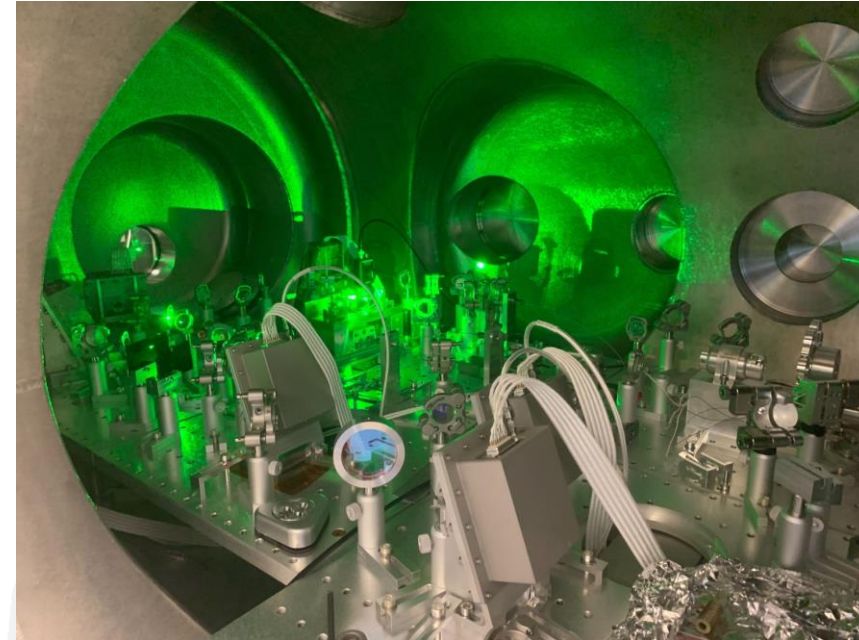
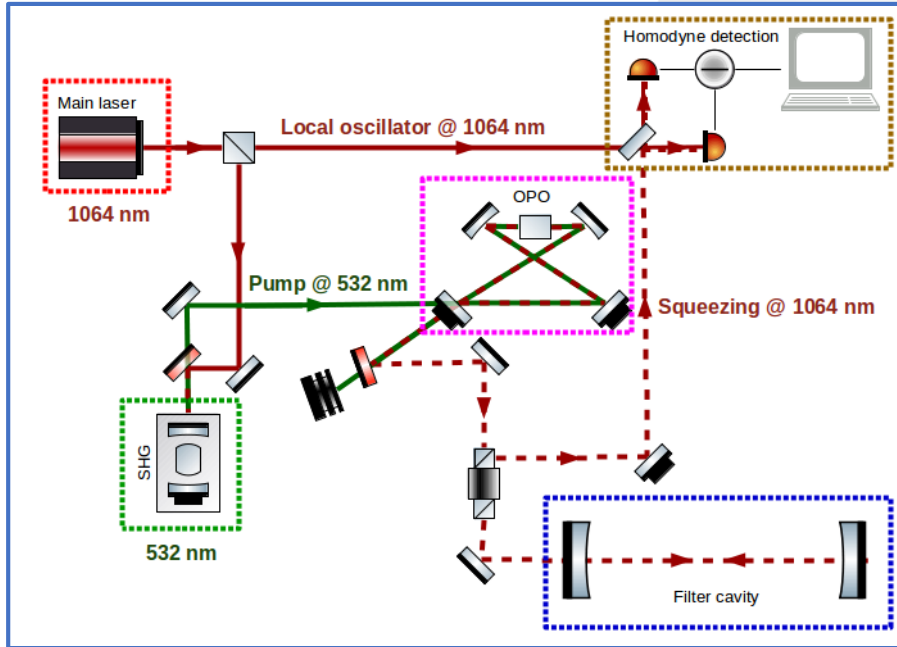


R&D at IJCLab



A. van de Walle, M. Andia, F. Glotin, A. Lartaux, N. Leroy, V. Loriette, P. Stevens

ANR EXSQUEEZ (2015-2019)
ANR QFILTER (2019-2024)

