New developments in EPOS :

Toward a global approach from Heavy Ions to Cosmic Rays

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Outline

Introduction

- Updates \rightarrow EPOS LHC-R
 - A real global approach to do hadronic interactions
- Impact of Hadronic Rescattering (HS)
- Predictions for air showers (EAS)
 - $\clubsuit X_{_{max}}$ and μ
- Muon puzzle
 - Why collective effects impact muon production ?

Recent LHC data provide new constraints on models changing X_{max} and the muon production if a global approach is used.

Global approach Impact of HS X_{max} and µ Muon puzzle

Sensitivity to Hadronic Interactions



- Air shower development dominated by few parameters
 - mass and energy of primary CR
 - cross-sections (p-Air and (π-K)-Air)
 - (in)elasticity
 - multiplicity
 - <u>charge ratio</u> and baryon production
- Change of primary = change of hadronic interaction parameters
 - cross-section, elasticity, mult. ...
- Model tuned to accelerator data

Theory AND data are important to constrain the hadronic model parameters.

Global approach Impact of HS X_{max} and μ Possible updates since EPOS LHC

- First LHC data lead to reduced differences between models
 But a number of new data since model release could be use to further improve the models :
 - Update of the p-p cross sections (ALFA)
 - Data at 13 TeV (CMS, ATLAS, LHCf)
 - More detailed p-Pb measurements (fluctuations) CMS
 - Particle yields as a function of multiplicity (ALICE, LHCb)
 - Very important to understand the mechanism behind particle production
- Update of EPOS LHC → EPOS LHC-R
 - New EPOS 4 available for heavy ion physics but not usable for air showers (yet)
 - Modify EPOS LHC to take into account new data and new knowledge accumulated with (and code from) EPOS 4
 - Almost final result (but still preliminary) including all <u>collective effects</u> !

Introduction



Introduction

What means global approach ?

Global approach is the key !

- Tuning models neglecting some physics process may lead to wrong parameters !
- Correct tune possible only if everything taken into account
- Even without a direct impact on the shower development (rare particle or not forward), it will change model parameters and the extrapolation (in energy or phase space)





String Fragmentation



- Common hadronization in all the models
- Parameters fixed on e+-e- only in EPOS
 - Other CR models tuned on p-p data
 - "Contamination" by beam remnant
- Very important for forward particle production (EAS)

Annihilation at high energy

Used for beam remnant hadronization



Used in dilute systems = CORONA







Generic "EPOS"

First attempt using theoretical constraints

Impose isospin symmetry (u=d) for pions, ρ s and nucleons BEFORE decay

- Fix ρ^0 and multiplicity



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Core-Corona

- Core hadronization = thermal hadronization of Quark Gluon Plasma
- Mixing of core and corona hadronization needed to achieve detailed description of p-p data (ref K.Werner)
 - Evolution of particle ratios from pp to PbPb
 - Particle correlations (ridge, Bose Einstein correlations)
 - Pt evolution, …

Both hadronizations are universal but the fraction of each change with particle density







Hadronic Rescattering (HS)

Missing effect in all CR models until now !

- Re-interaction of hadrons after parton hadronization (space-time evolution)
- "traditionally" used only for heavy ion collisions (until recently NOT in p-p)
- No direct impact on EAS development since forward particles escape
- But significant to large impact at midrapidity in heavy ion collisions !

Let's apply it to all system (from e+-e- to PbPb) !



Example with protons in p-p and Pb-Pb @ LHC



Example with protons in p-p and Pb-Pb @ LHC



Introduction Global approach Impact of HS X_{max} and μ Muon puzzle

Example with Lambda particle in p-p and Pb-Pb @ LHC



Example with Lambda particle in π -Air @ all energies



Impact on air showers

Changes applying core-corona (same results with hadronic rescattering)

- Same X_{max}
- Increase of the number of muons by up to 5% ... only !



Impact on air showers

Changes applying core-corona (same results with hadronic rescattering)

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- But what about e+-e- results after applying hadronic rescattering ?



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Impact of HS on light systems

If short hadronization time (<1fm/c), particles close enough to interact

Impact of HS

- Small but significant effect even in e+e- interactions
- Reduce ρs and nucleons and increase pions

Global approach



Introduction

Impact of HS

 m_{m} and μ

Impact of HS on light systems

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Retune basic parameters with HS

Impact of HS

Second attempt using experimental constraints

- Keep symmetry for pions and nucleons but allow asymmetry for ρ (higher mass)
- Increase contribution of ρs and nucleon to compensate the effect of HS



👄 EPOS LHC-R

Global approach

Introduction

Global approach Impact of HS

 $\sum_{m=1}^{\infty}$ and μ

Retune basic parameters with HS

Second attempt using experimental constraints

- Keep symmetry for pions and nucleons but allow asymmetry for ρ (higher mass)
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EPOS LHC-R

Impact of HS

Check ALICE data



Impact of HS

Impact on air showers (2)

Changes applying new tune taking into account hadronic rescattering

- Same X_{max} (not change applied to cross-section, multiplicity or elasticity)
- Increase the number of muons by ... 30 to 50% (different slope) !

