

Parallelizing radio-interferometric image reconstruction by baselines

Sunrise Wang, Simon Prunet, Shan Mignot, André Ferrari



Outline

- **Introduction**
- **Partitioning image reconstruction by baseline**
- **Parallelizing image reconstruction by baseline**

Introduction

Antenna arrays



SKA-Low

Radio dishes

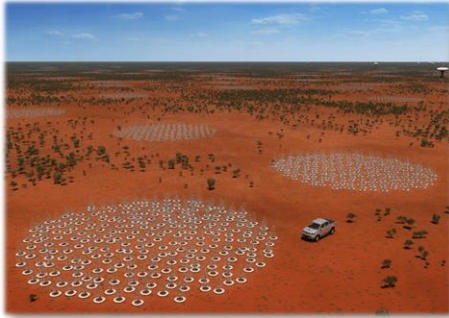


SKA-Mid

Introduction

SKA-Low

Antenna arrays



~0.7 TB/s

Correlator & Beamformer



~0.3 TB/s

Imaging supercomputer: SDP-Low



SKA-Mid

Radio dishes



~2.4 TB/s

Correlator



~1.1 TB/s

Imaging supercomputer: SDP-Mid



image sources: [1, 2, 3, 4, 5]
data source: [6, 7]

[1] Farnes, Jamie, et al. "Science pipelines for the square kilometre array." *Cdaxies* 6.4 (2018): 120.
[2] https://www.kit.edu/kit/en/analysis/ai_2021_059_kit-supercomputer-one-of-the-15-fastest-in-europe.php
[3] <https://www.technologyview.com/2023/09/21/1079909/whats-next-for-the-worlds-fastest-supercomputers/>
[4] <https://www.eso.org/public/france/images/eso1253c>
[5] https://www.atnf.csiro.au/projects/askap/news_commissioning_24022014.html
[6] Labate, Maria G., et al. "Highlights of the square kilometre array low frequency (SKA-Low) telescope." *Journal of Astronomical Telescopes, Instruments, and Systems* 8.1 (2022): 011024-011024.
[7] Swart, Gerhard P., Peter E. Dewdney, and Andrea Cremonini. "Highlights of the SKA-Mid telescope architecture." *Journal of Astronomical Telescopes, Instruments, and Systems* 8.1 (2022): 011021-011021.

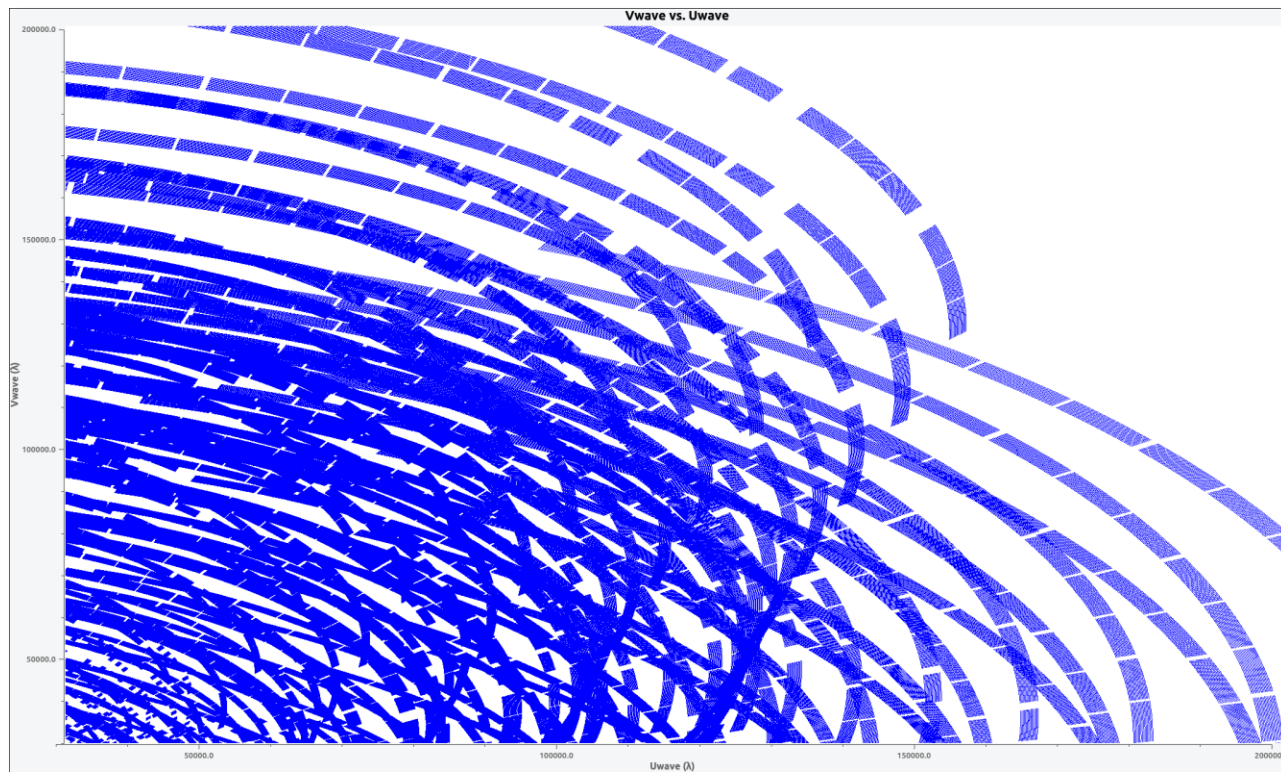
Scaling the computation

- Need to scale the computation, particularly de/gridding
 - Shown to take up to 94% of the time for a serial implementation[1]
- Partition visibilities, process separately

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Partition & Parallelization strategies:

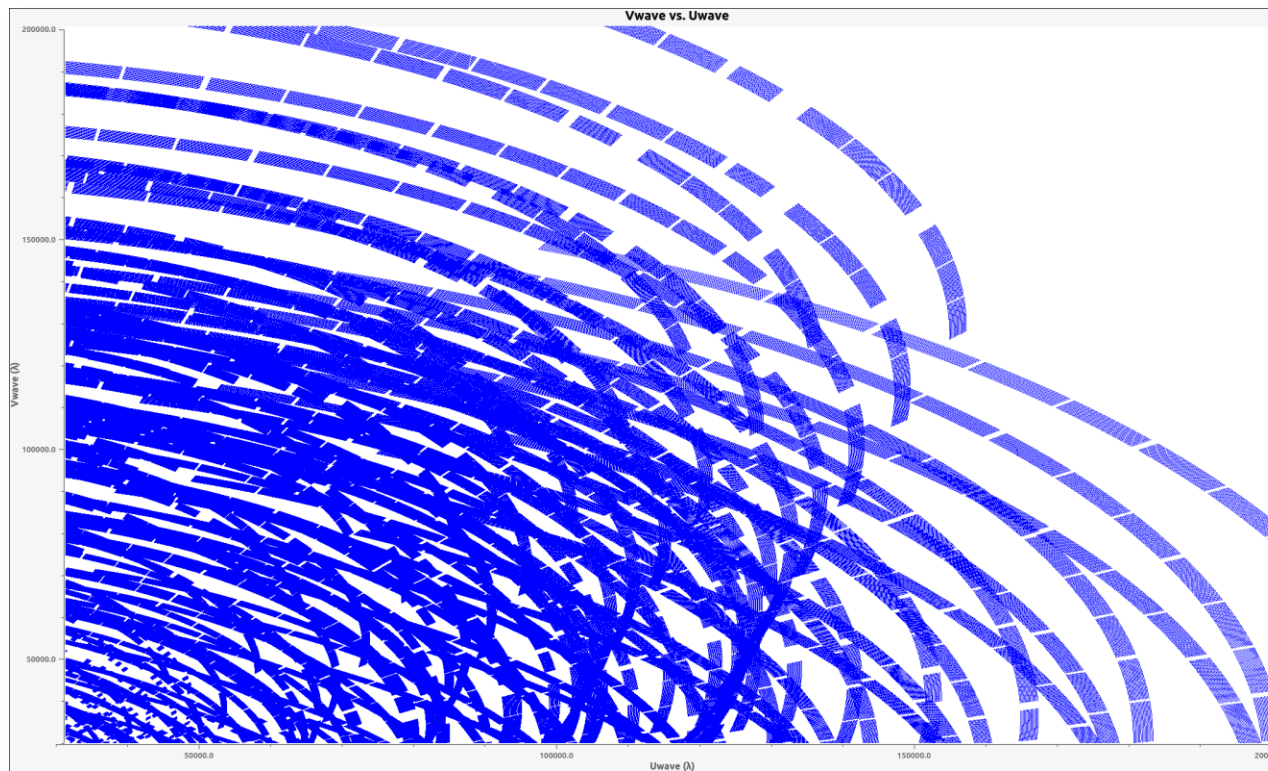


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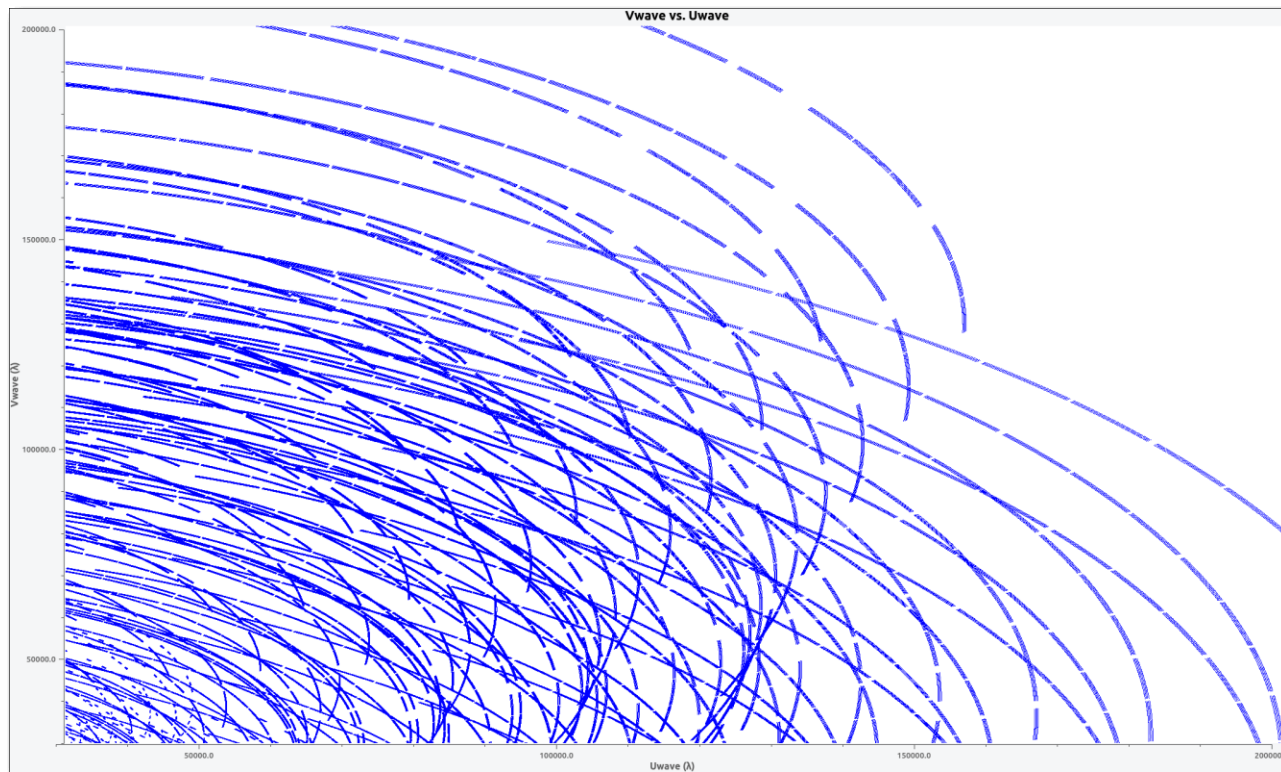


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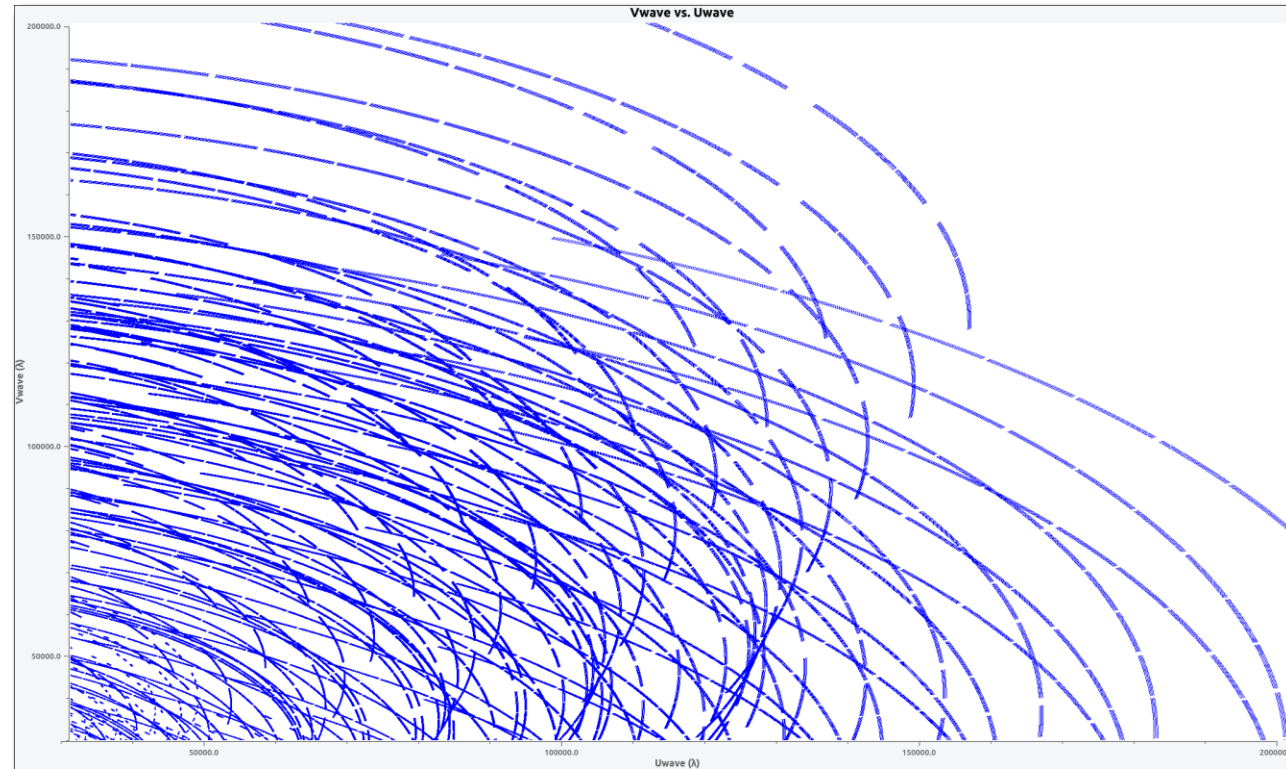


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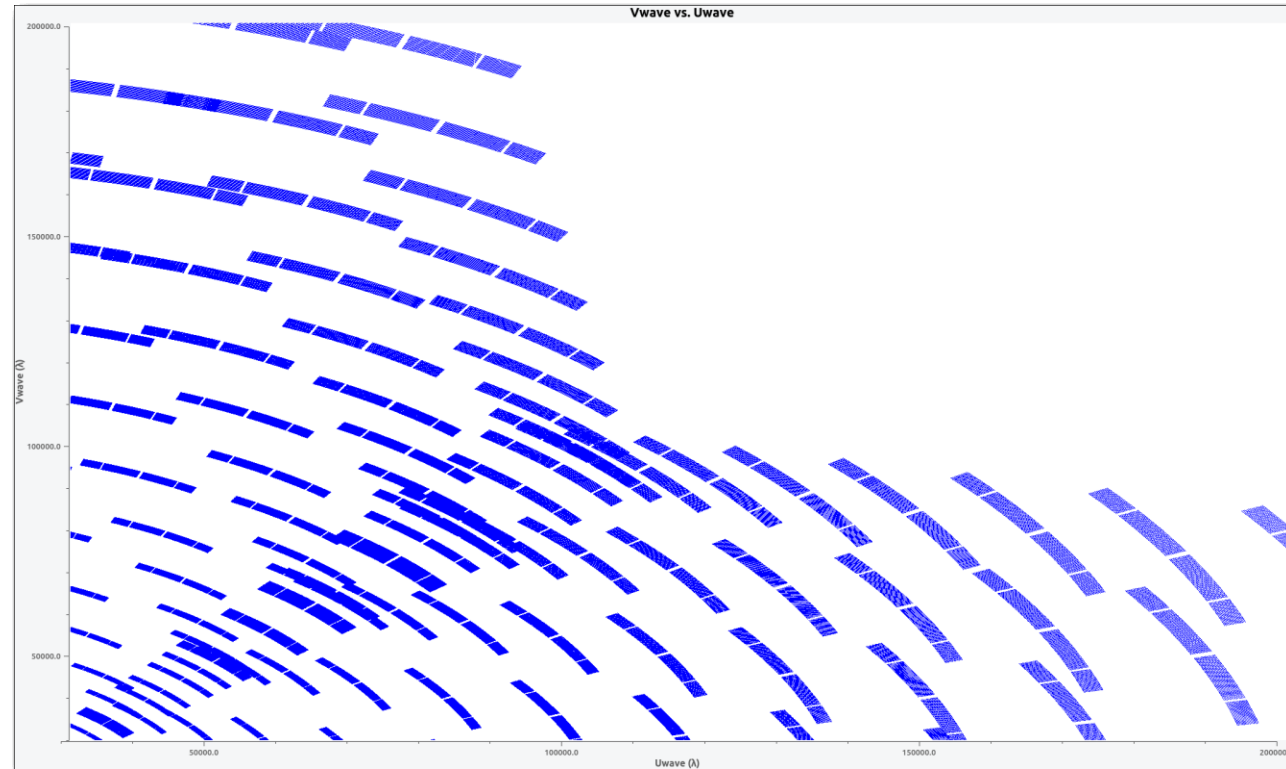


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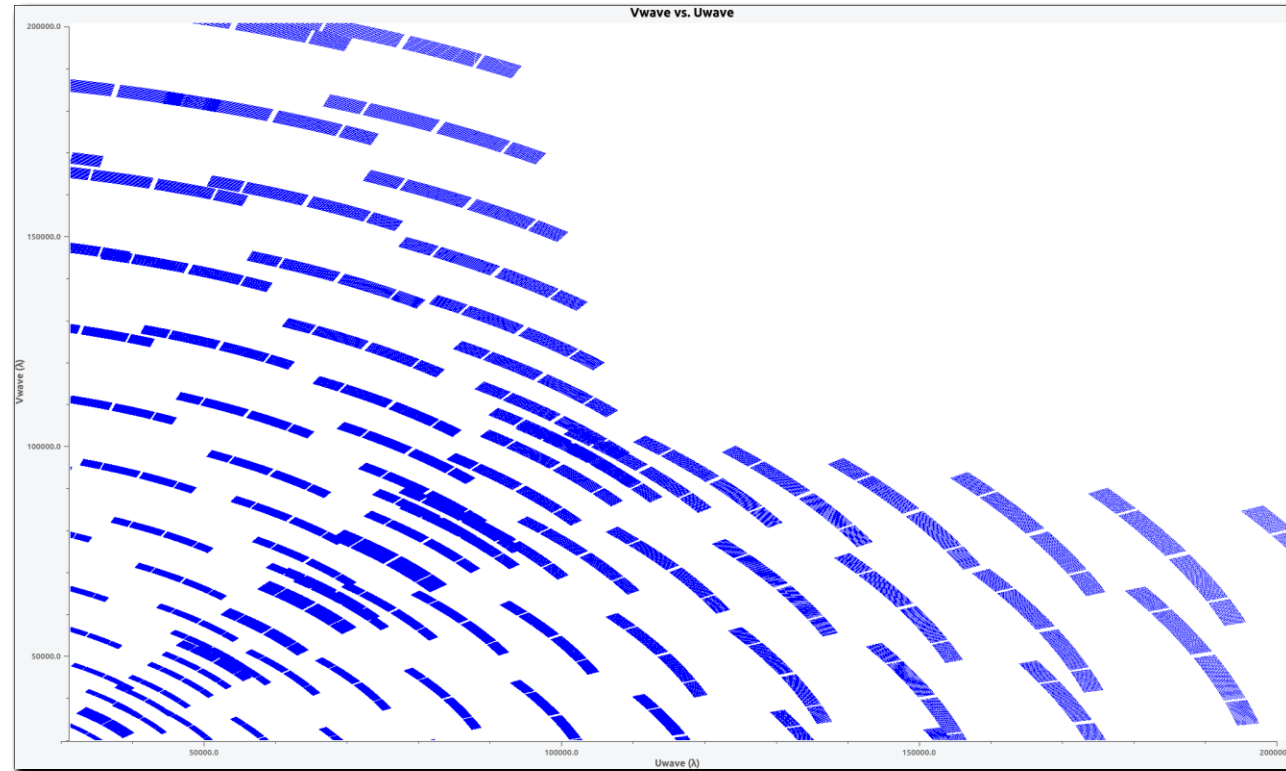


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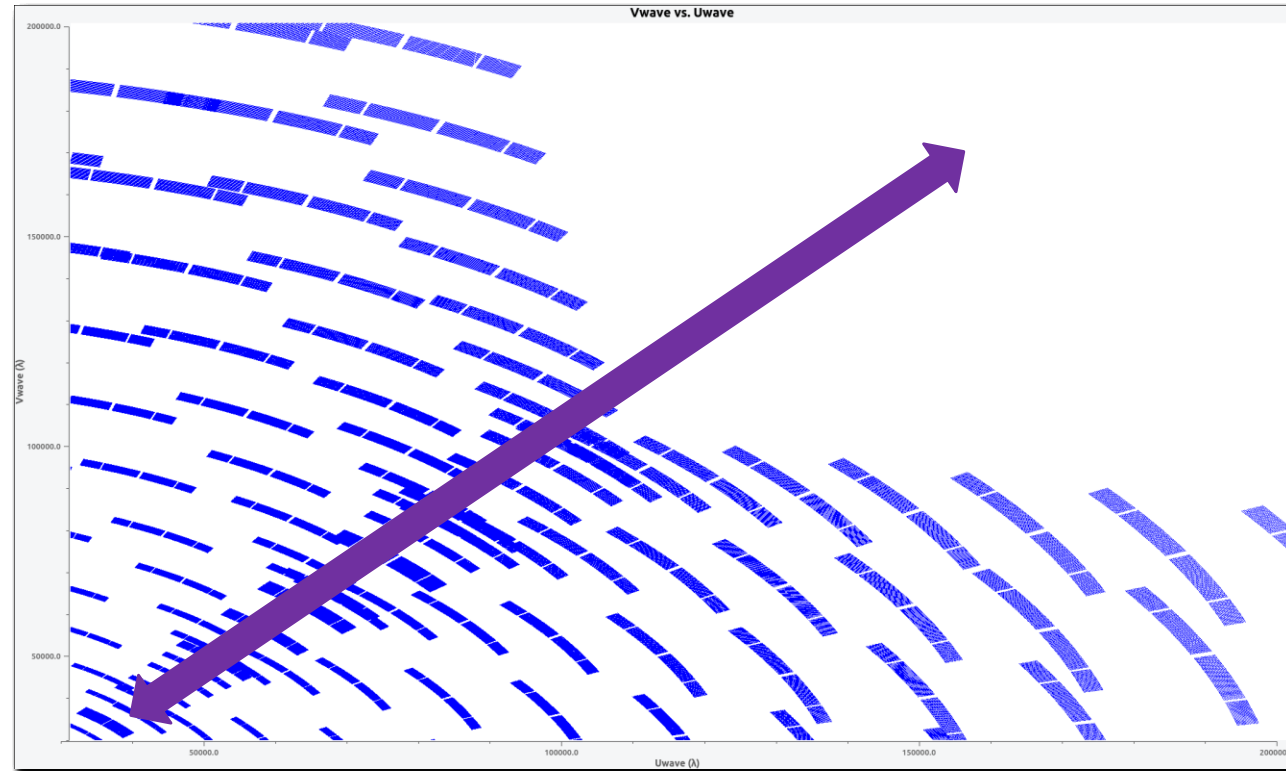


