Extensive air shower predictions: sufficiently constrained by accelerator data?

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Cosmic Rays and Neuronos in the Multi-Messenge Era

Paris\_Brecember 19-18-2024

arXiv: 2401.06202; 2403.16106; 2404.02085; 2409.05501

Jet production in MC generators: collinear factorization of pQCD

$$\frac{d\sigma_{pp}^{\text{jet}}}{dp_t^2} = \sum_{I,J=q,\bar{q},g} f_I \otimes \frac{d\sigma_{IJ}^{2\to 2}}{dp_t^2} \otimes f_J$$

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- for  $Q_0 \sim$  few GeV, soft physics irrelevant
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• are MC predictions trustworthy, without such a mechanism?

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### Dynamical higher twist effects in hadronic scattering

#### Hint: collinear factorization of pQCD valid at leading twist level

- perhaps higher twist effects do the job?
  - come into play at relatively small  $p_t$  [suppressed as  $1/p_t^n$ ]

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HT effects in  $\gamma^* A/pA$ : coherent multiple scattering on 'soft' gluons [Qiu & Vitev, PRL93 (2004) 262301; PLB632 (2006) 507]





 scattering involves any number of 'soft' gluon pairs (⇒ multiparton correlators)



[SO & Bleicher, Universe 5 (2019) 106; SO, arXiv: 2401.06202]

# Dynamical higher twist effects in hadronic scattering

NB: only moderate HT corrections allowed by HERA data



• HT corrections important at low  $Q^2$ 

•  $\Rightarrow$  too strong corrections at tension with  $Q^2$ -evolution of  $F_2$ 

# Dynamical higher twist effects in hadronic scattering

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#### $\pi$ - over $\rho$ -exchange dominance $\Rightarrow \sim 20\%$ increase of $N_{\mu}$

• why so?! • isospin symmetry:  $\rho^+: \rho^-: \rho^0 = 1:1:1$ •  $\Rightarrow \langle E_{\pi^{\pm}} \rangle: \langle E_{\pi^0} \rangle = 2:1$  in central production  $(\rho^{\pm} \rightarrow \pi^{\pm}\pi^0, \rho^0 \rightarrow \pi^{+}\pi^{-})$  $\sum_{\mu=1}^{1.4} \int_{\mu=1}^{\mu=1} \int_{\mu=1}^{$ 

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 $\pi$ -exchange process in  $\pi^+A$ : only  $\rho^+$  and  $\rho^0$  produced forward

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$$\Rightarrow \langle E_{\pi^{\pm}} \rangle : \langle E_{\pi^{0}} \rangle = 3 : 1$$

$$\pi^{+} \frac{\underline{u} \quad \underline{u}}{\overline{d}} \frac{\underline{u}}{d} | \begin{array}{c} \mu \\ \overline{d} \\ \pi^{0} \end{array} \rho^{+} \qquad \pi^{+} \frac{\underline{u} \quad \underline{u}}{\overline{d}} \frac{\underline{u}}{d} | \begin{array}{c} \mu \\ \overline{u} \\ \overline{u} \\ \pi^{+} \end{array} \rho^{0} |$$

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 ⇒ less energy channeled into e/m cascades!

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Energy-dependence: driven by absorptive corrections to the process

 high x production of ρ in π<sup>±</sup>p (π<sup>±</sup>A) or of neutrons in pp: only without additional inelastic rescatterings



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Energy-dependence: driven by absorptive corrections to the process

- high x production of ρ in π<sup>±</sup>p (π<sup>±</sup>A) or of neutrons in pp: only without additional inelastic rescatterings
- now can be tested in pp → nX thanks to LHCf data [backup slides]



#### Results for extensive air showers [SO, arXiv: 2403.16106]



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• up to  $\simeq 10 \text{ g/cm}^2$  shift of  $X_{\rm max}$  wrt QGSJET-II-04

• up to 
$$\simeq 5\%$$
 change of  $N_{\mu}$ 

What is the reason for the stability of the predictions?

- the model sufficiently constrained by LHC data?
- or a mere consequence of a particular model approach?

