

NUTRIG

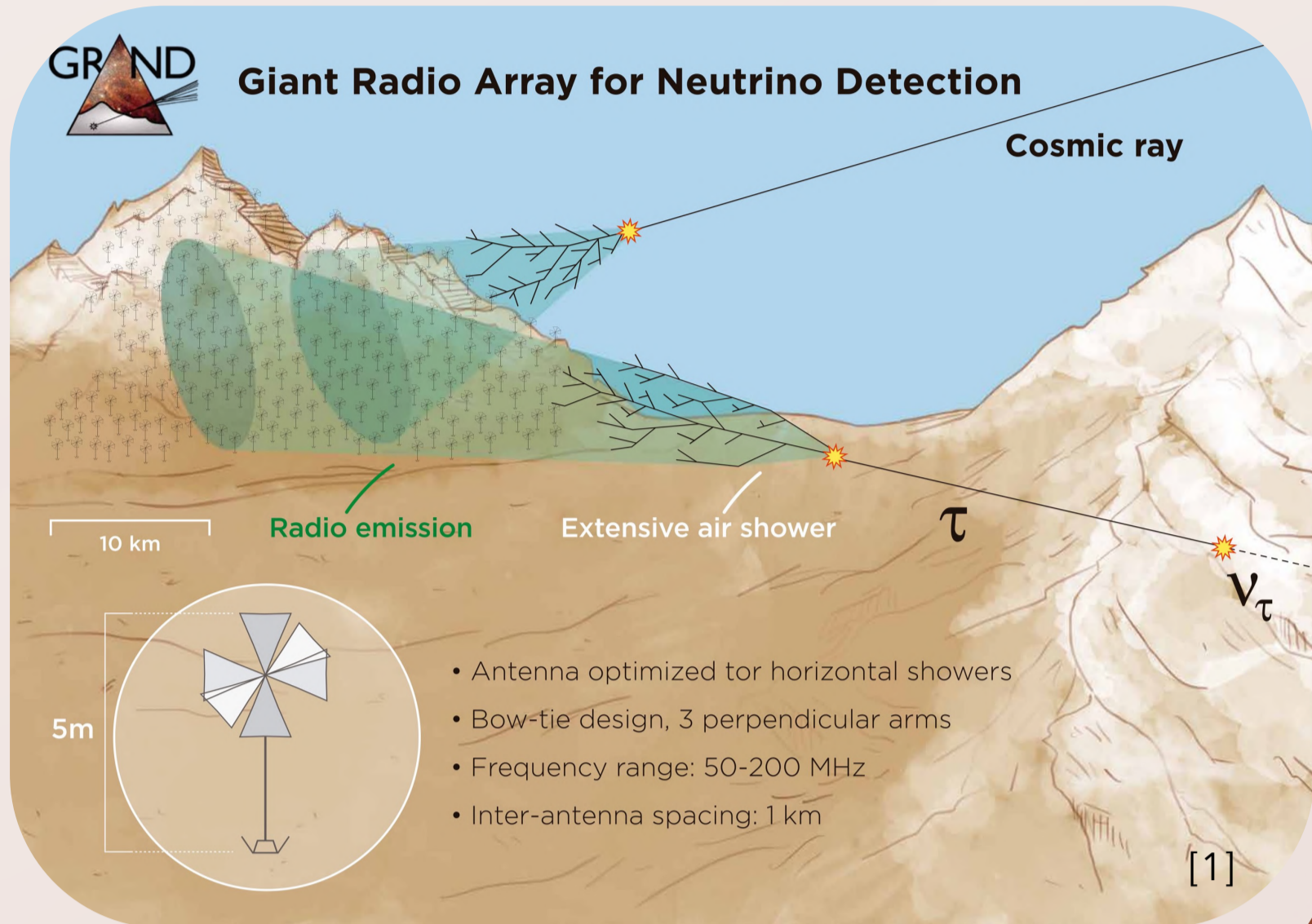
Towards an Autonomous Radio Trigger for GRAND

