

DESY Science Communication Lab

# OneHaLe: a One-zone Hadro-Leptonic code

Michael Zacharias

*Workshop on Numerical Multi-messenger Modeling*

Paris, France



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# Table of Content

- 1) Overview of the code
- 2) Scientific examples
- 3) Ongoing work
- 4) ExHaLe-Jet



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# OneHaLe is...

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50.0      : Bulk Lorentz factor for the blob
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1.50      : Electron spectral index
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50.0      : Multiplicative factor of the light travel time [eta*R/c]
1.8       : Mass of the supermassive black hole (1.0e+8 M_sol)
5.0e-4    : Eddington ratio
1.00e+18  : Initial location of the blob along jet axis [cm]
7.60e+16  : Radius of the BLR [cm]
1.0e+0    : Effective temperature of the BLR [K]
2.30e+24  : Effective luminosity of the BLR [erg/s]
4.20e+18  : Radius of the DT [cm]
5.0e+0    : Effective temperature of the DT [K]
3.0e+24   : Effective luminosity of the DT [erg/s]
0         : Calculate neutrino detection rate using ICeCube effective areas (0=no, 1=yes)
0.0       : Strength of the magnetic field perturbation [G]
0.0e+43   : Strength of the proton injection luminosity perturbation [erg/s]
0.0       : Strength of the acceleration time scale perturbation
0.0e+1    : Strength of the minimum proton Lorentz factor perturbation
0.0e+9    : Strength of the maximum proton Lorentz factor perturbation
0.0e+42   : Strength of the electron injection luminosity perturbation
0.0       : Strength of the proton index perturbation
0.0       : Strength of the electron index perturbation
```

Example of the parameter input file.

Gory details:

- Zacharias, M., 2021, Physics, 3, 1098
- Zacharias, M., et al., 2022, MNRAS, 512, 3948