Mostly due to asymmetry between charge and neutral ps !



Introduction Global approach

Impact of HS

(____and)

Other improvements in EPOS LHC-R

Number of limitations identified and solved compared in EPOS LHC

- Problem with nuclear fragments solved
 - Fluctuations of X_{max} for iron similar to others
- No more artificial symmetry neutron and proton
- Pion exchange and real Pomeron exchange
 - LHCf data
- Charm production
 - IceCube
- Indirect impact of core-corona (multiplicity) and hadronic rescattering (shape in pseudorapidity)
 - Higher elasticity due to smaller light cone momenta (see Sergey's talk)
- Lower cross-sections







max

Global changes

- Consequence of retuning, now EPOS shifted by +20 to 30 g/cm²
 - Still in full agreement with accelerator data !

X

Same elongation rate than QGSJETII-04 for protons



N_μ

Global changes

- Consequence of retuning, now EPOS shifted by +20 g/cm²
- Increase of the number of muons by about 10%
- Change in muon spectrum !



x 10 ⁴

Ε_μ

First simulations with full collective effect implementation:

Simulations without core-corona but ρ asymmetry already have more muons

Parallel shift changing all muon energies

- Pion-Air multiplicity impact muon energy between 10 and 100 GeV
- Better tune of kaons (indirect impact of core-corona)

Increase >100 GeV muons (Ice-Top/Ice-Cube)



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X_{max} and μ

First simulations with full collective effect implementation:

- Simulations without core-corona but ρ asymmetry already have more muons

E_u

Parallel shift changing all muon energies

- Pion-Air multiplicity impact muon energy between 10 and 100 GeV
- Better tune of kaons (indirect impact of core-corona)
- Very high energy muons from charm ! (background for neutrino analysis)



Muon Puzzle Solved ?

EPOS LHC-R, first model producing a deeper X_{max} and more muons and being compatible with measured accelerator data (better at LHC) :

- \blacksquare Deeper X_{max} give larger <InA> reducing the gap with measured muon content
- Increase of muons due to tuning taking into account collective effects further decrease the gap to reach Auger systematics
- → What about low energy ? Correlation Ne-N μ OK because of deeper X_{max} !



Impact of HS

Why?

Hadronic rescattering is important to tune properly the models !

- Change ratio between π and ρ in string fragmentation depending on phase-space
 - Forward particle production not the same than at mid-rapidity
- As seen before, if the effect is not taken into account
 - Either overestimate production compared to data ("bad tune")

Sibyll*

If ρ^0 or underestimate forward production of ρ^0 to get it right with mid-rapidity data



Impact of HS

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All models until now !

Outlook

- Updated results of cross-sections, multiplicity and diffraction
 Large impact on X_{max}
 - Larger <InA> (heavier primary mass → reduce "muon puzzle")
- Details of hadronization matters
 - Important role of resonances
 - \bullet ρ^0 impacted by hadronic rescattering, important to take it into account
 - Evolution of strangeness with multiplicity
 - Different type of hadronization in core = more muons

Combination of the 3 effects may solve the muon puzzle (to be confirmed) !

- Source of muon puzzle probably due to the fact that <u>hadron rescattering</u> was always neglected
 - Rescattering change the correlation between mid-rapidity (data and tuning) and forward particle production (EAS)

Updated EPOS LHC-R released in 2024 and then adapting EPOS 4 for CR

Recent LHC data provide new constraints on models, changing X_{max} and the muon production if a global approach is used.

Providing a possible solution to the "muon puzzle" !

Thank you !

Inelastic Cross-Section

- Probability for the particle to interact : directly related to X_{max}
- After TOTEM (CMS), new measurements by ALFA (ATLAS) with higher precision



Cross-Section Reduced

- Probability for the particle to interact : directly related to X_{max}
- After TOTEM (CMS), new measurements by ALFA (ATLAS) with higher precision
 - p-p cross-section slightly too high in all models
 - Change by up to -10% at the highest energy





Pseudorapidity

- Angular distribution of newly produced particles
- New data at 13 TeV in p-p
 - Test extrapolation with different triggers
 - Sibyll has a clear difference with other models (and data) : too narrow !
- Detailed data at 5 TeV for p-Pb
 - Wrong multiplicity distributions in all models (before retune)



Impact of HS

___ and μ

Muon puzzle

Improvements in EPOS LHC-R

- Number of limitations identified in EPOS LHC
- Problem with nuclear fragments
 - Double counting for single nucleons
 - Missing multifragment production
 - Now similar to other models
 - Significant impact on X_{max} fluctuations for nuclei
- Simplified high mass diffraction and pion \circ exchange replaced by real emission (IP or π)