Pablo Correa¹ for the GRAND collaboration

¹Laboratoire de Physique Nucléaire et de Hautes Energies (LPNHE), 4 place Jussieu, F-75252, Paris Cedex 5, France

Giant Radio Array for Neutrino Detection

- ▶ GRAND [1,2] targets **UHE neutrino astronomy** (>100 PeV)
- ▶ Aim to detect **transient radio pulses** of ν_τ -induced horizontal air showers
- ▶ **20 GRAND10k arrays** will cover **200,000 km²** across the globe
- ▶ **Pathfinders:** GRANDProto300 (China) and GRAND@Auger (Argentina) [2,3]
- ▶ Major requirement: **autonomous radio trigger** for air showers



NUTRIG Principle

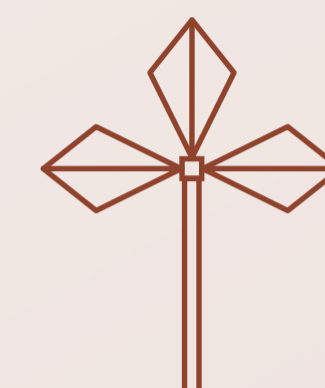
Radio-Emission Model

- ▶ Focus on **horizontal air showers** [4,5]
- ▶ Model emission in the **30-230 MHz** frequency band
- ▶ Apply model to **GRAND (prototype) sites**



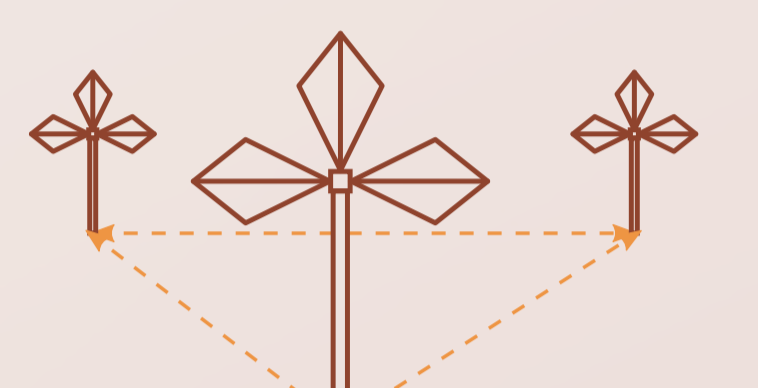
First-Level Trigger

- ▶ At the level of a single **detection unit**
- ▶ Exploit air-shower radio **pulse characteristics** [5,6]
- ▶ Target **FLT rate of 100 Hz** to limit data bandwidth

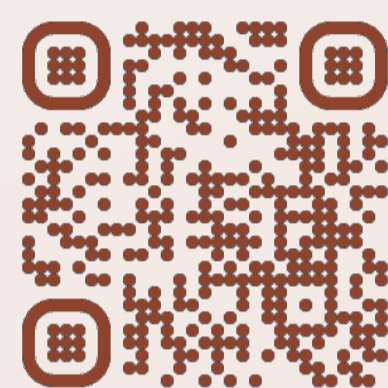


Second-Level Trigger

- ▶ At the level of a detection-unit **array**
- ▶ Use radio-emission model and **FLT information**
- ▶ Further **reduce noise** before disk storage

