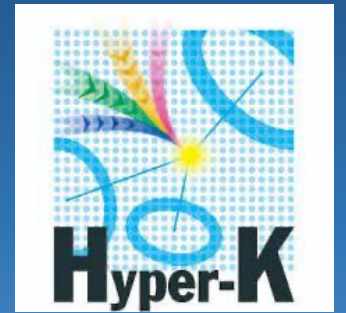


# Oscillation Analysis in the T2K and Hyper-K Experiments



**Denis Carabadjac**

dcarabadjac@llr.in2p3.fr



**Biennale LLR**

**02.02.2024**

**université  
PARIS-SACLAY**



**Probing the models beyond  
the Standard Model**

**Probing cosmology theories**

A flowchart diagram showing two arrows pointing from the top text blocks to a central box, and a single arrow pointing from the central box to the bottom text block.

Neutrino - pivotal indicator

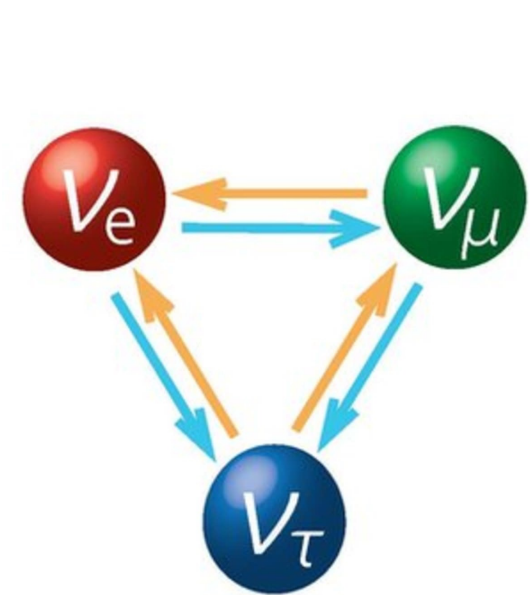
The study of neutrino oscillations is a pursuit of fundamental scientific knowledge

# Neutrino oscillations

The study of neutrino oscillations is a pursuit of fundamental scientific knowledge

Neutrino oscillations describe a physics phenomenon where a neutrino created with a specific lepton flavor (electron, muon, or tau) can later be measured to have a different flavor.

Appearance dis. define



Flavour eigenstates

$$\begin{pmatrix} e \\ \mu \\ \tau \end{pmatrix} \longleftrightarrow \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}$$

Participate in neutrino interactions and productions

Mass eigenstates

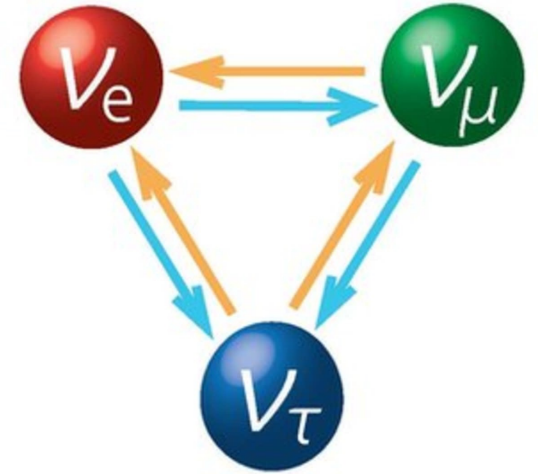
$$= U_{PMNS} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \longleftrightarrow \begin{pmatrix} m_1 \\ m_2 \\ m_3 \end{pmatrix}$$

Propagation in space-time

# Neutrino oscillations

Petkov will speak about it

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U_{PMNS} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$



$$U_{PMNS} = U(\delta_{CP}, \theta_{12}, \theta_{13}, \theta_{23})$$

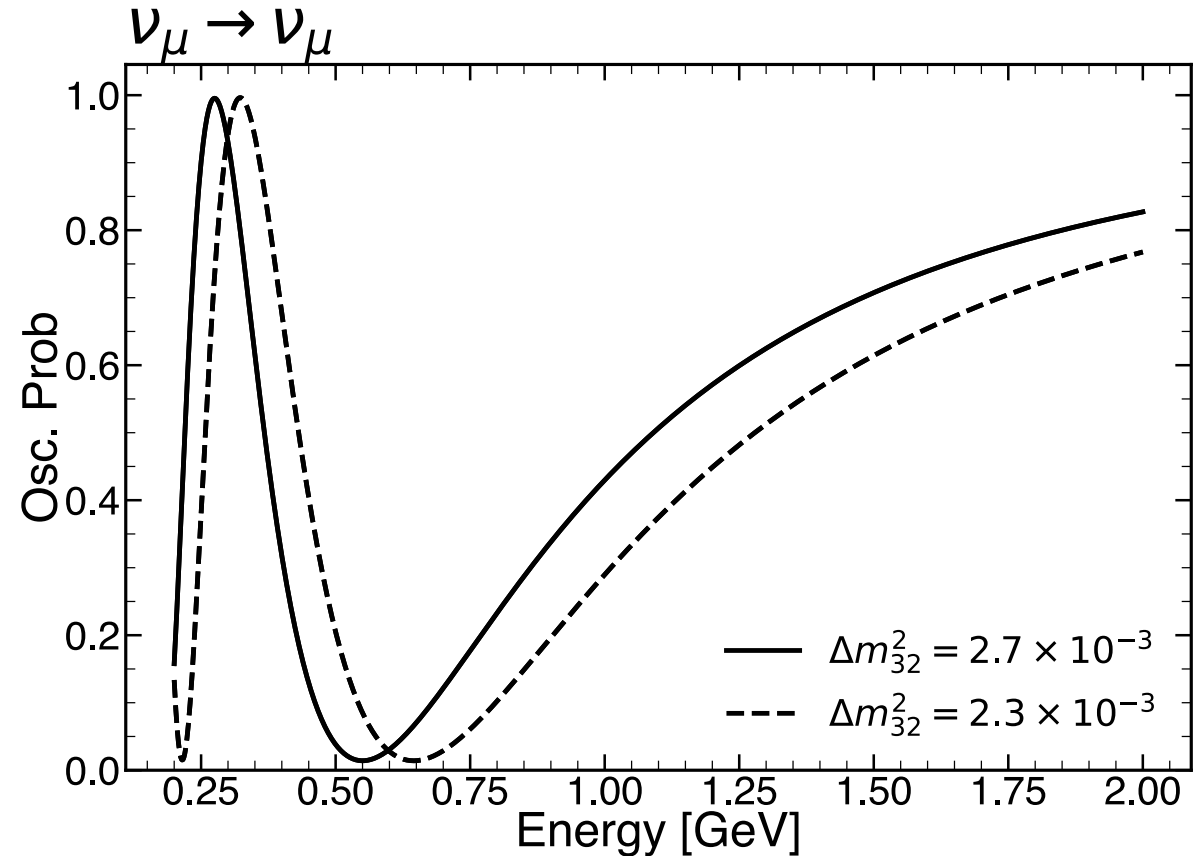
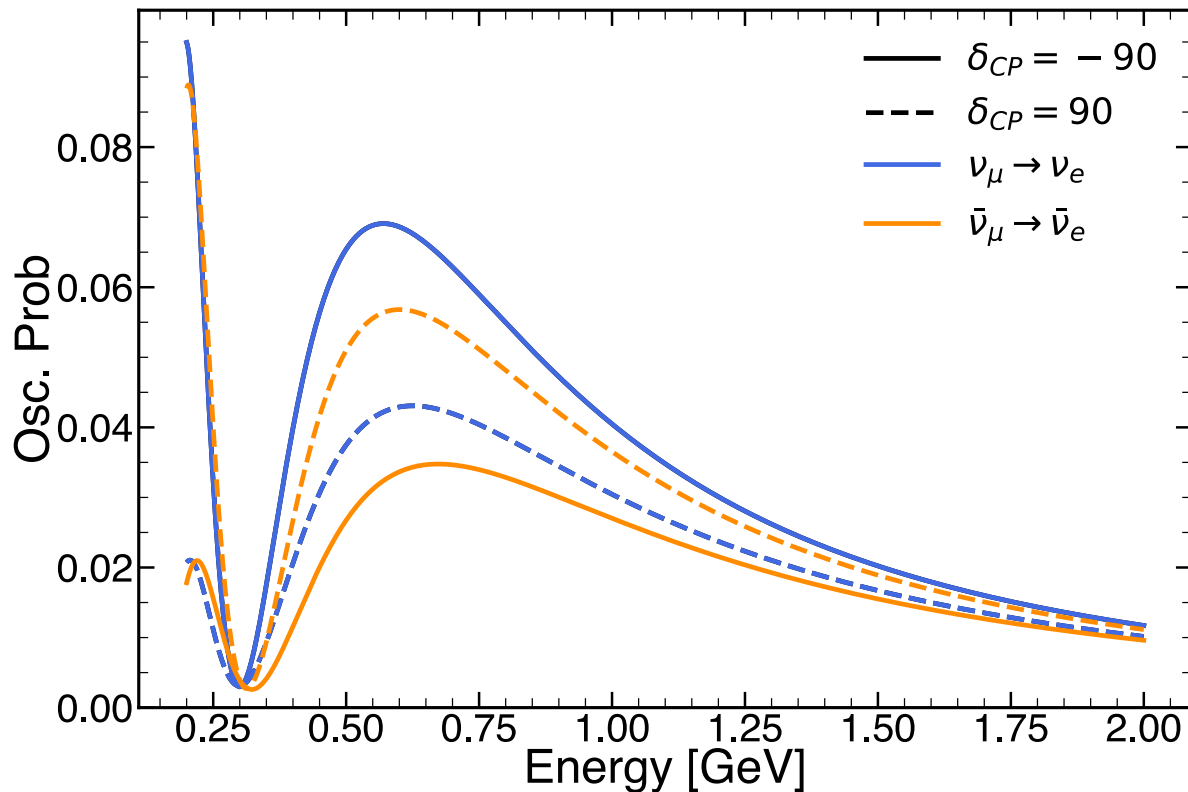
$$\Delta m_{ij}^2 = m_i^2 - m_j^2$$

Oscillation parameters:  $\delta_{CP}, \theta_{12}, \theta_{13}, \theta_{23} + \Delta m_{12}^2, |\Delta m_{32}^2|, \text{sign}(\Delta m_{32}^2)$  (Mass ordering)

With accelerator neutrino experiment we can measure  $\delta_{CP}, \theta_{23}, \theta_{13}, |\Delta m_{32}^2|$   
Small sensitivity to mass ordering (MO)\*

\*thus usually results are presented under different hypothesis of MO

# Neutrino oscillations



- **neutrino energy** measurement is **crucial** for oscillation inference
- **oscillation channel** (to electron or muon) and **mode** (neutrino or antineutrino) carries **complementary information**, on oscillation

➤ To infer the neutrino oscillation we need to **measure accurately** the **flavour** and **neutrino energy**

➤ **We would like to measure different oscillation channels and different modes**

# Requirements for an experiment



- We need neutrino which we can be produced with abundant flux
- We need to monitor this flux
- We need measure this flux before neutrino oscillations
- Finally, we need to measure this flux after neutrino oscillations

T2K experiment 

# T2K experiment

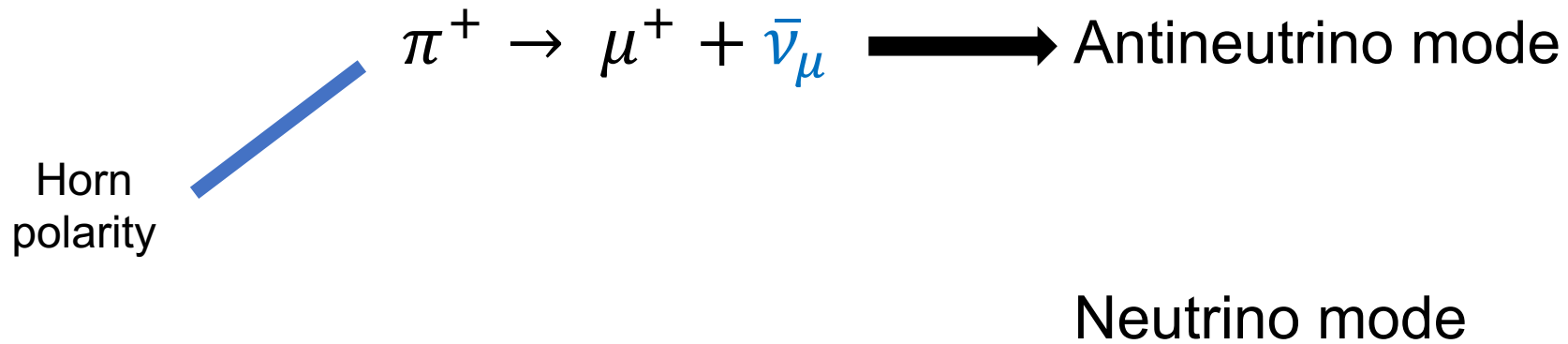




- ➔ • We need neutrino which we can be produce with abundant flux
- We need to monitor this flux
- We need measure this flux before neutrino oscillations
- Finally, we need to measure this flux after neutrino oscillations

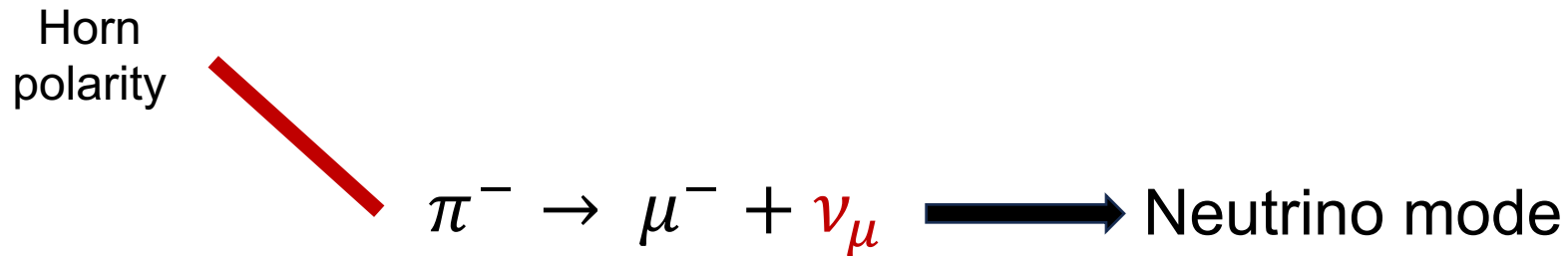
## Proton accelerator

- 30 GeV proton accelerator
- Carbon target for pions production
- 3 magnetic horns for pion focusing (positive or negative)



