Simulation-Based Inference Stakes and applications

Sacha Guerrini - Cosmology and Statistics Day 01/02/2024



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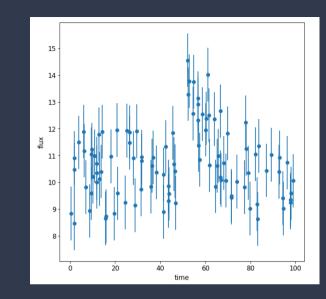
- I. Likelihood-Based Inference
- II. Simulation-Based Inference -Principles and Stakes
 III. Simulation-Based Inference -

Applications in cosmology

What is inference?

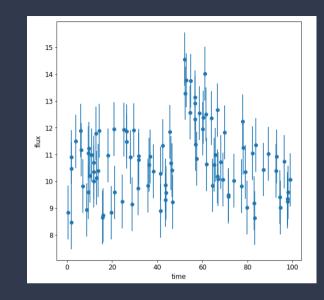


What is inference?



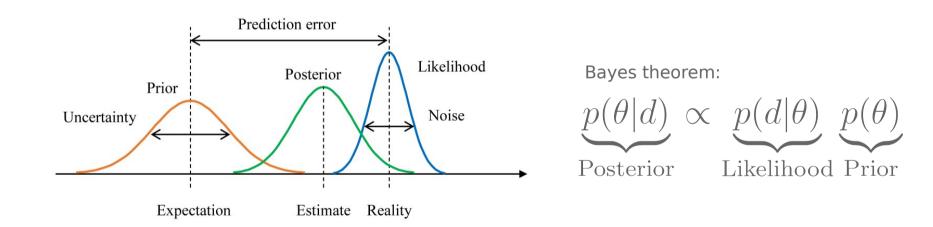


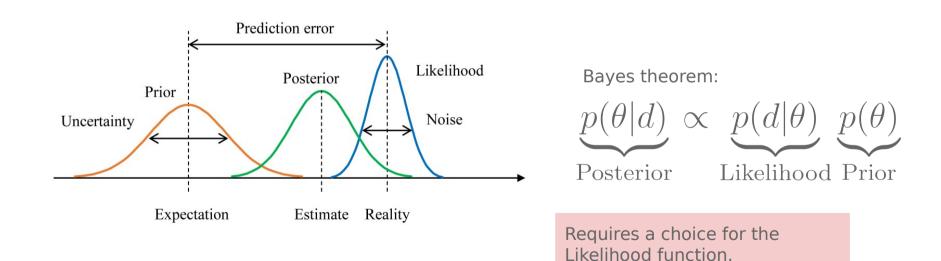
What is inference?











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Usual assumption: the Gaussian likelihood

 $L(\mathbf{d}|\boldsymbol{\theta}) = (2\pi)^{-m/2} |\mathbf{C}(\boldsymbol{\theta})|^{-1/2} \exp\left[-\frac{1}{2}\boldsymbol{\mu}(\boldsymbol{\theta})^T \mathbf{C}^{-1}(\boldsymbol{\theta})\boldsymbol{\mu}(\boldsymbol{\theta})\right]$

 $\mu = \mathbf{d} - \mathbf{y}(\boldsymbol{\theta})$ Data vector Model prediction

 $\mathbf{C}(\boldsymbol{\theta})$: Covariance matrix

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Covariance estimation methods:

- Jackknife/bootstrap resampling
- Sample variance from simulations
- Analytical model

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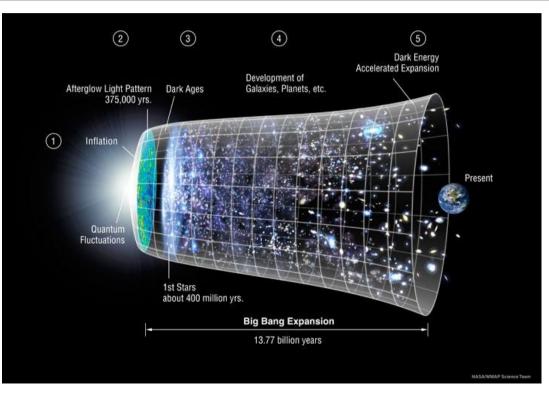
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Covariance estimation methods:

- Jackknife/bootstrap resampling
- Sample variance from simulations
- Analytical model

Can be cumbersome and/or computationally expensive.