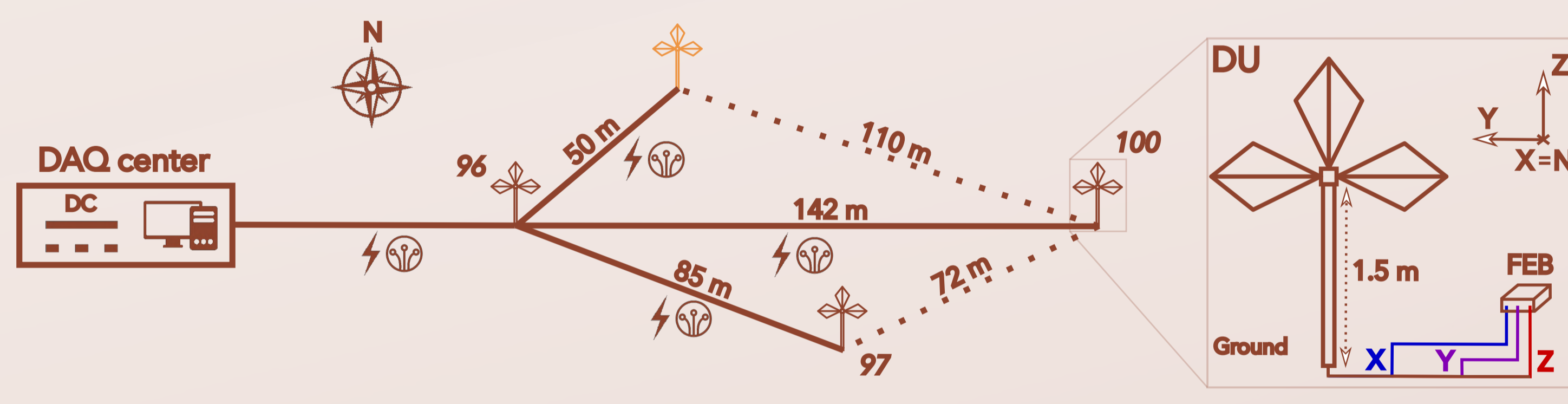


YouTube: @grand-observatory



GRAND@Nançay Prototype

- ▶ Dedicated to NUTRIG: **FLT development**
- ▶ Located at **Nançay Radio Observatory**
- ▶ Setup consists of **4 DUs** (3 currently deployed)
- ▶ **DAQ center** contains computer and DC power supply
- ▶ **Front-end electronics board** powered via cables
- ▶ Data transfer via **optical fibers** (no WiFi allowed)



Analysis of GRAND@Nançay Background

Front-End Chain + DAQ

1. **Low-noise amplifier:** 18 dB
2. **Band-pass filter:** 30-230 MHz
3. **Band-stop filter:** FM frequencies
4. **Variable-gain amplifier:** set to 23.5 dB
5. **ADC:** 14-bit digitizer at 500 Msamples/s
6. **FPGA+CPU:** trigger + event building
7. **DAQ center:** event storage + analysis

