

# Weak Lensing Mass Mapping with Uncertainty Quantification

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Co-supervised at CosmoStat, CEA DAp

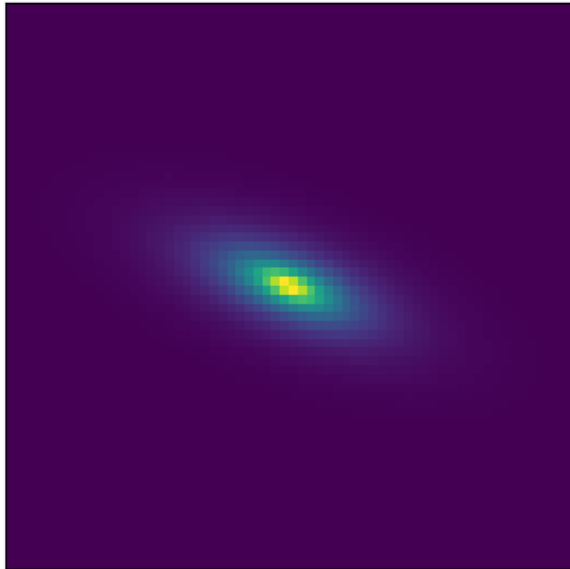
Joint ARGOS-TITAN-TOSCA workshop

2nd February 2024

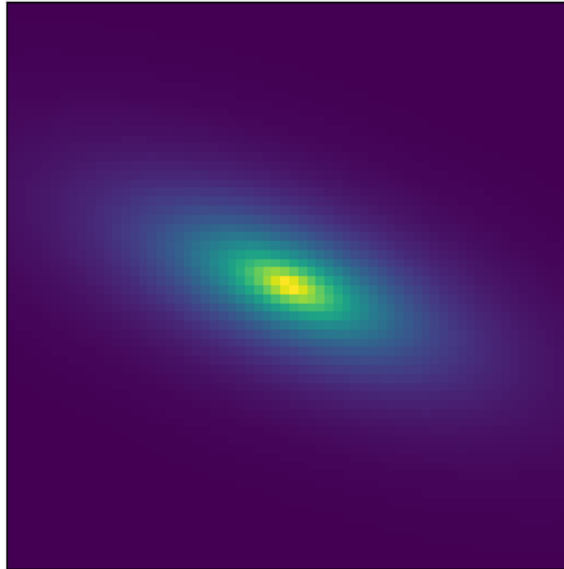


# Context and objectives

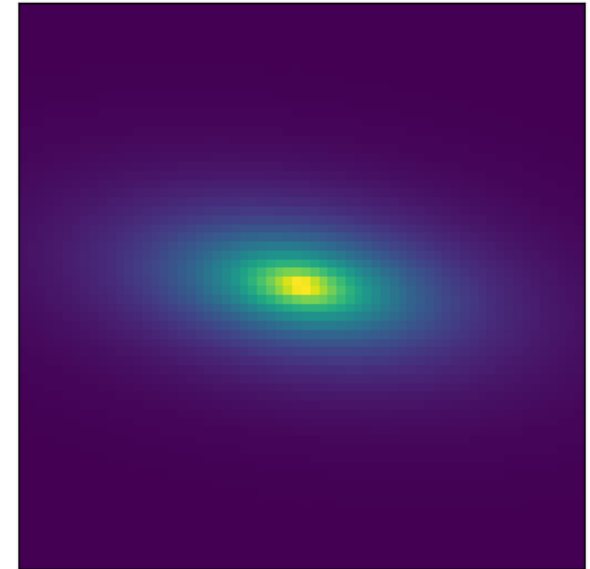
- **Gravitational lensing:** light rays emitted by distant objects (e.g., galaxies) are deflected by inhomogeneous matter density in the foreground.
- Dark matter cannot be detected by direct observations, but mass mapping can be performed from gravitational lensing observations.
- **Weak lensing regime:** two types of deformation:
  - **convergence:** isotropic dilation of the source;
  - **shear:** anisotropic stretching of the image.



Source galaxy, unlensed



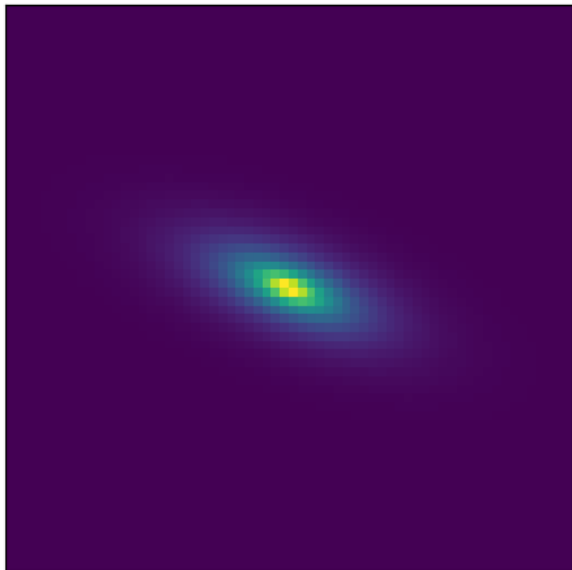
Convergence only  
 $\kappa = 1$



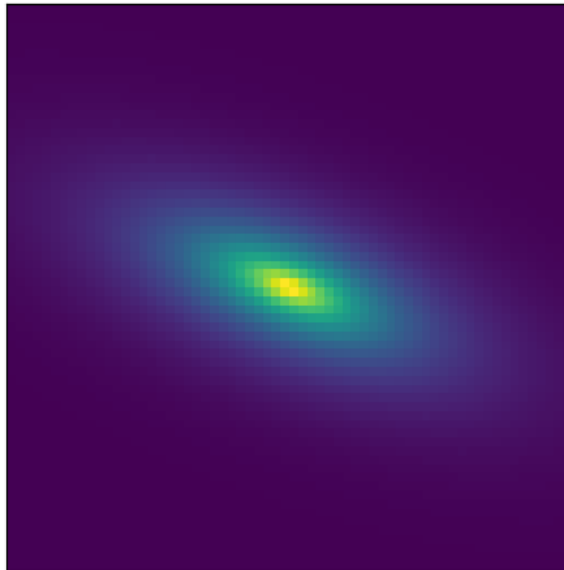
Convergence + shear  
 $\kappa = 1$  and  $\gamma = (0.1 - 0.3 i)$

# Context and objectives

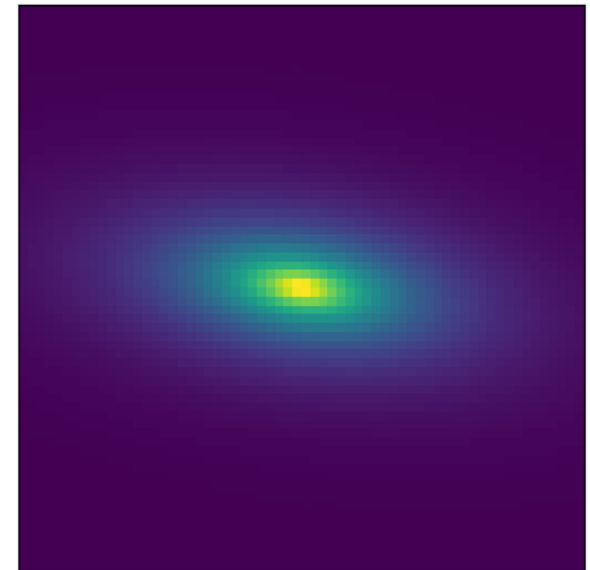
- Convergence map  $\kappa \in \mathbb{R}^K$  at a given redshift  $z$ : proportional to the projected mass along the line of sight  $\Rightarrow$  **variable of interest**.
- However,  $\kappa$  cannot be directly measured.
- Relationship between shear and convergence maps:  $\boldsymbol{\gamma} = \mathbf{A}\boldsymbol{\kappa}$ , with
  - $\boldsymbol{\gamma} \in \mathbb{C}^K$  true shear map (unknown);
  - $\mathbf{A} \in \mathbb{R}^{K \times K}$  Kaiser-Squires filter (known).