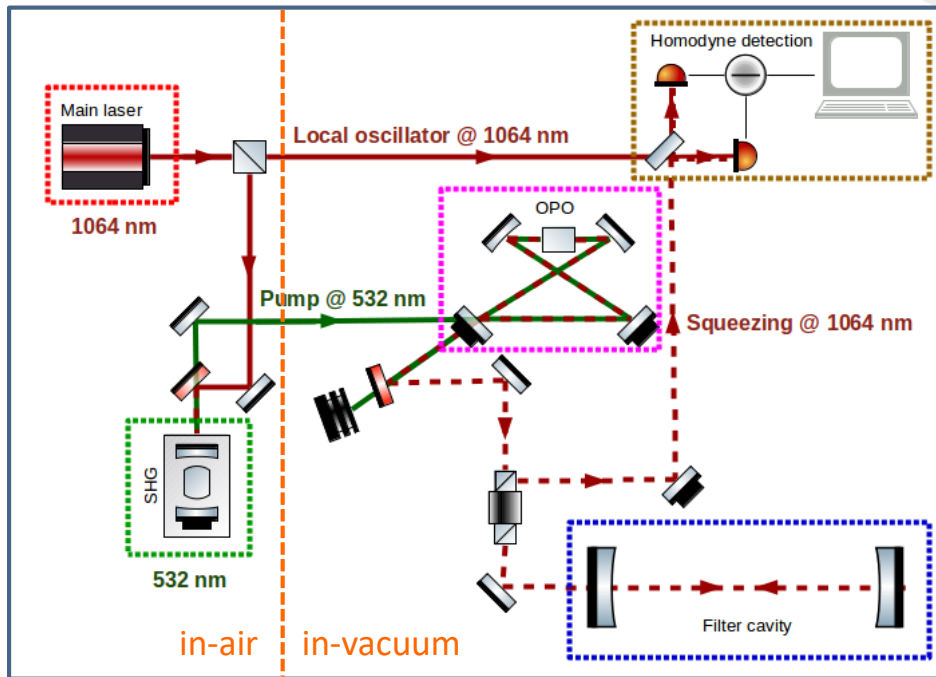
Goal :

- Demonstrate under vacuum frequency dependent squeezing (FDS) in the prospect of the Advanced Virgo detector O5 run and ET detector development => **improve GW detector sensitivity** in its whole bandwidth
- Adjust the squeezing angle corner frequency => **more flexibility**



- Squeezing source from LKB
- Existing 50-m filter cavity on the CALVA facility at IJCLab
- Optics coatings from IP2I/LMA
- Electronics from LAPP in the Virgo standards

- Demonstrate the impact of placing the squeezing under vacuum in the prospect of the Advanced Virgo detector O5 run and ET detector development.