Dataset

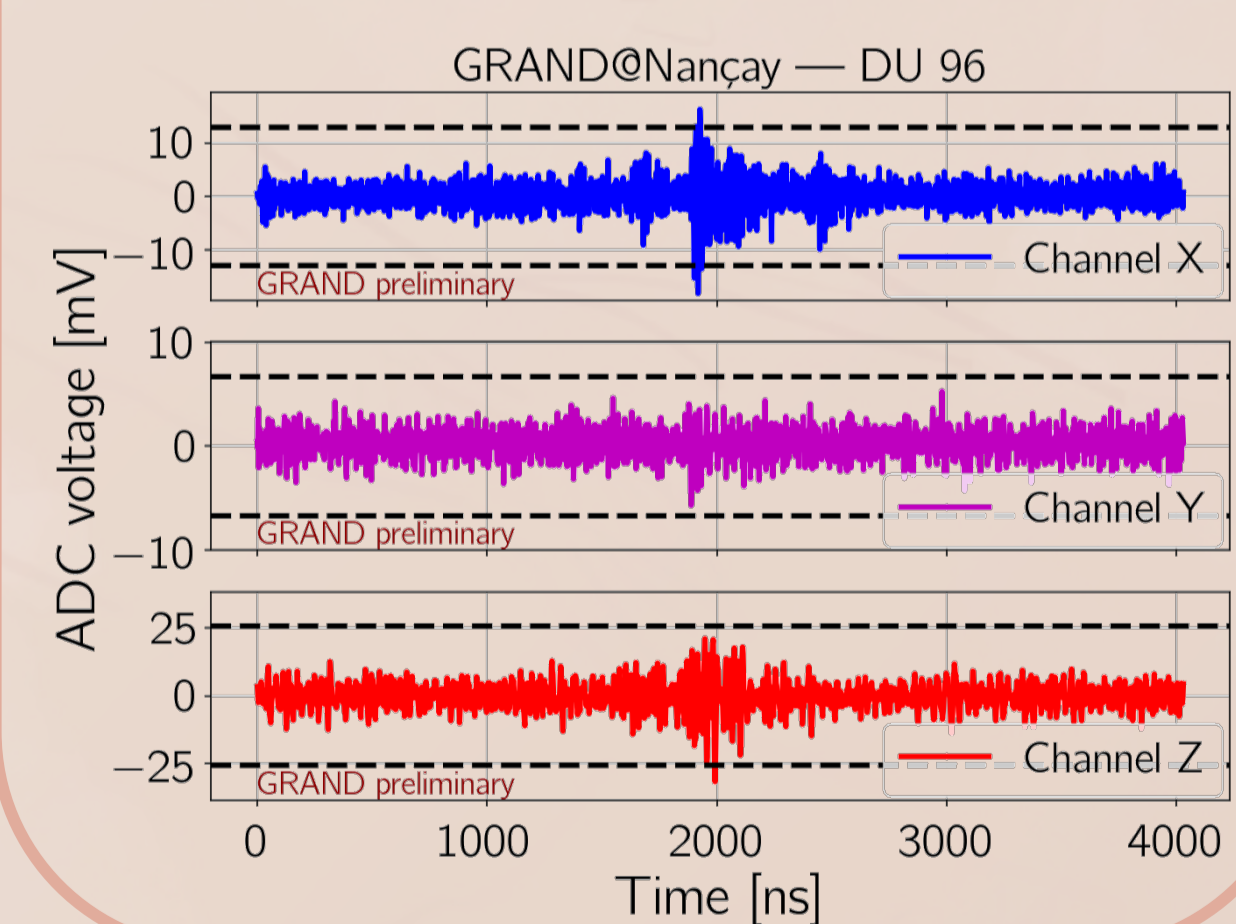
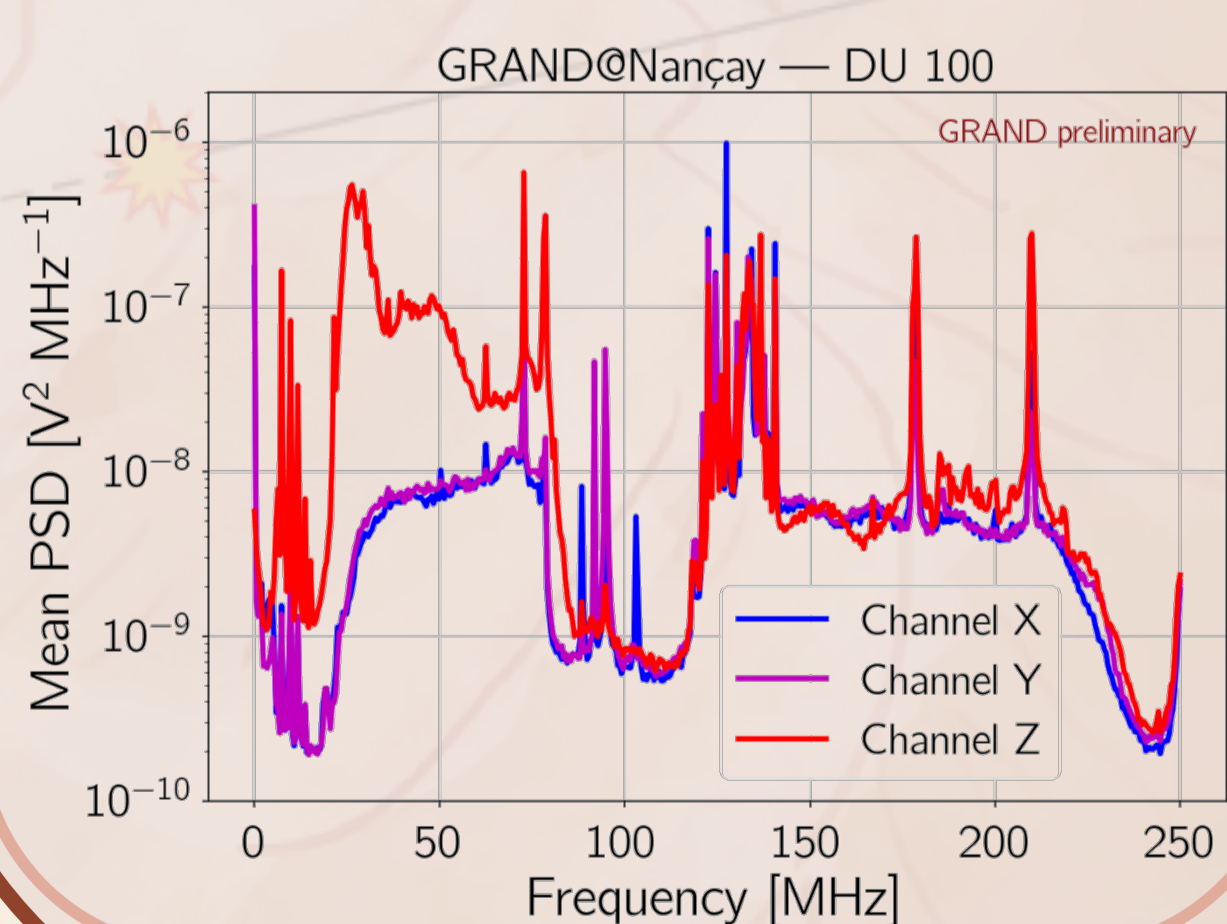
- ▶ **DAQ run:** night of 26-27 June 2023
- ▶ **DAQ mode:** force trigger every 10 s
- ▶ **Stable run** between 00:49-09:33 UTC+2