## Accelerator

- 30 GeV proton accelerator
- Carbon target for pions production
- 3 magnetic horns for pion focusing (positive or negative)



## Accelerator

- 30 GeV proton accelerator
- Carbon target for pions production
- 3 magnetic horns for pion focusing (positive or negative)

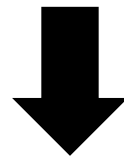


$\nu_{\mu}$

$\bar{\nu}_{\mu}$

$\nu_{\mu}$  &  $\bar{\nu}_{\mu}$

Accelerator neutrino experiment can measure properties on neutrino and antineutrino and to study the asymmetry between them



Can measure  $\delta_{CP}$  responsible for CP violation in lepton sector

- We need neutrino which we can be produce with abundant flux



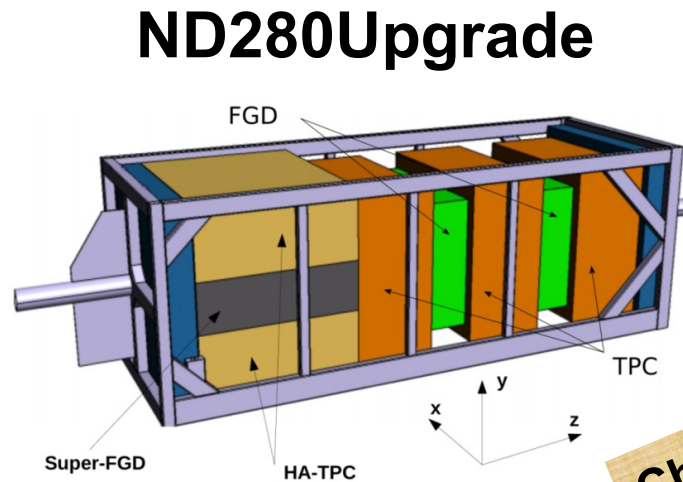
- We need to monitor this flux
- We need measure this flux before neutrino oscillations
- Finally, we need to measure this flux after neutrino oscillations

# Experimental measurement



- We need measure this flux before neutrino oscillations

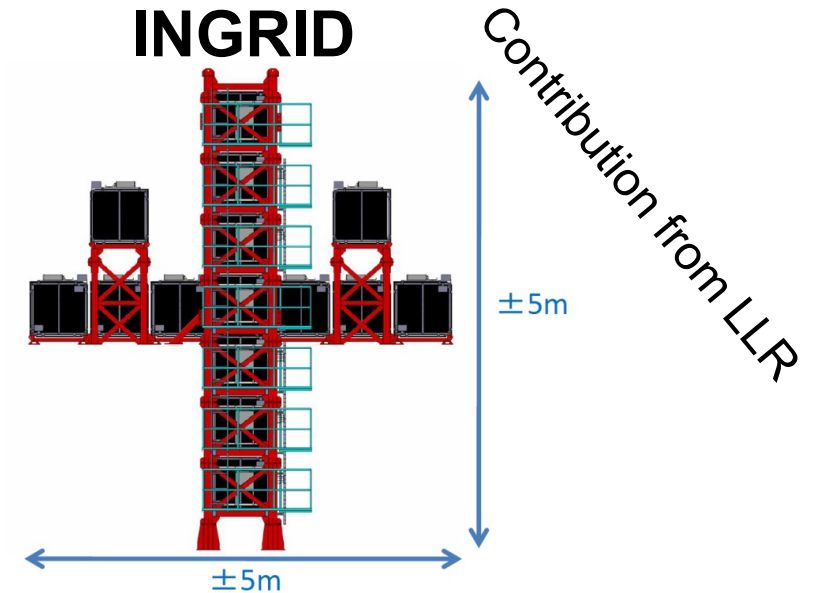
- We need to monitor this flux



ND280 Upgrade

Contribution from LLR to SFGD electronics

Check Viet's slides



INGRID

Contribution from LLR



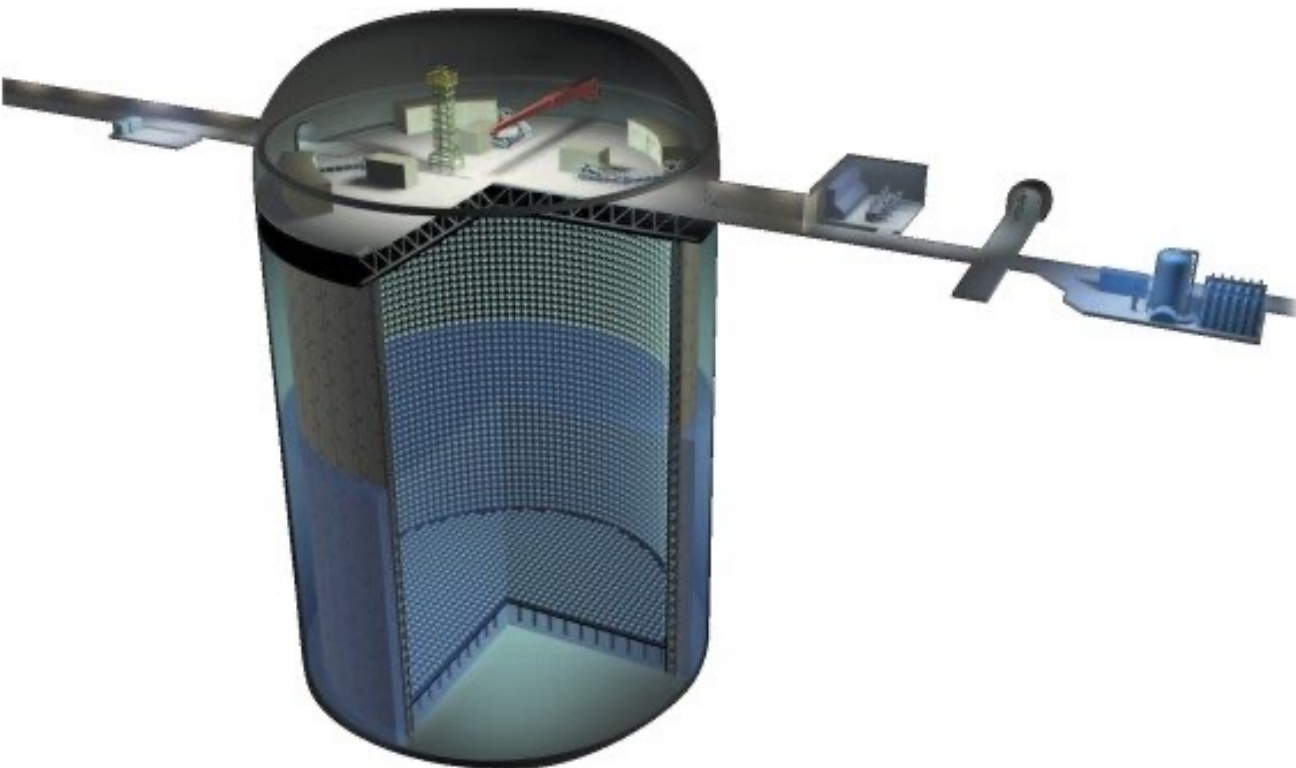
- We need neutrino which we can be produce with abundant flux ✓

- We need to monitor this flux ✓

- We need measure this flux before neutrino oscillations ✓

➔ • Finally, we need to measure this flux after neutrino oscillations

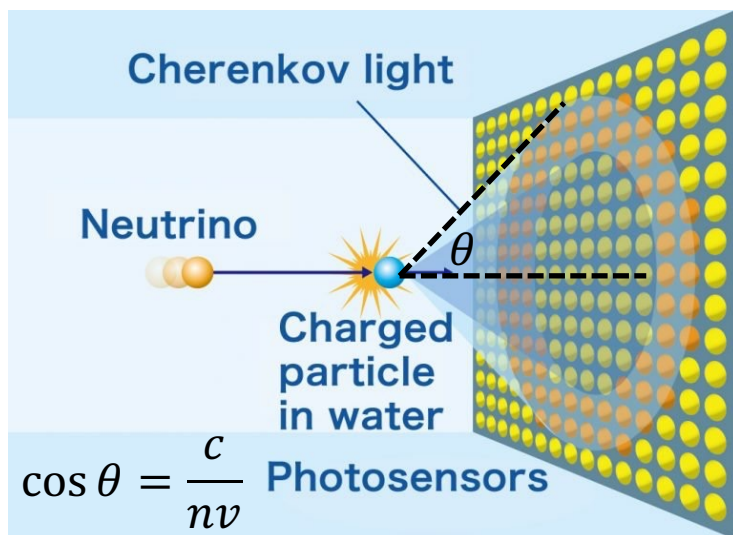
# T2K experiment



Detects water Cherenkov light from charged particle and reconstructs events with PMT charge & timing information.

	<b>Super-Kamiokande</b>
Site	Mozumi
Number of ID PMTs	11,129
Photo-coverage	40%
Mass/Fiducial Mass	50 kton/22.5 kton

# Super-Kamiokande events



Ring imaging water Cherenkov detector provides informative event reconstruction:

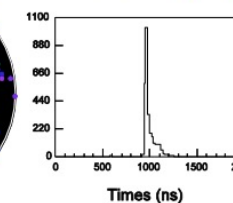
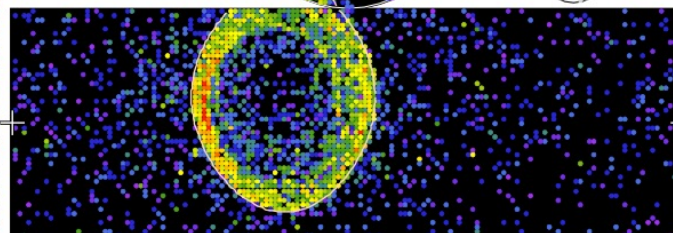
- 1) Number of Cherenkov photons  $\propto$  **momentum of the particle**
- 2) Arrival time  $\propto$  **interaction position**
- 3) Number of cherenkov rings  $\propto$  **number of the produced charged particle**

## $\mu$ -like event

Run 3962 Sub 125 Ev 965982  
97-05-01:15:32:29  
Inner: 2887 hits, 9607 pE  
Outer: 1 hits, 0 pE (in-time)  
Trigger ID: 0x03  
D wall: 1690.0 cm  
FC  $\mu$ -like, p = 1323.6 MeV/c

Charge (pe)

- >26.7
- 23.3-26.7
- 20.2-23.3
- 17.3-20.2
- 14.7-17.3
- 12.2-14.7
- 10.0-12.2
- 8.0-10.0
- 6.2-8.0
- 4.7-6.2
- 3.3-4.7
- 2.2-3.3
- 1.3-2.2
- 0.7-1.3
- 0.2-0.7
- < 0.2

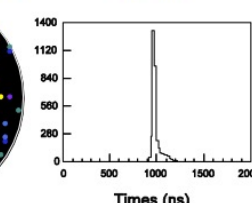
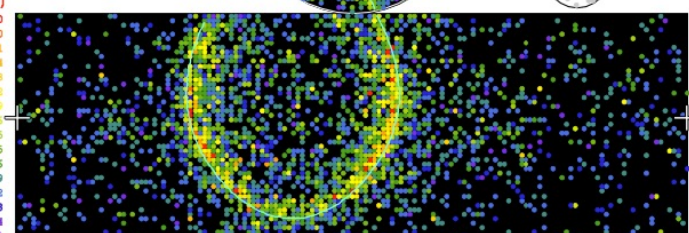


## e-like event

Run 5704 Event 3551590  
98-02-17:07:14:39  
Inner: 3397 hits, 7527 pE  
Outer: 0 hits, 0 pE (in-time)  
Trigger ID: 0x07  
D wall: 1089.6 cm  
FC e-like, p = 923.2 MeV/c

Charge (pe)

- >15.0
- 13.1-15.0
- 11.4-13.1
- 9.8-11.4
- 8.2-9.8
- 6.9-8.2
- 5.6-6.9
- 4.5-5.6
- 3.5-4.5
- 2.6-3.5
- 1.9-2.6
- 1.2-1.9
- 0.8-1.2
- 0.4-0.8
- 0.1-0.4
- < 0.1



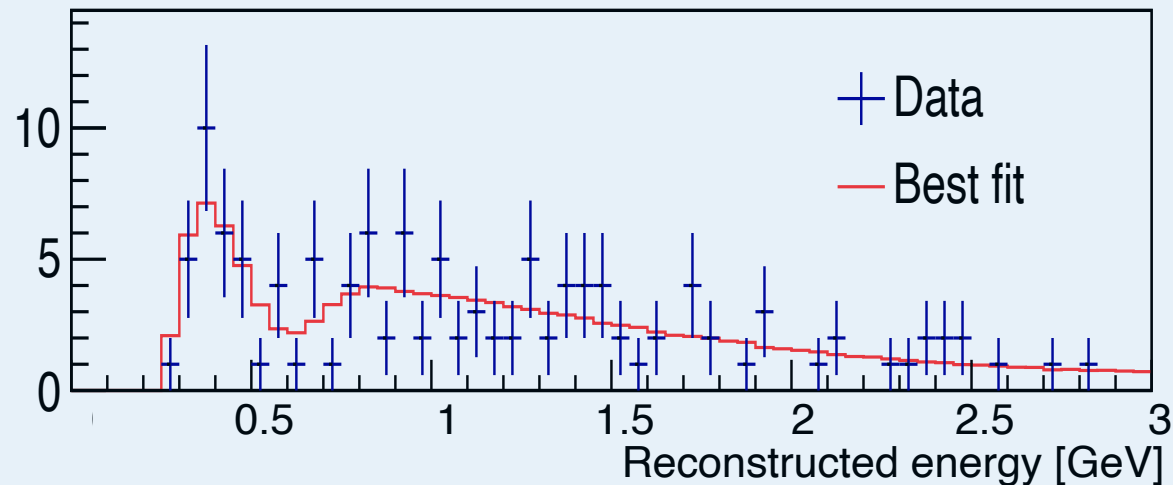
## 4) The ring shape $\propto$ **type of the particle**:

- electron generates electro-magnetic shower ring is diffused
- muon generates a sharp ring

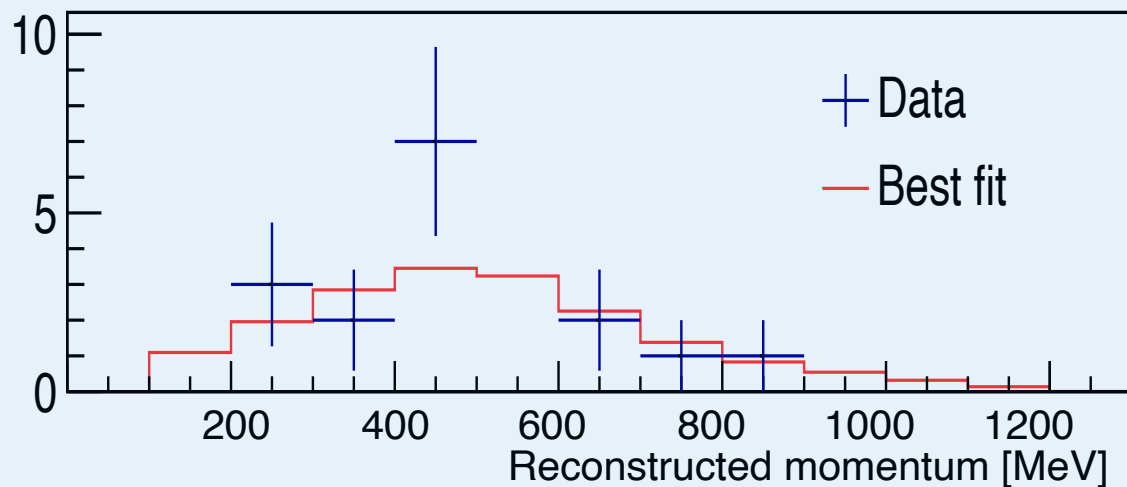
# Oscillation Analysis in T2K

## Antinumode

### Muon-like 1R

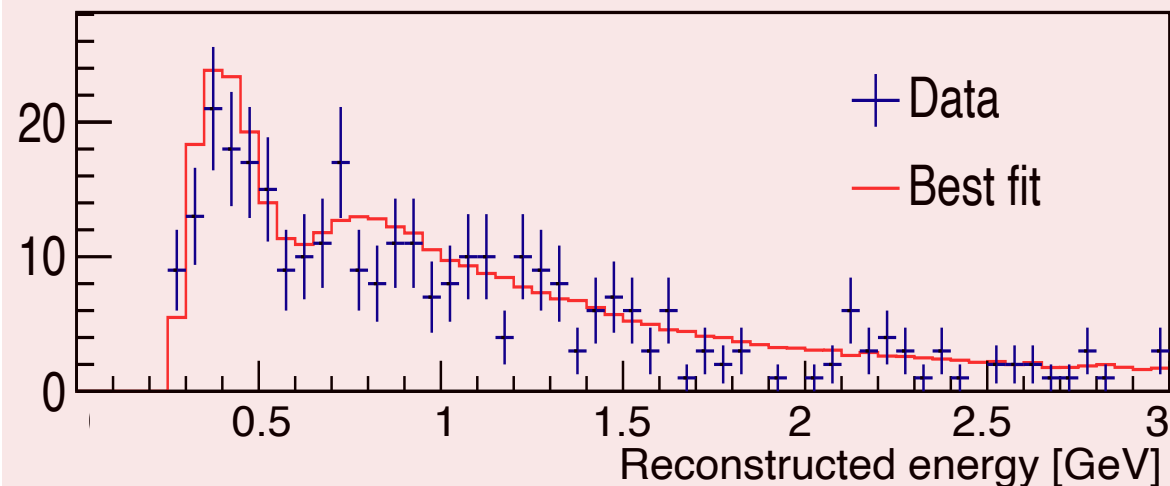


### Electron-like 1R

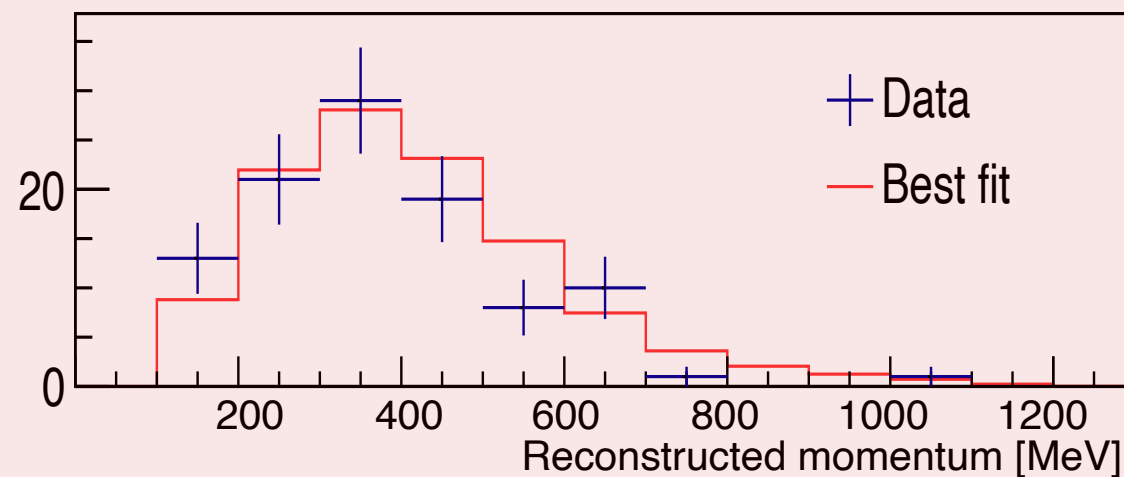


## Numode

### Change namings Muon-like 1R Numu dis



### Electron-like 1R

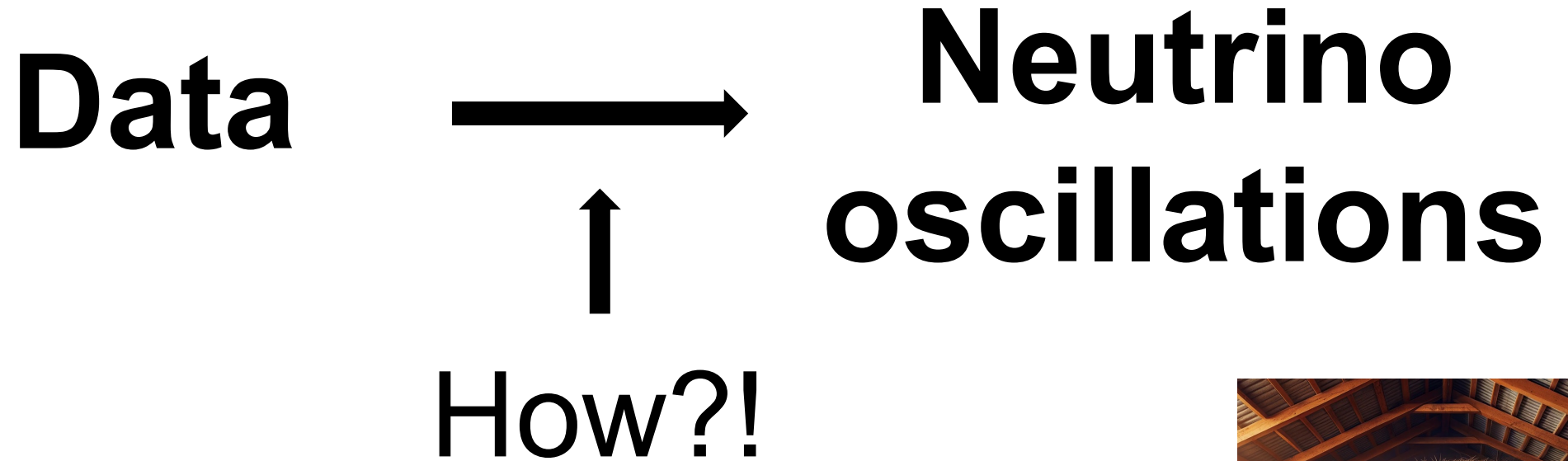


**Multi-ring samples are also used**

Spectra at ND

Comparison unosc/oscilated flux

Dis, app



# T2K Oscillation Analysis



$$N_{obs}^{\nu\alpha}(E_{\nu}^{true}) =$$

$$\Phi(E_{\nu}^{true}) \otimes \sigma(E_{\nu}^{true}) \otimes \epsilon(E_{\nu}^{true}) \otimes S(E_{\nu}^{true}, E_{\nu}^{reco}) \otimes P_{\nu_{\mu} \rightarrow \nu_{\alpha}}(E_{\nu}^{true}, \vec{\theta})$$

Neutrino flux

Interaction cross-section

Detector efficiency

Energy smearing matrix

Oscillation probability

We want to extract this!