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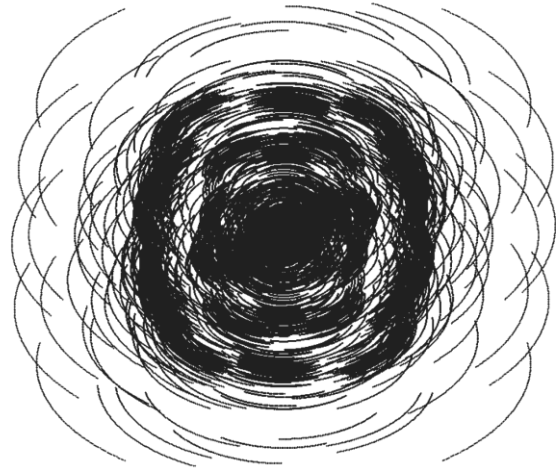
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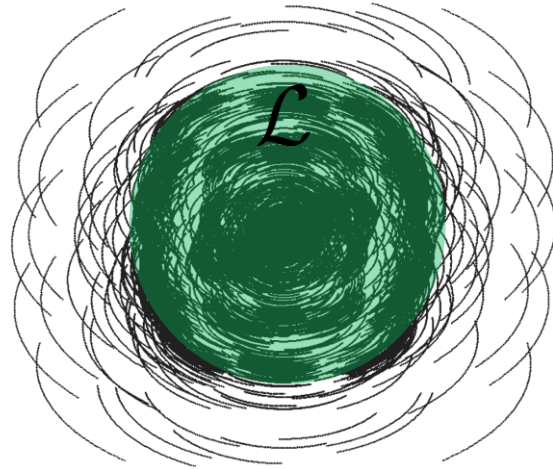
**Partitioning
image
reconstruction
by baseline**

Multi-step image reconstruction: General Framework

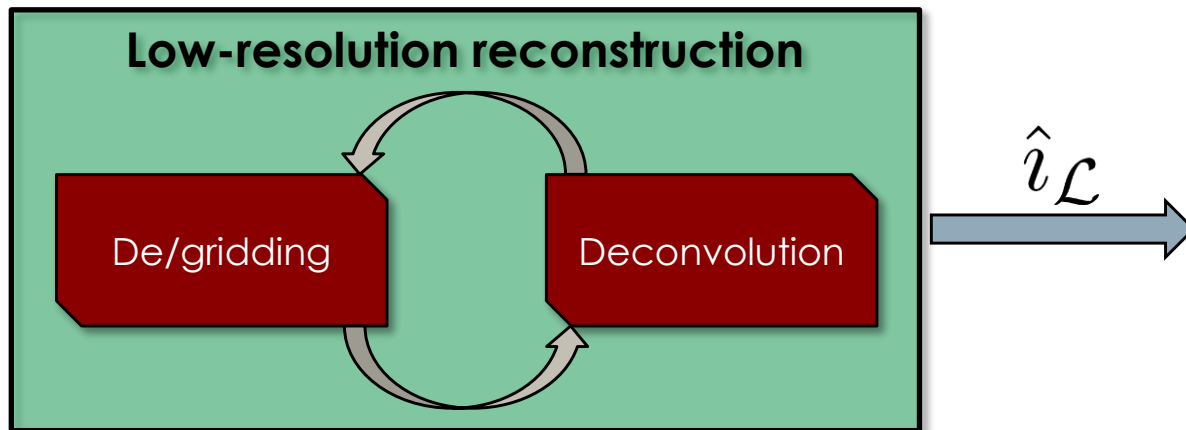
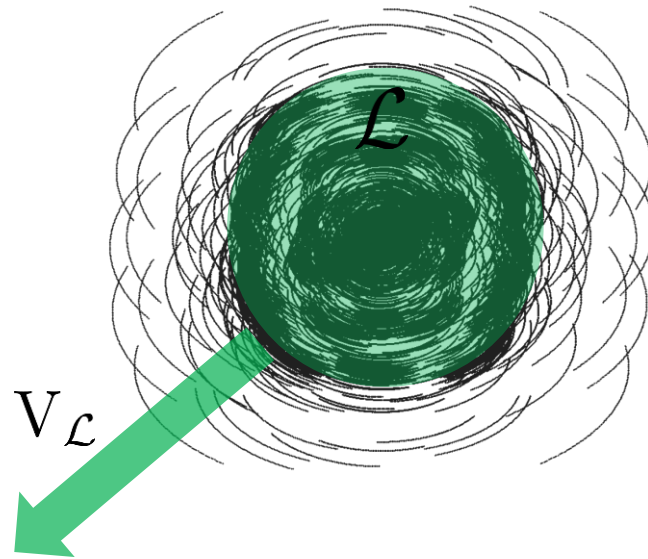
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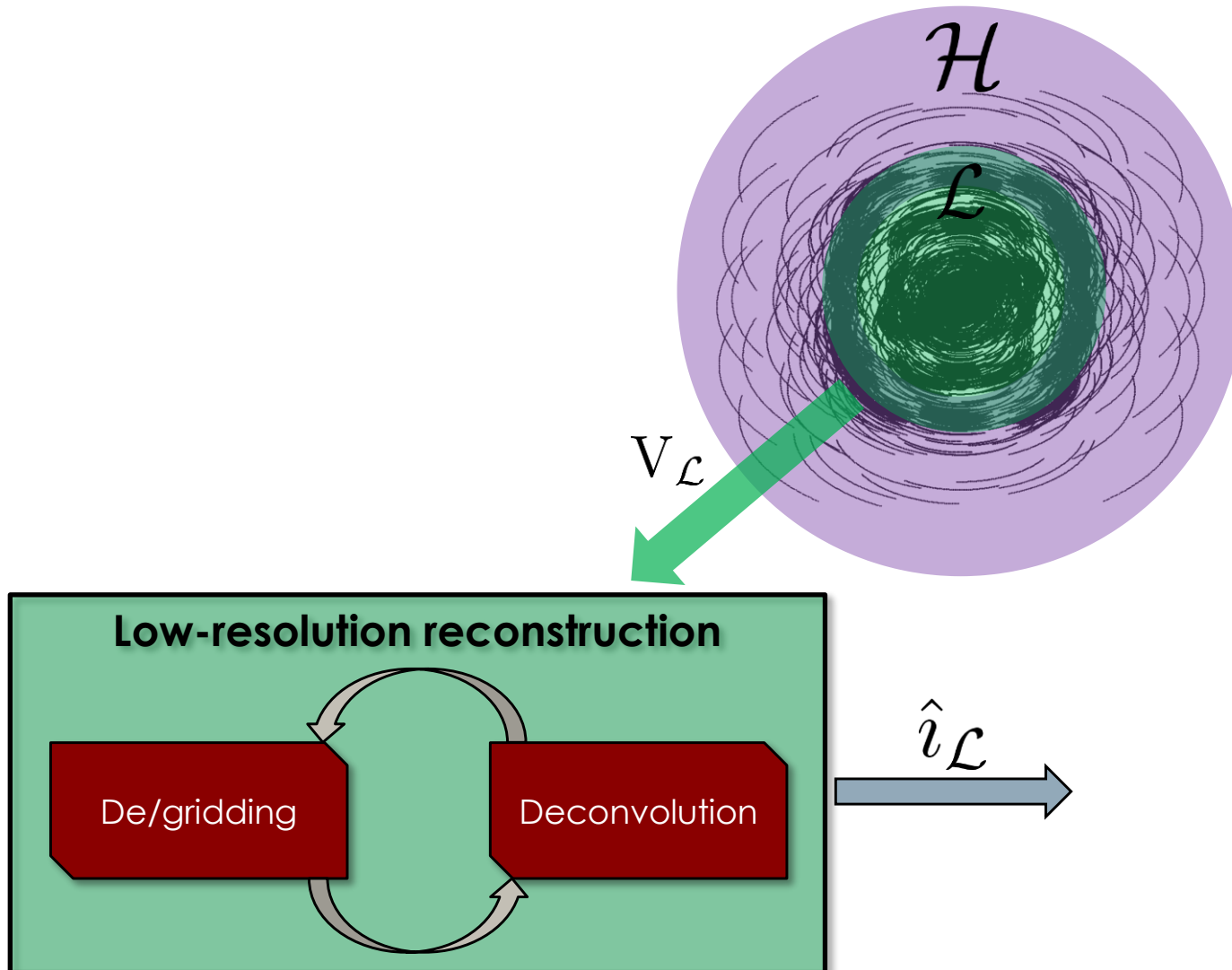
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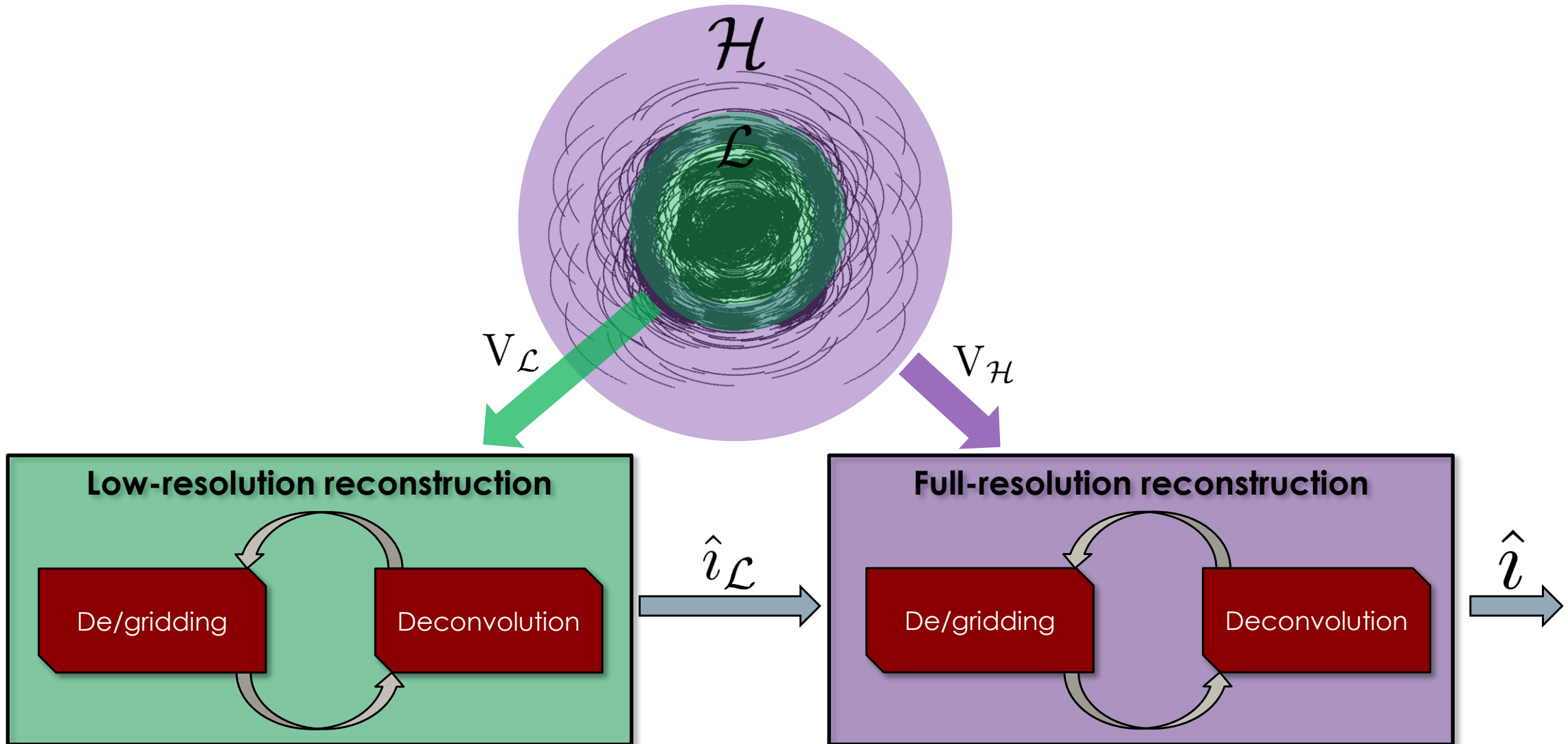
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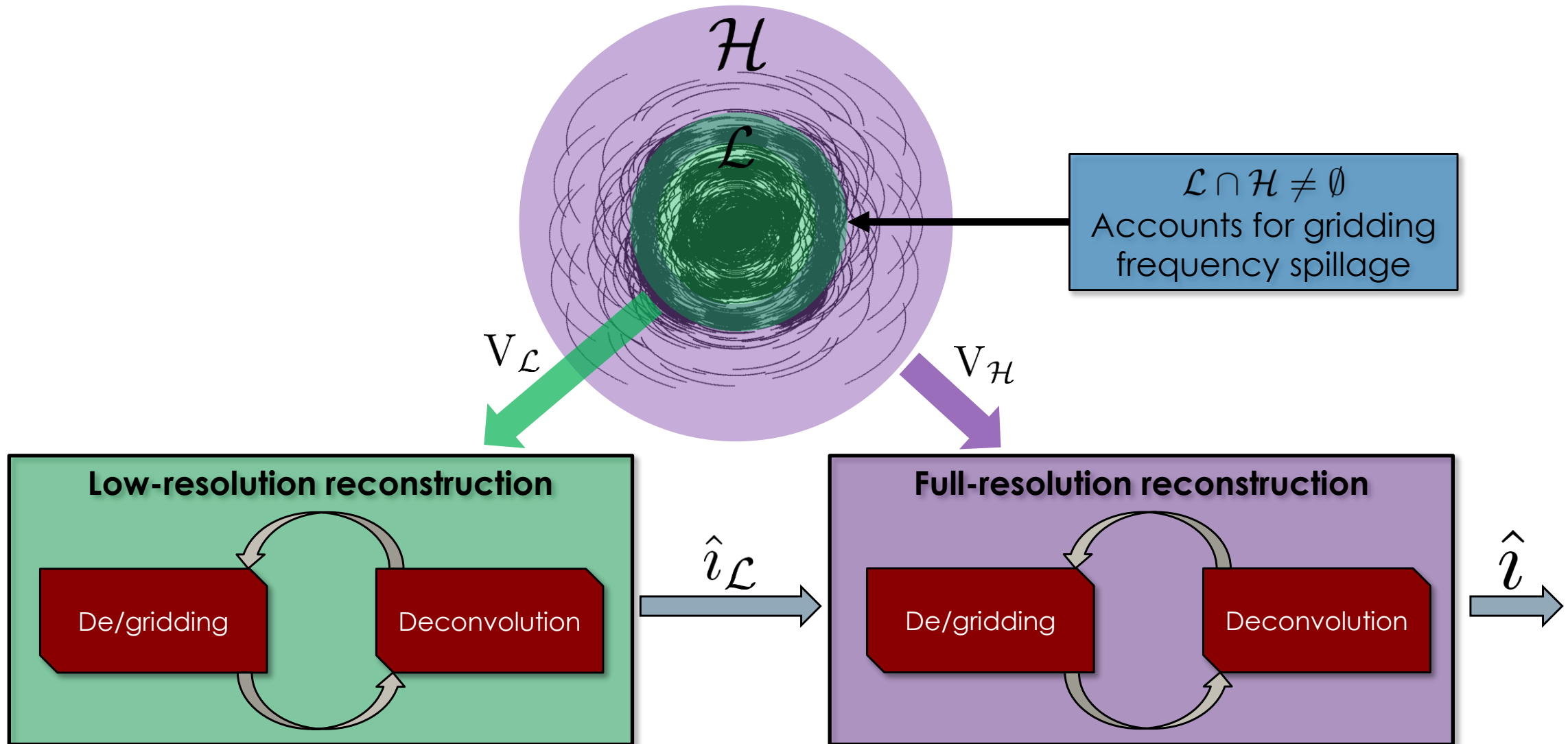
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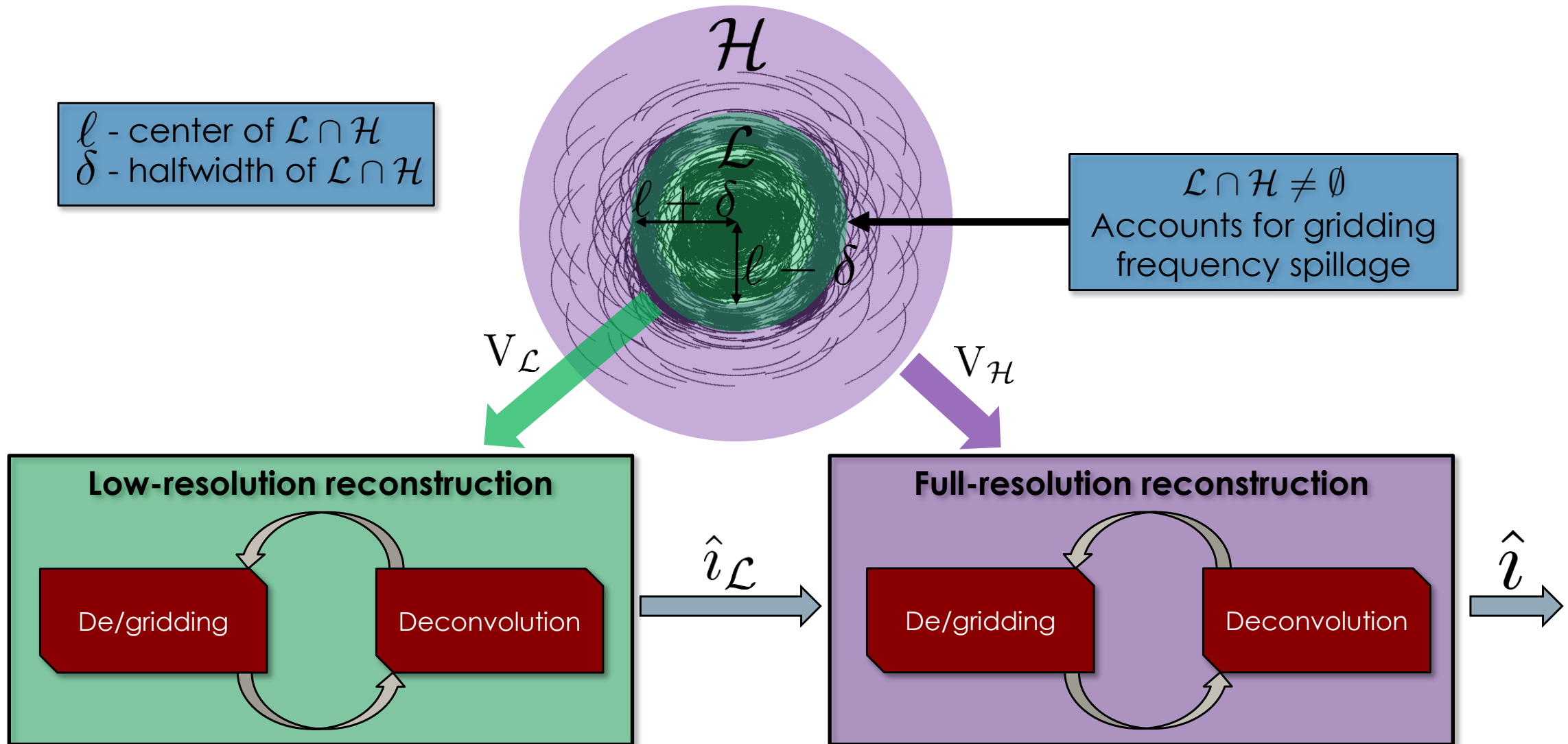
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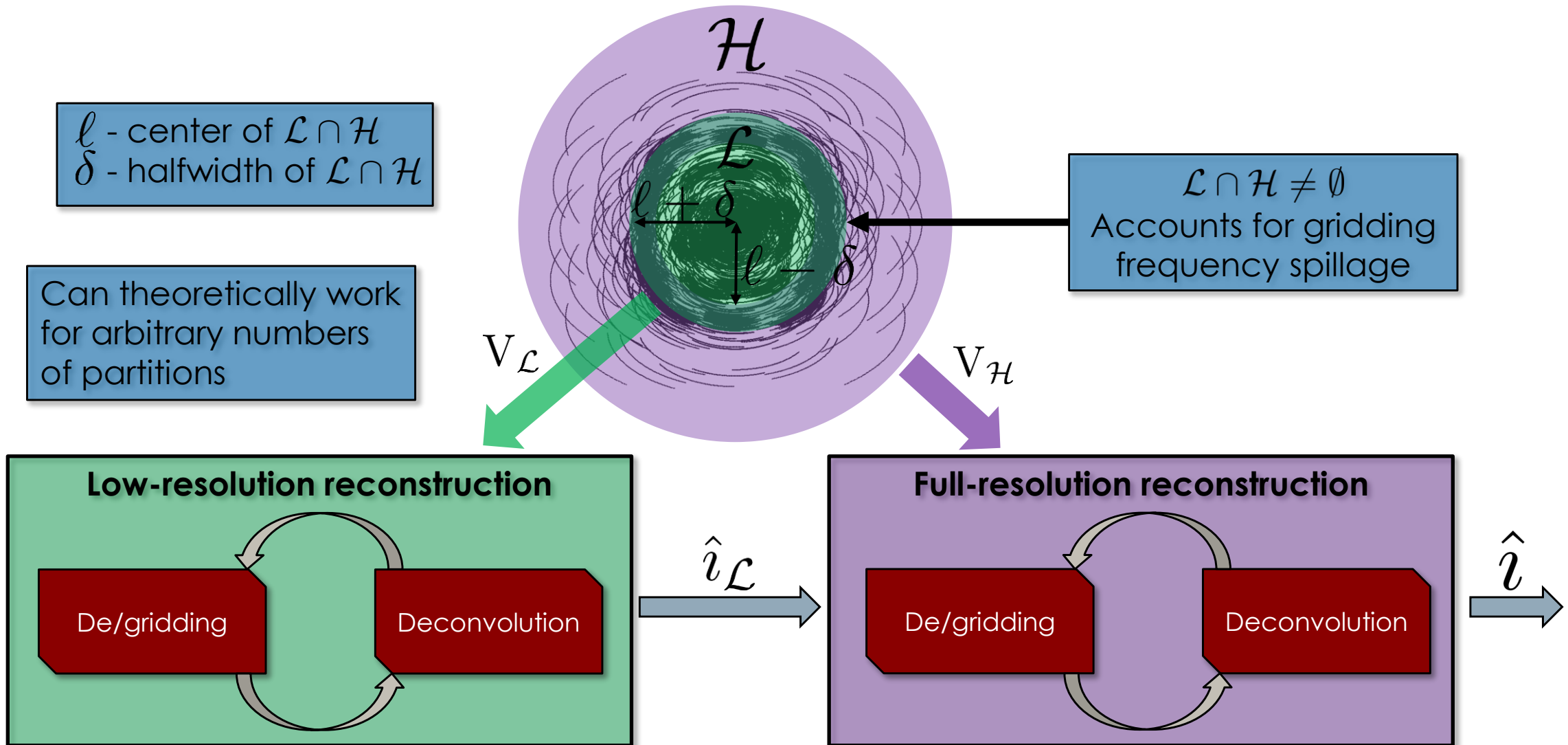
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Multi-step image reconstruction: General Framework



