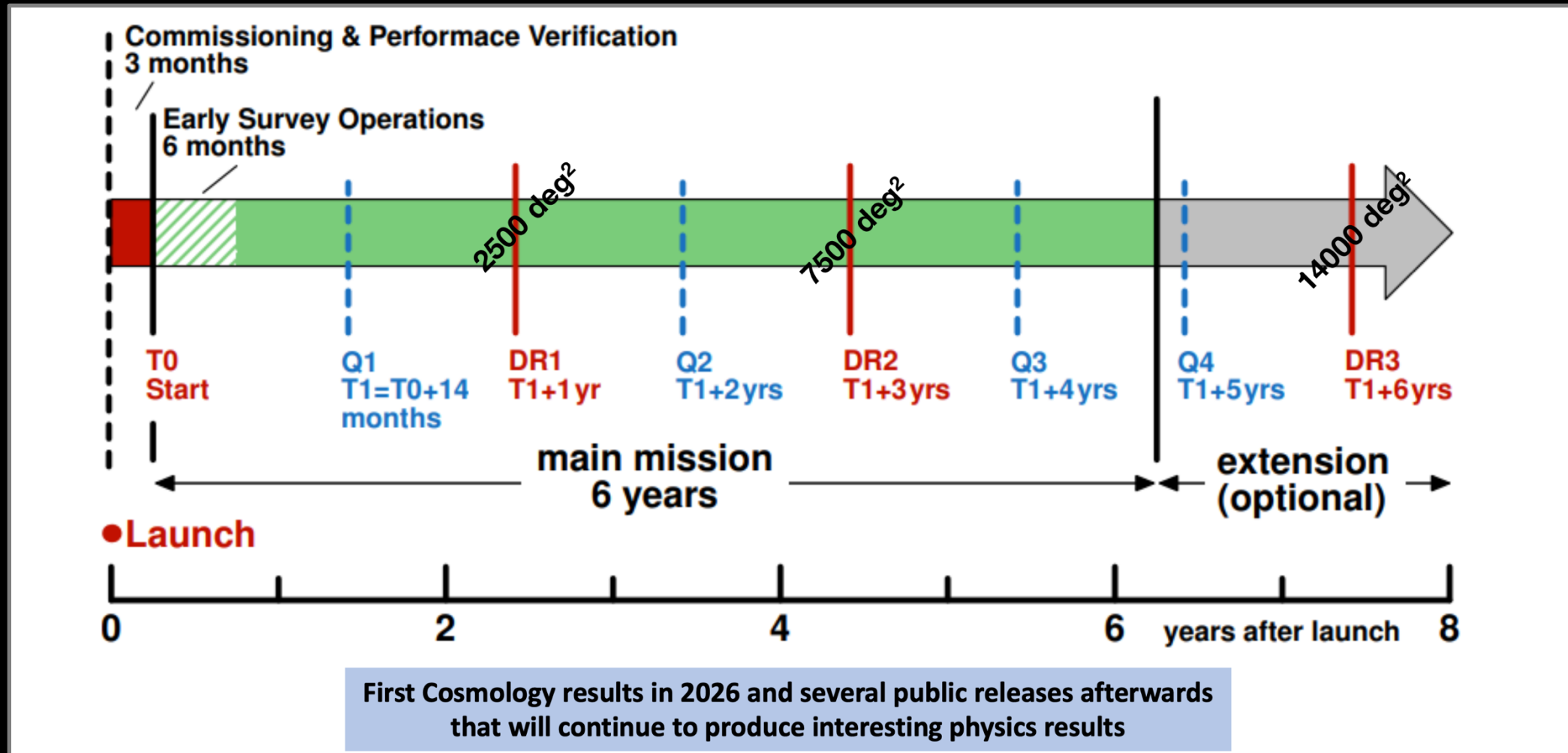
(credit: Elisabeth Krause)