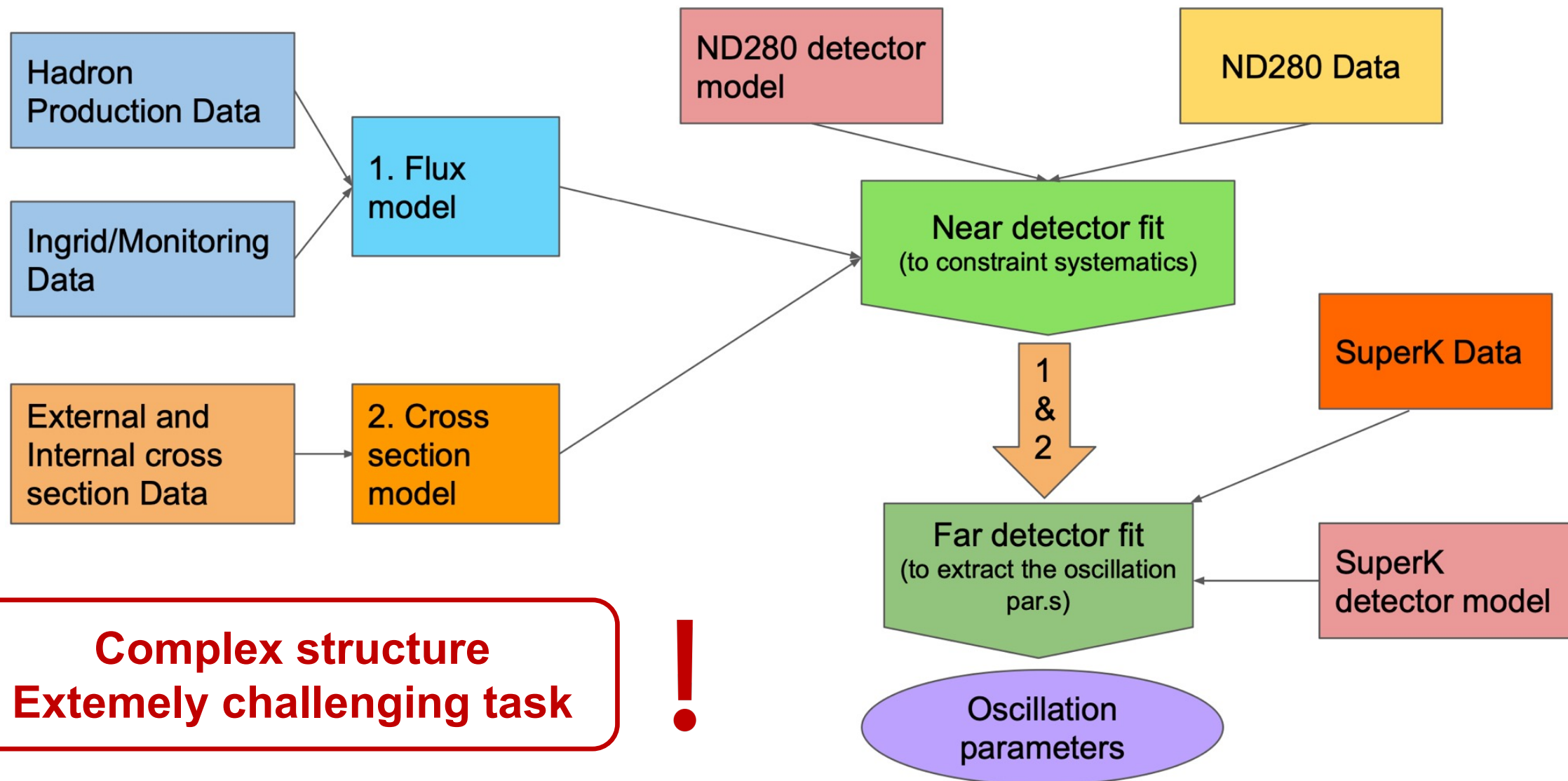
Systematics

$(\vec{f})$

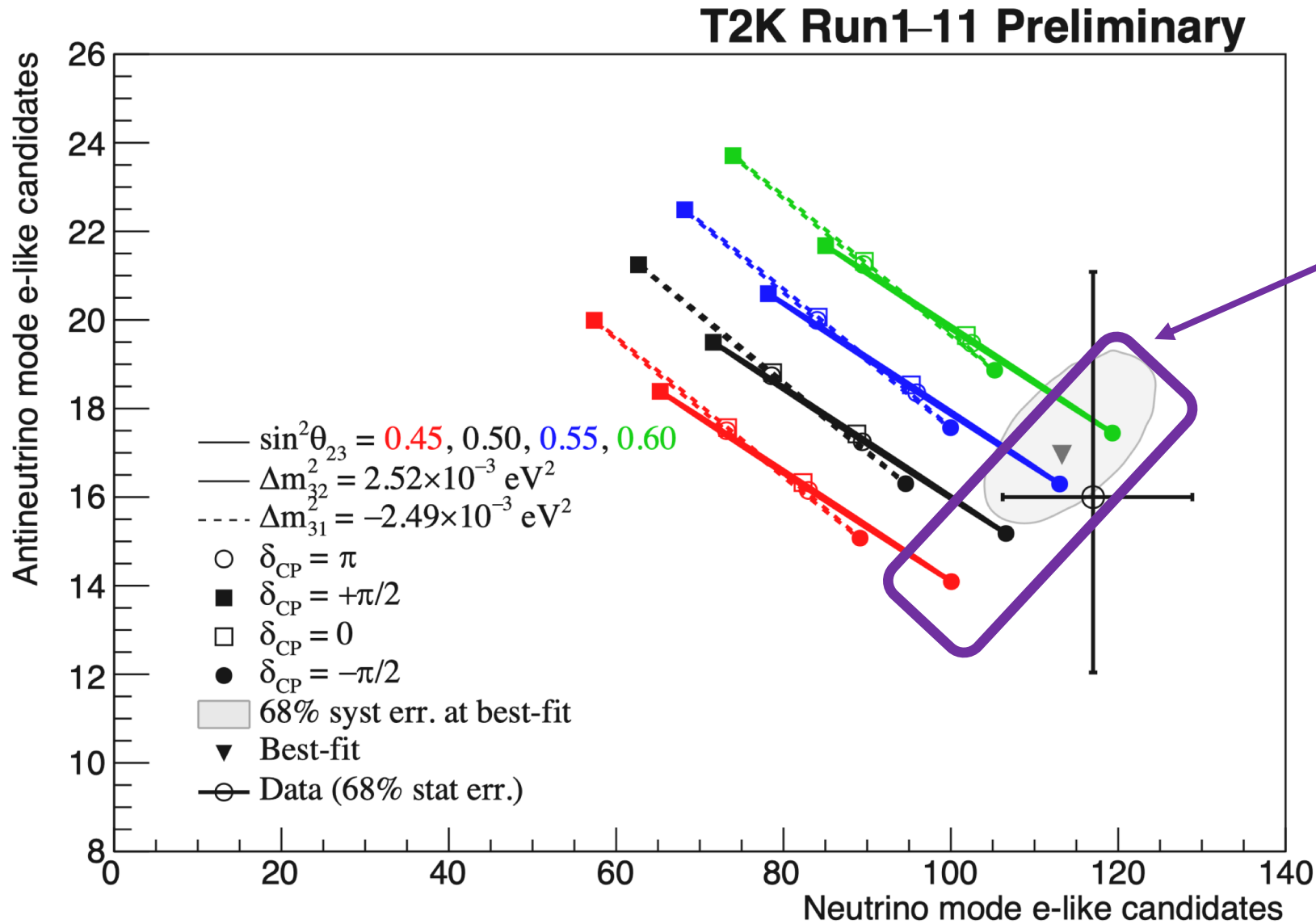
$\alpha = e, \mu$



# T2K Oscillation Analysis



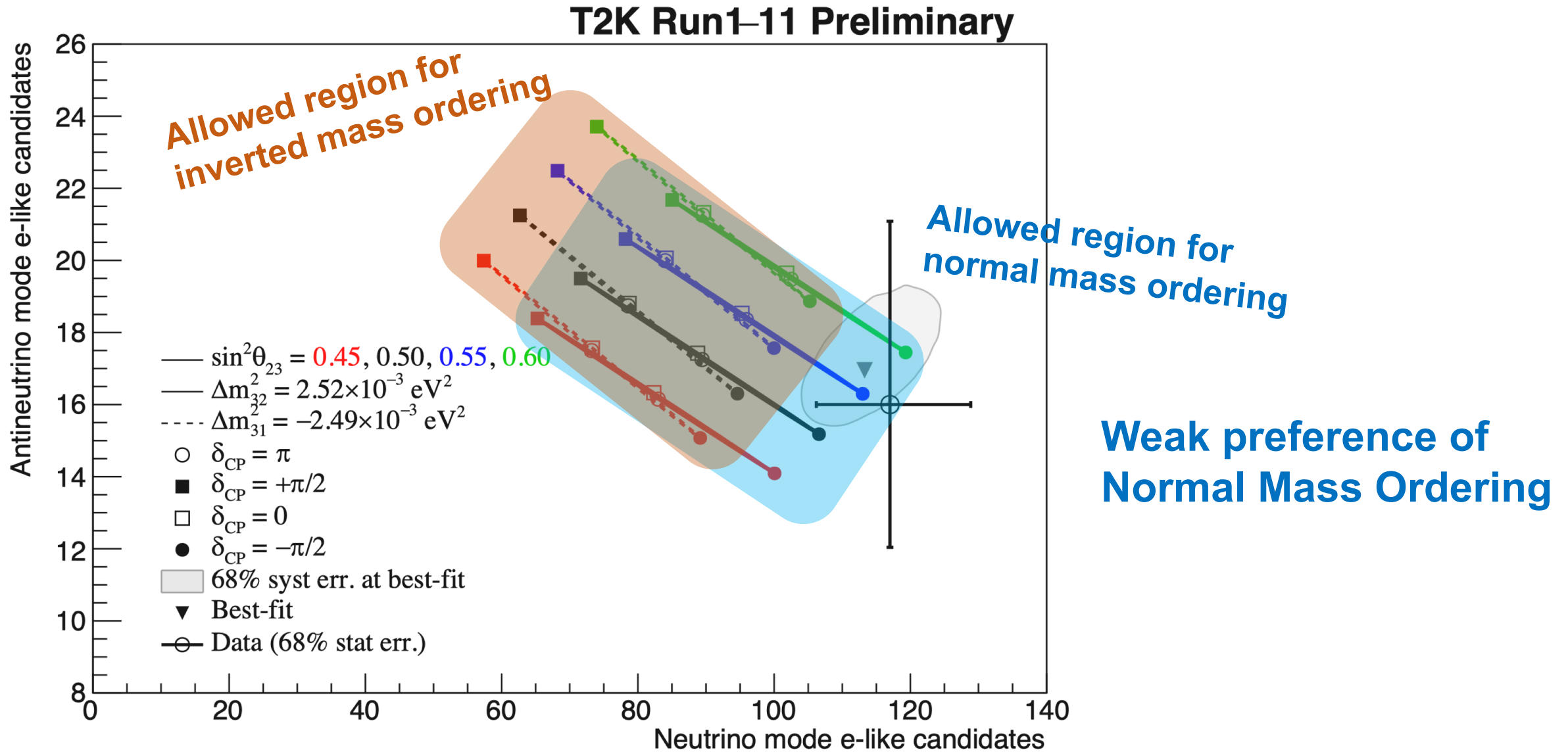
# T2K Oscillation Analysis



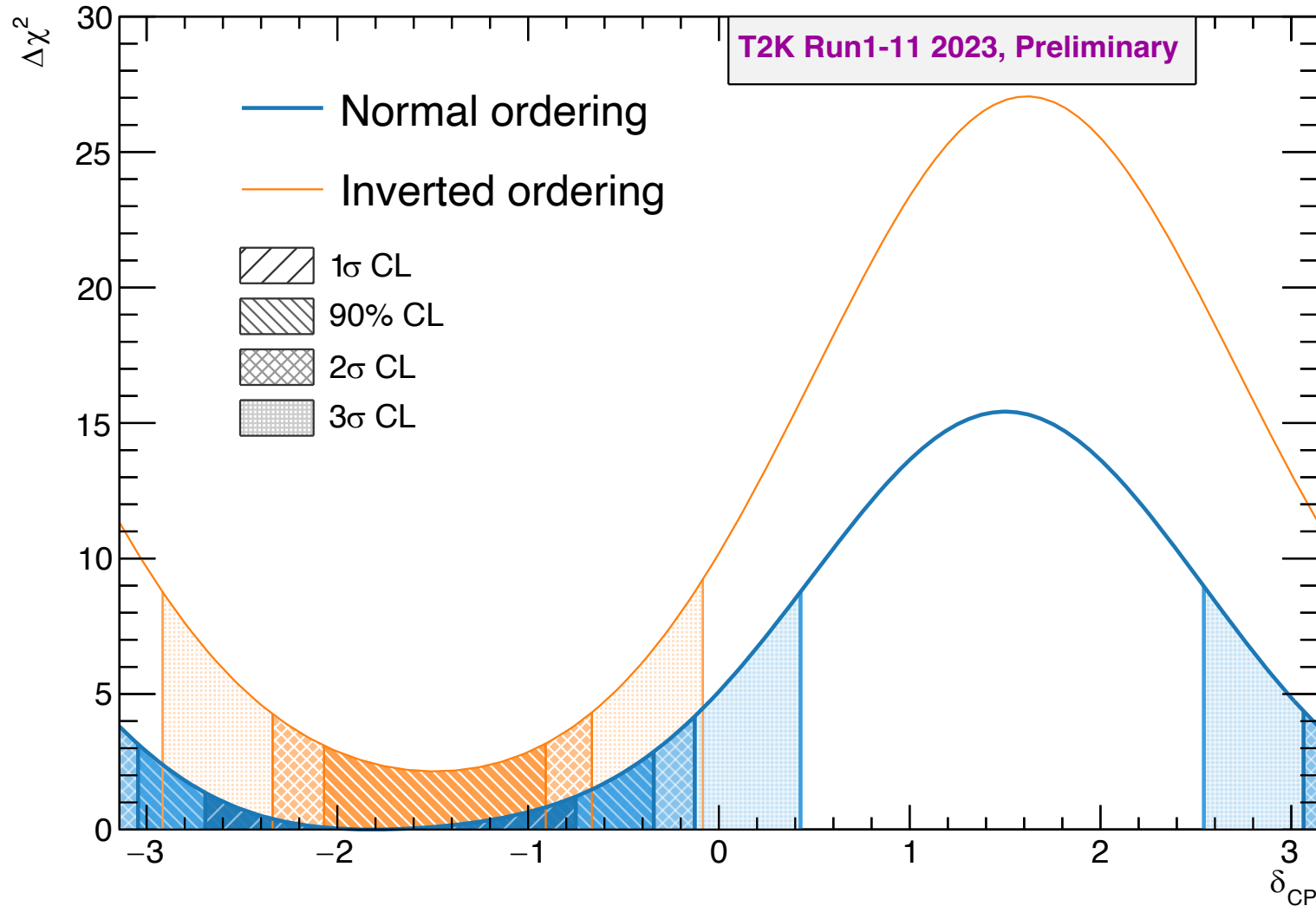
Max. CP-violation region  
 Best fit:  $\delta_{CP} \approx -\frac{\pi}{2}$

Weak preference of  
 Normal Mass Ordering

# T2K Oscillation Analysis







## Results:

$$\delta_{CP} = -2.08^{+1.33}_{-0.61}$$

CP cons. excluded at 90% level

Exc.  $\delta_{CP} = \pm\pi$  with 90% C.L.

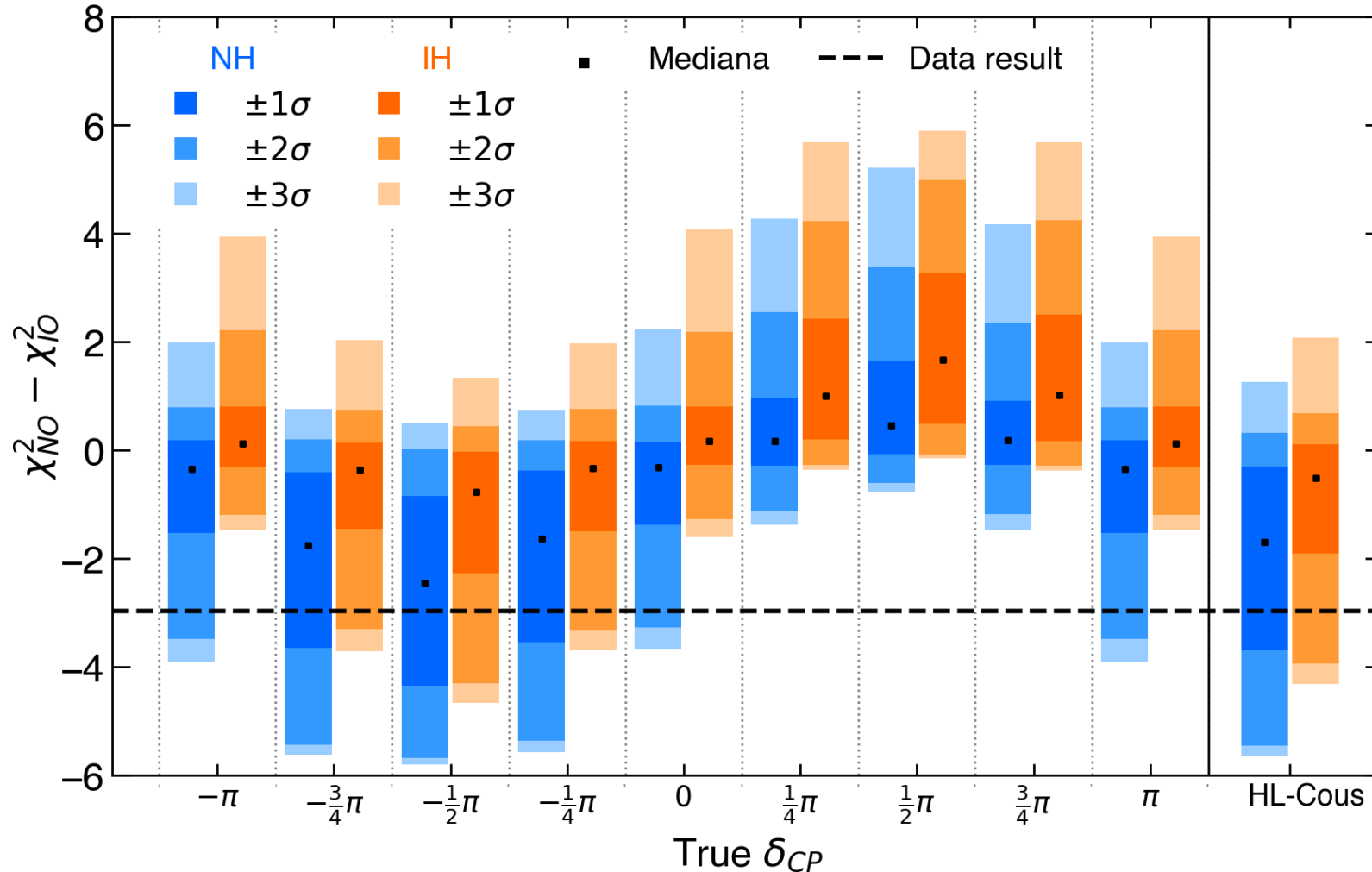
Exc.  $\delta_{CP} = 0$  with 2σ C.L.