- ... a time-dependent, parallelized, hadro-leptonic, one-zone C code
- ... very flexible to accommodate various kinds of variability
- ... in each time step solves the Fokker-Planck equation for every particle species using the Chang&Cooper routine, the radiation transport equation, and obtains the neutrino output
- ... Pion and Muon evolution explicitly calculated (following Hümmer+10)
- ... includes external photon fields: accretion disk, BLR, DT, CMB
- ... particle-photon interactions involve internal and external photon fields
- ... tests suggests good agreement with the “Hadronic Code Comparison” project

# Fokker-Planck equation

$$\frac{\partial n_i(\chi, t)}{\partial t} = Q_i(\chi) + \frac{\partial}{\partial \chi} \left[ \frac{\chi^2}{(a+2)t_{\text{acc}}} \frac{\partial n_i(\chi, t)}{\partial \chi} \right] - \frac{\partial}{\partial \chi} (\dot{\chi}_i n_i(\chi, t)) - \frac{n_i(\chi, t)}{t_{\text{esc}}} - \frac{n_i(\chi, t)}{\gamma t_{i,\text{decay}}^*}$$

with normalized momentum  $\chi = \gamma\beta$ .

Particle injection

Acceleration processes

Cooling processes

Particle escape

Particle decay

# Fokker-Planck equation

$$\frac{\partial n_i(\chi, t)}{\partial t} = Q_i(\chi) + \frac{\partial}{\partial \chi} \left[ \frac{\chi^2}{(a+2)t_{\text{acc}}} \frac{\partial n_i(\chi, t)}{\partial \chi} \right] - \frac{\partial}{\partial \chi} (\dot{\chi}_i n_i(\chi, t)) - \frac{n_i(\chi, t)}{t_{\text{esc}}} - \frac{n_i(\chi, t)}{\gamma t_{i,\text{decay}}^*}$$

with normalized momentum  $\chi = \gamma\beta$ .

## Particle injection

- Primary proton and electron (incl. positrons) distributions are injected in each time step according to  $Q_i(\chi) = q_{0,i}\gamma^{-s}$  between a minimum and maximum Lorentz factor and  $q_{0,i}$  depending on input parameters
- Charged pion injection following Hümmer+10
- Muon injection from pion decay
