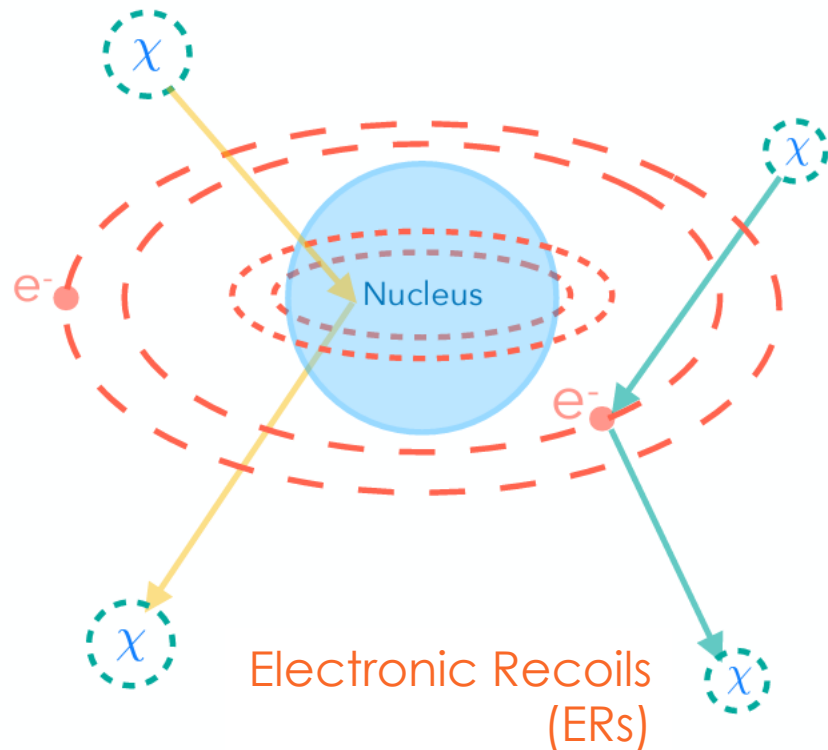


RECHERCHE DIRECTE VIA LES DÉTECTEURS À GAZ NOBLES

SARA DIGLIO, SUBATECH-NANTES

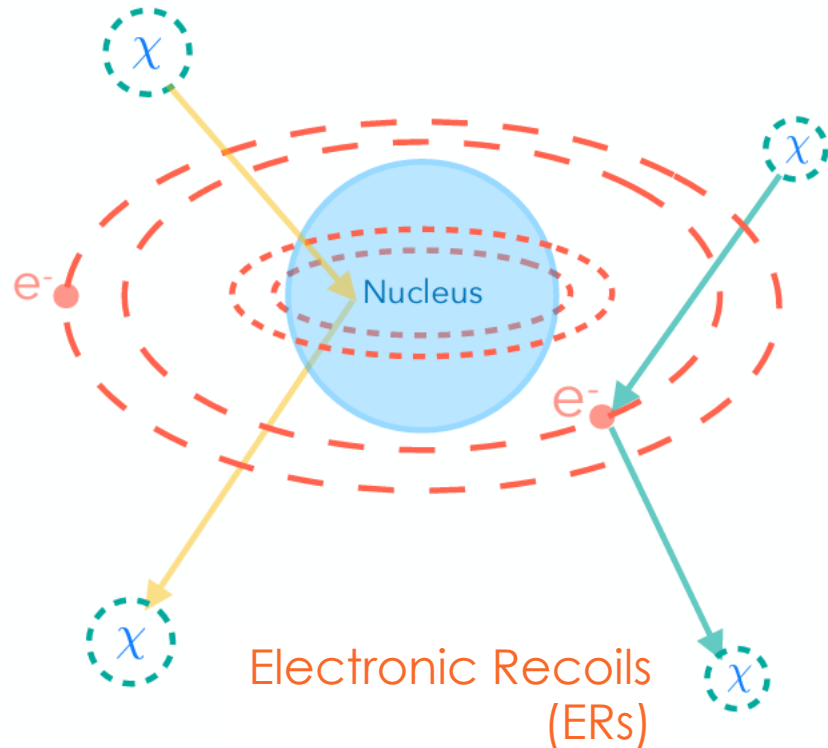
HOW TO DETECT A DM CANDIDATE ?

Nuclear recoils (NRs)

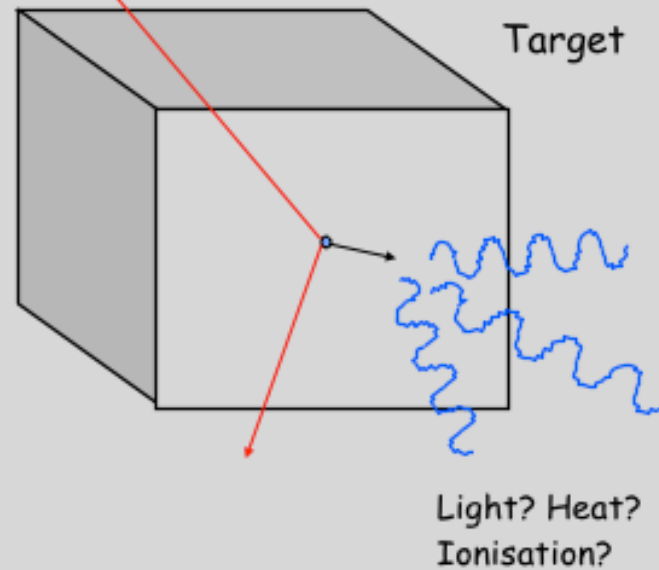


HOW TO DETECT A DM CANDIDATE ?

Nuclear recoils (NRs)

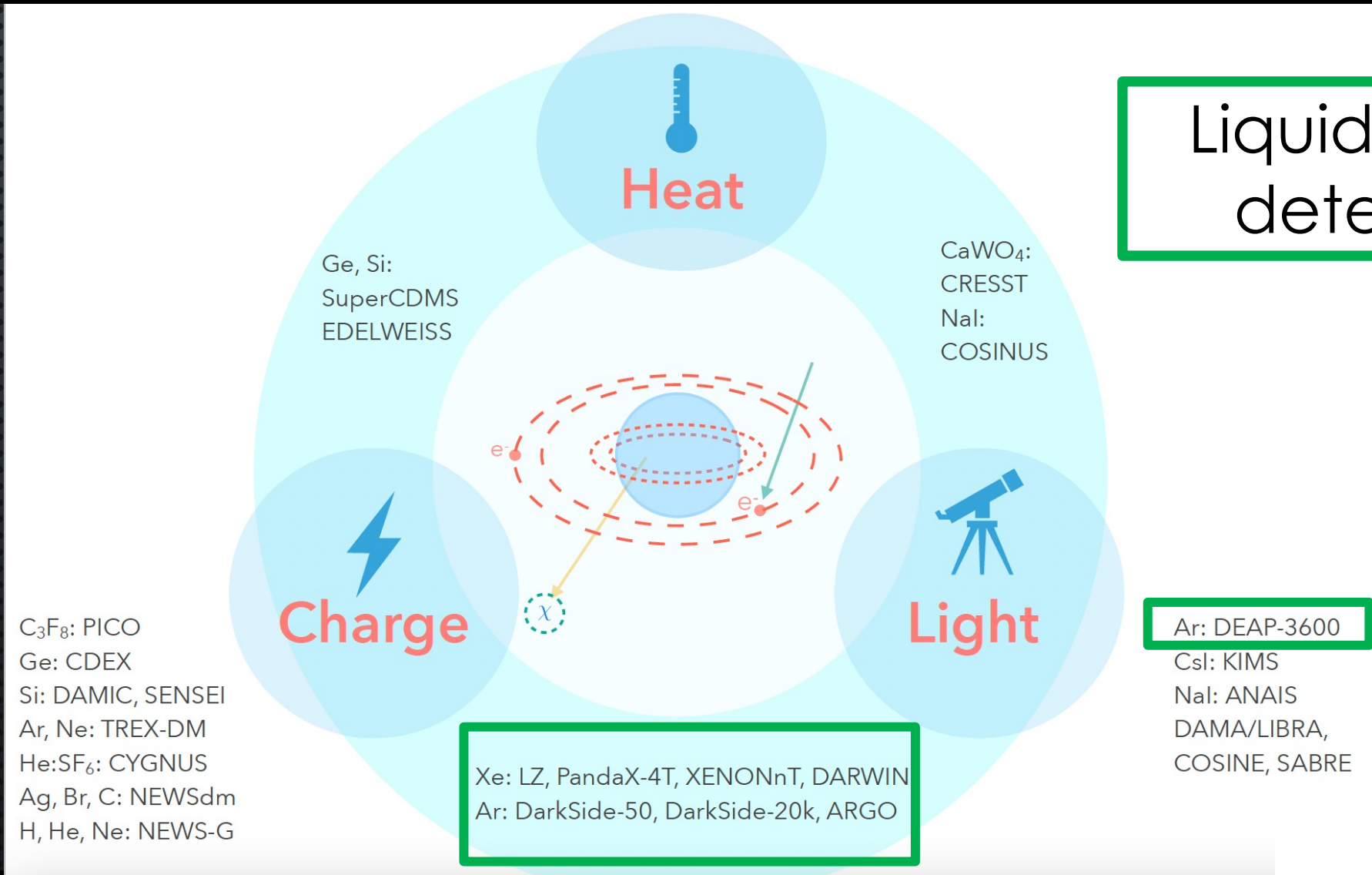


Particle (Dark Matter or Bkg)



A simple particle detector...