Source galaxy, unlensed



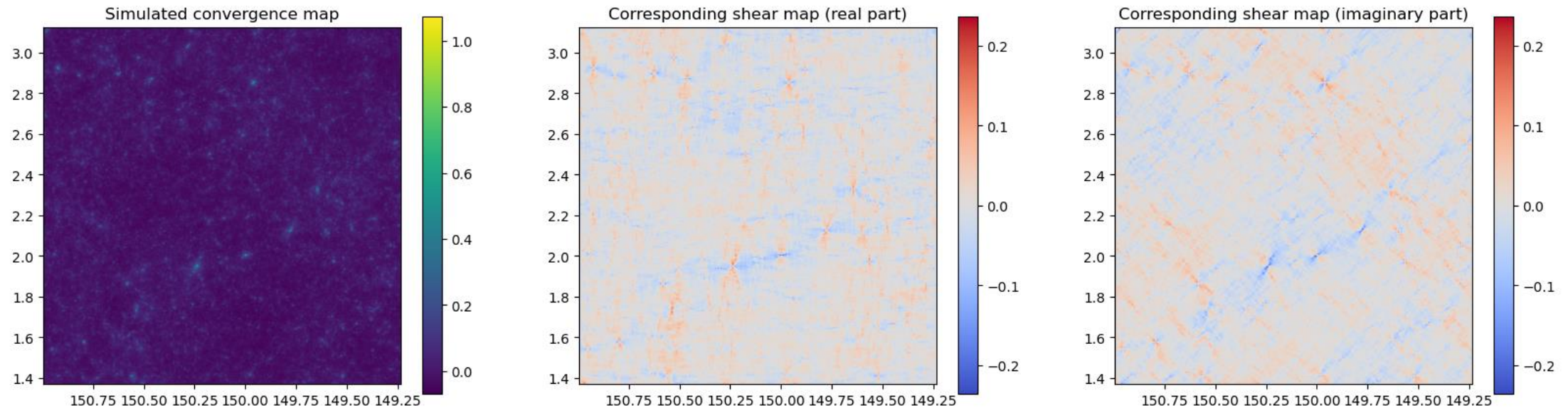
Convergence only  
 $\kappa = 1$



Convergence + shear  
 $\kappa = 1$  and  $\boldsymbol{\gamma} = (0.1 - 0.3 i)$

# Context and objectives

Example with the  $\kappa$ TNG simulated dataset<sup>1</sup>

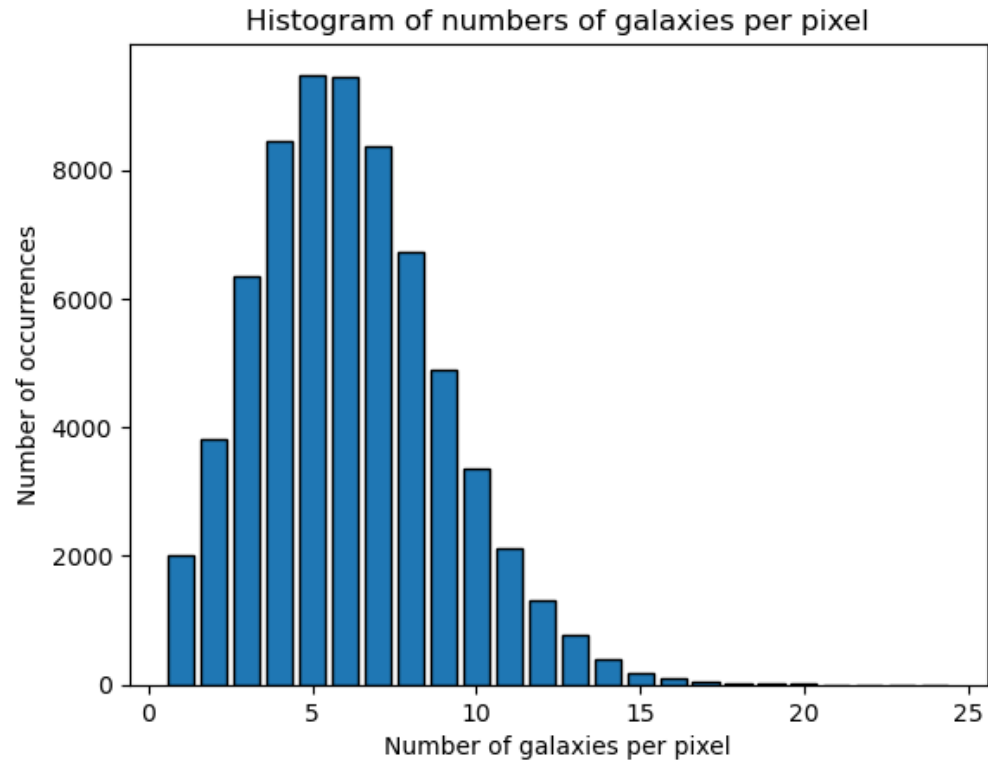
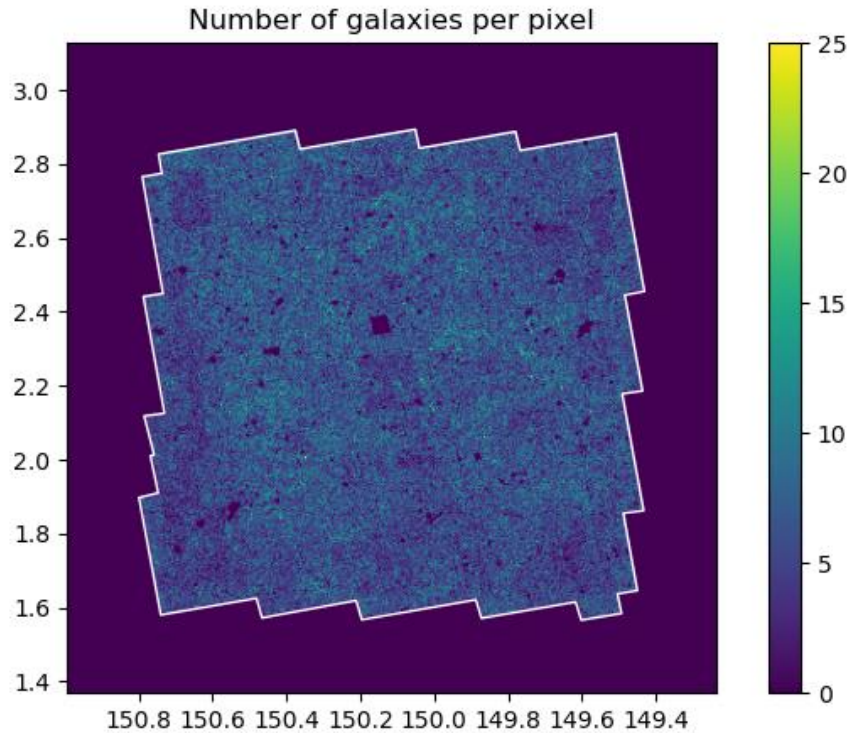


- As for  $\kappa$ , the true shear map  $\boldsymbol{\gamma}$  cannot be directly measured.
- Unbiased estimator of  $\boldsymbol{\gamma}$ , denoted by  $\hat{\boldsymbol{\gamma}}$ , obtained by measuring galaxy ellipticities.
- Relation between  $\hat{\boldsymbol{\gamma}}$  (observable) and  $\boldsymbol{\kappa}$  (quantity of interest):
$$\hat{\boldsymbol{\gamma}} = \mathbf{A}\boldsymbol{\kappa} + \mathbf{n}.$$
- Level of noise: depends on the number  $N_k$  of observed galaxies at a given pixel  $k$ .