Normal ordering preferred

At which level?  
See next slide

## Mass ordering results

T2K Run1-11 Preliminary

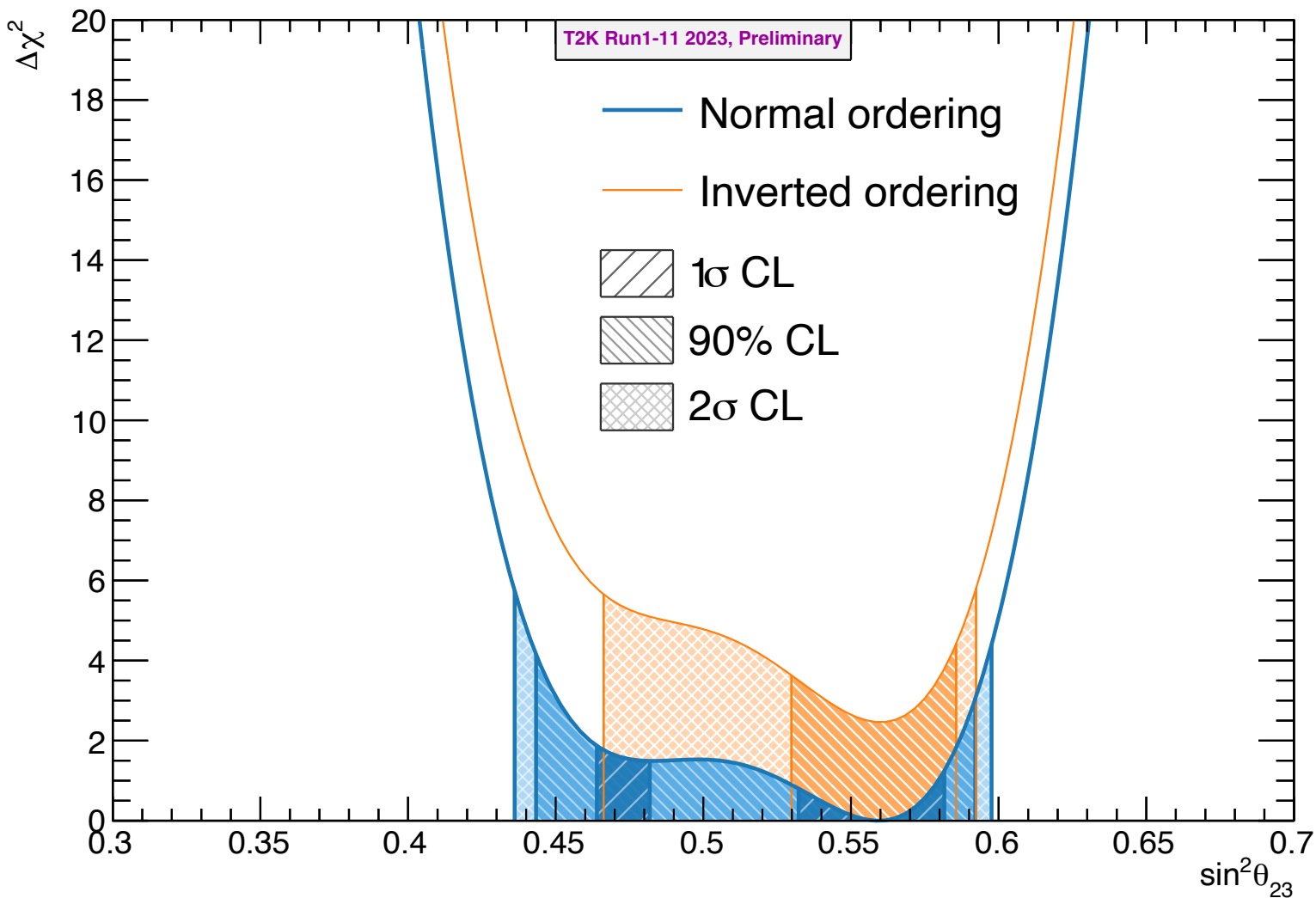


**Inverted Ordering  
exclusion**

p-value =  $1.69\sigma$   
 CLs =  $1.27\sigma$

$$CLs = \frac{p_{\text{value}}(\text{IH})}{1 - p_{\text{value}}(\text{NH})}$$

# T2K recent results



Confidence level	Interval (NH)
$1\sigma$	$[0.464, 0.482] \cup [0.532, 0.582]$
90%	$[0.443, 0.592]$
$2\sigma$	$[0.436, 0.598]$

Confidence level	Interval (IH)
$1\sigma$	$[0.530, 0.586]$
90%	$[0.530, 0.586]$
$2\sigma$	$[0.466, 0.592]$

$\epsilon(\sin^2 \theta_{23}) = 9\%$  (assuming upper octant)

Week preference of upper octant

**T2K experiment**



after 2027

**HK experiment**



# HyperK has an extensive physics program

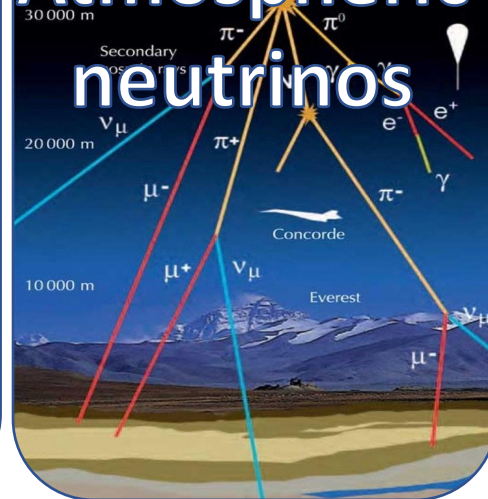
## Solar neutrinos

- MSW effect in the sun
- NSI in the Sun

## Supernova neutrinos

- Direct SNv :
- Relic SNv :

## Atmospheric neutrinos



## Accelerator neutrinos



## Proton decay



# HyperK has an extensive physics program

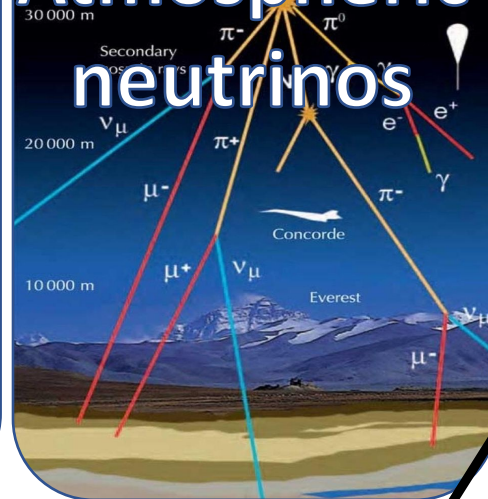
## Solar neutrinos

- MSW effect in the sun
- NSI in the Sun

## Supernova neutrinos

- Direct SNv :
- Relic SNv :

## Atmospheric neutrinos



## Proton decay



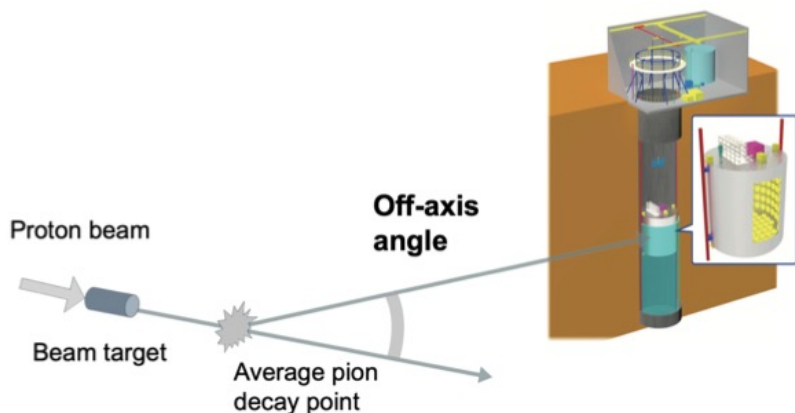
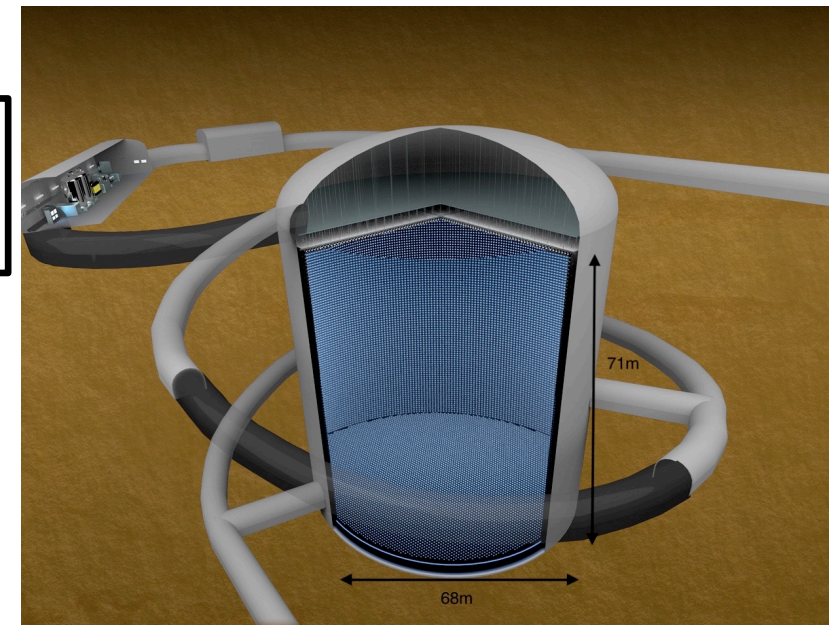
Focusing on that

# Hyper-K experiment

Next generation of long-baseline experiments

- New far detector (x8 of fiducial mass)
- Increase of beam power (500kW -> 1300kW)
- Additional near detector - Intermediate Water Cherenkov Detector (IWCD)
- **ND280 Upgrade - Done**

**x21 in stat!**



	Super-Kamiokande	Hyper-Kamiokande
Site	Mozumi	Tochibora
Number of ID PMTs	<b>11,129</b>	<b>20,000</b>
Photo-coverage	40%	20% ( <b>x2 sensitivity</b> )
Mass/Fiducial Mass	50 kton / <b>22.5 kton</b>	260 kton / <b>187 kton</b>

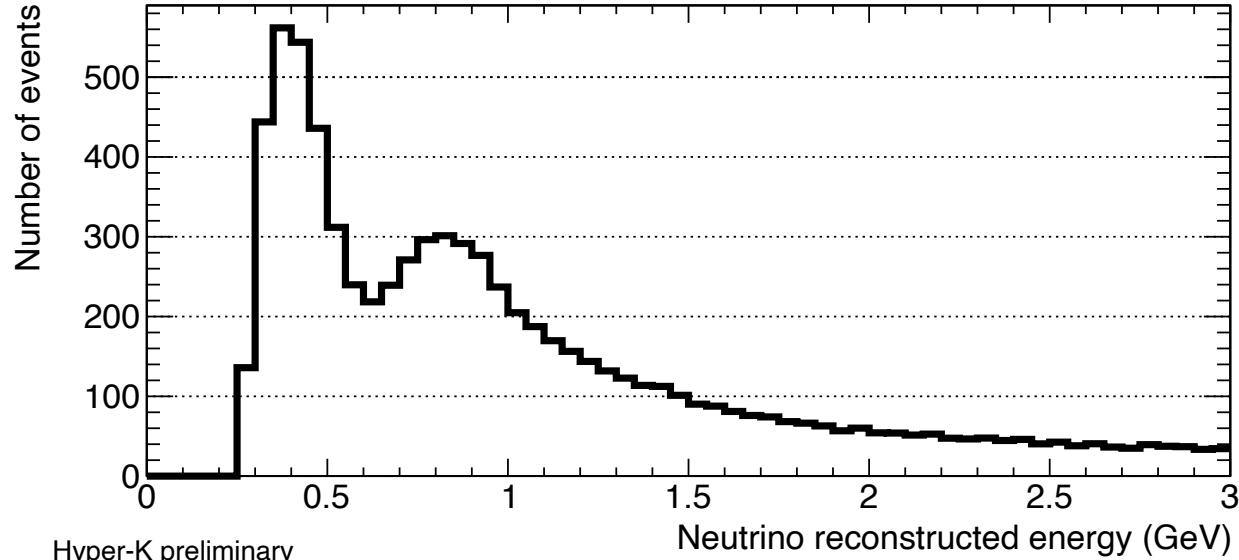
# Hyper-Kamiokande sensitivity studies

# HK sensitivity studies



After 10 HK years

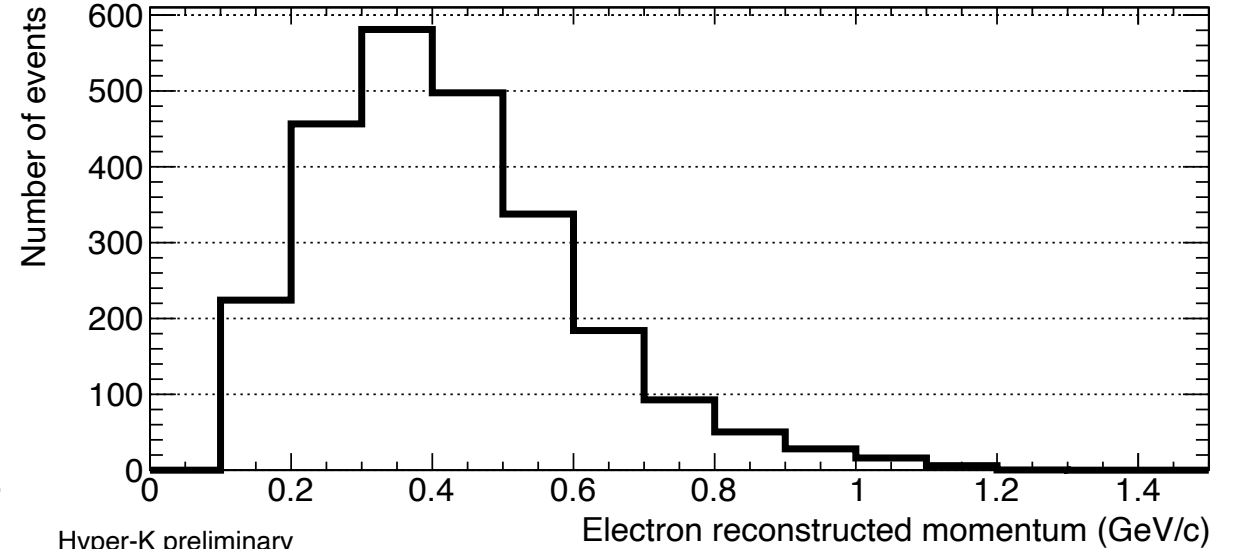
Far Detector,  $\nu$  mode, 1-ring  $\mu$ -like



