

Vanishing of Quadratic Love Numbers of Schwarzschild Black Holes

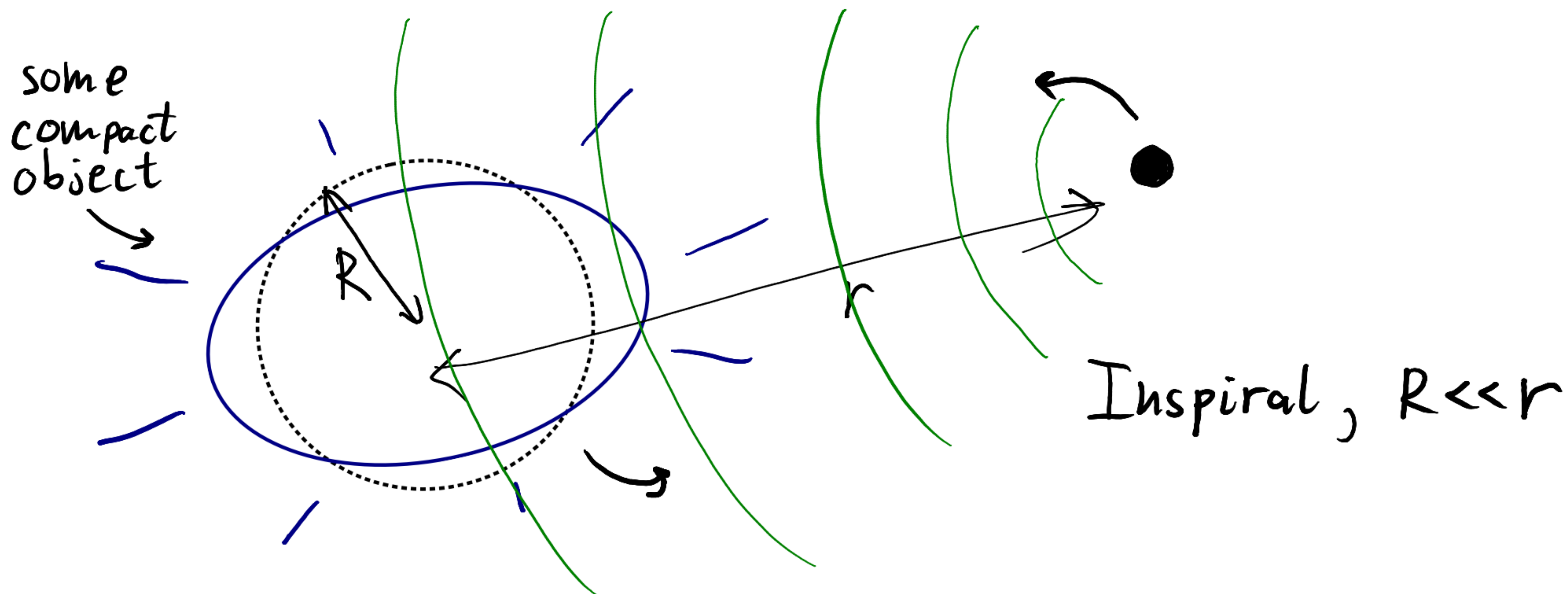
Théorie, Universe et Gravitation
November 2024, Annecy-le-Vieux

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[based on 2312.05065 and 2410.03542 with
M.M.Riva, L.Santoni, F.Vernizzi and S.Iteanu]



Reminder: Tidal effects and Love Numbers



- $R \ll r \Rightarrow$ "small" tidal forces \Rightarrow [Fang & Lovelace '05] [...]
 \Rightarrow Tidal effects encoded in **Love numbers λ**

Linear response: $(\text{Induced response}) = \lambda^{\text{lin}} \times (\text{Background Tidal field})$

ie. deformations/
induced multipoles

Why do we care about Love Numbers?

- Constrained for GW170817 \rightarrow constraint on EoS of neutron star [Abbott et al '18] [Annala et al '17]
- Black holes have $\lambda_{\text{BH}}^{\text{lin}} = 0!$ [Fang & Lovelace, Damour & Nagar, Le Tiec, Casals, ...]
(pure GR, $D=4$, asymp. flat)

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Study subleading tidal effects

- for BHs „leading = 0“

- to put leading order prediction on firmer ground

- Hidden symmetries

$$\left(\text{Tidal deformations} \right) \sim \underbrace{\lambda_{\text{BHs}}^{\text{lin}}}_{=0} \left(\text{Tidal field} \right) + \lambda^{\text{quad}} \left(\text{Tidal field} \right)^2 + \dots$$

Quadratic Love numbers

- 0th order in frequency (static)
- See [Gürlebeck '15] and [Poisson '20] on beyond linear tides

Love numbers

Worldline Effective Field Theory WEFT

* Clean and robust framework for general tidal effects
(systematic account of nonlinearities and free of gauge ambiguity)

[Goldberger & Rothstein '04] [...]

