





Modular Symmetry: a new perspective on the Flavor Puzzle

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$3\,\nu\,{\rm Paradigm}$



 $c_{ij} \equiv \cos \theta_{ij}$, $s_{ij} \equiv \sin \theta_{ij}$ Pontecorvo - Maki - Nakagawa - Sakata

Flavor puzzle and neutrino mixing





Flavor puzzle and neutrino mixing



Flavor puzzle and neutrino mixing



















$$\int d^4x \, d^6y \, \mathscr{L}_{10D} \implies \int d^4x \, \mathscr{L}_{EFT}$$

What is a modular form?

•
$$Y(\gamma(\tau)) = (c\tau + d)^k Y(\tau)$$

Holomorphic in: $\{\tau \in \mathbb{C} \mid \operatorname{Im}(\tau) > 0\}$

$$c, d \in \begin{pmatrix} a & b \\ c & d \end{pmatrix} \equiv \gamma$$
$$a, b, c, d \in \mathbb{Z} \quad , \quad ad - bc = 1$$

"Weight" k>0



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$$\blacktriangleright \mathcal{W}(\Phi) = \sum (Y_{I_1...I_n}(\tau) \varphi^{(I_1)} ... \varphi^{(I_n)})_{\mathbf{1}}$$

 $\mathscr{W}(\Phi)$ modular invariant if:

$$\begin{cases} \rho \otimes \rho_{I_1} \otimes \rho_{I_2} \dots \otimes \rho_{I_n} \supset \mathbf{1} & \longrightarrow & \text{Usual} \\ k_Y = k_{I_1} + k_{I_2} + \dots + k_{I_n} & \longrightarrow & \text{Novelty} \end{cases}$$

$$Y_{I_1...I_n}(\tau) \to (c\tau + d)^{k_Y} \rho(\gamma) Y_{I_1...I_n}(\tau)$$

Yukawa: modular forms of weight k_{y}

 $\varphi^{(I)} \to (c\tau + d)^{-k_I} \rho^{(I)}(\gamma) \varphi^{(I)}$

Superfields with modular charges $-k_I$

 $(c\tau + d)^{k_{Y}}(c\tau + d)^{-\sum k_{I_{n}}} = 1$

Symmetry breaking





$$\mathsf{M}_{e} \sim \sum_{i} \alpha_{i} \begin{pmatrix} f_{11}(\tau) & f_{12}(\tau) & \dots \\ \dots & \dots & \dots \\ \dots & \dots & f_{33}(\tau) \end{pmatrix}$$

 $f_{ij} \equiv$ pre-determined functions of τ







Let's try an example!

Weight: $k_Y = 2$

Let's choose
$$N = 3$$
 ($\Gamma_3 \cong A_4$)
Neutrino sector: Weinberg $\mathcal{W}_{\nu} \supset \frac{1}{\Lambda} (H_u H_u L L Y)_1$
Low weights:
 $k_L = +1$, $k_u = 0$ F. Feruglio [1706.08749]
 $L \sim 3 \equiv (L_e, L_\mu, L_\tau)$ $L_a \sim 2_{-1/2}$ under $SU(2)_L \times U(1)_Y$
 $\begin{pmatrix} Y_1(\tau) \\ Y_2(\tau) \\ Y_3(\tau) \end{pmatrix} \sim 3$
 $L \otimes L \otimes Y_3^{(2)} \sim (3 \otimes 3) \otimes 3 \supset 1$

Weights: $-k_L - k_L + k_Y = 0 \rightarrow -1 - 1 + 2 = 0$



Let's try an example!

Neutrino sector: Weinberg

$$\begin{pmatrix} Y_1(\tau) \\ Y_2(\tau) \\ Y_3(\tau) \end{pmatrix} \sim \mathbf{3}$$

 $\mathcal{W}_{\nu} \supset \frac{1}{\Lambda} (H_u H_u L L Y)_1$

$$m_{\nu} = \begin{pmatrix} 2Y_1 & -Y_3 & -Y_2 \\ -Y_3 & 2Y_2 & -Y_1 \\ -Y_2 & -Y_1 & 2Y_3 \end{pmatrix} \frac{v_u^2}{\Lambda}$$

