

# J1048+7143: A SUPERMASSIVE BLACK HOLE BINARY CANDIDATE

RUB

Ilja Jaroschewski

Cosmic Rays and Neutrinos  
in the Multi-Messenger Era

Collaborators:

Emma Kun,  
Julia Becker Tjus,  
Silke Britzen,

...

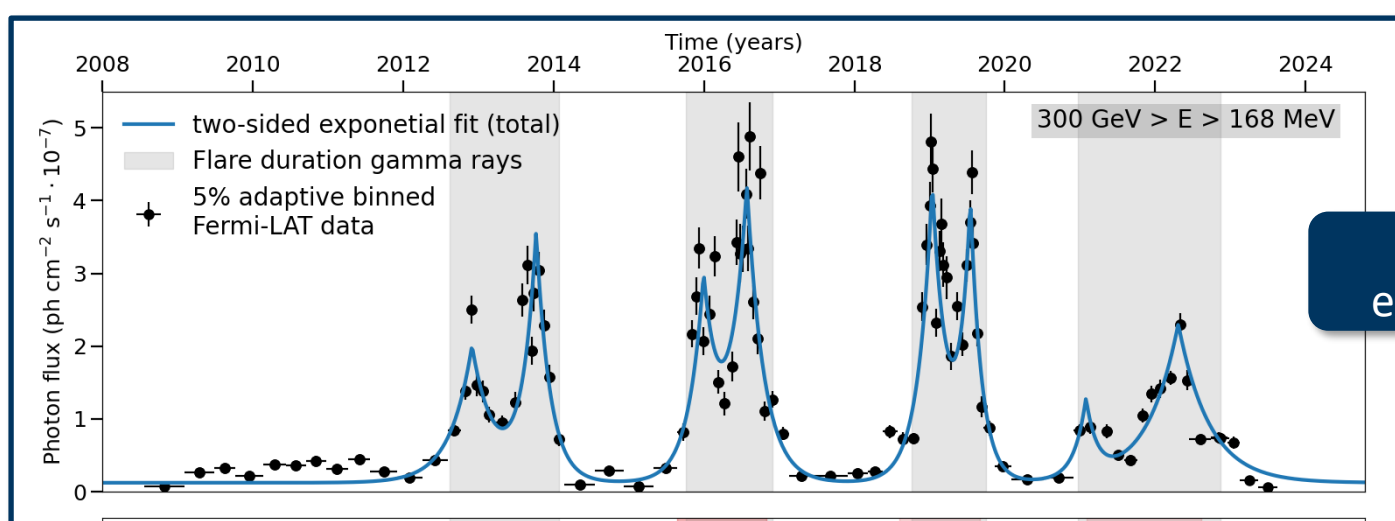
DFG

MICRO ● SFB1491

cea

irfu

# Gamma-Ray Light Curve



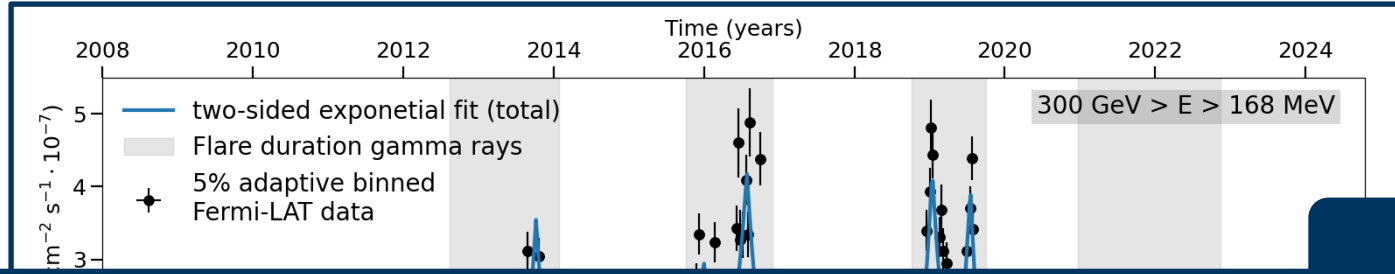
**J1048+7143**

Kun, **IJ+** in prep.

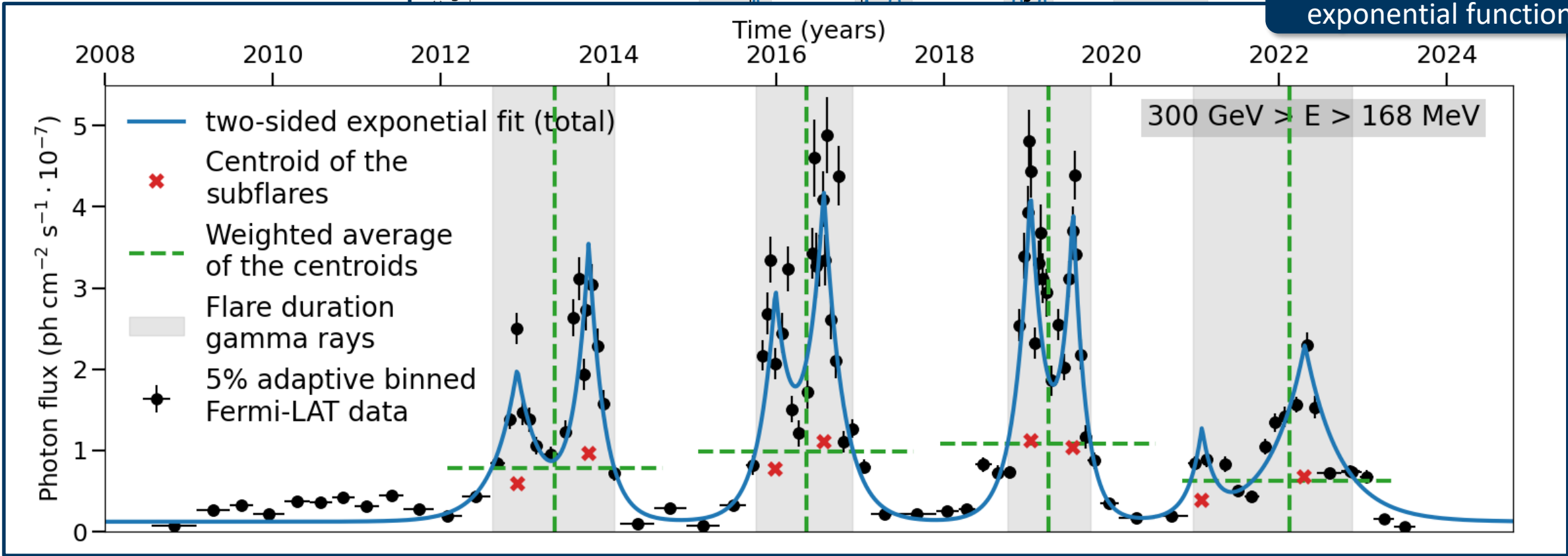
cea

irfu

# Gamma-Ray Light Curve



Fit with two-sided exponential function

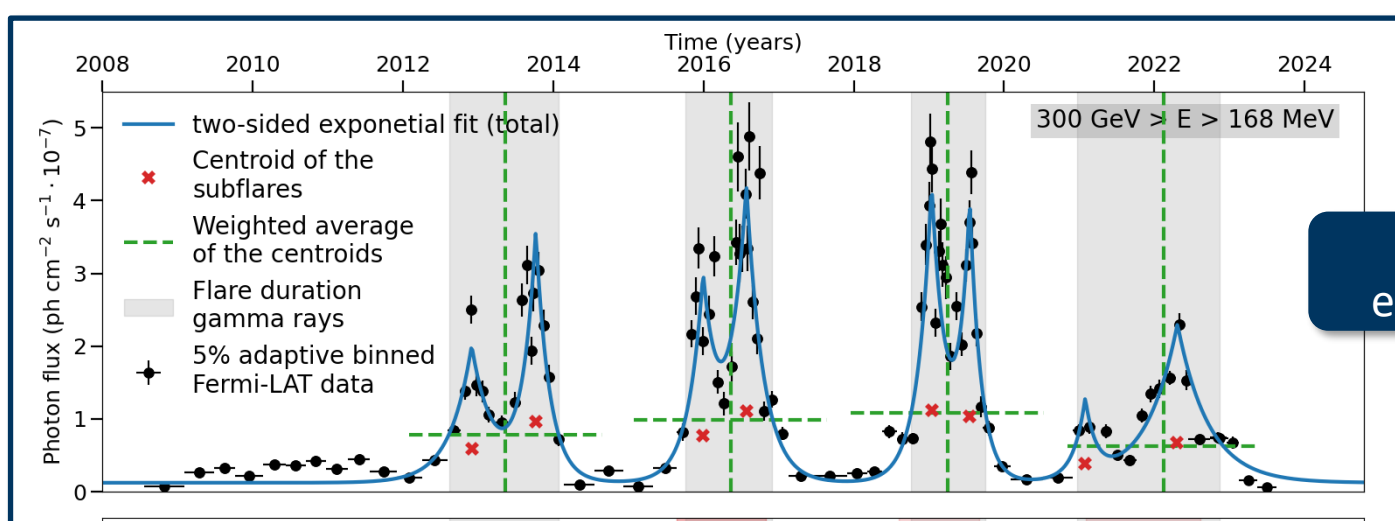


Kun, IJ+ in prep.

**J1048+7143**



# Gamma-Ray Light Curve + Optical + Radio



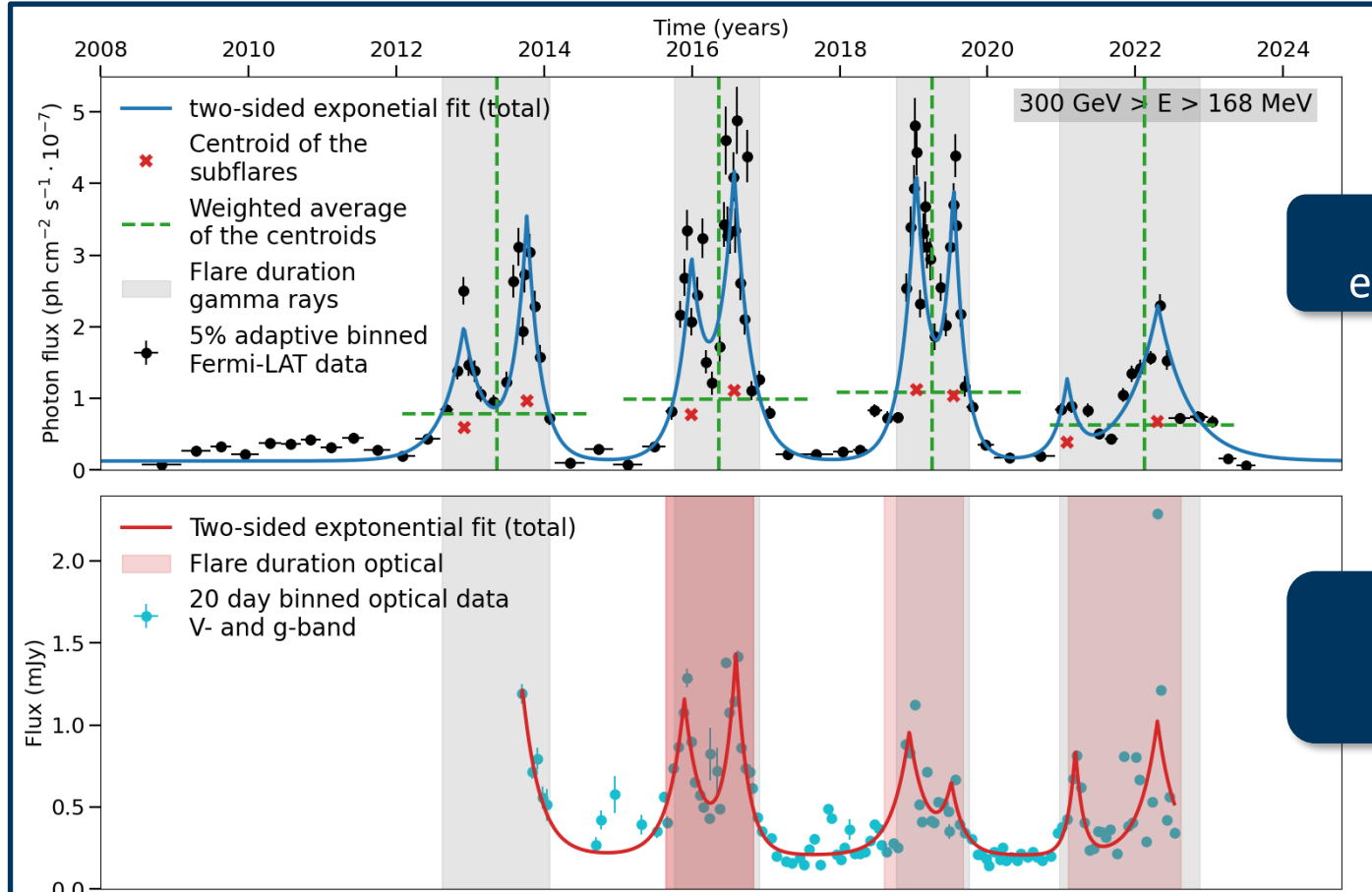
Fit with two-sided exponential function

J1048+7143

Kun, IJ+ in prep.



# Gamma-Ray Light Curve + Optical



Fit with two-sided exponential function

ASAS  
AAVSO  
ZTF in progress

Kun, IJ+ in prep.