Multi-step image reconstruction: Deconvolution framework

Deconvolution framework for every major cycle n , similar to [1, 2]

$$\alpha_n = \arg \min_{\alpha} \|\tilde{i}_n - HW\alpha\|_2^2 + \lambda_n \|\alpha\|_1$$

$$\bar{i}_n = W\alpha_n$$

Multi-step image reconstruction: Deconvolution framework

n_{th} major cycle residual

$$\tilde{i}_n = F^\dagger G(v - G^\dagger F \sum_{n=1}^N \bar{i}_{\mathcal{L}_n})$$

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Used over $\|v - G^\dagger FW\alpha\|_2^2$ for efficiency
Assumes $Hi \approx F^\dagger GG^\dagger Fi$
Errors corrected in major cycle

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Deconvolved
image

Final reconstructed image: $\hat{i} = \tilde{i}_{N+1} + \sum_{n=1}^N \bar{i}_n$

Multi-step image reconstruction: The low and full-resolution steps

Low-resolution deconvolution

$$\alpha_{\mathcal{L}_n} = \arg \min_{\alpha} \|\tilde{i}_{\mathcal{L}_n} - H_{\mathcal{L}} W \alpha\|_2^2 + \lambda_{\mathcal{L}_n} \|\alpha\|_1$$

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Full resolution deconvolution

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First data fidelity term
contains only high-resolution
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Second data fidelity term

$$l_{\mathcal{L}_n} = \hat{i}_{\mathcal{L}} - \sum_{j=1}^{n-1} \bar{i}_j$$

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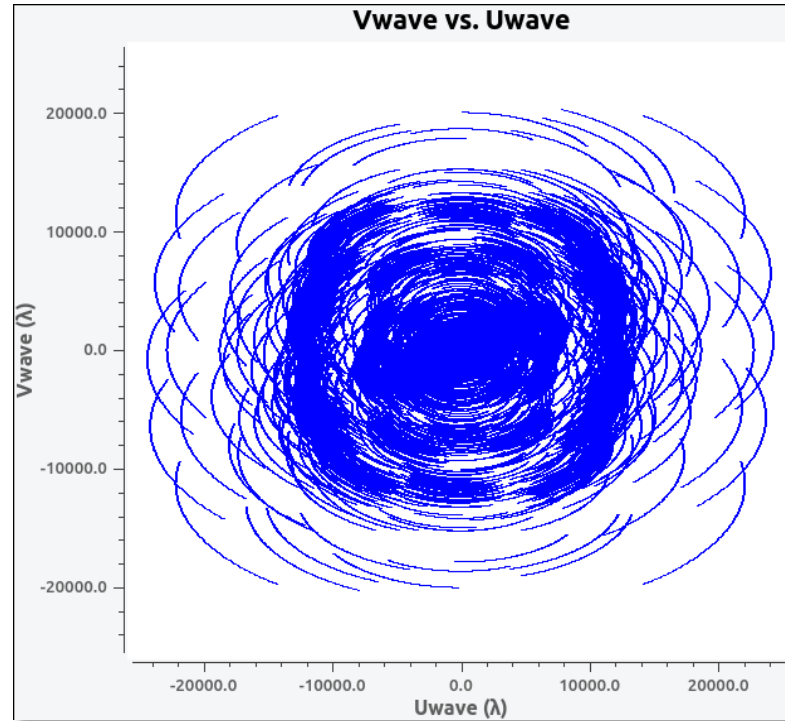
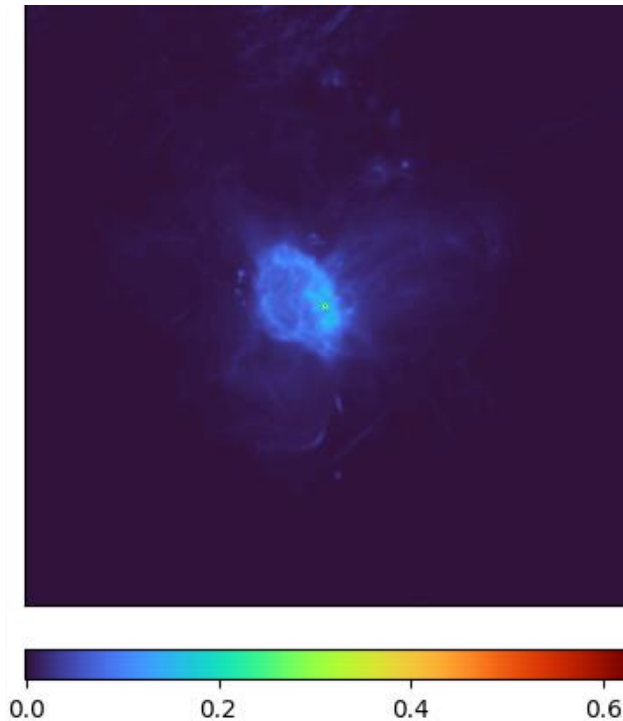
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High and low resolution filters
Normalizes data fidelity terms
Ensures visibilities in $\mathcal{L} \cap \mathcal{H}$ sum to 1

Second data fidelity term
 $l_{\mathcal{L}_n} = \hat{i}_{\mathcal{L}} - \sum_{j=1}^{n-1} \bar{i}_j$

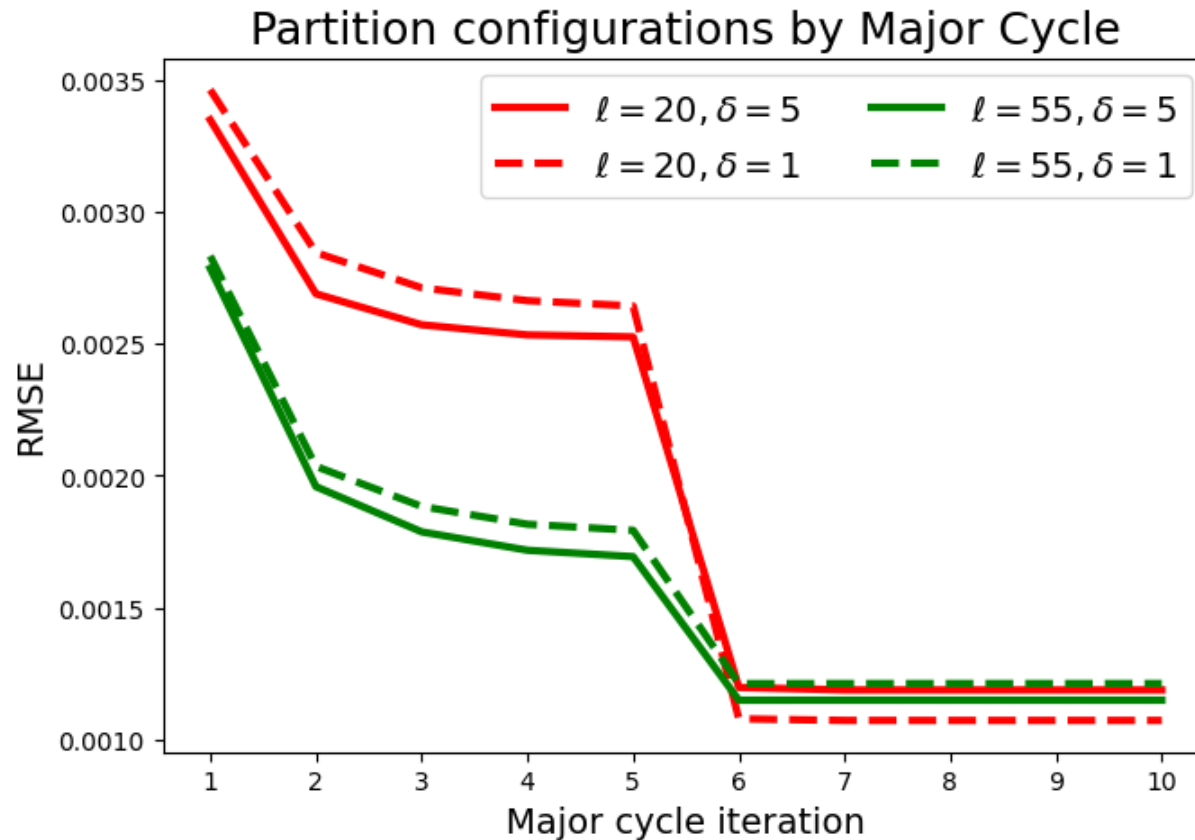
Results – Measurement set



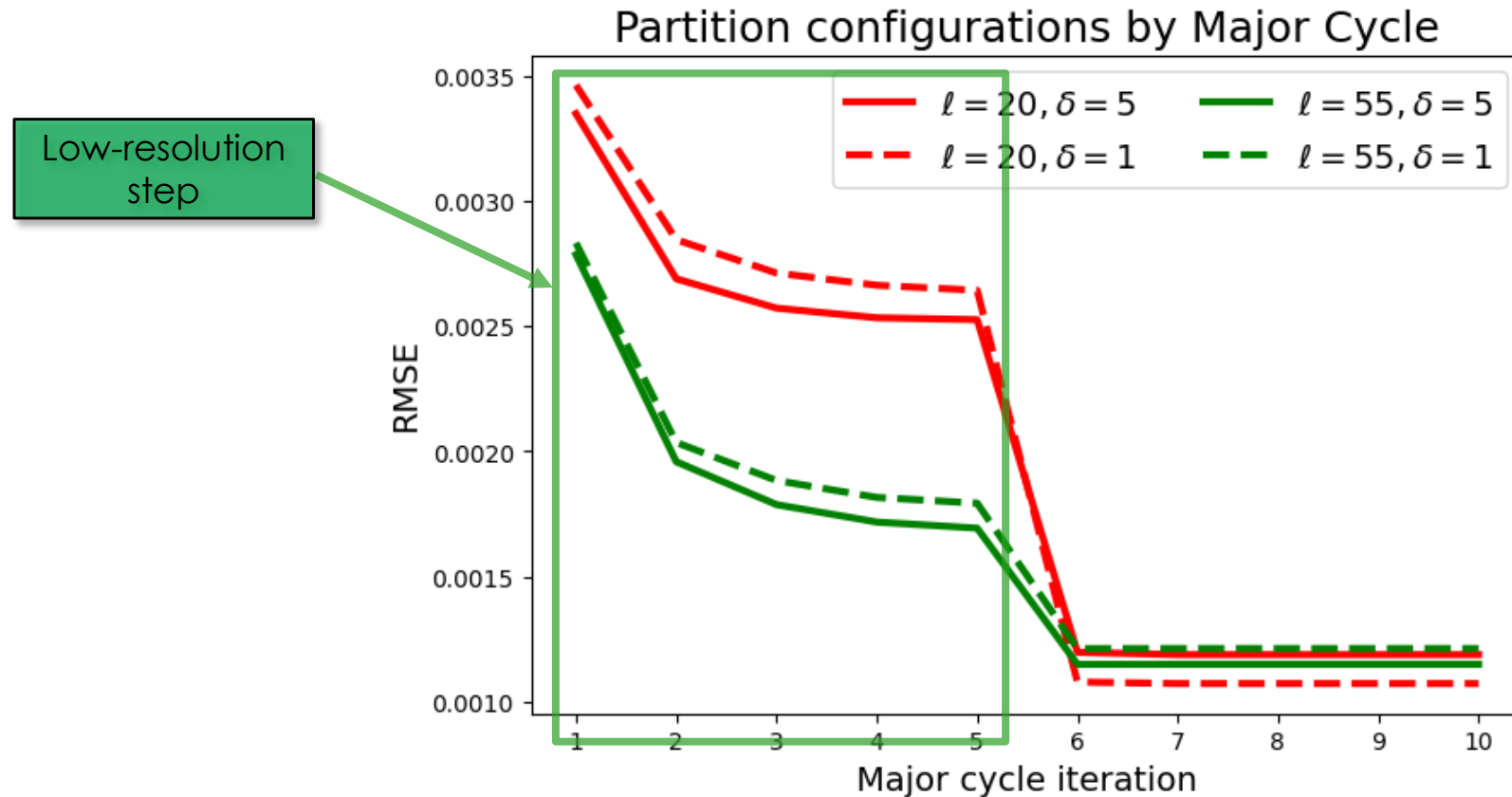
- Initial images tapered and cutout from 1.28GHz mosaic produced in [1]
- Visibilities generated with MeerKAT configuration
- Exposure time of 4h, samples every 120s for to generate visibility positions
- Degrid to get visibility values
- Visibility noise artificially added

Results: Partition configuration affect on reconstruction accuracy and speed

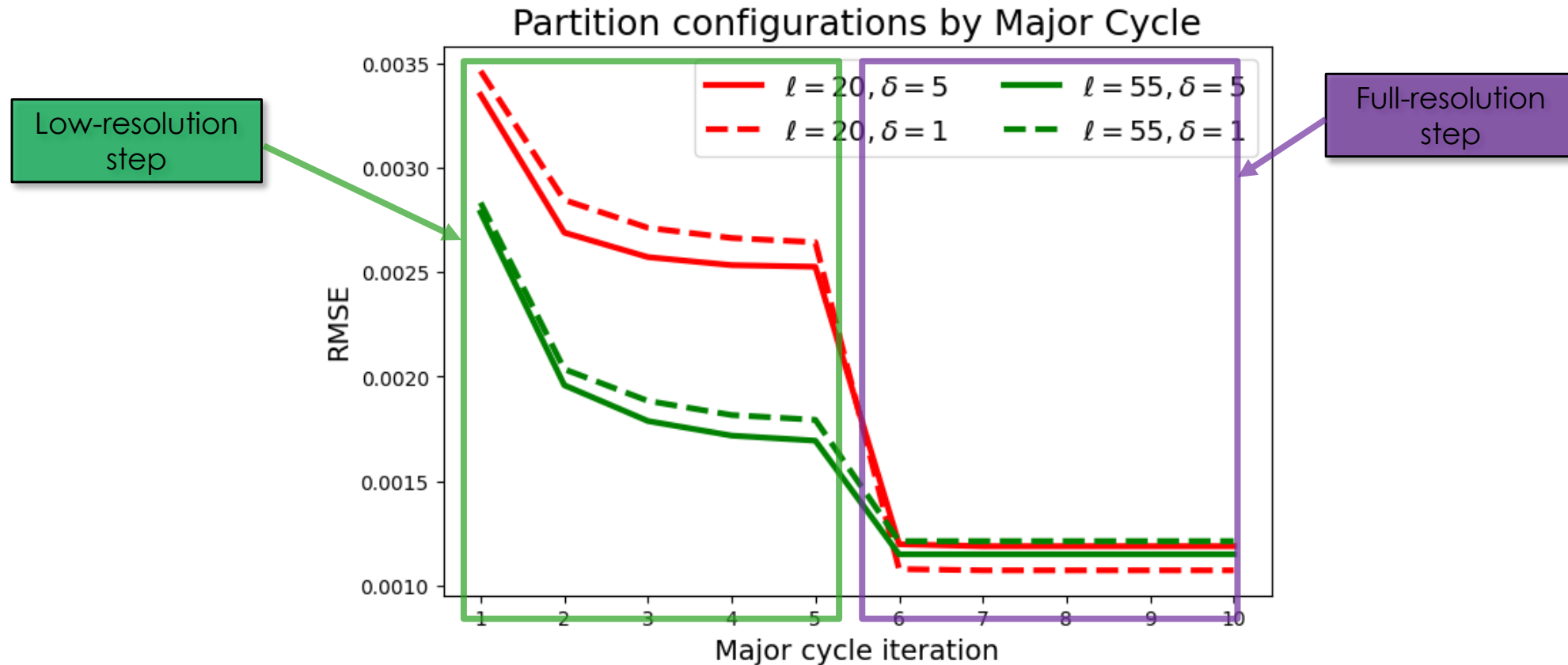
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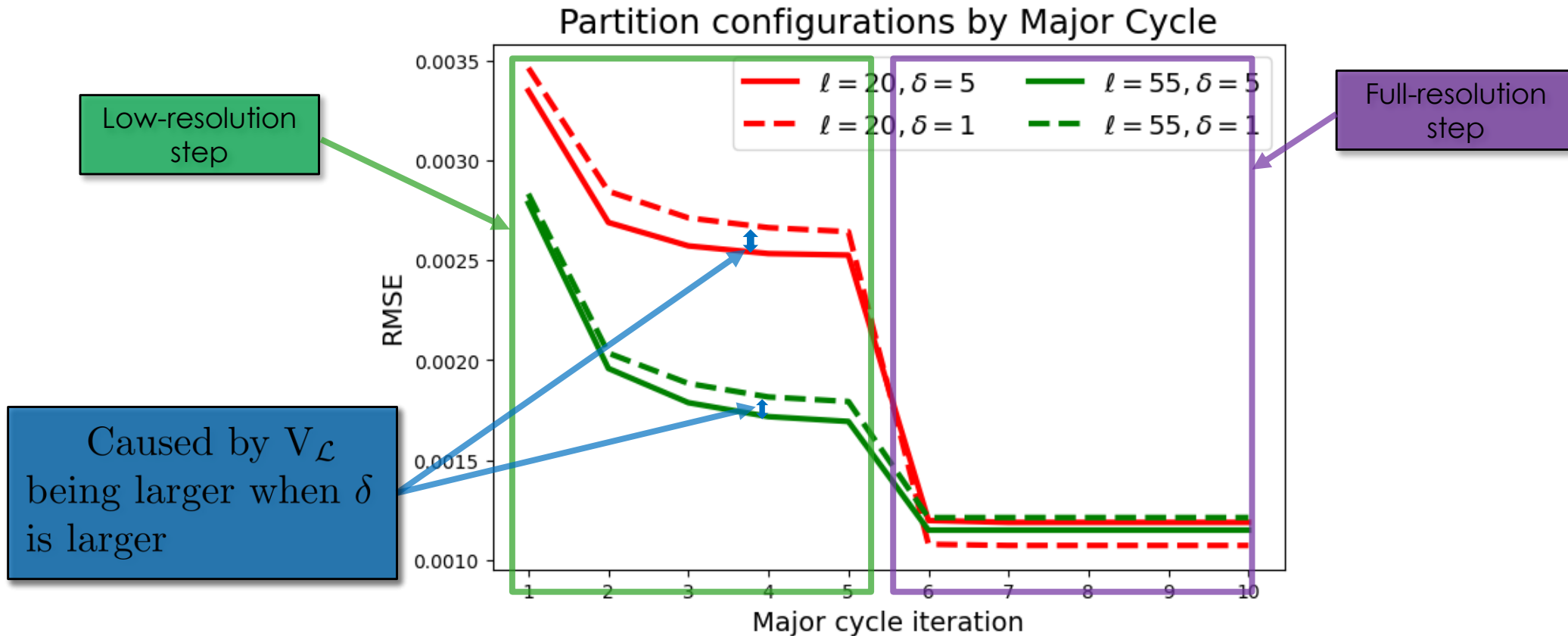
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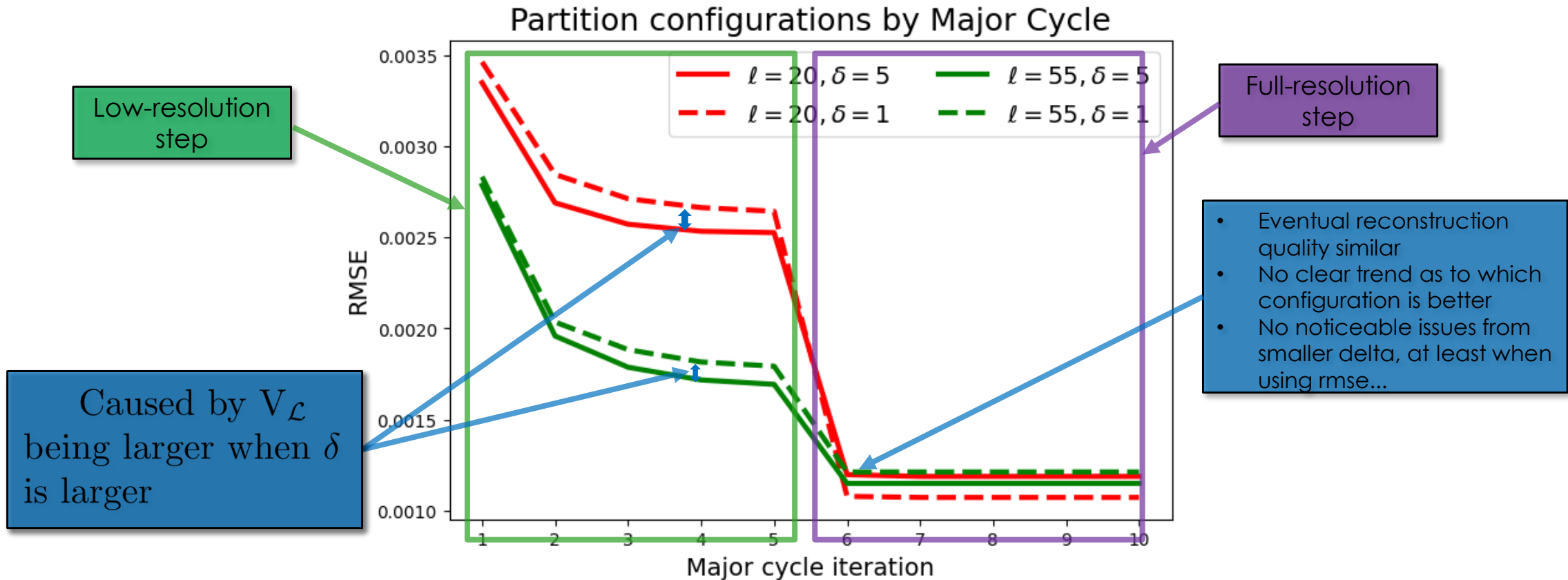
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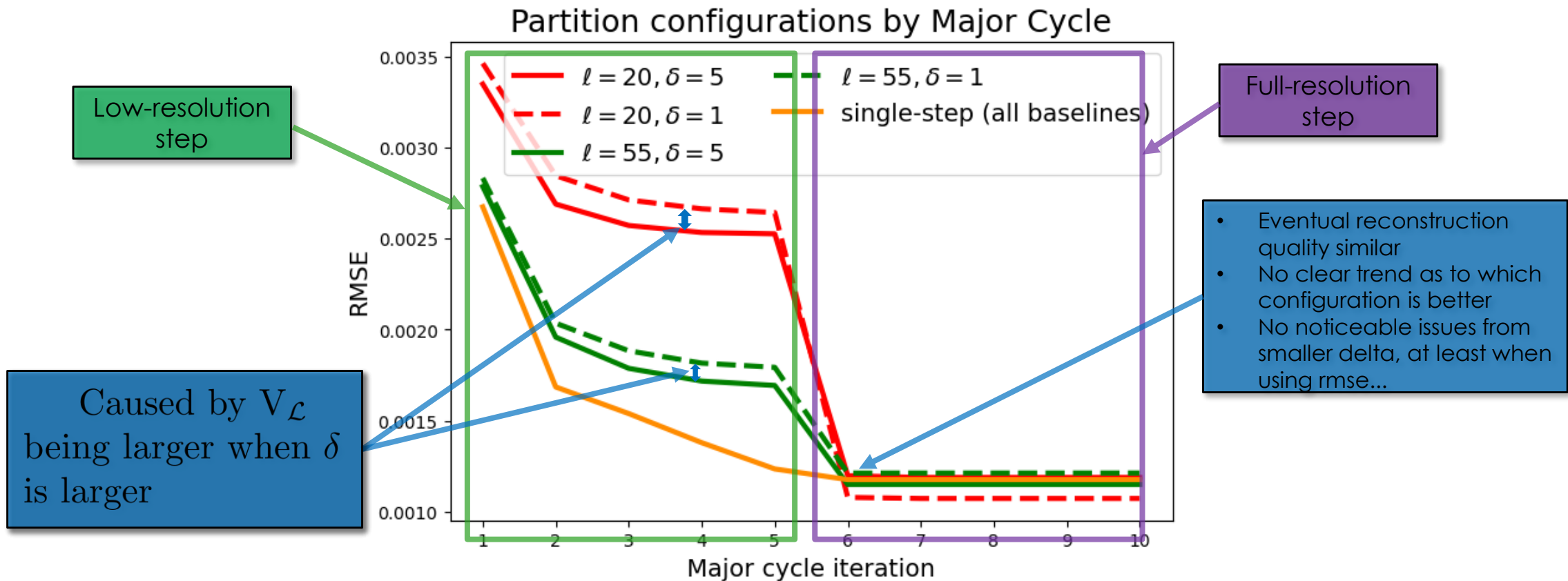
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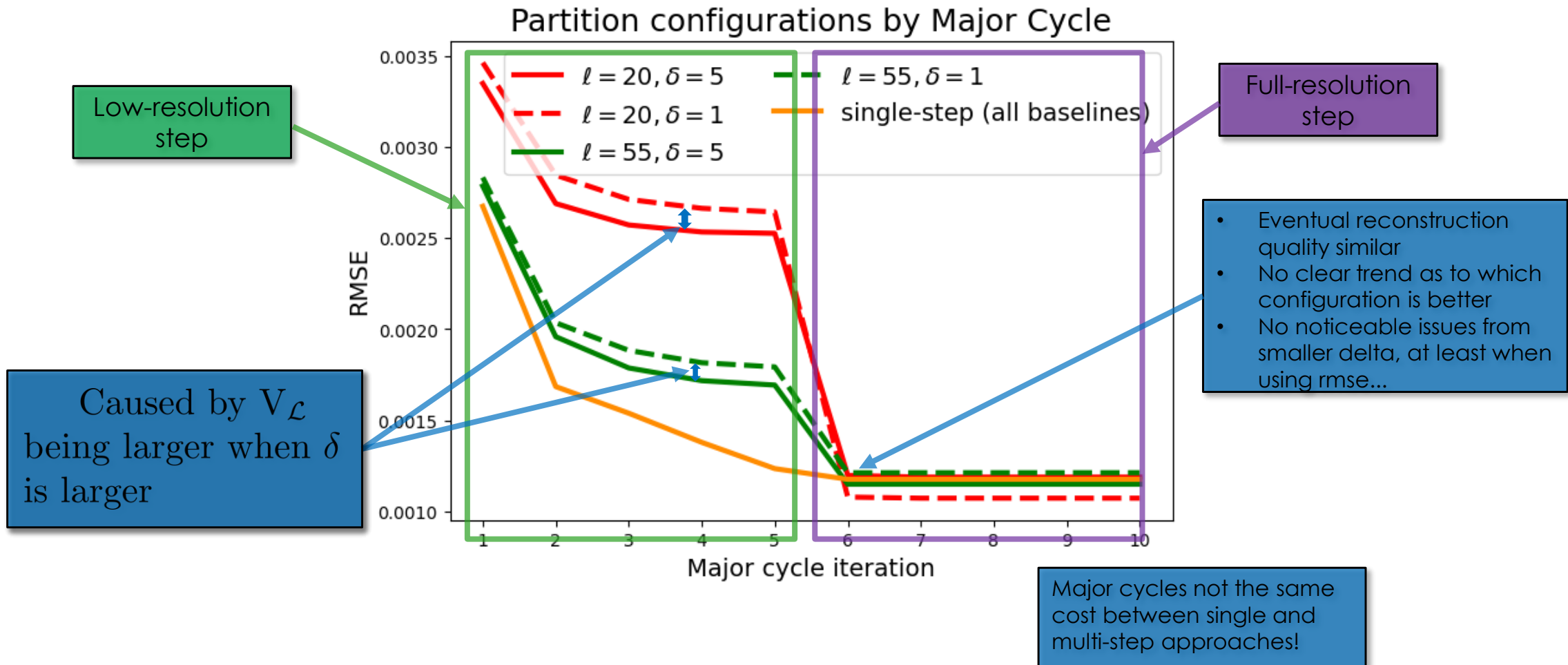
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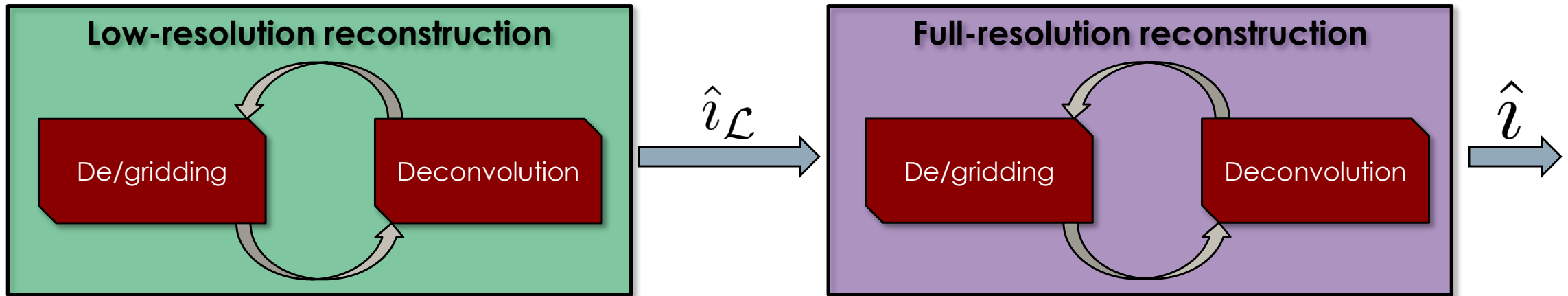