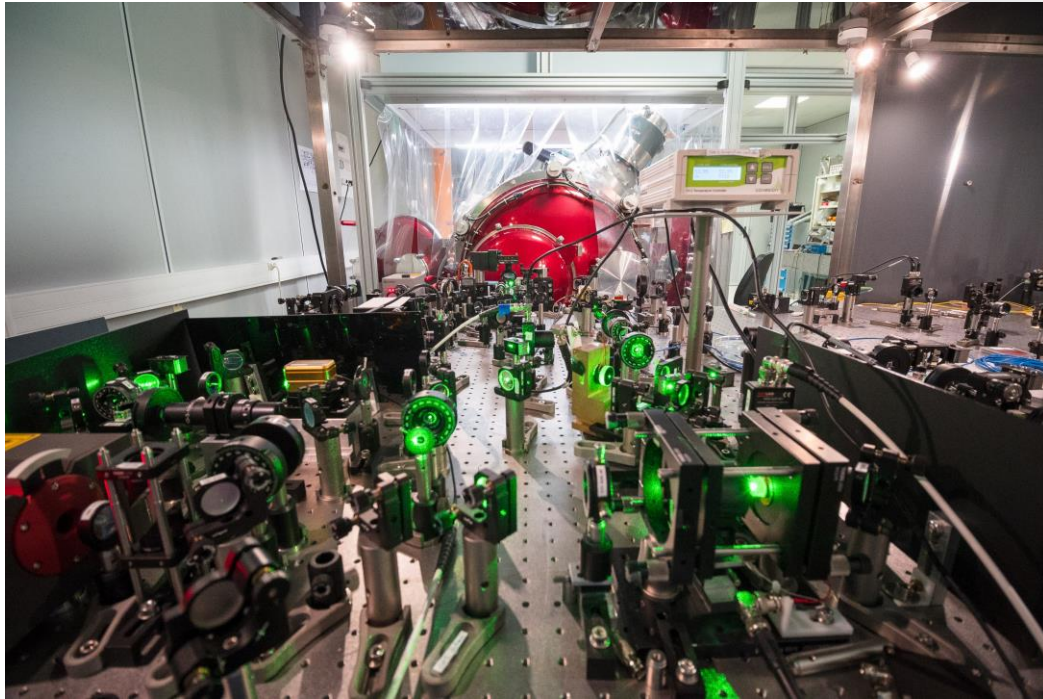


Simplified scheme

Others main interest points :

- OPO bow-tie type like LIGO
 - Impact on the stray light
- Bench under vacuum but not suspended
- SHG done in air and send under vacuum by a window and not a fiber.
- Modified Coherent Locking (MCL) scheme :
 - Using an frequency shifted infrared beam that will share a phase relation with the squeezing to lock the squeezing angle
- A 50-m long suspended cavity

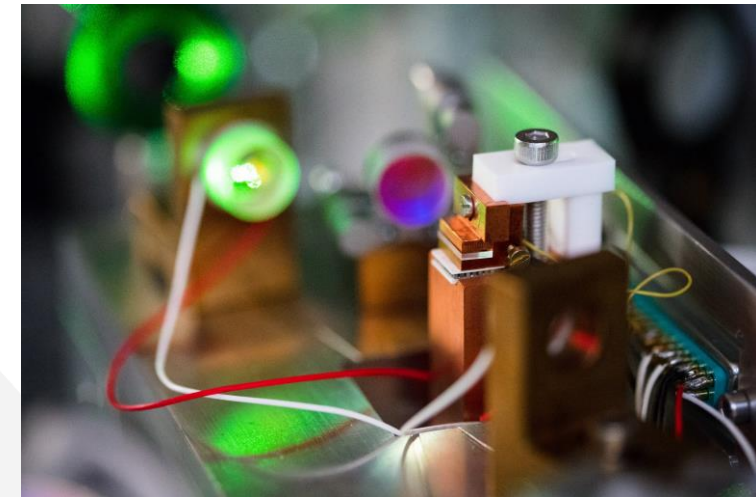
- Demonstrate the impact of placing the squeezing under vacuum in the prospect of the Advanced Virgo detector O5 run and ET detector development.



In-air bench experiment setup

First results :

- Experimental setup has been implemented
- OPO characterized
- Anti-squeezing observed