Weight: $k_Y = 2$

$$Y_1(\tau) = 1 + 12q + 36q^2 + 12q^3 + \dots$$

$$Y_2(\tau) = -6q^{1/3}(1 + 7q + 8q^2 + \dots)$$

$$Y_3(\tau) = -18q^{2/3}(1 + 2q + 5q^2 + \dots)$$

$$q \equiv e^{2\pi i \tau}$$



But poor fit, too constrained

Non-Abelian discrete flavour symmetry

$$\mathcal{W}_L \supset \frac{\alpha}{\Lambda} E_i^c (L_j \varphi_k)_1 H_d$$

$$\varphi \sim \mathbf{2} = \begin{pmatrix} \varphi_1 \\ \varphi_2 \end{pmatrix} \rightarrow \mathsf{Flavon}$$

Free parameters

Modular flavour symmetry

$$\mathcal{W}_L \supset \alpha E_i^c [L_j(Y_2(\tau))_k]_1 H_d$$

$$Y_{2}(\tau) \sim \mathbf{2} = \begin{pmatrix} Y_{1}(\tau) \\ Y_{2}(\tau) \end{pmatrix}$$

Completely fixed

Structure dependent on the symmetry breaking sector

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The only unknown is the complex VEV of $\ \tau$

$$, j, k$$
, flavour indices





$\tau = \operatorname{Re} \tau + i \operatorname{Im} \tau$

Only source of CPV in the model is the VEV of $\boldsymbol{\tau}$

The fundamental domain of τ and CP



$$\mathcal{D} = \left\{ \tau \in \mathbb{C} : \operatorname{Im} \tau > 0, |\operatorname{Re} \tau| \le \frac{1}{2}, |\tau| \ge 1 \right\}$$

Every $\tau \notin \mathscr{D}$ can be mapped in $\tau' \in \mathscr{D}$ through Γ transformation



Only source of CPV in the model is the VEV of $\boldsymbol{\tau}$





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Finite modular groups: searching for simplicity



A basis for modular forms of weight k

$$\eta(\tau) \equiv q^{1/24} \prod_{n=1}^{\infty} (1-q^n)$$

$$P_1(\tau), Y_2(\tau) \sim 2$$

$$Lowest weight: k=2$$

$$\begin{pmatrix} Y_1(\tau) \\ Y_2(\tau) \end{pmatrix}_2 \rightarrow (c\tau+d)^2 \rho(\gamma)_2 \begin{pmatrix} Y_1(\tau) \\ Y_2(\tau) \end{pmatrix}_2$$

$$Very limited number!$$

$$N \quad d_k(\Gamma(N)) \quad \Gamma_N \\ 2 \quad k/2+1 \quad S_3$$

$$1' \otimes 1' = 1 \quad , \quad 1' \otimes 2 = 2 \quad , \quad 2 \otimes 2 = 1 \oplus 1' \oplus 2$$

Guiding principles



Extra flavons besides τ (modulus)





$$Y_{1}(\tau) = \frac{7}{100} + \frac{42}{25}q + \frac{42}{25}q^{2} + \frac{168}{25}q^{3} + \dots$$

$$Y_{2}(\tau) = \frac{14\sqrt{3}}{25}q^{1/2}(1 + 4q + 6q^{2} + \dots)$$

$$Im \tau \gtrsim 1$$
Fourier expansion $(q \equiv e^{2\pi i \tau})$

Smallest modular finite group: S_3

Following guiding principles.

$$D_{\ell} \equiv \begin{pmatrix} \text{electron} \\ \text{muon} \end{pmatrix} \sim 2 \qquad \qquad \ell_3 \equiv \text{tau} \sim 1'$$

		E_1^c	E_2^c	E_3^c	D_ℓ	ℓ_3	$H_{d,u}$
	$SU(2)_L \times U(1)_Y$	(1, +1)	(1, +1)	(1, +1)	(2, -1/2)	(2, -1/2)	$(2, \mp 1/2)$
Irreps	$\Gamma_2 \cong S_3$	1	1′	1′	2	1′	1
Veights	k_I	4	0	-2	2	2	0

