

Ugo de Noyers

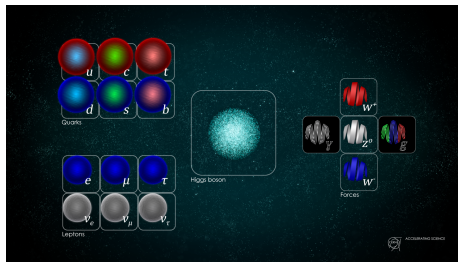
Unraveling Dark Matter and neutrino mysteries with a scotogenic approach

in collaboration with Maud Sarazin
and Björn Herrmann

April 15, 2024



Standard Model and its limitations



$$G_{\text{SM}} = SU(3)_C \times SU(2)_L \times U(1)_Y$$

Advantages

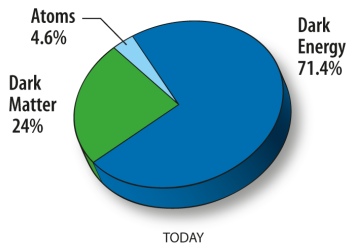
- Higgs boson prediction
- Observables well tested experimentally

Major problems

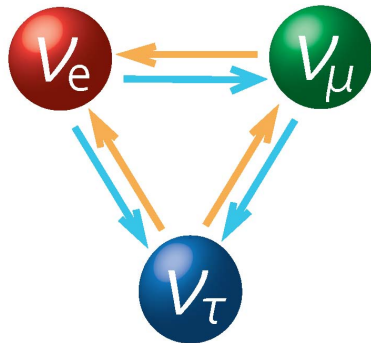
- Gravity not included
- Description of the visible matter only
- Neutrinos remain massless

Unexplained phenomena

Dark Matter nature unknown despite its abundance



Neutrino oscillations linked to them having a mass



Planck measurements^a :

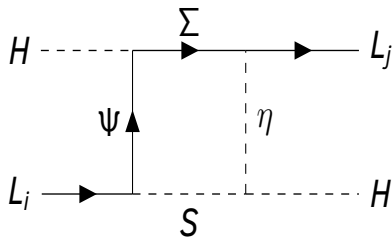
$$\Omega h^2 = 0.1200 \pm 0.0012$$

^a[arXiv:1807.06209v4](https://arxiv.org/abs/1807.06209v4)

$$P(\nu_i \longleftrightarrow \nu_j) \propto \sin^2 \left(\frac{\Delta m_{ij}^2 L}{4E} \right)$$

Scotogenic models: a possible extension of the SM

T12G topology



	Ψ_1	Ψ_2	Σ_1	Σ_2	S	η
$SU(3)_c$	1	1	1	1	1	1
$SU(2)_L$	2	2	3	3	1	2
$U(1)_Y$	1	-1	0	0	0	1
$[M]$	3/2	3/2	3/2	3/2	1	1

Same gauge symmetry
as SM

Addition of an extra
symmetry \mathbb{Z}_2 :

SM particles are even
under this symmetry

BSM particles are odd

Lagrangian of our model

$$\begin{aligned} -\mathcal{L}_{\text{scalar}} &\supset M_H^2 |H|^2 + \lambda_H |H|^4 + \frac{1}{2} M_S^2 S^2 + \frac{1}{2} \lambda_{4S} S^4 + M_\eta^2 |\eta|^2 + \lambda_{4\eta} |\eta|^4 \\ &+ \frac{1}{2} \lambda_S S^2 |H|^2 + \lambda_\eta |\eta|^2 |H|^2 + \frac{1}{2} \lambda_{S\eta} S^2 |\eta|^2 + \lambda'_\eta |\eta^\dagger H|^2 \\ &+ \frac{1}{2} \lambda''_\eta \left((\eta^\dagger H)^2 + \text{h.c.} \right) + \kappa \left(S \eta^\dagger H + \text{h.c.} \right) \\ -\mathcal{L}_{\text{fermion}} &\supset \frac{1}{2} M_{\Sigma_1} \bar{\Sigma}_1 \Sigma_1 + \frac{1}{2} M_{\Sigma_2} \bar{\Sigma}_2 \Sigma_2 \\ &+ M_\Psi \Psi_1 \Psi_2 + y_{1j} \Psi_1 \Sigma_j H + y_{2j} \bar{\Psi}_2 \Sigma_j H \\ &+ g_\Psi^k \Psi_2 L_k S + g_{\Sigma_j}^k \eta \Sigma_j L_k + g_R^k \tilde{\eta} \Psi_1 e_k^c + \text{h.c.} \end{aligned}$$

Why this topology?