DIRECT DETECTION TECHNIQUES



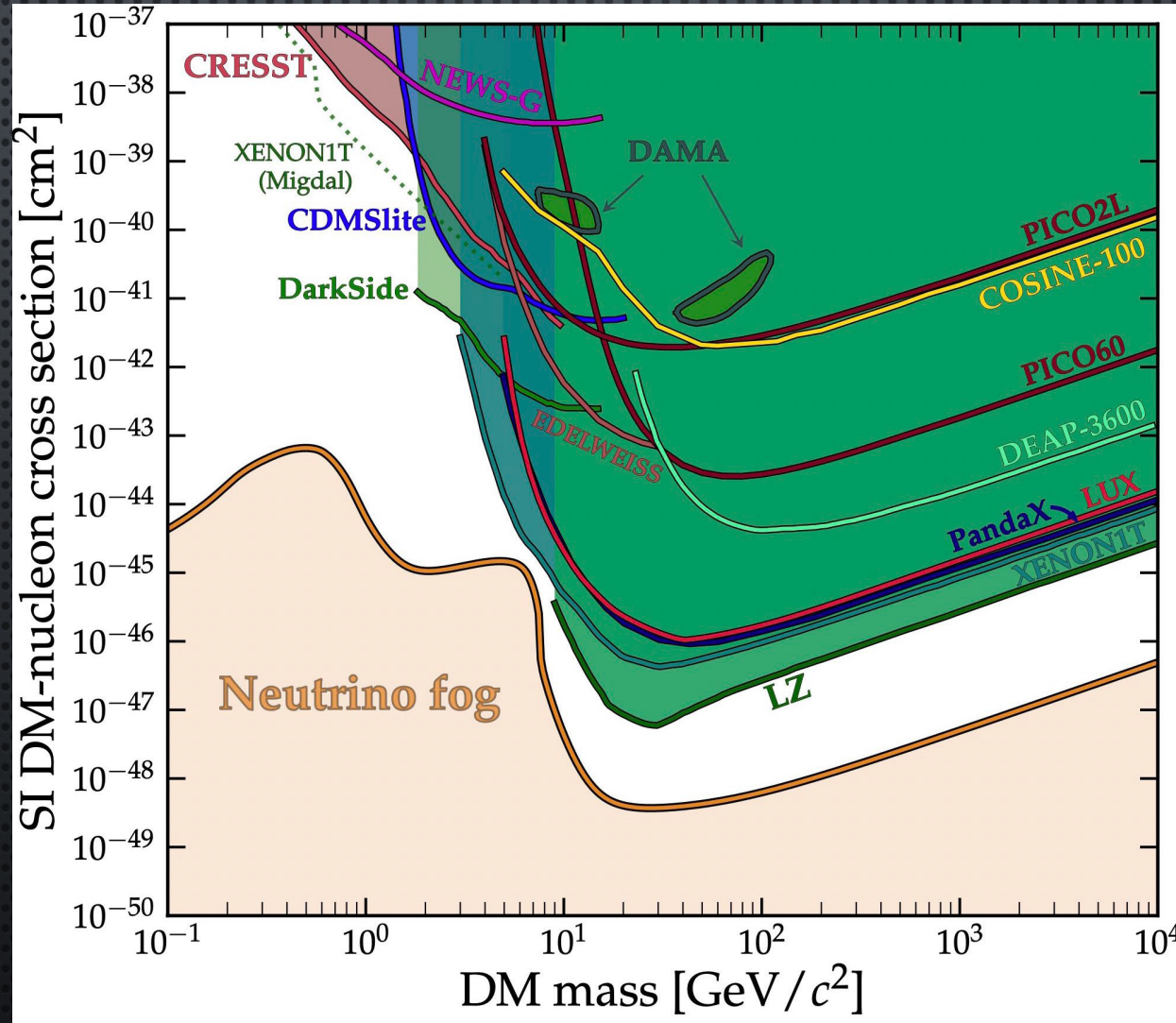
Liquid Noble detectors

Ar: DEAP-3600

CsI: KIMS
NaI: ANAIS
DAMA/LIBRA,
COSINE, SABRE

Credit: L.Baudis

DARK MATTER DIRECT DETECTION STATE OF THE ART



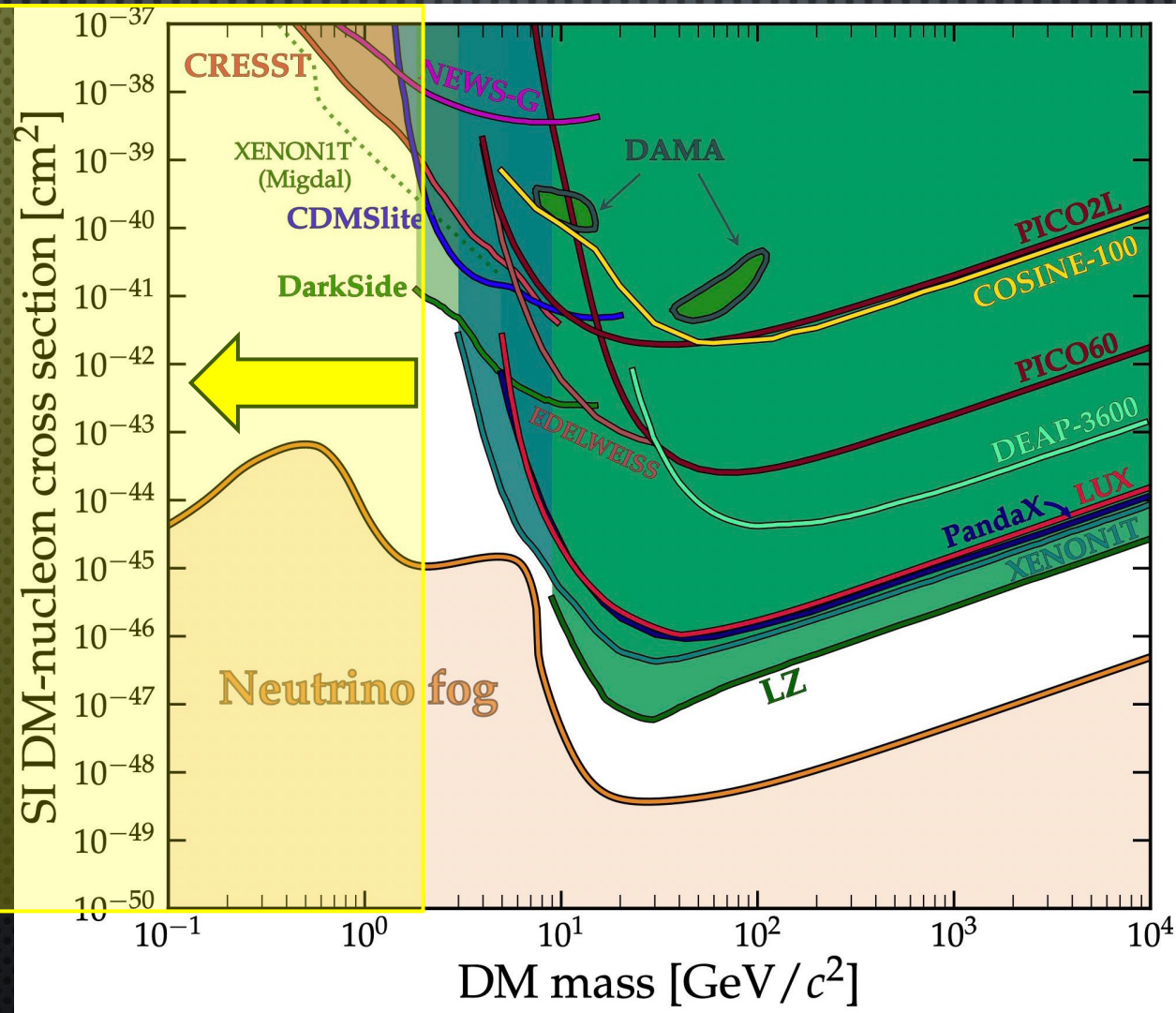
DARK MATTER DIRECT DETECTION STATE OF THE ART

Low Mass Region
 $m_{\text{DM}} < \sim 1 \text{ GeV}$

Region dominated by
Cryogenic detectors

Low energy threshold

Talk by V. Novati



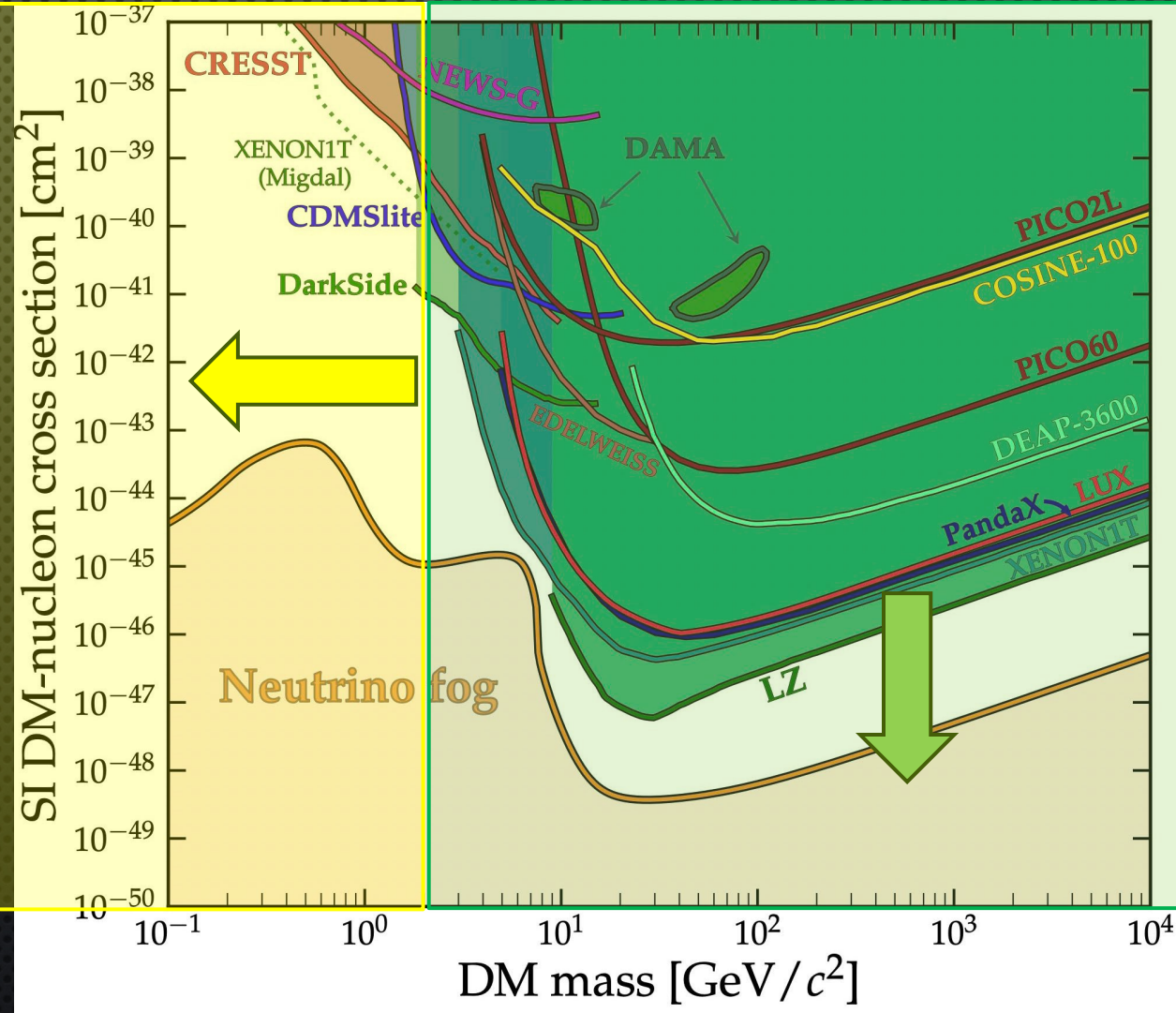
DARK MATTER DIRECT DETECTION STATE OF THE ART

Low Mass Region
 $m_{\text{DM}} < \sim 1 \text{ GeV}$

Region dominated by
Cryogenic detectors

Low energy threshold

Talk by V. Novati



High Mass Region
 $m_{\text{DM}} > 1 \text{ GeV}$

Region dominated by
**Noble liquids
detector**

Large mass detectors

This Talk

LIQUID NOBLE TECHNOLOGY

WIMP Limits vs Time: principal detector categories

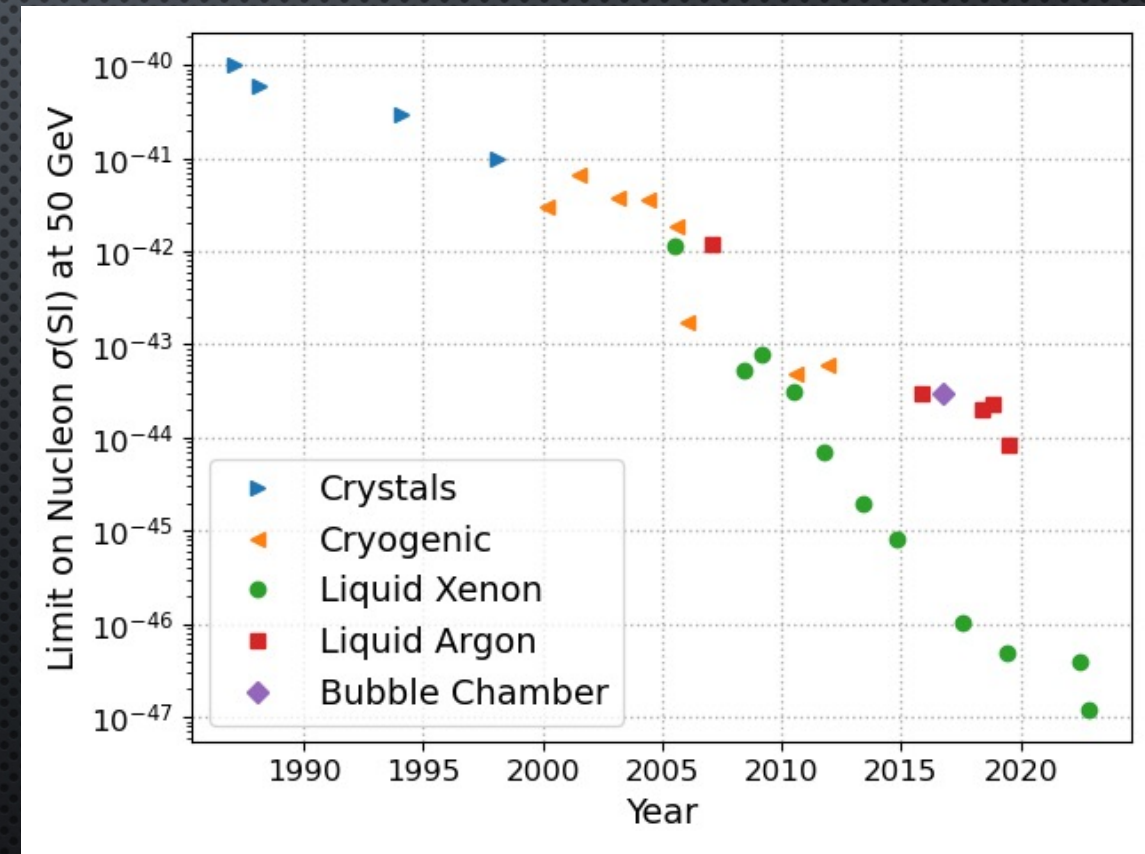


Figure by Tom Shutt, SLAC

LIQUID NOBLE TECHNOLOGY

- Leading sensitivity at intermediate/high DM masses
- **Scalability** → large target masses
- **High Density** → Self shielding
- **Readily purified** → ultra-low bkg
- **Easy cryogenic** : 170 K (LXe), 87 K (LAr)
- **SI and SD** (^{129}Xe , ^{131}Xe) interactions
- Many other science opportunities:
 - Double weak decays
 - Solar and SN ν
 - ...