now



Unprecedented volume coverage with objects up to redshift 2

Euclid Data Releases



Cosmological inference



Theoretical Predictions

COMPUTE EUCLID PRIMARY OBSERVABLES ACCORDING TO THE THEORY



Non-linear cosmology

MODEL NON-LINEAR SCALES FOR POWER SPECTRA



Euclid data as input

DATA VECTORS ARE PROVIDED BY THE EUCLID SCIENCE GROUND SEGMENT



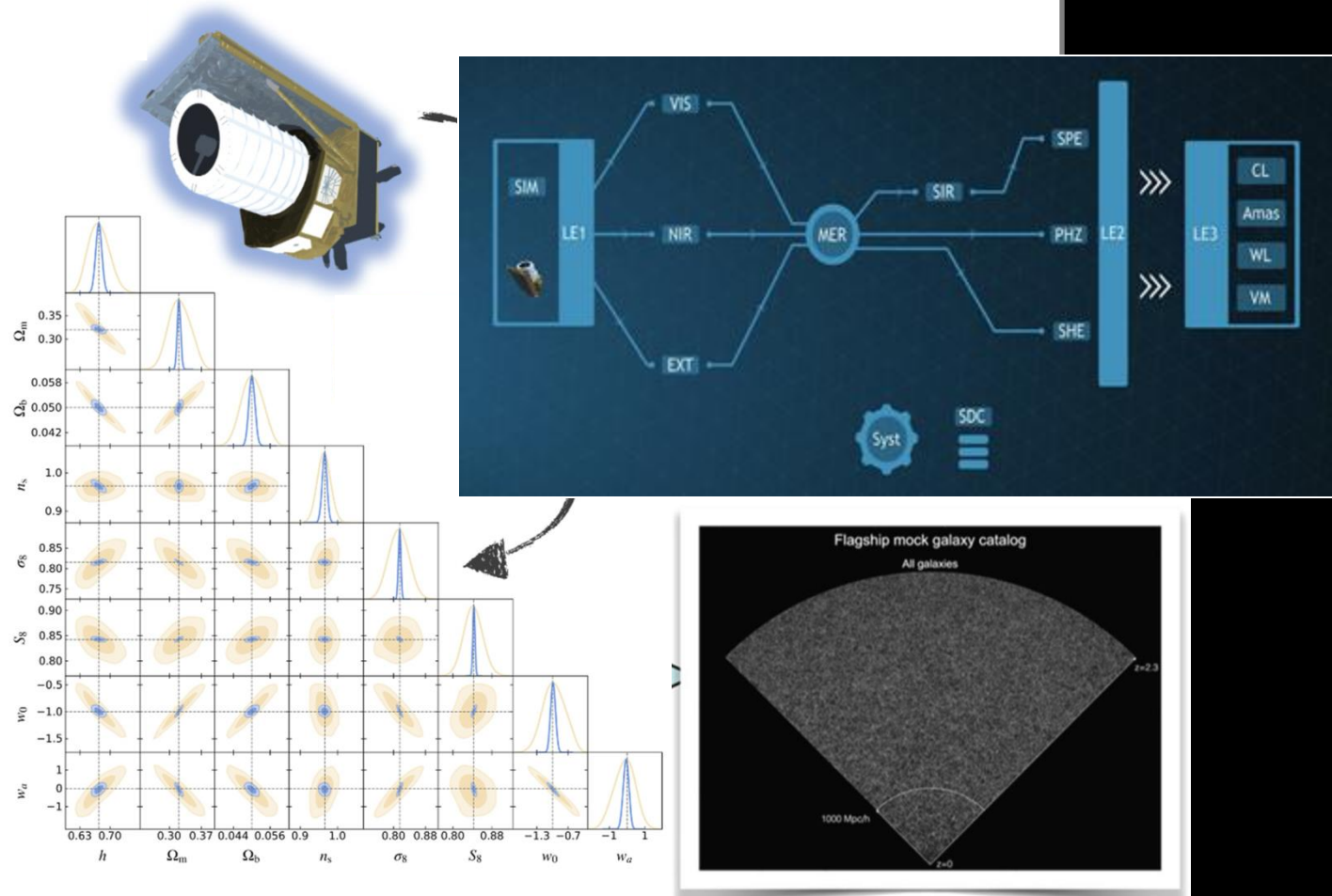
Likelihood computation