- T12A topology already been studied by Björn Herrmann and Maud Sarazin
- T12A model only allows for 2 neutrino masses to be generated
- Variant of T12A with extra fermion singlet can generate 3 neutrino masses
- Phenomenology of a fermion singlet F has been studied so now we wanted to study the one of a fermion triplet

MCMC algorithm

36 parameters to scan in the MCMC algorithm \implies consequent computational time to scan a 36-D hypercube

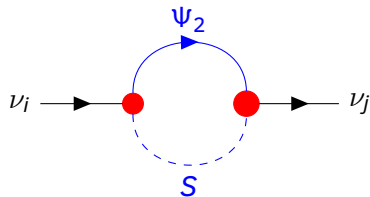
Use of Casas-Ibarra parametrization to take in account experimental constraints in input

Likelihood computed as $\mathcal{L}_n = \prod_i \mathcal{L}_i^n$ with $\ln(\mathcal{L}_i^n) = -\frac{(\mathcal{O}_i^n - \mathcal{O}_i^{\text{exp}})^2}{2\sigma_i^2}$

33 experimental constraints were taken in account

Jump to next point if $\mathcal{L}_n > \mathcal{L}_{\text{old}}$, if not then jump with a $p = \mathcal{L}_n / \mathcal{L}_{\text{old}}$ probability

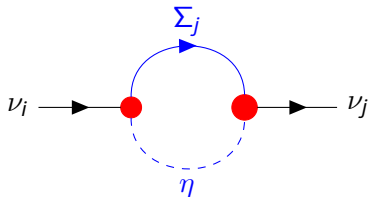
Advantages: mass generation of 3 neutrinos



Couplings determined thanks

to Casas-Ibarra
parametrization

$$M_\nu = g^t M_L \mathcal{G}$$



$$\mathcal{G} = \begin{pmatrix} g_\Psi^e & g_\Psi^\mu & g_\Psi^\tau \\ g_{\Sigma_1}^e & g_{\Sigma_1}^\mu & g_{\Sigma_1}^\tau \\ g_{\Sigma_2}^e & g_{\Sigma_2}^\mu & g_{\Sigma_2}^\tau \end{pmatrix}$$

Casas-Ibarra parametrization

- Diagonalization of the neutrino mass matrix by the PMNS matrix

$$M_\nu = V_{\text{PMNS}}^\dagger D_\nu V_{\text{PMNS}}^*$$

Pontecorvo-Maki-Nakagawa-Sakata (PMNS) matrix expresses the mismatch between the rotations of the LH charged leptons and the neutrinos

$$V_{\text{PMNS}} = V_{eL}^\dagger V_{\nu L}$$

- Decomposition of neutrino mass matrix in different terms

$$M_\nu = \mathcal{G}^t M_L \mathcal{G}$$

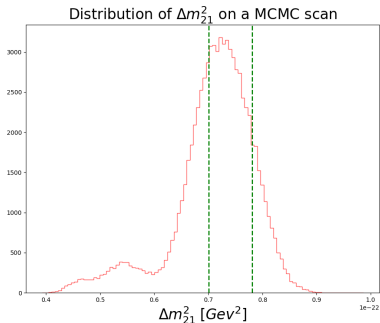
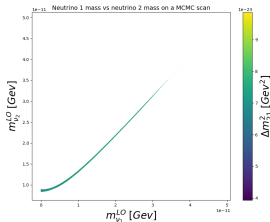
Combining those two statements leads to

$$\mathcal{G} = U_L D_L^{-1/2} R D_\nu^{1/2} V_{\text{PMNS}}^*$$

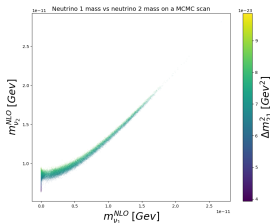
$$\text{with } R^t R = R R^t = \mathbb{I}_3$$