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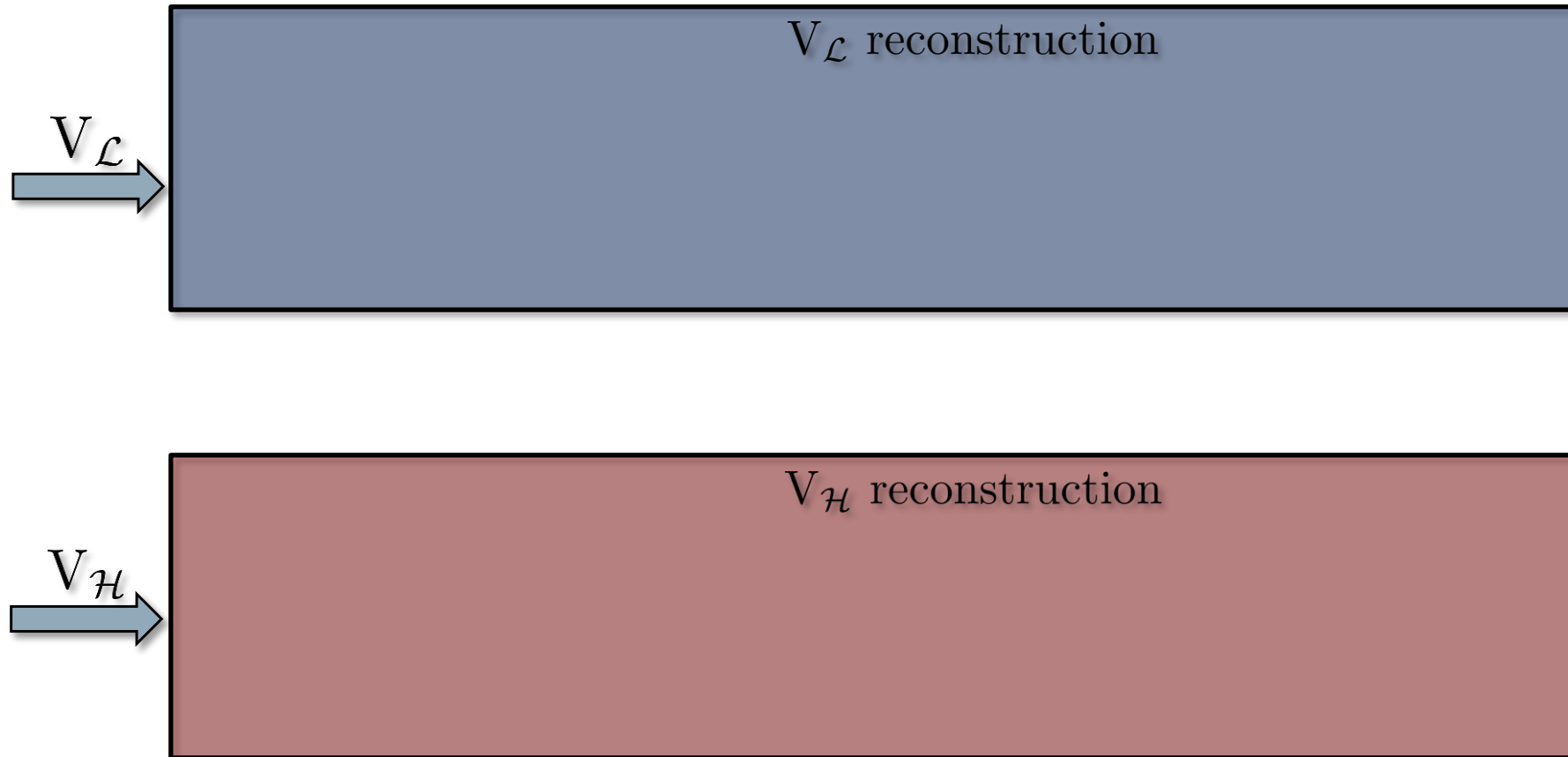
Parallelizing image reconstruction by baseline

Parallelizing the Multi-step Image Reconstruction

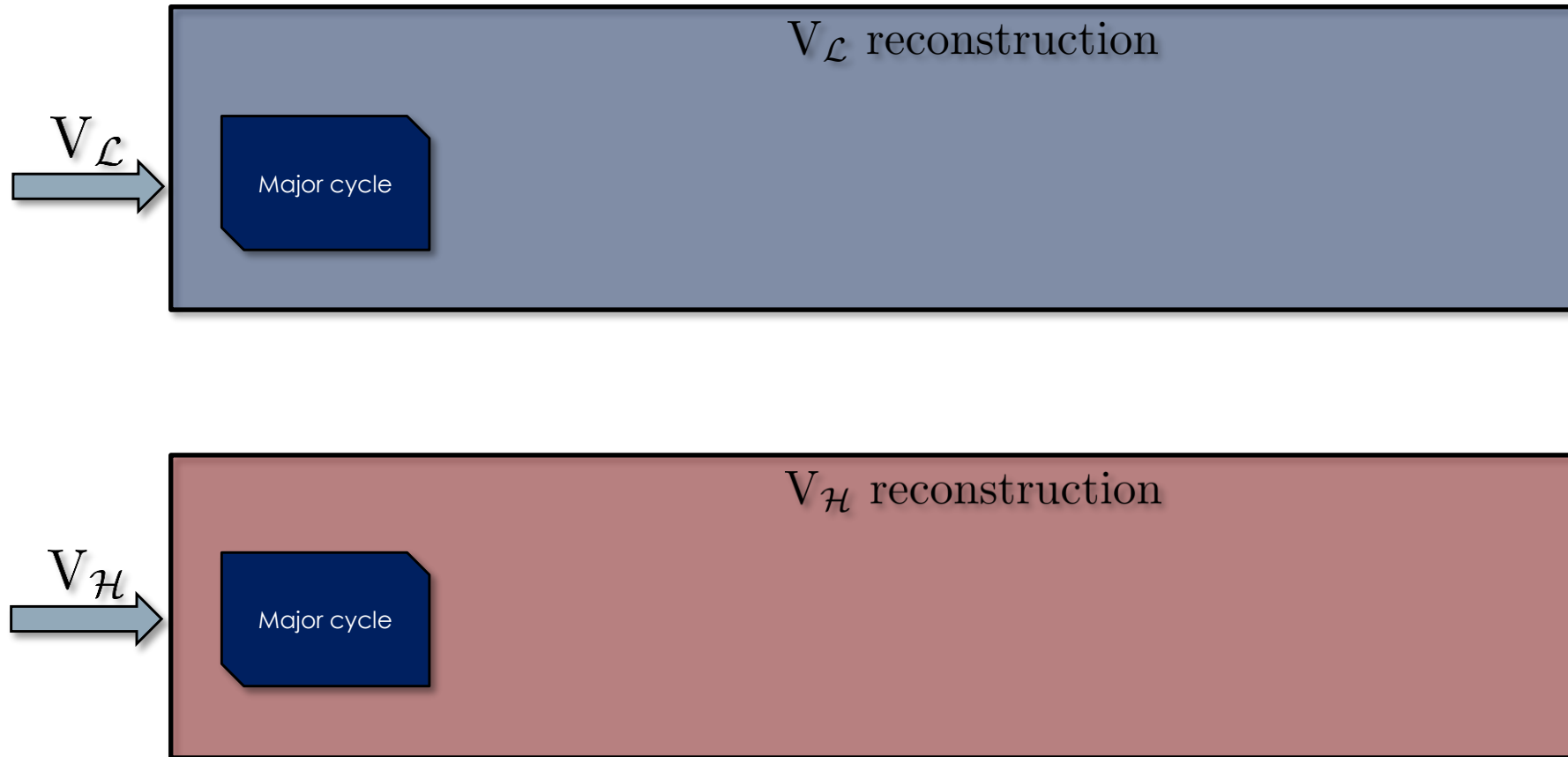


- Described approach allows for better visibility distribution, but is still serial
- Framework can be modified for parallelization