Hyper-K preliminary  
10 years ( $2.7 \times 10^{22}$  POT 1:3  $\nu\bar{\nu}$ )  
True Normal Ordering,  $\sin^2\theta_{13}=0.0218$ ,  $\sin^2\theta_{23}=0.528$ ,  $\Delta m_{32}^2=2.509 \times 10^{-3} \text{eV}^2/c^4$ ,  $\delta_{CP}=-1.601$

~8800 events

Far Detector,  $\nu$  mode, 1-ring e-like + 0 decay e



Hyper-K preliminary  
10 years ( $2.7 \times 10^{22}$  POT 1:3  $\nu\bar{\nu}$ )  
True Normal Ordering,  $\sin^2\theta_{13}=0.0218$ ,  $\sin^2\theta_{23}=0.528$ ,  $\Delta m_{32}^2=2.509 \times 10^{-3} \text{eV}^2/c^4$ ,  $\delta_{CP}=-1.601$

~2400 events

Total number of events in all **samples** more than 25000!

For comparison: T2K has recorded 600 events by 2024

**Sensitivity studies are performed for three syst. models:**

- Statistics only (zero systematic uncertainties)
- T2K 2020 systematics model (where we are now)
- Improved systematics

The Improved systematics model was produced by scaling the T2K-2020 error model by

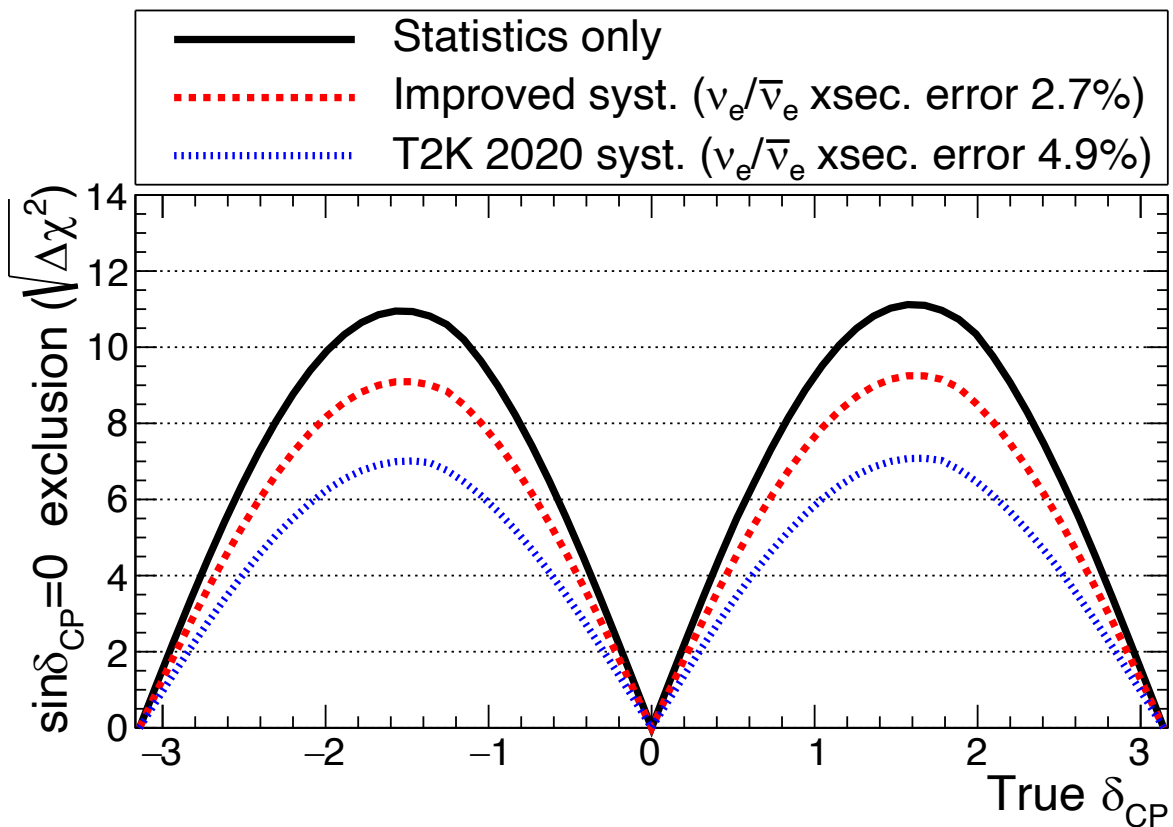
- scaling the uncertainties by  $\frac{1}{\sqrt{N}}$ , where  $N = 7.5$  is the relative increase in neutrino beam exposure from T2K to Hyper-K
- putting additional constraints to the cross-section model uncertainties (coming from higher expected performance of upgraded ND280 and IWCD)







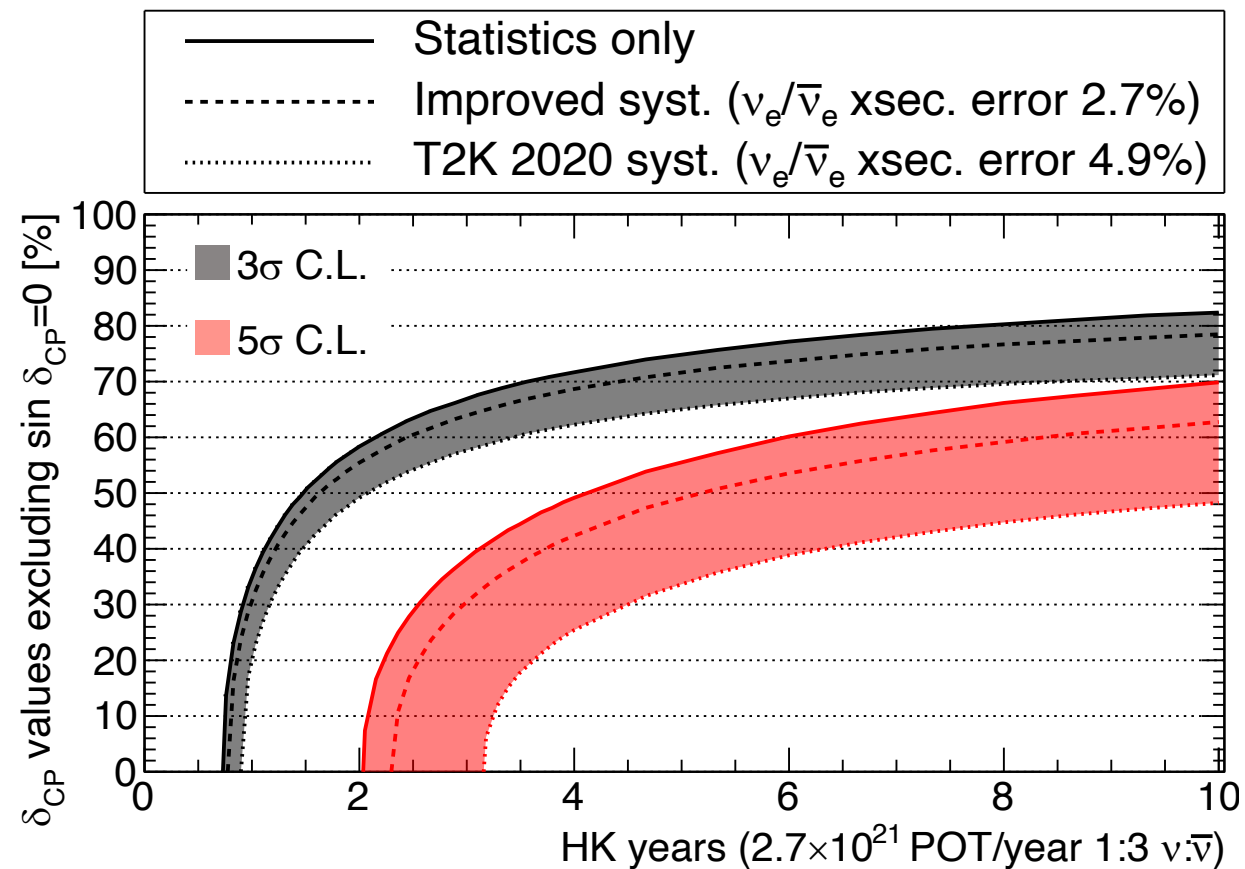
# HK sensitivity to $\delta_{CP}$



Hyper-K preliminary

True normal ordering (known), 10 years ( $2.7 \times 10^{22}$  POT 1:3  $\nu:\bar{\nu}$ )

$\sin^2\theta_{13}=0.0218\pm 0.0007$ ,  $\sin^2\theta_{23}=0.528$ ,  $\Delta m_{32}^2=2.509\times 10^{-3}\text{eV}^2/c^4$



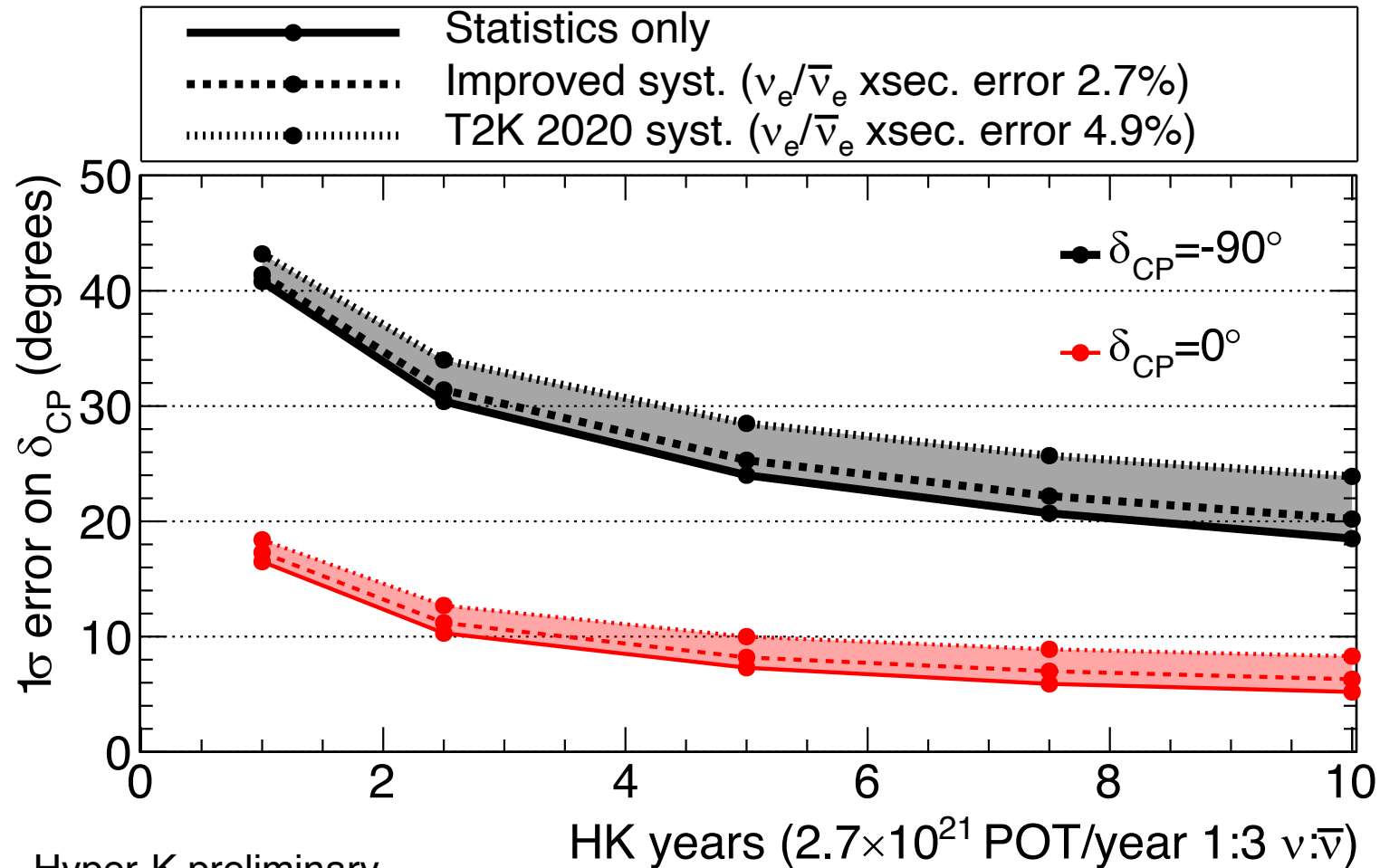
Hyper-K preliminary

True normal ordering (known)

$\sin^2\theta_{13}=0.0218\pm 0.0007$ ,  $\sin^2\theta_{23}=0.528$ ,  $\Delta m_{32}^2=2.509\times 10^{-3}\text{eV}^2/c^4$