Satisfying neutrino mass differences

At LO level



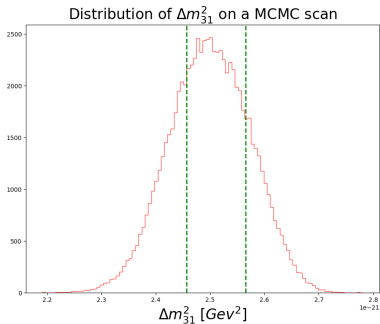
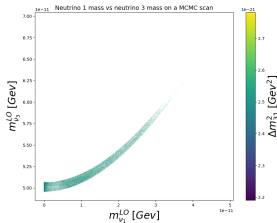
At NLO level



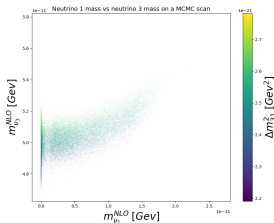
Second bump might be linked to loop corrections

Satisfying neutrino mass differences

At LO level



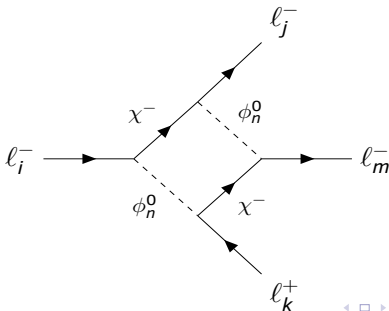
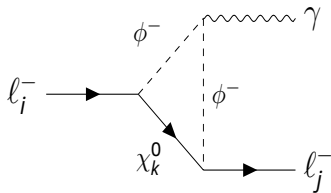
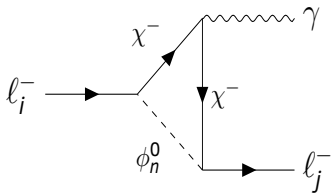
At NLO level



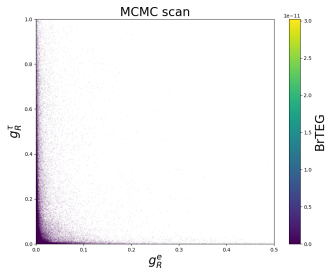
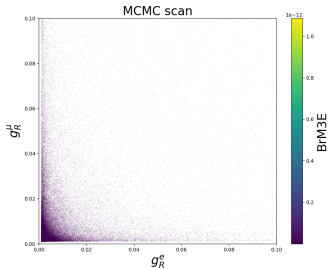
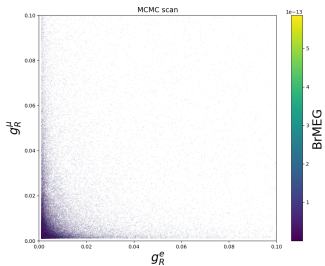
Clean peak

Constraints: Lepton Flavor Violation

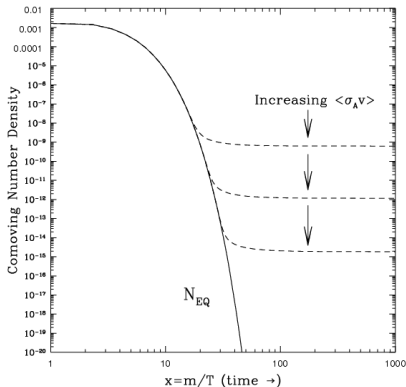
$$-\mathcal{L}_{\text{fermion}} \supset g_R^k e_k^c \tilde{\eta} \Psi_1 + (g_\Psi^k \Psi_2 L_k S + g_{\Sigma_j}^k \eta \Sigma_j L_k)$$



Constraints: Lepton Flavor Violation



Details on Freeze-out



Features

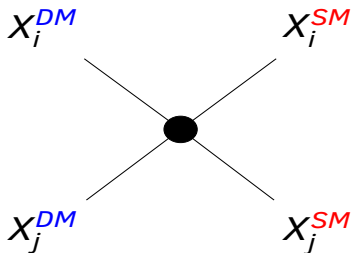
- DM in thermal equilibrium with thermal bath deep within the radiation-dominated epoch
- as $\Gamma \lesssim H$ DM decouples first chemically and then kinematically from the thermal bath
- Correct relic density can be reached by different ways