Charged-leptons

$$\mathcal{W}_{e}^{H} = \alpha E_{1}^{c} H_{d} (D_{\ell} Y_{2}^{(3)})_{1} + \beta E_{2}^{c} H_{d} (D_{\ell} Y_{2})_{1'} + \gamma E_{3}^{c} H_{d} \ell_{3} + \alpha_{D} E_{1}^{c} H_{d} \ell_{3} Y_{1'}^{(3)}$$

$$(m_{\tau}, m_{\mu}, m_{e}) \sim m_{\tau} (1, |Y_{1}|, |Y_{1}^{3}|) \qquad |Y_{1}| \sim \mathcal{O}(10^{-2})$$

Smallest modular finite group: S_3



$$\mathcal{W}_{e}^{H} = \alpha E_{1}^{c} H_{d} (D_{\ell} Y_{2}^{(3)})_{1} + \beta E_{2}^{c} H_{d} (D_{\ell} Y_{2})_{1'} + \gamma E_{3}^{c} H_{d} \ell_{3} + \alpha_{D} E_{1}^{c} H_{d} \ell_{3} Y_{1'}^{(3)}$$

$$(m_{\tau}, m_{\mu}, m_{e}) \sim m_{\tau} (1, |Y_{1}|, |Y_{1}^{3}|) |Y_{1}| \sim \mathcal{O}(10^{-2})$$

$$M_{\ell}^{\dagger} = \begin{pmatrix} \alpha(Y_{2}^{(3)})_{1} & \alpha(Y_{2}^{(3)})_{2} & \alpha_{D}Y_{1'}^{(3)} \\ \beta Y_{2} & -\beta Y_{1} & 0 \\ 0 & 0 & \gamma \end{pmatrix} v_{d}$$
Charged-leptons masses reproduced with:
$$\frac{\beta}{\alpha} \sim \frac{\gamma}{\alpha} \sim \frac{\alpha_{D}}{\alpha} \approx \mathcal{O}(1)$$

Modular invariance: texture zeros!

Smallest modular finite group: S_3

Neutrino sector from Weinberg operators

$$\mathcal{W}_{\nu}^{k_{\ell}=2} \supset \frac{g}{\Lambda} H_{u} H_{u} (D_{\ell} D_{\ell})_{2} Y_{2}^{(2)} + \frac{g'}{\Lambda} H_{u} H_{u} D_{\ell} \ell_{3} (Y_{2}^{(2)}) + \\ + \frac{g''}{\Lambda} H_{u} H_{u} (D_{\ell} D_{\ell})_{1} Y_{1}^{(2)} + \frac{g_{p}}{\Lambda} H_{u} H_{u} \ell_{3} \ell_{3} (Y_{1}^{(2)}) .$$

$\blacktriangleright \Lambda \rightarrow \text{Scale of new physics}$

► {g'/g, g''/g, g_p/g } $\in \mathbb{R}$ free dimensionless parameters





JHEP 09 (2023) 043 D. Meloni, M.Parriciatu



Numerical results with modular S_3

Fit VS 6 dimensionless observables (no CP) $q_j = \{\sin^2 \theta_{12}, \sin^2 \theta_{13}, \sin^2 \theta_{23}, m_e/m_\mu, m_\mu/m_\tau, r\}$





Parameter	Best-fit value and 1σ range			
$\Delta m_{\rm sol}^2/(10^{-5}~{\rm eV}^2)$	$7.36\substack{+0.16 \\ -0.15}$			
	NO	ΙΟ		
$ \Delta m^2_{ m atm} /(10^{-3}~{ m eV}^2)$	$2.485\substack{+0.023\\-0.031}$	$2.455\substack{+0.030\\-0.025}$		
$r\equiv \Delta m_{ m sol}^2/ \Delta m_{ m atm}^2 $	0.0296 ± 0.0008	0.0299 ± 0.0008		
$\sin^2 heta_{12}$	$0.303\substack{+0.013\\-0.013}$	$0.303\substack{+0.013\\-0.013}$		
$\sin^2 heta_{13}$	$0.0223\substack{+0.0007\\-0.0006}$	$0.0223\substack{+0.0006\\-0.0006}$		
$\sin^2 heta_{23}$	$0.455\substack{+0.018\\-0.015}$	$0.569\substack{+0.013\\-0.021}$		
$\delta_{ m CP}/\pi$	$1.24\substack{+0.18 \\ -0.13}$	$1.52\substack{+0.14 \\ -0.15}$		
m_e/m_μ	0.0048 ± 0.0002			
$m_{\mu}/m_{ au}$ 0.0565 ± 0.0045		± 0.0045		
	F. Capozzi et al.			

