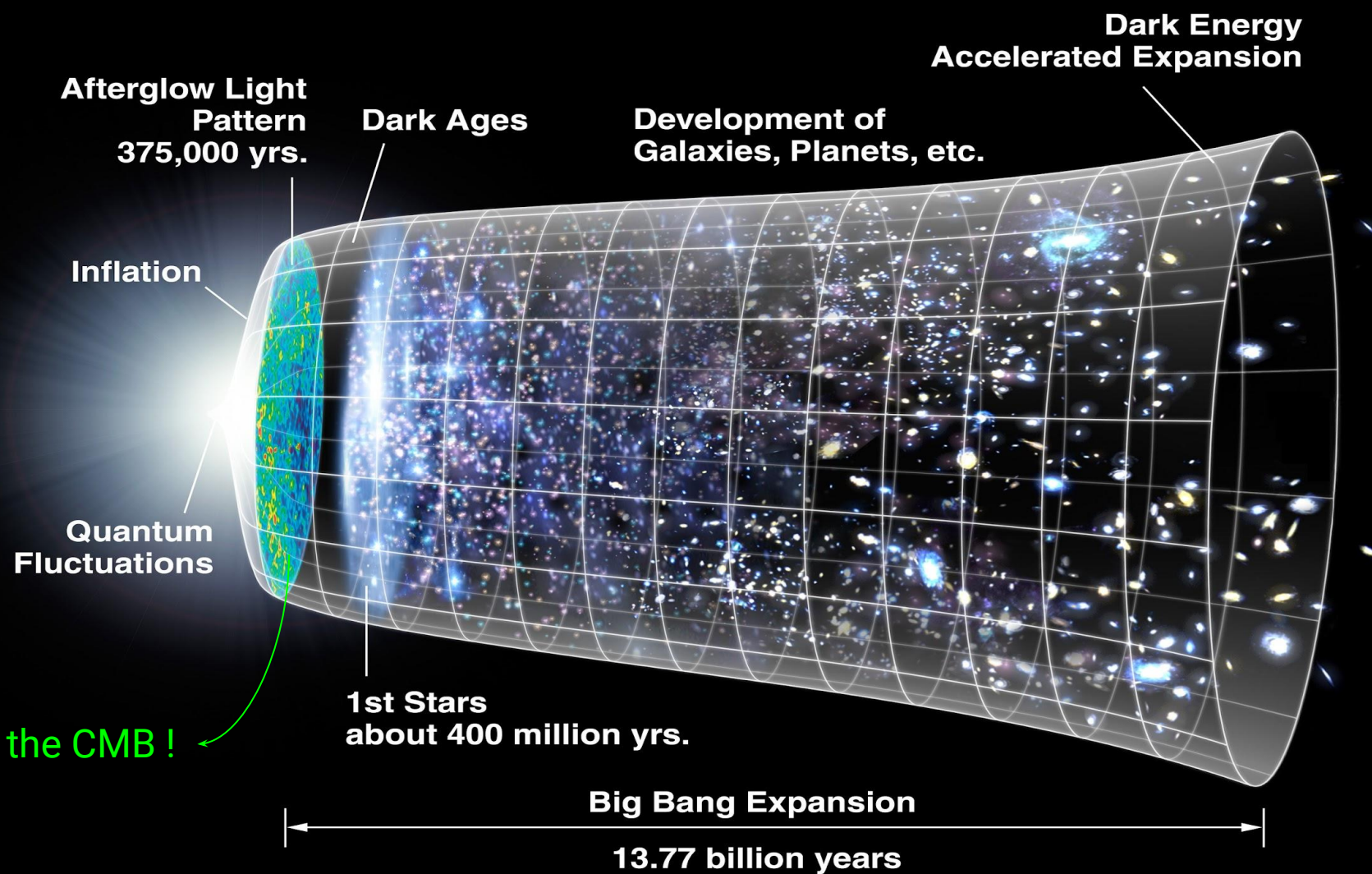


Map-making of the polarized Cosmic
Microwave Background with next-generation
ground-based observatories, Simons
Observatory and CMB-S4