WE COMPARE THEORY AGAINST DATA DOING A STATISTICAL ANALYSIS



Cosmological parameters

PRODUCE THE BAYESIAN STATISTICAL ANALYSIS TO OBTAIN CONSTRAINTS ON COSMOLOGICAL PARAMETERS

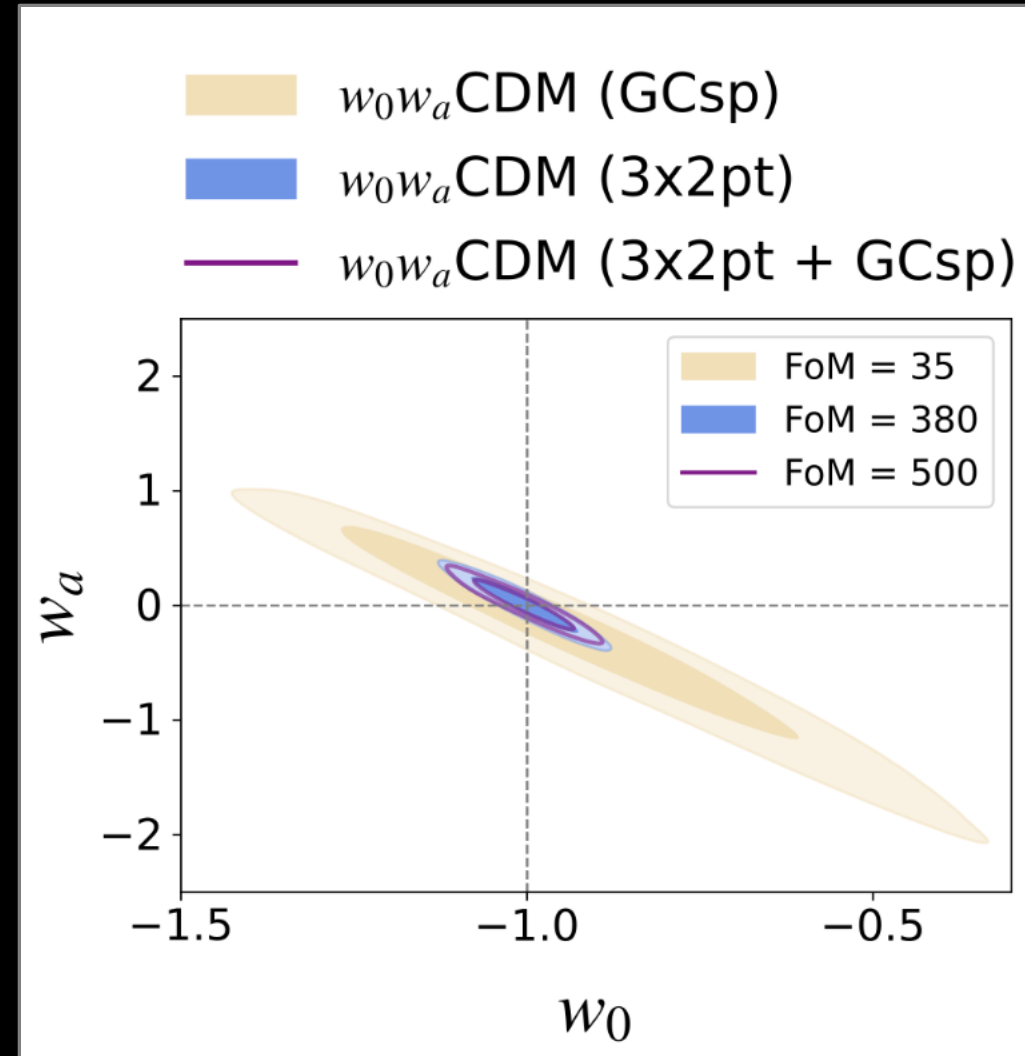


Euclid: the physics of cosmic acceleration

The dark energy equation of state $w \neq -1$ indicates a departure from the cosmological constant, requiring a dynamical dark energy scenario.

Parameterised evolution of $w(z)$:

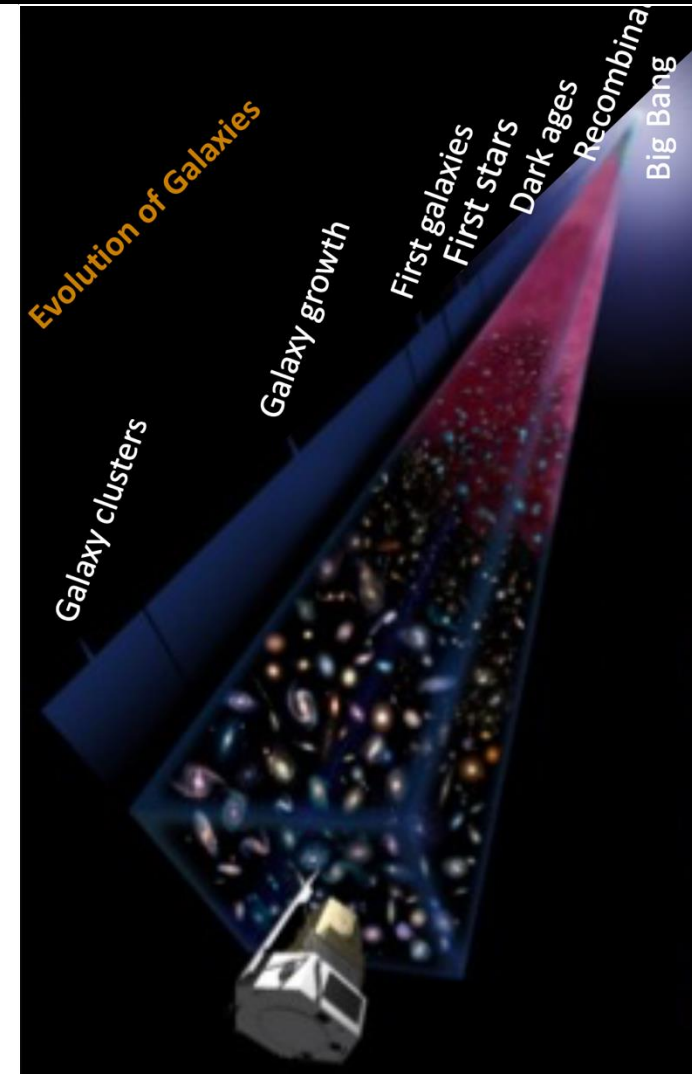
$$w(z) = w_0 + w_a \frac{z}{1+z}$$



Euclid Additional Science

- 10^5 galaxy clusters
- **Cosmic Voids**
- **Cross-correlations with CMB** temperature and lensing
- 10^5 **strong gravitational lenses**
- Transients in Deep fields
 - ~ 50 **Super-luminous SNe** / year (Inserra+17)
- Galaxy formation and evolution
 - Census of **AGN** at $1 < z < 3$
 - Galaxy **morphologies** at $z > 1$
 - **Lyman break galaxies** at $z > 7$
 - High- z **quasars**
- Milky Way
 - Census of **brown dwarf** stars
 - **Satellites & environs**

x10 in number of sources compared to previous surveys in most cases



Outlook

Euclid launched successfully from Cape Canaveral on July 1st 2023 with Space-X Falcon9: very efficient orbit insertion into L2, >10 yrs lifetime expected

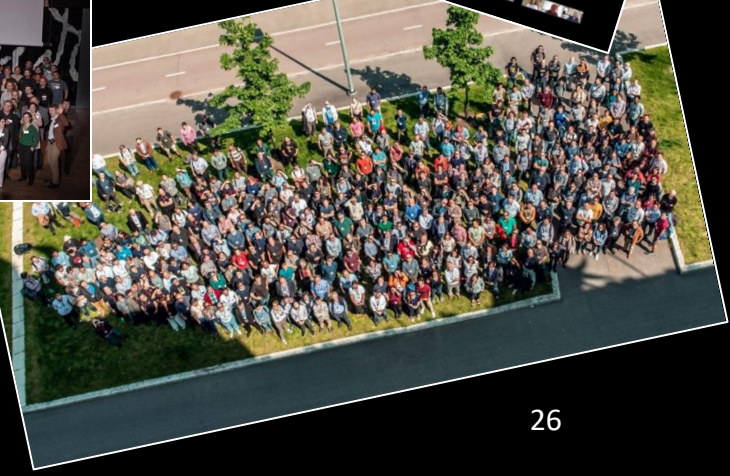
Unique optical quality telescope: a terapixel machine. Early issues (guiding system, stray light and ice deposition) solved. X-ray flares are handled successfully on a regular basis

Currently surveying 10 deg² per day: goal is to map one third of the sky (14000 deg²) and unveil the nature of DM and DE by measuring weak gravitational lensing and galaxy clustering to exquisite precision and unmatched accuracy (probe combination)

Immense potential for legacy science: Early Release Observations provide a glimpse of the unprecedented information content of even a single Euclid snapshot.

Results will provide high precision results on the equation of state of dark energy and other cosmological parameters.

First data release after one year of operations (2500 deg², mid 2026)