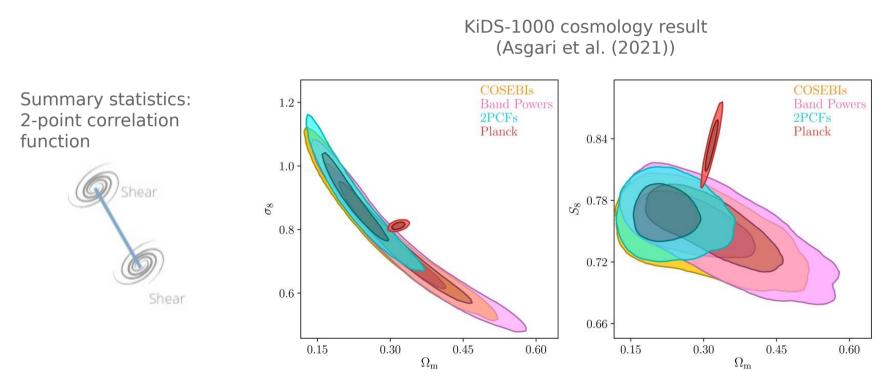
Application: Cosmological parameters inference



The history of the Universe is well described by a few parameters:

- H₀: Current expansion rate
- Ω_m : Matter density
- Ω_b : Baryon density
- Ω_{Λ} : Dark energy density
- σ_8 : Clumpiness
- n_s: Scale index of initial density fluctuations
- w: Evolution of dark energy

Application: Cosmological parameters inference



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What is problematic with LBI?

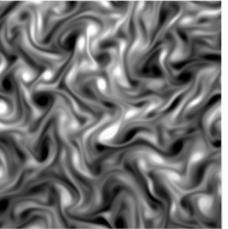
Caveats of Likelihood-Based Inference

- Covariance estimate can be cumbersome.
- Latent variables are intractable and can add systematic effects that needs to be taken into account.
- The Likelihood is not necessarily Gaussian or analytical. (Gaussian assumption do not capture all the information, namely the interactions between scales.)

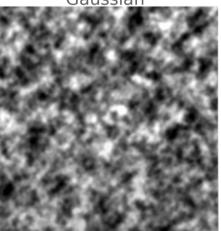
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Turbulent flow



Gaussian



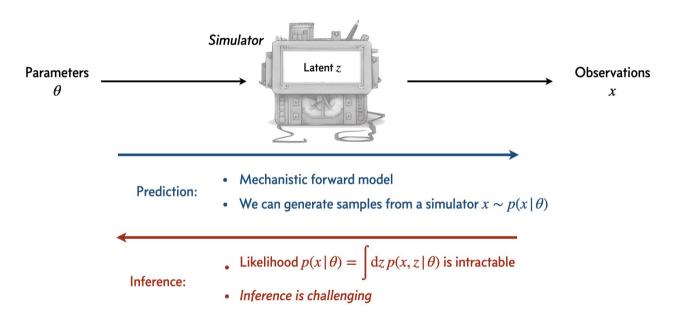


Image Credit: Siddharth Mishra-Sharma

Number of Simulation-based Inference Papers by Year 180 160 140 120 Number of papers 100 80 60 40 20 0 2001-2002-2003-2004 2005 2006-2007 2008-2009. 2010-2011 2012-2013-2014-2015-2016-2017-2018-2019-2020-2021-2022-2023-2024

year

Tools and Resources

Likelihood approximation networks (LANs) for fast inference of simulation models in cognitive neuroscience

Alexander Fengler 🖱, Lakshmi N Govindarajan, Tony Chen, Michael J Frank 🍧

Department of Cognitive, Linguistic and Psychological Sciences, Brown University, United States; Carney Institute for Brain Science, Brown University, United States; Psychology and Neuroscience Department, Boston College, United States

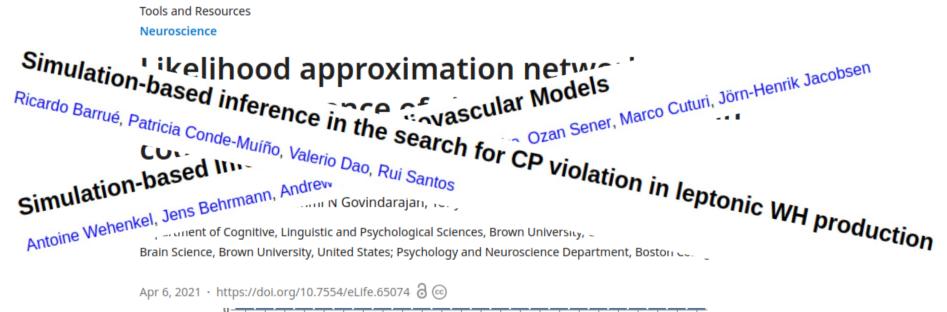
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for fast inference of
Cognitie
Simulation-based Inference for Cardiovascular Models
                    Tools and Resources
  Antoine Wehenkel, Jens Behrmann, Andrew C. Miller, Guillermo Sapiro, Ozan Sener, Marco Cuturi, Jörn-Henrik Jacobsen
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Tools and Resources Neuroscience