Phys. Rev. D **104** (Oct, 2021)

10 (6) parameters

VS

12 (9) observables

Numerical results with modular S_3

Model predictions

Normal ordering for neutrinos



Model predictions

Precise predictions on neutrino masses and observables of exp. interest!





What S_3 provides in the literature











BACKUP SLIDES



But what about a $\Gamma_2 \simeq S_3$ seesaw version?

Recent work with S.Marciano, D.Meloni: arxiv:2402.18547

Reproduces low-energy CP-violation and matter-antimatter asymmetry of the Universe through Leptogenesis



Introduce Minimal seesaw scenario with only 2 RHN transforming as ~ 2 under S_3 with weight 2

$$\mathcal{W}_{\nu} = gH_{u}N^{c}D_{\ell}Y_{\mathbf{2}}^{(2)} + g'H_{u}(N^{c}Y_{\mathbf{2}}^{(2)})_{\mathbf{1}'}\ell_{3} + g''H_{u}(N^{c}D_{\ell})_{\mathbf{1}}Y_{\mathbf{1}}^{(2)} + \\ + \Lambda[(N^{c}N^{c})_{\mathbf{2}}Y_{\mathbf{2}}^{(2)} + \lambda(N^{c}N^{c})_{\mathbf{1}}Y_{\mathbf{1}}^{(2)}],$$

$$M_{D} = gv_{u} \begin{pmatrix} -(Y_{2}^{2} - Y_{1}^{2}) + \frac{g''}{g}(Y_{1}^{2} + Y_{2}^{2}) & 2Y_{1}Y_{2} & \frac{g'}{g}(2Y_{1}Y_{2}) \\ 2Y_{1}Y_{2} & (Y_{2}^{2} - Y_{1}^{2}) + \frac{g''}{g}(Y_{1}^{2} + Y_{2}^{2}) & -\frac{g'}{g}(Y_{2}^{2} - Y_{1}^{2}) \end{pmatrix}_{\mathrm{RL}}$$

$$\mathcal{M}_R = \Lambda \begin{pmatrix} -(Y_2^2 - Y_1^2) + \lambda(Y_1^2 + Y_2^2) & 2Y_1Y_2 \\ 2Y_1Y_2 & (Y_2^2 - Y_1^2) + \lambda(Y_1^2 + Y_2^2) \end{pmatrix}_{\mathrm{RR}}.$$

$$m_{
u} = -M_D^T \mathcal{M}_R^{-1} M_D$$
.



Excellent fit: $\chi^2 \sim 0.98$ (but now $\delta_{\rm CP}$ is fitted)

But what about a $\Gamma_2 \simeq S_3$ seesaw version?

Recent work with S.Marciano, D.Meloni: arxiv:2402.18547

Reproduces low-energy CP-violation and matter-antimatter asymmetry of the Universe through Leptogenesis



 $\Lambda \rightarrow r\Lambda$ Majorana mass scale not completely fixed by low-energy

Backup slides: Clebsch-Gordan

 \blacktriangleright Clebsch-Gordan coefficients for S_3

$$\mathbf{2}\otimes\mathbf{2} = \mathbf{1}\otimes\mathbf{1}'\otimes\mathbf{2} egin{cases} \mathbf{1} & \sim & \psi_1arphi_1 + \psi_2arphi_2\ \mathbf{1}' & \sim & \psi_1arphi_2 - \psi_2arphi_1\ \mathbf{2} & \sim & \psi_1arphi_2 - \psi_2arphi_1\ \mathbf{2} & \sim & \left(egin{array}{c} \psi_1\ \psi_2\ \mathbf{2}\ \mathbf{2}\ \mathbf{2}\ \mathbf{2}\ \mathbf{2}\ \mathbf{1}'\ \mathbf{2}\ \mathbf{2}\$$

$$egin{aligned} \mathbf{1}' \otimes \mathbf{1}' &= \mathbf{1} & \sim y_1 y_2 \ \mathbf{1}' \otimes \mathbf{2} &= \mathbf{2} & \sim & egin{pmatrix} -y_1 \psi_2 \ y_1 \psi_1 \end{pmatrix} \end{aligned}$$