<sup>1</sup> <http://columbialensing.org/>

# Context and objectives

Number of measured galaxies per pixel + mask (from the COSMOS shape catalog<sup>1</sup>):

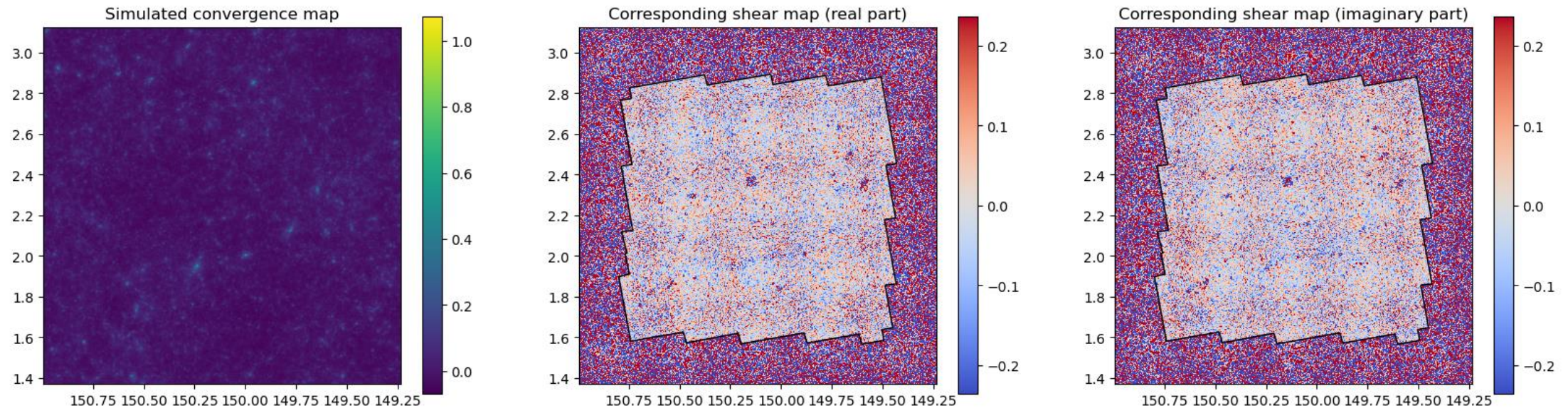


- At pixel  $k$ ,  $\sigma_k \propto \frac{1}{\sqrt{N_k}}$  ( $N_k$  number of measured galaxies).
- Masked data: noise set to an arbitrary large value.

<sup>1</sup> <https://astro.uni-bonn.de/en/m/schrabba/research>

# Context and objectives

Noisy shear maps (ellipticities)



**Objective:** given  $\hat{\gamma}$ , estimate  $\hat{\kappa}^-$  and  $\hat{\kappa}^+$  such that

$$P\{\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]]\} \leq \alpha$$

for a given confidence level  $\alpha \in ]0, 1[$ , for any pixel  $k$ .

**Requirement:** we seek a “fast” mass mapping method (i.e., non-iterative).

# Proposed approach

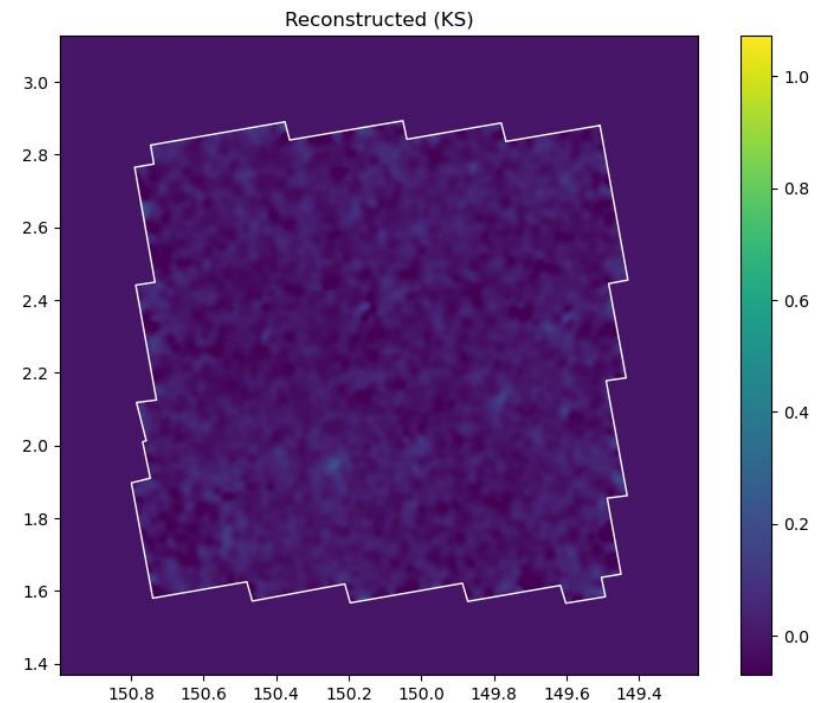
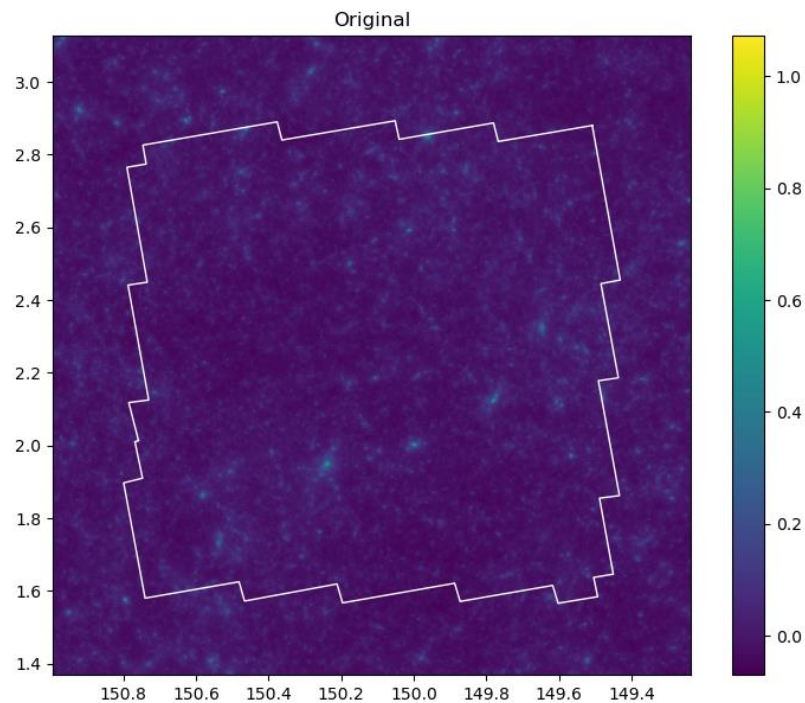
1. Compute heuristic bounds  $\hat{\kappa}^-$  and  $\hat{\kappa}^+$  (Kaiser-Quires, MSE solution, Wiener solution, plug-and-play algorithms, Glimpse, MCAIens...).
2. Post-processing: adjust bounds  $\hat{\kappa}^-$  and  $\hat{\kappa}^+$  using a **calibration set**.

→ Distribution-free UQ, does not assume any prior distribution.

→ Works for any heuristic prediction method, including deep learning.