WIMP Limits vs Time: principal detector categories

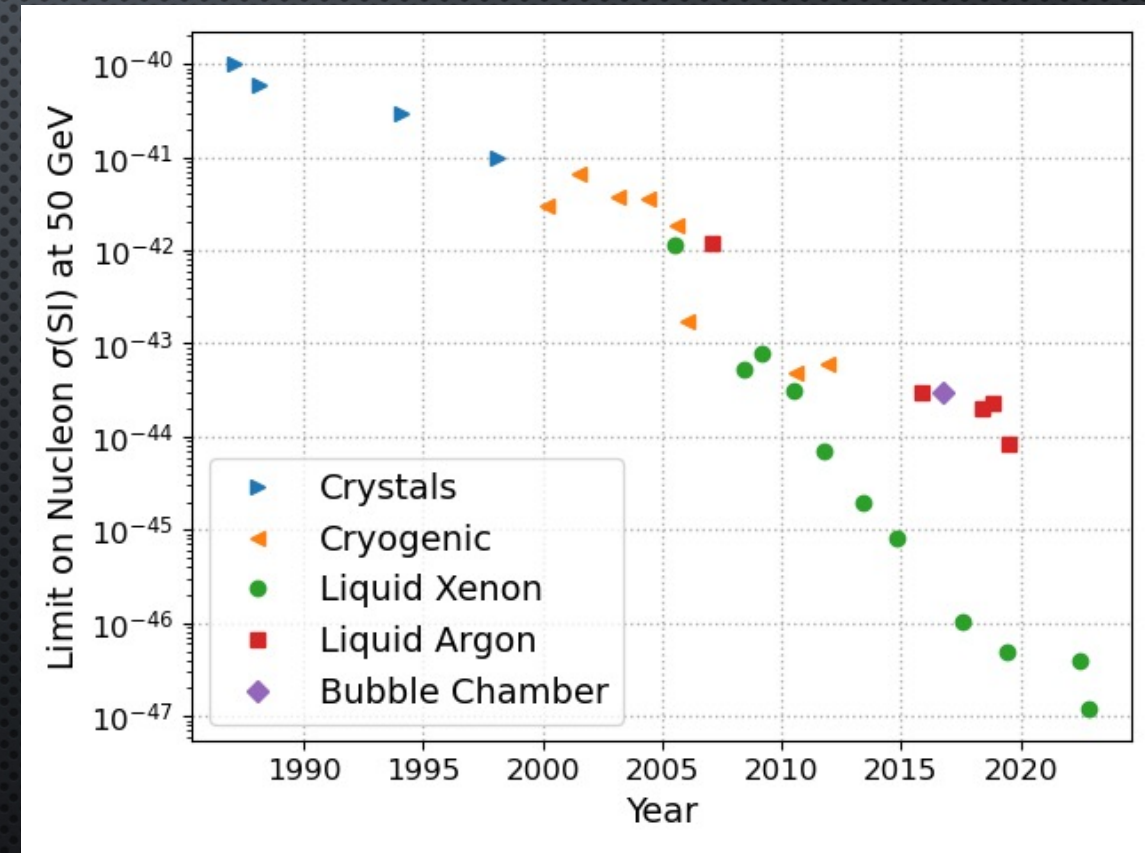
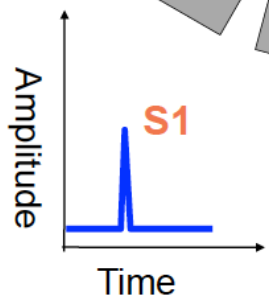
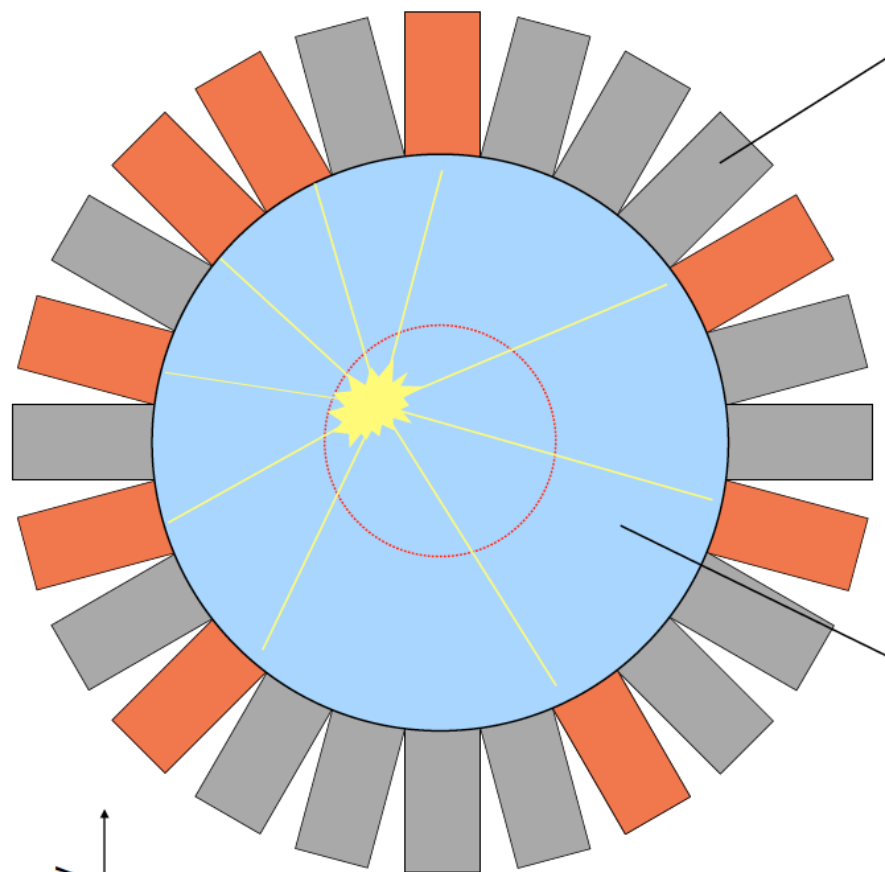


Figure by Tom Shutt, SLAC

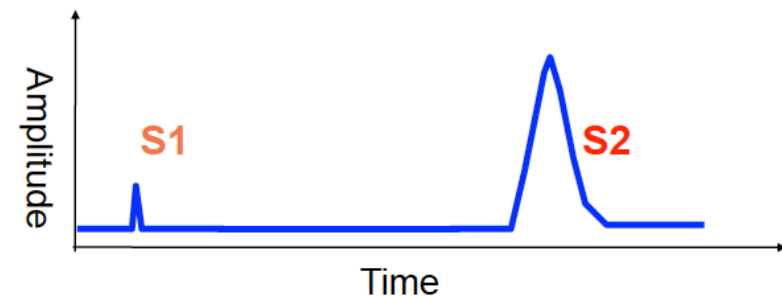
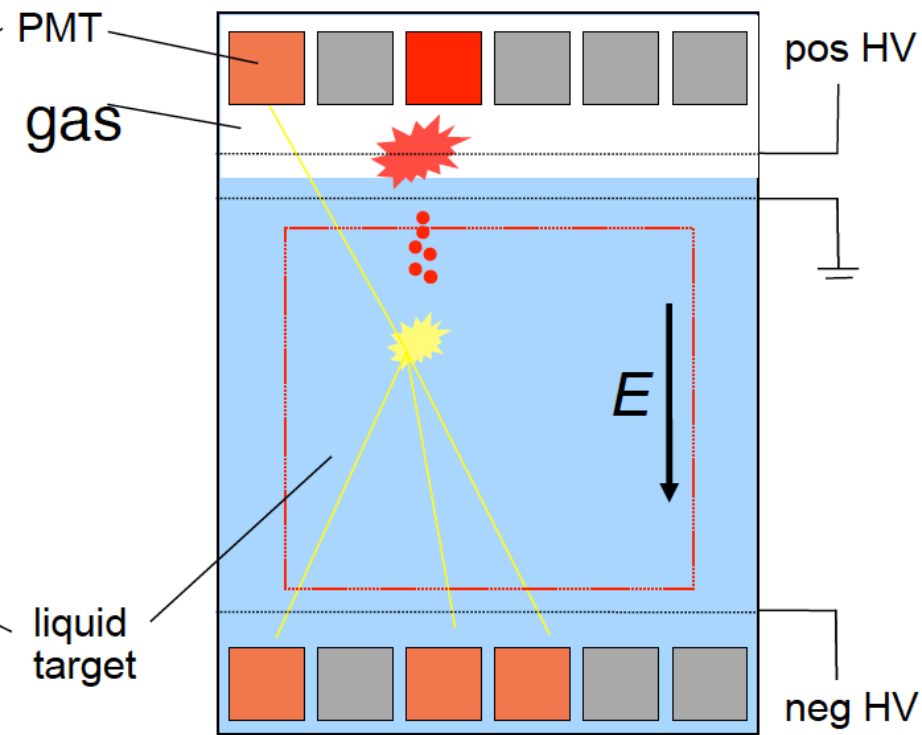
DETECTOR TYPES