#### Kinematic range for hadron production, relevant for $N_{\mu}$ predictions

• let us restrict ourselves with pion production only:

$$N_p^{\mu}(E_0) \simeq \int dx \, \frac{dN_{p-\mathrm{air}}^{\pi^-}(E_0,x)}{dx} \, N_{\pi^{\pm}}^{\mu}(xE_0)$$

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Using an ansatz:  $dN_{p-\mathrm{air}}^{\pi^{\pm}}(E_0,x)/dx \propto x^{-1-\Delta}(1-x)^{\beta}$ 

$$N_{p}^{\mu}(E_{0}) \propto E_{0}^{\alpha_{\mu}} \int_{x_{\min}}^{1} dx \, x^{\alpha_{\mu}-1-\Delta} (1-x)^{\beta} \overset{\#}{\underset{x}{\overset{\mu}{=}}} \overset{10^{2}}{\underset{10^{-1}}{\overset{\mu}{\underset{x}{\overset{\mu}{=}}}} \overset{\mu}{\underset{10^{-1}}{\overset{\mu}{\underset{x}{\overset{\mu}{=}}}} \overset{\mu}{\underset{10^{-1}}{\overset{\mu}{\underset{10^{-1}}{\overset{\mu}{\underset{10^{-1}}}}} \overset{\mu}{\underset{10^{-1}}{\overset{\mu}{\underset{10^{-1}}}} \overset{\mu}{\underset{10^{-1}}{\overset{\mu}{\underset{10^{-1}}}} \overset{\mu}{\underset{10^{-1}}{\overset{\mu}{\underset{10^{-1}}}} \overset{\mu}{\underset{10^{-1}}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}} \overset{\mu}{\underset{10^{-1}}} \overset{$$

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• largest contribution from  $\langle x_{\pi} \rangle \simeq \frac{\alpha_{\mu} - \Delta}{\alpha_{\mu} + \beta - 1 - \Delta} \sim 0.1$  $(\Delta \simeq 0.4, \ \alpha_{\mu} \simeq 0.9, \ \beta \simeq 4.5)$ 



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- relevant (x<sub>π</sub>) for π-air interactions follows similarly



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Accounting for all 'stable' hadrons ( $\pi^{\pm}$ , kaons, (anti)nucleons)

• relevant quantity for EAS muon content:

$$\sum_{h=\text{stable}} \langle (x_E^h)^{\alpha_{\mu}} \rangle = \sum_{h=\text{stable}} \int dx_E \, x_E^{\alpha_{\mu}} \, \frac{dN_{\pi^{\pm} \text{air}}^n(E_0, x_E)}{dx_E}$$

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• NB: 
$$\sum_{h=\text{stable}} \langle x_E^h \rangle \simeq 1 - \langle x_E^{\pi^0} \rangle$$

#### How to increase $\sum_{h=\text{stable}} \langle x_E^h \rangle$ (decrease $\langle x_E^{\pi^0} \rangle$ )?

• change the energy dependence of the pion exchange process  $\Rightarrow$  larger forward yield of  $\rho$ -mesons  $\Rightarrow$  higher  $\langle E_{\pi^{\pm}} \rangle / \langle E_{\pi^{0}} \rangle$ 

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- change the model calibration (e.g. based on NA61 data): more kaons & (anti)nucleons





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Neglecting absorptive corrections to the  $\pi$ -exchange process: larger forward production of  $\rho$ -mesons at higher energies





#### In such a case: large contribution of pion elastic scattering



• 
$$\sigma_{\pi-\text{air}}^{\text{el}} \rightarrow \frac{1}{2} \sigma_{\pi-\text{air}}^{\text{tot}}$$
 at  $E_0 \rightarrow \infty$
# Model uncertainties for predicted $N_{\mu}$ [SO & Sigl, arXiv: 2404.02085]

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 $\pi^{*} + N \rightarrow \rho^{0} (10^{6} \text{ GeV/c})$ 

OGSJET-III

no absorption

pion exchange (QGSJET-III)