# A simple example: the Kaiser-Squires solution

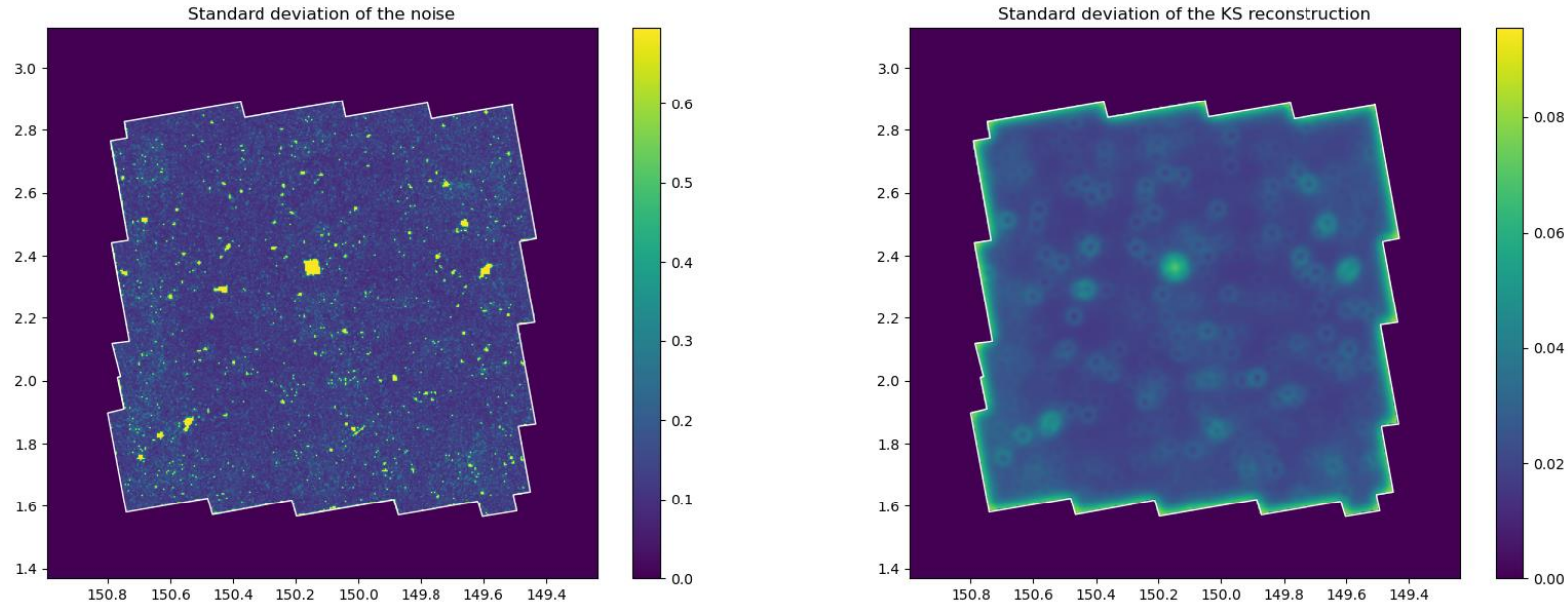
- Exact solution if  $\hat{\boldsymbol{\gamma}} = \boldsymbol{\gamma}$  (no noise, no mask):  $\hat{\boldsymbol{\kappa}} = \mathbf{A}^\dagger \hat{\boldsymbol{\gamma}}$ .
- In practice, the KS filter is followed by Gaussian smoothing:  $\hat{\boldsymbol{\kappa}} = \mathbf{S} \mathbf{A}^\dagger \hat{\boldsymbol{\gamma}}$ .





# Kaiser-Squires bound estimation

Estimation of  $\hat{\kappa}^-$  and  $\hat{\kappa}^+$ : we have  $\hat{\kappa} \sim N(\mathbf{S}\kappa, \mathbf{S}\mathbf{A}^\dagger \Sigma \mathbf{A}^\dagger \mathbf{S}^*)$ .

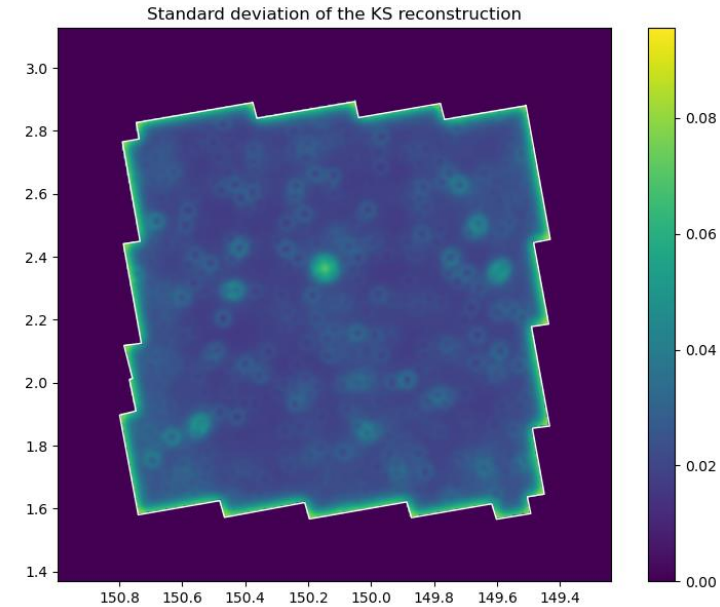
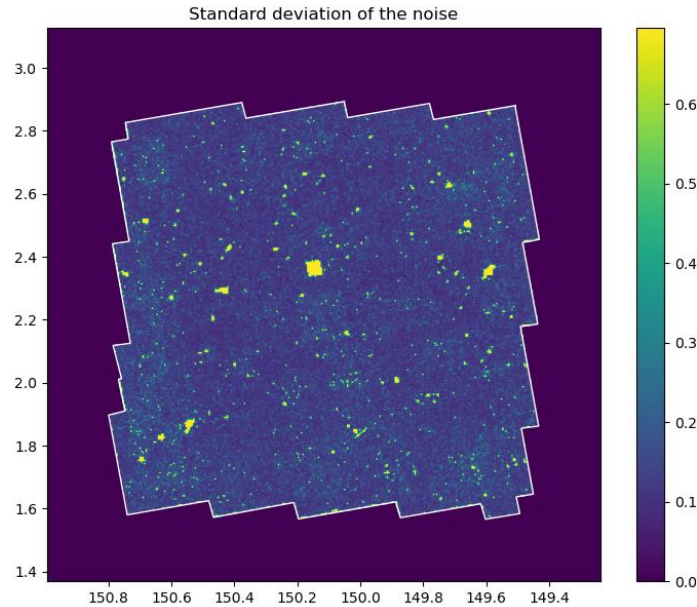


- $\mathbf{r}[k] := \Phi_k^{-1}(1 - \alpha/2)$  with  $\Phi_k$  Gaussian CDF at pixel  $k$ ;
- $\hat{\kappa}^- := \hat{\kappa} - \mathbf{r}$  and  $\hat{\kappa}^+ := \hat{\kappa} + \mathbf{r}$ ;  
 $\rightarrow \mathbb{P}\{\mathbf{S}\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]]\} \leq \alpha$  (approximately).
- $\hat{\kappa}$  unbiased estimator of  $\mathbf{S}\kappa$ , but biased estimator of  $\kappa$ . So are the bounds  $\hat{\kappa}^-$  and  $\hat{\kappa}^+$   
 $\rightarrow \mathbb{P}\{\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]]\} \leq \alpha??$

# Kaiser-Squires bound estimation

Random variable centered in  $\mathbf{S}\boldsymbol{\kappa}$

Estimation of  $\hat{\boldsymbol{\kappa}}^-$  and  $\hat{\boldsymbol{\kappa}}^+$ : we have  $\hat{\boldsymbol{\kappa}} \sim N(\mathbf{S}\boldsymbol{\kappa}, \mathbf{S}\boldsymbol{\Lambda}^\dagger \boldsymbol{\Sigma} \boldsymbol{\Lambda}^\dagger \mathbf{S}^*)$ .