**J1048+7143**



# Gamma-Ray Light Curve

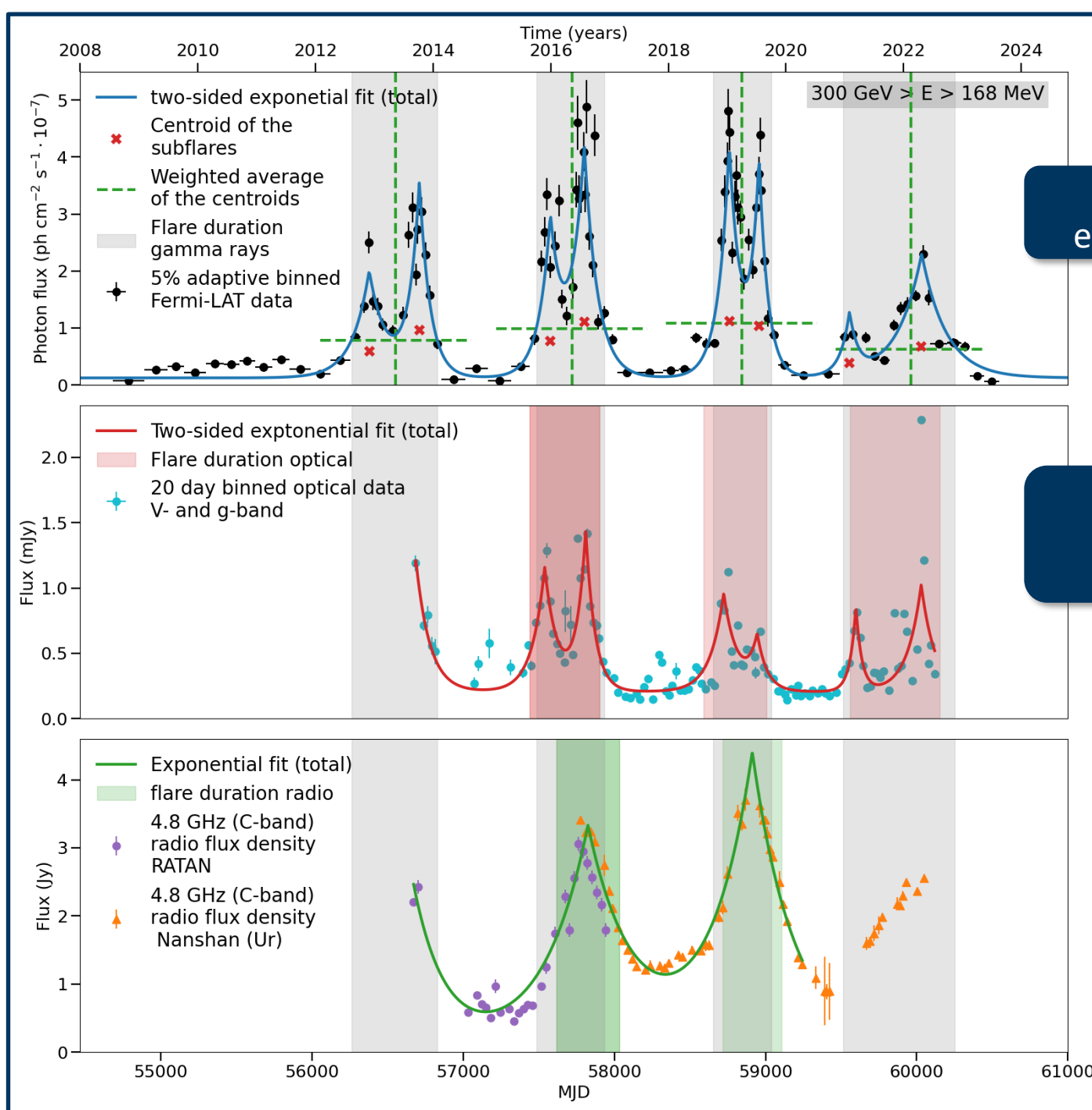
+ Optical

+ Radio

**J1048+7143**

1

Ilja Jaroschewski



Fit with two-sided exponential function

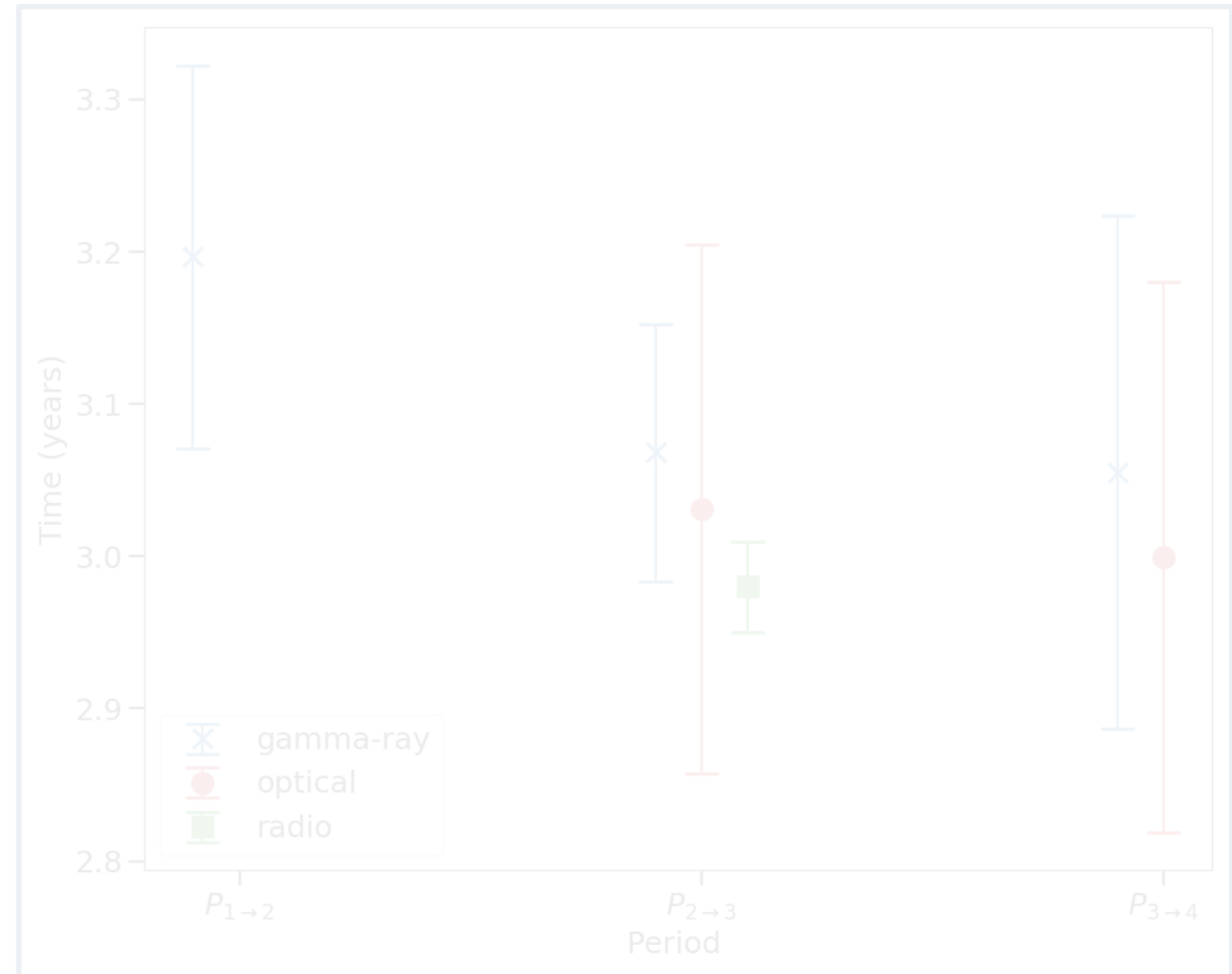
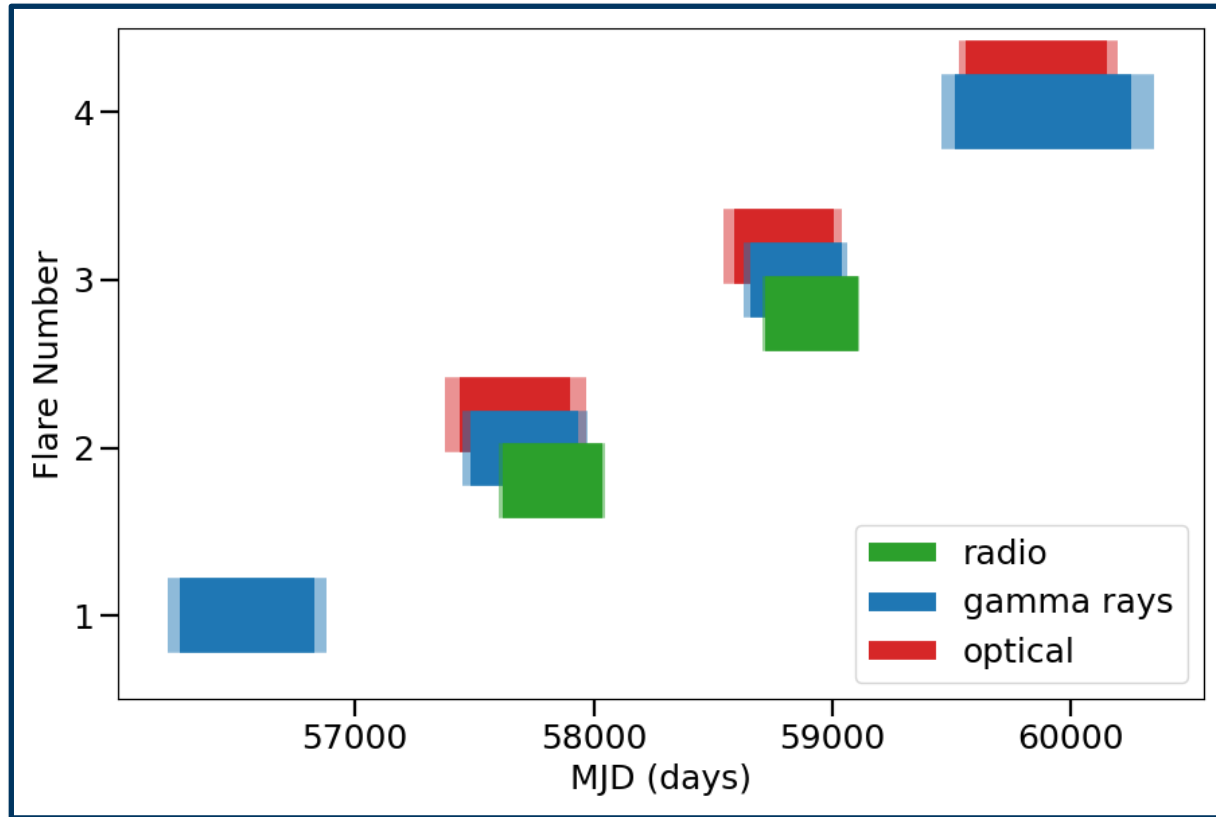
ASAS  
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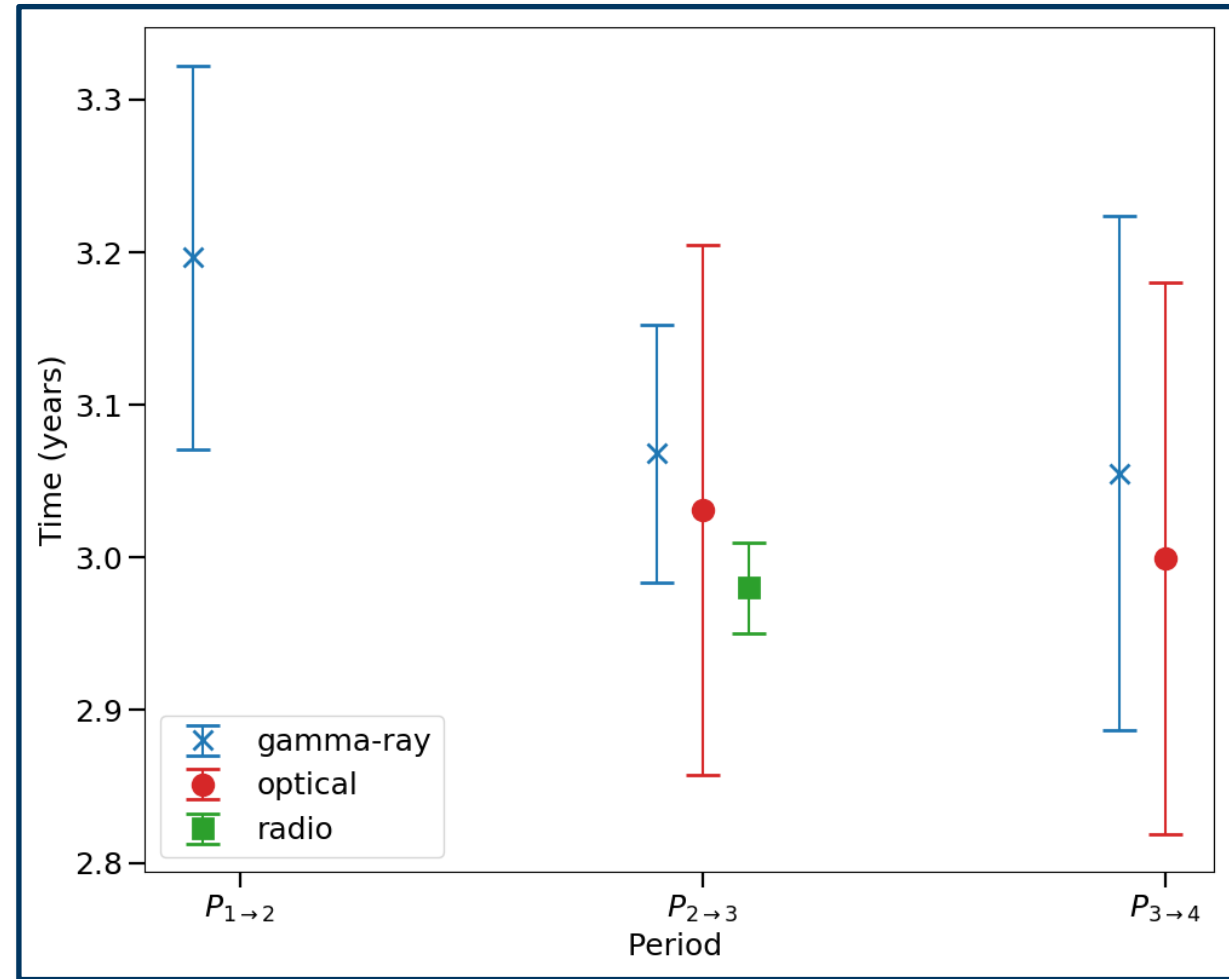
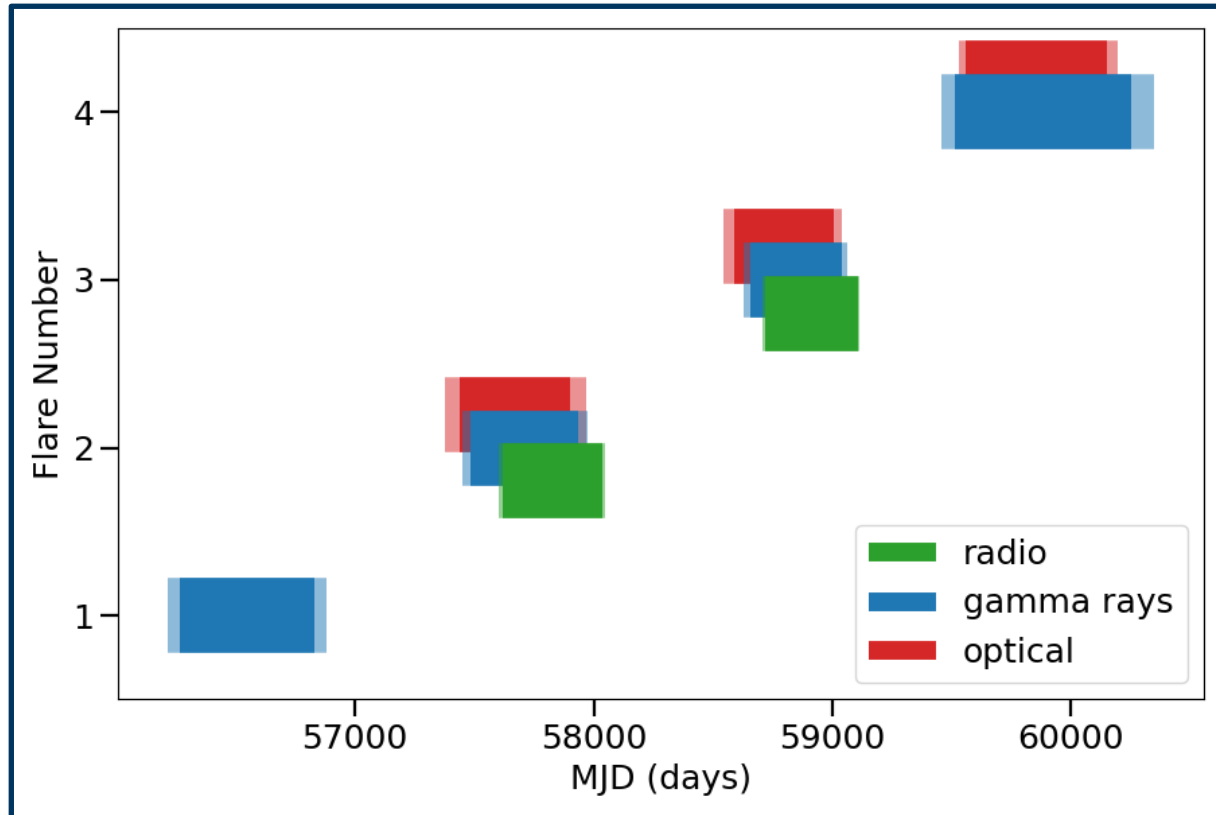
# Flare Durations

Not on Poster!