- Secondary electron injection from muon decay, Bethe-Heitler pair production and  $\gamma$ - $\gamma$  pair production

## Acceleration processes

## Cooling processes

## Particle escape

## Particle decay



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with normalized momentum  $\chi = \gamma\beta$ .

Particle injection

Acceleration processes

- Fermi I/II, but only as a “re-acceleration”
- Main acceleration through a generic injection term
- $t_{\text{acc}}$  is a free parameter

Cooling processes

Particle escape

Particle decay

# Fokker-Planck equation

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with normalized momentum  $\chi = \gamma\beta$ .

Particle injection

Acceleration processes

Cooling processes

- Protons: synchrotron, adiabatic, p- $\gamma$ , Bethe-Heitler
- Charged pions / muons: synchrotron, adiabatic
- Electrons: synchrotron, adiabatic, inverse Compton

Particle escape

Particle decay

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with normalized momentum  $\chi = \gamma\beta$ .

Particle injection

Acceleration processes

Cooling processes

Particle escape

- Escape mimics an advective motion of the plasma through the emission region:

$$t_{\text{esc}} = \eta_{\text{esc}} R/c$$

with free parameter  $\eta_{\text{esc}} \geq 1$

Particle decay



# Fokker-Planck equation

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with normalized momentum  $\chi = \gamma\beta$ .

Particle injection

Acceleration processes

Cooling processes

Particle escape

Particle decay

- For muons and pions only with the decay time given in the proper frame of each particle

# Radiation transport equation

$$\frac{\partial n_{\text{ph}}(\nu, t)}{\partial t} = \frac{4\pi}{h\nu} j_{\nu}(t) - n_{\text{ph}}(\nu, t) \left( \frac{1}{t_{\text{esc,ph}}} + \frac{1}{t_{\text{abs}}} \right)$$

Photon production

Photon escape

Photon absorption processes