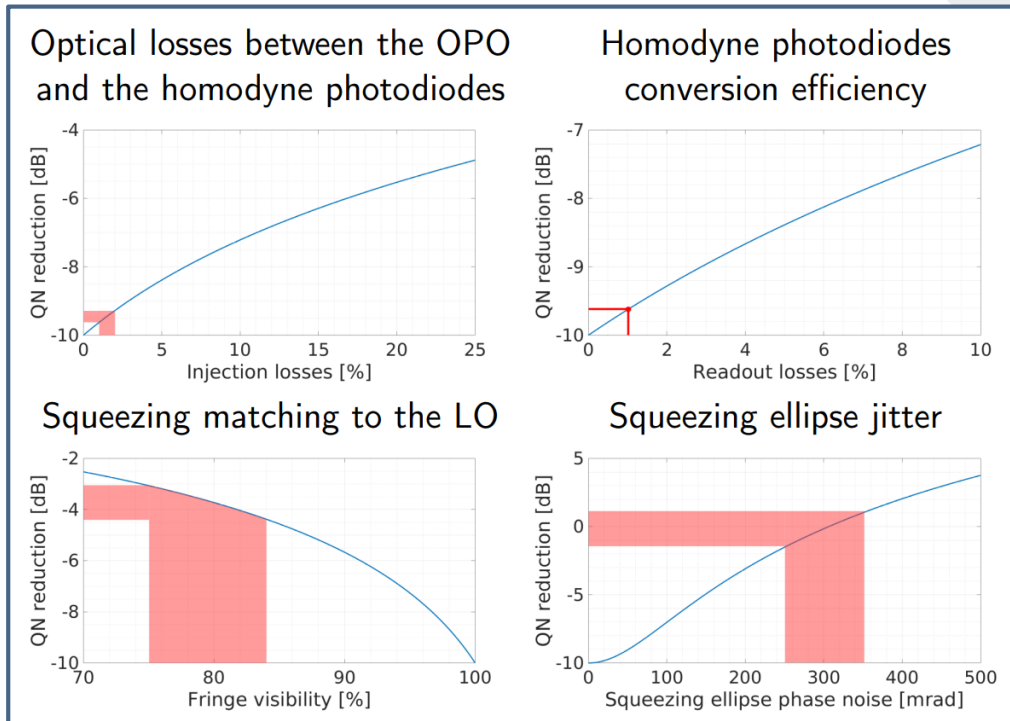


Bow-tie OPO

- Demonstrate the impact of placing the squeezing under vacuum in the prospect of the Advanced Virgo detector O5 run and ET detector development.

No squeezing observed due to mainly 2 limitations :

- Matching of the LO and squeezing on homodyne detection
- Squeezing ellipse angle control

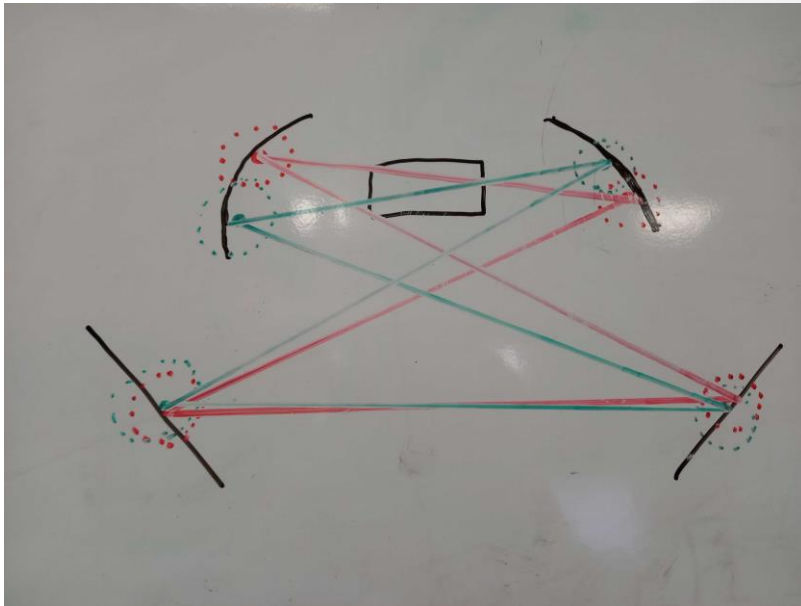


Squeezing loss sources

- Demonstrate the impact of placing the squeezing under vacuum in the prospect of the Advanced Virgo detector O5 run and ET detector development.

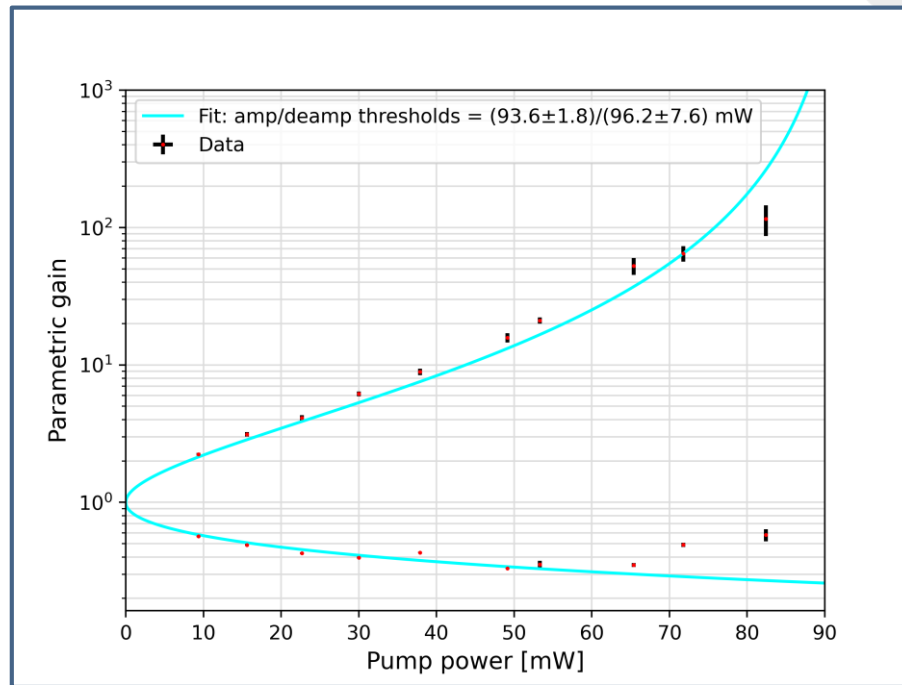
Some problems during the last years :