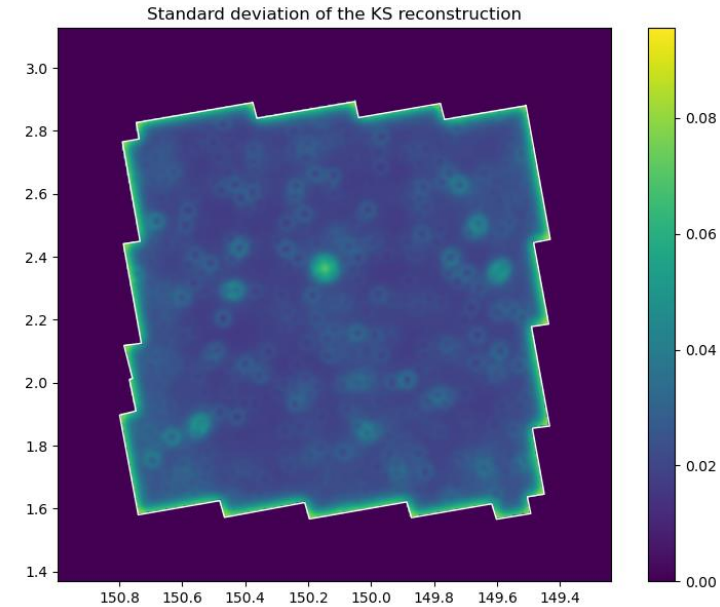
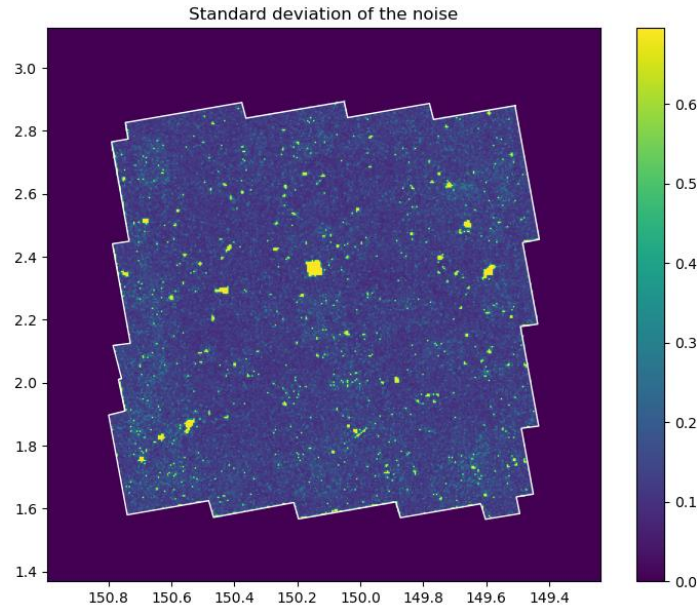


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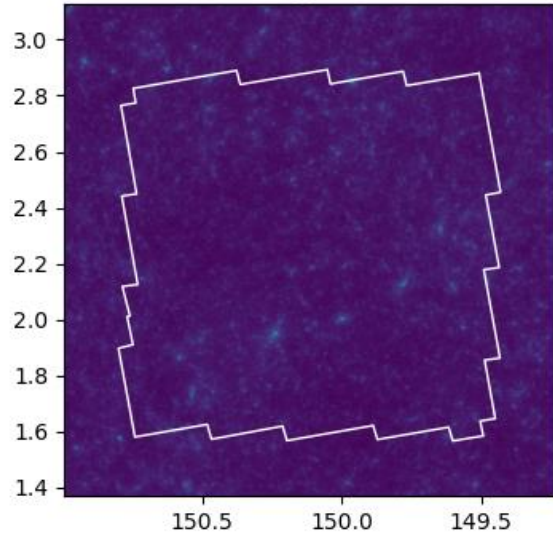
$$\rightarrow \mathbb{P}\{\boldsymbol{\kappa}[k] \notin [\hat{\boldsymbol{\kappa}}^-[k], \hat{\boldsymbol{\kappa}}^+[k]]\} \leq \alpha??$$

Calibration needed!

# Kaiser-Squires bound estimation

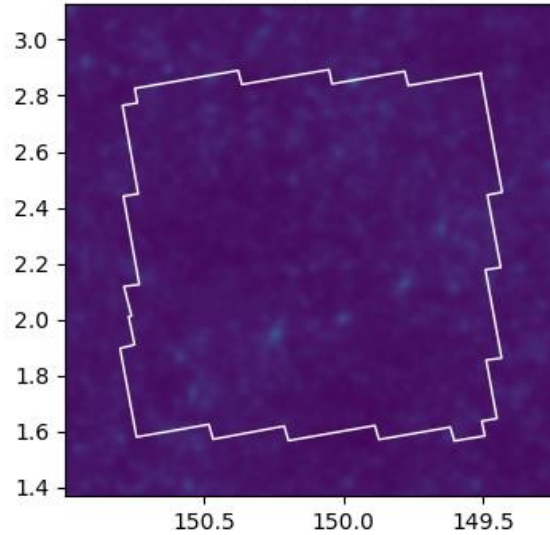
Target:  $\alpha \approx 4,6\%$  ( $2\sigma$ -confidence)

Original



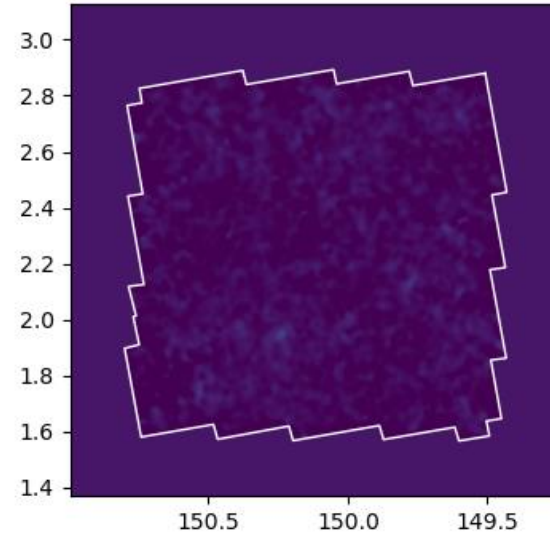
$\kappa$

Original (smoothed)



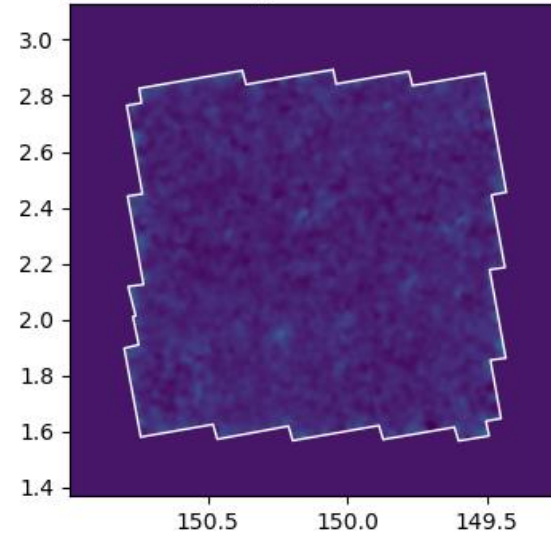
$S\kappa$

Lower bound



$\hat{\kappa}^-$

Upper bound

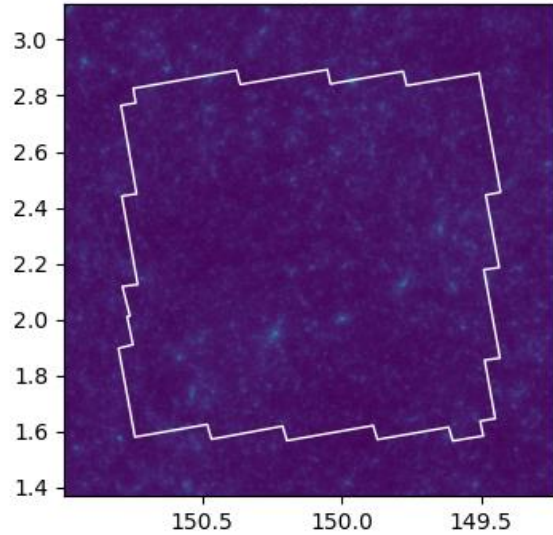


$\hat{\kappa}^+$

# Kaiser-Squires bound estimation

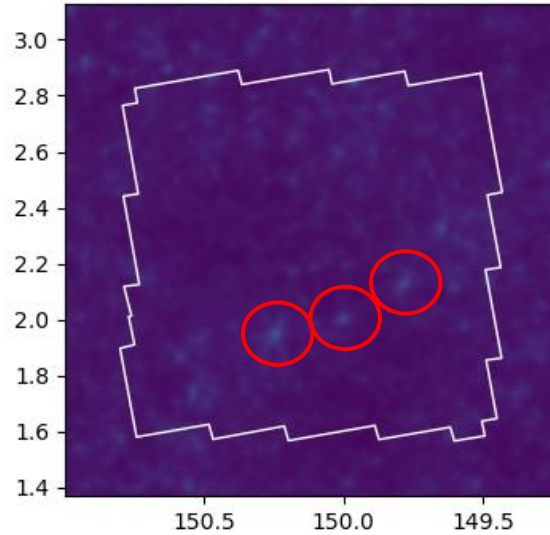
Target:  $\alpha \approx 4,6\%$  ( $2\sigma$ -confidence)