# Radiation transport equation

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## Photon production

- Synchrotron (all particles)
- Inverse-Compton (electrons) on all radiation fields (external photon fields: angle-averaged in the comoving frame after boosting + delta-function approximation to one of the integrals for each IC/ext process)
- Neutral pions decay directly to  $\gamma$ 's

## Photon escape

## Photon absorption processes



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# Radiation transport equation

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Photon production

Photon escape

- Photons leave the source with average escape time

$$t_{\text{esc,ph}} = 0.75 R/c$$

Photon absorption processes

# Radiation transport equation

$$\frac{\partial n_{\text{ph}}(\nu, t)}{\partial t} = \frac{4\pi}{h\nu} j_{\nu}(t) - n_{\text{ph}}(\nu, t) \left( \frac{1}{t_{\text{esc,ph}}} + \frac{1}{t_{\text{abs}}} \right)$$

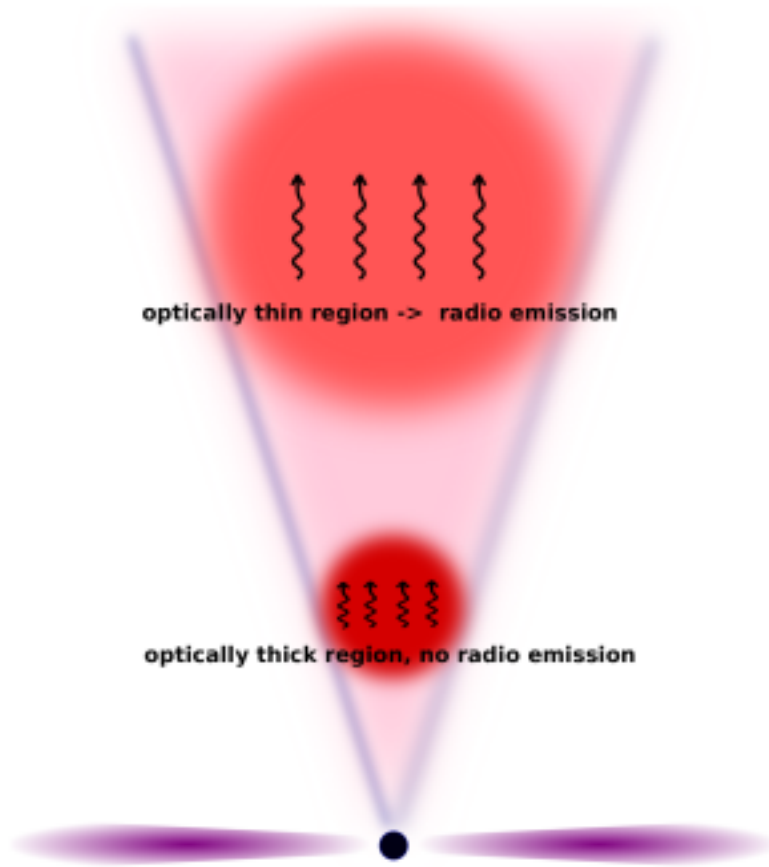
Photon production

Photon escape

Photon absorption processes

- Bethe-Heitler and  $\gamma$ - $\gamma$  pair production processes using all photon fields (external ones angle-averaged in the comoving frame after boosting)
- Synchrotron-self absorption
- Photons that left the emission region, are also absorbed in the BLR and DT fields (but no EBL or CMB absorption considered) if applicable

# Scientific Example 1: Moving and expanding blob

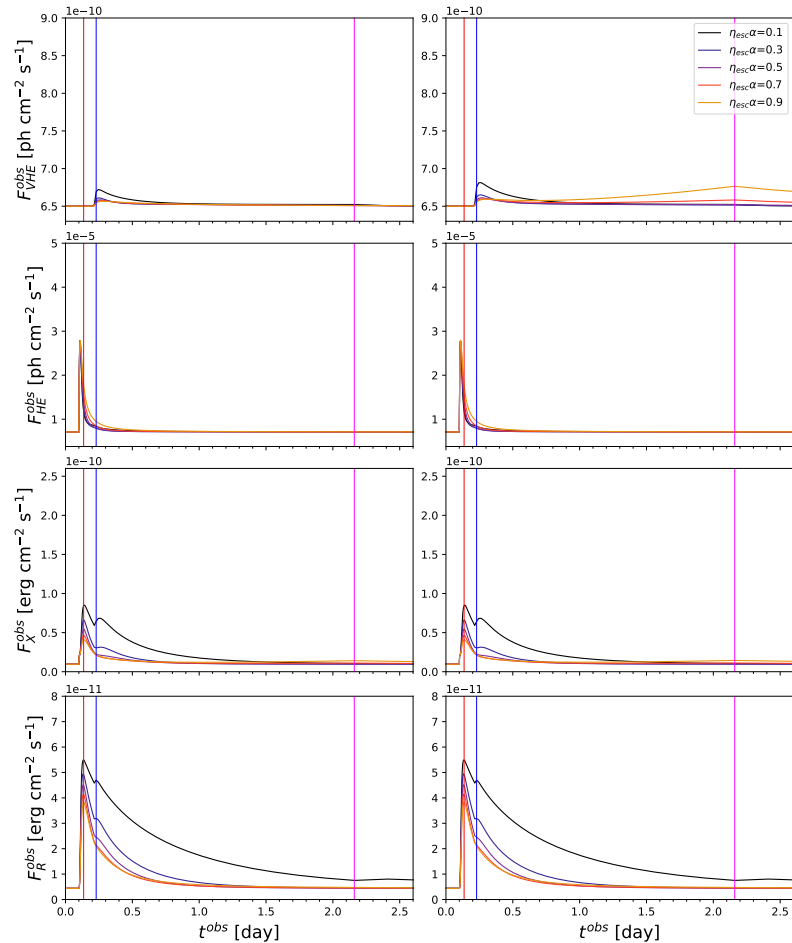


- The blazar one-zone model typically assumes a constant radius of the emission region
- If the emission region moves, the blob should expand adiabatically due to its higher pressure compared to the jet medium
- Expansion:  $R(t) = R_0 + \alpha ct$
- Escape of particles:  
 $t_{\text{esc}}(t) = \eta_{\text{esc}} R(t)/c = t_0 + \eta_{\text{esc}} \alpha t$  with  $t_0 = \eta_{\text{esc}} R_0/c$