Single Phase

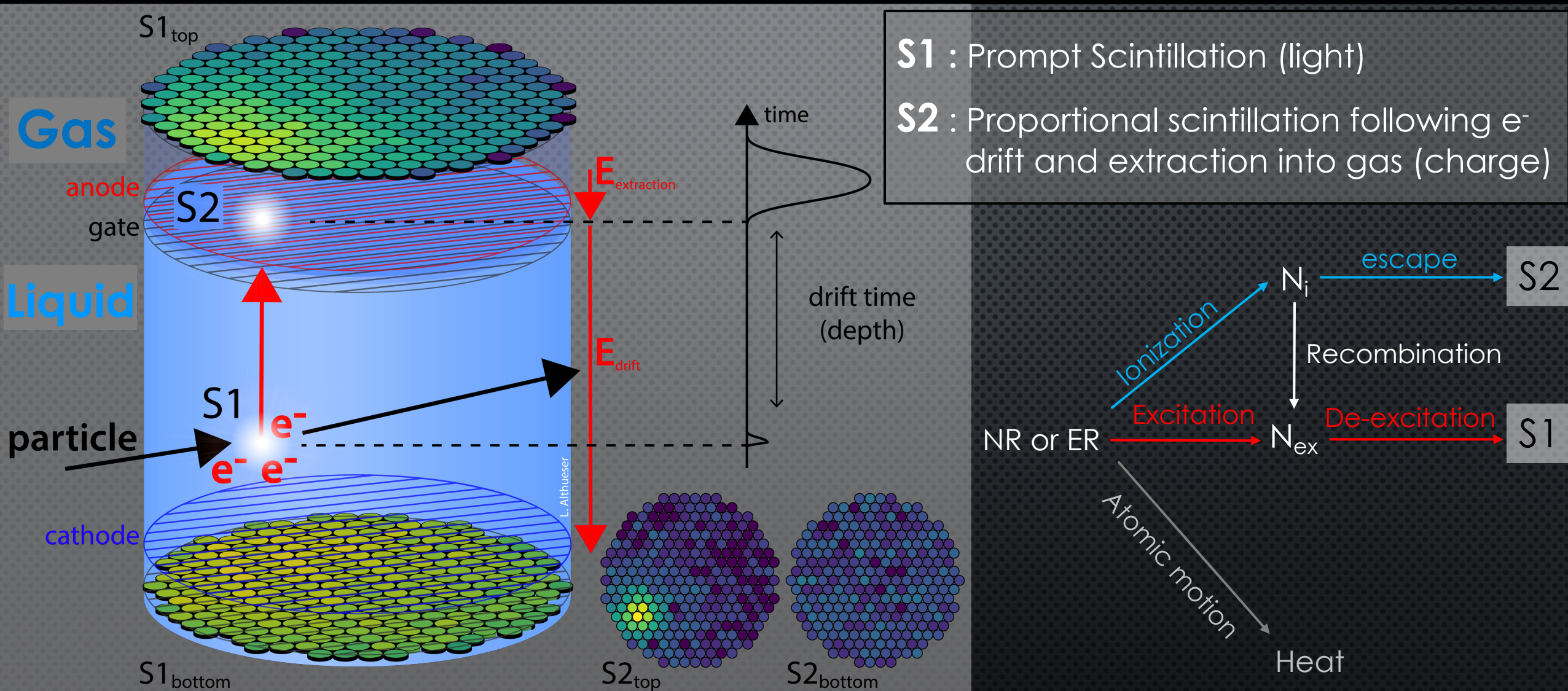


- High light yields,
- simple geometry
- no E-fields

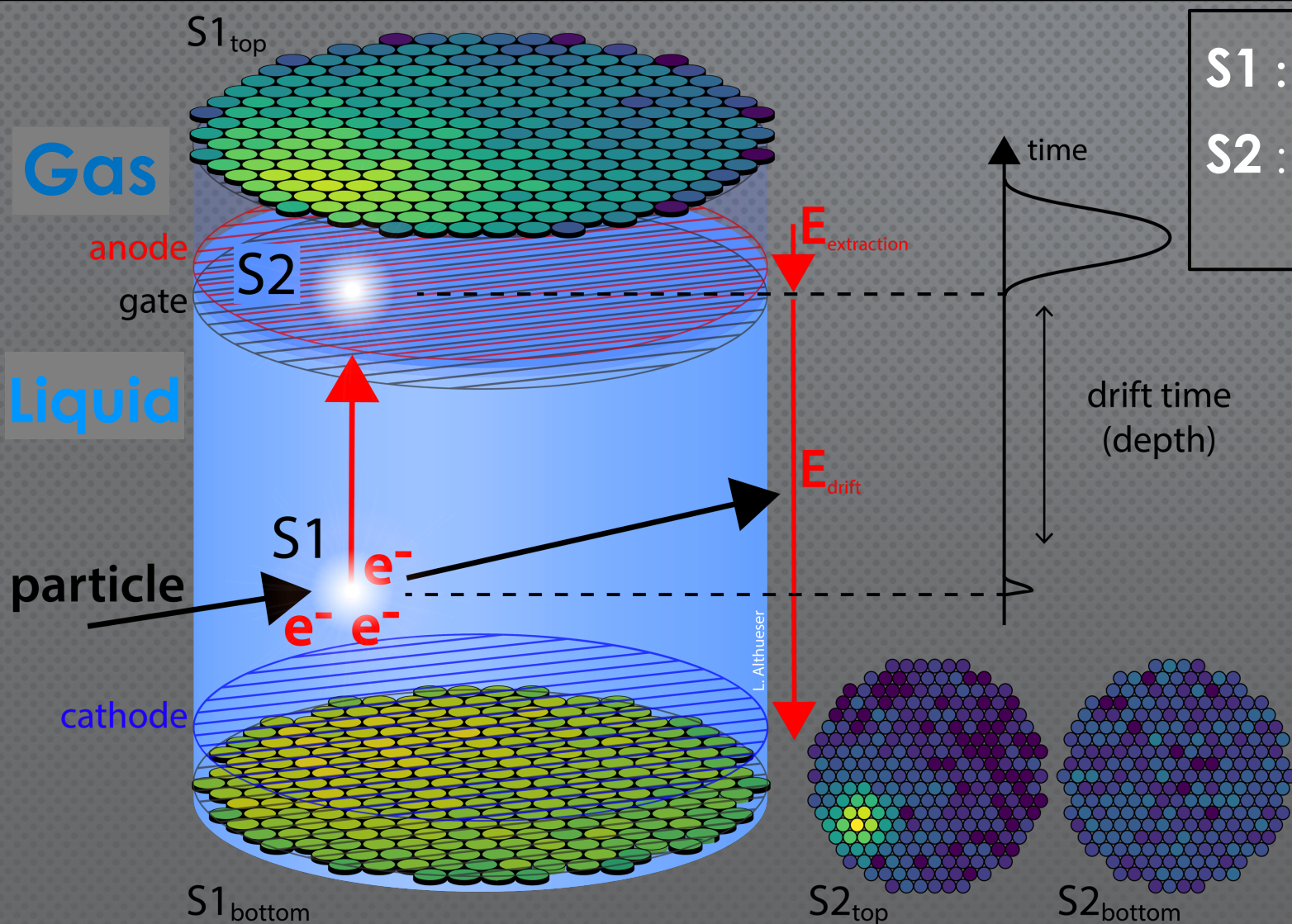
Dual Phase



DUAL PHASE TIME PROJECTION CHAMBERS

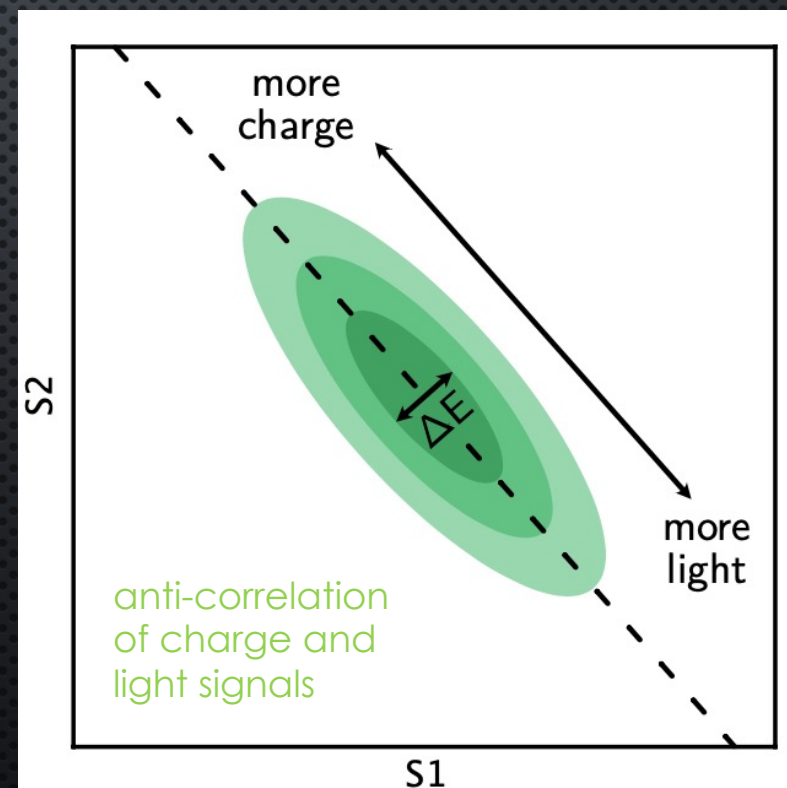


DUAL PHASE TIME PROJECTION CHAMBERS

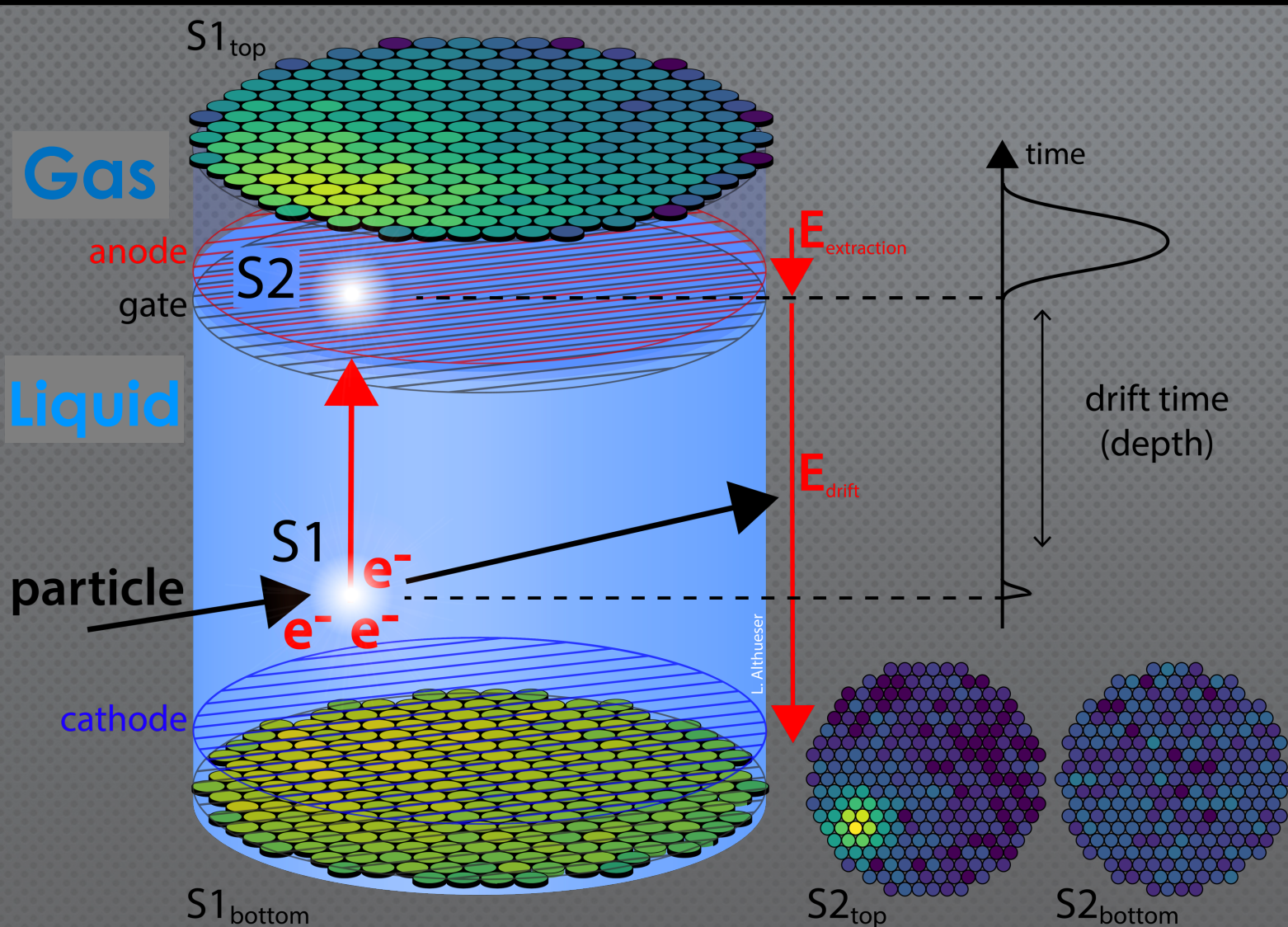


S1 : Prompt Scintillation (light)

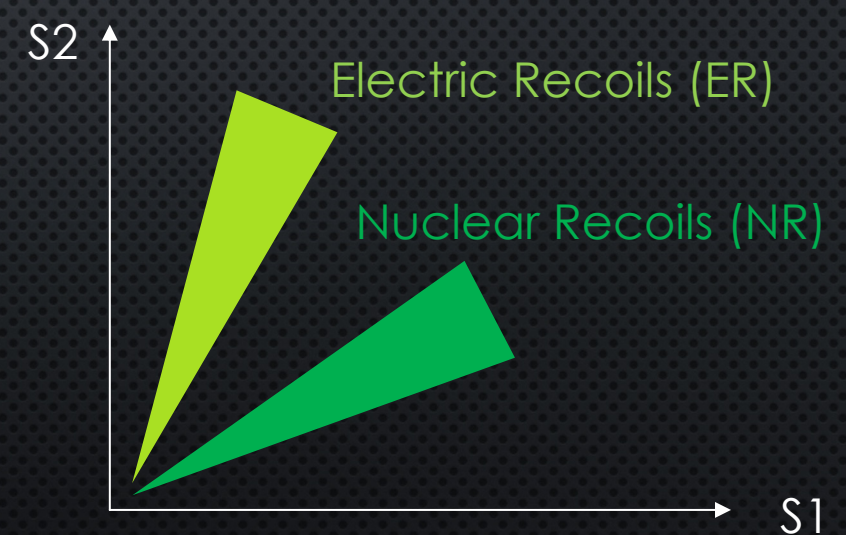
S2 : Proportional scintillation following e^- drift and extraction into gas (charge)



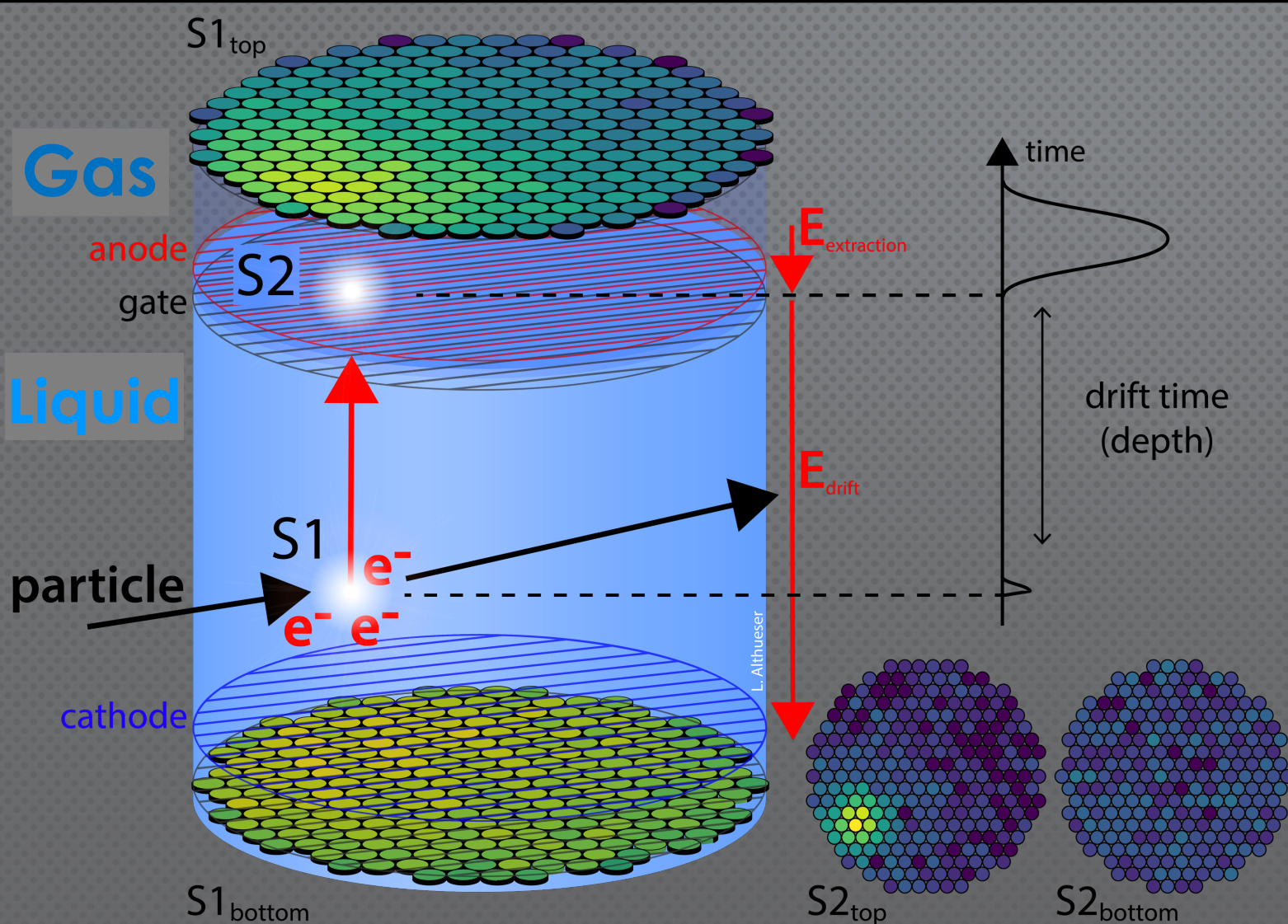
DUAL PHASE TIME PROJECTION CHAMBERS



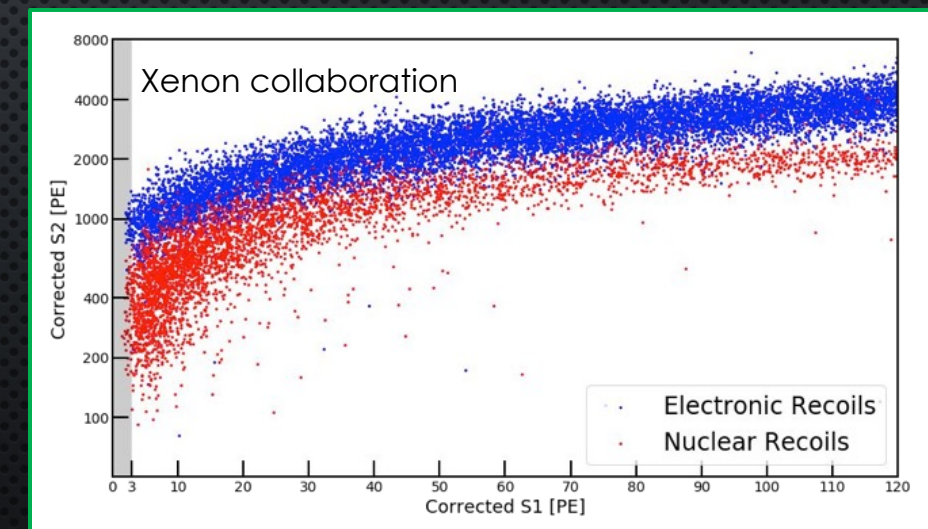
- Energy from S1 and S2
- Separation ER vs NR
- 3D event reconstruction:
 - X, Y from S2 hit pattern on top PMTs
 - Z from electrons drift time