Original



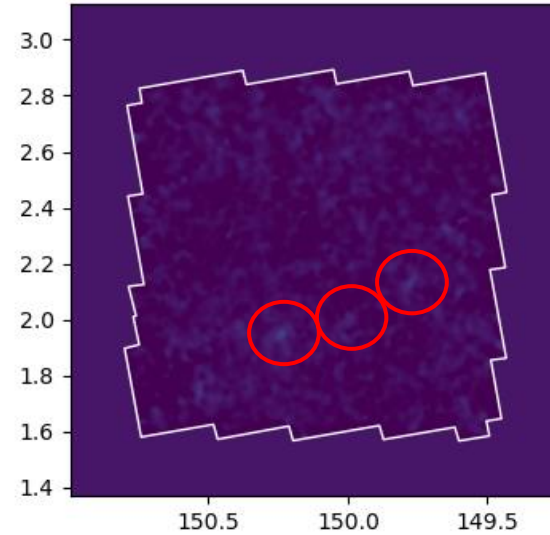
$\kappa$

Original (smoothed)



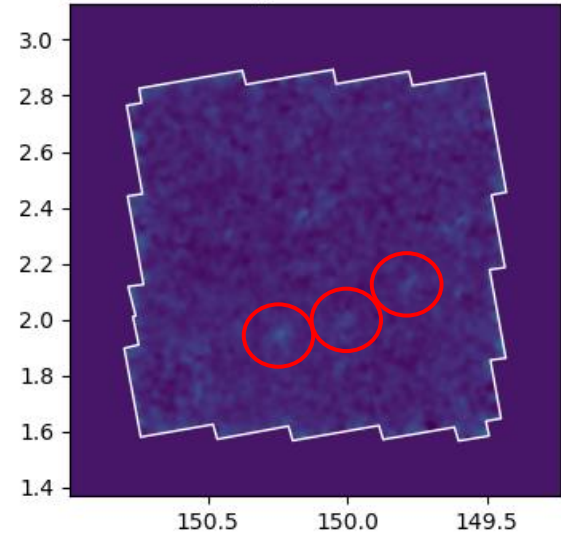
$S\kappa$

Lower bound



$\hat{\kappa}^-$

Upper bound



$\hat{\kappa}^+$

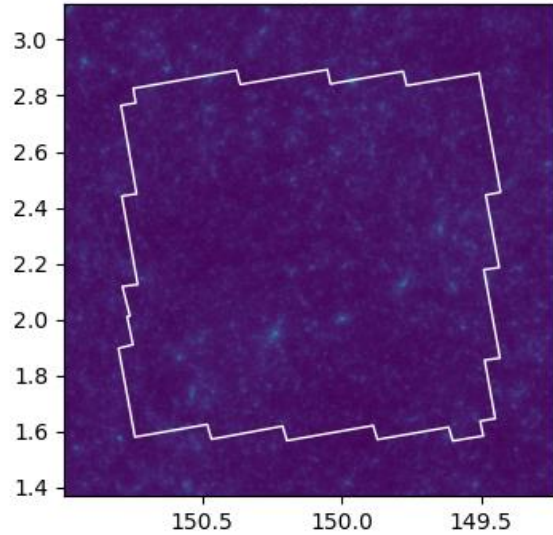
Correct identification of  
mass overdensities

# Kaiser-Squires bound estimation

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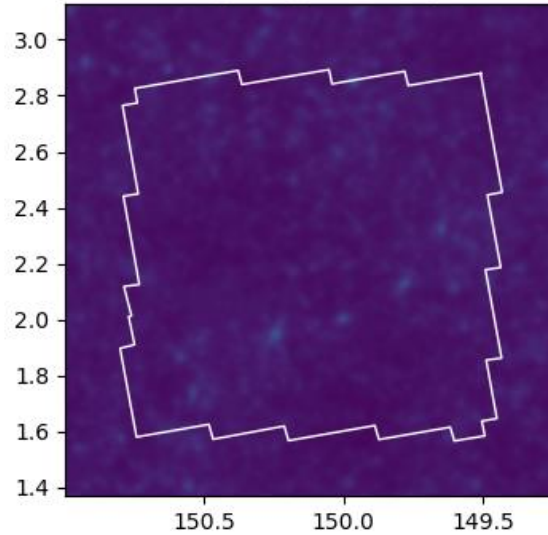
High level of uncertainty  
due to masked data

Original



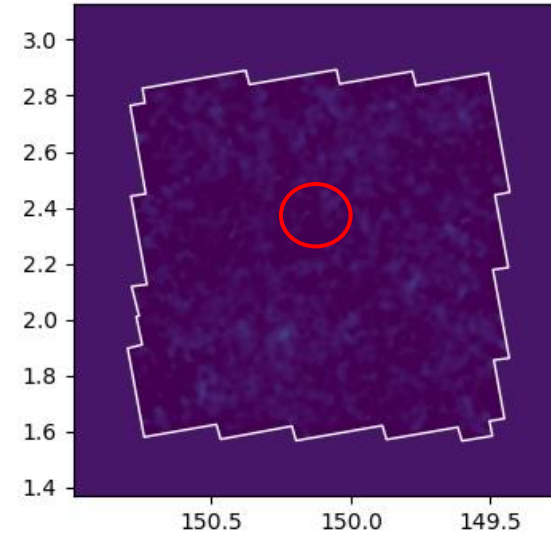
$\kappa$

Original (smoothed)