- If the blob expands rapidly ( $\eta_{\text{esc}} \alpha \rightarrow 1$ ), particles are trapped efficiently and particles accumulate
- For constant bulk flow  $\Gamma$ , *time* and *distance from BH* are related linearly  $z \propto \beta_{\Gamma} ct$  (comoving frame!)

Expansion of blobs in a conical jet

(taken from Boula&Mastichiadis22)

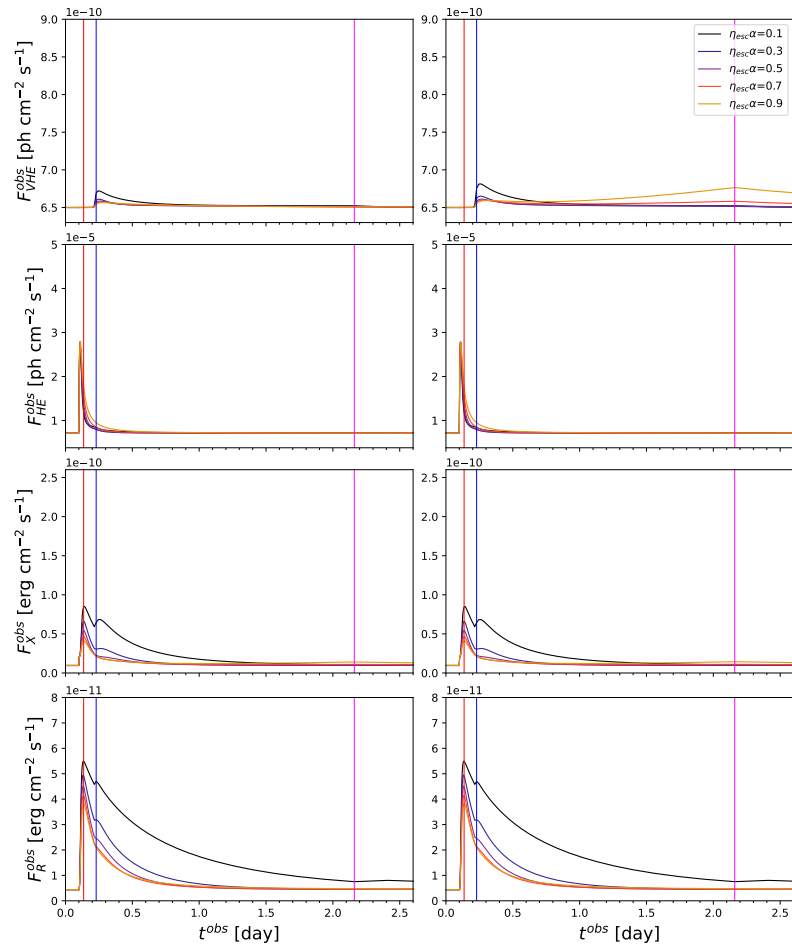
# Simulating a cascade based on PKS 1510-089



- Using a parameter set based on PKS 1510-089, an FSRQ at  $z = 0.361$  with bright AD, BLR, DT
- Magnetic field evolution:  $B(z) = B_0 R_0 / R(z)$  (assuming dominating toroidal structure)
- Curves are shown for increasing opening angles (*dark to light*):  
 $\eta_{\text{esc}} \alpha \in [0.1, 0.3, 0.5, 0.7, 0.9]$
- Vertical lines mark:  $t_0$  (red), passing BLR (blue), passing DT (magenta)
- Note:  $\Delta t^{\text{obs}} = \Delta z' / (\delta \Gamma \beta_{\Gamma} c)$

Simulations **without** cascade.

# Simulations without cascade



**VHE** emission absorbed within the BLR; shows a secondary bump in hadronic sims for large opening angles at  $t_{DT}$

**HE** shows a quick rise to peak at  $t_0$ ; flare over at  $t_{BLR}$ ; minor difference in the decay pattern

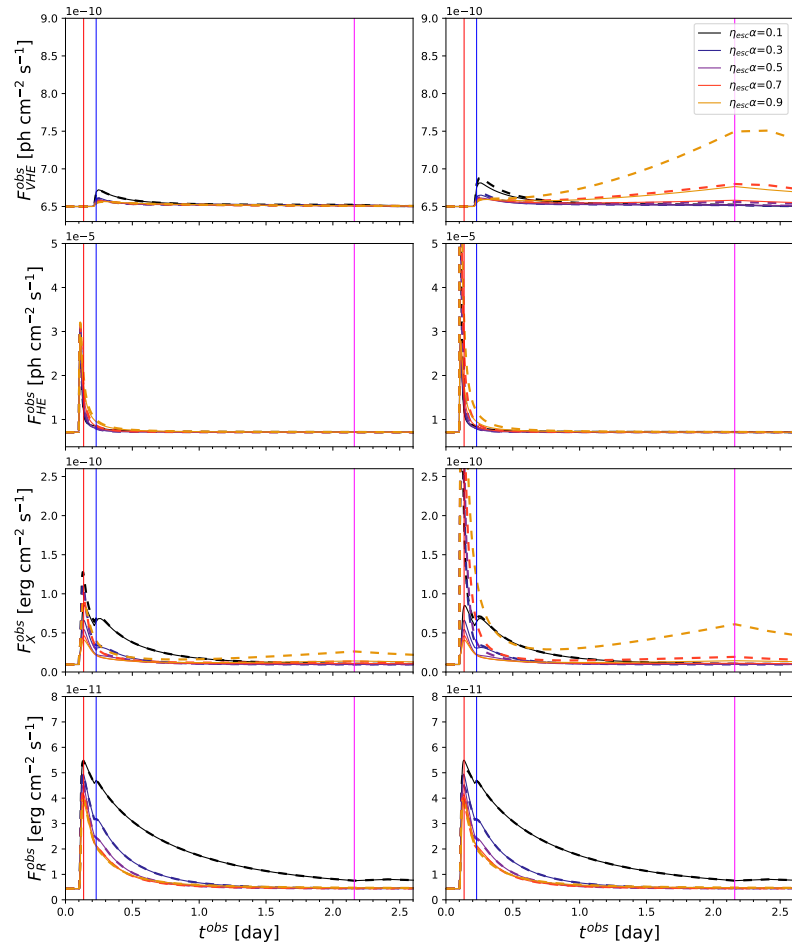
**X** rise a bit slower than HE; secondary peak at  $t_{BLR}$  for small opening angles; this influences the decay pattern; very minor secondary flare at  $t_{DT}$

**R** similar to X-ray

Simulations **without** cascade.



# Simulations with cascade



**VHE** similar to sim w/o cascade, but secondary bump in hadronic sims for large opening angles at  $t_{DT}$  much stronger

**HE** leptonic sim similar to w/o cascade except for slightly higher peak flux; much higher peak fluxed (off scale) in hadronic sim, but also quick decline

**X** leptonic sim with higher peak flux, but similar decay pattern except for more pronounced secondary flare at  $t_{DT}$  for large opening angles; hadronic sim with much higher peak flux (off scale) and much stronger secondary peak at  $t_{DT}$  for large opening angles

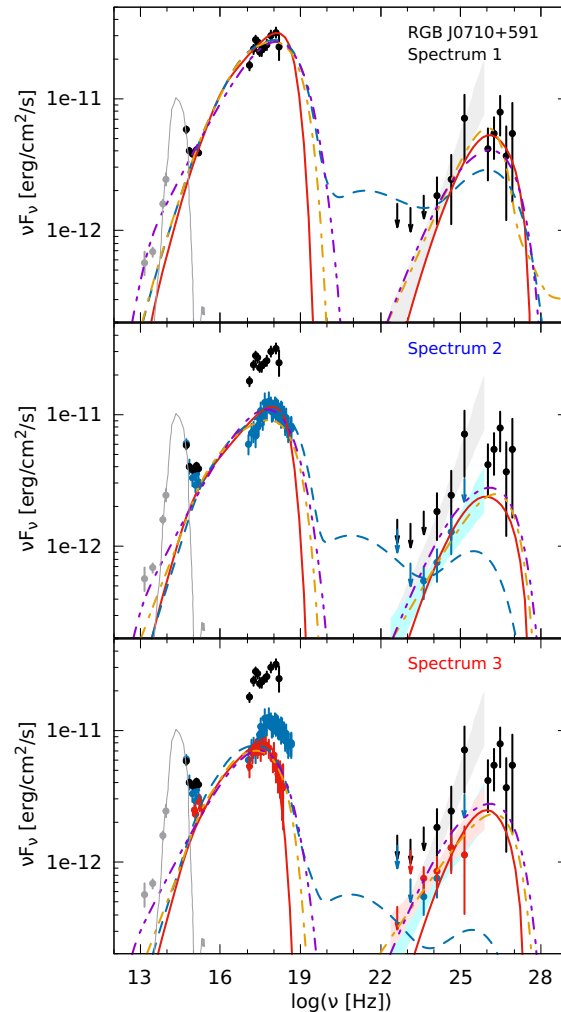