Advantages: stability of Dark Matter particle

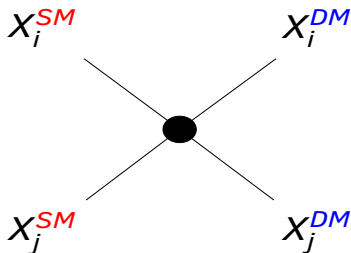
Boltzmann equation in FLRW model

$$\frac{dn}{dt} + 3Hn = -\langle\sigma v\rangle \left(n^2 - (n^{\text{eq}})^2\right)$$

$$\text{with } \Omega h^2 = \frac{nm_{\text{DM}}}{\rho_c}$$

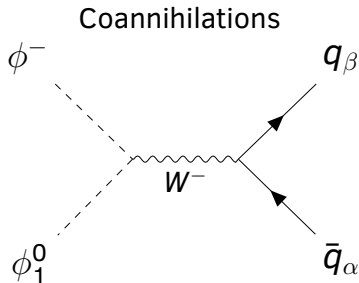
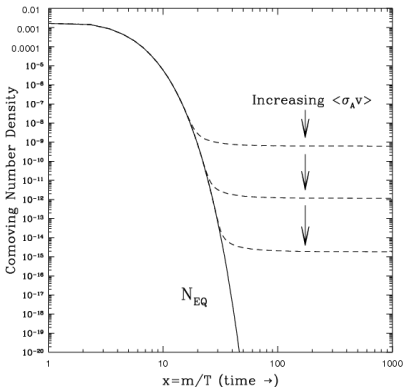


Freeze-out process



Freeze-in process

Satisfying DM relic density



$$(\phi_1^0, \phi_2^0, A^0) = U_\phi (S, \eta^0, A^0)$$

$$\frac{dn}{dt} = -3Hn - \sum_{i,j=1}^N \langle \sigma_{ij} v_{ij} \rangle (n_i n_j - n_i^{\text{eq}} n_j^{\text{eq}})$$

Dark matter candidate?

Diagonalization of mass matrices

$$\mathcal{M}_{\chi^0} = \begin{pmatrix} M_{\Sigma_1} & 0 & \frac{v}{\sqrt{2}}y_{11} & \frac{v}{\sqrt{2}}y_{21} \\ 0 & M_{\Sigma_2} & \frac{v}{\sqrt{2}}y_{12} & \frac{v}{\sqrt{2}}y_{22} \\ \frac{v}{\sqrt{2}}y_{11} & \frac{v}{\sqrt{2}}y_{12} & 0 & M_{\Psi} \\ \frac{v}{\sqrt{2}}y_{21} & \frac{v}{\sqrt{2}}y_{22} & M_{\Psi} & 0 \end{pmatrix}$$

$$\Rightarrow (\chi_1^0, \chi_2^0, \chi_3^0, \chi_4^0) = U_{\chi} (\Sigma_1^0, \Sigma_2^0, \Psi_1^0, \Psi_2^0)$$

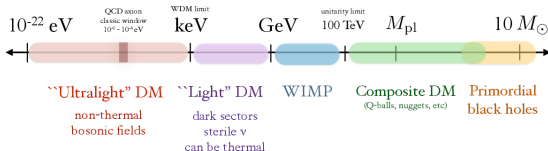
$$\mathcal{M}_{\phi^0}^2 = \begin{pmatrix} M_S^2 + \frac{1}{2}v^2\lambda_S & v\kappa & 0 \\ v\kappa & M_{\eta}^2 + \frac{1}{2}v^2\lambda_L & 0 \\ 0 & 0 & M_{\eta}^2 + \frac{1}{2}v^2\lambda_A \end{pmatrix}$$

$$\Rightarrow (\phi_1^0, \phi_2^0, A^0) = U_{\phi} (S, \eta^0, A^0)$$

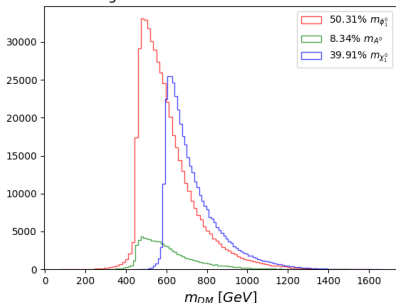
Dark matter candidate

Mass scale of dark matter

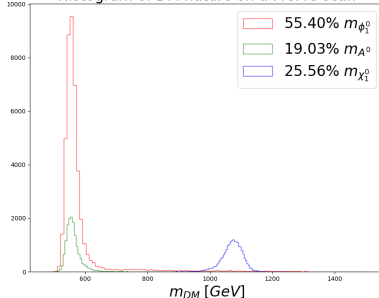
(not to scale)