- Climatization failure :
 - Summer 2021 : Experiment that unaligned itself in the span of few hours
 - Full alignment of the experiment needed at the end of the problem
- Main laser failure :
 - May to September 2022 : internal fan failure
 - Back in January 2023 with Gaussian parameter changed which entails experiment adaptation
- Unexpected OPO alignment difficulty :
 - A mirror unglued itself : Full realignment after change
 - Strange chromatic alignment dependence : Infrared and Green both resonant but not superposed



Scheme of OPO with
not superposed but resonant beam

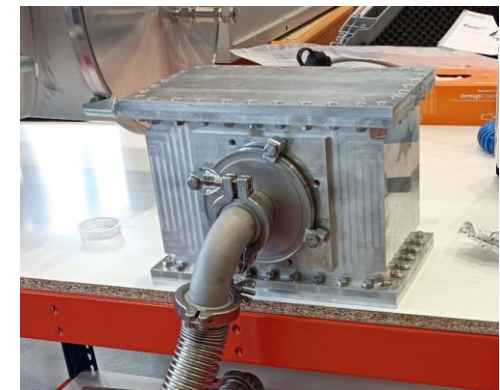
- Demonstrate the impact of placing the squeezing under vacuum in the prospect of the Advanced Virgo detector O5 run and ET detector development.



OPO threshold measurement

Improvements since then :

- Hardware and software development for squeezing angle control loop
- OPO has been realigned and characterized
- Preparation of a vacuum sub-tank to decorrelate squeezing measurement improvement between :
 - vacuum squeezing generation
 - vacuum homodyne detection (HD)