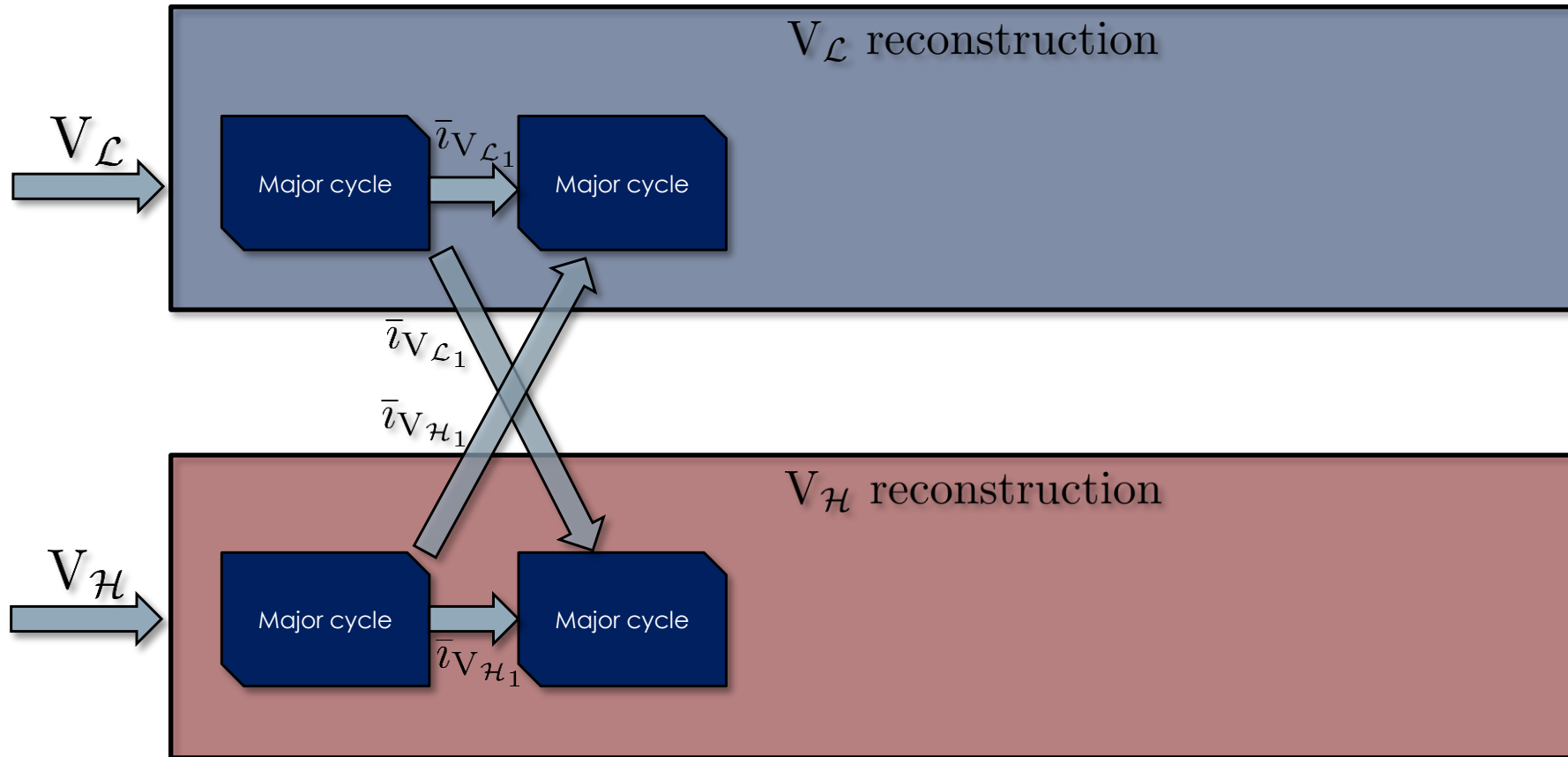
Interleaved parallelization strategy



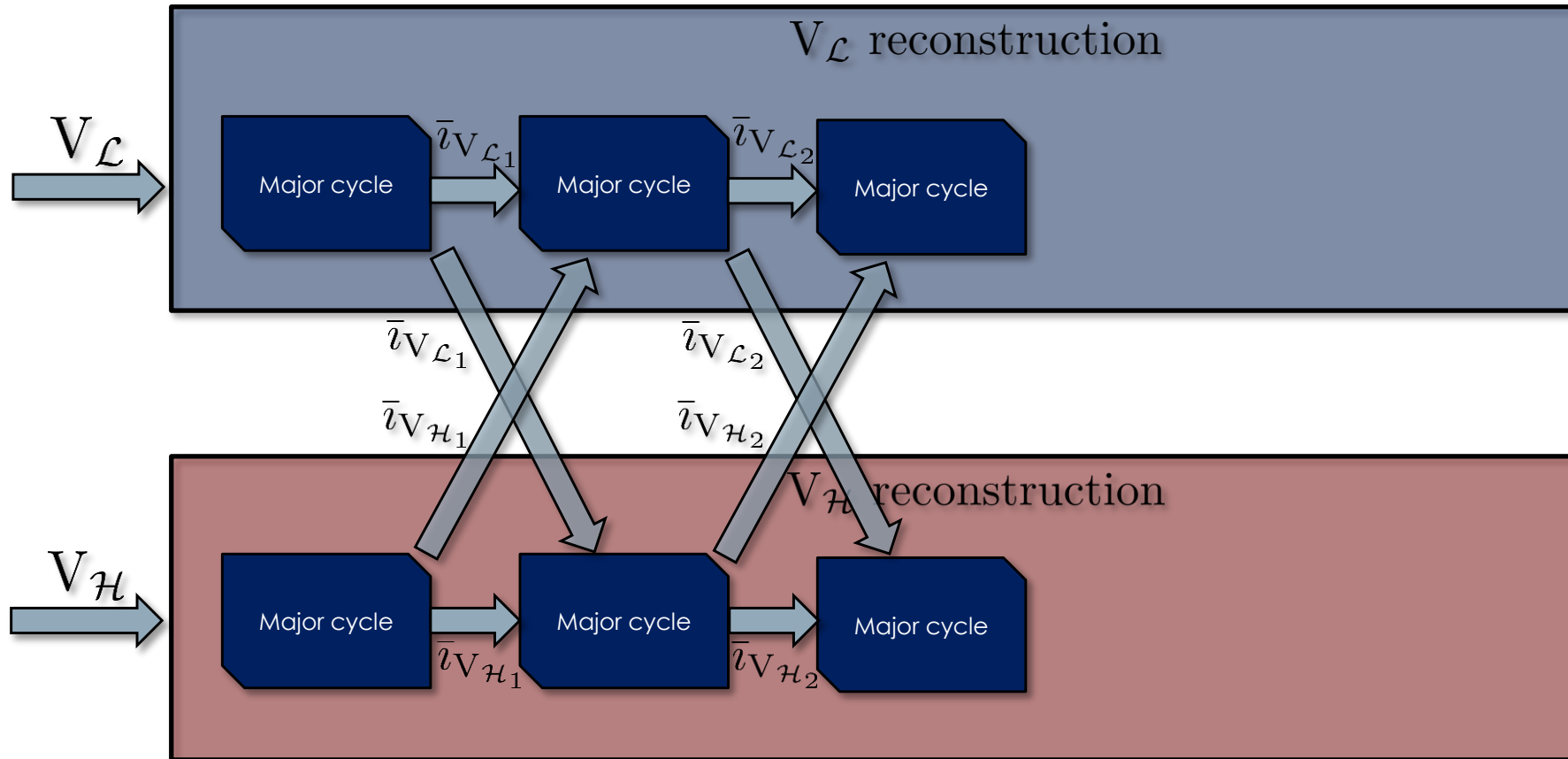
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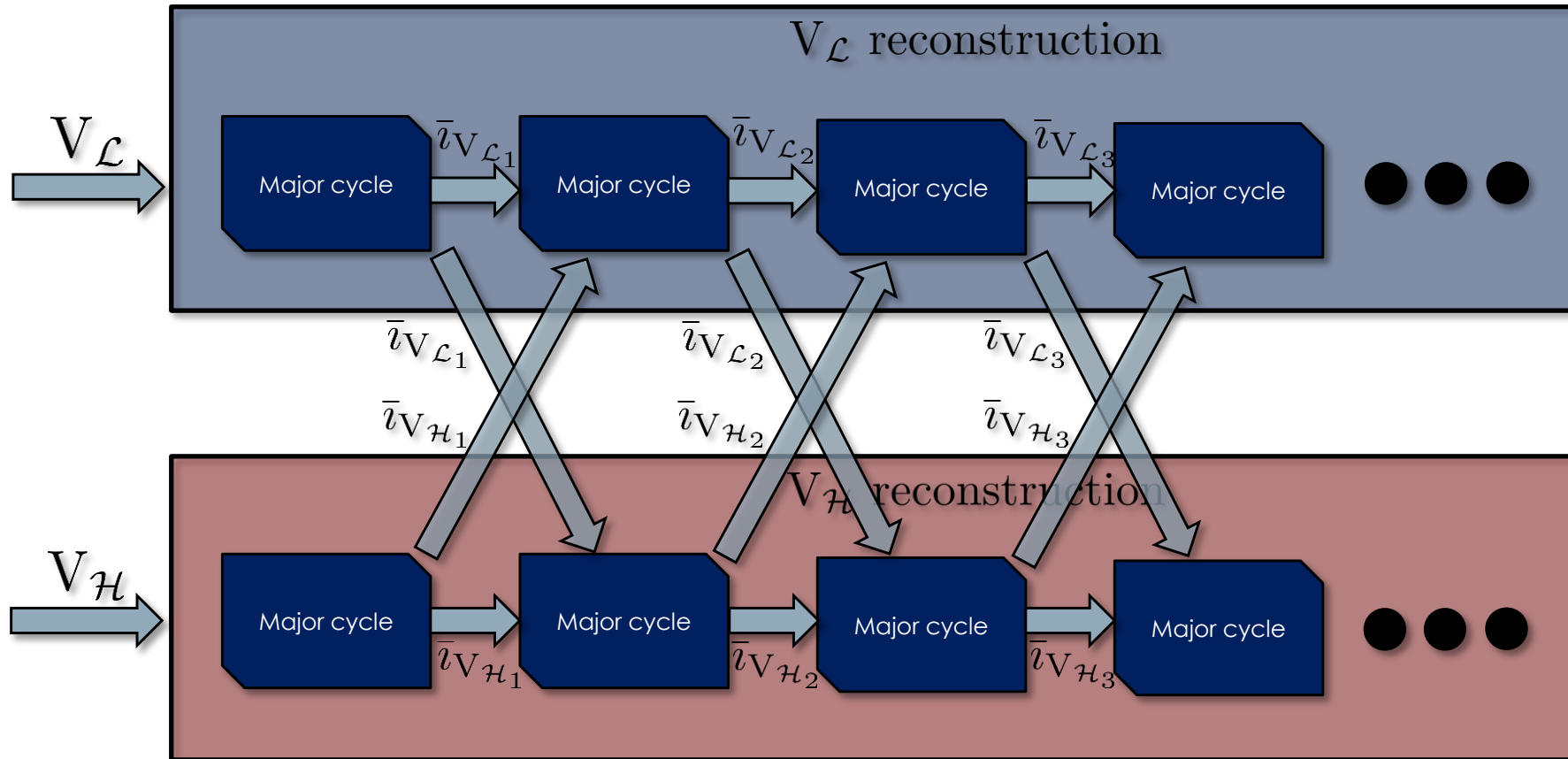
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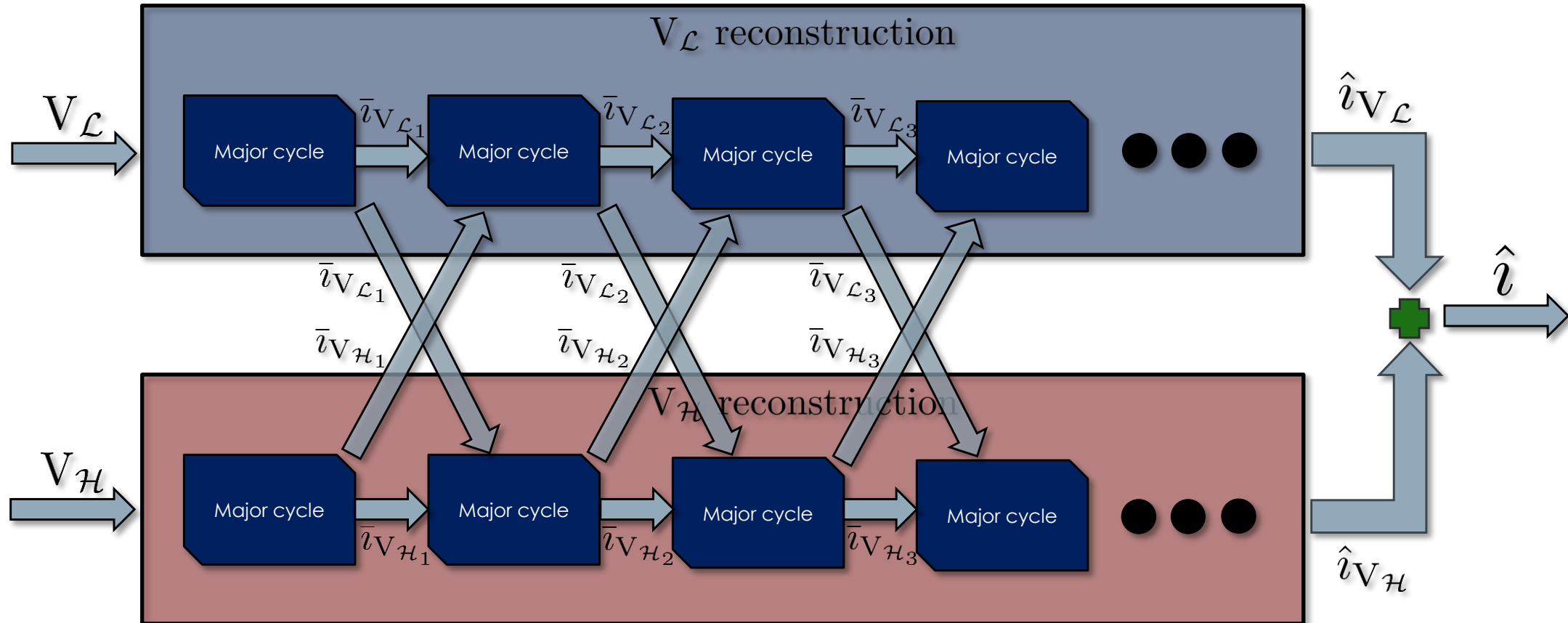
Interleaved parallelization strategy



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Interleaved parallelization strategy

$V_{\mathcal{L}}$ deconvolution

$$\alpha_{V_{\mathcal{L}_n}} = \arg \min_{\alpha} \|G_{\mathcal{L}}(\tilde{v}_{\mathcal{L}_n} - H_{\mathcal{L}}W\alpha)\|_2^2 + \|G_{\mathcal{H}}(h_n - W\alpha)\|_2^2 + \lambda_{V_{\mathcal{L}_n}} \|\alpha\|_1$$

$$\bar{v}_{V_{\mathcal{L}_n}} = W\alpha_{V_{\mathcal{L}_n}}, h_n = \sum_{j=1}^{n-1} \bar{v}_{V_{\mathcal{H}_j}} - \sum_{j=1}^{n-1} \bar{v}_{V_{\mathcal{L}_j}} = \hat{v}_{V_{\mathcal{H}_{n-1}}} - \hat{v}_{V_{\mathcal{L}_{n-1}}}$$

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Visibility data-fidelity term

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Visibility data-fidelity term

Additional data-fidelity term, only from 2nd major cycle onwards

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Interleaved parallelization strategy

$V_{\mathcal{L}}$ deconvolution

$$\alpha_{V_{\mathcal{L}_n}} = \arg \min_{\alpha} \left(\|G_{\mathcal{L}}(\tilde{v}_{\mathcal{L}_n} - H_{\mathcal{L}}W\alpha)\|_2^2 + \|G_{\mathcal{H}}(h_n - W\alpha)\|_2^2 + \lambda_{V_{\mathcal{L}_n}} \|\alpha\|_1 \right)$$

$$\bar{v}_{V_{\mathcal{L}_n}} = W\alpha_{V_{\mathcal{L}_n}}, h_n = \sum_{j=1}^{n-1} \bar{v}_{V_{\mathcal{H}_j}} - \sum_{j=1}^{n-1} \bar{v}_{V_{\mathcal{L}_j}} = \hat{v}_{V_{\mathcal{H}_{n-1}}} - \hat{v}_{V_{\mathcal{L}_{n-1}}}$$

Visibility data-fidelity term

Additional data-fidelity term, only from 2nd major cycle onwards

$V_{\mathcal{H}}$ deconvolution

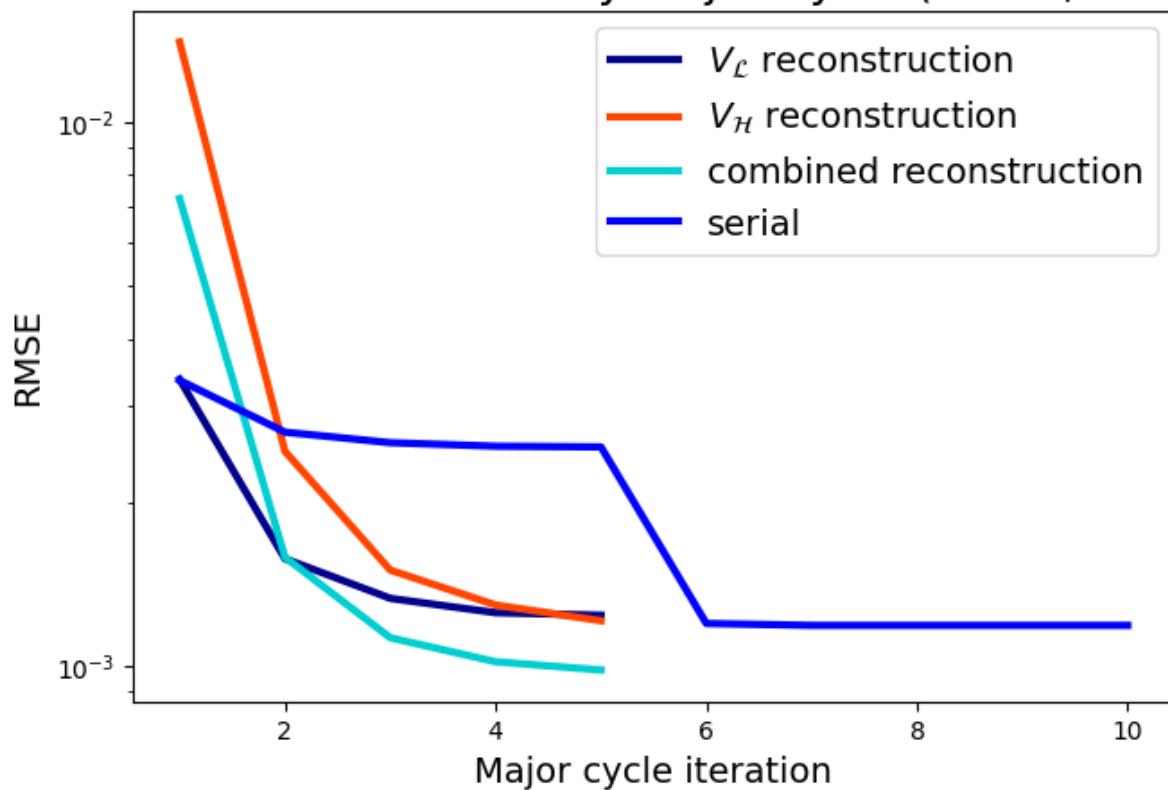
$$\alpha_{V_{\mathcal{H}_n}} = \arg \min_{\alpha} \left(\|G_{\mathcal{H}}(\tilde{v}_{\mathcal{H}_n} - H_{\mathcal{H}}W\alpha)\|_2^2 + \|G_{\mathcal{L}}(l_n - W\alpha)\|_2^2 + \lambda_{V_{\mathcal{H}_n}} \|\alpha\|_1 \right)$$

$$\bar{v}_{V_{\mathcal{H}_n}} = W\alpha_{V_{\mathcal{H}_n}}, l_n = \sum_{j=1}^{n-1} \bar{v}_{V_{\mathcal{L}_j}} - \sum_{j=1}^{n-1} \bar{v}_{V_{\mathcal{H}_j}} = \hat{v}_{V_{\mathcal{L}_{n-1}}} - \hat{v}_{V_{\mathcal{H}_{n-1}}}$$

- Two full-resolution images instead of one low and one full
- Roughly symmetric reconstruction costs
- Combined image should be of higher quality
- Quadratic communication with number of partitions
- At least 4 major-cycles needed for an image

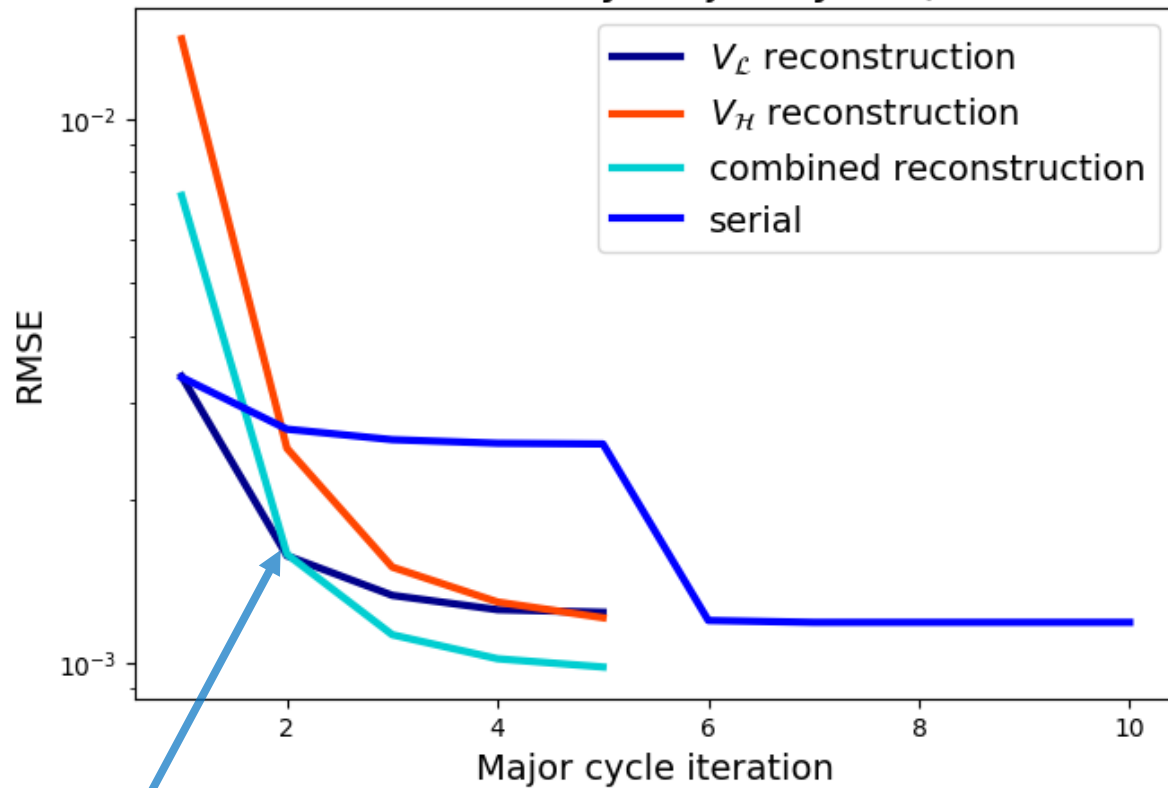
Results: Reconstruction quality of interleaved method

Serial vs Interleaved by Major Cycle ($\ell = 20, \delta = 5$)



Results: Reconstruction quality of interleaved method

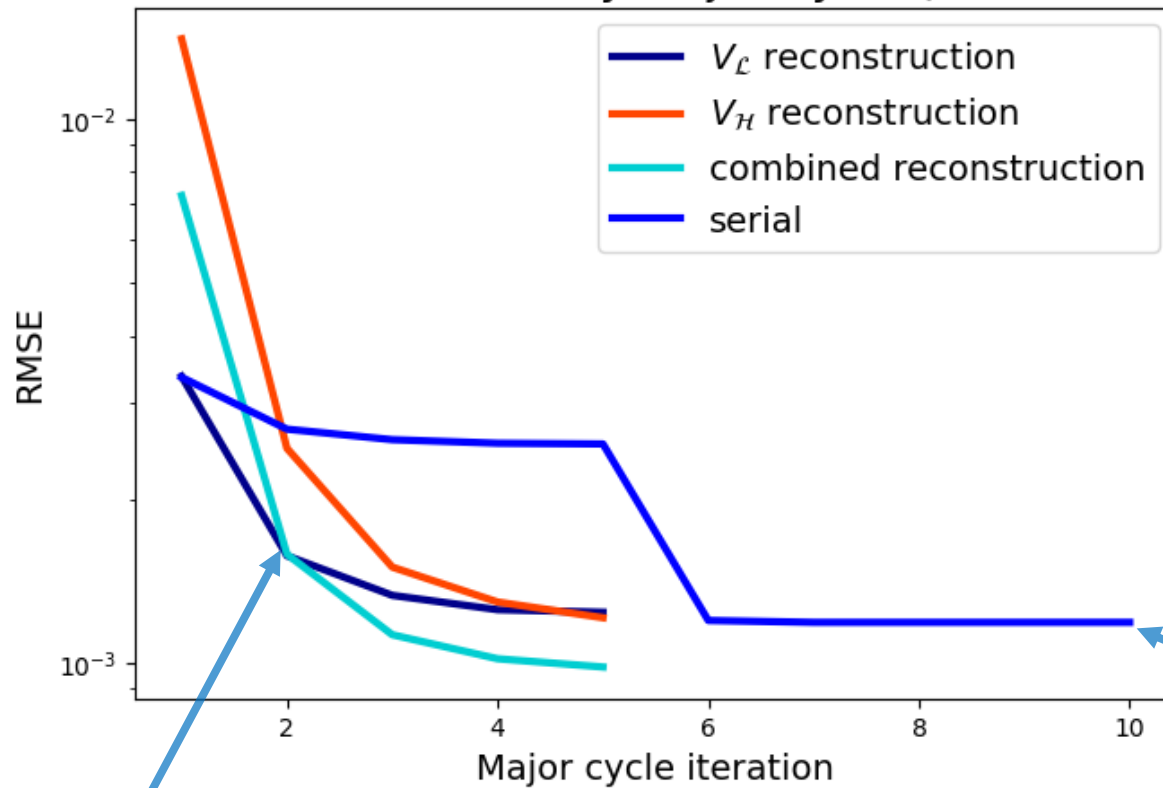
Serial vs Interleaved by Major Cycle ($\ell = 20, \delta = 5$)



Combined reconstruction equal to or outperforms others after 2nd major cycle

Results: Reconstruction quality of interleaved method

Serial vs Interleaved by Major Cycle ($\ell = 20, \delta = 5$)



Combined reconstruction equal to or outperforms others after 2nd major cycle

Serial reconstruction worse than parallel after roughly same amount of computation

Conclusions and Future Work

➤ To conclude:

- Method to partition visibilities by baseline length when reconstructing images
 - ❖ Shown to have similar cost and quality to processing all baselines together
- Parallelization strategy for given method
 - ❖ Better reconstruction quality than serial for same cost

Future work:

- More partitions
- Implementation and benchmarking on large cluster
- Different metrics for reconstruction quality
- Investigate the framework with other deconvolution frameworks

Thank you!
Questions?