EPOS LHC-R interaction with Air

(preliminary)



Hadronization in Simulations

- Historically (theoretical/practical reasons) string fragmentation used in high energy models (Pythia, Sibyll, QGSJET, ...) for proton-proton.
 - Light system are not "dense"
 - Works relatively well at SPS (low energy)
 - ➡ But problems already at RHIC, clearly at Fermilab, and serious at LHC :
 - Modification of string fragmentation needed to account for data
 - Various phenomenological approaches :
 - Color reconnection
 - String junction
 - ✤ String percolation, …
 - Number of parameters increased with the quality of data ...
- Statistical model only used for heavy ion (HI) in combination with hydrodynamical evolution of the dense system : QGP hadronization
 - Account for flow effects, strangeness enhancement, particle correlations...

Core-Corona appoach and CR

To test if a QGP like hadronization can account for the missing muon production in EAS simulations a core-corona approach can be artificially apply to any model

- Particle ratios from statistical model are known (tuned to PbPb) and fixed : core
- Initial particle ratios given by individual hadronic interaction models : corona
- → Using CONEX, EAS can be simulated mixing corona hadronization with an arbitrary fraction ω_{core} of core hadronization: $N_i = \omega_{core} N_i^{core} + (1 \omega_{core}) N_i^{corona}$



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Introduction

Constraints from Correlated Change

- One needs to change energy dependence of muon production by ~+4%
- To reduce muon discrepancy
 β has to be change
 - X_{max} alone (composition) will not change the energy evolution
 - β changes the muon energy evolution but not X_{max}

•
$$\beta = \frac{\ln(N_{mult} - N_{\pi^0})}{\ln(N_{mult})} = 1 + \frac{\ln(1 - \alpha)}{\ln(N_{mult})}$$

• +4% for β -> -30% for $\alpha = \frac{N_{\pi^0}}{N_{mult}}$

$$N_{\mu} = A^{1-\beta} \left(\frac{E}{E_0}\right)^{\beta}$$

 $X_{max} \sim \lambda_e \ln \left(E_0 / (2.N_{mult} \cdot A) \right) + \lambda_{ine}$





17%

 10^{3}

 10^{4}

 10^{5}

 10^{6}

 10^{7}

 10^{8}

 10^{2}

 10^{8}

 10^{9}

E (GeV)

 10^{10}

 10^{7}

105

 10^{3}

 10^{2}

 10^{4}

 10^{6}

0.40

 $\begin{array}{c} \mathbf{R} = \mathbf{E}_{em} / \mathbf{E}_{had} \\ = \mathbf{B}_{em} / \mathbf{E}_{had} \\ = \mathbf{0.30} \end{array}$

0.25

0.20

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 10^{10}

Plot by M. Perlin

16%

 10^{9}

E (GeV)

Evolution of hadronization from core to corona

The relative fraction of π^0 depends on the hadronization scheme \Rightarrow Change of ω_{core} with energy change $\alpha = \frac{N_{\pi^0}}{N_{\text{mult}}}$ or $R(\eta) = \frac{\langle dE_{\text{em}}/d\eta \rangle}{\langle dE_{\text{had}}/d\eta \rangle}$

which define the muon production in air showers.



Evolution of hadronization from core to corona



Possible Particle Physics Explanations

A 30% change in particle charge ratio ($\alpha = \frac{N_{\pi^0}}{N_{mult}}$) is huge ! \rightarrow Possibility to increase N_{mult} limited by X_{max}

- New Physics ?
 - Chiral symmetry restoration (Farrar et al.) ?
 - Strange fireball (Anchordoqui et al., Julien Manshanden) ?
 - String Fusion (Alvarez-Muniz et al.) ?

Problem : no strong effect observed at LHC (~10¹⁷ eV)

- Unexpected production of Quark Gluon Plasma (QGP) in light systems observed at the LHC (at least modified hadronization)
 - **Reduced** α is a sign of QGP formation (enhanced strangeness and baryon) production reduces relative π^0 fraction. Baur et al., arXiv:1902.09265)
 - α depends on the hadronization scheme
 - How is it done in hadronic interaction models ?

Impact of HS

LHC acceptance and Phase Space



- p-p data mainly from "central" detectors
 - → pseudorapidity η =-ln(tan(θ /2))
 - \bullet $\theta=0$ is midrapidity
 - \bullet θ >>1 is forward
 - •• $\theta < <1$ is backward
- Different phase space for LHC and air showers
 - most of the particles produced at midrapidity
 - important for models
 - most of the energy carried by forward (backward) particles
 - important for air showers

A 3rd way : the core-corona approach

Consider the local density to hadronize with strings OR with QGP:

First use string fragmentation but modify the usual procedure, since the density of strings will be so high that they cannot possibly decay independently : core

