IRN Terascale 2024 - Laboratori Nazionali di Frascati

Probing dark matter models with DarkPACK: present state and future developments

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DARKPACK



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- Cross sections
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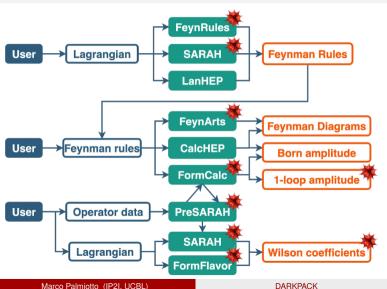
- Cross sections
- Decay rates
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maybe at 1 loop

To compute

- Relic density
- Direct and indirect detection observables

Some solutions



- Many codes are required
- Several passages of input
- Mathematica
 dependencies

DarkPACK is conceived to have a **unique** and **modular** workflow

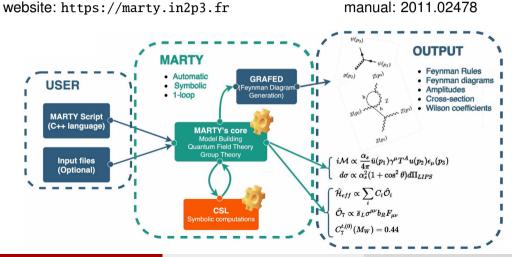
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Modular

- Possibility of stopping at any point of the chain...
- ... to link it with external software
- More ease in writing custom functionalities \leftarrow Object-oriented structure



Marco Palmiotto (IP2I, UCBL)

DARKPACK

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 - $\sum |M|^2$, Γ
 - Wilson coefficients
 - Feyman diagrams

With MARTY the user can

- Write a Lagrangian symbolically in a C++ source file
 - By defining the gauge symmetries of the model
 - By defining the fields of the model
 - By adding potential terms
 - By performing SSB if that's in the model
- Symbolically get quantities such as
 - $\sum |\overline{M}|^2, \Gamma$
 - Wilson coefficients

 \rightarrow up to 1 loop level

- Feyman diagrams
- Output those results in a **numerical** C++ library

DarkPACK and its documentation can be downloaded at

https://gitlab.in2p3.fr/darkpack/darkpack-public

(2211.10376 Palmiotto, Arbey, Mahmoudi)

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Why the MSSM?

- Numerical tests ← existence of many other tools
- Performance check \leftarrow lots of particles and Feynman rules

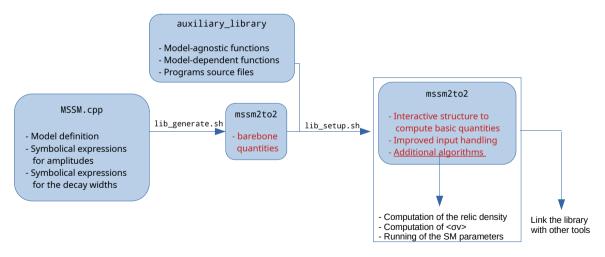
It relies on two script

- lib_generate.sh to generate the library
- lib_setup.sh to copy the files in auxiliary_library in the needed paths and to compile the final library

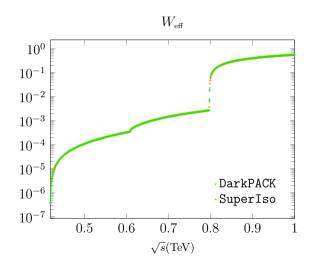
Detailed instructions on the scripts can be found in the README.md

You need to have MARTY installed, and define the environmental variable INSTALLMARTYPATH as the path where it is built

How it works



Some output



Capabilities

Observables:

- $\sum |M|^2$, $\Gamma \rightarrow$ up to 1-loop (LO)
- $W_{\text{eff}}, \langle \sigma v \rangle \rightarrow \text{improved stability at low } T$
- $\Omega h^2 \rightarrow$ from SuperIso Relic well-tested, reliable in MSSM, NMSSM

Implementation:

- user-friendly
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- native parallelisation ← avoiding global variables
 → good portability
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released MSSM

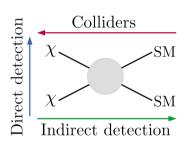
- performance
- consistency

release of new models

- stability
- ease of use

Development roadmap

- Releasing new models
- Improving the model-agnostic algorithms
- More general forms of the Boltzmann equation
 - \rightarrow Solving a system of equations: one for every species
 - \rightarrow Supporting models with multiple DM candidates
 - \rightarrow Considering more general scenarios, i.e. freeze-in
- Native functions for direct searches
 → MARTY provides Wilson coefficients
- Native functions for indirect searches
 - \rightarrow required amplitudes already provided
 - \rightarrow already possible to link it with external software
- Improving portability with UFO files
 - \rightarrow such a feature is among the future developments of <code>MARTY</code>



Conclusions

Today DarkPACK allows to

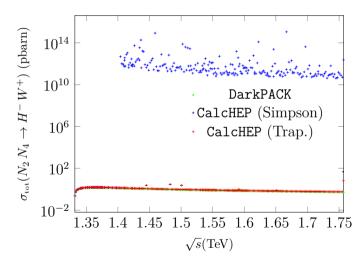
- Compute $\sum |M|^2$ and Γ at LO in many NP scenarios
- Compute $\langle \sigma v \rangle$, Ωh^2 for coannihilation
- Have a library easy to link with other software
- Have a framework portable and performance-oriented
- \rightarrow validated in the MSSM

Currently, we are working on the implementation of a new model Next, we will

- Release the source code for the new model
- Follow the development roadmap
- use DarkPack to verify if specific NP models can help to explain DM observables

Thank you for the attention!

Simpson rule vs trapezoidal rule pt. 1



Simpson rule vs trapezoidal rule pt. 2