$$S[g_{\mu\nu}, x^{\mu}] \stackrel{R \ll r}{=} \underbrace{\frac{1}{16\pi G} \int d^4x \sqrt{-g} R}_{\text{Einstein-Hilbert}} - \underbrace{M \int d\tau}_{\text{point particle}} + \int d\tau \sum_{\ell \geq 2} \left[E_{\mu_2} Q_E^{\mu_2} + B_{\mu_2} Q_B^{\mu_2} \right] \leftarrow \text{Tidal effects}$$

↑ ↑
induced multipoles

• F_{μ_2} $\equiv \sum_{\mu_1, \dots} \nabla_{\mu_2-2} C_{\mu_2-1 \alpha \mu_2 \beta} u^\alpha u^\beta \rangle_{\text{STF}}$
derivatives of electric part of the Weyl tensor
(even)

B_{μ_2} $\equiv \frac{1}{2} \nabla_{\mu_1} \dots \nabla_{\mu_2-2} \epsilon_{\delta \mu_2-1} C_{\mu_2 \delta \alpha \beta} u^\alpha u^\beta$
magnetic part
(odd)

Love numbers

Worldline Effective Field Theory WEFT

$$S[g_{\mu\nu}, x^\mu] = \frac{1}{16\pi G} \int d^4x \sqrt{-g} R - M \int d\tau + \int d\tau \sum_{l \geq 2} \left[E_{\mu\nu} Q_E^{\mu\nu} + B_{\mu\nu} Q_B^{\mu\nu} \right]$$

multipoles

Tidal effects

Non linear response

$$Q_{\mu\nu}^E = \lambda^{\text{lin}} E_{\mu\nu} + \lambda_{LL\mu\nu}^{E^3} (E_{\mu\nu} E_{\mu\nu})_{\mu\nu} + \lambda_{LL\mu\nu}^{EB^2} (BB)_{\mu\nu} + O(E^3, EB^2)$$

(Linear) Love numbers Quadratic Love Numbers

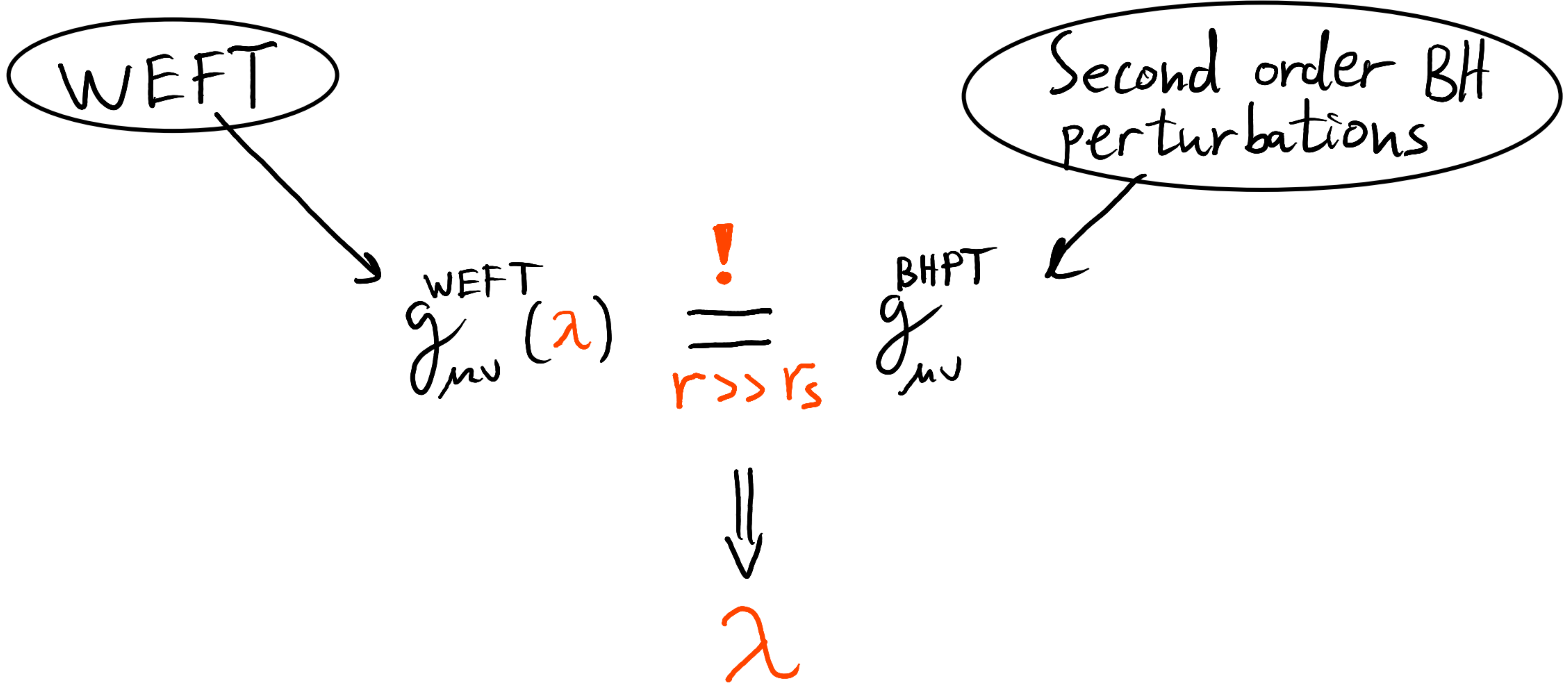
• Similar for $Q^B = \dots$
[Bern et al '20]

[Riva, Santoni,
NS, Vernizzi '23]

[Iteanu, Riva, Santoni,
NS, Vernizzi '24]

How to find Quadratic Love numbers?