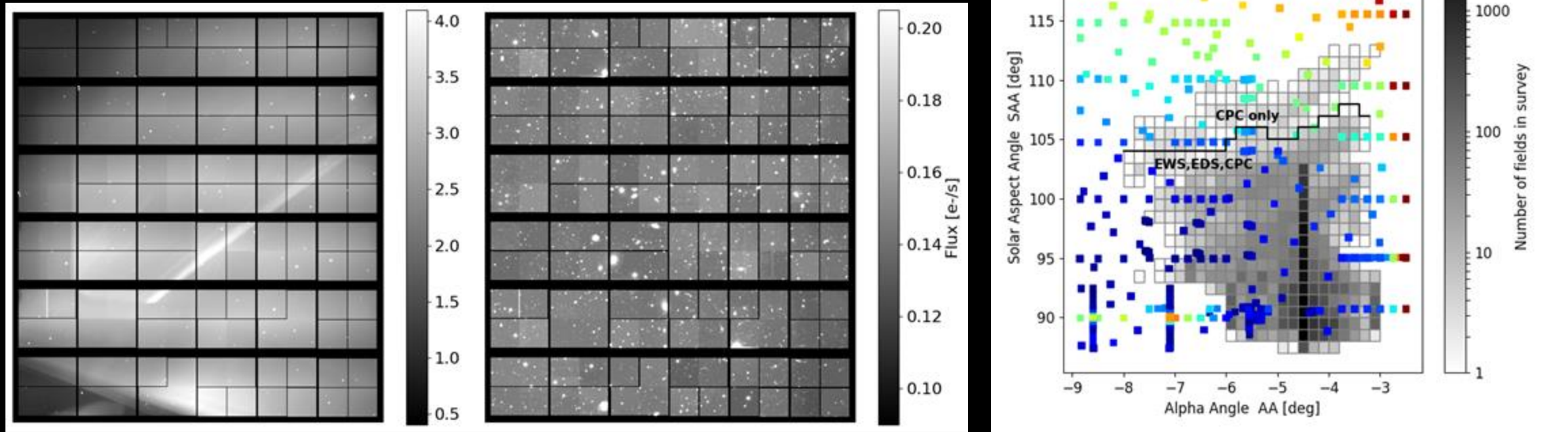
euclid
consortium

EXTRA SLIDES

Internal straylight



Mellier et al 2024



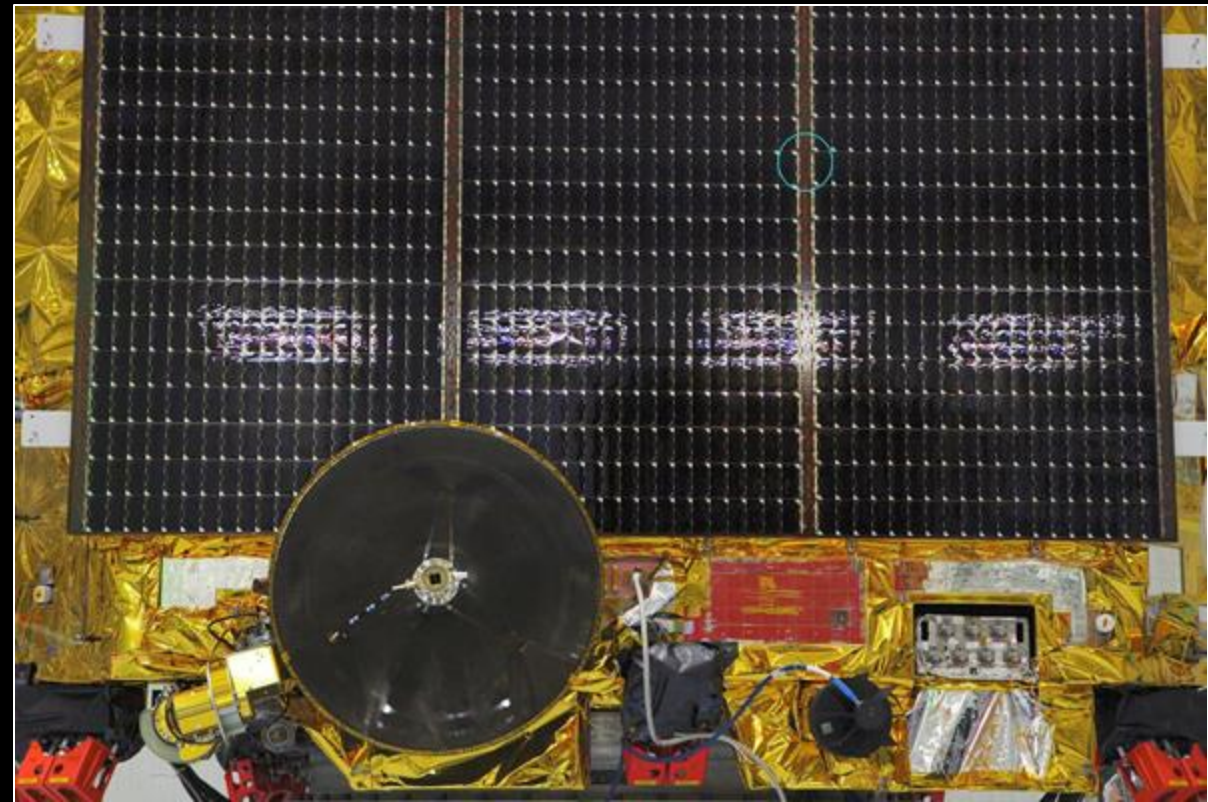