**Significant impact from systematics**

# HK sensitivity to $\delta_{CP}$



$1\sigma$  error on  $\delta_{CP}$

For Improved syst. model:  $20.2^\circ$   
 For T2K 2020 syst. model:  $23.9^\circ$

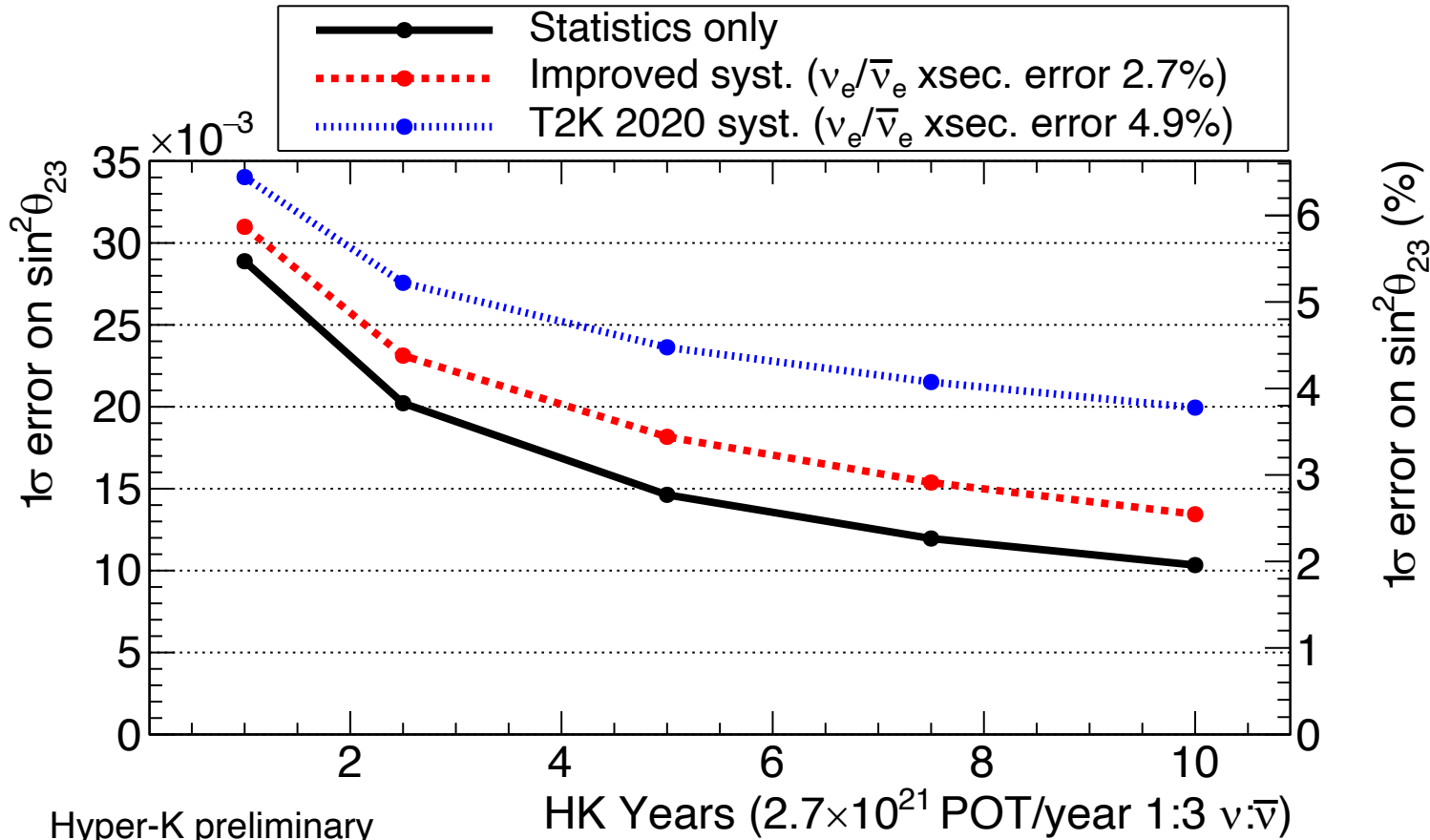
Hyper-K preliminary

True normal ordering (known)

$$\sin^2\theta_{13}=0.0218\pm 0.0007, \sin^2\theta_{23}=0.528, \Delta m_{32}^2=2.509\times 10^{-3}\text{eV}^2/c^4$$

For comparison: T2K  $\delta_{CP}$  resolution is around  $70^\circ$

# HK sensitivity to $\sin^2 \theta_{23}$



## Relative $1\sigma$ resolution:

For Improved syst. model:  $\varepsilon = 2.5\%$   
 For T2K 2020 syst. model:  $\varepsilon = 3.8\%$

Hyper-K preliminary

True normal ordering (known)

$\sin^2 \theta_{13} = 0.0218 \pm 0.0007$ ,  $\sin^2 \theta_{23} = 0.528$ ,  $\Delta m_{32}^2 = 2.509 \times 10^{-3} \text{ eV}^2/c^4$ ,  $\delta_{CP} = -1.601$

**For comparison: T2K  $\delta_{CP}$  resolution is around 9%**

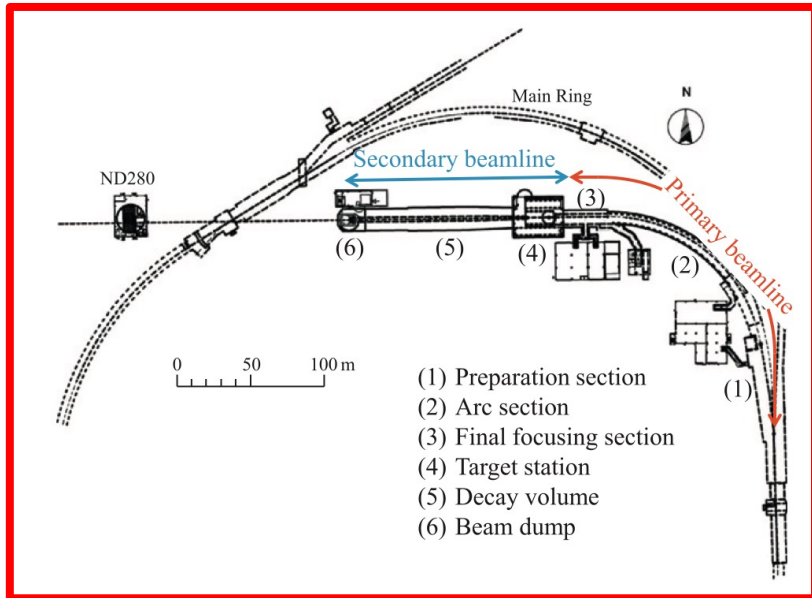
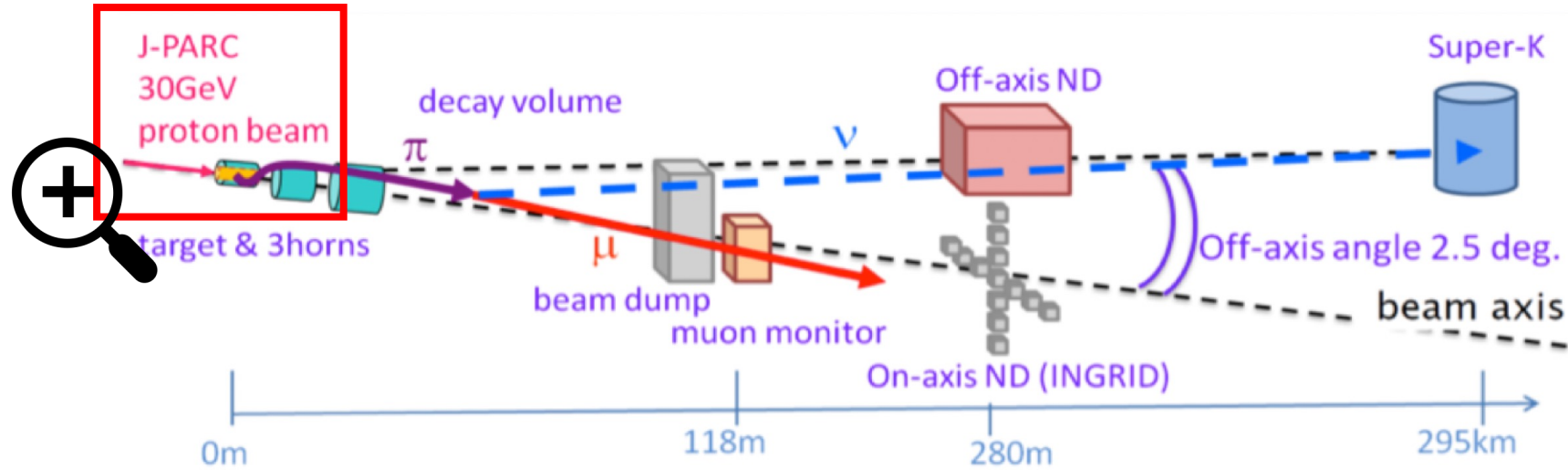
- Neutrino is a way to find out the hints of fundamental physics and oscillations help to study the neutrino
- Oscillation measurement faces many challenges from experimental and theoretical point nevertheless T2K could achieve fascinating results
- Hyper-K being very ambitious project it comes with its own set of challenges (neutrino interaction and flux modelling, significant hardware challenges with various complex detectors, etc).

**But it is very exciting as HK era promises to be a ground breaking period in neutrino research.**



**BACKUP**

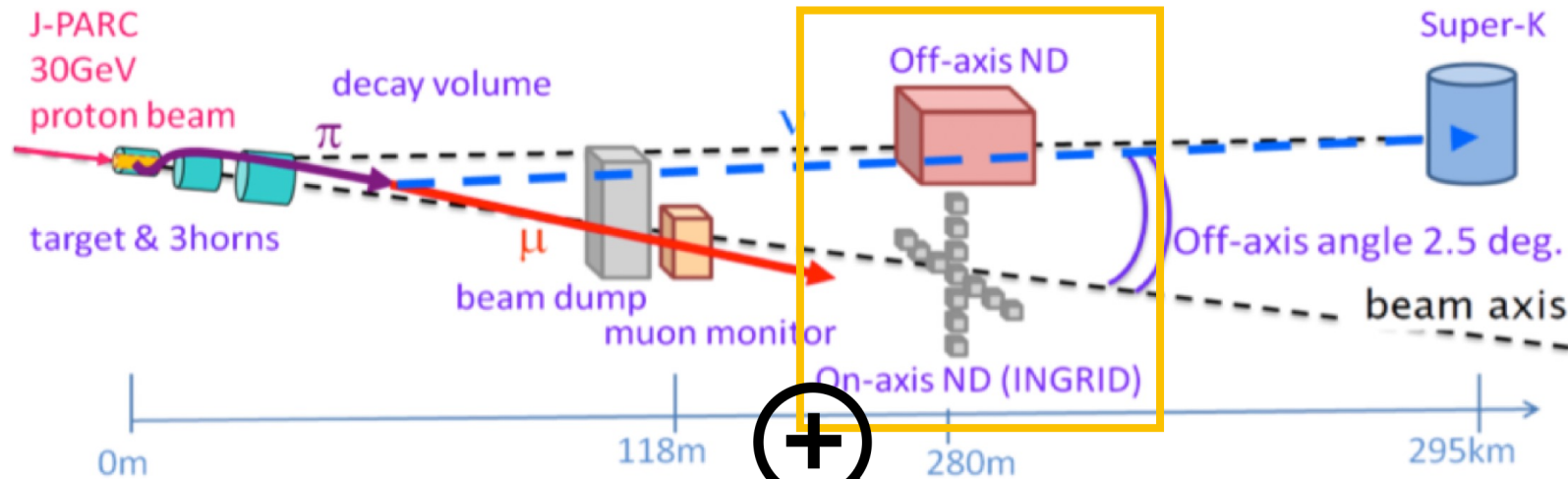
# T2K experiment



- 30 GeV proton accelerator
- Carbon target for pions production
- 3 magnetic horns for pion focusing (positive or negative)



# T2K experiment



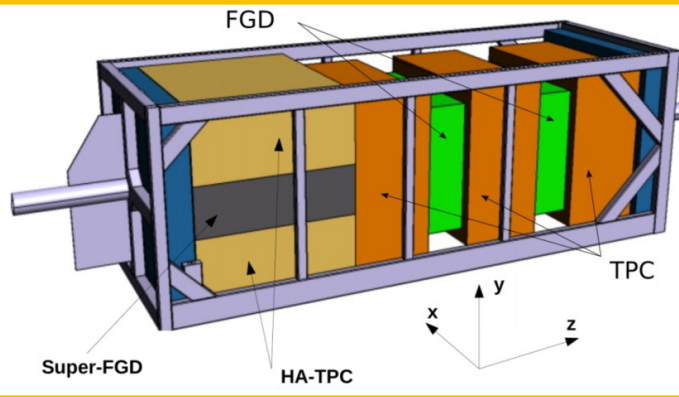
## ND280Upgrade

## INGRID

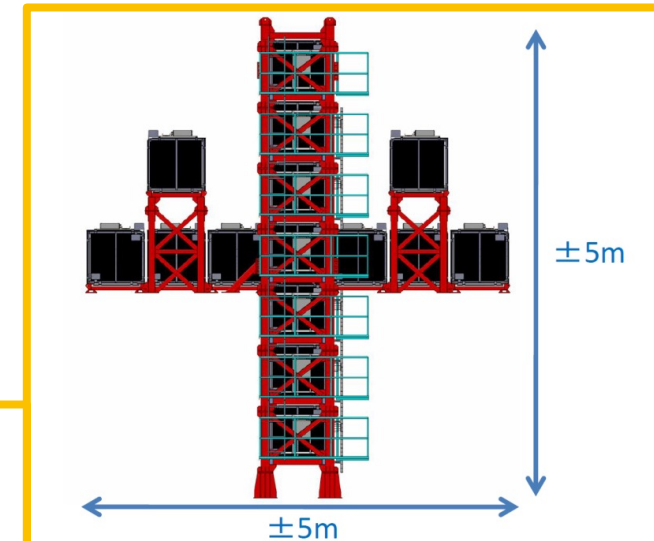
- Off-axis detector
- Consists of multiple subdetectors
- Constrains neutrino flux and cross-section interaction