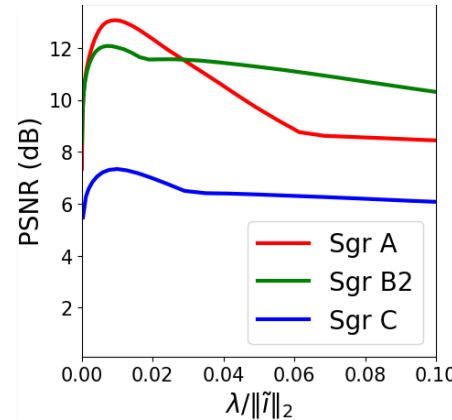
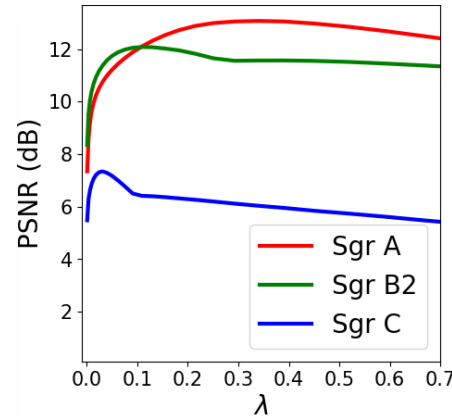




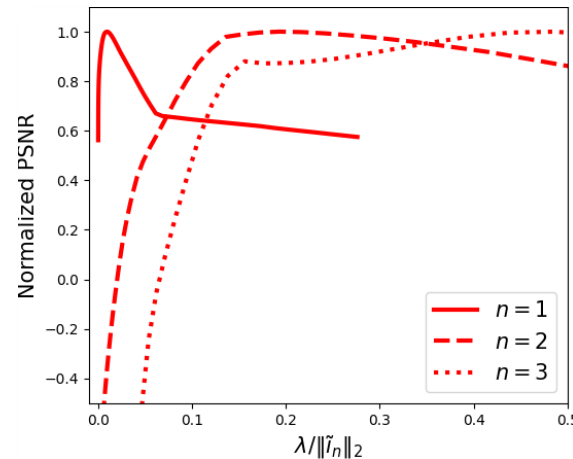
Appendices

Selection of λ

Across different datasets for first major-cycle:



Across first three major-cycles for Sgr A dataset:



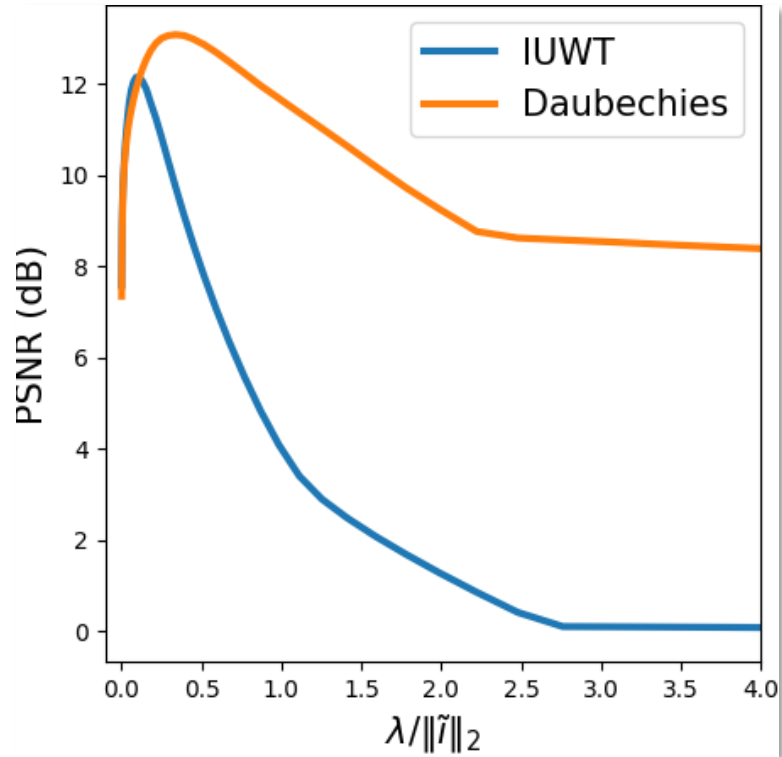
Results suggest that lambda should be normalized by the norm of the image, and be increased as the major-cycles progress to maximize RMSE/PSNR. We use:

$$\lambda_{\mathcal{L}_n} = 0.05 \|\tilde{I}_{\mathcal{L}_n}\|_2 \times 2^n$$

$$\lambda_{\mathcal{H}_n} = 0.05 \|\tilde{I}_{\mathcal{H}_n}\|_2 \times 2^n$$

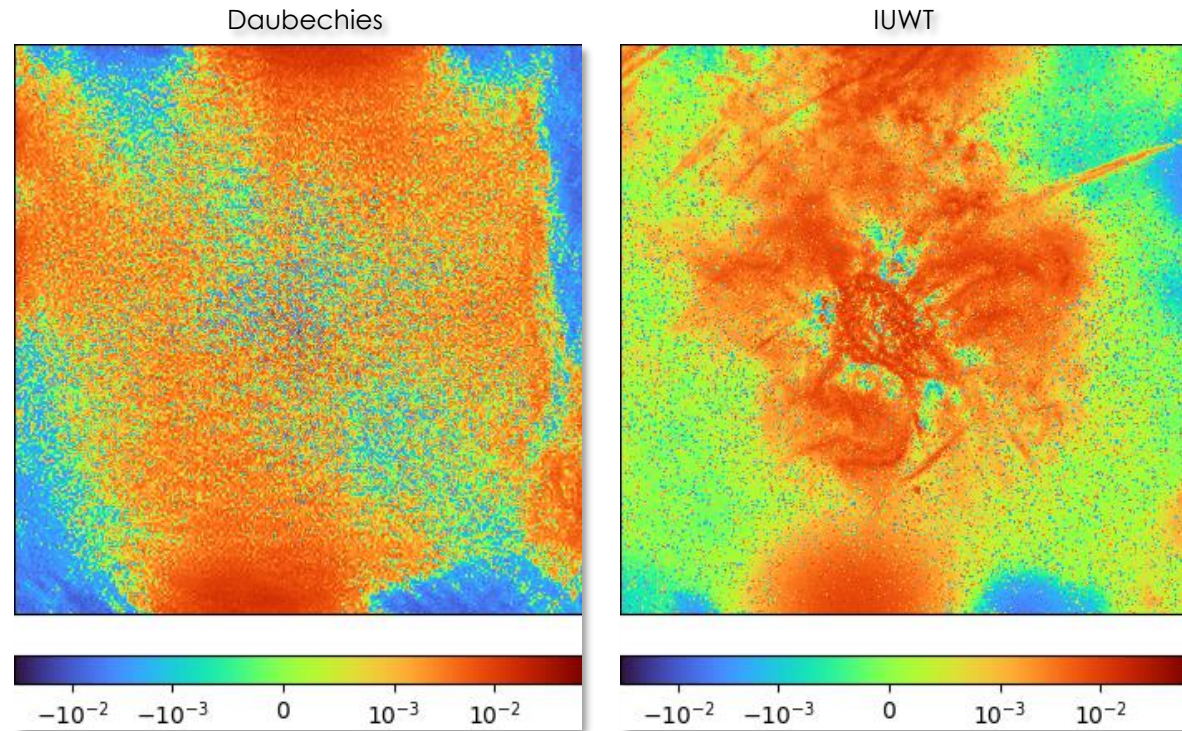
$$\lambda_n = 0.01 \|\tilde{I}_n\|_2 \times 2^n$$

IUWT vs Daubechies



IUWT seems worse at reconstructing large-scale extended emissions, possibly due to its isotropic nature.

First major-cycle residuals for Sgr A



Filters

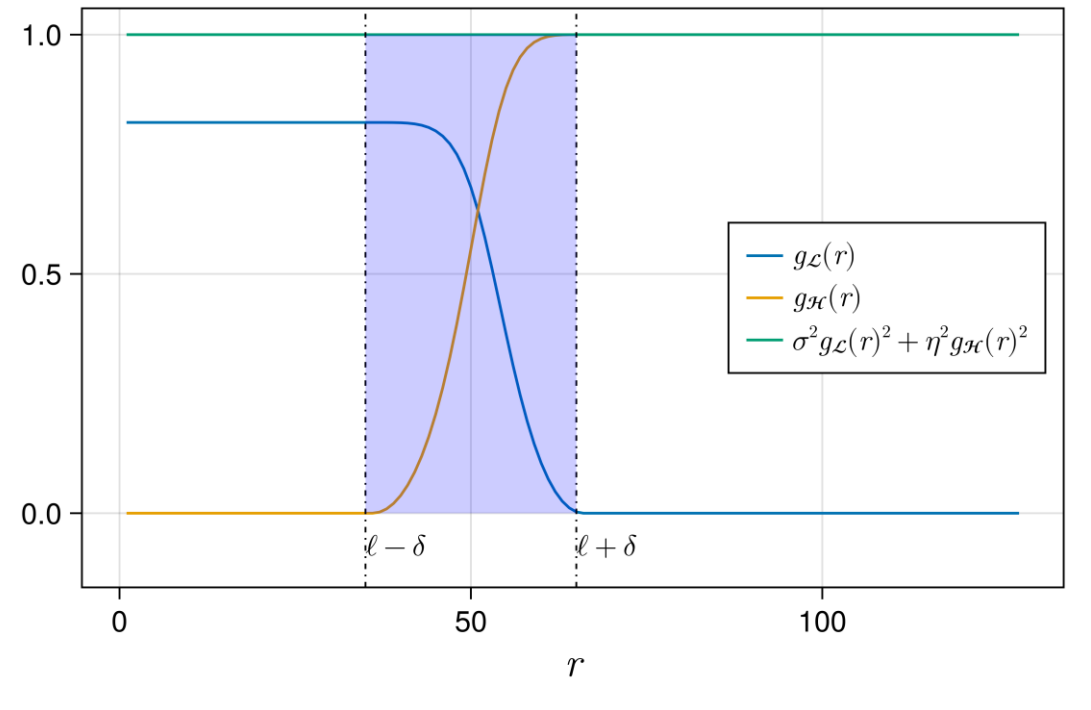
$$r > \ell + \delta : |g_{\mathcal{H}}(r)|^2 = 1/\sigma^2, g_{\mathcal{L}}(r) = 0$$

$$r < \ell - \delta : g_{\mathcal{H}}(r) = 0, |g_{\mathcal{L}}(r)|^2 = 1/\eta^2$$

$$\ell - \delta < r < \ell + \delta : \sigma^2 |g_{\mathcal{H}}(r)|^2 + \eta^2 |g_{\mathcal{L}}(r)|^2 = 1$$

$$g_{\mathcal{L}}(r) = \alpha(r) \left(1 - \sin \left(\frac{\pi}{2\delta} (r - \ell) \right) \right)$$

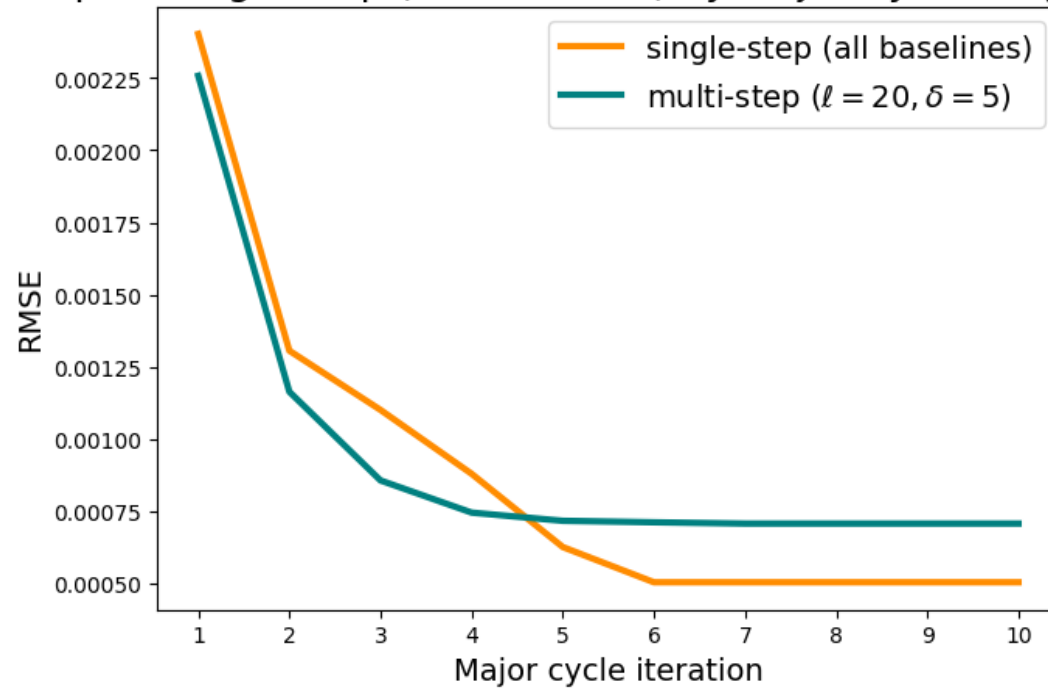
$$g_{\mathcal{H}}(r) = \alpha(r) \left(1 + \sin \left(\frac{\pi}{2\delta} (r - \ell) \right) \right)$$



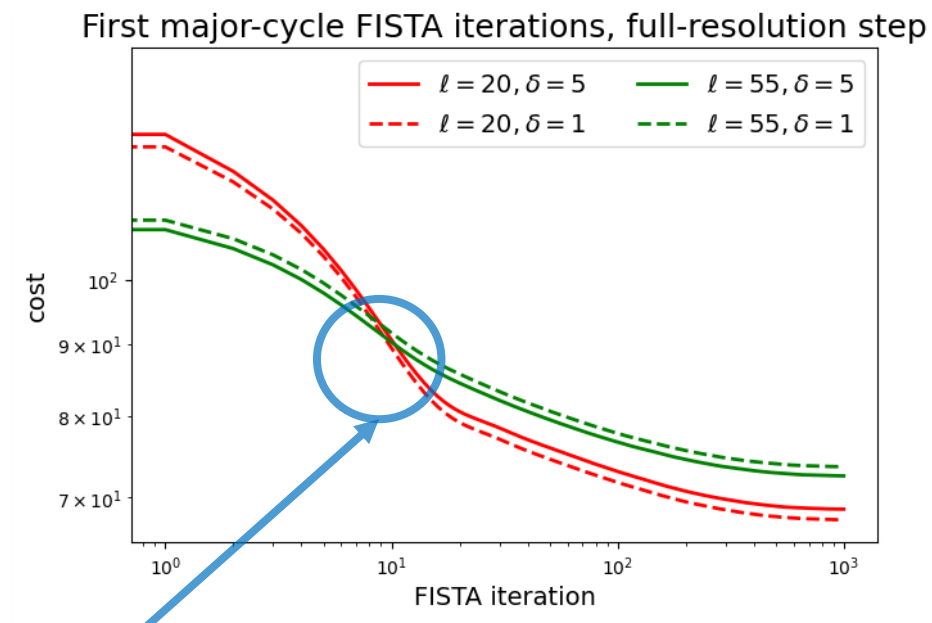
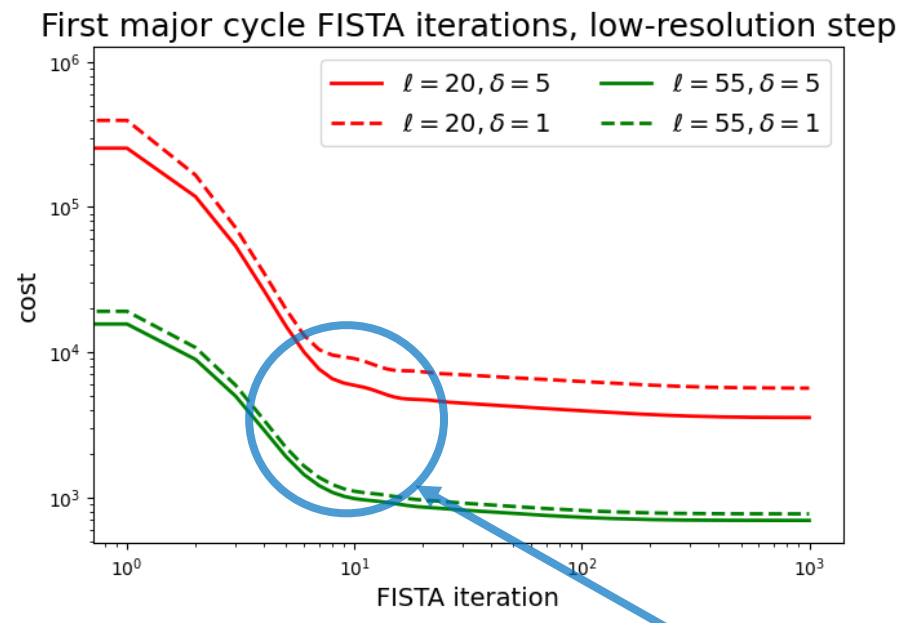
Reconstruction of large-scale features

Full-resolution step does not seem to change the large-scales

Multi-step vs Single-step (all baselines) by Major Cycle, large-scale error

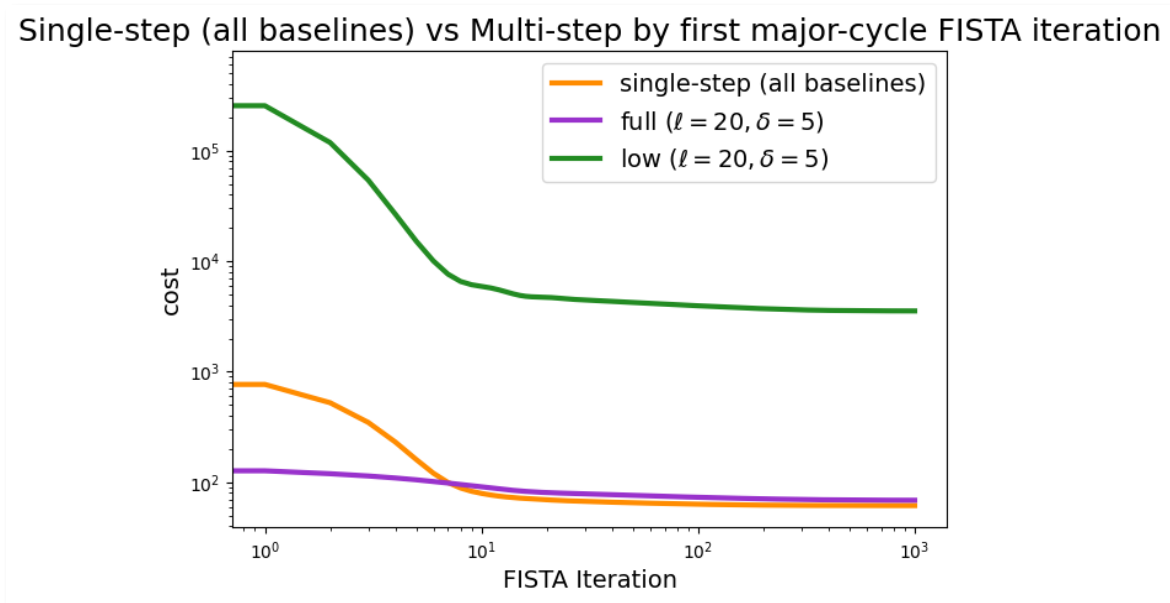
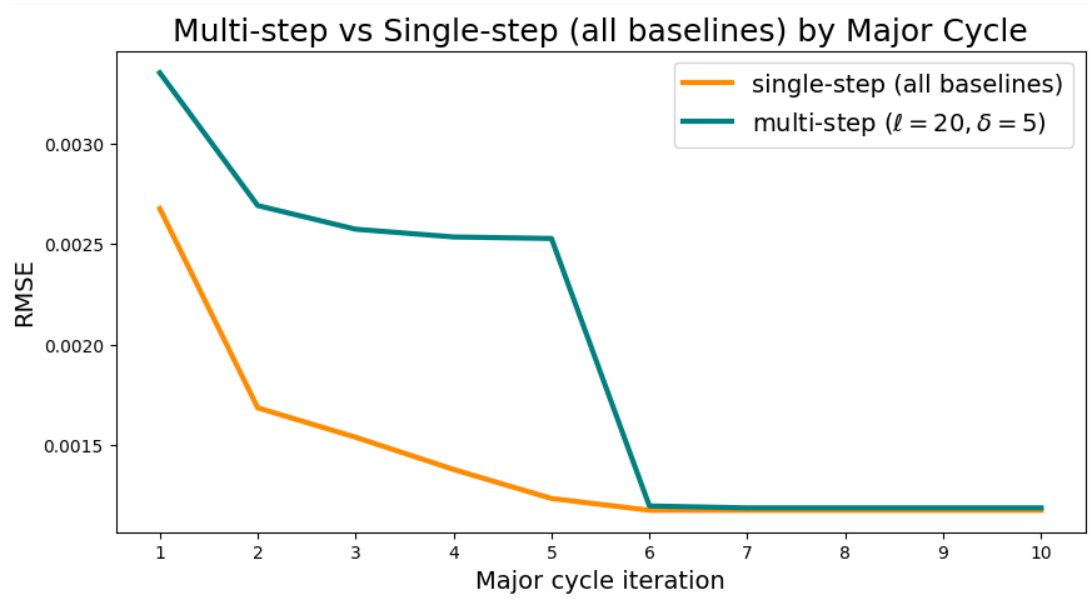


Results: Partition configuration affect on reconstruction accuracy and speed, FISTA

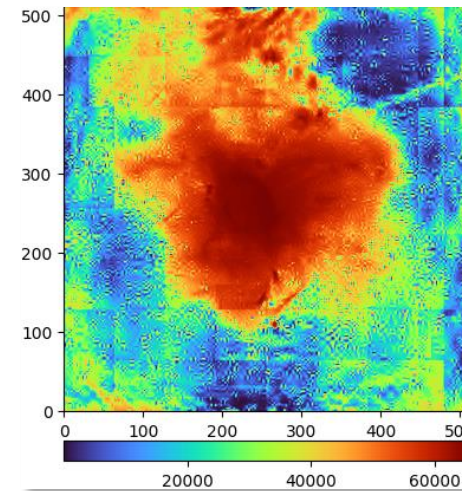
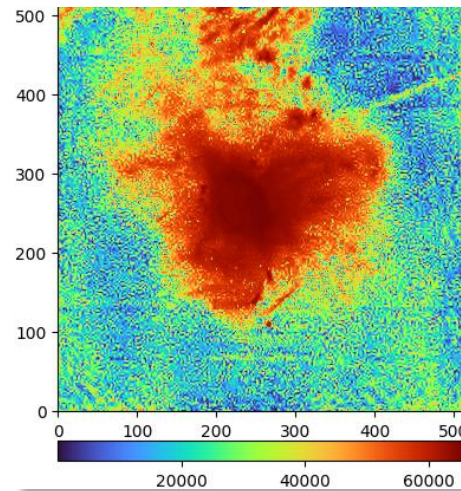
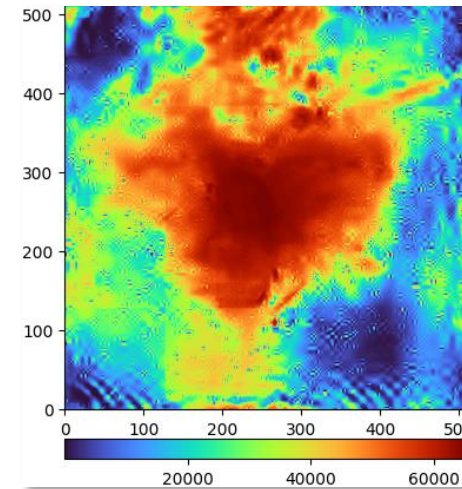
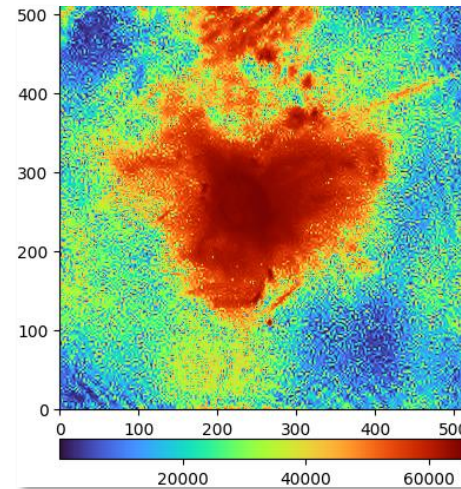
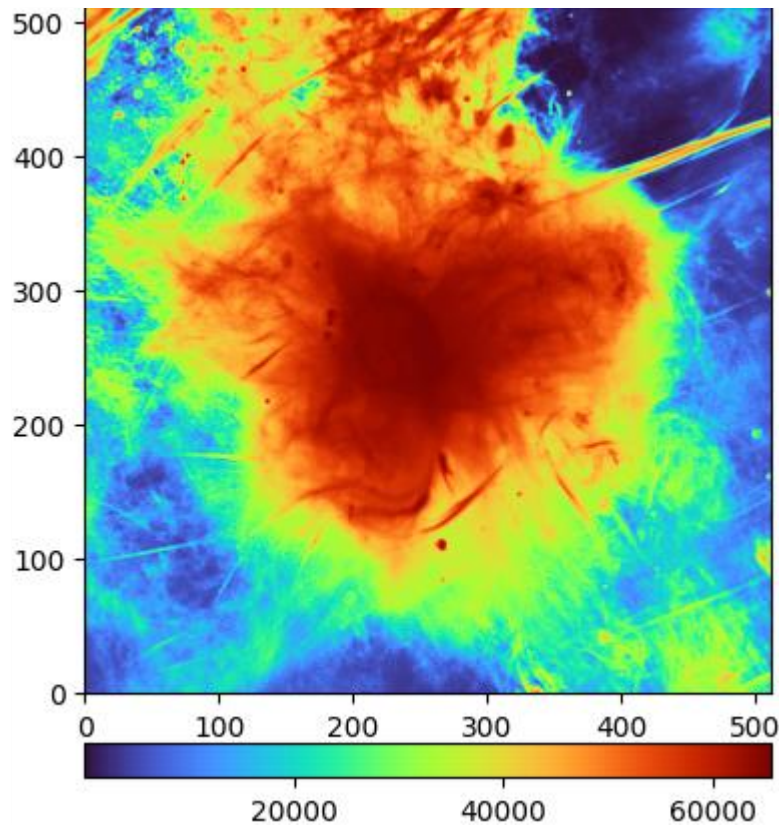


Convergence roughly the same, gradient of cost function remains very similar between partition configurations and starts tapering off around 10-25 iterations.