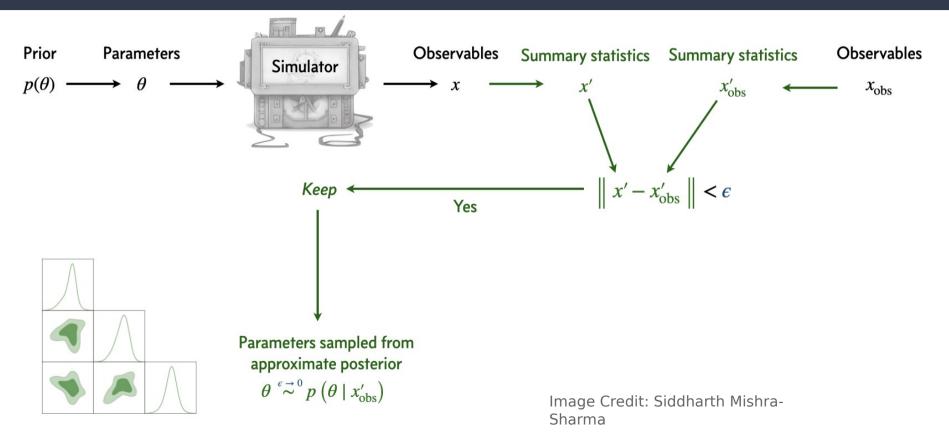
Simulation-based inference and approximation network ----- Cuturi, Jörn-Henrik Jacobsen Ricardo Ram Simulation-Based Inference of Strong Gravitational Lensing Parameters

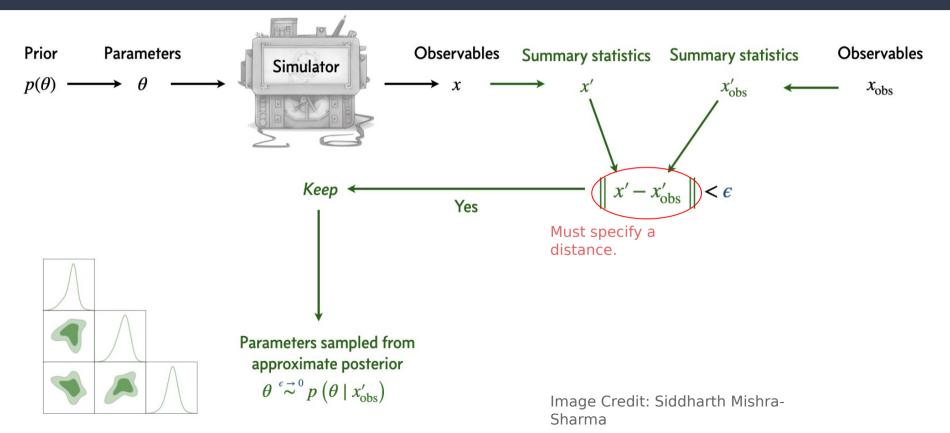
Ronan Legin, Yashar Hezaveh, Laurence Perreault Levasseur, Benjamin Wandelt