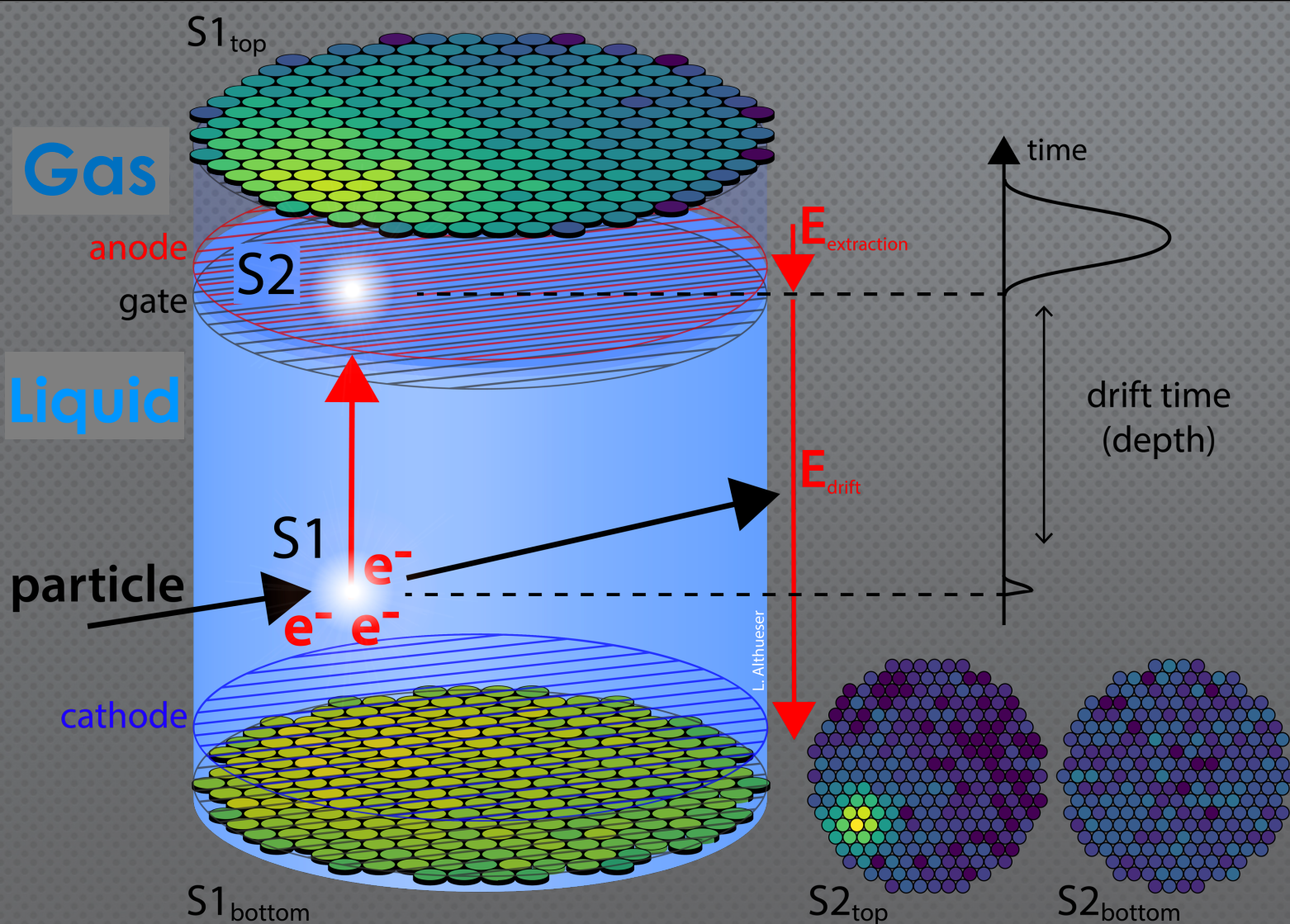
DUAL PHASE TIME PROJECTION CHAMBERS



- Energy from S1 and S2
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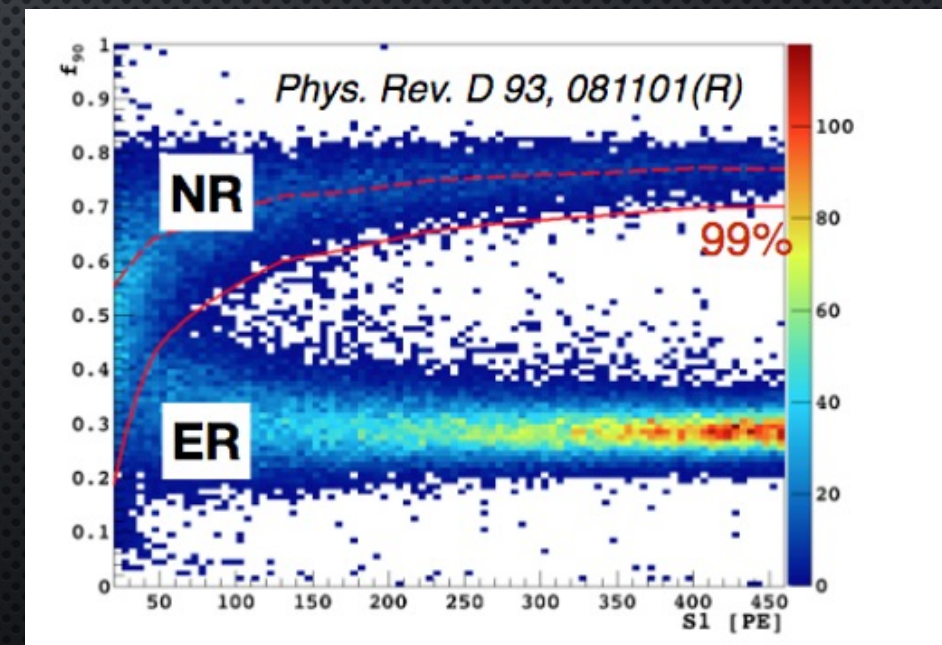


DUAL PHASE TIME PROJECTION CHAMBERS



Additional ER/NR discrimination in LAr TPCs

- Pulse Shape Discrimination (PSD) parameter f_{90} : fraction of light seen in the first 90 ns



PROPERTIES OF NOBLE GASES

Complementarity between LXe & LAr

- Higher mass number in **Xe** → higher WIMP rate
- Better ER/NR discrimination in **Ar**
- Higher density in **Xe** → stronger self-shielding
- ...

Noble Gas

LXe

LAr

Atomic mass [g/mol]

131.3

39.95

Density [g/cm³]

3.06

1.40

Wavelength [nm]

178

128

Average ionization energy W [eV]

15.6

23.3

Ionization Yield [e-/keV]

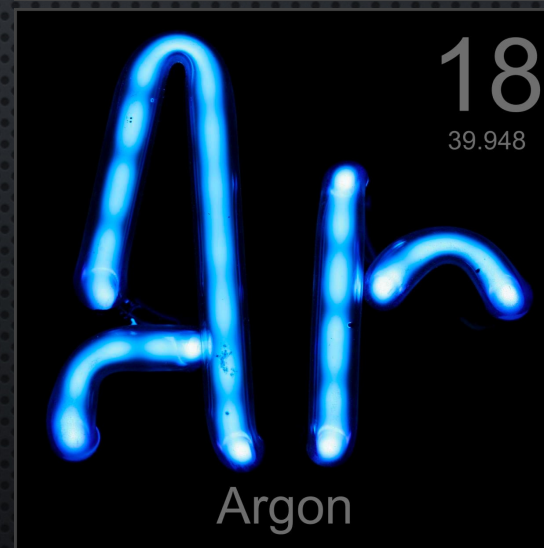
64

42

Scintillation Yield [photon/keV]

46

40



CURRENT GENERATION LIQUID NOBLE DETECTORS

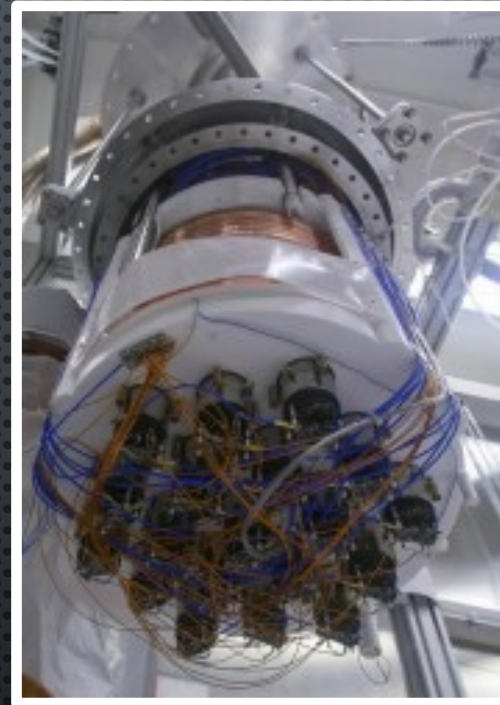
LIQUID ARGON DETECTORS

○ **DarkSide50**

- Target: 50 kg (fiducial 31 kg) (depleted in Ar39)
- 38 PMTs
- Data taking: 2013 - 2019

○ **DEAP 3600**

- Target : 3.3 t (fiducial 1 t)
- 255 PMTs
- Data taking: 2016 - 2020



DarkSide50 @LNGS
Dual phase



DEAP3600 @Snolab
(Canada)
Single phase

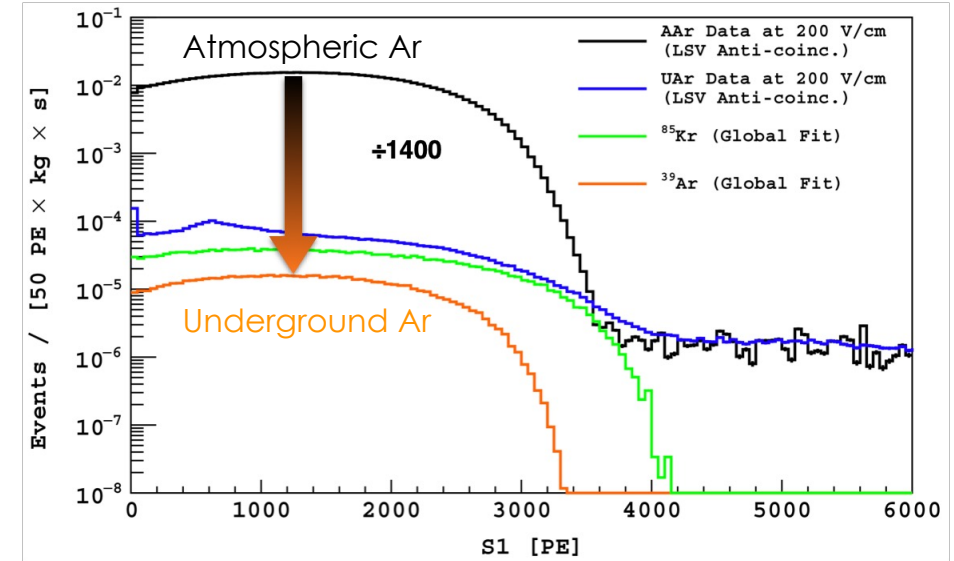
LIQUID ARGON DETECTORS

○ DarkSide50

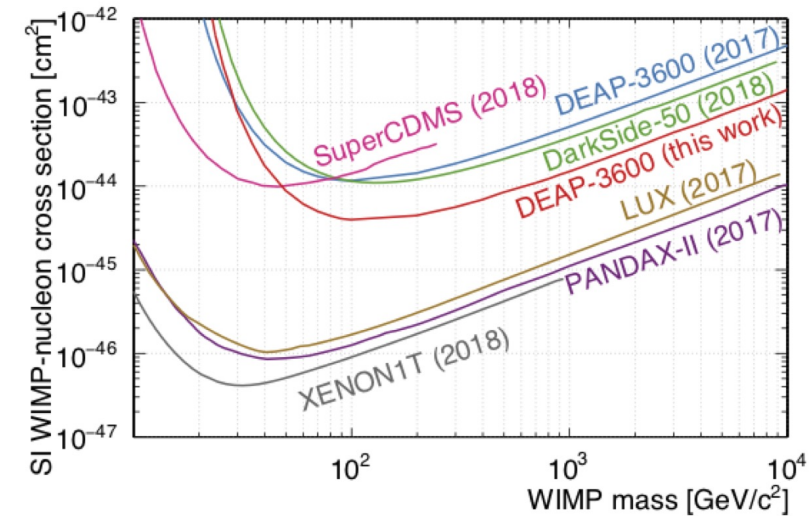
- Target: 50 kg (fiducial 31 kg) (depleted in Ar39)
- 38 PMTs
- Data taking: 2013 - 2019
- Exposure: 532 live-days (0.045 t x year)