# Flare Durations + Periods Between

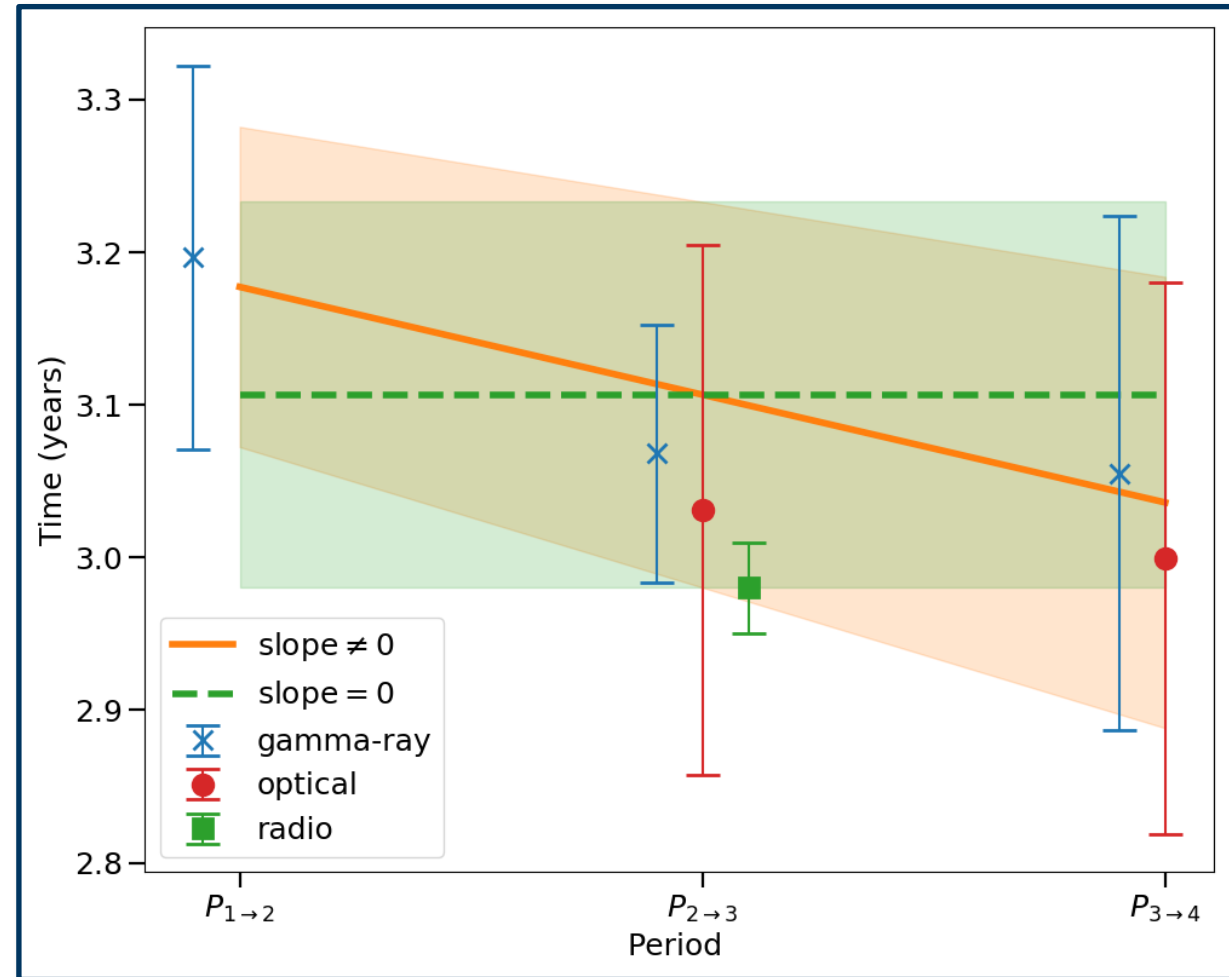
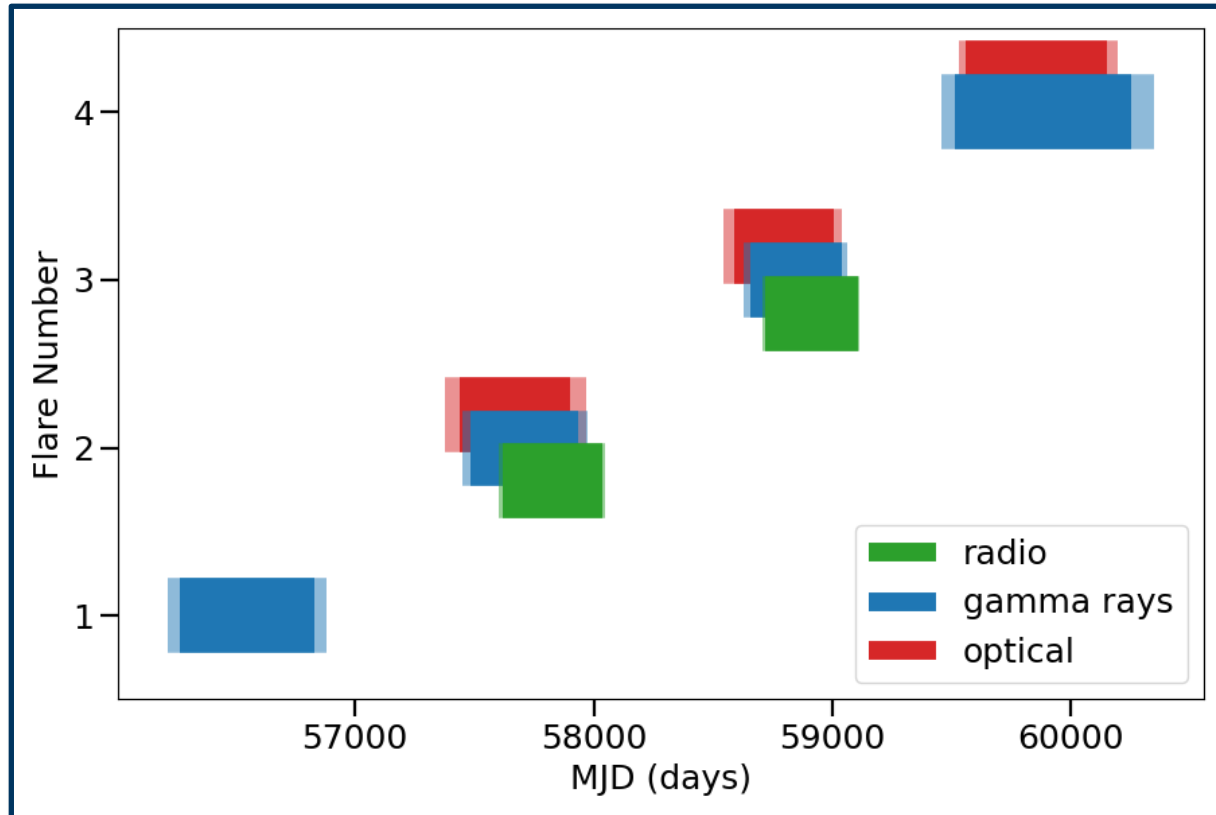
Not on Poster!





# Flare Durations + Periods Between

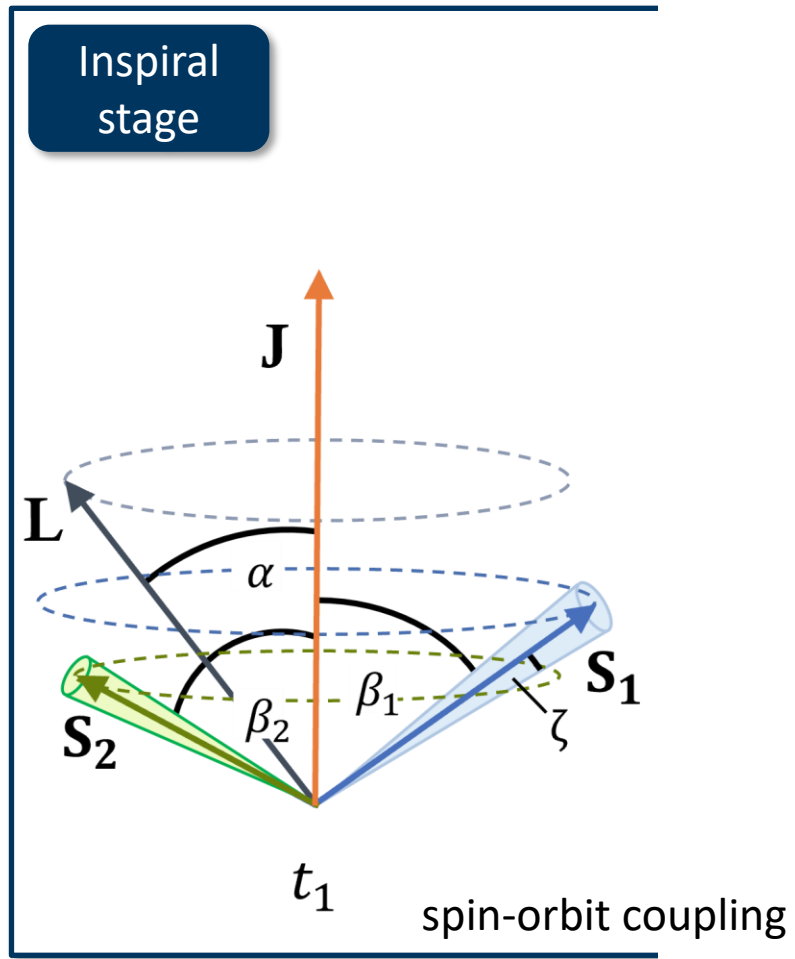
Not on Poster!



# The Jet Precession Model

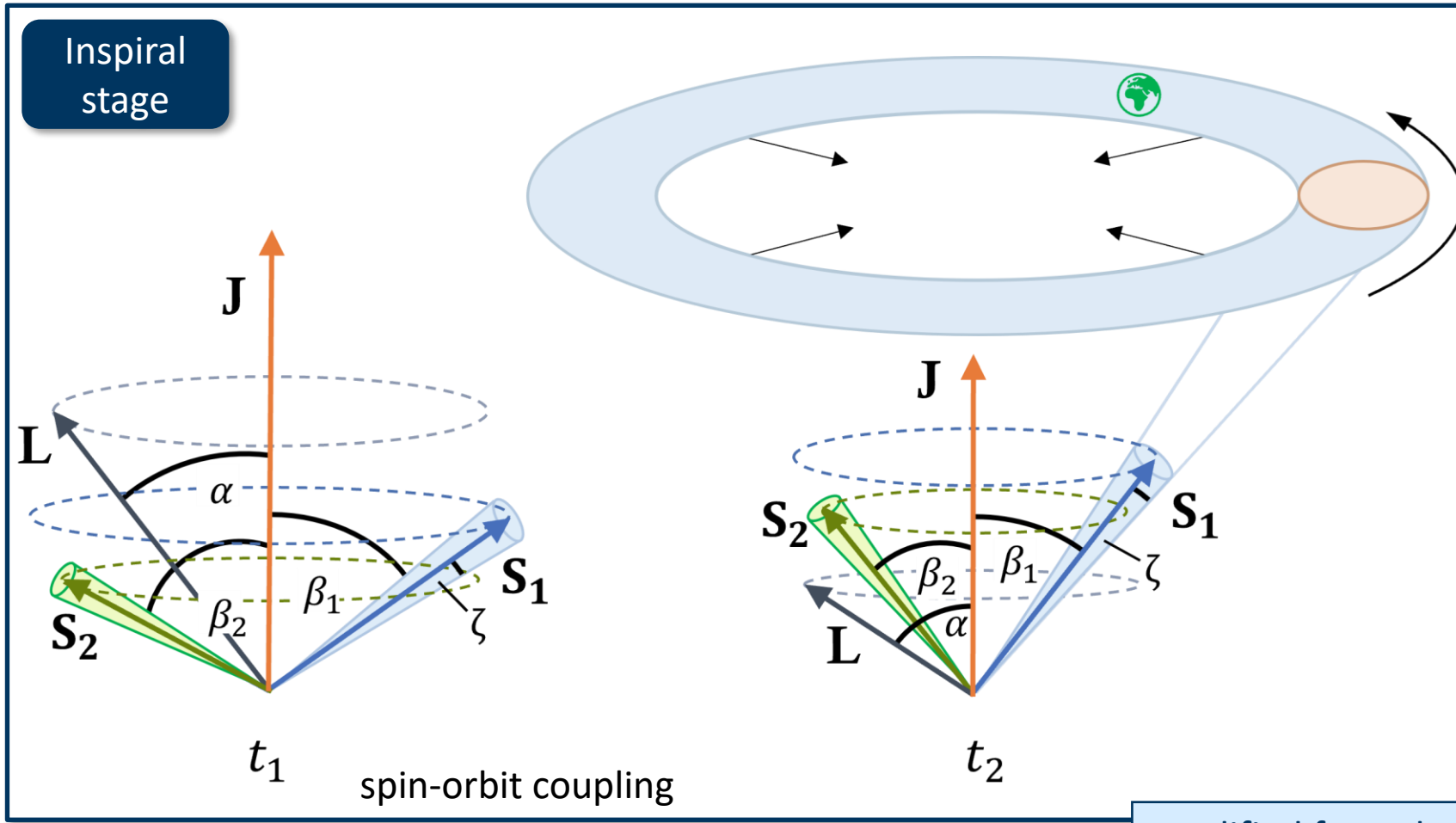
modified from de Bruijn et al. (2020)