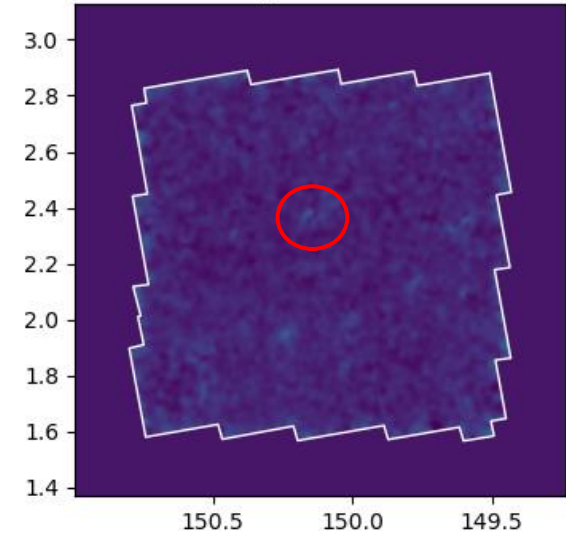
$S\kappa$

Lower bound



$\hat{\kappa}^-$

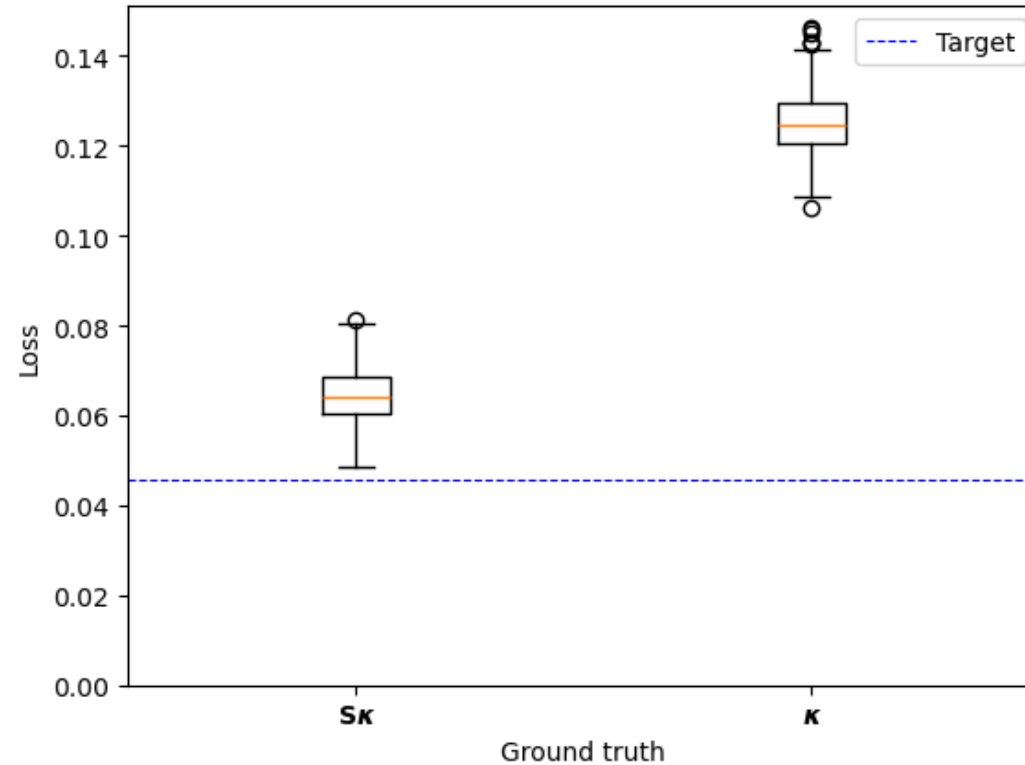
Upper bound



$\hat{\kappa}^+$

# Kaiser-Squires bound estimation

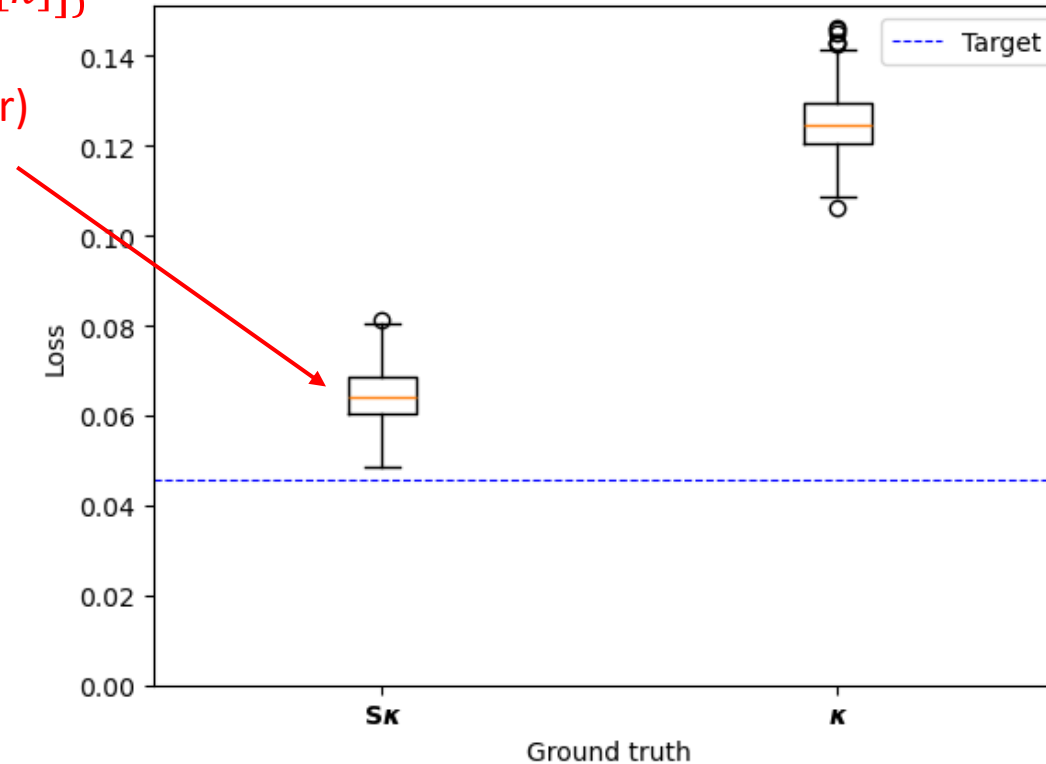
$\alpha \approx 4,6\%$  ( $2\sigma$ -confidence)



# Kaiser-Squires bound estimation

$\alpha \approx 4,6\%$  ( $2\sigma$ -confidence)

$P\{\mathbf{S}\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]]\}$   
(empirical)  
(unbiased estimator)