Transient RFI Pulses

- ▶ **Background** for autonomous radio arrays is transient RFI
- ▶ **Search for transient RFI pulses** in GRAND@Nançay data
- ▶ **Transient definition:** pulse that exceeds 5x the standard deviation of time trace
- ▶ **Observed RFI pulses** are typically wider (>100 ns) than air-shower pulses
- ▶ **Transient RFI rate** estimated to be several 100 Hz (low statistics)

Mean RFI Spectrum

- ▶ Measure **spectrum of radio-frequency interference** at GRAND@Nançay
- ▶ Compute **time-averaged power spectral density** of all collected data
- ▶ **Features:** short waves, FM, local emitters, aeronautic comms, DAB

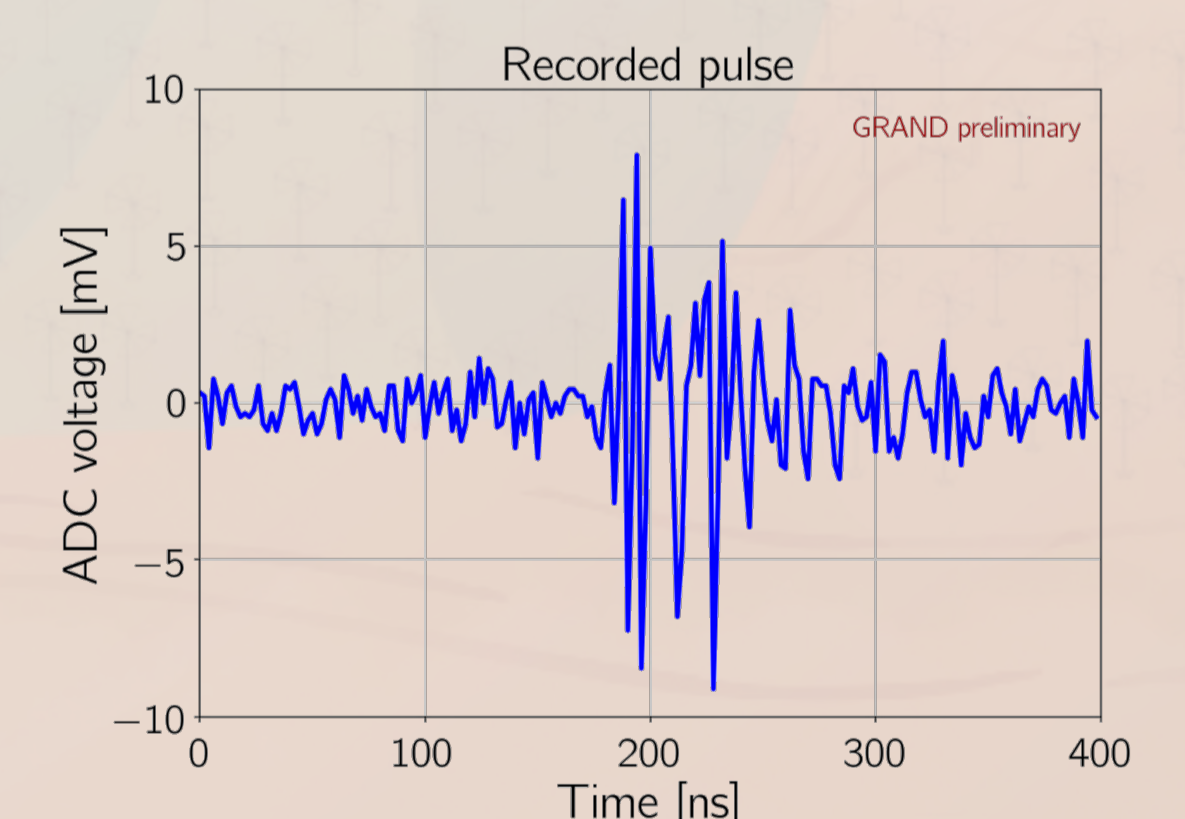
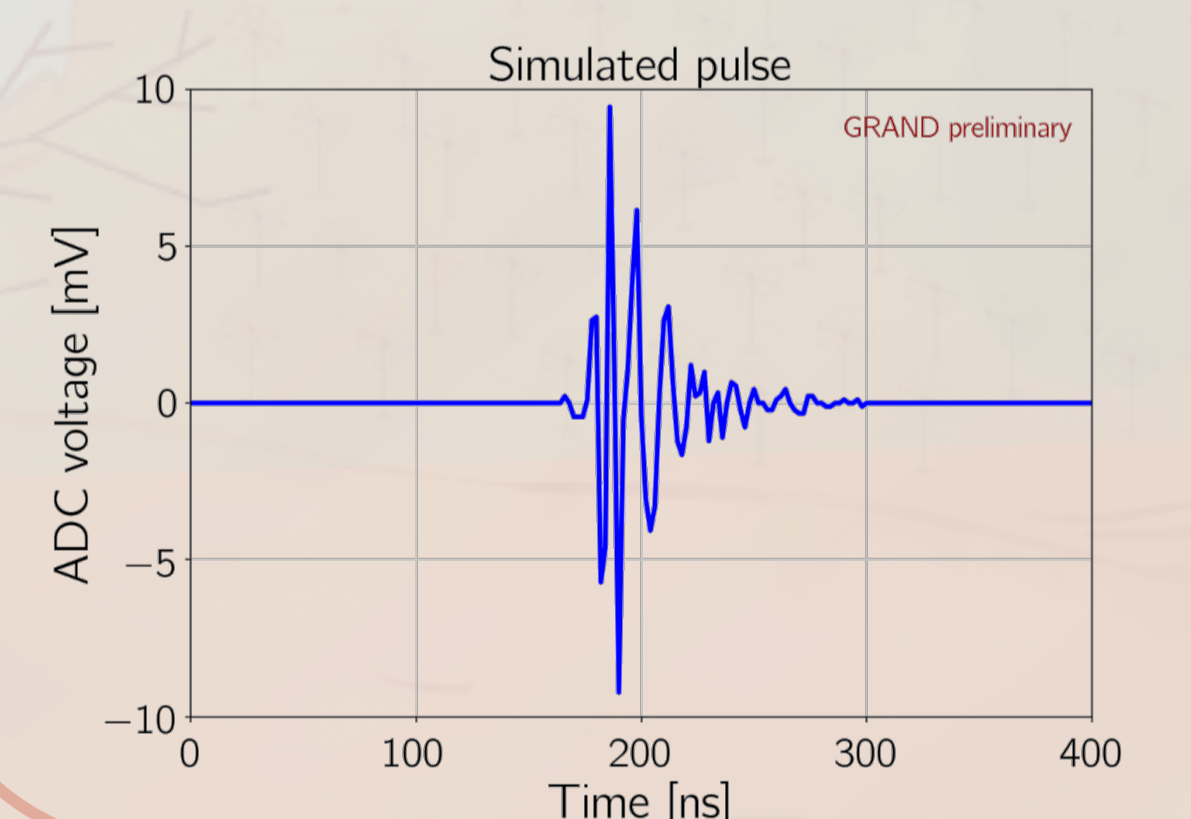
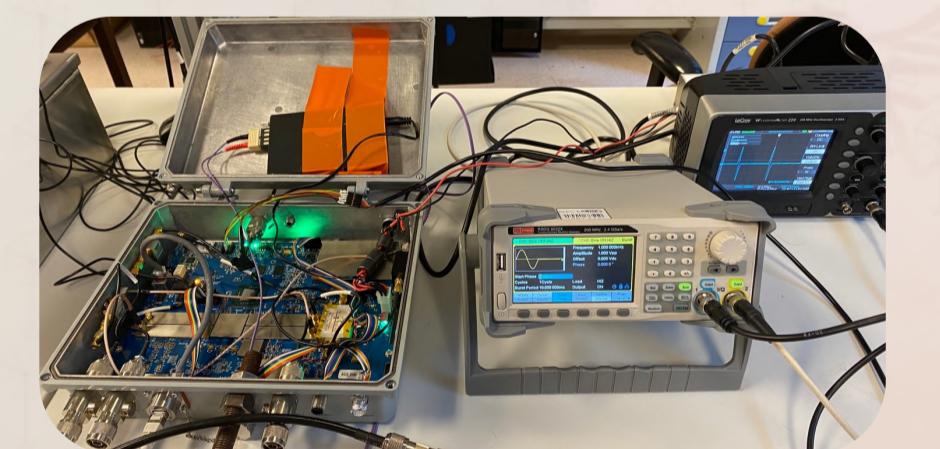