# The Jet Precession Model



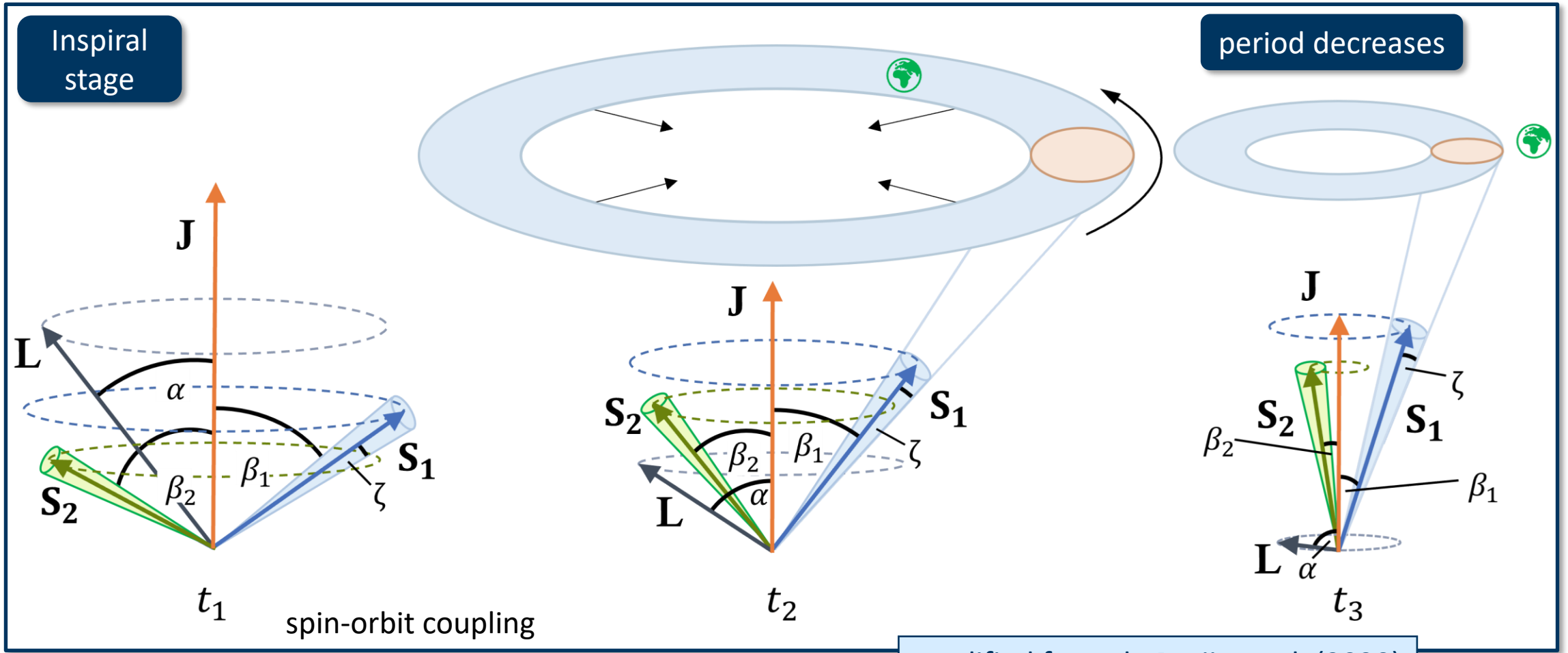
modified from de Bruijn et al. (2020)

# The Jet Precession Model



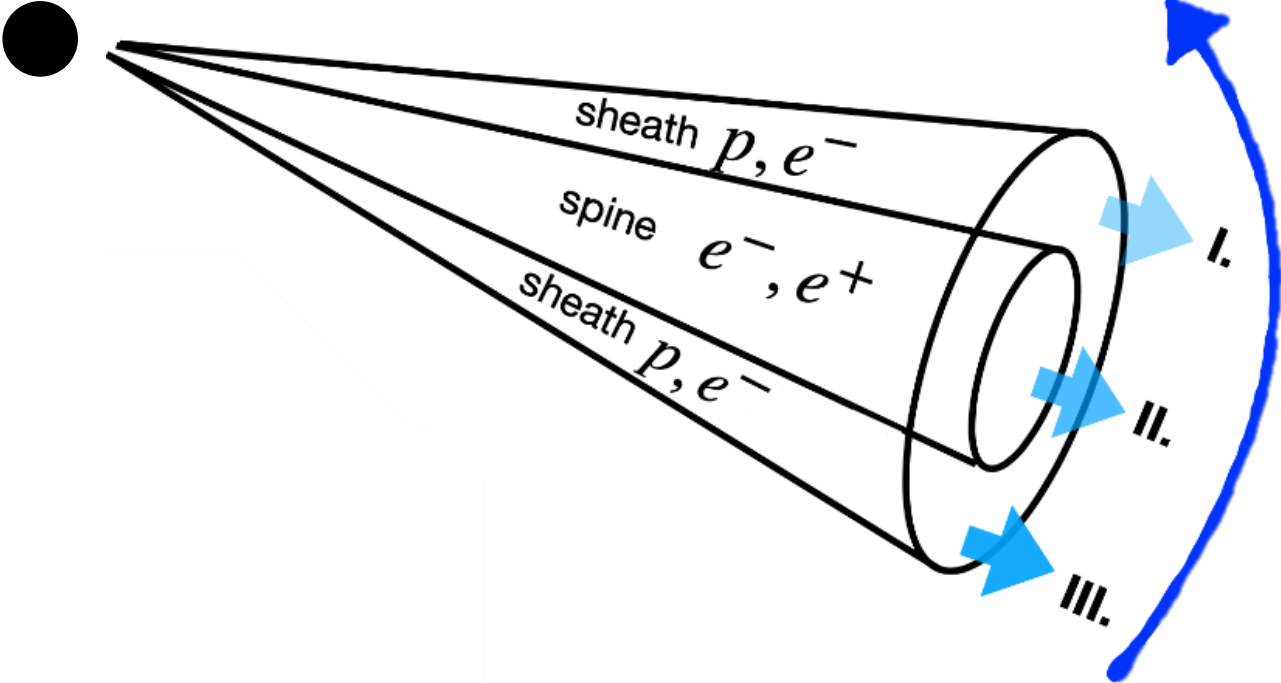
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# The Jet Precession Model



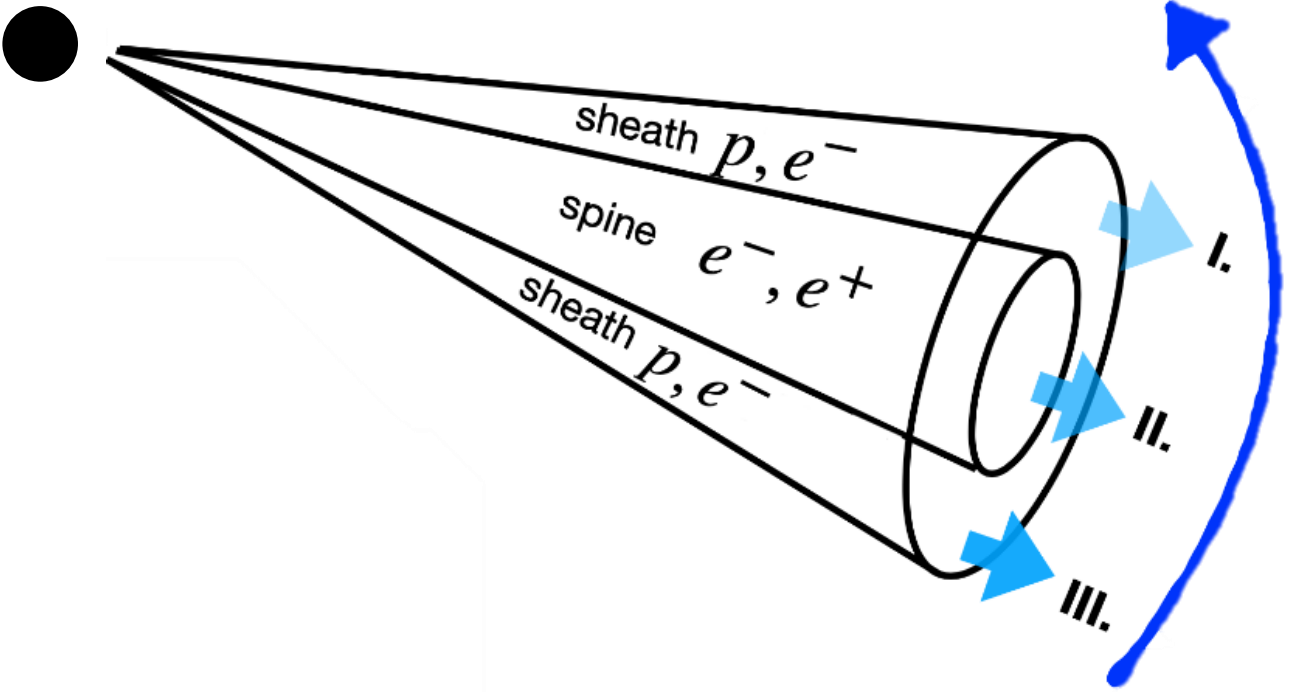
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# Double Peak Structure – an Explanation



Observer

# Double Peak Structure – an Explanation

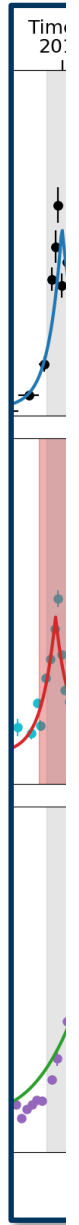
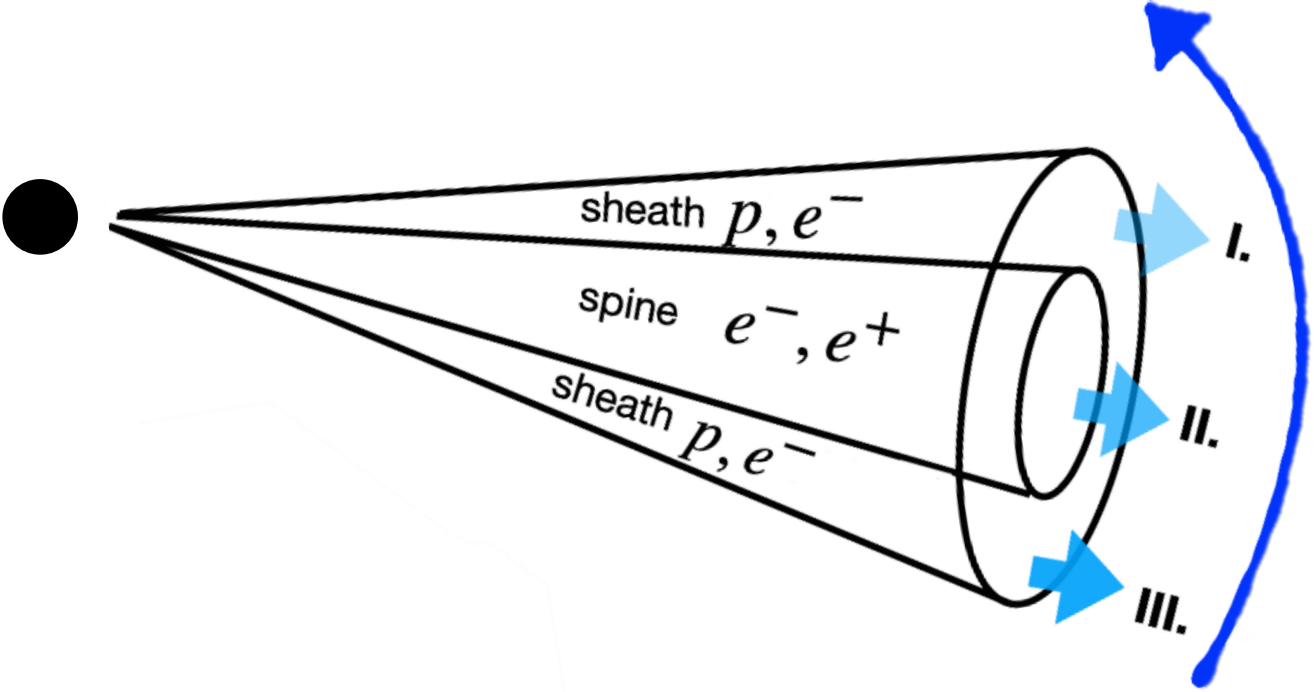


Target



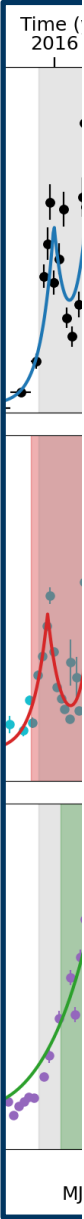
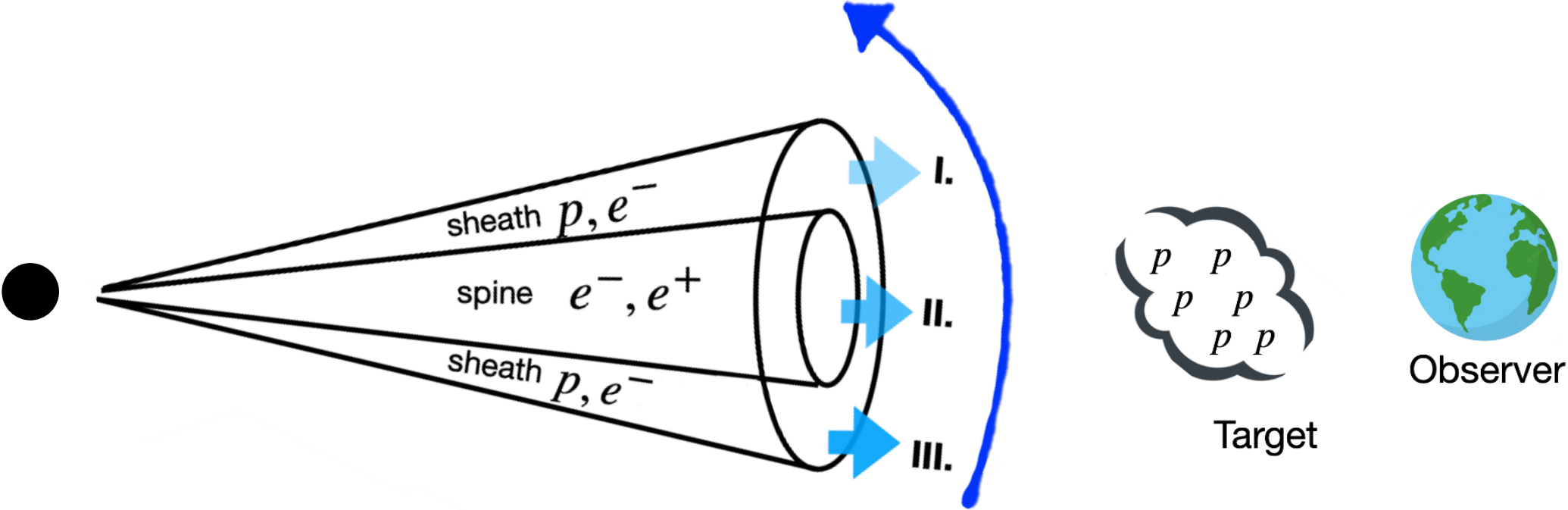
Observer