 $y_1, y_2 \equiv \text{pseudo-singlets} (1')$

e.g.
$$(D_{\ell}Y_2(\tau))_{1'} = (D_{\ell})_1(Y_2(\tau))_2 - (D_{\ell})_2(Y_2(\tau))_1 \sim \mathbf{1}'$$

The Modular symmetry approach

 $\sigma \equiv$

Modular-invariant SUSY action

$$S = \int d^4x \int d^2\theta d^2\bar{\theta} \ K(\Phi,\bar{\Phi}) + \left[\int d^4x \int d^2\theta \ W(\Phi) + \text{h.c.} \right]$$

$$K\ddot{a}hler \text{ potential}$$

$$\sigma \equiv \Lambda_{\tau} \tau$$

$$Superpotential$$

$$Gives the kinetic terms after the modulus acquires a VEV \\A minimalistic form is chosen$$

$$A \min anistic form is chosen$$

$$Fhe superfields transform as:$$

$$\left\{ \tau \to \gamma(\tau) = \frac{a\tau + b}{c\tau + d} \\ \varphi^{(I)} \to (c\tau + d)^{-k_I} \rho^{(I)}(\gamma) \varphi^{(I)} \end{aligned}
, with \ \gamma = \begin{pmatrix} a & b \\ c & d \end{pmatrix} \in \Gamma_N$$

$$\left\{ \begin{array}{c} W(\Phi) \to W(\Phi) \\ W(\Phi) \to W(\Phi) \\ K(\Phi,\bar{\Phi}) \to \underbrace{K(\Phi,\bar{\Phi}) + f(\Phi) + f(\bar{\Phi})} \\ W(\Phi) \to W(\Phi) \\ K(\Phi,\bar{\Phi}) \to \underbrace{K(\Phi,\bar{\Phi}) + f(\Phi) + f(\bar{\Phi})} \\ Kahler transformation \\ \end{array} \right\}$$

The Kähler potential...



In a bottom-up approach, this is unjustified



M.-C. Chen, S. Ramos-Sánchez, and M. Ratz, "A note on the predictions of models with modular flavor symmetries," *Physics Letters B* 801 (Feb, 2020) 135153.

Th

This question is an open one

The Modular symmetry approach

The group generators

Finite modular group can be defined: $\Gamma_N \equiv \overline{\Gamma} / \overline{\Gamma}(N)$

$$\Gamma(N) = \left\{ \begin{pmatrix} a & b \\ c & d \end{pmatrix} \in \operatorname{SL}(2, \mathbb{Z}) \, \middle| \, \begin{pmatrix} a & b \\ c & d \end{pmatrix} \equiv \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \pmod{N} \right\}$$

subgroups of Γ N=1,2,3...called "level"

 $\overline{\Gamma} \equiv \Gamma / \{ \pm \mathbb{I} \}$ $\overline{\Gamma}(N) \equiv \Gamma(N) / \{ \pm \mathbb{I} \}$

Generators S and T of the modular group Γ_N

$$\tau \xrightarrow{\mathbf{S}} -\frac{1}{\tau} \qquad \tau \xrightarrow{\mathbf{T}} \tau + 1 \qquad S^2 = T^N = (ST)^3 = \mathbb{I}$$
$$S = \begin{pmatrix} 0 & 1\\ -1 & 0 \end{pmatrix} \qquad , \qquad T = \begin{pmatrix} 1 & 1\\ 0 & 1 \end{pmatrix}$$

 S_3 Generators S and T satisfy: $S^2 = T^2 = (ST)^3 = \mathbb{I}$

$$\rho(S) = \frac{1}{2} \begin{pmatrix} -1 & -\sqrt{3} \\ -\sqrt{3} & 1 \end{pmatrix}, \qquad \rho(T) = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$
$$(\rho(S))^2 = \mathbb{I}, \qquad (\rho(S)\rho(T))^3 = \mathbb{I}, \qquad (\rho(T))^2 = \mathbb{I},$$