**R** not influenced by the cascade

Gory details:

- Zacharias, M., 2023, A&A, 669, A151

Simulations **with** cascade (*dashed lines*).

## Scientific Example 2: Steady-state modeling of eHBLs



- eHBLs exhibit the most extreme peak frequencies among blazars
- Can exhibit (long-term) variability
- Study to model 4 eHBLs with various models (SSC, *e-p-shock*, LH $\rho$ , LH $\pi$ )
- Used OneHaLe in steady-state for the LH models

Modeling various states of RGB J0710+591

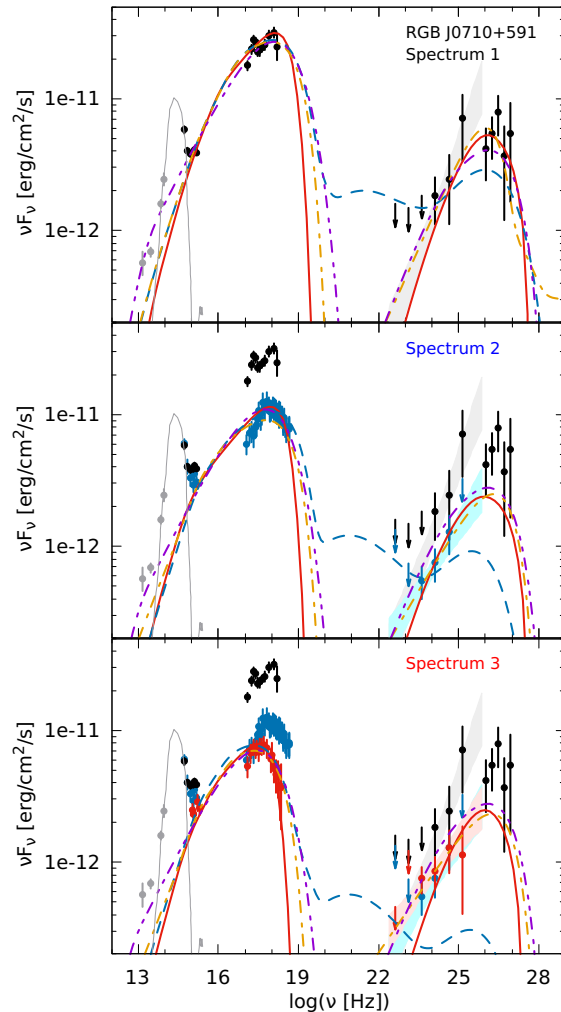


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## Scientific Example 2: Steady-state modeling of eHBLs



- SSC, *e-p-shock* and LHp with good fits
- SSC with least power consumption
- *e-p-shock* with best physical setup and good power demand
- LHp with excessive power demand
- $LH\pi$  parameters chosen such to suppress SSC, no good fits, excessive power demand
- Interestingly, upper limits on AD suggest that power output of eHBL is above the AD power (irrespective of model)

Gory details:

- Goswami, P., et al., 2024, A&A, 682, A134

Modeling various states of RGB J0710+591



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# Ongoing development

Version 1.1 (current version)

- Available upon reasonable request to me
- hdf5 usage optional (but Bethe-Heitler only with hdf5)
- Output written to individual ascii files
- Variability limited to certain shapes, unless one digs into the code



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## Version 2.0 (in development)

- ✓ Full hdf5 (output written to single hdf5 file with python script for first look plots and to produce ascii files if wanted)
- ✓ Variability patterns easy to change for the user