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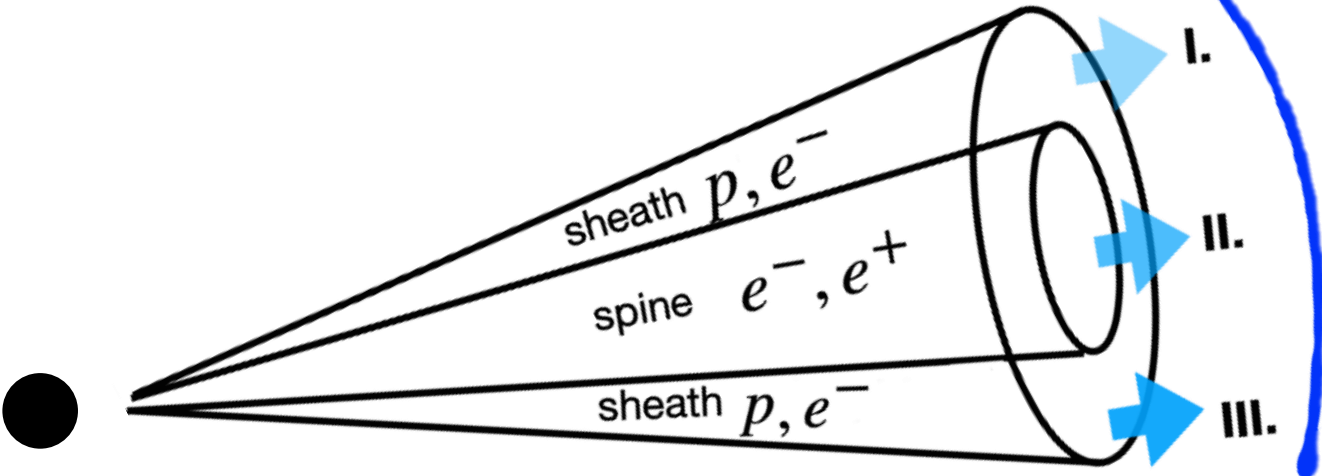




# Double Peak Structure – an Explanation



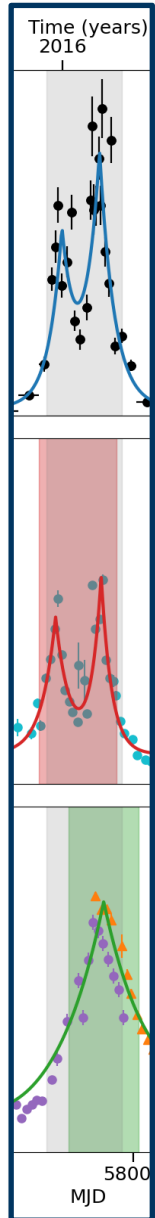
# Double Peak Structure – an Explanation



Target



Observer



# Applying the Jet Precession Model in 2 Steps

## 1. Step

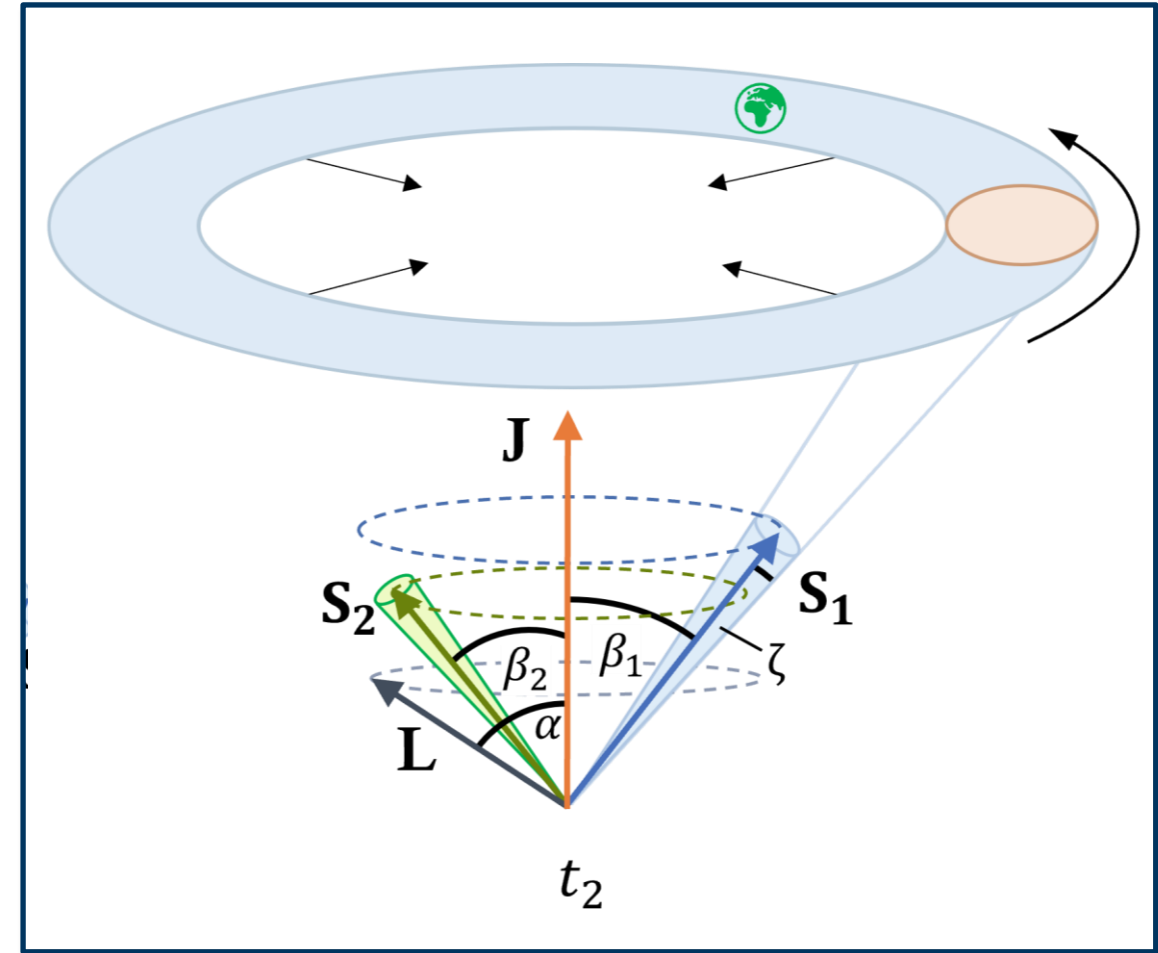
Input from  
Observation



$$\phi(\Delta T_{GW}, M, q) = \phi(\Delta T_{GW} - P_{1 \rightarrow 2}, M, q) \pm \zeta + 360^\circ$$

## 2. Step

$$\phi(\Delta T_{GW}, M, q) = \phi(\Delta T_{GW} - P_{1 \rightarrow 2} - P_{2 \rightarrow 3}, M, q) \pm 2\zeta + 720$$



Kun, JJ et al. (2022)

# Applying the Jet Precession Model in 2 Steps

## 1. Step

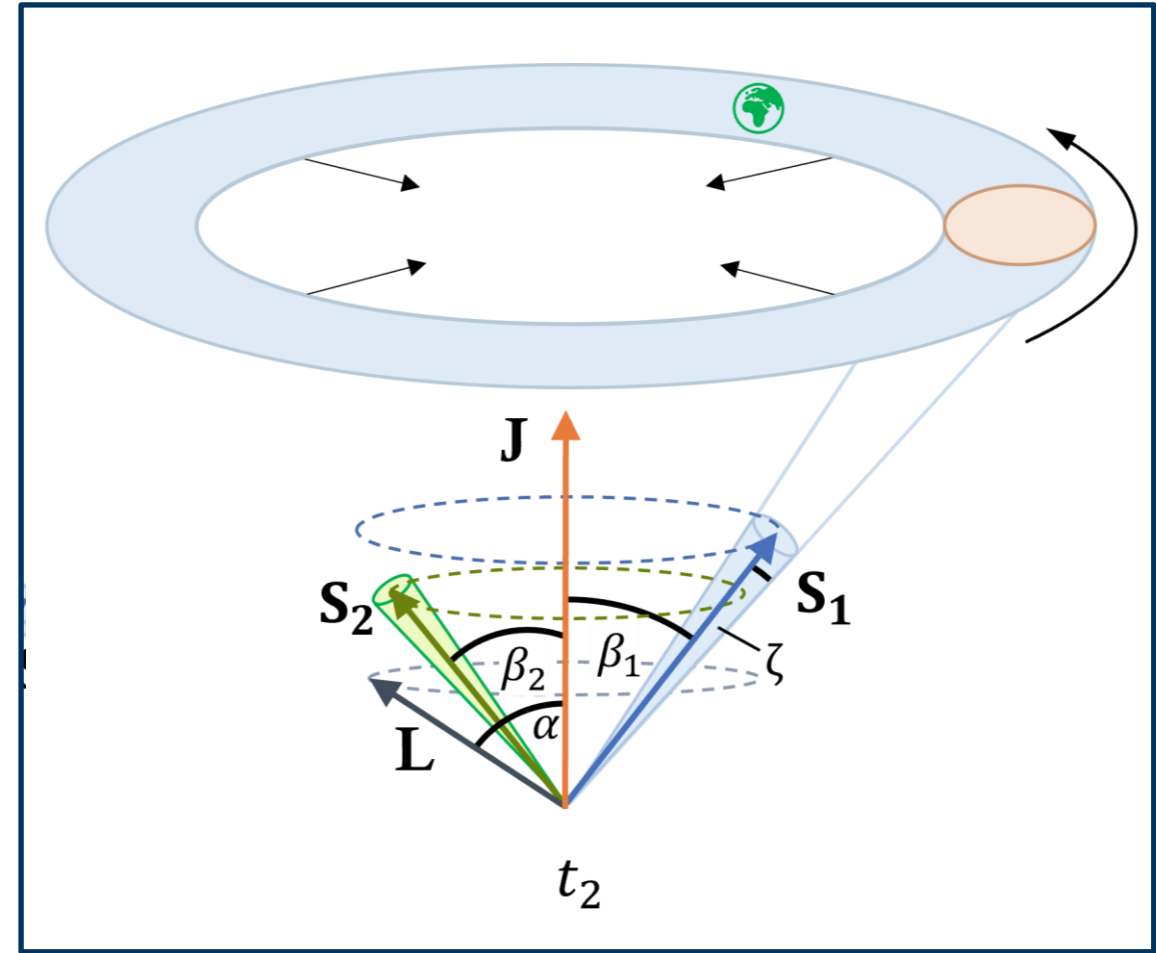
Input from  
Observation



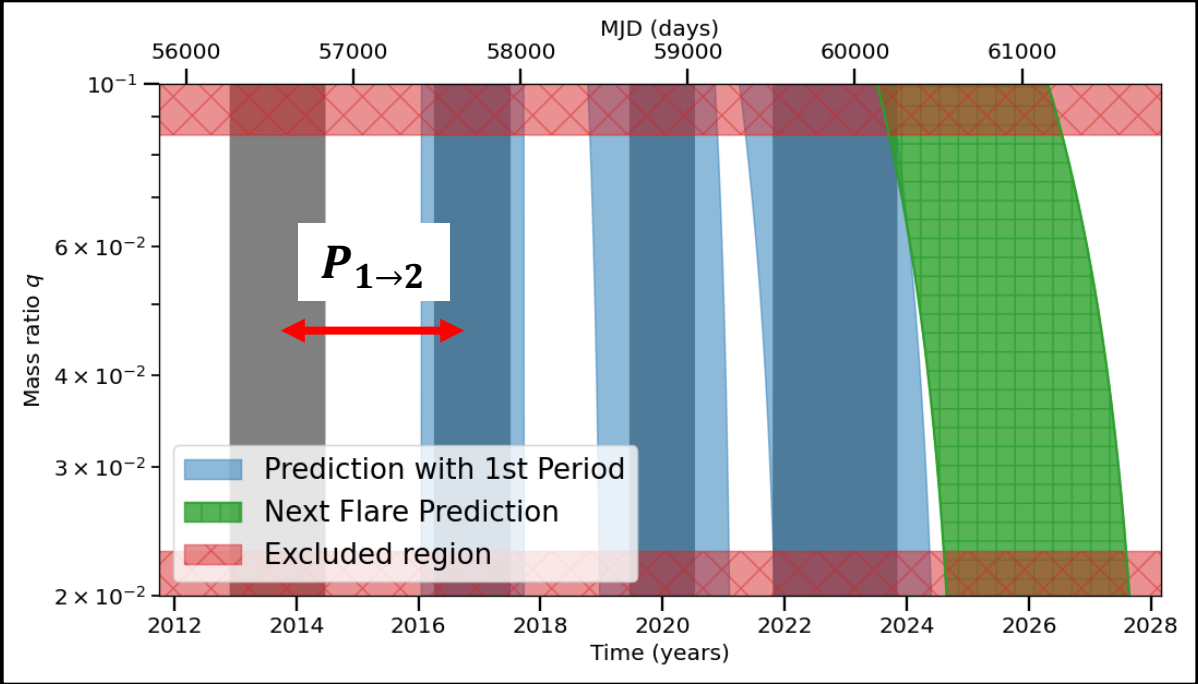
$$\phi(\Delta T_{GW}, M, q) = \phi(\Delta T_{GW} - P_{1 \rightarrow 2}, M, q) \pm \zeta + 360^\circ$$

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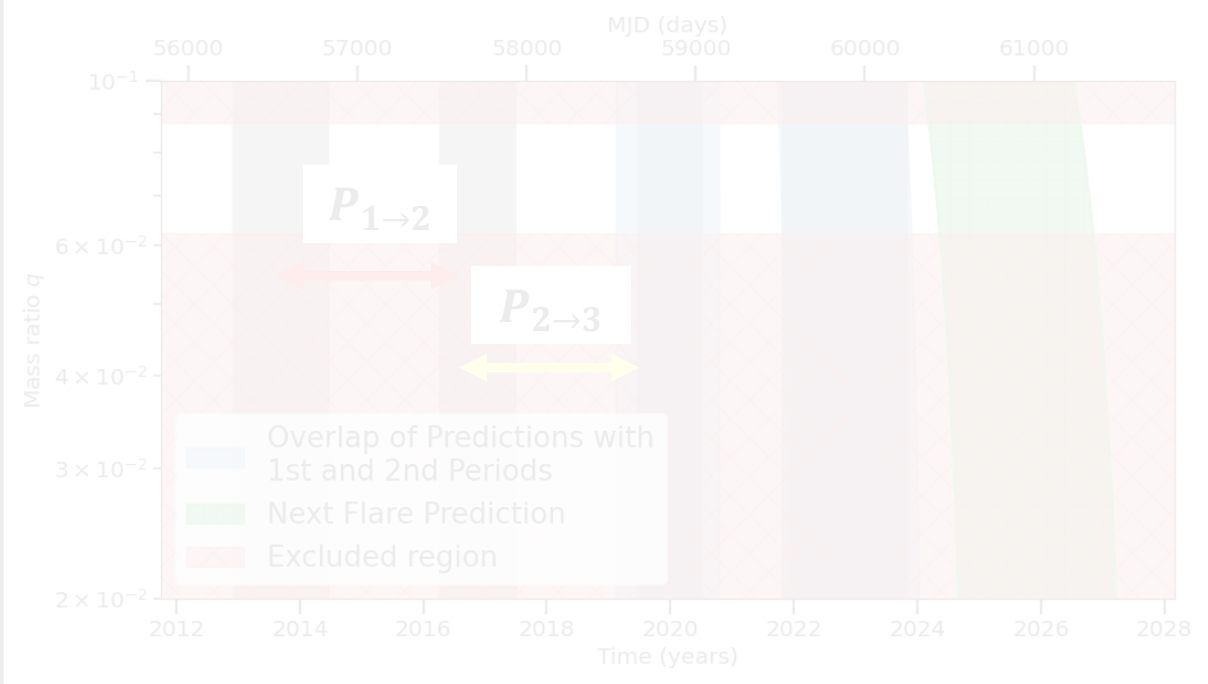
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# Flare Prediction in Gamma Rays – 4th Flare agrees!