CDD 2024
Simon Biquard

*Credit:
ESA / Planck
collaboration*

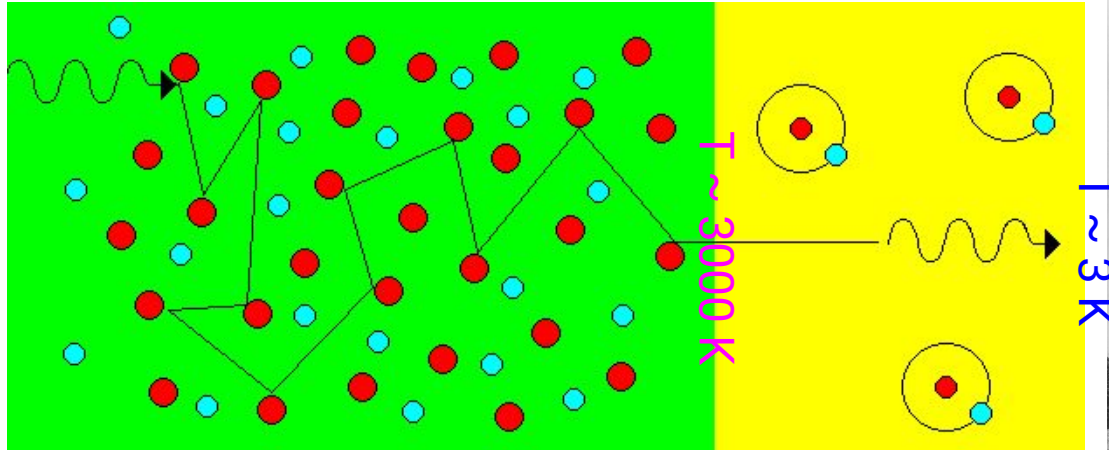
Cosmic



Background (+ microwave)

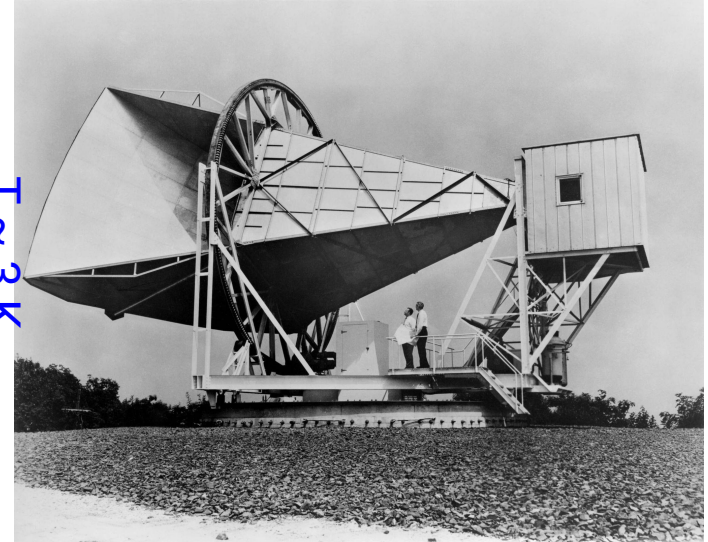
Origin of the CMB

hot opaque $\xrightarrow{\text{recombination}}$ colder transparent



credit: J. Schombert ([website](#))

1940s : prediction (Alpher, Herman)
1964 : detection (Penzias, Wilson)
1978 : Nobel prize

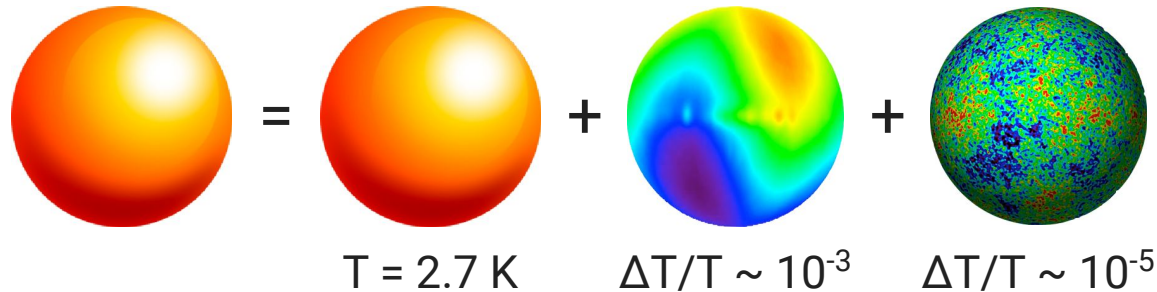


CMB anisotropies

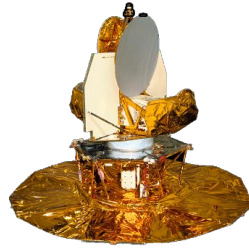
Universe is **not homogeneous** on small scales (galaxies, clusters, ...)

Today's inhomogeneities result from tiny fluctuations in the very early U.