The Modular S_3 model: lowest weights

Level 2 modular forms of lowest weight (2) constructed from Dedekind's Eta "seed function"

Closed set under the modular group

$$Y(\alpha, \beta, \gamma | \tau) = \frac{d}{d\tau} [\alpha \log \eta(\tau/2) + \beta \log \eta((\tau+1)/2) + \gamma \log \eta(2\tau)]$$

 $\alpha + \beta + \gamma = 0$

This fixes the constants

$$\begin{cases} Y_1(\tau) = \frac{C}{2} \left[\frac{\eta'(\tau/2)}{\eta(\tau/2)} + \frac{\eta'\left(\frac{\tau+1}{2}\right)}{\eta\left(\frac{\tau+1}{2}\right)} - \frac{8\eta'(2\tau)}{\eta(2\tau)} \right] \\ Y_2(\tau) = \frac{C}{2} \sqrt{3} \left[\frac{\eta'(\tau/2)}{\eta(\tau/2)} - \frac{\eta'\left(\frac{\tau+1}{2}\right)}{\eta\left(\frac{\tau+1}{2}\right)} \right] \end{cases}$$

"C" arbitrary

Impose transformation properties under S_3 generators

$$\begin{split} \rho(S) &= \frac{1}{2} \begin{pmatrix} -1 & -\sqrt{3} \\ -\sqrt{3} & 1 \end{pmatrix}, \qquad \rho(T) = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}, \\ (\rho(S))^2 &= \mathbb{I}, \qquad (\rho(S)\rho(T))^3 = \mathbb{I}, \qquad (\rho(T))^2 = \mathbb{I}, \end{split}$$

$$\left(\begin{array}{c}Y_{1}(\tau)\\Y_{2}(\tau)\end{array}\right)_{2} \rightarrow (c\tau+d)^{2}\rho(\gamma)_{2} \left(\begin{array}{c}Y_{1}(\tau)\\Y_{2}(\tau)\end{array}\right)_{2}$$

The Modular S_3 model: the normalisation

Level 2 modular forms of lowest weight (2) constructed from Dedekind's Eta "seed function"

$$\{\eta(\tau/2), \eta\left(\frac{\tau+1}{2}\right), \eta(2\tau)\}$$

Closed set under the modular group

"C" arbitrary

$$\begin{cases} Y_1(\tau) = \frac{C}{2} \left[\frac{\eta'(\tau/2)}{\eta(\tau/2)} + \frac{\eta'\left(\frac{\tau+1}{2}\right)}{\eta\left(\frac{\tau+1}{2}\right)} - \frac{8\eta'(2\tau)}{\eta(2\tau)} \right] \\ Y_2(\tau) = \frac{C}{2} \sqrt{3} \left[\frac{\eta'(\tau/2)}{\eta(\tau/2)} - \frac{\eta'\left(\frac{\tau+1}{2}\right)}{\eta\left(\frac{\tau+1}{2}\right)} \right] \\ \downarrow \\ \downarrow \\ \downarrow \\ In our case, this is true if \\ C is purely imaginary \\ C = \frac{7i}{2} \\ The choice made in this \end{cases}$$

work

 25π

 Impose CP symmetry on the model
 Superpotential parameters must be real: less free parameters

P. Novichkov, J. Penedo, S. Petcov, A. Titov Journal of High Energy Physics **2019** no. 7, (Jul, 2019)

$$Y(\tau) \xrightarrow{\mathrm{CP}} Y(-\tau^*) = Y^*(\tau)$$

Only source of CPV is the VEV of modulus ${\cal T}$

The Modular S_3 model: charged-leptons sector

Found two viable choices for modular charges and weights



Backup slides

Numerical procedure

Define a "figure of merit", i.e. chi-square for every set of parameters $l(p_i) \equiv \sqrt{\chi^2(p_i)}$

Define a "potential" with a given temperature T and a threshold

$$V(p_i) = \begin{cases} l(p_i) &, \quad l(p_i) \le l_{\max} \\ +\infty &, \quad \text{otherwise} \end{cases}$$