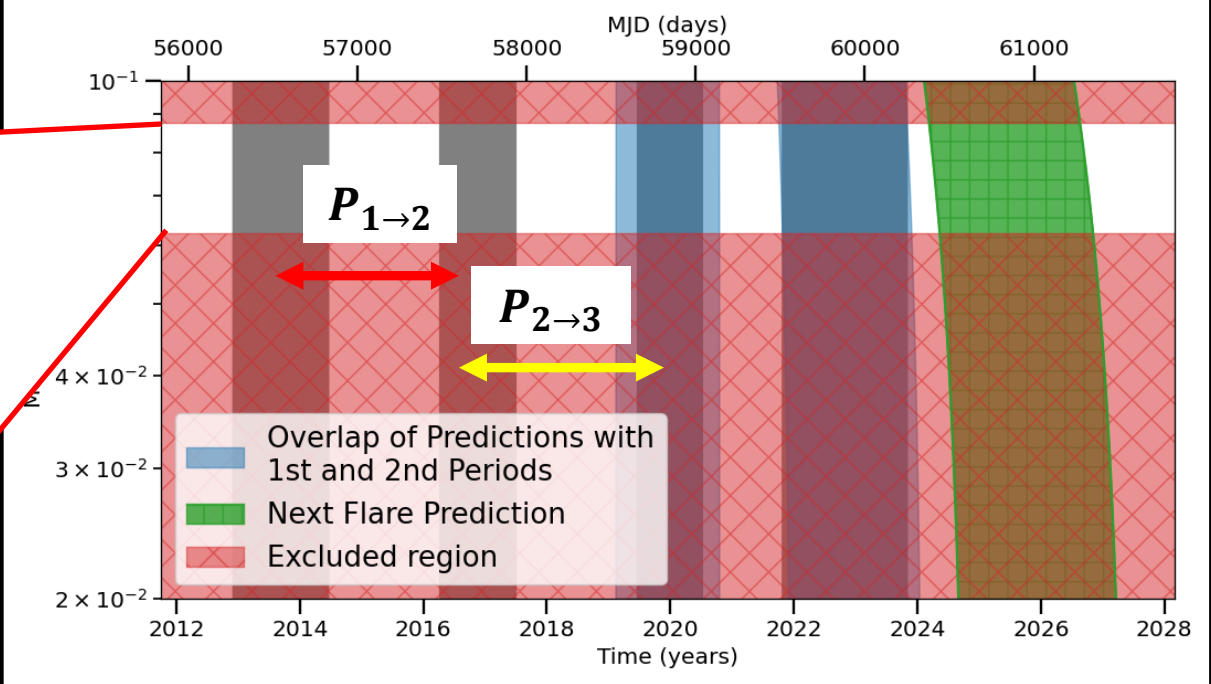
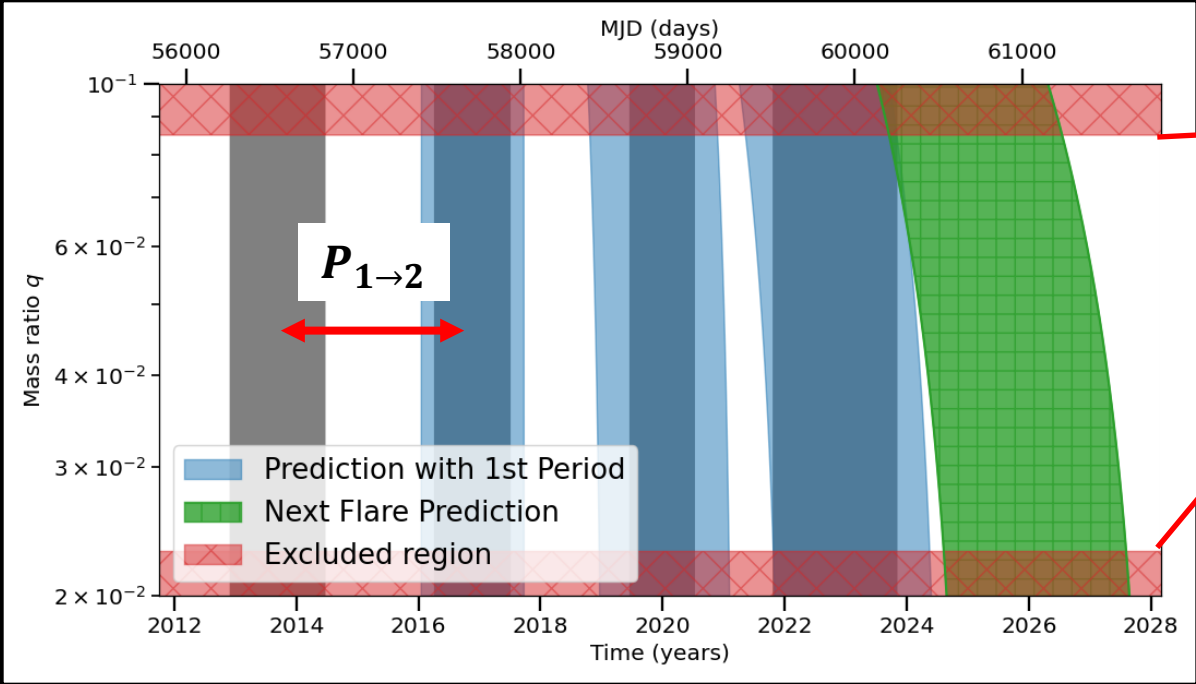


half-opening angle:  $\zeta \sim 5.73^\circ$



half-opening angle:  $\zeta \sim 8.34^\circ$

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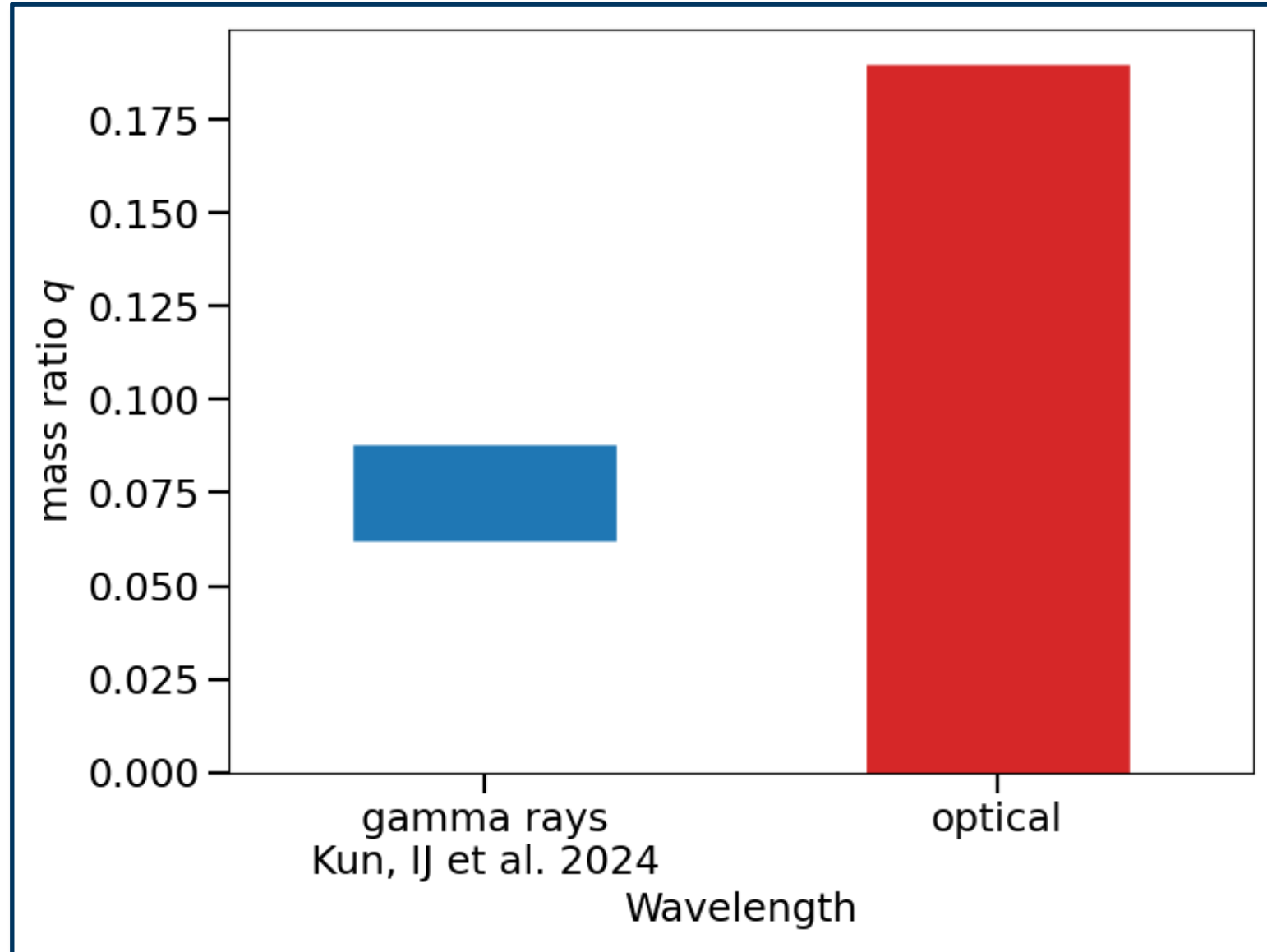
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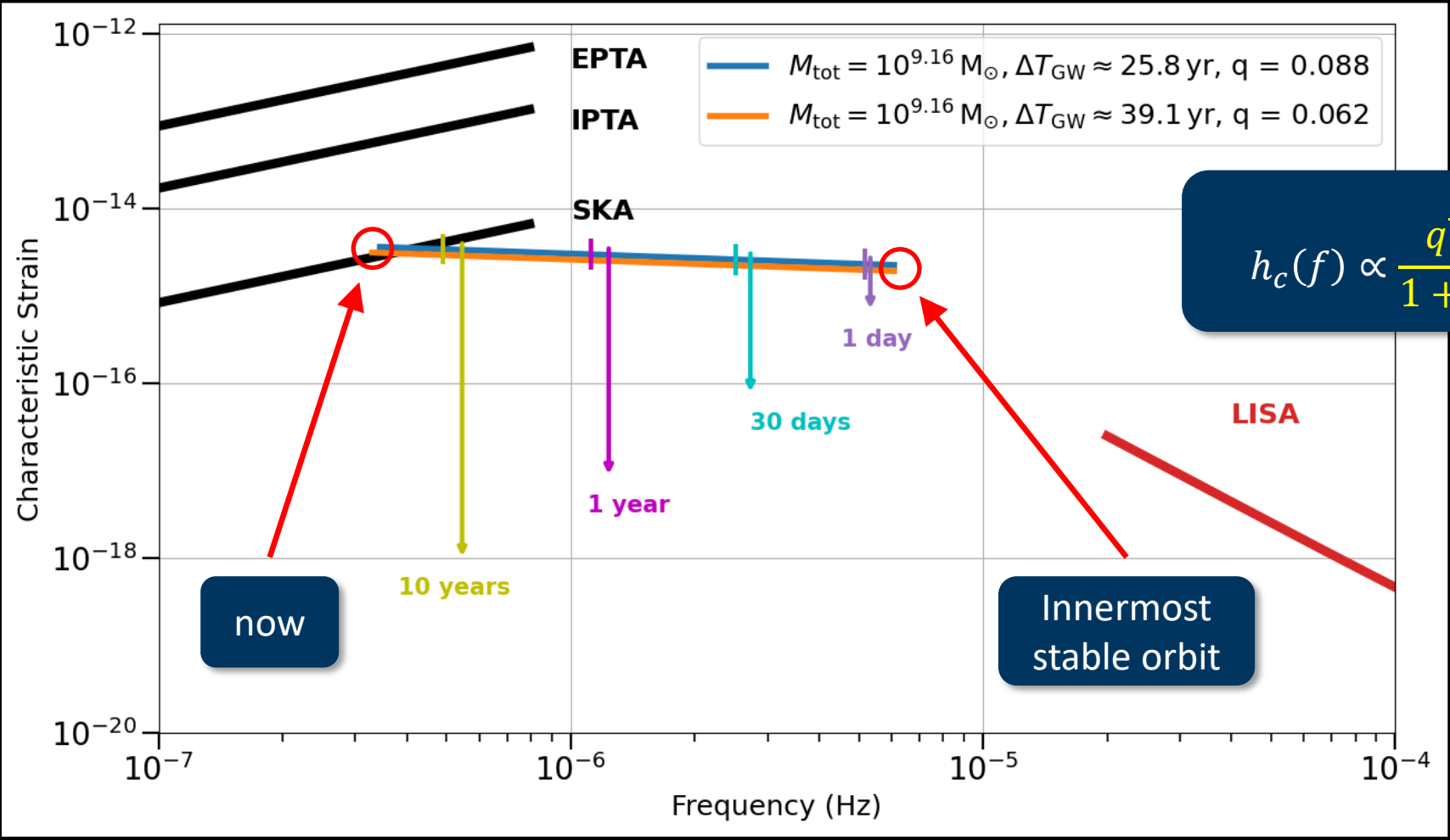
$$0.062 \leq q \leq 0.088$$

# Constrained Binary Mass Ratios

Not on Poster!



# Expected Gravitational Wave Signal



$$h_c(f) \propto \frac{q^{\frac{1}{2}}}{1+q} M^{\frac{5}{6}} \cdot \frac{1}{r(z)(1+z)^{1/2}} f^{-1/6}$$

$r(z)$ : comoving distance



# Summary – Blazar J1048+7143

analytical Jet Precession model applied with:

1.

optical

- 3 flares observed
- Mass ratio constrained from above

$$q \lesssim 0.19$$

- 4 flares observed
- 4<sup>th</sup> flare (successfully) predicted
- Mass ratio constrained from above **AND** below

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- 2 flares observed
- 3<sup>rd</sup> flare: more data necessary

Time range of next flare, if the jet will point  
at Earth once more:

2024 March 10 - 2026 November 6

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2024 March 10 - 2026 November 6

# Summary – Blazar J1048+7143

analytical Jet Precession model applied with:



in combination with Spine-Sheath jet model

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double-peak structure in gamma rays and optical light curve explainable

More coming:

- ZTF optical light curve → combined optical light curve
- Expected Neutrino upper limits
- Swift X-Ray light curve

Goal → Combined MM picture!

Stay Tuned!!!