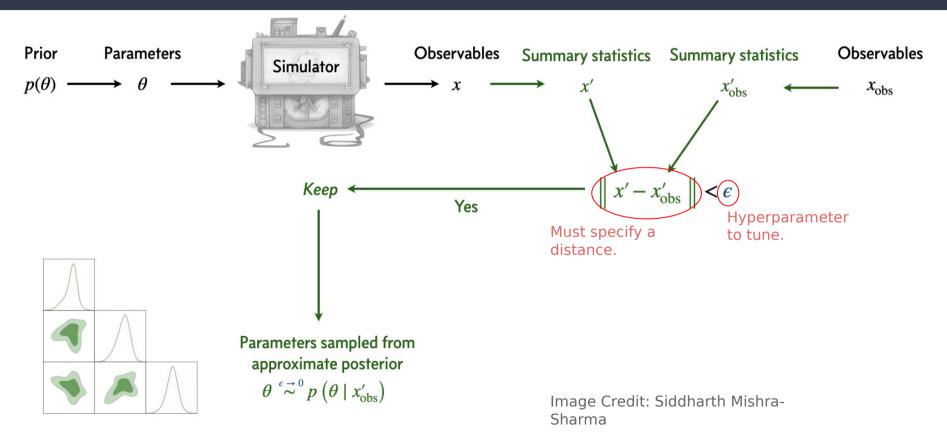
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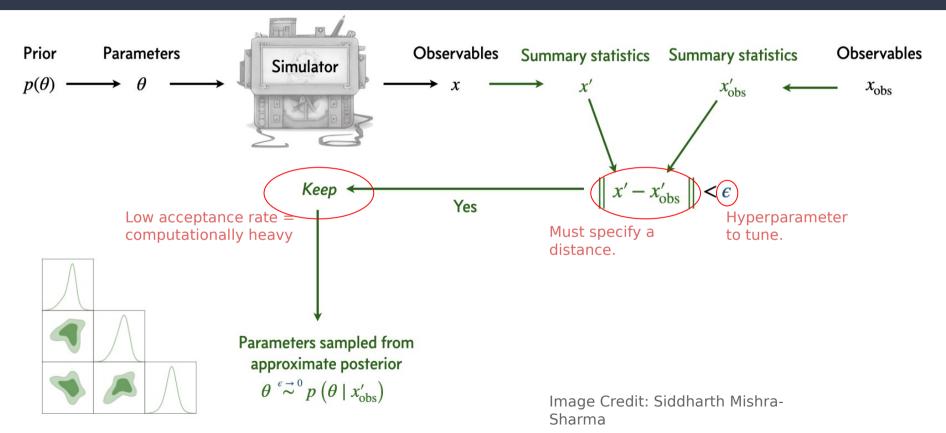
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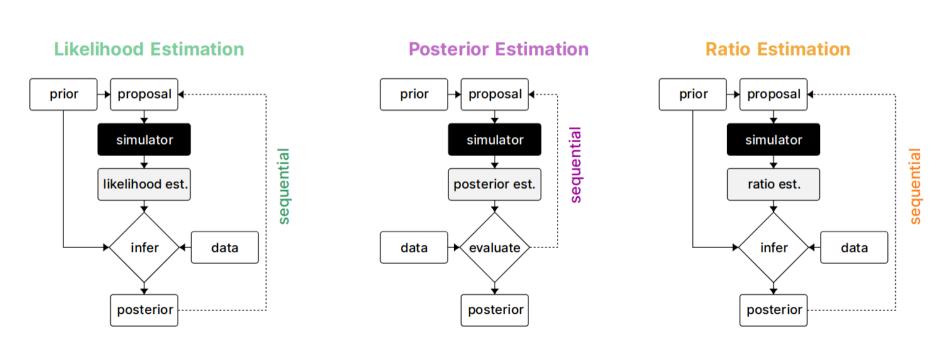


Image Credit: Lueckmann et al.

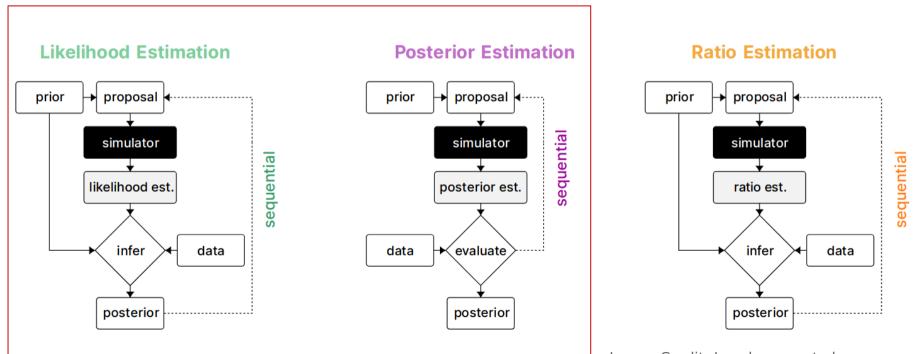
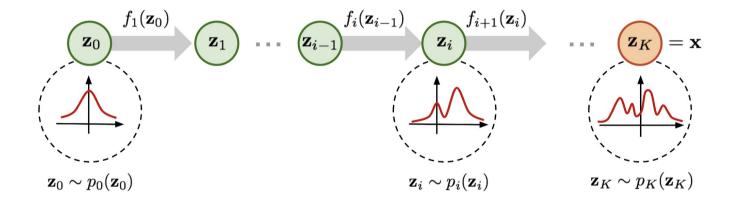


Image Credit: Lueckmann et al.