- match BH perturbations on WEFT calculation



Second order BH perturbation theory

$$g_{\mu\nu} = \bar{g}_{\mu\nu}^{\text{Sch}} + \delta g_{\mu\nu}$$

- Solving Einstein's eqs for $\delta g_{\mu\nu}$ perturbatively:

$$\delta g_{\mu\nu} = {}^{(1)}\delta g_{\mu\nu} + {}^{(2)}\delta g_{\mu\nu}$$

1st order (linear): $\bar{D}^{(1)}\delta g_{\mu\nu} = 0 \xrightarrow{\text{solve}} {}^{(1)}\delta g_{\mu\nu}$

2nd order: $\bar{D}^{(2)}\delta g_{\mu\nu} = \bar{\Gamma}_{\mu\nu}^{(1)}[{}^{(1)}\delta g_0, {}^{(1)}\delta g]$ ← plug into source

- Parametrisation and Regge-Wheeler gauge [Regge & Wheeler '57]

$$\delta g_{\mu\nu} = \delta g_{\mu\nu}^{(\text{even})} + \delta g_{\mu\nu}^{(\text{odd})}$$

master variable
↓

- focus on even: $\delta g_{\mu\nu}^{(\text{even})} \sim \delta g_{tt} \sim H_0(r, \theta, \varphi) = \sum_{\ell m} H_0^{\ell m}(r) Y_{\ell m}^e(\theta, \varphi)$

Second order BH perturbation theory

Linear solutions: ${}^{(1)}H_0^{lm} = \sum_+ P_l^{(2)} \left(2\frac{r}{r_s} - 1 \right)$

$$\left(\partial_r^2 + \frac{2r-r_s}{f(r)r^2} \partial_r - \frac{\lambda r^2 f(r) + r_s^2}{r^4 f^2(r)} \right) H_0^{lm}(r) = \sum_{\substack{l_1 m_1 \\ l_2 m_2}} \sum_{m, m_1, m_2} H_0^{l l_1 l_2}(r)$$

sources, known functions
of linear solutions

- Impose regularity of solution at the horizon

- Due to specific form of the source we find that all regular solutions at 2nd order are

finite polynomials in (r/r_s) !

[Iteanu, Riva, Santoni,
NS, Vernizzi 124]

Matching to WEFT

[Iteanu, Riva, Santoni,
NS, Vernizzi '24]

• BHPT gives ${}^{(2)}H_0^{lm} = \mathcal{E}_+^z \left[A^{(BH)} r^n + \#r_s r^{n-1} + \dots + \#r_s^{n-1} r + \#r_s^n \right]$

• We calculate ${}^{(2)}H_0$ and ${}^{(2)}h_0$ to leading order in r_s in WEFT

WEFT: ${}^{(2)}H_0^{lm} = A^{(EFT)} \mathcal{E}_+^z r^n + \lambda^{E^3} \frac{\#}{r^{l+1}}$

Matching to WEFT

• BHPT gives ${}^{(2)}H_0^{lm} = \mathcal{E}_+^z \left[A^{(BH)} r^n + \#r_s r^{n-1} + \dots + \#r_s^{n-1} r + \#r_s^n \right]$

• We calculate ${}^{(2)}H_0$ and ${}^{(2)}h_0$ to leading order in r_s in WEFT

WEFT: ${}^{(2)}H_0^{lm} = \mathcal{E}_+^z A^{(EFT)} r^n + \lambda^{E^3} \frac{\#}{r^{l+1}}$

$A^{(BH)} = A^{(EFT)}$
 → we matched correctly

for BHs there are no $\frac{1}{r^{l+1}}$
 ⇓ just positive powers

$\lambda_{123}^{E^3} = 0$
 $\lambda_{123}^{EB^2} = 0$
 for all multipoles!

[Iteanu, Riva, Santoni, NS, Vernizzi '24]

Conclusions

- * Quadratic static tidal response of Schwarzschild black holes vanishes for all multipoles

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Hidden symmetries beyond linear perturbations?

• Linear perturbations in near zone have $SO(4,2)$ $\xrightarrow{\text{static}}$ $SO(3,1)$ for Schwarzschild
[Hwu et al '21, '22] [Charalambous et al '21, '22] \Downarrow
 $\lambda^{\text{lin}} = 0$

• [Combaruzier-Szeitszneider et al '24] and [Kehagias & Riotto '24] found symms for non-linear even sector with axial symm.

• Symmetries including odd sector?

Conclusions

- * Quadratic static tidal response of Schwarzschild black holes vanishes for all multipoles
- * Simplicity of 2nd order BHPT \rightarrow smarter way of organising calculation?
- * Extend to all orders, Kerr, RN ; consider $D > 4$, AdS ...