○ DEAP 3600

- Target : 3.3 t (fiducial 1 t)
- 255 PMTs
- Data taking: 2016 - 2020
- Exposure: 231 live-days (2019) (0.63 t x year)
- Detector upgrade in progress



Phys. Rev. D. (2018) 98: 102006



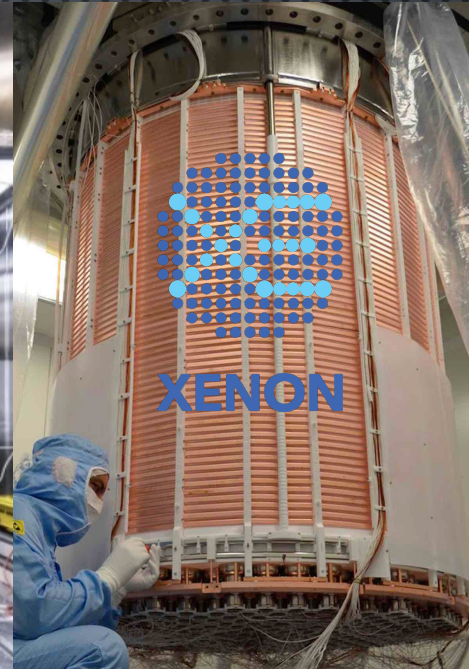
Phys. Rev. D. (2019) 100: 022004

LIQUID XENON DETECTORS

- **XENONnT**
 - Target: 5.9 t (fiducial 4.2 t)
 - Data taking since 2021
- **Lux-Zeplin (LZ)**
 - Target : 7 t (fiducial 5.5 t)
 - Data taking since 2022
- **PandaX-4T**
 - Target: 3.7 t (fiducial 2.7 t)
 - Data taking since 2021



LZ @SURF (US)
Dual phase



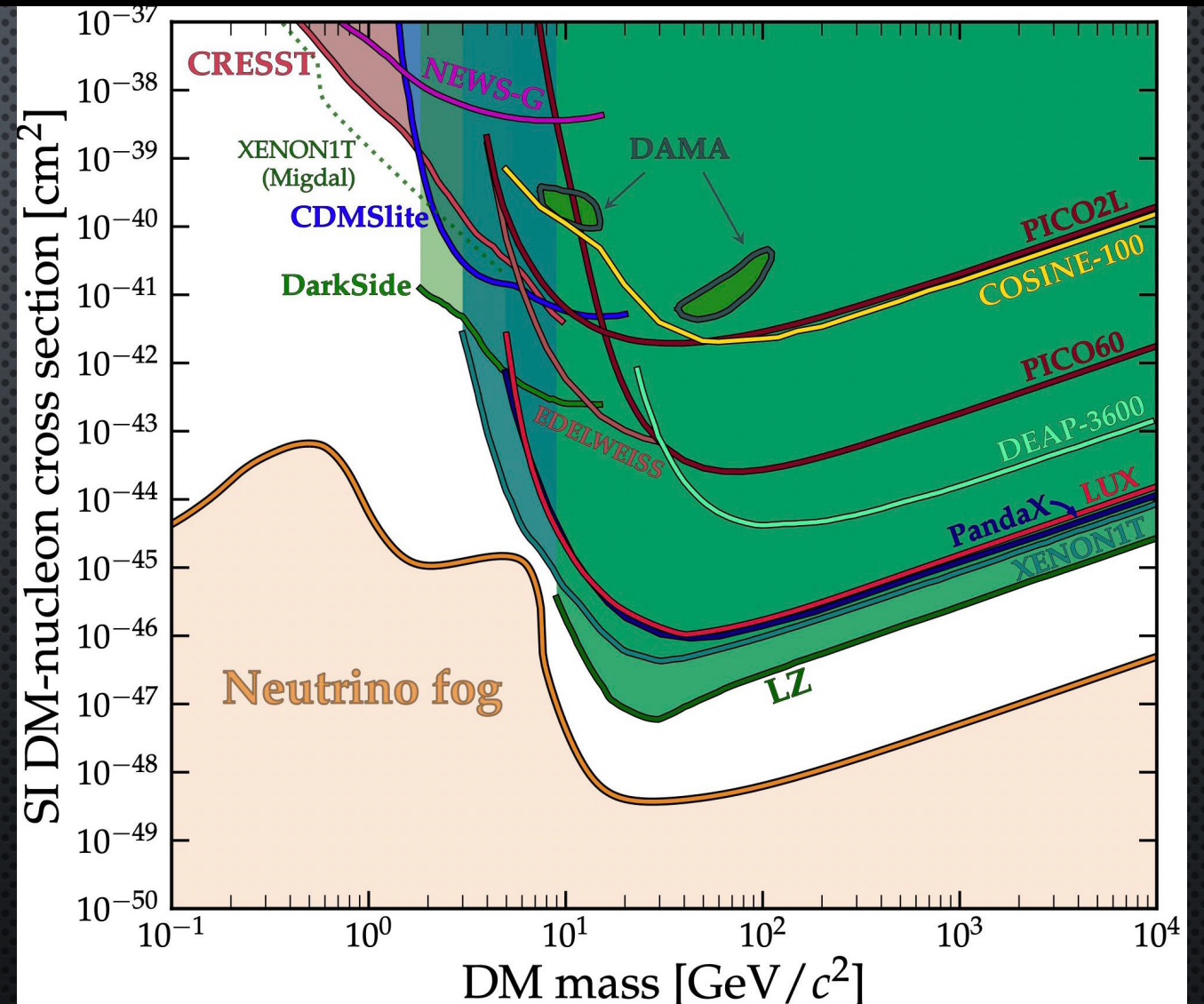
XENONnT @LNGS
Dual phase



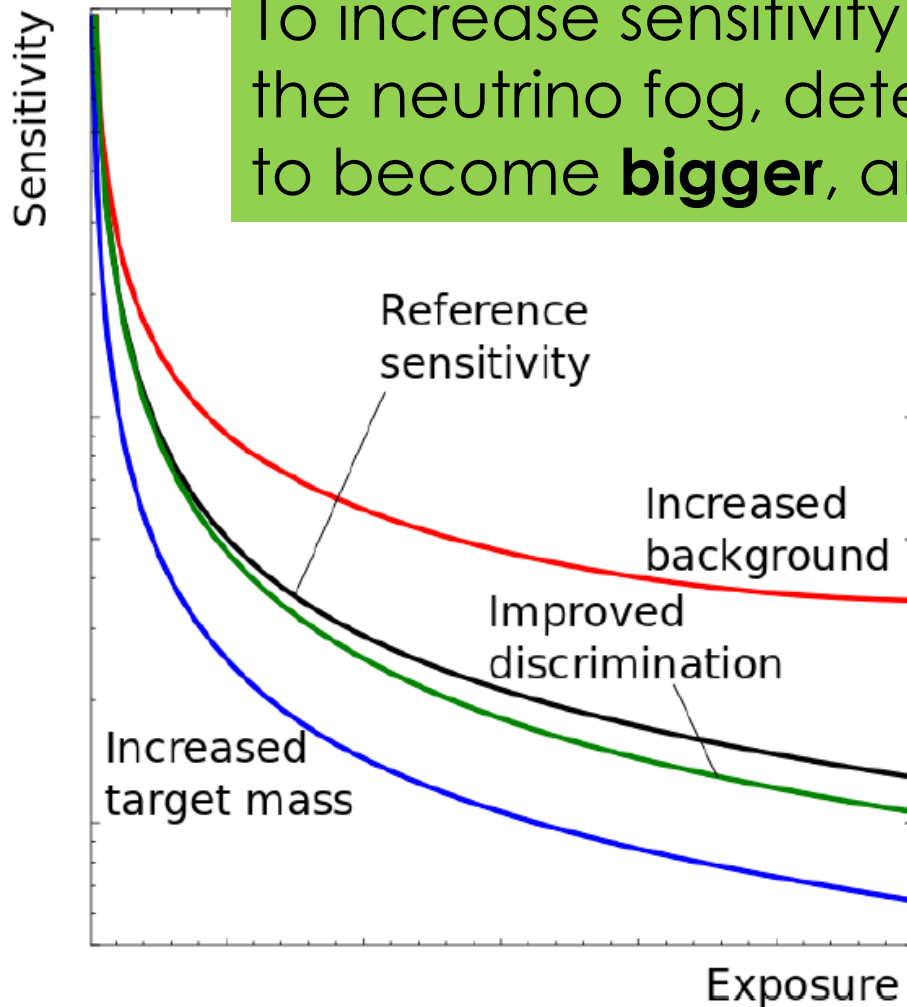
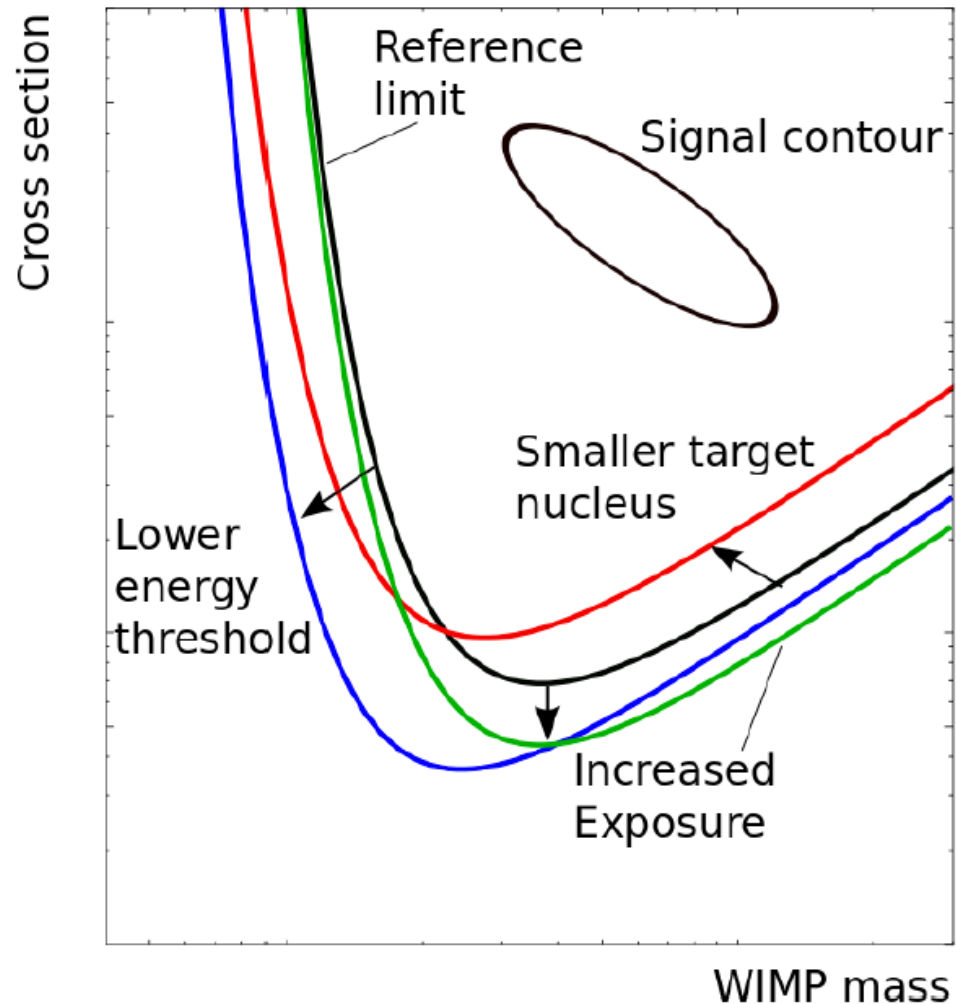
PandaX-4T @CJPL
(China)
Dual phase

LIQUID XENON DETECTORS

- **XENONnT**
 - Target: 5.9 t (fiducial 4.2 t)
 - Data taking since 2021
 - Exposure: 97.1 live-days (1.1 t x year)
- **Lux-Zeplin (LZ)**
 - Target : 7 t (fiducial 5.5 t)
 - Data taking since 2022
 - Exposure: 60 live-days (0.9 t x year)
- **PandaX-4T**
 - Target: 3.7 t (fiducial 2.7 t)
 - Data taking since 2021
 - Exposure: 86.0 live-days (0.63 t x year)



HOW TO IMPROVE WIMP SENSITIVITY?