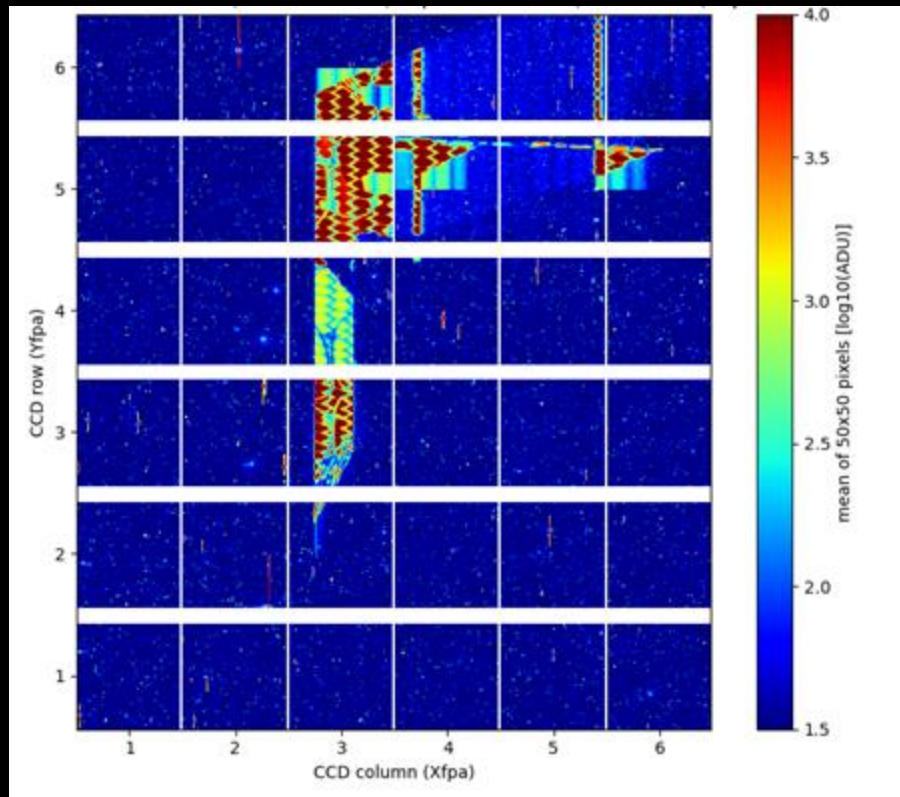
Left: Sunlight hitting a thruster nozzle reaches VIS detectors. NISP is unaffected.