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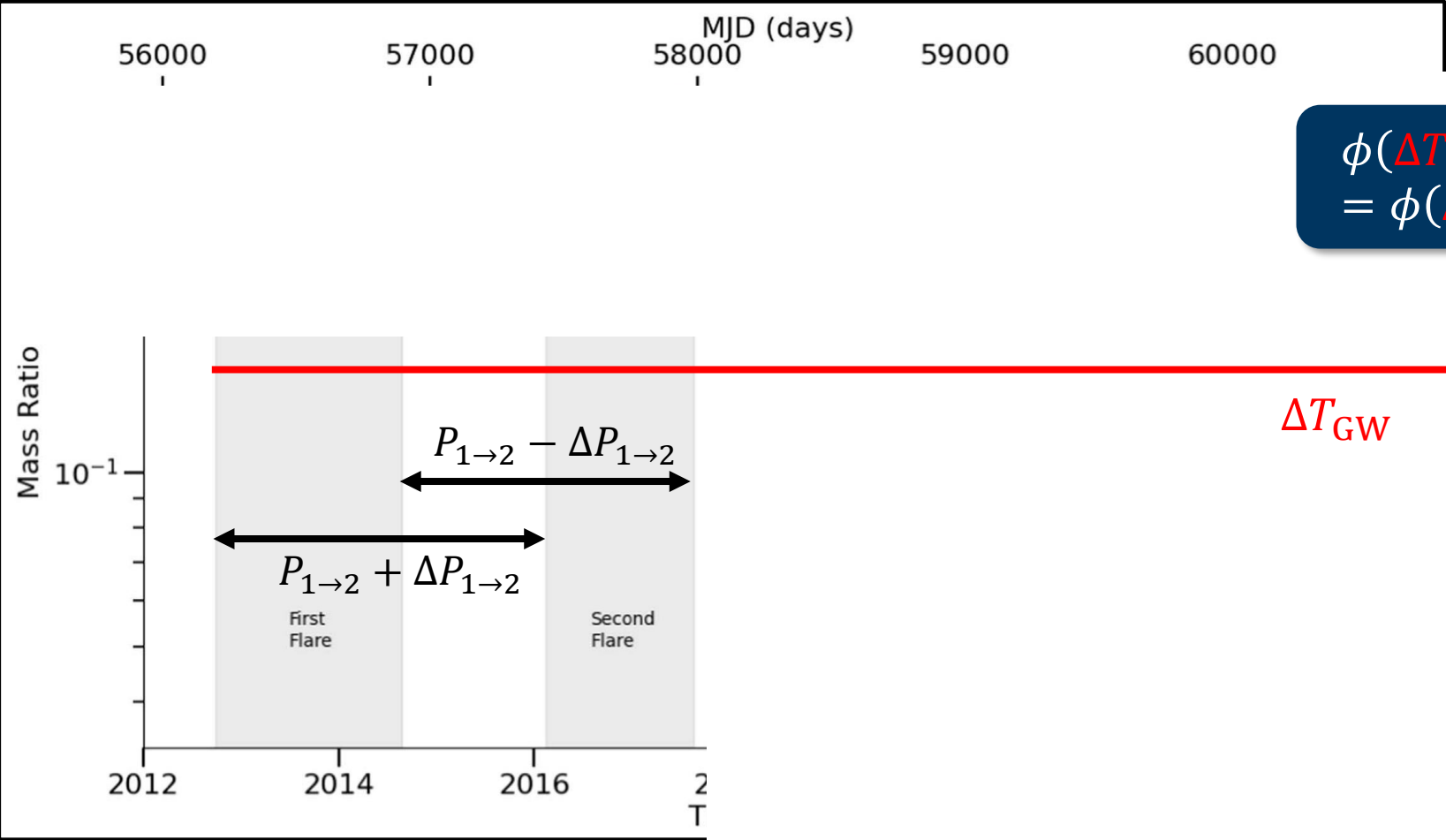
2024 March 10 - 2026 November 6

# Appendix

# Flare Prediction in Gamma Rays

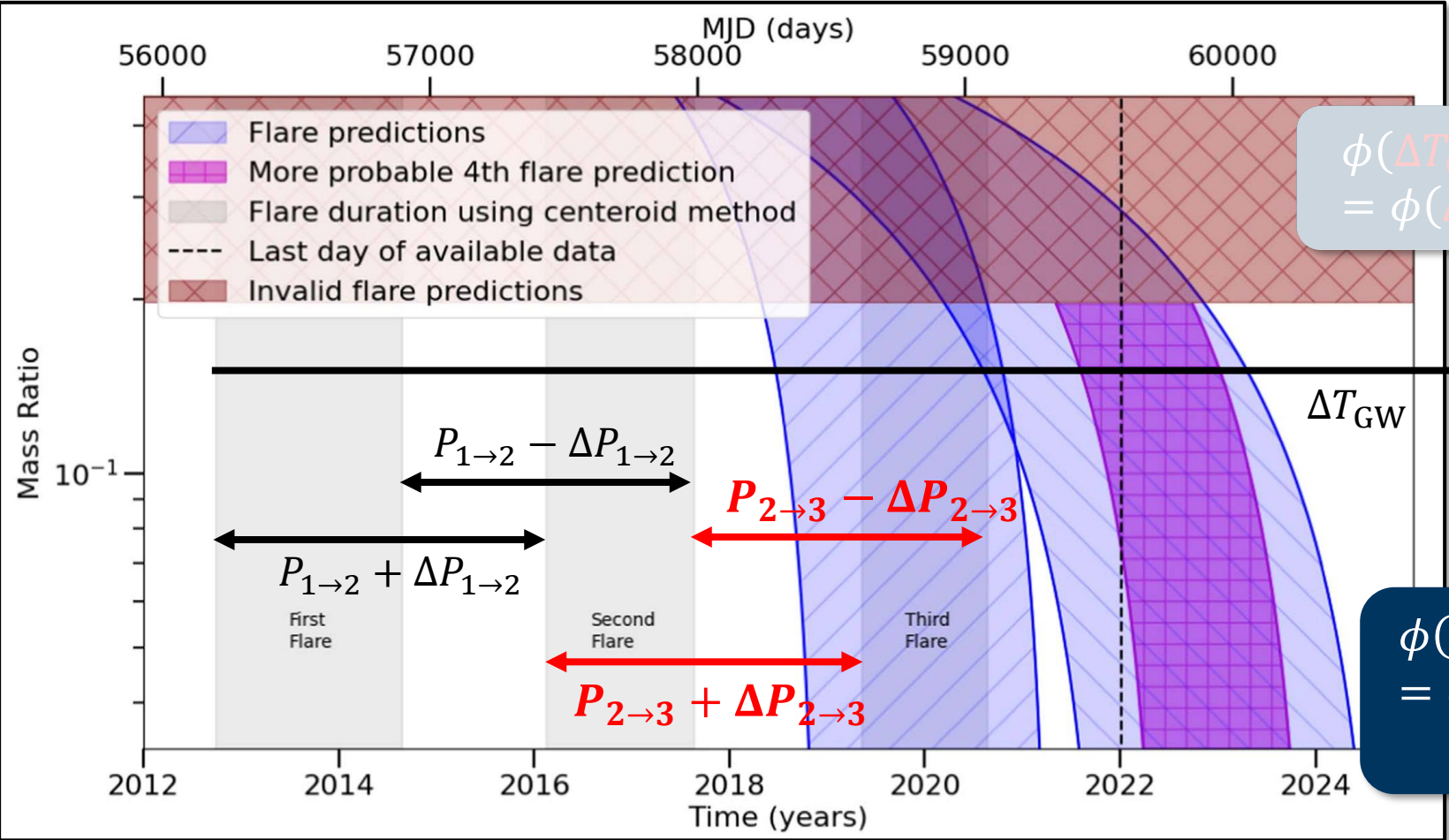
**1. Step**

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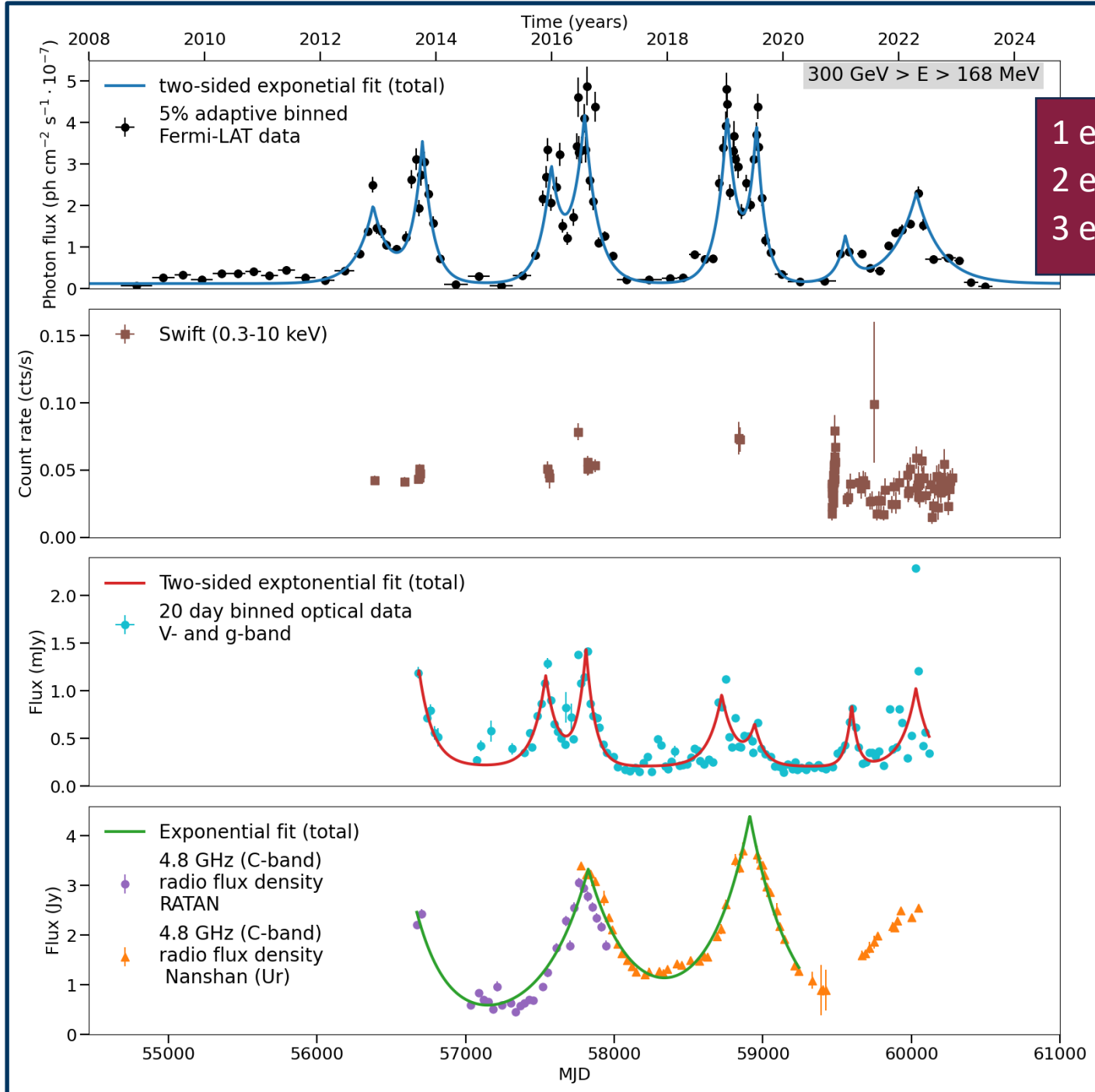
2. Step

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half-opening angle:  $\zeta \sim 5.73^\circ$



# Backup: Gamma-Ray Light Curve + X-Ray + Optical + Radio



1 exp func:  $\chi^2_{\text{red}} \approx 22.57$   
2 exp func:  $\chi^2_{\text{red}} \approx 8.54$   
3 exp func:  $\chi^2_{\text{red}} \approx 10.43$

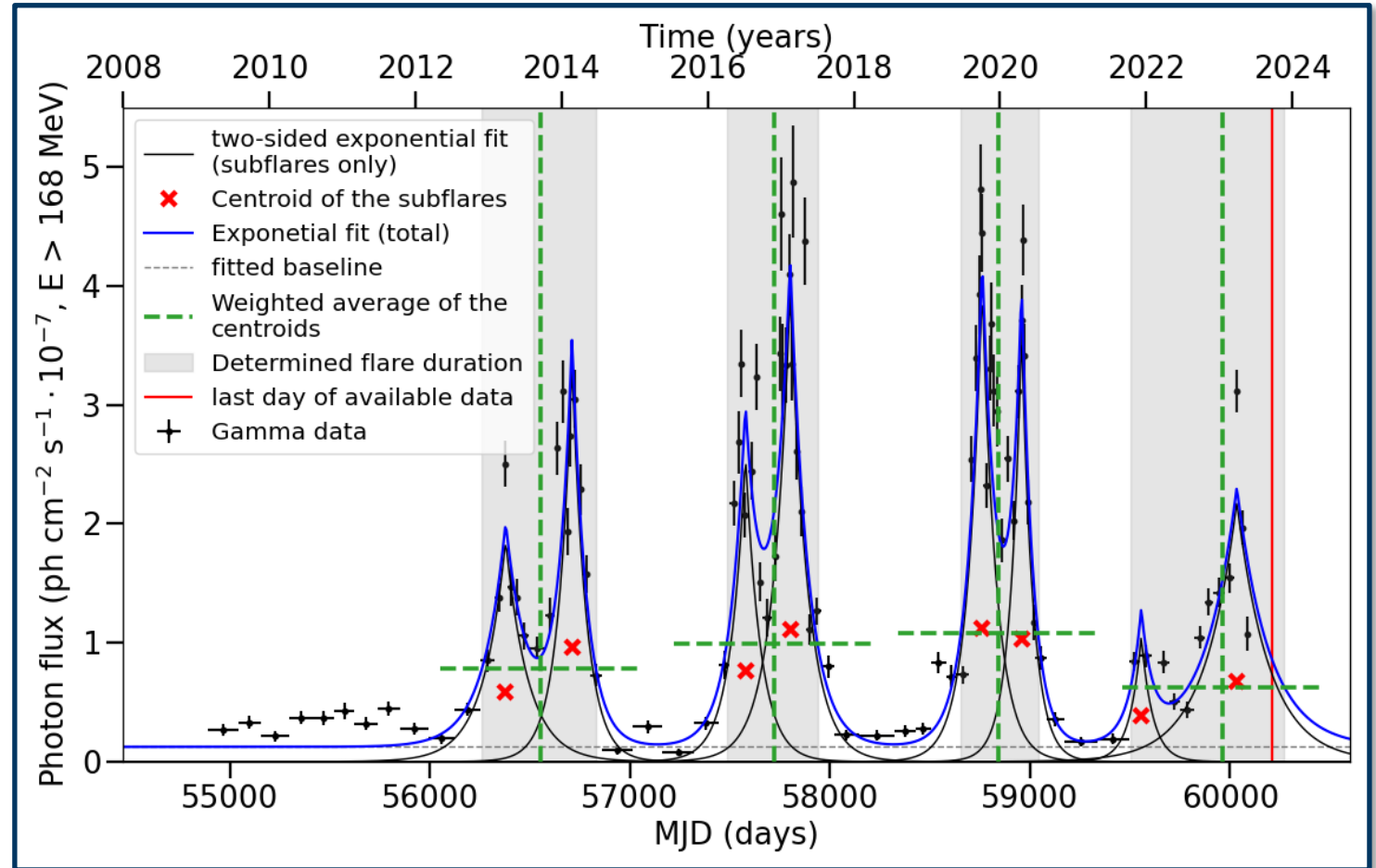
# Backup: J1048+7143 – Centroid Method (Kun, IJ et al. 2022)