 $\chi^{2}(p_{i}) = \sum_{j=1}^{6} \left(\frac{q_{j}(p_{i}) - q_{j}^{\text{b-f}}}{\sigma_{j}}\right)^{2}$ $p_{i} = \{\tau, \beta/\alpha, \gamma/\alpha, \dots, g'/g, g_{p}/g, \dots\}$ $q_{j} = \{\sin^{2}\theta_{12}, \sin^{2}\theta_{13}, \sin^{2}\theta_{23}, m_{e}/m_{\mu}, m_{\mu}/m_{\tau}, r\}$

P. P. Novichkov, J. T. Penedo, S. T. Petcov, and A. V. Titov, "Modular S_4 models of lepton masses and mixing," (2019)

 At iteration "t", generate a new
 point from a Gaussian centred on the previous one Accept the new point with a probability given by: $P_{\alpha} = \min[1, \exp(V(p_i^{(t)}) - V(p_i'))/T]$ A measure of fine-tuning: Altarelli-Blankenburg

Fine-tuning =
$$\frac{\sum_{i} \left| \frac{\text{par}_{i}}{\delta \text{par}_{i}} \right|}{\sum_{i} \left| \frac{\text{obs}_{i}}{\sigma_{i}} \right|}$$







Backup slides

Model I [7]		
	Best-fit and 1σ range	
$\operatorname{Re} \tau$	$\pm 0.0895^{+0.0032}_{-0.0055}$	
$\operatorname{Im}\tau$	$1.697\substack{+0.023\\-0.016}$	
eta/lpha	$14.33\substack{+0.58\\-0.38}$	
$\gamma/lpha$	$17.39\substack{+1.38 \\ -0.87}$	
g'/g	$31.57\substack{+27.59\\-10.29}$	
g''/g	$7.17\substack{+6.36 \\ -2.36}$	
g_p/g	$8.51\substack{+7.99 \\ -3.03}$	
$v_d \alpha [\text{MeV}]$	102.14	
$v_u^2 g / \Lambda \left[\mathrm{eV} \right]$	0.47	
$\sin^2 heta_{12}$	$0.300\substack{+0.013\\-0.006}$	
$\sin^2 heta_{13}$	$0.0223\substack{+0.0004\\-0.0006}$	
$\sin^2 heta_{23}$	$0.452\substack{+0.015\\-0.009}$	
r	$0.0295\substack{+0.0007\\-0.0006}$	
m_e/m_μ	$0.0048\substack{+0.0001\\-0.0002}$	
$m_\mu/m_ au$	$0.0578\substack{+0.0023\\-0.0040}$	
Ordering	NO	
δ/π	$\pm 1.594^{+0.007}_{-0.010}$	
$m_1 [{ m eV}]$	$0.0182\substack{+0.0018\\-0.0014}$	
$m_2 [{\rm eV}]$	$0.0201\substack{+0.0017\\-0.0013}$	
$m_3 \; [\mathrm{eV}]$	$0.0537\substack{+0.0006\\-0.0005}$	
$\sum_i m_i [\text{eV}]$	$0.092\substack{+0.004\\-0.003}$	
$\langle m_{etaeta} angle~[{ m meV}]$	$18.89^{+1.90}_{-1.47}$	
$m_{eta}^{\mathrm{eff}} \; \mathrm{[meV]}$	$20.26^{+1.69}_{-1.30}$	
$lpha_1/\pi$	$\pm 1.124^{+0.014}_{-0.017}$	
α_2/π	$\pm 0.949^{+0.005}_{-0.005}$	
Fine-tuning	12.2	
$\chi^2_{ m min}$	0.16	

Model II [8]