0.6

0.8

- $\Rightarrow$  scarce hadron production!
- $\Rightarrow$  decrease of  $N_{\mu}$ (instead of an enhancement)

# Model uncertainties for predicted $N_{\mu}$ [SO & Sigl, arXiv: 2404.02085]





•  $\simeq 40\%$  more kaons &  $\simeq 60\%$  more (anti)protons required

 NB: such enhancements create a tension with other data on kaon & (anti)proton production in πp & pp interactions



### Relative changes of the calculated $N_{\mu}$ : < 10%



- why?
- small impact of the the considered enhancements on  $\sum_{h=\text{stable}} \langle x_E^h \rangle$ (changes mostly affect central hadron production)

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- $\Rightarrow$  accelerator data allow one to enhance  $N_{\mu}$  by up to  $\sim 10\%$

### Predictions for EAS maximum depth $X_{max}$ : 3 main 'switches'

- inelastic proton-air cross section  $(\sigma_{p-air}^{inel})$
- inelastic diffraction rate  $(\sigma_{p-air}^{diffr}/\sigma_{p-air}^{inel})$
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- even smaller difference for pA:  $\sigma_{pp}^{\text{inel}} \propto R_p^2$ ,  $\sigma_{pA}^{\text{inel}} \propto (R_p + R_A)^2$



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- NB: 1% change of  $\sigma_{p-\text{air}}^{\text{inel}} \Rightarrow \Delta X_{\text{max}} \simeq 1 \text{ g/cm}^2 \text{ at } 10^{19} \text{ eV}$



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 $\sigma_{pp}^{\text{inel}} \propto R_p^2$ ,  $\sigma_{pA}^{\text{inel}} \propto (R_p + R_A)^2$ 

• NB: 1% change of 
$$\sigma_{p-\text{air}}^{\text{inel}} \Rightarrow \Delta X_{\text{max}} \simeq 1 \text{ g/cm}^2 \text{ at } 10^{19} \text{ eV}$$

$$R_{pp} = 2 R_p$$

$$\mathbf{R}_{pA} = \mathbf{R}_{p} + \mathbf{R}_{A}$$

Diffraction uncertainties:  $\Delta X_{\text{max}} \lesssim 5 \text{ g/cm}^2$  [SO, PRD89 (2014) 074009]

### The only significant freedom left: inelasticity for p - air

• higher energy  $\Rightarrow$  higher multiple scattering  $\Rightarrow$  higher  $K_{p-\text{air}}^{\text{inel}}$ 

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How to give less energy away to secondary hadrons?

- central rapidity density of secondaries: constrained by data
- main 'switch': constituent parton (string end) momentum distribution  $(x^{-\alpha_q})$  [SO, J.Phys. G29 (2003) 831]



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- $\alpha_q \rightarrow 1$ : approximate Feynman scaling for forward spectra
- NB: may not work for semihard scattering (minijet production)



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 NB: higher discrimination power expected from combined studies with central & forward detectors (e.g. LHCf & ATLAS) [SO, Bleicher, Pierog & Werner, PRD94 (2016) 114026]





### Choice of string end distribution $(x^{-\alpha_q})$ : impact on $X_{\text{max}}$



• up to  $\simeq 10~{
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• why a moderate effect on particle production & X<sub>max</sub>?

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- why a moderate effect on particle production & X<sub>max</sub>?
- 'warranted' scaling violation due to semihard scattering (energy fraction taken by perturbatively generated partons ⇒ lower bound on K<sup>inel</sup><sub>p-air</sub>)
   <sup>10</sup> [see backup slides]



#### Exotic: modification of the hadronization by 'collective effects'

 assuming this is modified by 'collective effects' & neglecting parton cascades: strings are formed between constituent partons & highest p<sub>t</sub> partons



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- rather nonphysical: collective effects may be strong in central (small b) collisions only
  - ⇒ should not have large impact on the average parton production pattern (dominated by peripheral collisions)





- energy-dependence of K<sup>inel</sup><sub>p-air</sub> reduced drastically
  - $\alpha_q = 0.9$ :  $K_{p-\mathrm{air}}^\mathrm{inel}$  doesn't depend on energy above 1 PeV

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 ⇒ (mini)jet production has no impact on the inelasticity in the
 α<sub>q</sub> → 1 limit



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### Impact of string end distribution on $X_{max}$ (no parton cascades)



- up to  $\simeq 30 \text{ g/cm}^2$  shift of  $X_{\text{max}}$ ( $\alpha_q = 0.9$ )
- can be refuted/constrained by LHC data?