$$X_{i,j} = \frac{\int t \cdot F_{i,j}(t) dt}{\int F_{i,j}(t) dt}$$

$$Y_{i,j} = \frac{1}{2} \frac{\int F_{i,j}^2(t) dt}{\int F_{i,j}(t) dt}$$

$$X_i = \frac{A_i}{A_i + B_i} X_{i,1} + \frac{B_i}{A_i + B_i} X_{i,2}$$

$$Y_i = \frac{A_i}{A_i + B_i} Y_{i,1} + \frac{B_i}{A_i + B_i} Y_{i,2}$$



# Backup: J1048+7143 – Flare Characteristics 1/3

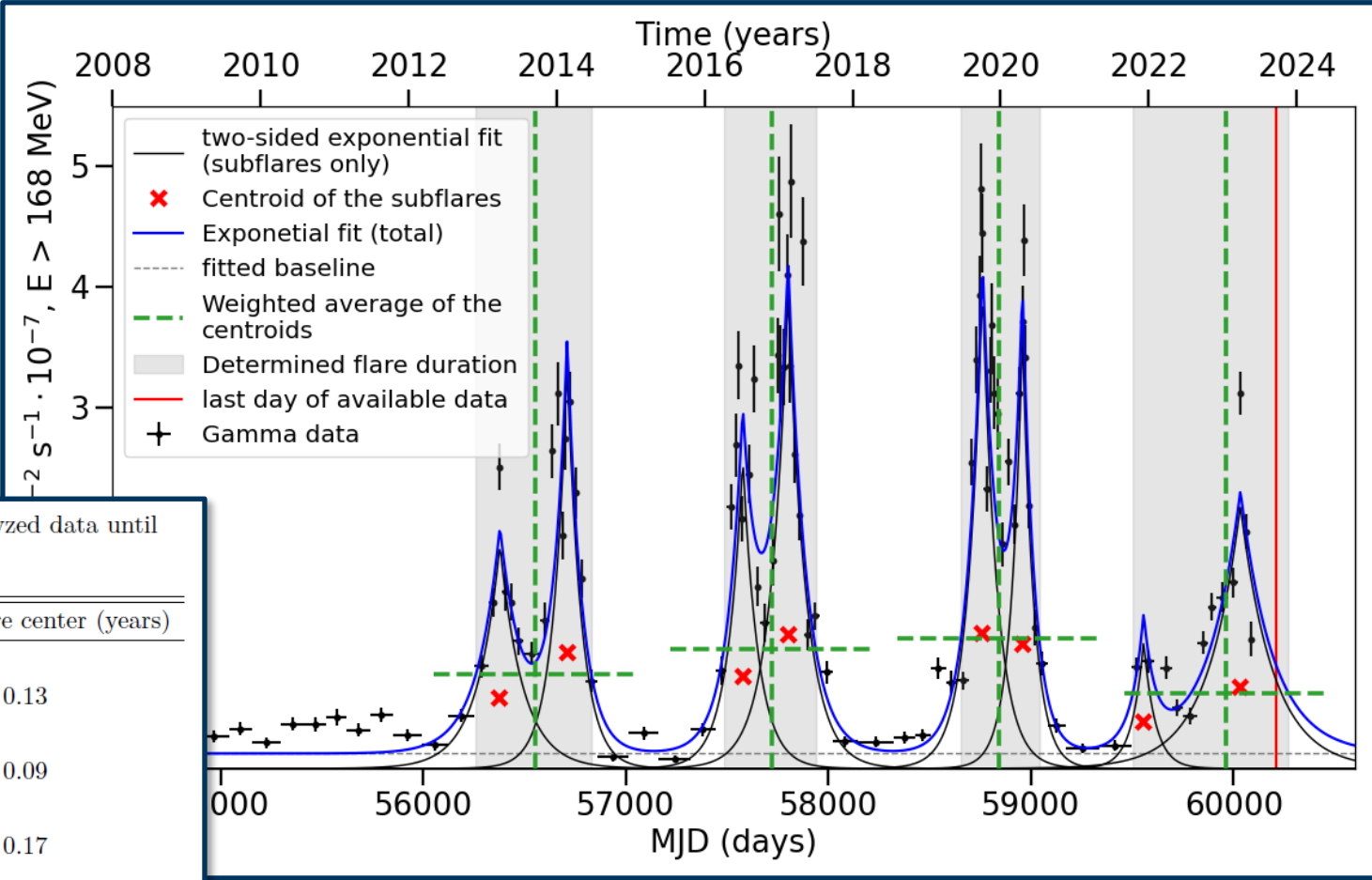
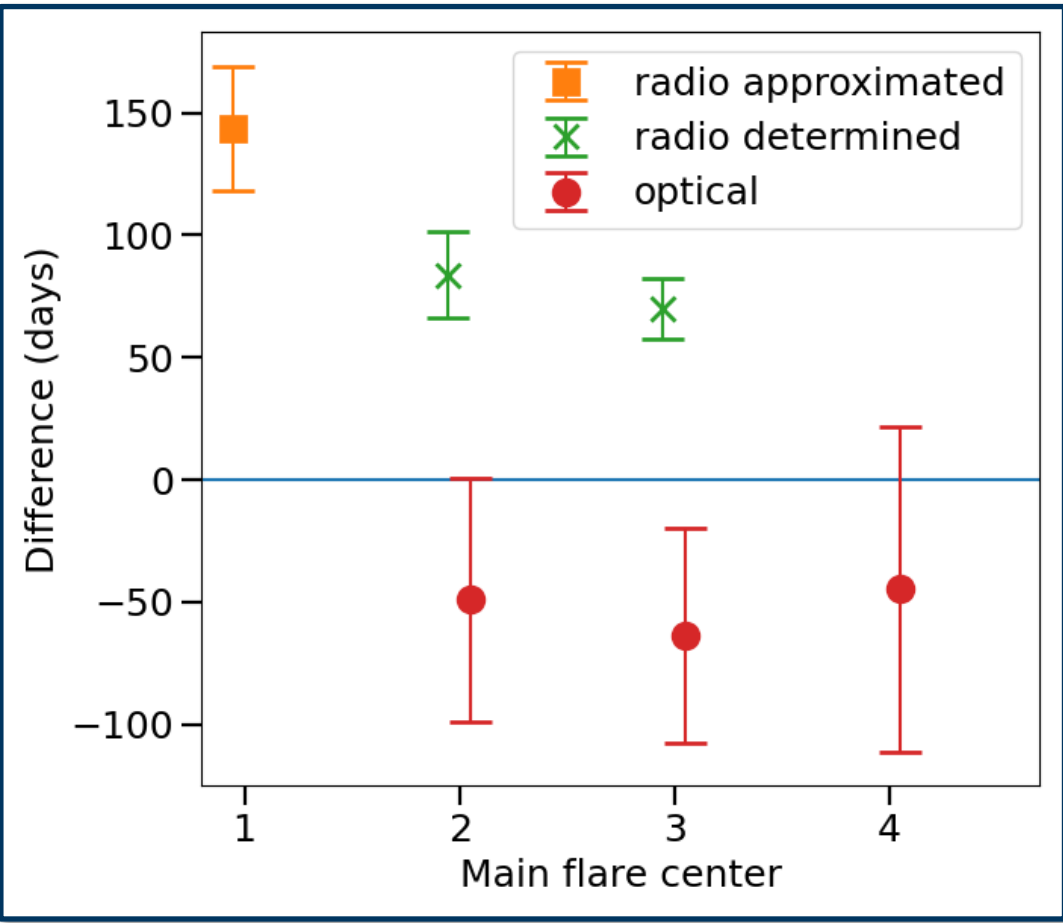


Table 5.1: Characteristics of the gamma-ray light curve of J1048+7143 (analyzed data until MJD 60099) applying the centroid method.

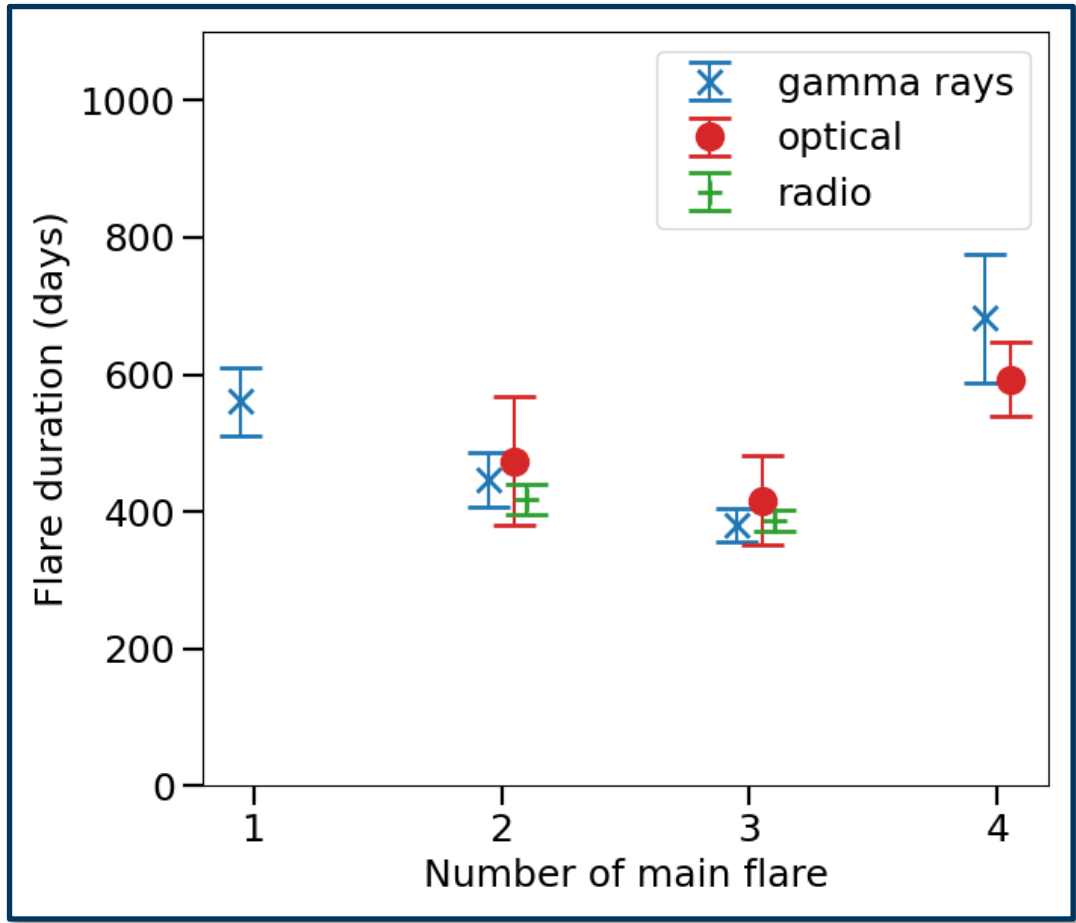
Parameter	Flare center (MJD)	Flare duration (days)	Time till next flare center (years)
F1	56554 ± 38	565 ± 78	
$P_{1 \rightarrow 2}$			3.20 ± 0.13
F2	57722 ± 25	450 ± 61	
$P_{2 \rightarrow 3}$			3.07 ± 0.09
F3	58842 ± 18	383 ± 38	
$P_{3 \rightarrow 4}$			3.06 ± 0.17
F4	59958 ± 59	756 ± 113	

# Backup: J1048+7143 – Flare Characteristics 2/3

Difference of main flare centers in the gamma-ray and radio light curve:

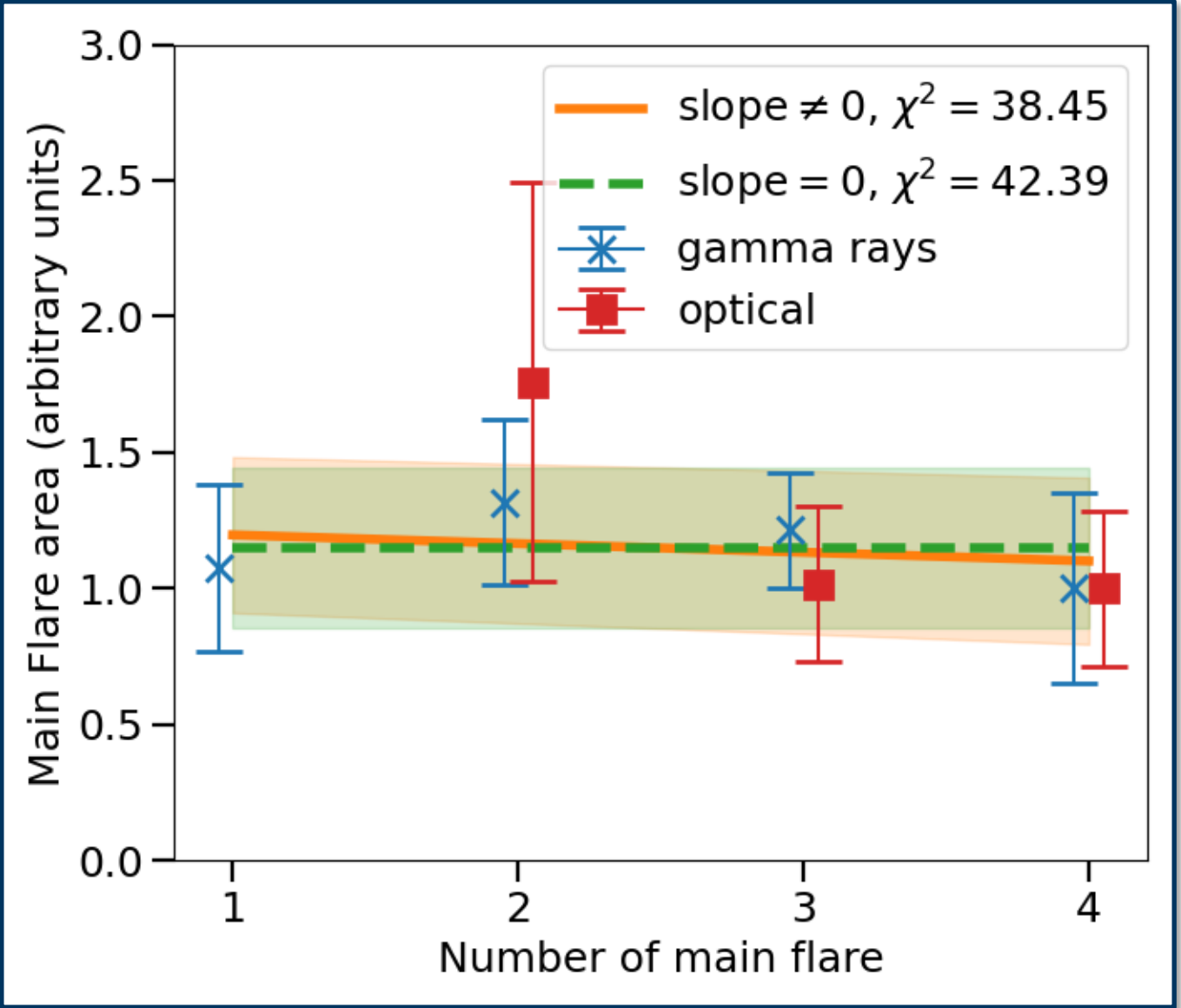


Main Flare durations:



# Backup: J1048+7143 – Flare Characteristics 3/3

Flare areas normalized to last flare



# Backup: J1048+7143 – Possible Nutation in Gamma Rays + Optical

Earth's path through the jet