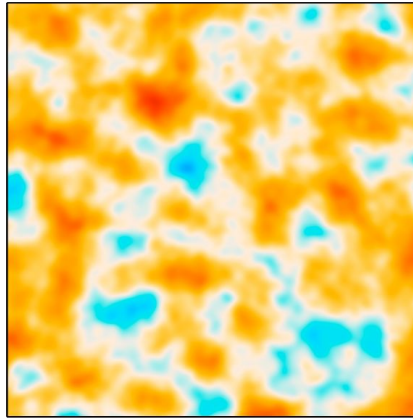
These fluctuations are visible on the CMB surface!



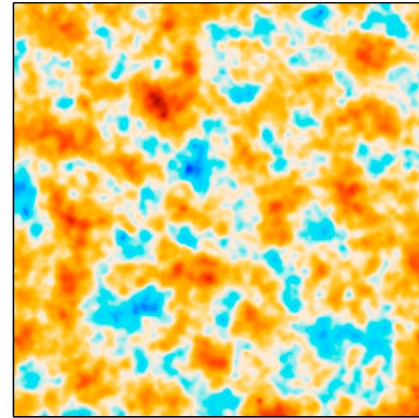
Credits: Josquin Errard



COBE



WMAP



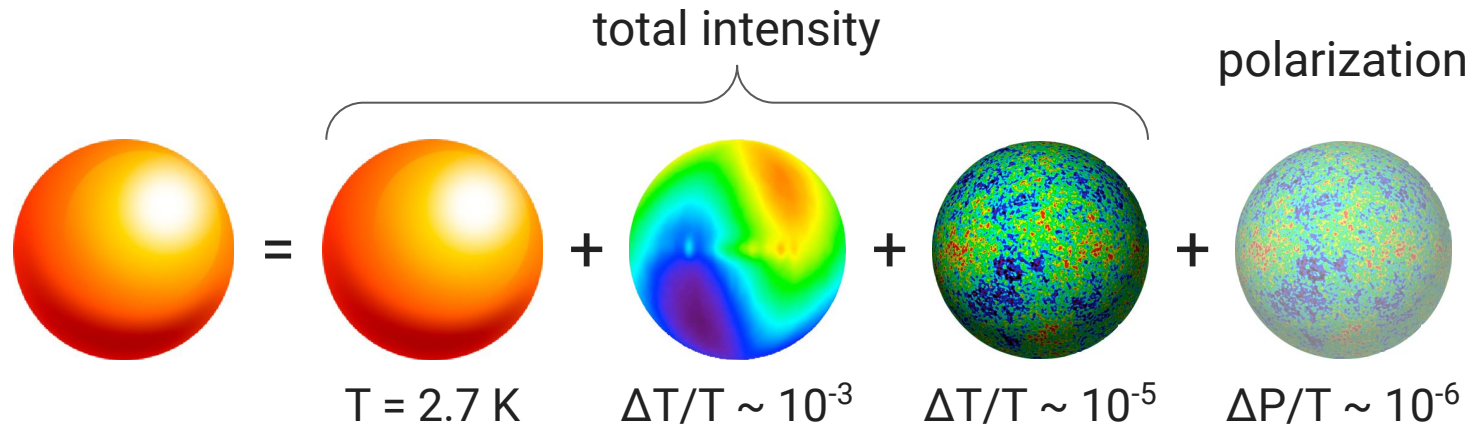
Planck

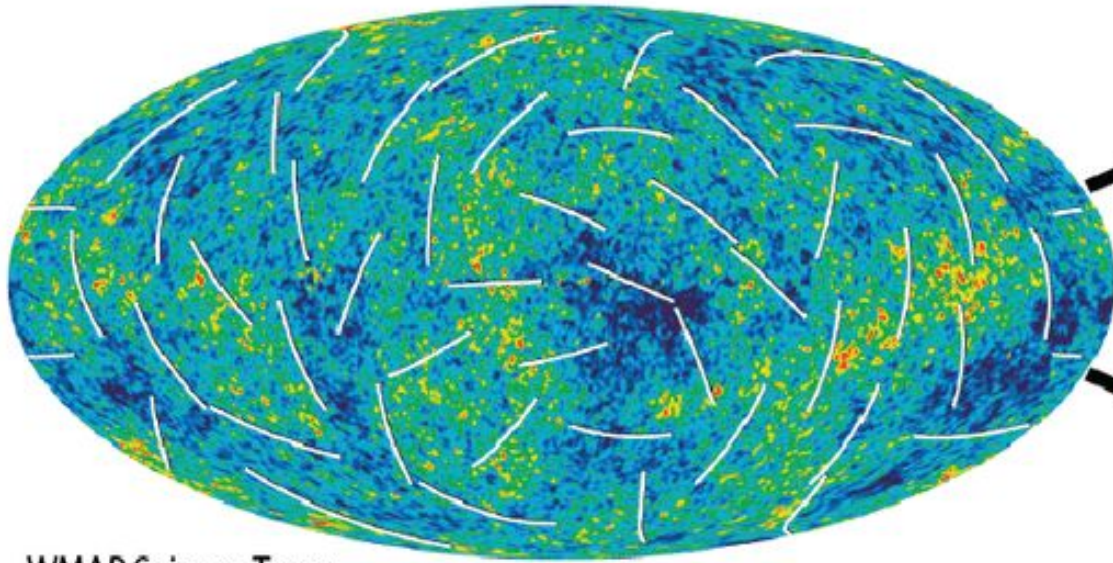
Increased angular resolution with successive generations of instruments – Credit: NASA/JPL-Caltech/ESA