Middle: Straylight is largely avoided by orienting Euclid so that nozzle is in shadow

Right: The survey was fine-tuned to select low-straylight conditions, only.

X-ray contamination

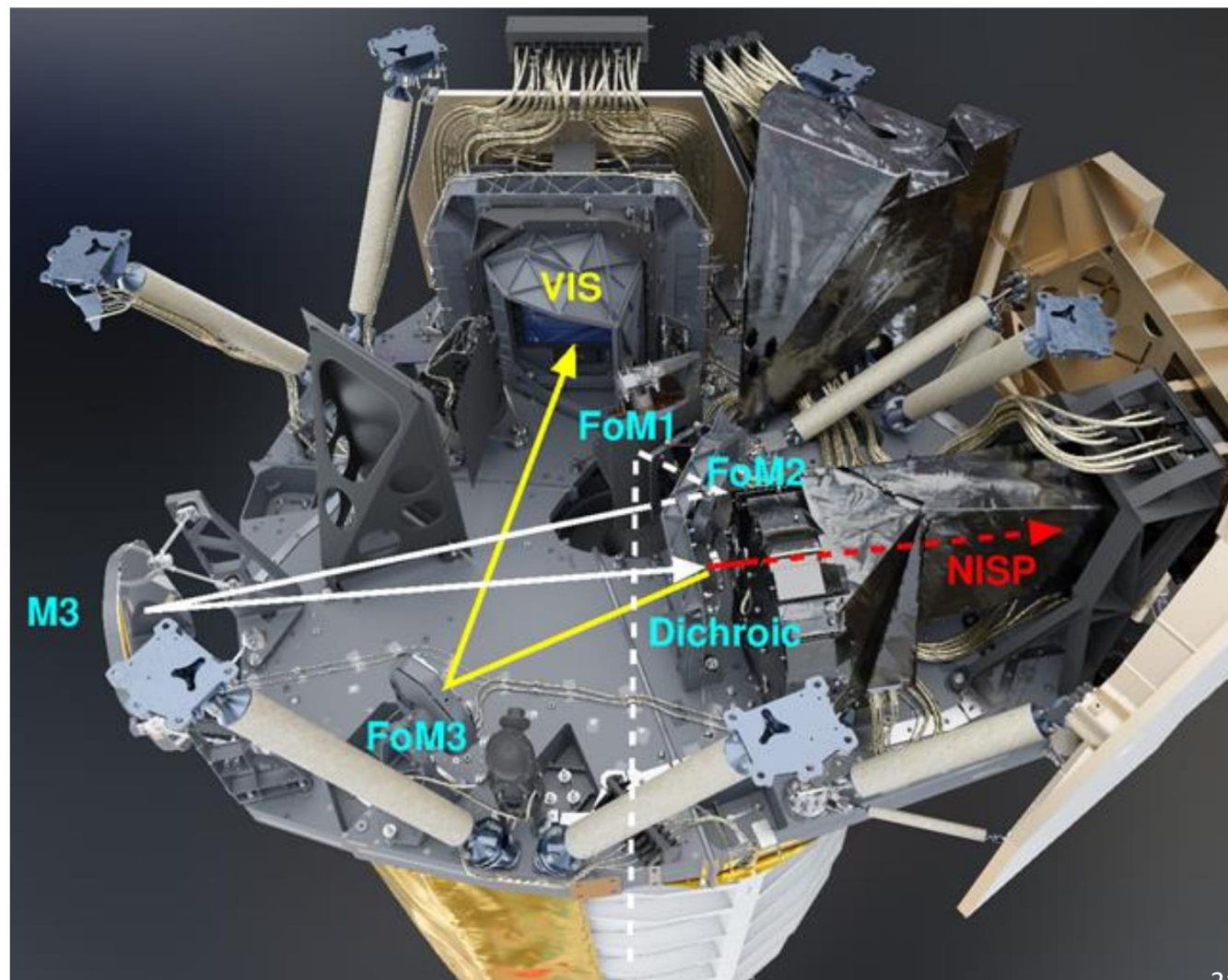
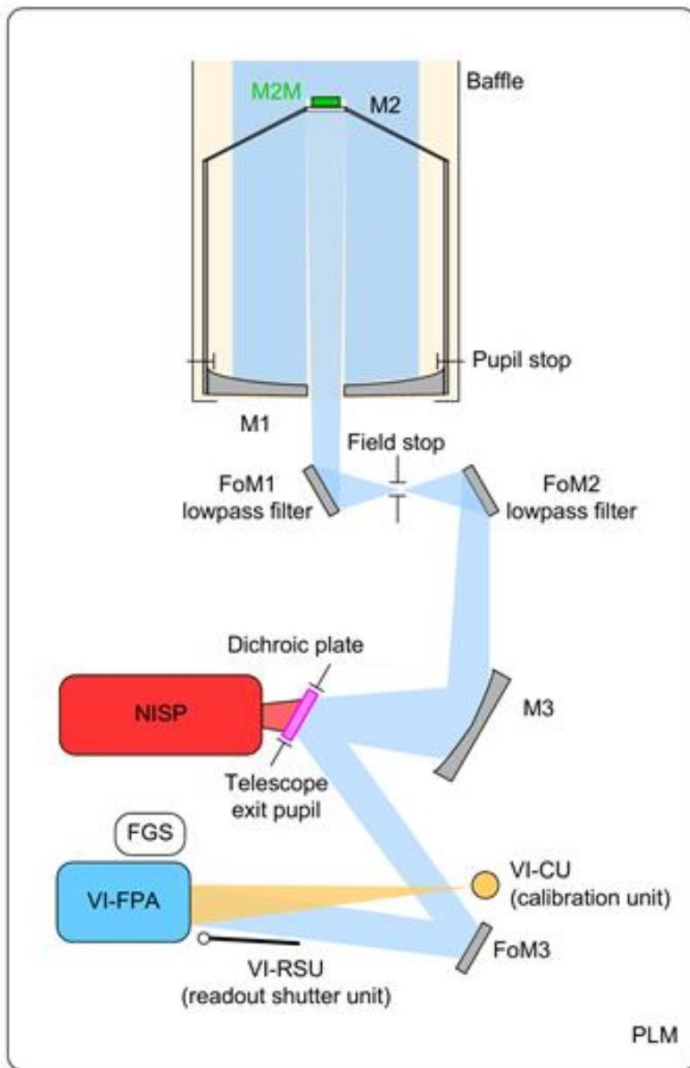
X-rays from Solar flares penetrate sunshield and reach VIS. Average area loss during Solar maximum: 3-4%