NUTRIG Implementation Strategy

- ▶ **Develop FLT + SLT algorithms** using simulated air-shower signals + background data [7]
- ▶ Port FLT algorithm to FPGA on FEB and perform **checks at LPNHE testbench** in Paris
- ▶ **Test FLT algorithm at GRAND@Nançay** in real data-taking conditions
- ▶ **Implement FLT + SLT at GrandProto300** and optimize for scalability to GRAND10k arrays

Generation of Simulated Air-Shower Pulse at LPNHE

- ▶ **Simulate air-shower pulse** at antenna-nut output
- ▶ **Reproduce pulse** with custom-wave generator
- ▶ **Inject** generated pulse to FEB and **record**
- ▶ **Compare** to pulse simulation at ADC level



References

- [1] GRAND collaboration, J. Álvarez-Muñiz et al. (2020), *Sci. China Phys. Mech. Astron.* **63** 219501
- [2] GRAND collaboration, J. Torres de Mello Neto (these proceedings), *PoS ICRC2023* 1050
- [3] GRAND collaboration, P. Ma (these proceedings), *PoS ICRC2023* 304
- [4] T. Huege, F. Schlüter, and L. Brenk (2021), *PoS ICRC2019* 294
- [5] F. Schlüter and T. Huege (2023), *JCAP* **01** 008
- [6] D. Charrier et al. (2019), *Astropart. Phys.* **110** 15-29
- [7] GRAND collaboration, S. Le Coz (these proceedings), *PoS ICRC2023* 224