Polarization

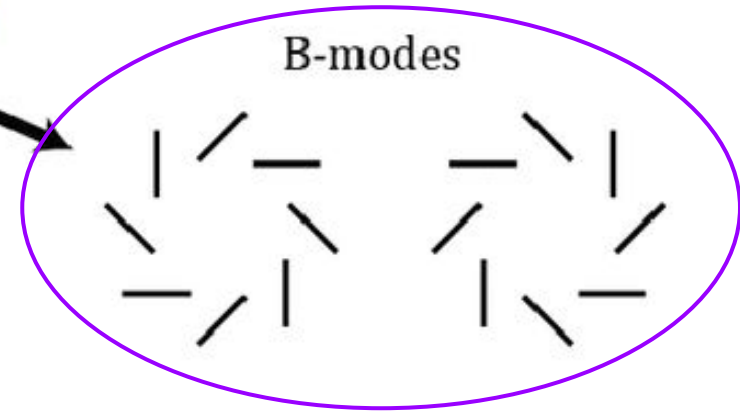
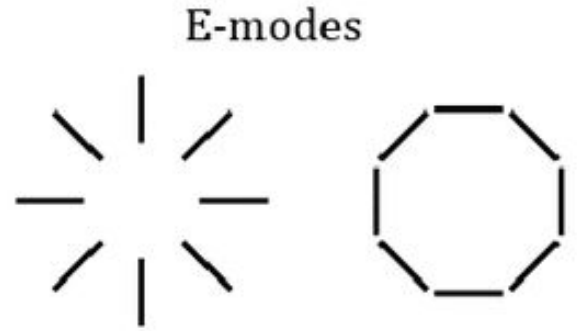
What about polarization?

Thomson scattering + inhomogeneities = CMB photons are polarized





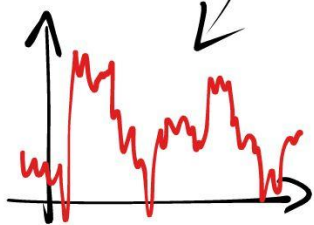
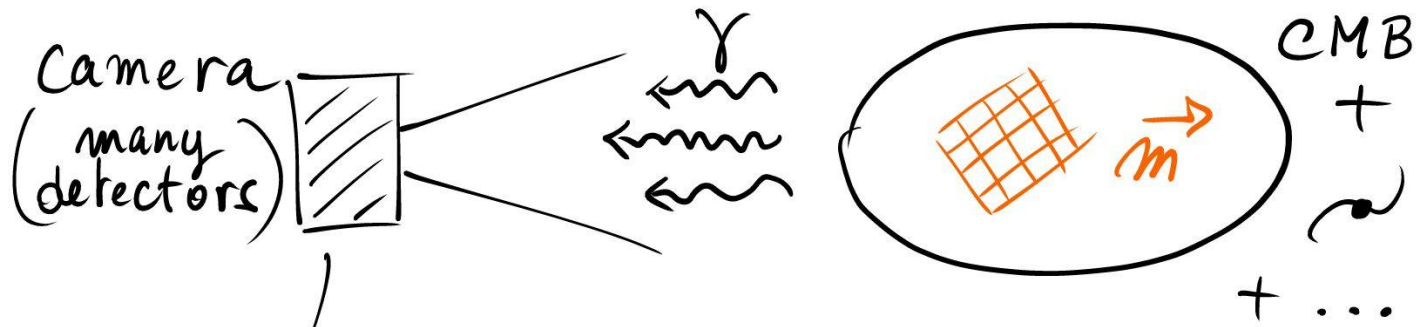
WMAP Science Team



Potential trace of *primordial gravitational waves* !