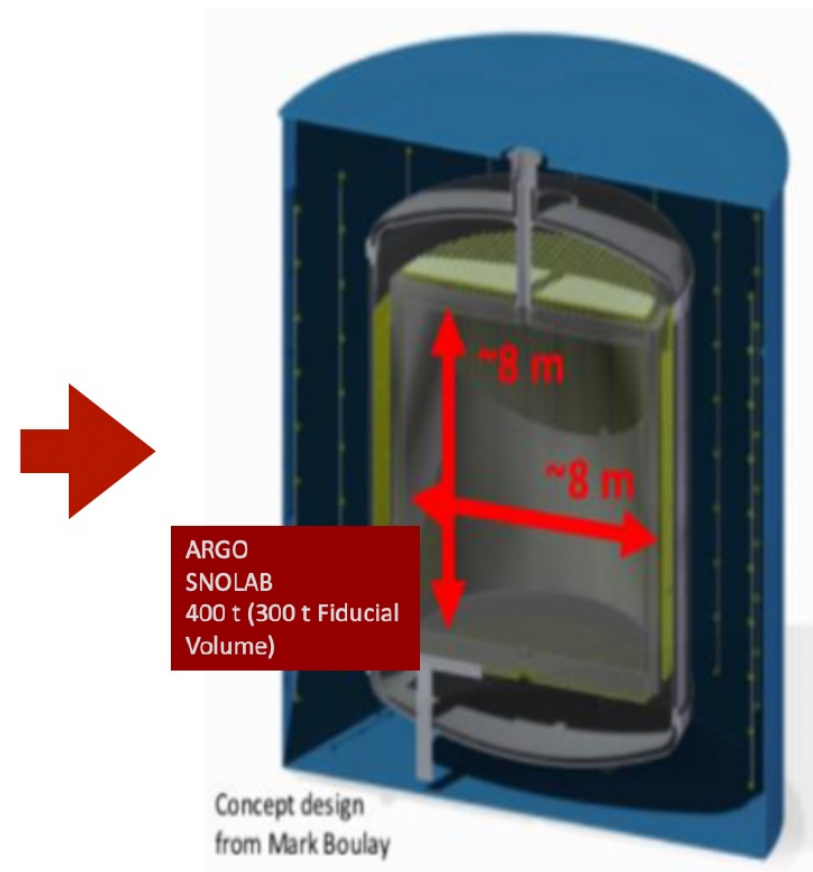
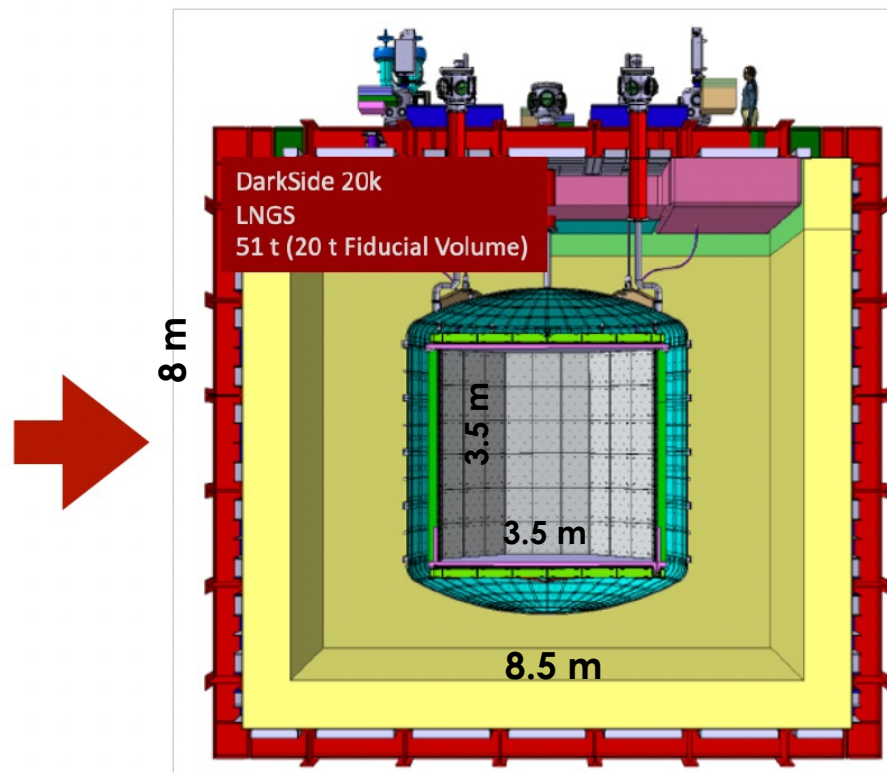


To increase sensitivity down to the neutrino fog, detectors have to become **bigger**, and **quieter**

NEXT GENERATION LIQUID NOBLE DETECTORS

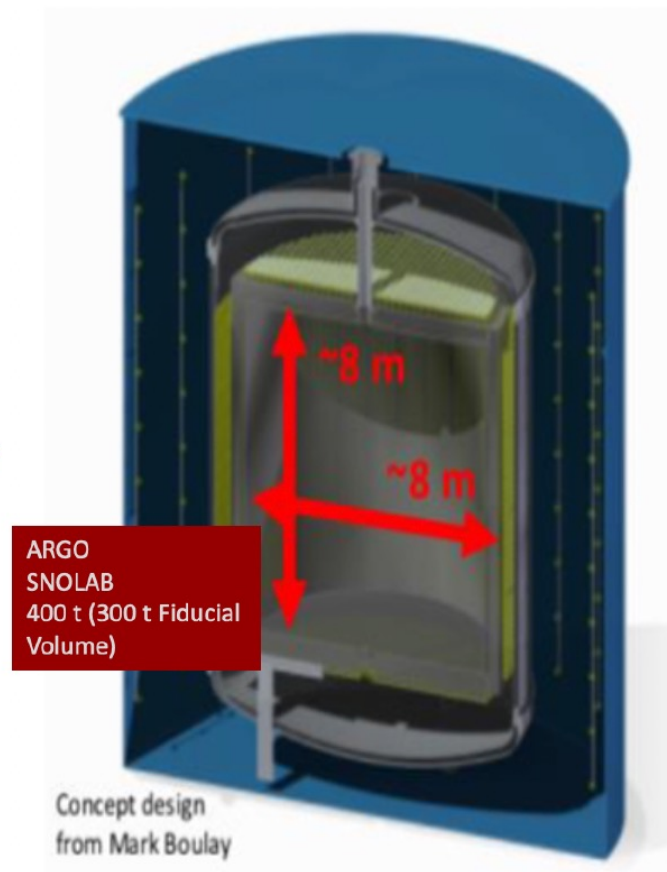
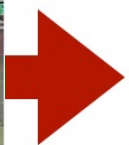
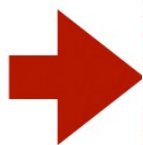
THE GLOBAL ARGON DARK MATTER COLLABORATION - GADMC

GADMC includes over 400 researchers from 69 institutions in 14 countries



THE GLOBAL ARGON DARK MATTER COLLABORATION - GADMC

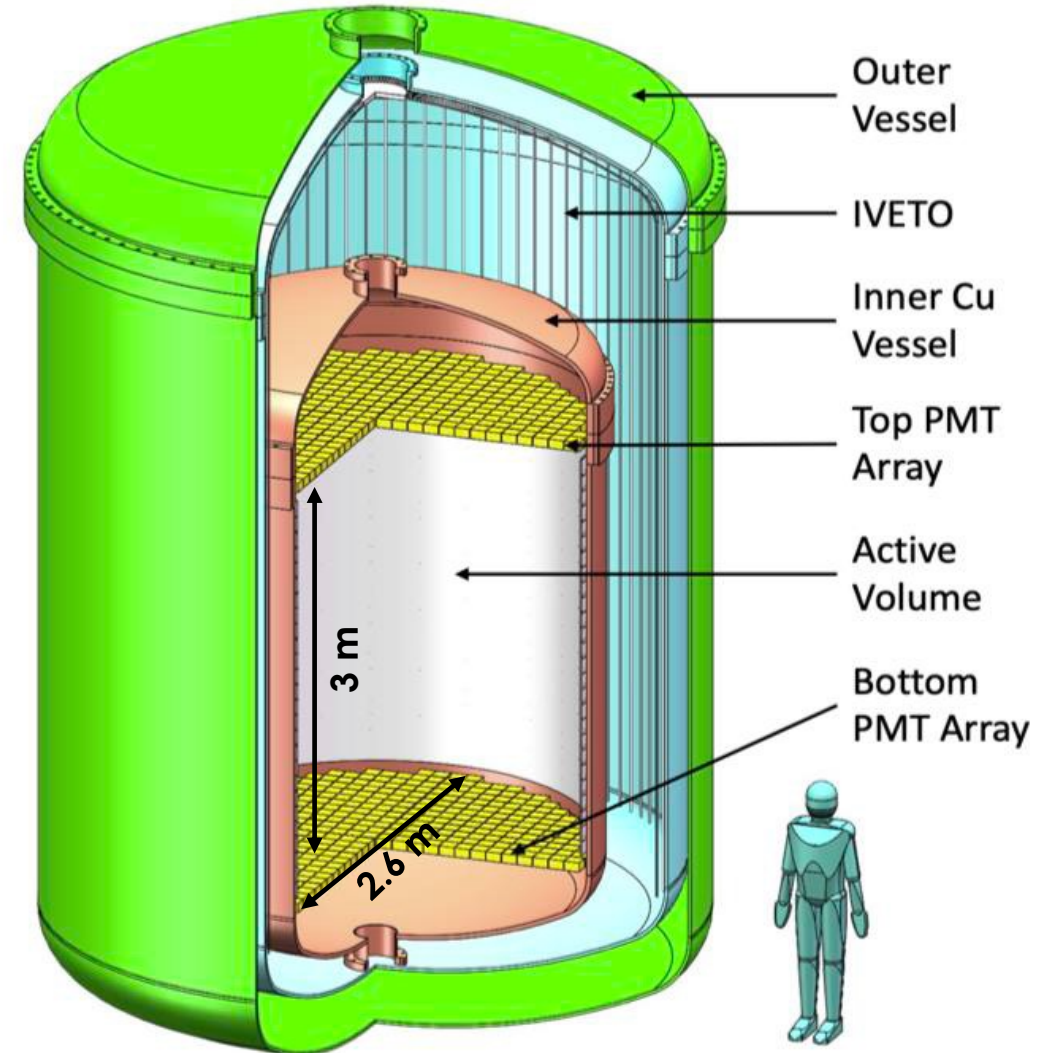
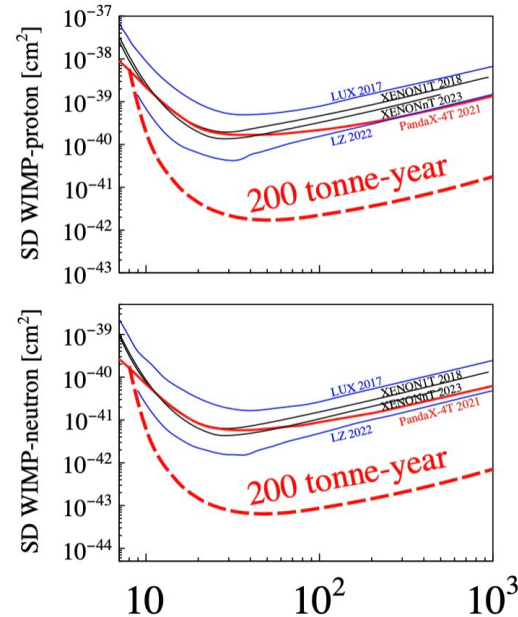
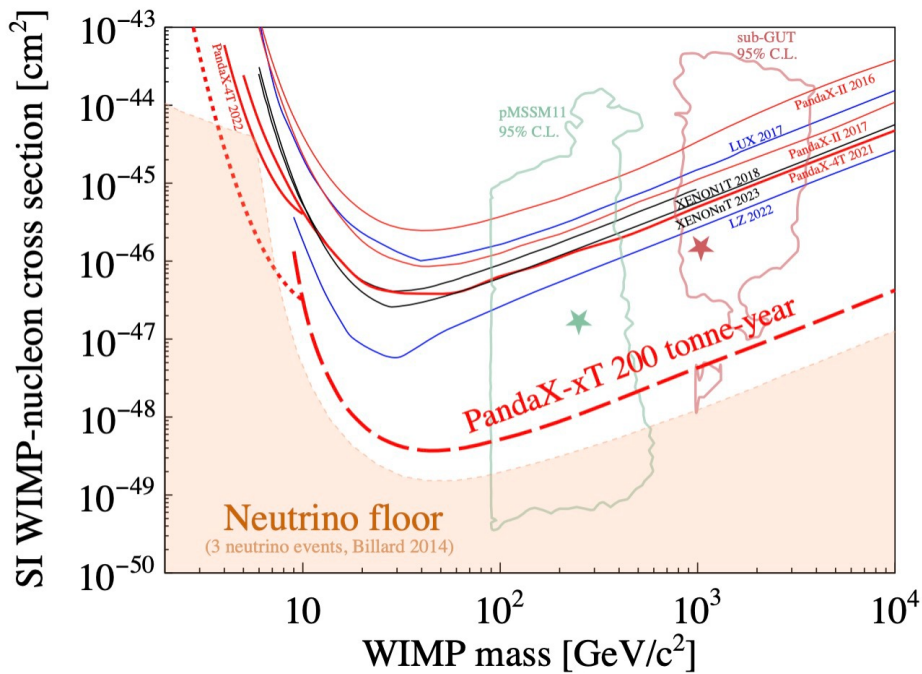
GADMC includes over 400 researchers from 69 institutions in 14 countries