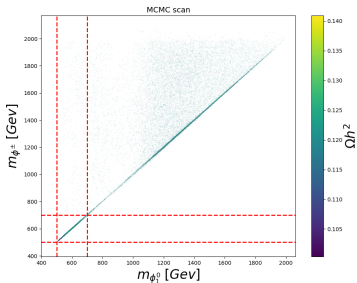
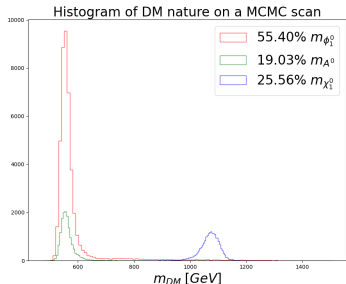
Histogram of DM nature on a random scan



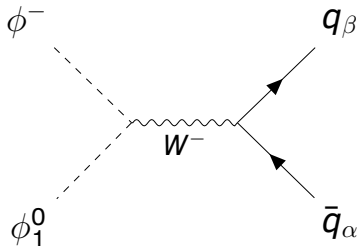
Histogram of DM nature on a MCMC scan



Satisfying DM relic density



Coannihilations

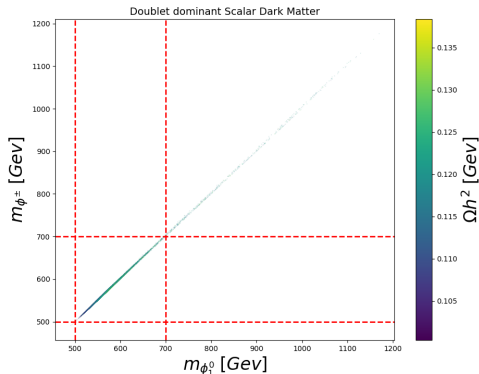
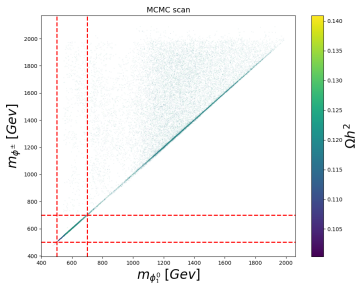
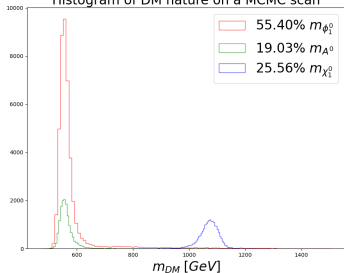


$$(\phi_1^0, \phi_2^0, A^0) = U_\phi (S, \eta^0, A^0)$$

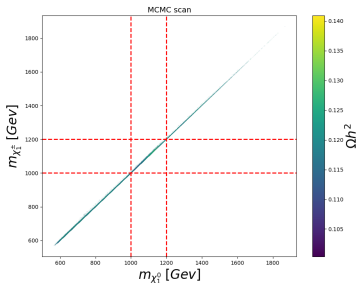
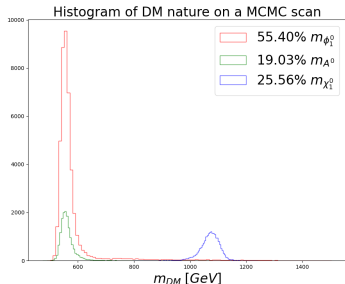
$$\eta = \begin{pmatrix} \eta^+ \\ \frac{1}{\sqrt{2}} (\eta^0 + iA^0) \end{pmatrix}$$

Satisfying DM relic density

Histogram of DM nature on a MCMC scan



Satisfying DM relic density



Coannihilations

$$(\chi_1^0, \chi_2^0, \chi_3^0, \chi_4^0) = U_{\chi^0} (\Sigma_1^0, \Sigma_2^0, \Psi_1^0, \Psi_2^0)$$