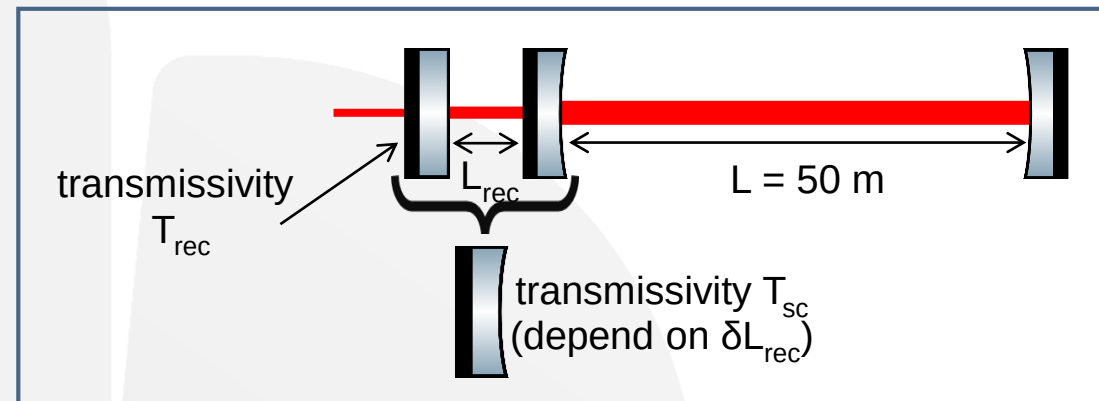
Vacuum tank for HD

- Study the design constrains of a 3-mirrors cavity



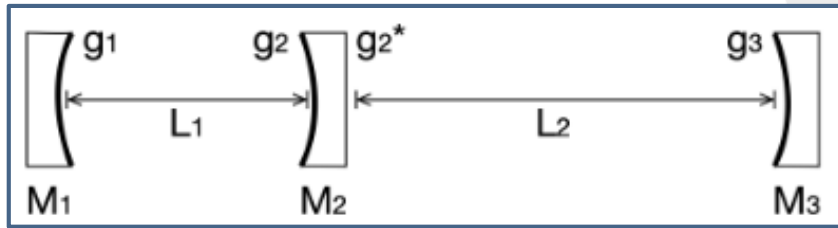
50m vacuum cavity at CALVA facility

- Use a “small cavity” as a variable transmission mirror to tune the squeezing angle corner frequency without change of mirrors :
 - + Faster adaptation implementation
 - + No supplementary cost once installed
 - ? Stability and control of coupled cavities



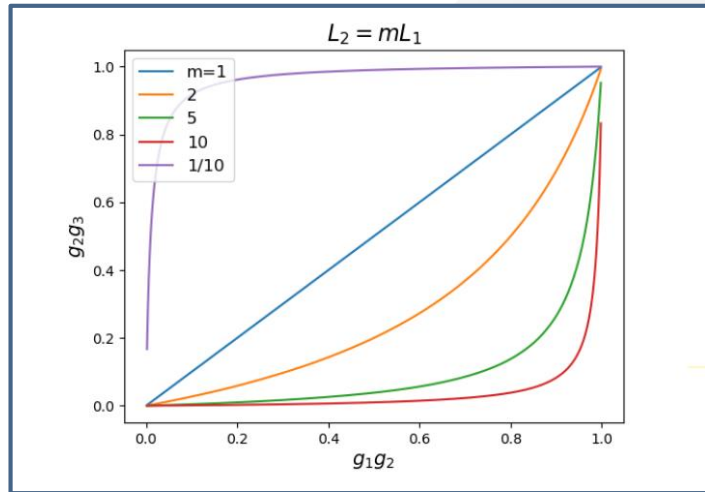
Qfilter principle scheme

- Study the design constrains of a 3-mirrors cavity

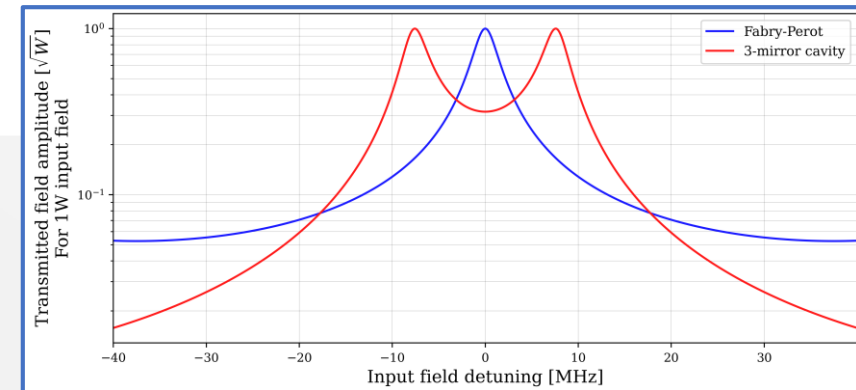


Coupled cavity notation

- Characterize the exotic properties of coupled cavities and the impact on cavity design