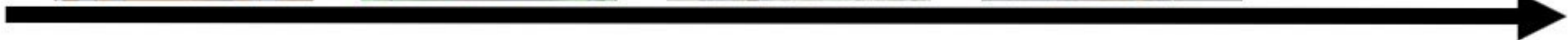
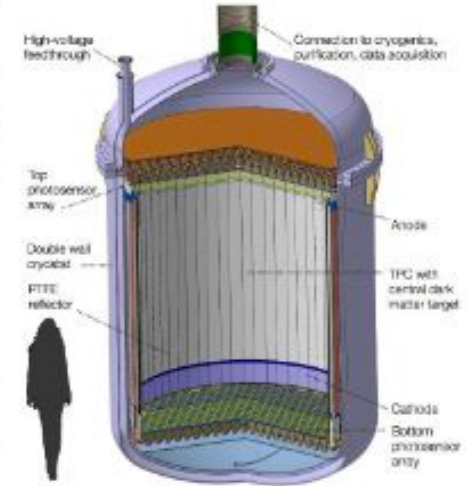
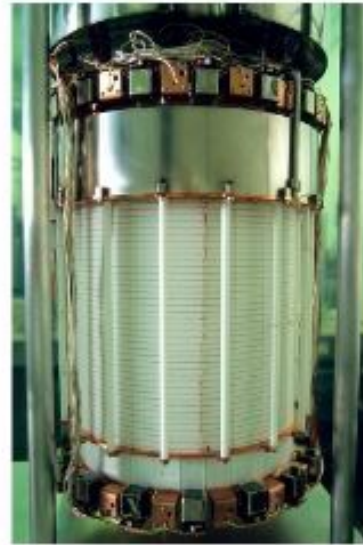
Construction started at LNGS
First physics run foreseen in 2027

PANDAX - XT

- @CJPL (China)
- Multiple phases: gradually update the detector based on the Xe possession
- Total target mass 43 t
- Operation for >10 years



THE DARWIN PROJECT



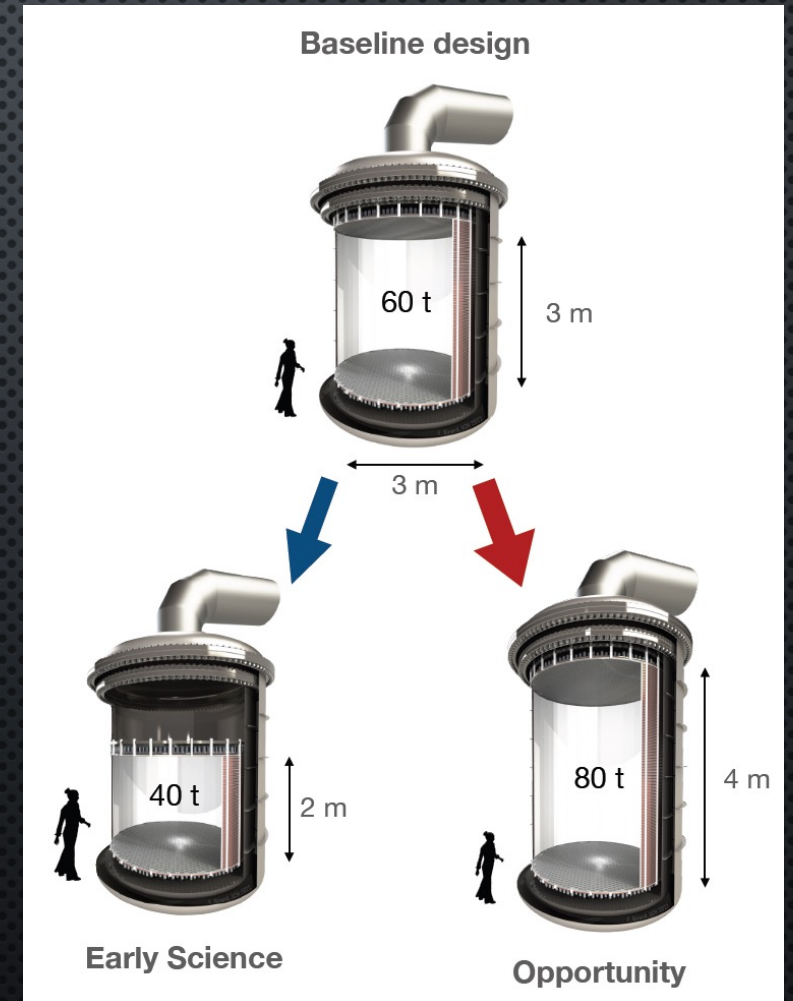
	XENON10	XENON100	XENON1T	XENONnT	DARWIN
Operation period	2005-2007	2008-2016	2012-2019	2020-2026	2030
Xenon mass	14 kg Xe target	62 kg Xe target	2 t Xe target	5.9 t active Xe 8.5 t total Xe	~40 t active Xe ~50 t total Xe
Height	15 cm	30 cm	96 cm	148 cm	~2.6 m
Diameter	20 cm	30 cm	97 cm	133 cm	~2.6 m

XLZD CONSORTIUM



XLZD consortium (xlzd.org) to design and build a common multi-ton xenon experiment

- Merger of DARWIN/XENON and LUX-ZEPLIN collaborations
 - **XENONnT** and **LZ**: ongoing science programs, technology progenitors
 - **DARWIN**: initiated R&D and design studies
- 104 group-leaders in 16 countries : MoU signed in July 2021
- joint “white paper” on physics reach : 600 authors, 141 institutions



PHYSICS REACH

Dark Matter

- Dark photons
- Axion-like particles
- Planck mass

WIMPs

- Spin-independent
- Spin-dependent
- Sub-GeV

Sun

- Solar pp neutrinos
- Solar Boron-8 neutrinos

Big Bang

- Neutrinoless double beta decay
- Double electron capture

Supernova

- Supernova neutrinos
- Multi-messenger

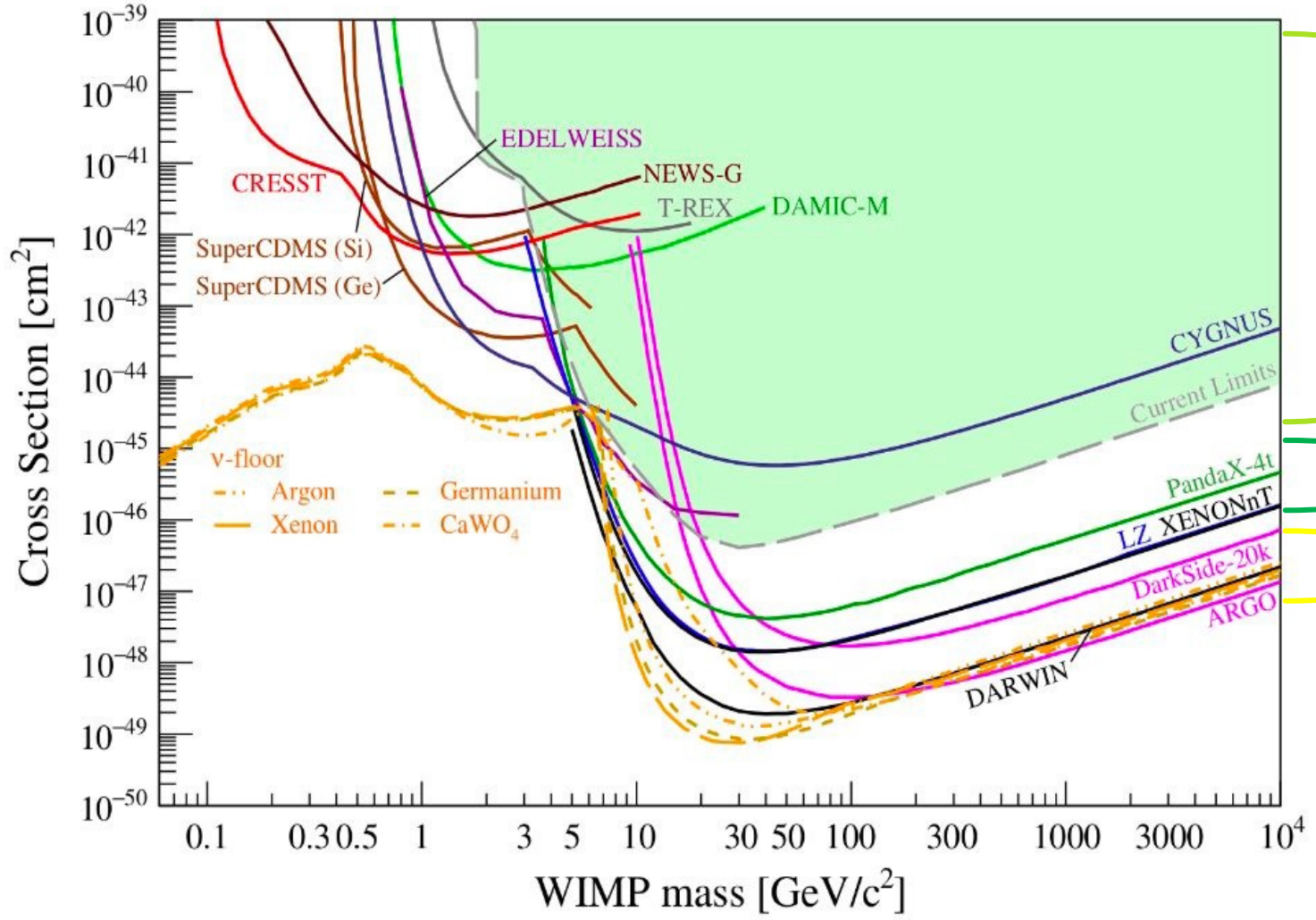
Cosmic Rays

- Atmospheric neutrinos



A Xenon Observatory with a broad science program

ULTIMATE WIMP SENSITIVITY



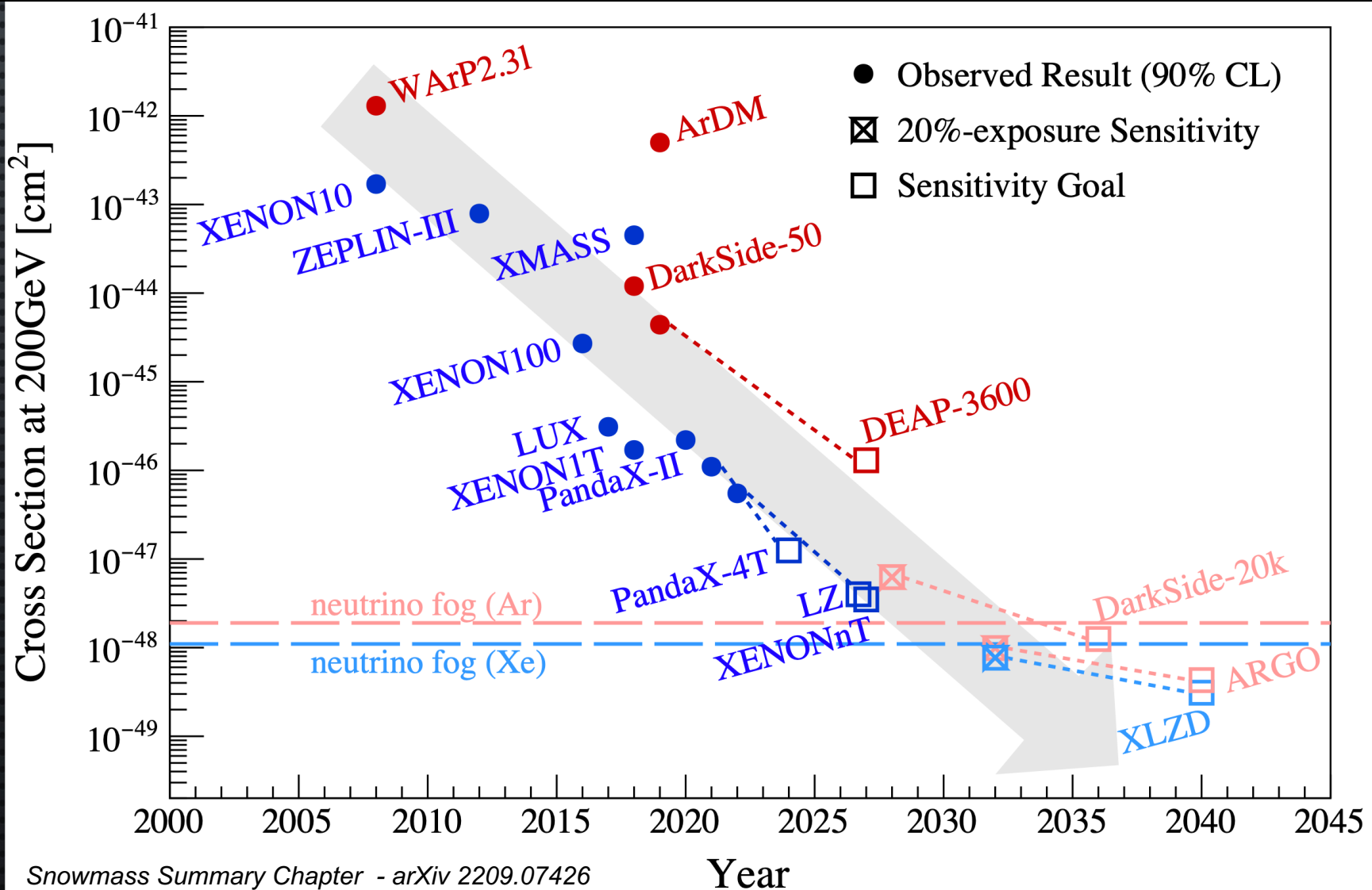
Ultimate sensitivity dominated by neutrino interactions

Past

Present

Future:
Darkside 20k
ARGO
DARWIN

ULTIMATE WIMP SENSITIVITY



CONCLUSION & OUTLOOK

- LXe and LAr TPCs have demonstrated to be the leading technologies to exploit WIMPs searches at masses $> 1 \text{ GeV} / c^2$
- Current generation of noble liquid TPCs recently presented first results and continue to take data over the coming years
- They proved to be sensitive to additional rare events too
- An effort is in place in both Xe and Ar communities towards big consortia (GADMC & XLZD) to build giant detectors aiming to explore the WIMP parameter space up to the neutrino fog + other rare events channels
- R&D and design of these next-generation detectors is ongoing