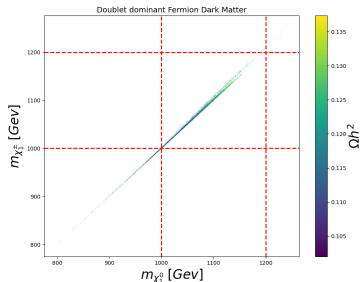
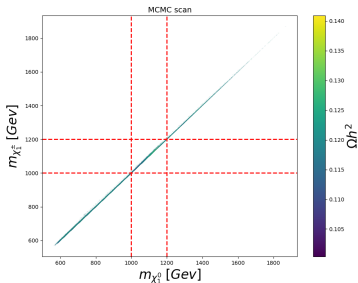
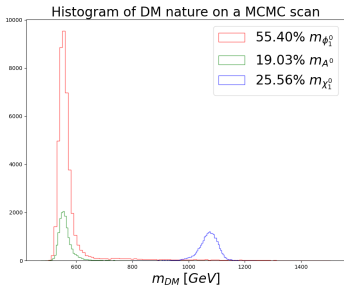
$$(\chi_1^+, \chi_2^+, \chi_3^+) = U_{\chi^+} (\Sigma_1^+, \Sigma_2^+, \Psi_2^+)$$

$$(\chi_1^-, \chi_2^-, \chi_3^-) = U_{\chi^-} (\Sigma_1^-, \Sigma_2^-, \Psi_1^-)$$

$$\Sigma_j = \begin{pmatrix} \frac{\Sigma_j^0}{\sqrt{2}} & \Sigma_j^+ \\ \Sigma_j^- & -\frac{\Sigma_j^0}{\sqrt{2}} \end{pmatrix}$$

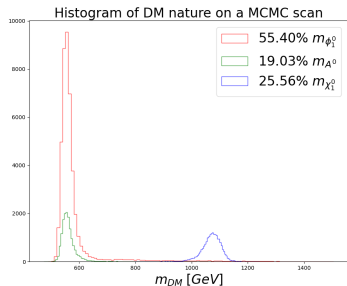
$$\Psi_1 = \begin{pmatrix} \Psi_1^0 \\ \Psi_1^- \end{pmatrix}, \Psi_2 = \begin{pmatrix} -\Psi_2^+ \\ \Psi_2^0 \end{pmatrix}$$

Satisfying DM relic density

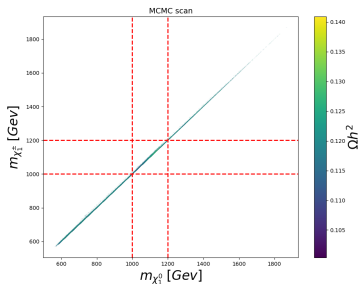
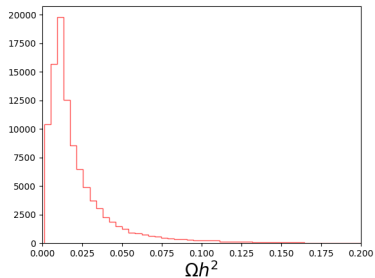


Fermionic DM only comes from the doublets

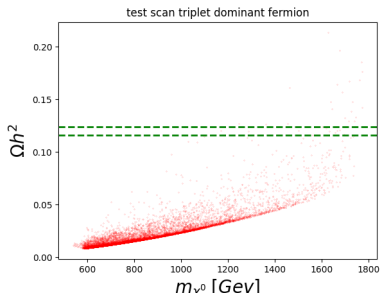
Satisfying DM relic density



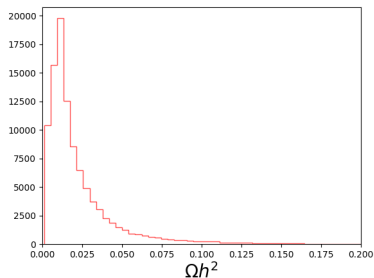
Fermionic DM only comes from the doublets



Satisfying DM relic density



Fermionic DM only comes from the doublets



- Extension of SM that deals with DM and neutrino masses
- Regions in parameter space satisfying major constraints (DM relic density and neutrino mass differences)
- Mass of BSM particles are reachable at the LHC

Thanks for your attention