	Best-fit and 1σ range
$\operatorname{Re} \tau$	$\pm 0.090^{+0.004}_{-0.004}$
$\operatorname{Im}\tau$	$1.688\substack{+0.015\\-0.018}$
eta/lpha	$1.03\substack{+0.04\\-0.04}$
$\gamma/lpha$	$1.26\substack{+0.12 \\ -0.08}$
$lpha_D/lpha$	$1.33\substack{+1.51 \\ -1.05}$
g'/g	$41.9^{+73.7}_{-12.8}$
$g^{\prime\prime}/g$	$9.55\substack{+16.81 \\ -2.91}$
g_p/g	$11.5^{+21.2}_{-3.8}$
$v_d \alpha \; [\text{MeV}]$	1404.6
$v_u^2 g / \Lambda [\text{eV}]$	0.35
$\sin^2 heta_{12}$	$0.305\substack{+0.009\\-0.015}$
$\sin^2 heta_{13}$	$0.0222\substack{+0.0007\\-0.0006}$
$\sin^2 heta_{23}$	$0.454\substack{+0.007\\-0.008}$
r	$0.0295\substack{+0.0007\\-0.0007}$
m_e/m_μ	$0.0048\substack{+0.0002\\-0.0002}$
$m_\mu/m_ au$	$0.0570\substack{+0.0034\\-0.0048}$
Ordering	NO
δ/π	$\pm 1.597^{+0.009}_{-0.006}$
$m_1 [{\rm eV}]$	$0.0174\substack{+0.0011\\-0.0014}$
$m_2 [{\rm eV}]$	$0.0194\substack{+0.0010\\-0.0012}$
$m_3 \; [\mathrm{eV}]$	$0.0535\substack{+0.0004\\-0.0004}$
$\sum_i m_i [\text{eV}]$	$0.090\substack{+0.002\\-0.003}$
$\langle m_{\beta\beta} \rangle \; [{ m meV}]$	$18.14\substack{+1.17 \\ -1.48}$
$m_{eta}^{\mathrm{eff}}~\mathrm{[meV]}$	$19.60\substack{+1.02 \\ -1.25}$
$lpha_1/\pi$	$\pm 1.129^{+0.019}_{-0.013}$
$lpha_2/\pi$	$\pm 0.946^{+0.004}_{-0.004}$
Fine-tuning	11.2
$\chi^2_{ m min}$	0.074

Backup slides

	Best-fit and 1σ range
$\operatorname{Re}\tau$	$\pm 0.244^{+0.012}_{-0.067}$
$\operatorname{Im}\tau$	$1.132\substack{+0.027\\-0.297}$
eta/lpha	$0.92\substack{+0.85\\-0.03}$
$\gamma/lpha$	$-1.20\substack{+0.06\\-2.14}$
$\log_{10}(lpha_D/lpha)$	$-13.4^{+13.2}_{-76.3}$
g'/g	$2.76\substack{+0.21 \\ -0.23}$
$g^{\prime\prime}/g$	$-2.53\substack{+0.13\\-0.03}$
$\log_{10}(\lambda)$	$-12.2^{+10.9}_{-59.2}$
$v_d \alpha$, [GeV]	$1.08\substack{+0.06\\-0.69}$
$v_u^2g^2/\Lambda~[{\rm eV}]$	$3.46\substack{+0.55\\-1.65}$
$\sin^2 heta_{12}$	$0.305\substack{+0.011\\-0.011}$
$\sin^2 heta_{13}$	$0.0221\substack{+0.0006\\-0.0005}$
$\sin^2 heta_{23}$	$0.448\substack{+0.014\\-0.016}$
r	$0.0296\substack{+0.0006\\-0.0008}$
m_e/m_μ	$0.0048\substack{+0.0001\\-0.0002}$
$m_\mu/m_ au$	$0.0574\substack{+0.0032\\-0.0050}$
Ordering	NO
$J_{ m CP}$	$-0.018\substack{+0.002\\-0.002}$
$lpha_1/\pi$	0
α_2/π	$\pm 0.112^{+0.792}_{-0.014}$
$m_1 \; [{ m meV}]$	0
$m_2 [{ m meV}]$	$8.620\substack{+0.095\\-0.123}$
$m_3 [{ m meV}]$	$50.806\substack{+0.016\\-0.021}$
$\sum_i m_i [\text{eV}]$	$0.0594\substack{+0.0001\\-0.0001}$
$ m_{etaeta} $ [meV]	$3.61\substack{+0.09\\-0.09}$
$m_\beta^{\rm eff}~[{\rm meV}]$	$8.90\substack{+0.10 \\ -0.09}$
$d_{ m FT}$	3.03
$\chi^2_{ m min}$	0.98

Minimal seesaw model: arxiv:2402.18547 with S.Marciano, D.Meloni



Non-standard interactions



from Feruglio's slides at Mod. Symmetry Bethe Workshop