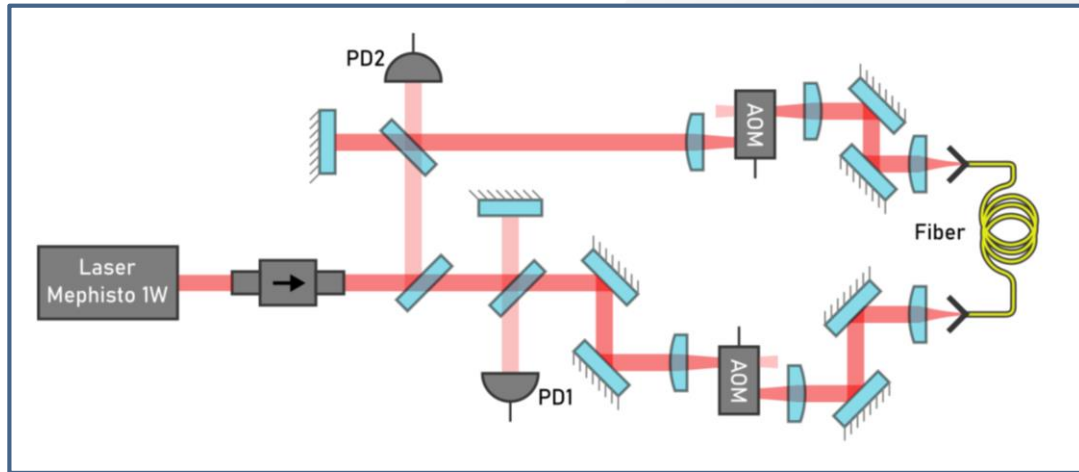
Effect of cavity length dissymmetry on stability



Double resonance peaks of a 3-mirrors cavity VS "classical" resonance peak of 2-mirrors cavity

- GW detectors are large scale facilities and it's necessary for several application to distribute laser beams and theirs phase safely at different place of the interferometer
=> Fibers are a convenient transport method but the phase noise developed inside them is problematic

Squeezing on Virgo is one of the sub-system impacted by this problem

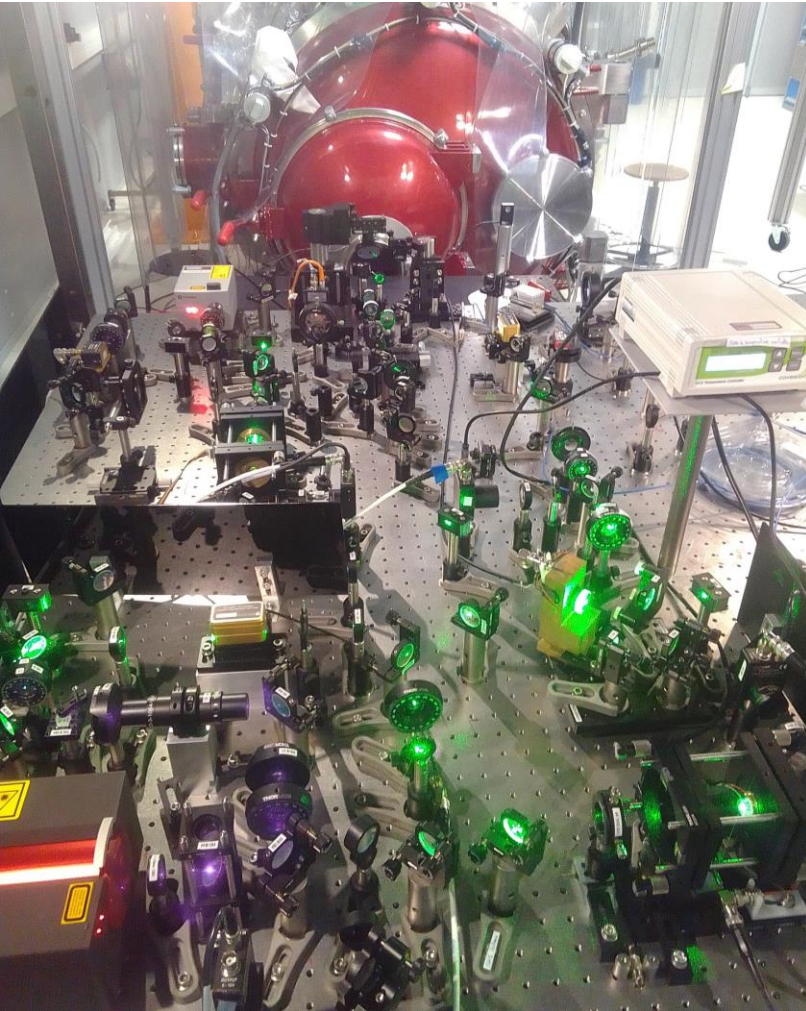


Fiber noise cancellation experiment planned at IJCLab (based on SYRTE)

A first meeting has taken place and more are planned.

Several teams plan to work on this subject :

- Padova
- Nikhef
- EGO
- IJCLAB



Any questions ?