- () Include neutrons
- () Further user-friendliness improvements
- () Upload to GitHub

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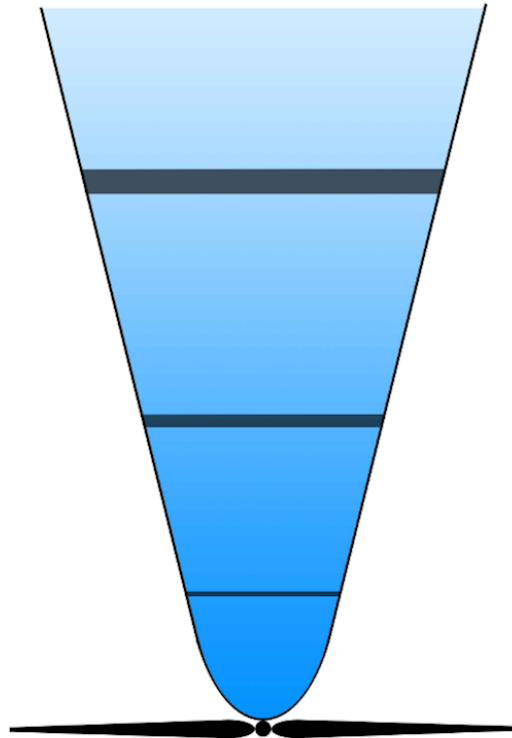
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## Further plans:

- Tabulate more integrals, remove delta-approximations
- Switch  $\gamma\text{-}\gamma$  pair production cross section from Aharonian+83 to Böttcher&Schlickeiser97
- Allow for restart of sim after certain checkpoints
- Any suggestions?



# ExHaLe-jet: An Extended Hadro-Leptonic jet model



Sketch: jet cut into numerous slices (dark), in which the kinetic equations for each particle species are solved Figure:

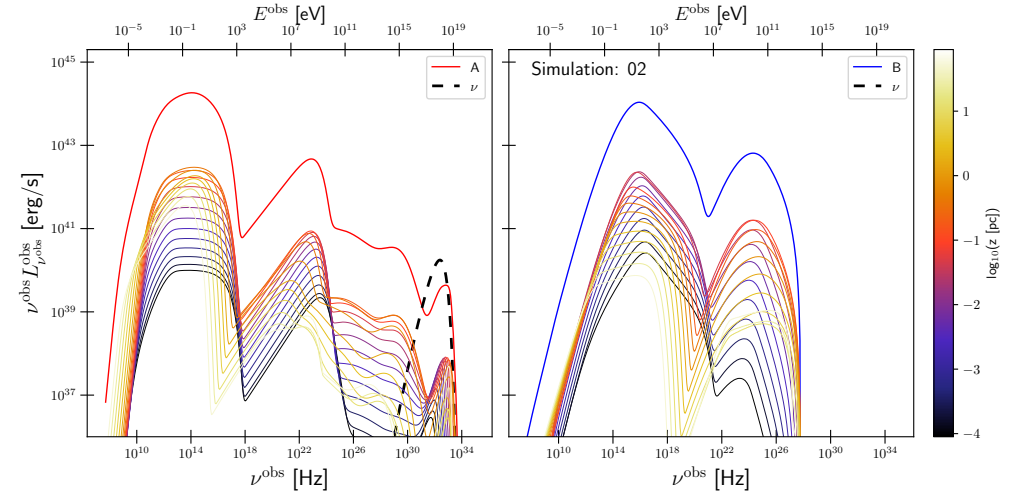
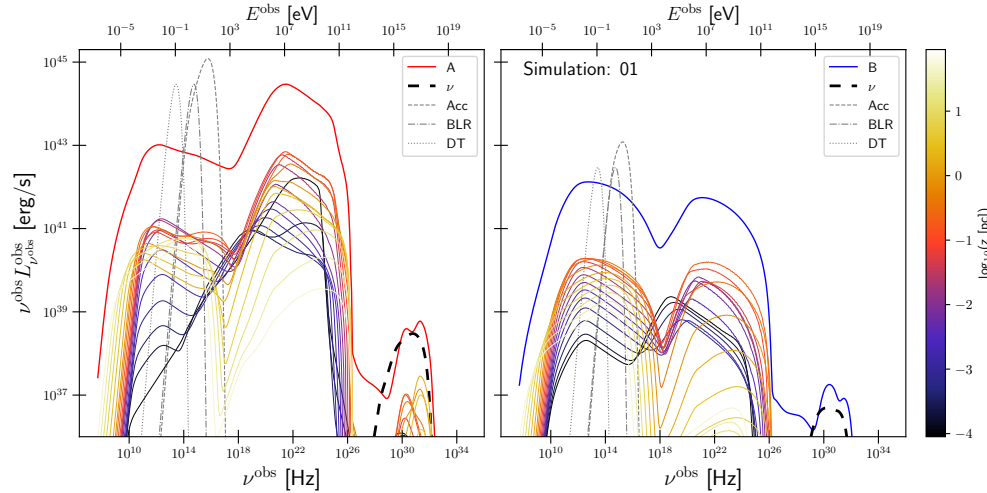
courtesy of Jonathan Heil

- Core functionality as OneHaLe
- Jet length cut into numerous slices, where kinetic equation is solved for each species
  - Injection of primary proton and electron distribution at the base; evolved self-consistently along the jet
  - Injection of secondaries (pions, muons, pairs) in each slice
  - Pairs propagated along with primaries
  - Radiation and neutrino output for each slice
- Geometry currently fixed as
  - Parabolic acceleration region:  $\Gamma_b(z) \propto \sqrt{z}$
  - Conical coasting region  $\Gamma_b(z) = \text{const.}$
  - Radius:  $R(z) \propto \tan [0.26/\Gamma_b(z)]$
  - Magnetic field derived with Bernoulli equation
- Code not for public use as of now

Gory details:

- Zacharias, M., et al., 2022, MNRAS, 512, 3948

# ExHaLe-jet: Total spectra



## Strong external fields

- High Compton dominance
- Most flux  $\sim z_{\text{acc}}$
- Total power sub-Eddington
- Moderate neutrino number

## Weak external fields

- Low Compton dominance
- Most flux  $\sim z_{\text{acc}}$
- Total power sub-Eddington
- Low neutrino number

## P-syn solution

- P flux  $< 0.5z_{\text{acc}}$
