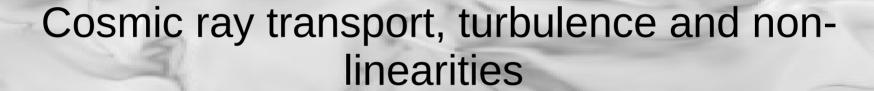
## Matthieu BOUCHET



With Yoann GENOLINI at LAPTh







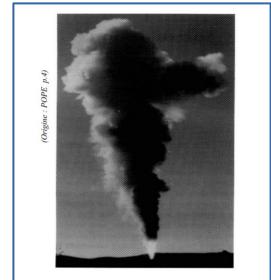
### A brief talk about turbulence

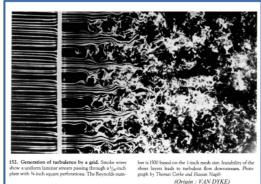
### **Matthieu BOUCHET**

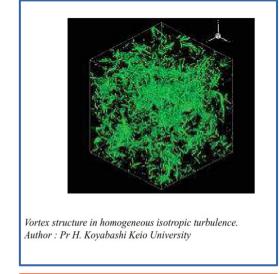
Cosmic ray transport, turbulence and non-linearities

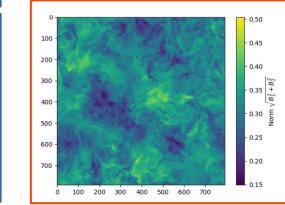
### In Hydrodynamics :

- Chaotic movement of particles in a fluid
- Caracterized by vorticies, strong instabilities









### In Magnetohydrodynamics

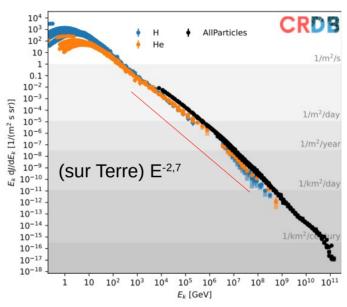
- → Turbulent stream of plasma
- Consequences on electromagnetic field and transport of particles in the interstellar medium

# Introduction to cosmic rays

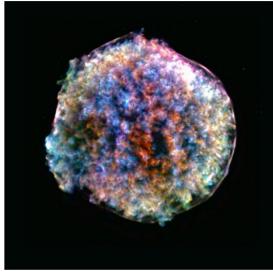
### Matthieu BOUCHET

Cosmic ray transport, turbulence and non-linearities

- Relativistic particles
  - → Power law of the energy spectra
- Galactic and extra-galactic origins
  - → Supernovae remnants
  - ultra-high-energy : not very well known



Spectra Flux VS Energy of cosmic rays, observed from Earth



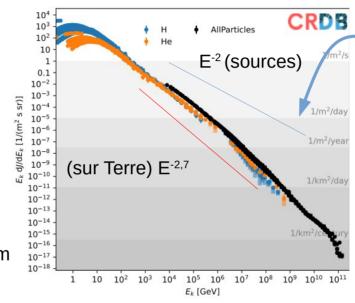
Tycho (Tycho Brahé, 1572), observed in X rays by Chandra telescope

# Introduction to cosmic rays

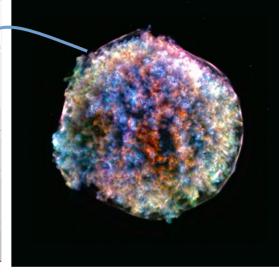
#### **Matthieu BOUCHET**

Cosmic ray transport, turbulence and non-linearities

- Relativistic particles
  - → Power law of the energy spectra
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  - → ultra-high-energy : not very well known
- Principal problem : what phenomenon influence their propagation?
  - Spectra observed at sources different from the one observed on Earth



Spectra Flux VS Energy of cosmic rays, observed from Earth



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# Introduction to cosmic rays

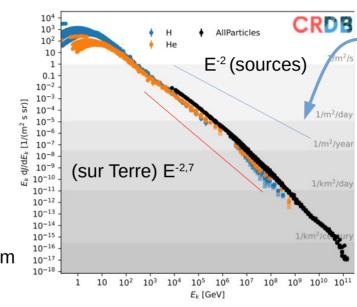
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Cosmic ray transport, turbulence and non-linearities

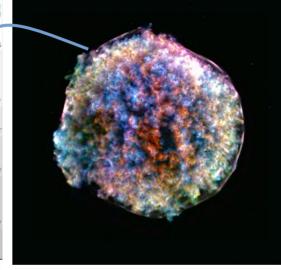
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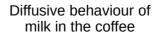
- Influenced by magnetic fields
- Complexe trajectories
- diffusion of cosmic rays

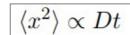


Spectra Flux VS Energy of cosmic rays, observed from Earth



Tycho (Tycho Brahé, 1572), observed in X rays by Chandra telescope







## **Problematics**

#### **Matthieu BOUCHET**

Cosmic ray transport, turbulence and non-linearities

 Turbulent interstellar medium plasma

- The magnetic field serves as a guide (or rails) for cosmic rays
- → Turbulence induces magnetic perturbations



Perturbed magnetic field due to turbulence:

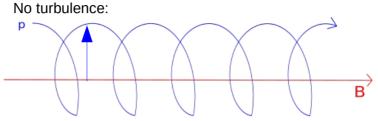
$$R_L = \frac{\gamma m v}{qB}$$

### **Problematics**

#### **Matthieu BOUCHET**

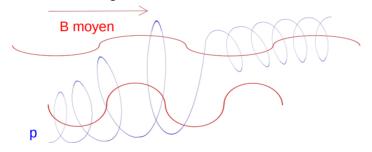
Cosmic ray transport, turbulence and non-linearities

- Turbulent interstellar medium plasma
  - The magnetic field serves as a guide (or rails) for cosmic rays
  - Turbulence induces magnetic perturbations
- Cosmic ray trajectories are perturbed
  - Interactions between particles and magnetic field fluctuations
  - → Field line random walk (FLRW)

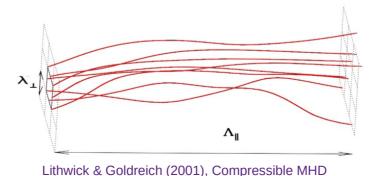


 $R_L = \frac{\gamma m v}{q B}$ 

Perturbed magnetic field due to turbulence :



Gyro-resonance: the particle jumps from one field line to another



Turbulence in Interstellar Plasmas

Field line random walk

# Conclusion

### **Matthieu BOUCHET**

Cosmic ray transport, turbulence and non-linearities

Goal: develop a theoretical framework for cosmic ray transport to learn more about present and futur observables

Thank you for your attention!