Results: Comparison to all-baselines reconstruction

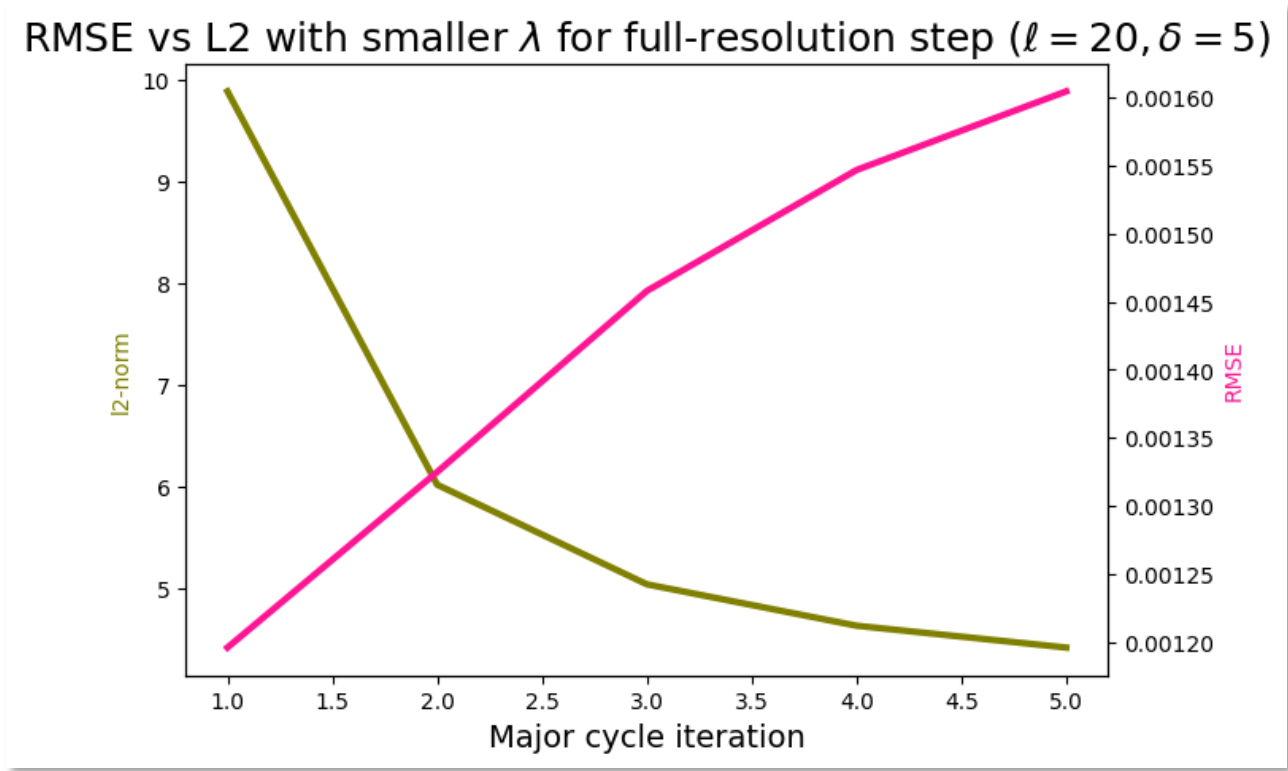


Example image reconstructions (Histogram equalized)

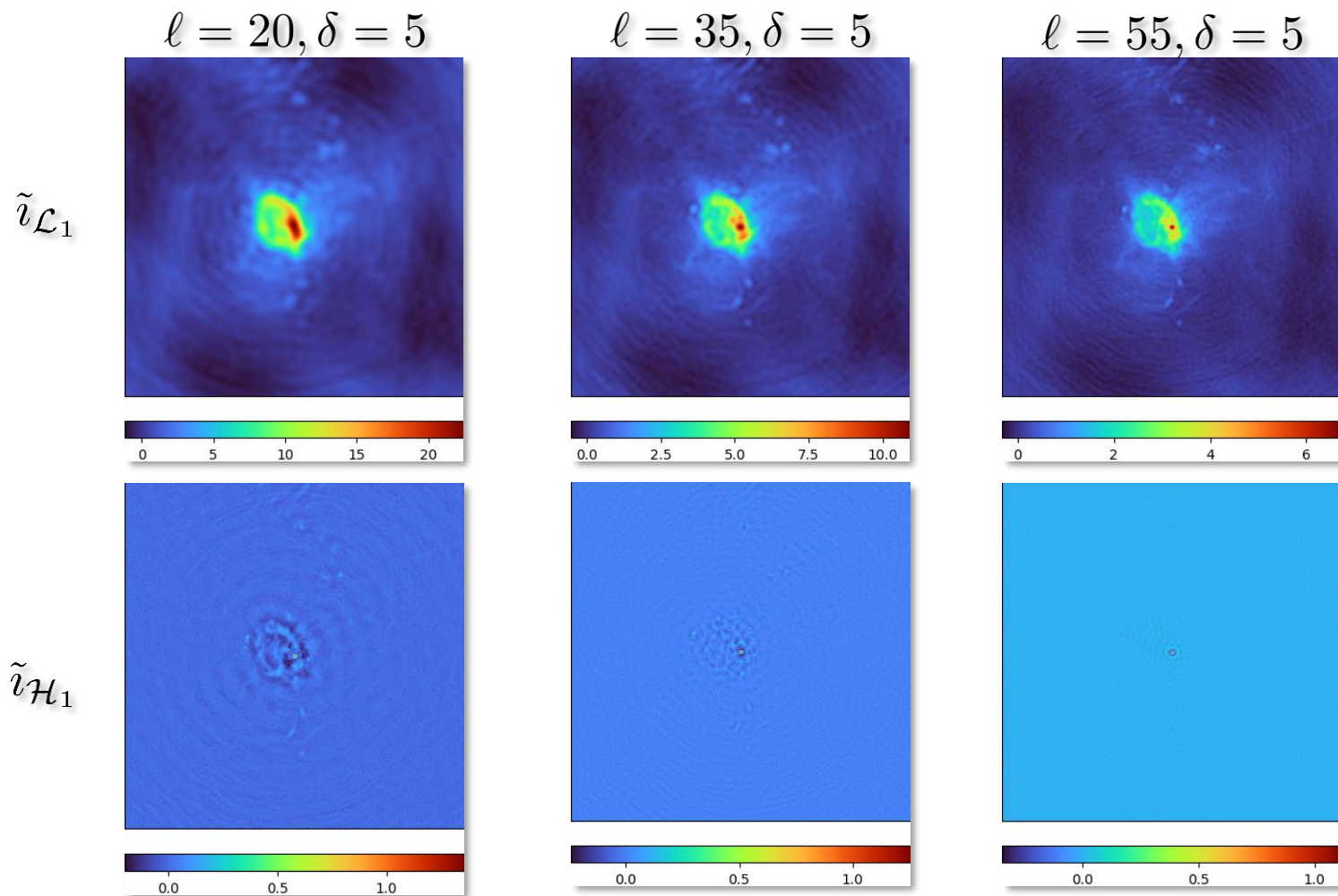


Using a less aggressive lambda for the full-resolution step

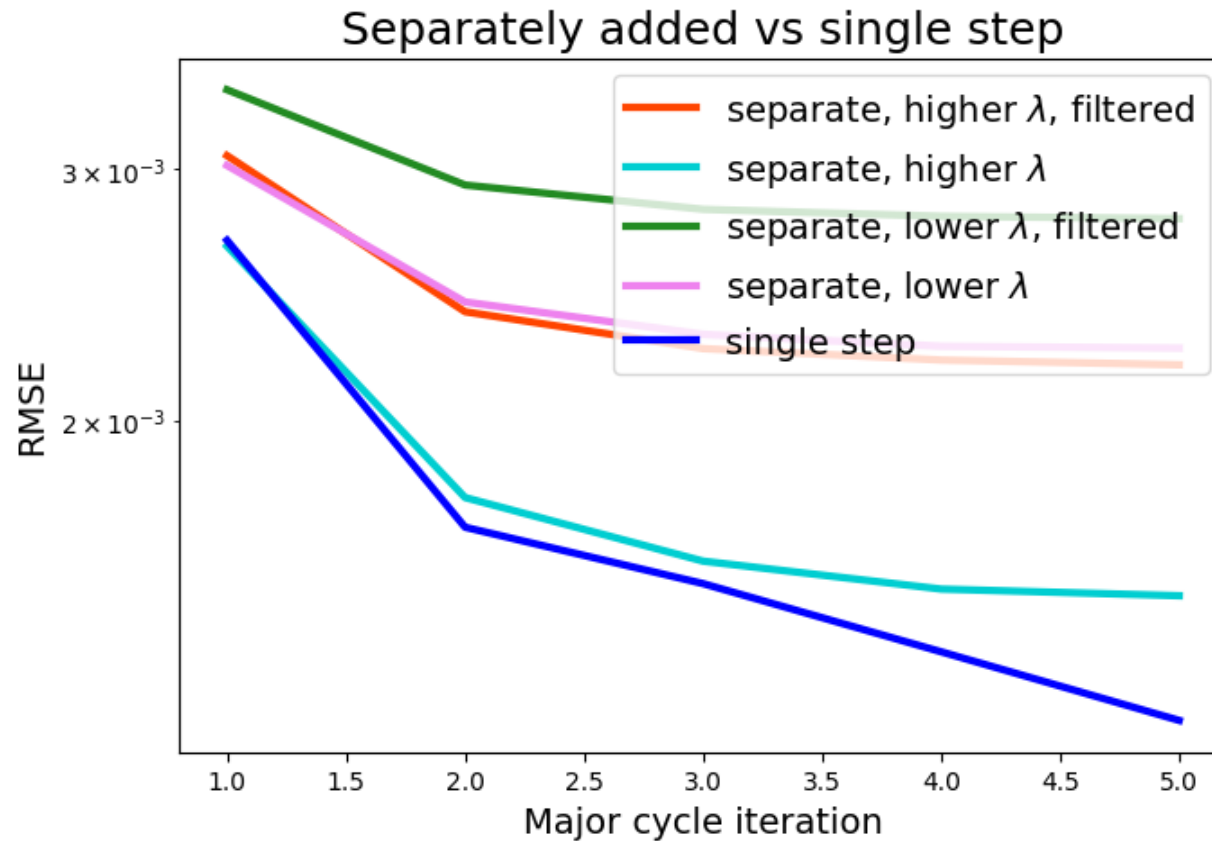
A less aggressive lambda can result in gradual reduction of residual. This can overfit to noise, ergo the worsening RMSE.



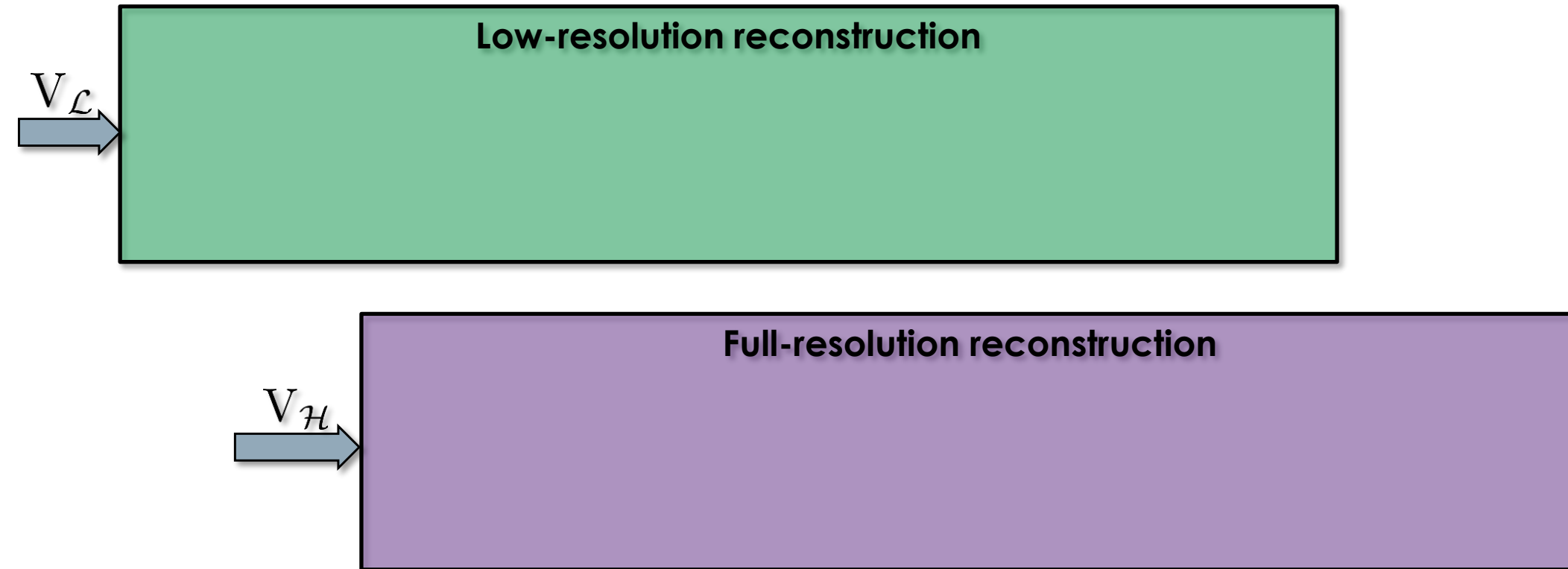
Visualization of partition configurations



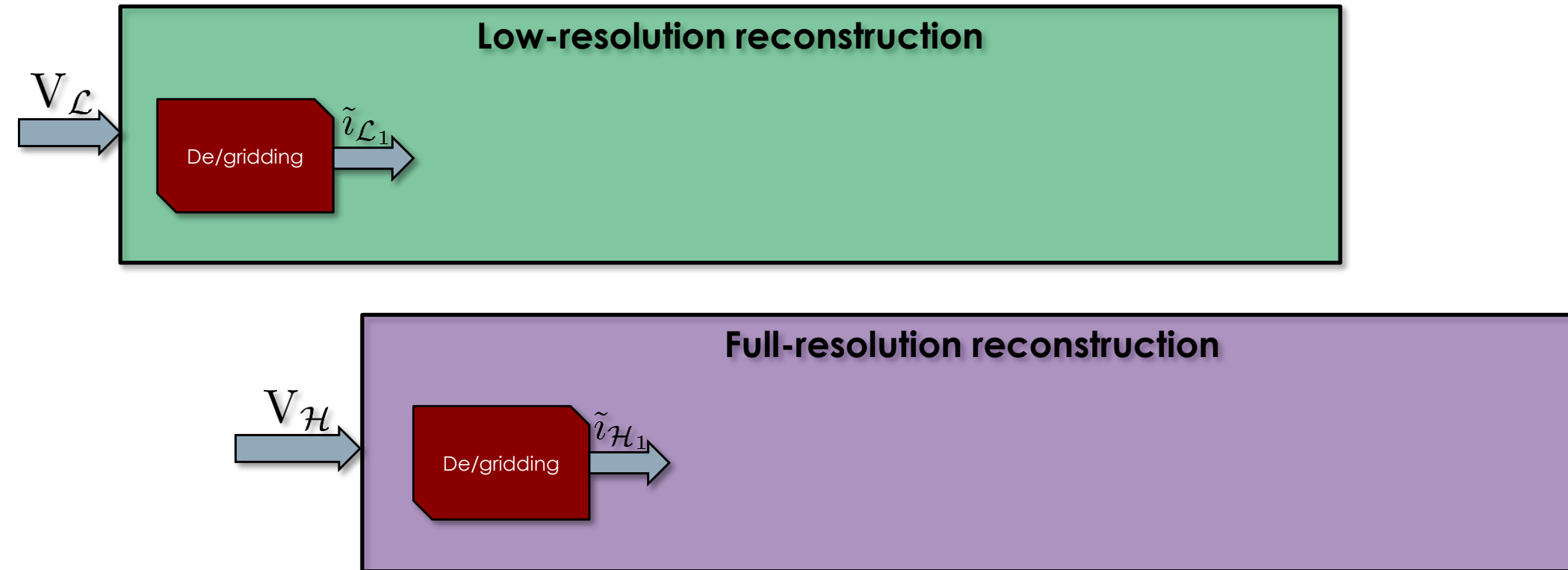
Just adding separately deconvolved images