Image: A mathematical states of the state

### The limit $\alpha_a \rightarrow 1$ : disfavored by LHCf data on forward neutrons



•  $\alpha_q \rightarrow 1$ : forward neutron yield exceeds the measured one

### The limit $\alpha_q \rightarrow 1$ : disfavored by LHCf data on forward neutrons







# Outlook

- Major development in QGSJET-III: phenomenological treatment of HT corrections to hard scattering processes
  - tames the low  $p_t$  rise of (mini)jet rates
  - ${\, \bullet \, }$  reduces the model dependence on the low  $p_t$  cutoff  $Q_0$
- Technical improvement: treatment of  $\pi$ -exchange process
  - energy-dependence: due to absorptive corrections (probability not to have additional inelastic rescattering)
- Rather small changes for EAS characteristics (wrt QGSJET-II) • up to  $\simeq 10 \text{ g/cm}^2$  shift of  $X_{\text{max}}$  and up to  $\simeq 5\%$  change of  $N_{\mu}$
- Model uncertainties for  $N_{\mu}$ : only up to  $\sim 10\%$  enhancement
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- ${\rm \bullet}\,$  More exotic: 'collective effects'  $\Rightarrow \Delta X_{max}$  up to  $\simeq 30~{\rm g/cm^2}$ 
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### Extra slides follow

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### (1) Technical improvement: $\pi$ -exchange process



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#### And moving over 6 energy decades to 13 TeV c.m.



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- high energies ⇒ quick rise of (mini)jet production
  - small  $\alpha_s(p_t^2)$  compensated by infrared and collinear logs (arising from parton cascading):  $\ln(x_i/x_{i+1})$ ,  $\ln(p_{t_{i+1}}^2/p_{t_i}^2)$

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- no: x-distribution of those gluons is weighted with the hard scattering!



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Virtual gluons emitted by protons are indeed soft:  $\propto x^{-1-\Delta_g}$ 

• but the probability for hard scattering: convolution with  $\sigma^{hard}_{gg}$ 

$$w_{\text{hard}}(s) \propto \int dx^+ dx^- f_g(x^+, Q_0^2) f_g(x^-, Q_0^2) \,\mathbf{\sigma}_{gg}^{\text{hard}}(x^+ x^- s, Q_0^2)$$

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- $\sigma^{\rm hard}_{gg}(\hat{s},Q^2_0) \propto \hat{s}^{\Delta_{
  m hard}}$  contribution of the DGLAP 'ladder'
- $\Rightarrow$  gluons which succeed to interact have large x:  $\propto x^{\Delta_{hard} \Delta_g 1}$ (iff  $\Delta_{hard} \simeq 0.3 > \Delta_g$ )
  - i.e., first partons in a perturbative cascade are 'valence-like' (independently on our assumptions for string end distribution)

### Changing $X_{\text{max}}$ implies equal or larger changes for $X_{\text{max}}^{\mu}$

• any change of the primary interaction ( $\sigma_{p-\text{air}}^{\text{inel}}$ ,  $\sigma_{p-\text{air}}^{\text{diffr}}$ ,  $K_{p-\text{air}}^{\text{inel}}$ ) impacts only the initial stage of EAS development



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- ⇒ parallel up/down shift of the cascade profile (same shape)
- $\Rightarrow$  same effect on  $X_{\text{max}}$  &  $X_{\text{max}}^{\mu}$
- additionally: the corresponding change of physics impacts π-air interactions at all the steps of the cascade development
  - $\Rightarrow$  cumulative effect on  $X^{\mu}_{\max}$



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