- E flux  $< 10z_{\text{acc}}$
- Total power sub-Eddington and p dominated
- High neutrino number

## SSC solution

- Low Compton dominance
- Most flux  $\sim z_{\text{acc}}$
- SSC drops faster than syn
- Total power sub-Eddington and e dominated



# Famous last words...

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- It includes (almost) all relevant processes incl. external photon fields
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Anton Dmitriiev, Patrick Kilian, Andreas Zech, Anita Reimer, Catherine Boisson, Markus Böttcher, and the various people who have already used and tested the code

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0.0       : Strength of the magnetic field perturbation [G]
0.0e+43   : Strength of the proton injection luminosity perturbation [erg/s]
0.0       : Strength of the acceleration time scale perturbation
0.0e+1    : Strength of the minimum proton Lorentz factor perturbation
0.0e+9    : Strength of the maximum proton Lorentz factor perturbation
0.0e+42   : Strength of the electron injection luminosity perturbation
0.0       : Strength of the proton index perturbation
0.0       : Strength of the electron index perturbation
```

Example of the parameter input file.

Gory details:

- Zacharias, M., 2021, Physics, 3, 1098
- Zacharias, M., et al., 2022, MNRAS, 512, 3948

- OneHaLe is a flexible, time-dependent, lepto-hadronic one-zone code
- It includes (almost) all relevant processes incl. external photon fields
- Version 1.1 available upon reasonable request to me
- Version 2.0 in development; will be uploaded to GitHub
- Big THANKS to:  
Anton Dmitriiev, Patrick Kilian, Andreas Zech, Anita Reimer, Catherine Boisson, Markus Böttcher, and the various people who have already used and tested the code

**Thank you for your attention!**



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