Normalizing Flows



$$p_i(z_i) = p_{i-1}(f^{-1}(z_i)) \left| \det \frac{df^{-1}}{dz_i} \right|$$

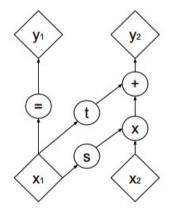
An example: Real NVP

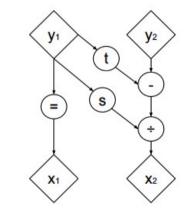
Training data: $(heta, d_{ ext{sim}})$

Loss function: $\ln U = -\sum_{i} \ln p(d_i | \theta_i; w)$

Minimizes KL-divergence w.r.t a target density:

$$D_{\mathrm{KL}}(p^*(d|\theta)||p(d|\theta;w)) = \int p^*(d|\theta) \log\left(\frac{p^*(d|\theta;w)}{p(d|\theta;w)}\right)$$

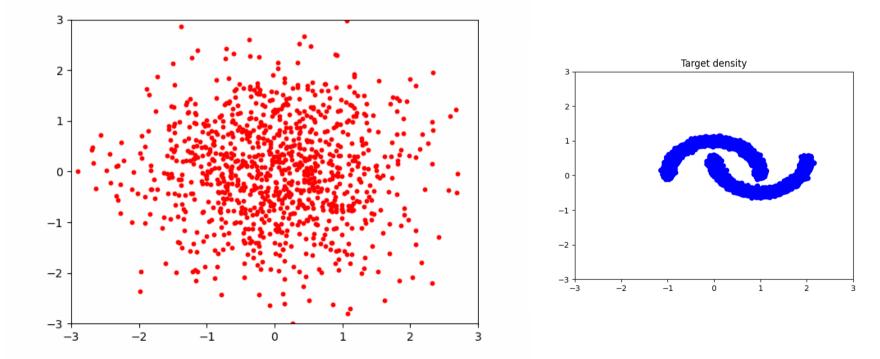




(a) Forward propagation

(b) Inverse propagation

An example: Real NVP



Simulation-Based Inference: Summary

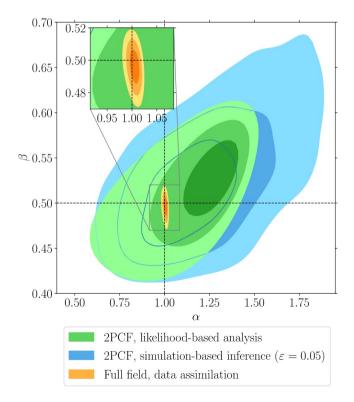
Advantages:

- Does not require any assumption on the form of the likelihood, you learn it.
- Systematics effects can be included in the forward model.
- Straightforward method.
- Can be trained and exploited on GPUs and the model is fully differentiable.

Drawbacks:

- Requires high-quality simulated data. The latter is difficult to assess.
- Uncertainty with neural network is difficult to quantify.
- Training can be long and cumbersome (e.g. compression and pre-training steps)

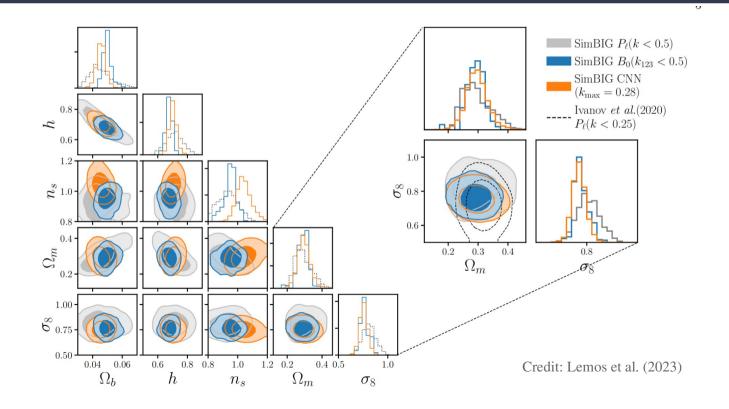
Application: Cosmological parameters inference



Leclercq and Heavens (2021):

- Simplified toy model for cosmology application.
- Likelihood-based analysis does not peak at the accurate mode.

Application: Cosmological parameters inference



Conclusion and summary

Simulation-based inference allows:

- To sample parameters from the posterior without any assumption on the form of the likelihood.
- To include all systematics effects in the forward model.
- To efficiently perform those computations on GPUs.

But:

- One must assess the realism of the simulations used.
- Uncertainty quantification is difficult.

Ressources:

- https://github.com/smsharma/awesome-neural-sbi
- https://simulation-based-inference.org/