E- and B-mode patterns on the CMB polarization map
Credit: WMAP collaboration

Map-making



time-ordered data (> petabyte!)

$$\vec{d} = P \vec{m} + \vec{n}$$

reconstruct observed sky + characterize statistical properties

$$\hat{\vec{m}} = ??? \times \vec{d}$$

Next-generation observatories

Simons Observatory

Atacama desert (Chile)

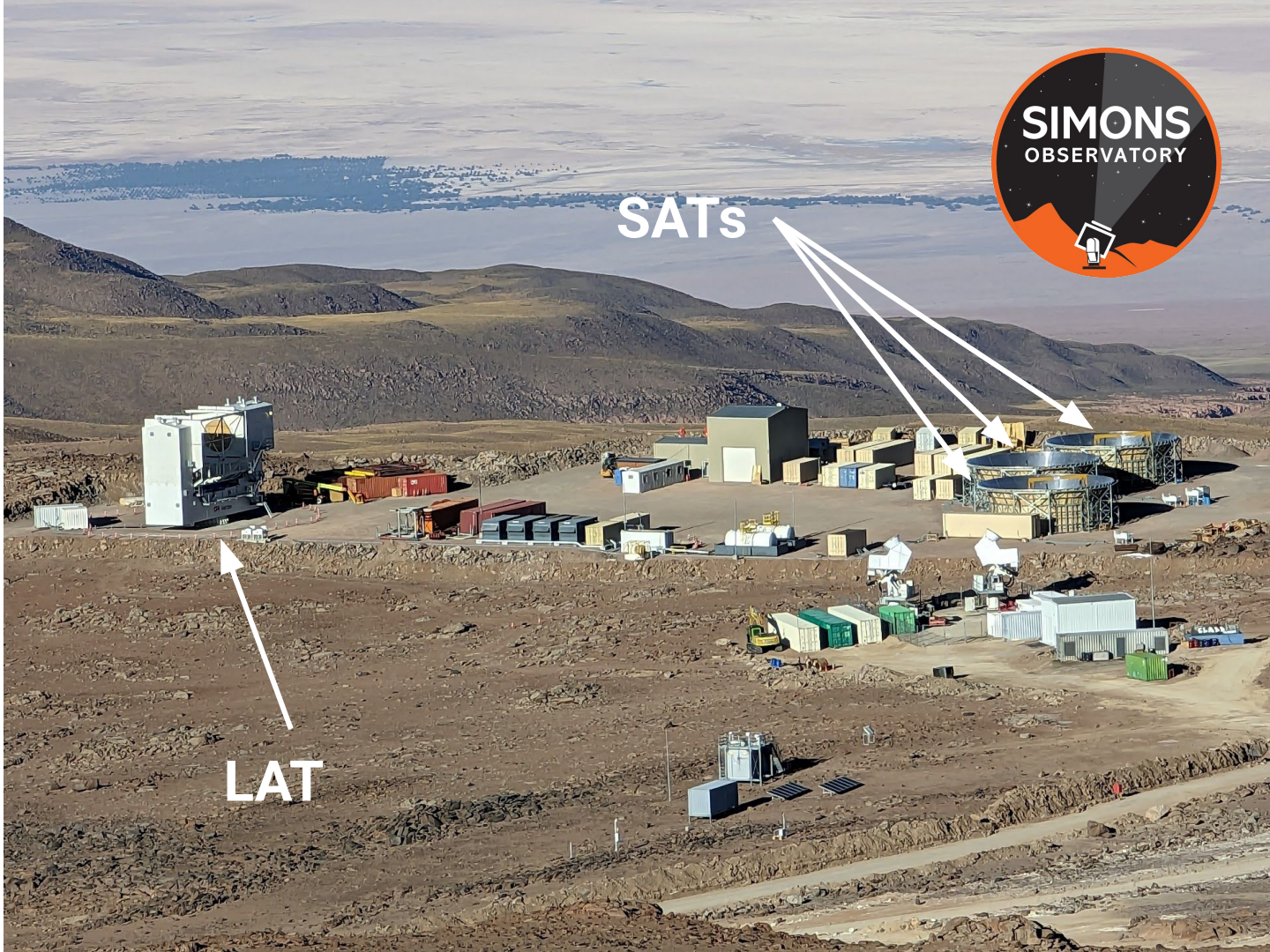
5200 m above sea level

3 small (0.4 m) telescopes

1 large (6 m) telescope

60,000 TES detectors

Many science cases !



SATs

LAT

Questions ?