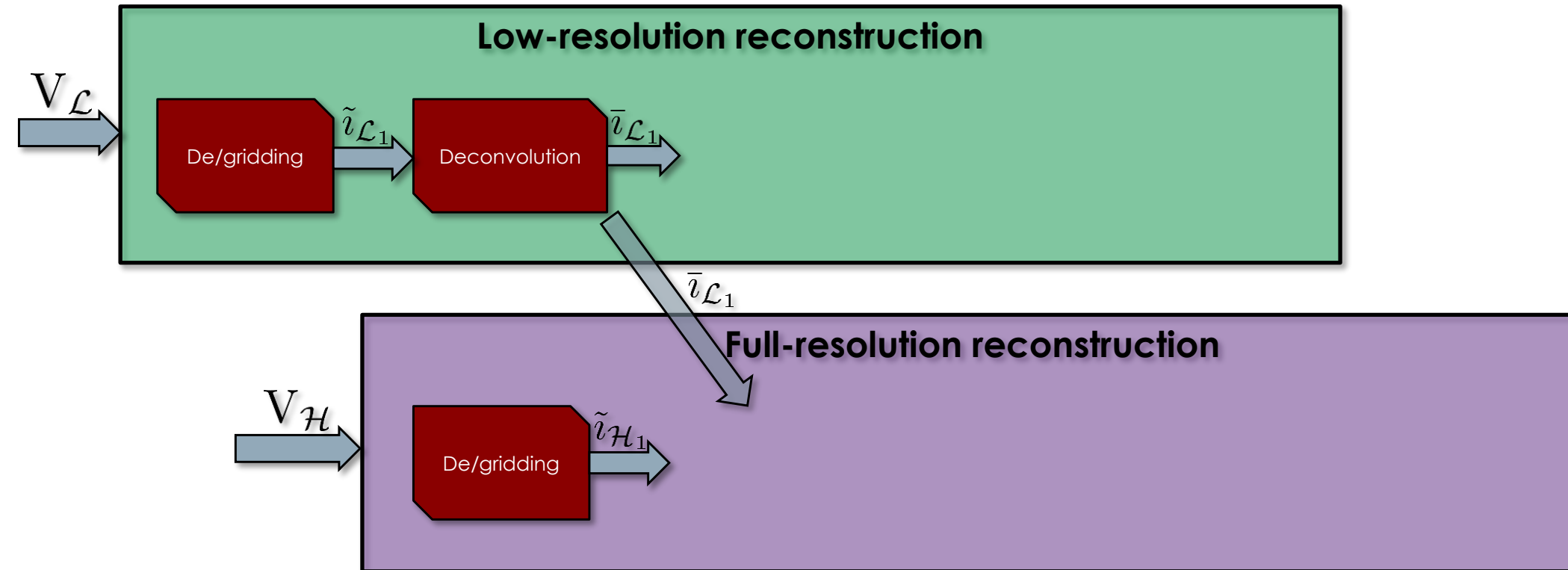
Pipelined parallelization strategy



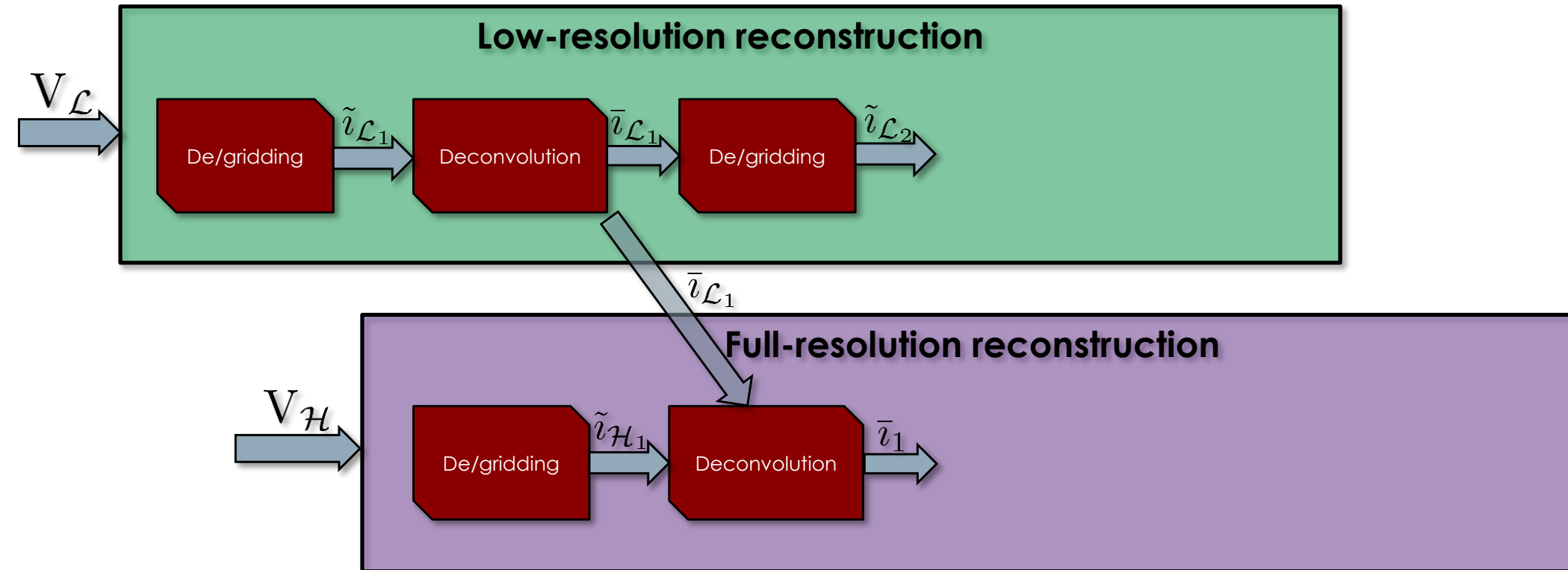
Pipelined parallelization strategy



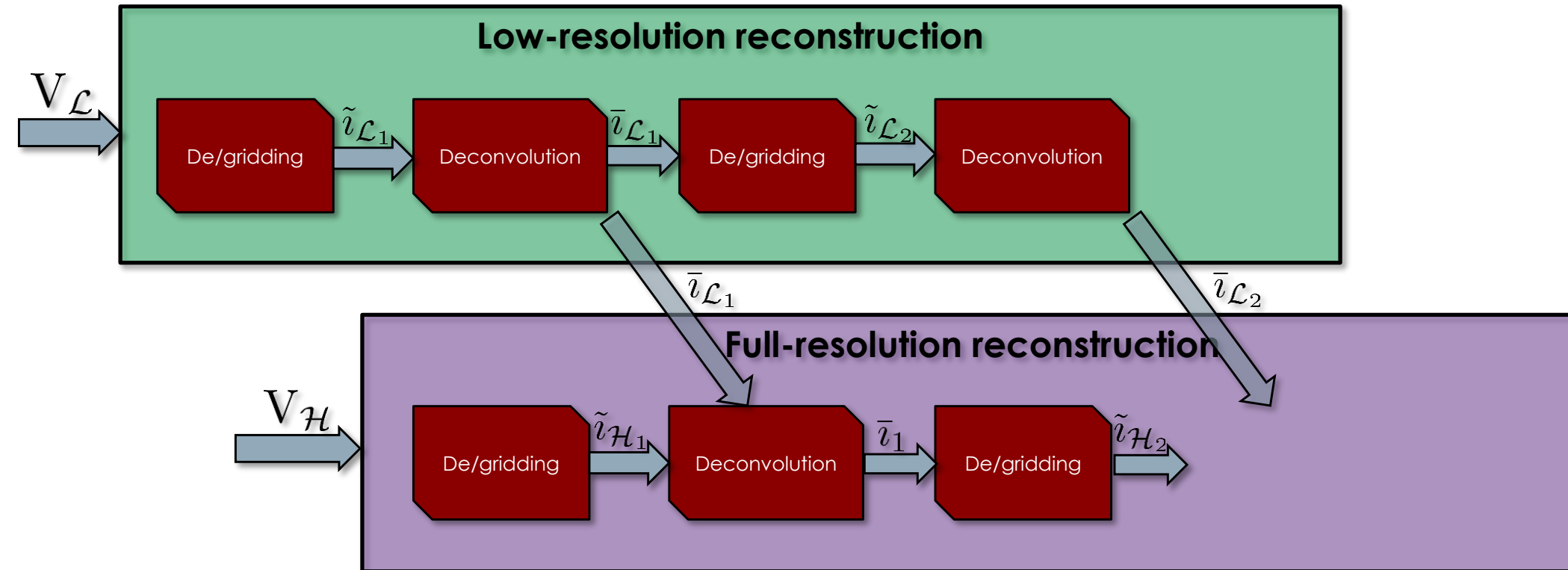
Pipelined parallelization strategy



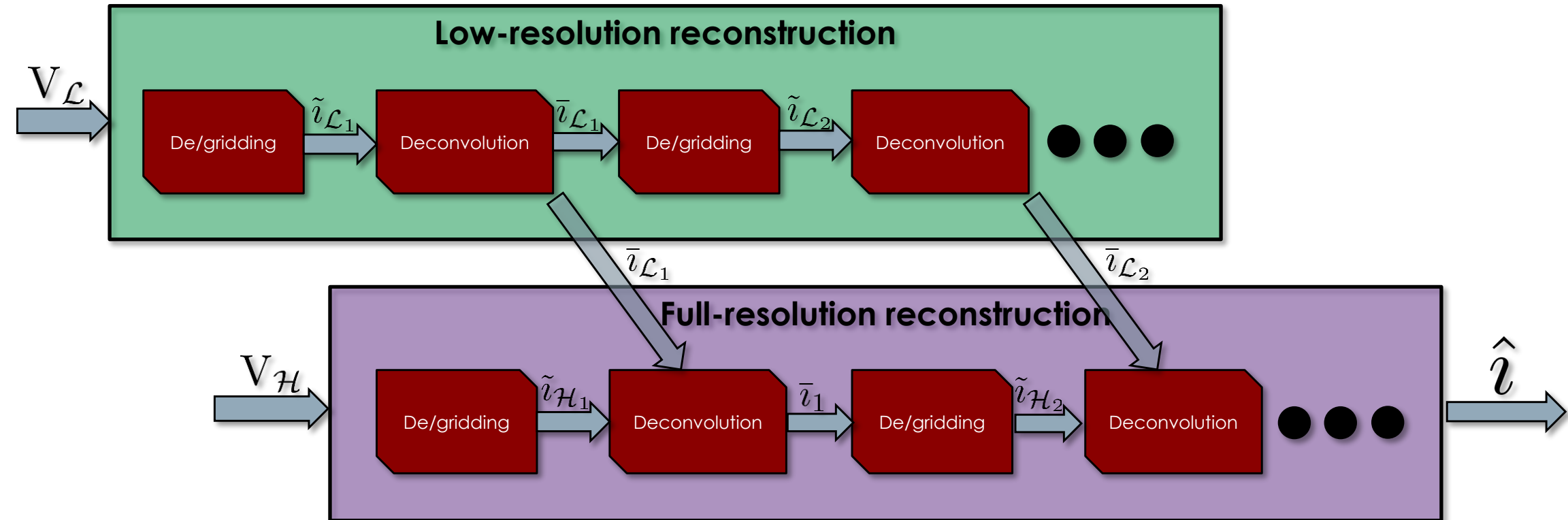
Pipelined parallelization strategy



Pipelined parallelization strategy



Pipelined parallelization strategy



Pipelined parallelization strategy

- Low-resolution step remains the same
- For the full-resolution step:

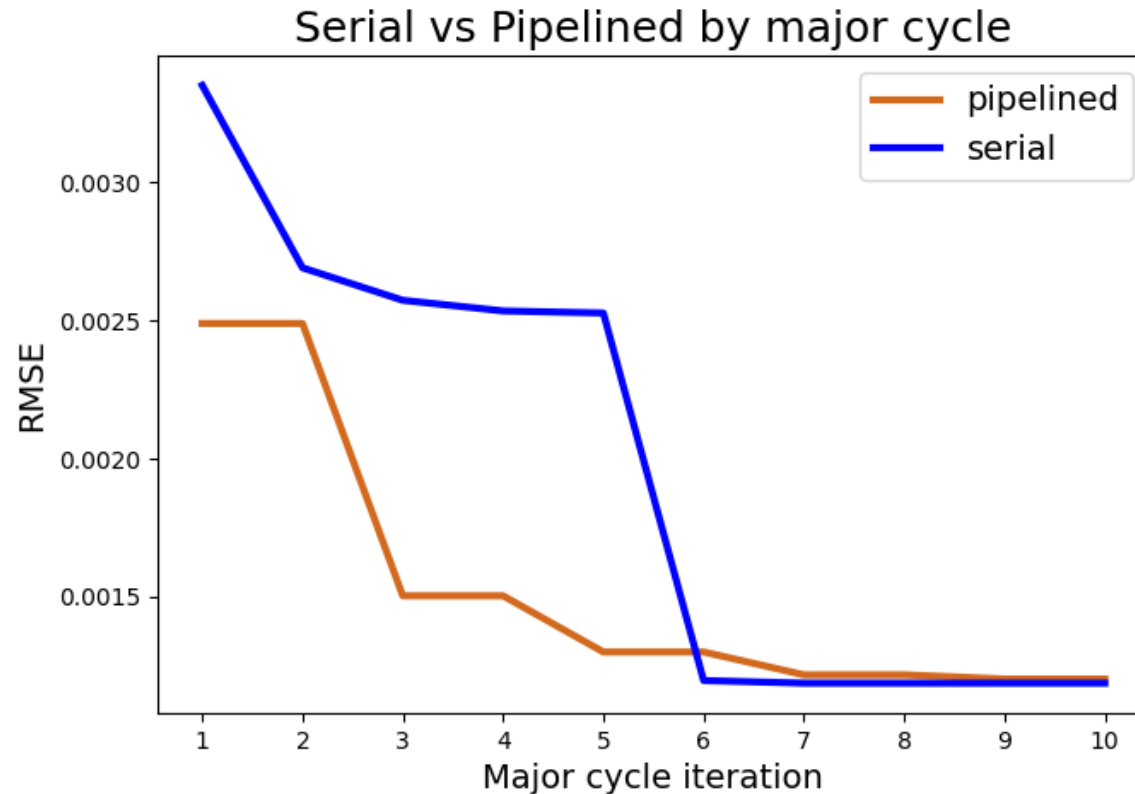
$$l_n = \sum_{j=1}^N \bar{v}_{\mathcal{L}_j} - \sum_{j=1}^{n-1} \bar{v}_j = \hat{v}_{\mathcal{L}} - \hat{v}_{n-1}$$

changes to:

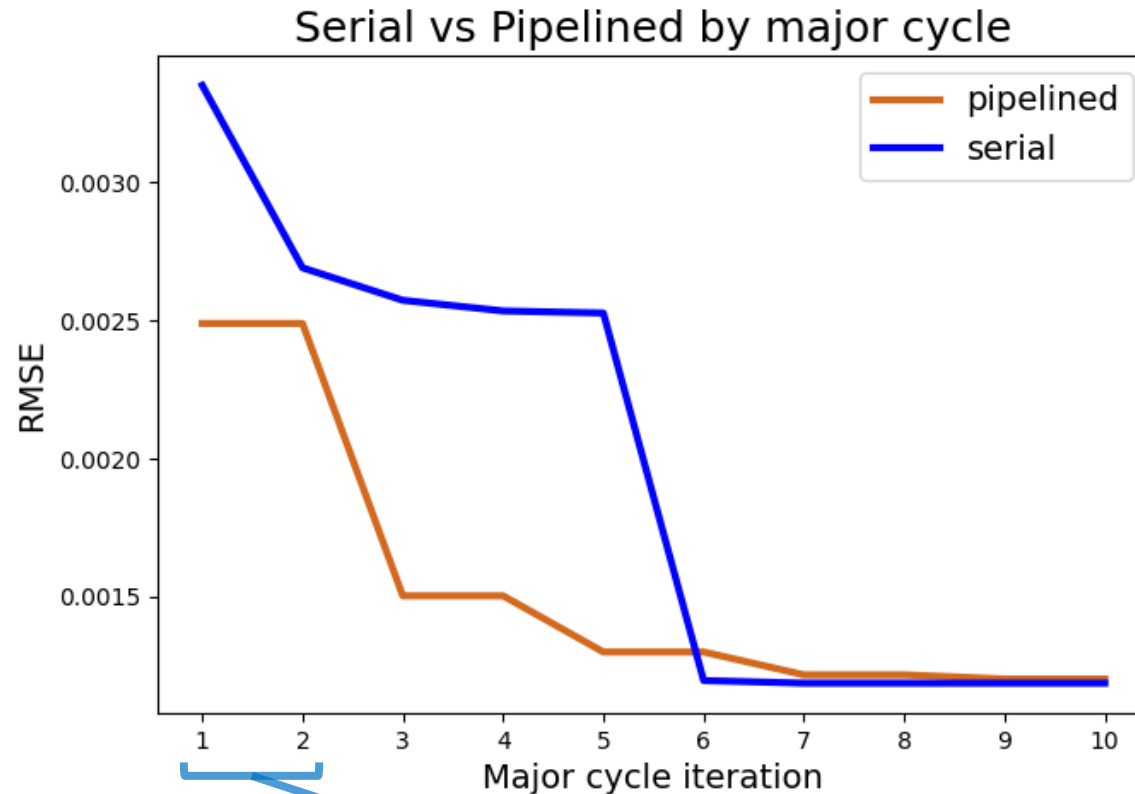
$$l_n = \sum_{j=1}^n \bar{v}_{\mathcal{L}_j} - \sum_{j=1}^{n-1} \bar{v}_j = \hat{v}_{\mathcal{L}_{n-1}} - \hat{v}_{n-1} + \bar{v}_{\mathcal{L}_n}$$

- Can result in waiting if computation costs of each step not similar
 - ❖ Asynchronous strategy can alleviate this somewhat
- Quality upper bound of serial (but most likely slightly worse)
- One image transmitted per major cycle

Results: Reconstruction quality of pipelined parallel method

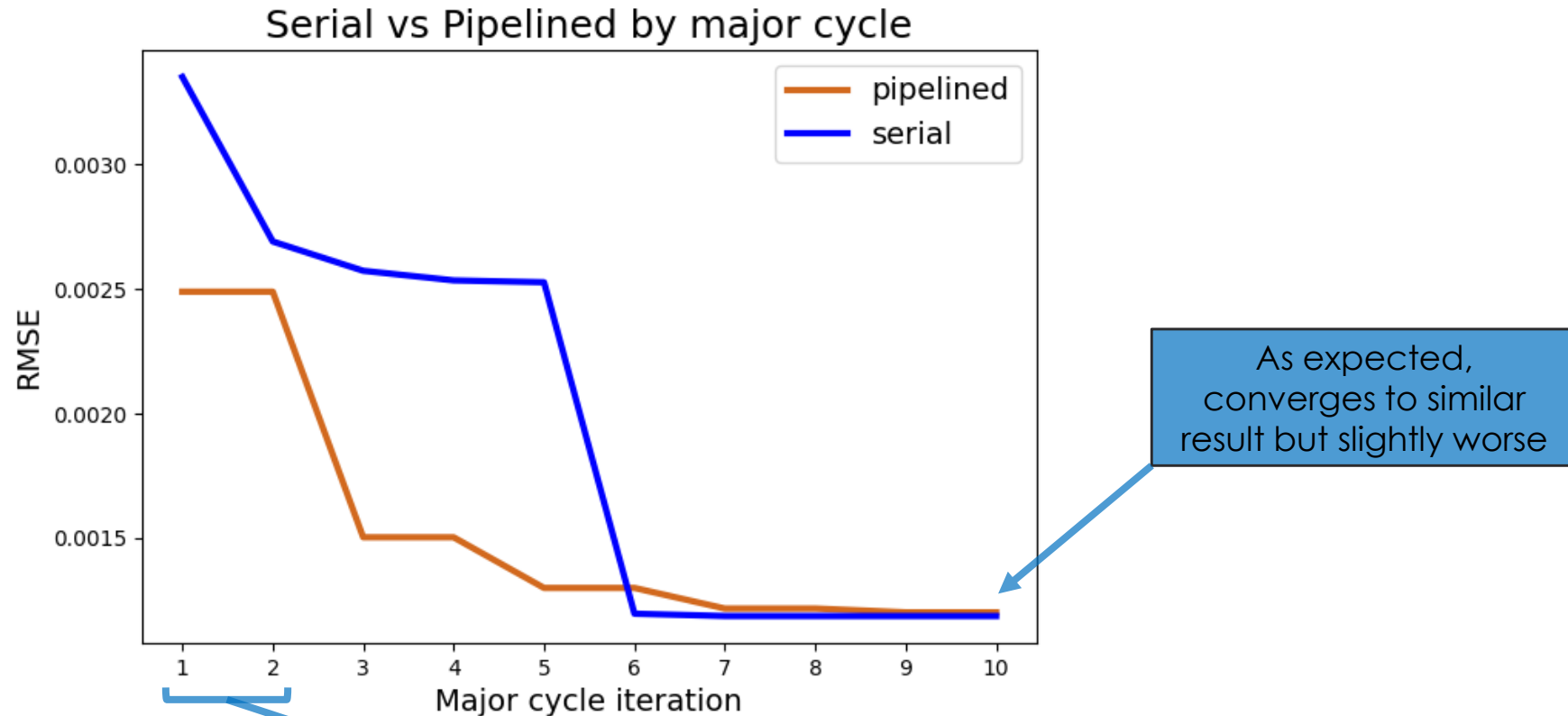


Results: Reconstruction quality of pipelined parallel method



For pipelined, only showing full-resolution reconstruction after 1 low and 1 full-resolution major-cycle

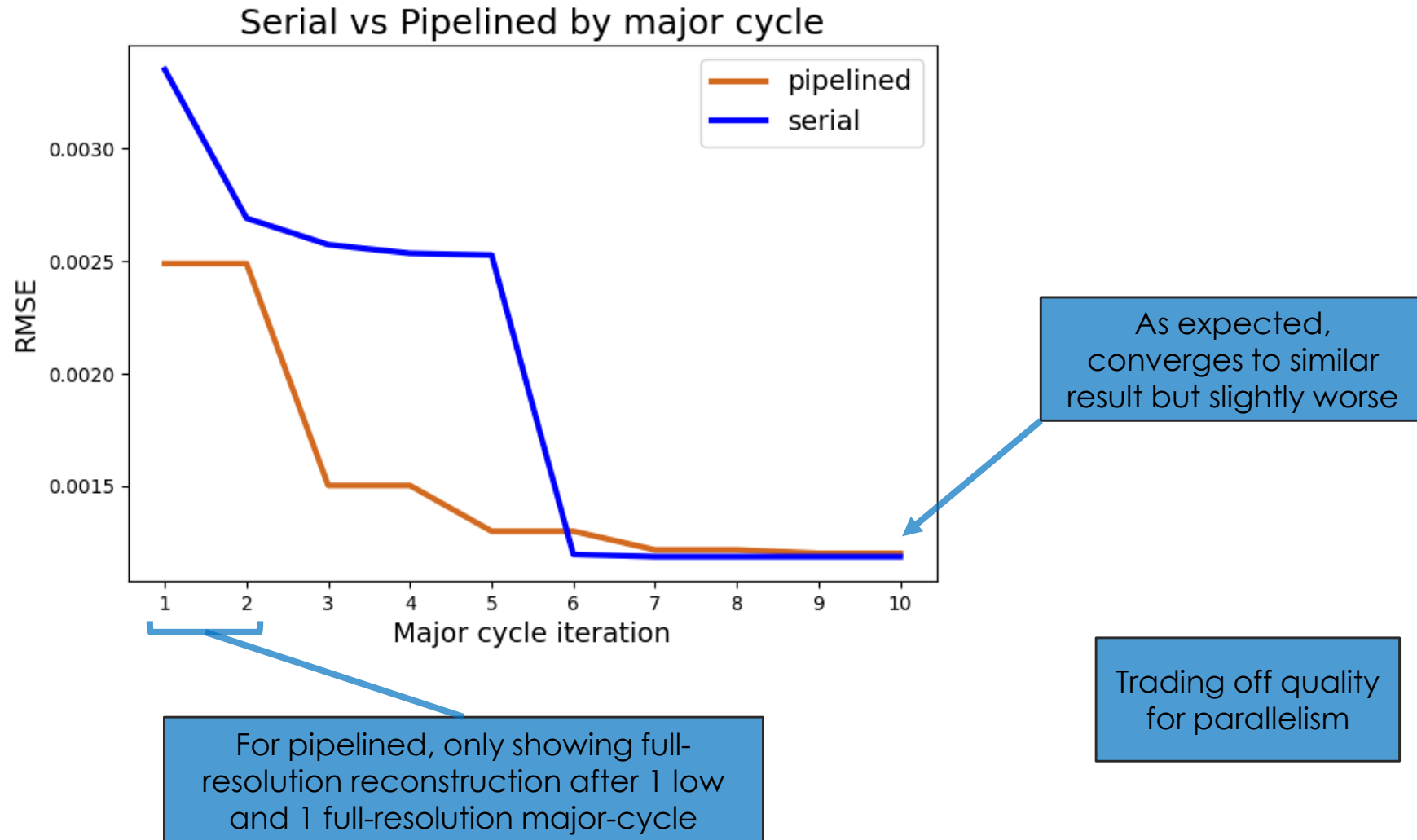
Results: Reconstruction quality of pipelined parallel method



As expected, converges to similar result but slightly worse

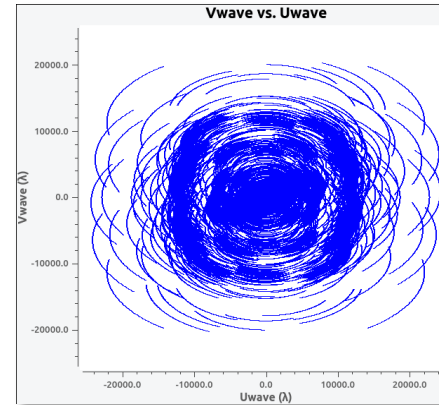
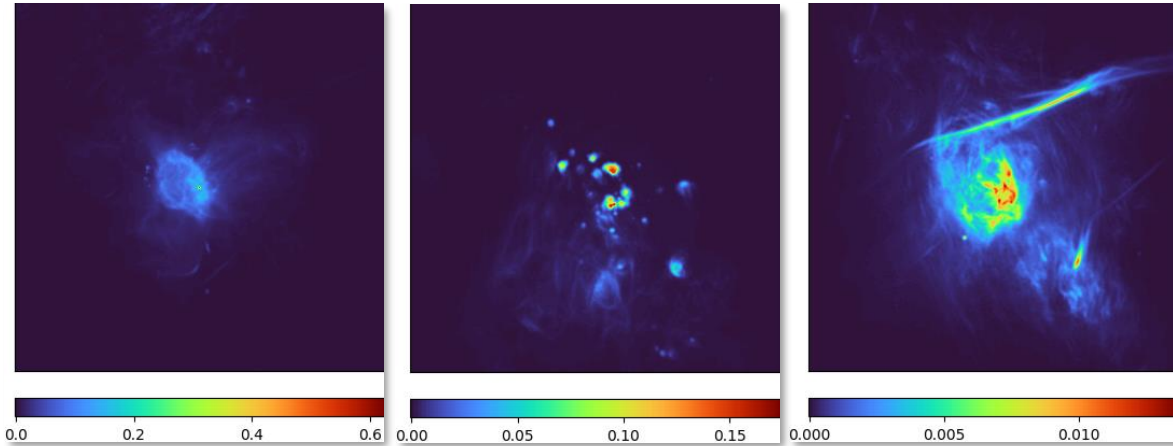
For pipelined, only showing full-resolution reconstruction after 1 low and 1 full-resolution major-cycle

Results: Reconstruction quality of pipelined parallel method



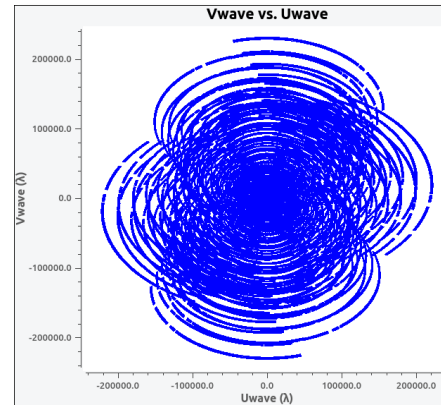
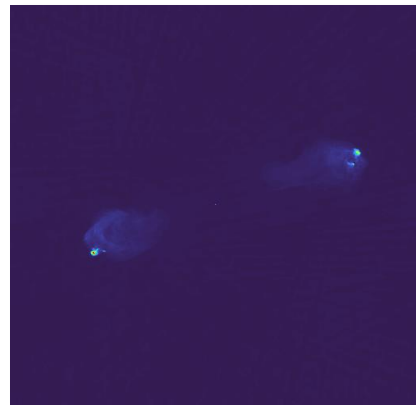
Results – Measurement sets

Simulated



- Initial images tapered and cutout from 1.28GHz mosaic produced in [1]
- Visibilities generated with MeerKAT configuration
- Exposure time of 4h, samples every 120s for to generate visibility positions
- Degrid to get visibility values
- Visibility noise artificially added

Real



- Measurement set fully described in [2]
- VLA, all 4 configurations, S-band
- Rather than the full dataset, we only use visibilities from 1988.5 MHz – 2020.5 MHz

[1] Heywood, I., et al. "The 1.28 GHz MeerKAT galactic center mosaic." *The Astrophysical Journal* 925.2 (2022): 165.
[2] Sebokolodi, M. Lerato L., et al. "A Wideband Polarization Study of Cygnus A with the Jansky Very Large Array. I. The Observations and Data." *The Astrophysical Journal* 903.1 (2020): 36.