### Have been recently upgraded providing:

- Lower down proton reconstruction threshold
- Neutron tagging.
- 4pi coverage of particles tracking

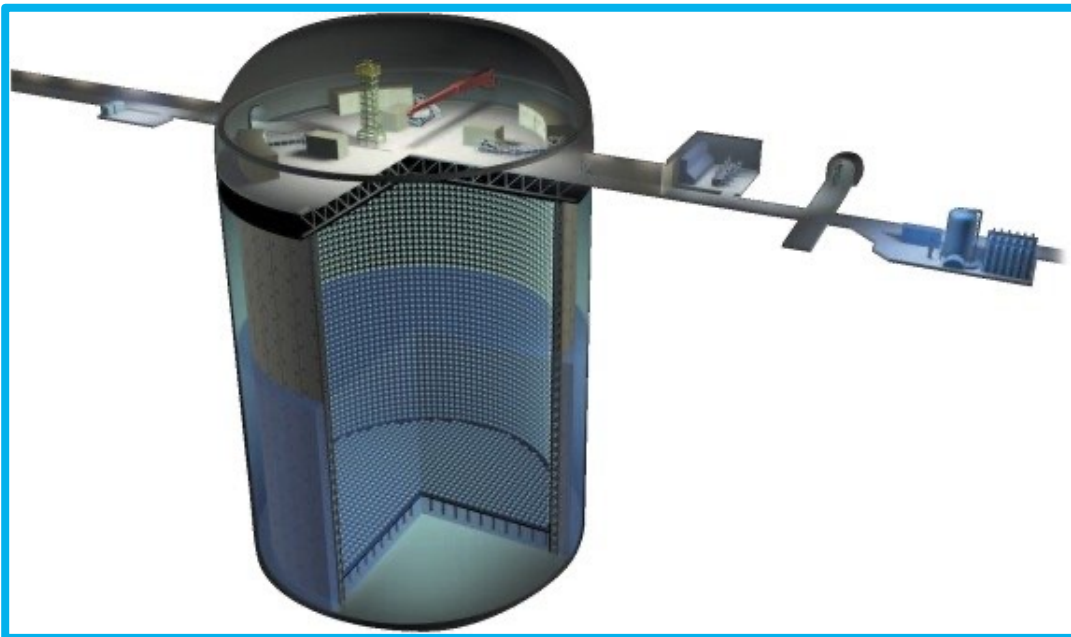
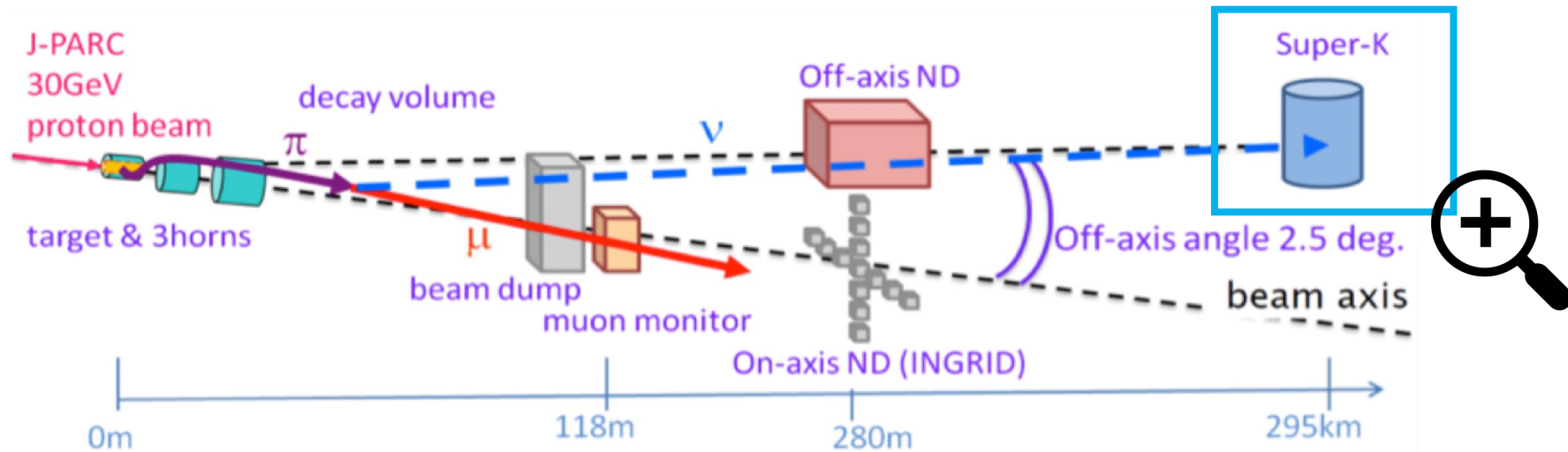


- Monitors the neutrino beam direction and intensity



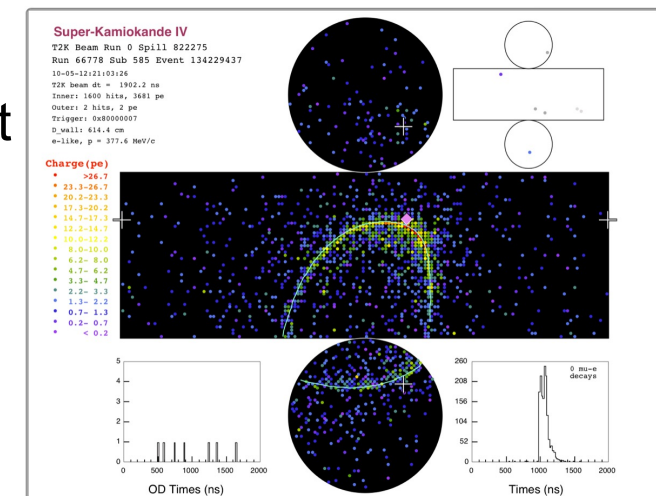


# T2K experiment

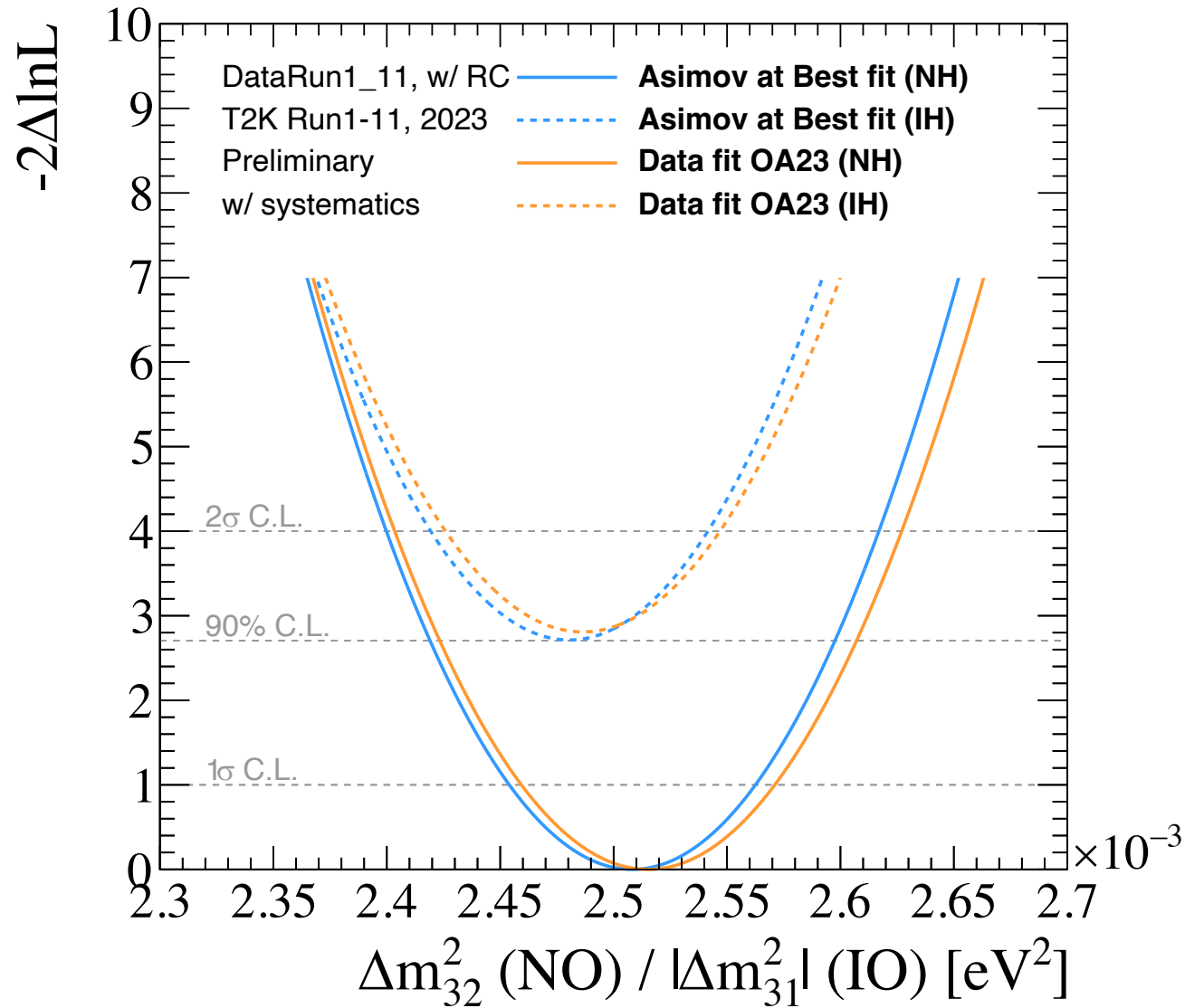


- Fiducial mass = 27.2 kton
- Detects water Cherenkov light from charged particle and reconstructs events with PMT charge & timing information.

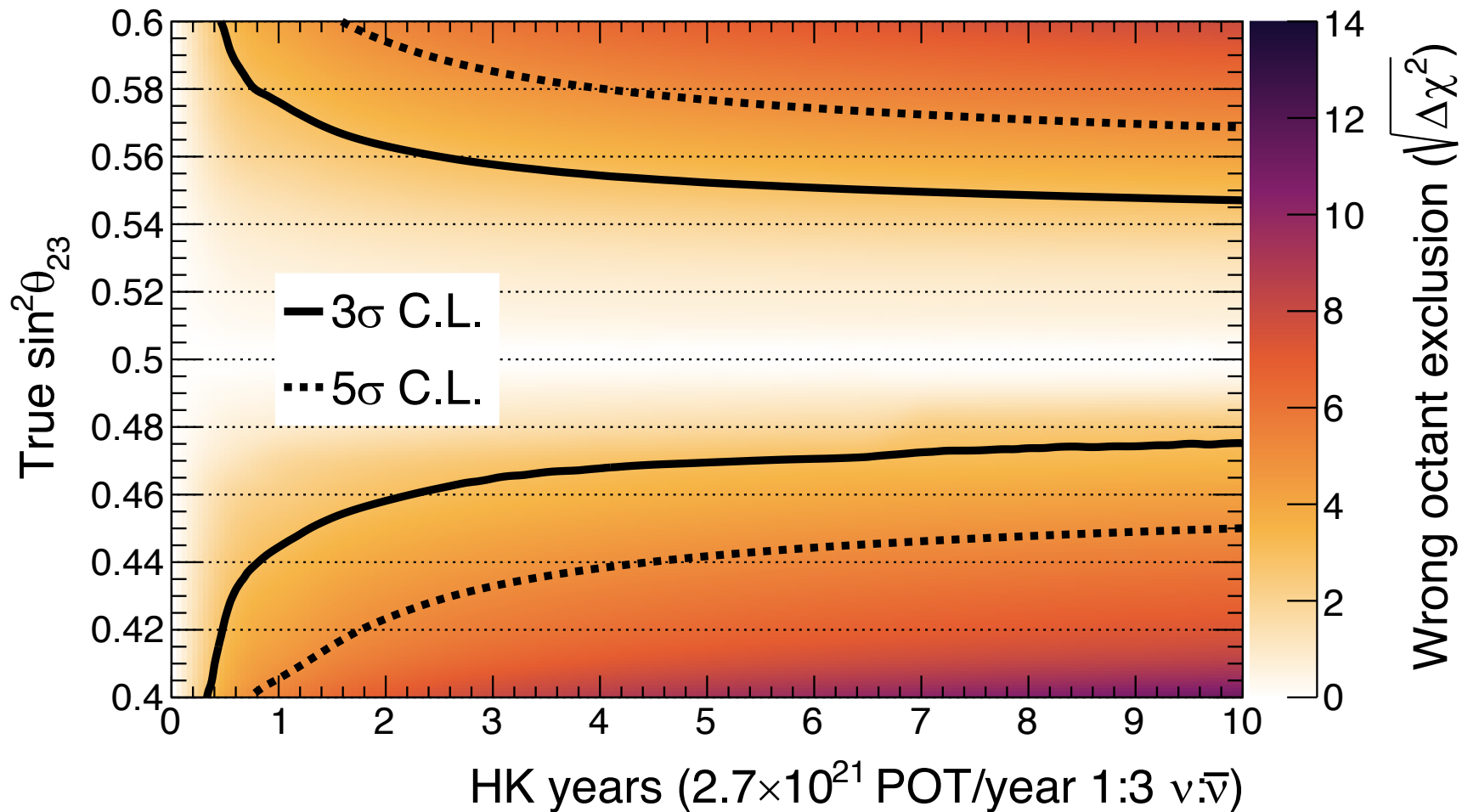
## Example of data event



# T2K recent results

 $\Delta m_{32}^2$ 

# HK sensitivity to $\sin^2 \theta_{23}$ octant



Hyper-K preliminary

True normal ordering (known), Improved systematics

$\sin^2 \theta_{13} = 0.0218 \pm 0.0007$ ,  $\delta_{CP} = -1.601$ ,  $\Delta m_{32}^2 = 2.509 \times 10^{-3} \text{ eV}^2/c^4$