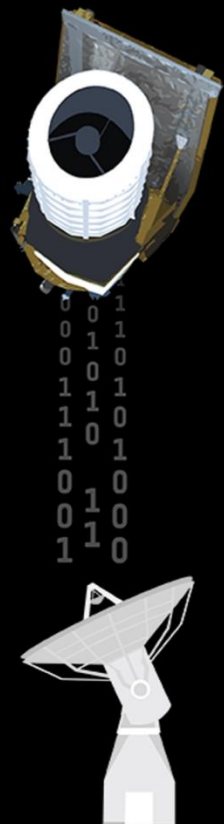
VIS image taken during an X-class flare. Weak flares cause isolated cosemics. Strong flares result in contiguous area loss.

X-rays enter through the major gaps between solar cells.

Payload module, instruments and path



Data Transfer



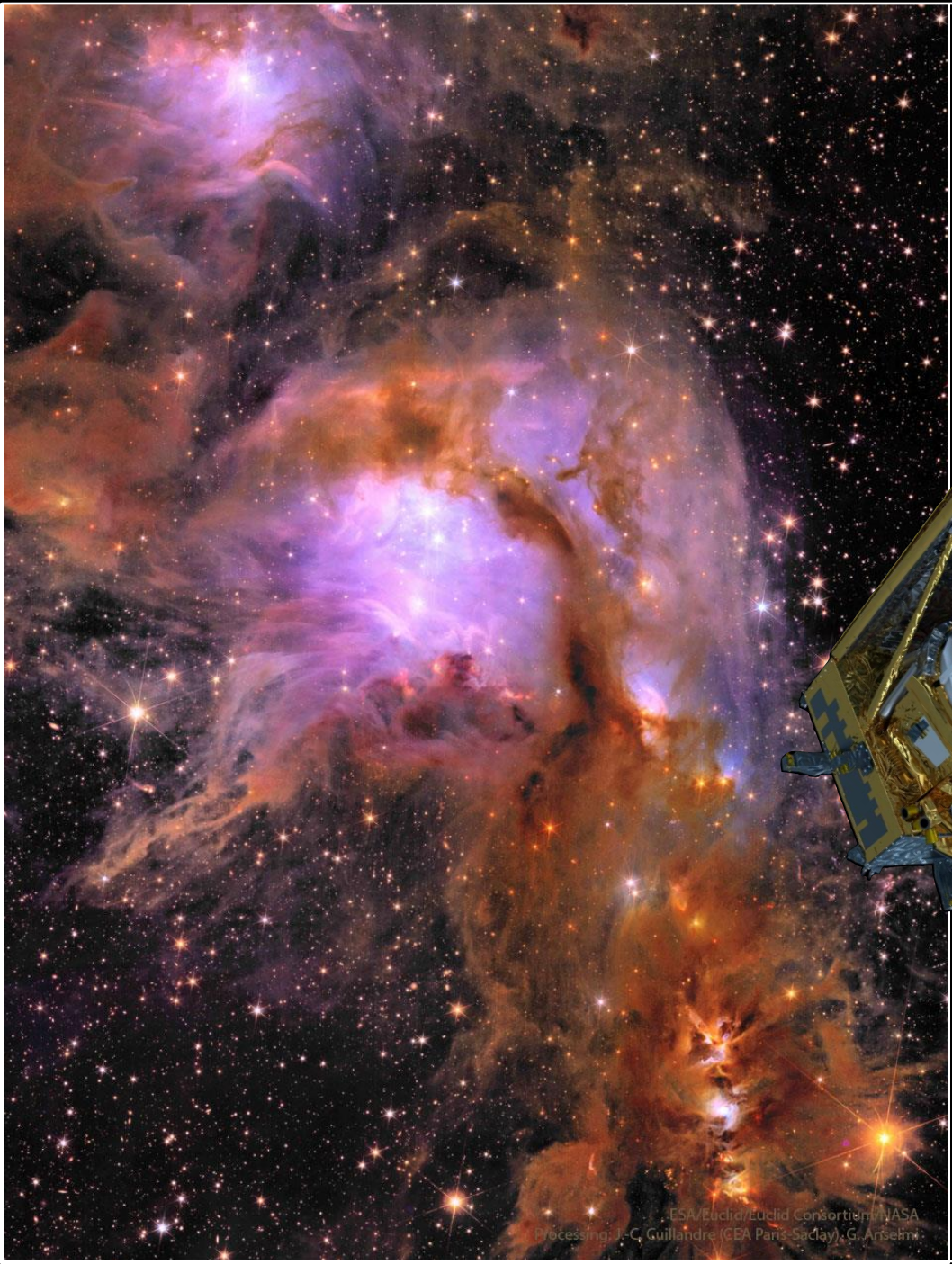
- Collecting 30 Petabytes of unprocessed data
- Communications from L2, to ESOC (Germany), to ESAC (Madrid), once per day
- Distribution along 9 dedicated data centres

Euclid Top Level Science Requirements

Dark Energy	Test of Gravity
<ul style="list-style-type: none"> • Measure the cosmic expansion history to better than 10% in redshift bins $0.9 < z < 1.8$. • Look for deviations from $w_0 = -1$, indicating dynamical Dark energy. • Euclid primary probes to give $\text{FoM}_{DE} > 400$ (1-sigma errors on w_0 and w_a of 0.02 and 0.1 respectively) 	<ul style="list-style-type: none"> • Measure the growth rate to better than 0.02 in redshift bins between $0.9 < z < 1.8$. • Measure the growth index, γ, with a precision better than 0.02. • Separately constrain the two relativistic potentials. ψ and ϕ. • Test the cosmological principle.
Dark Matter	Initial conditions
<ul style="list-style-type: none"> • Detect Dark matter halos on a mass scale $10^8 < M/M_\odot < 10^{15}$. • Measure the Dark matter mass profiles on cluster and galactic scales. • Measure the sum of neutrino masses with an accuracy of 0.03 eV. 	<ul style="list-style-type: none"> • Constraint σ_8 and n_s to a 1-sigma accuracy of 0.01. • For extended models, improve constraints on spectral indices compared to Planck alone by a factor ~ 2. • Measure non-Gaussianity: $\Delta f_{NL} = \pm 2$.

- *DE equation of state: $P/\rho = w$ with $w(a) = w_p + w_a(a_p - a)$*
- *Growth rate of structure formation: $f \sim \Omega^\gamma$;*
- *$\text{FoM} = 1/(\Delta w_a \times \Delta w_p) > 400 \rightarrow \sim 2\%$ precision on w_p*

Euclid Redbook 2011



ESA/Euclid/Euclid Consortium/WASA
Processing: J.-C. Guillemandre (CEA Paris-Saclay), G. Anselmi

THE END