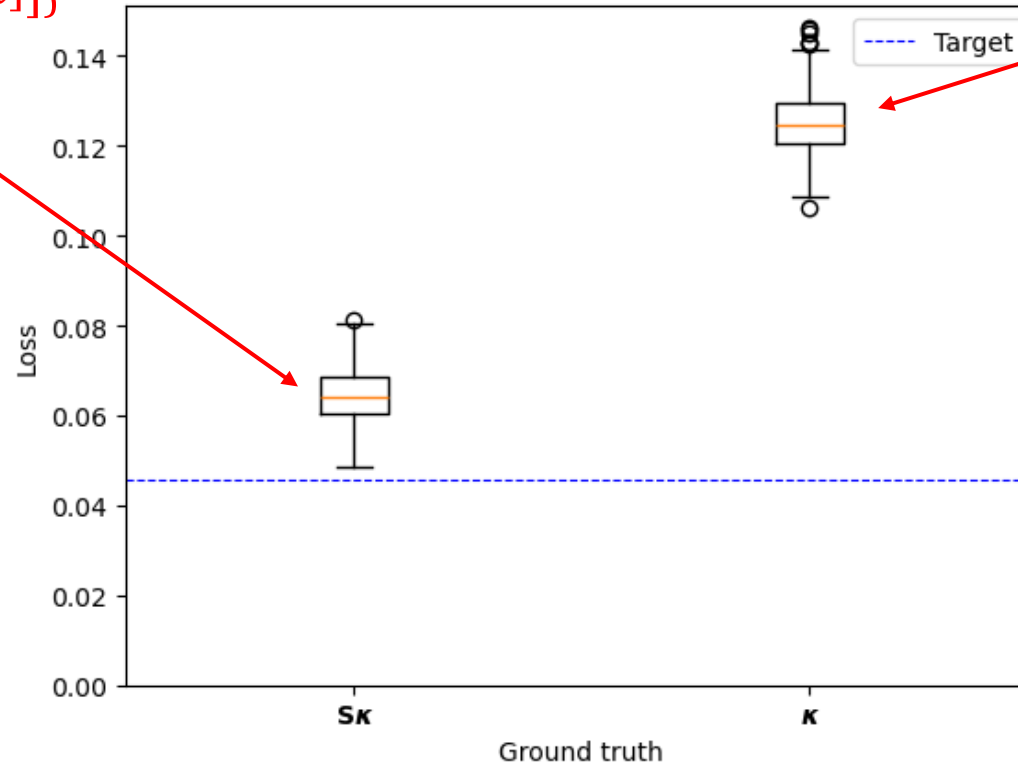




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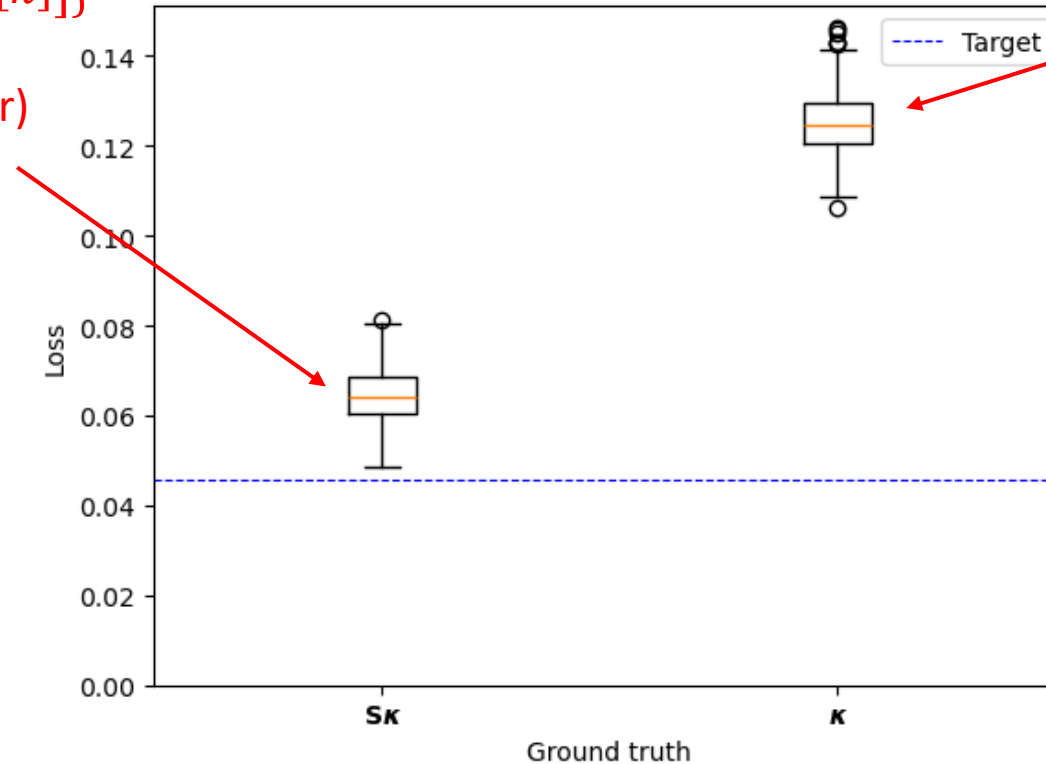


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$\alpha \approx 4,6\%$  ( $2\sigma$ -confidence)

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(empirical)  
(unbiased estimator)



$P\{\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]]\}$   
(empirical)  
(biased estimator)

→ Undercoverage

# Calibration procedure

**Objective (reminder):** given  $\hat{\gamma}$ , estimate  $\hat{\kappa}^-$  and  $\hat{\kappa}^+$  such that

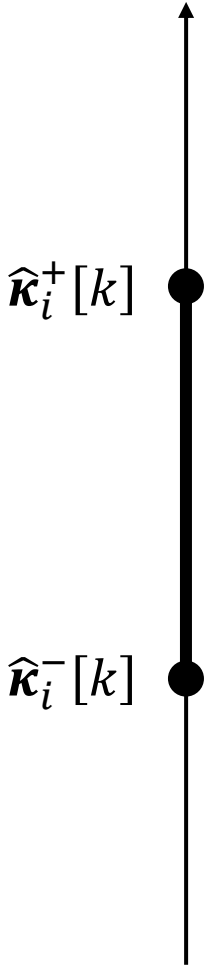
$$P\{\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]]\} \leq \alpha.$$

**Conformalized quantile regression (CQR):**<sup>1</sup> performed on a **calibration set**  $(\hat{\gamma}_i, \kappa_i)_{i=1}^n$ . For each pixel  $k$ :

1. compute a prediction score on each calibration example:

$$e_i[k] := \max\{\hat{\kappa}_i^-[k] - \kappa_i[k], \kappa_i[k] - \hat{\kappa}_i^+[k]\};$$

2. get the  $(1 - \alpha)$ -quantile of  $(e_i[k])_{i=1}^n$ , denoted by  $\mathbf{q}_{(1-\alpha)}[k]$ ;
3. adjust the bounds: set  $\hat{\kappa}^- \leftarrow \hat{\kappa}^- - \mathbf{q}_{(1-\alpha)}$  and  $\hat{\kappa}^+ \leftarrow \hat{\kappa}^+ + \mathbf{q}_{(1-\alpha)}$ .



<sup>1</sup> Y. Romano, E. Patterson, and E. Candès, “Conformalized Quantile Regression,” in *NeurIPS*, 2019.

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Ground truth inside prediction bounds  $\rightarrow e_i[k] < 0$

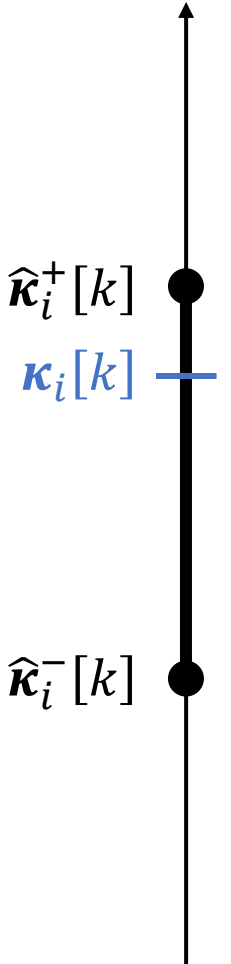
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# Calibration procedure

Ground truth outside prediction bounds  $\rightarrow e_i[k] > 0$

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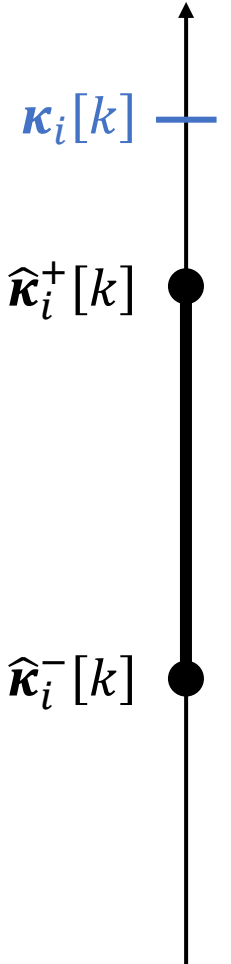
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2. get the  $(1 - \alpha)$ -quantile of  $(e_i[k])_{i=1}^n$ , denoted by  $q_{(1-\alpha)}[k]$ ;  
Well-calibrated  $\rightarrow q_{(1-\alpha)}[k] = 0$
3. adjust the bounds: set  $\hat{\kappa}^- \leftarrow \hat{\kappa}^- - q_{(1-\alpha)}$  and  $\hat{\kappa}^+ \leftarrow \hat{\kappa}^+ + q_{(1-\alpha)}$   
Undercoverage  $\rightarrow q_{(1-\alpha)}[k] > 0$   
Overcoverage  $\rightarrow q_{(1-\alpha)}[k] < 0$

$\hat{\kappa}_i^+[k]$

$\hat{\kappa}_i^-[k]$

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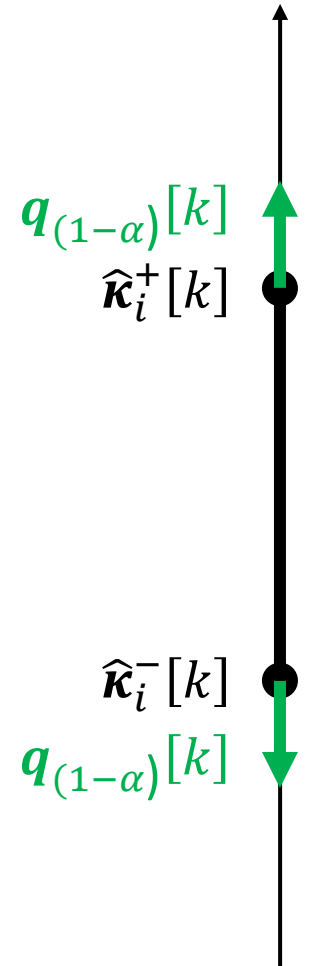
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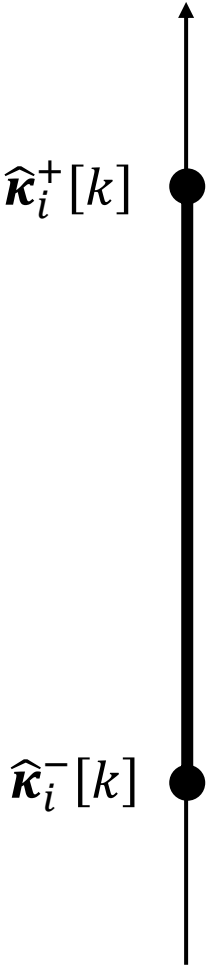
**Conformalized quantile regression (CQR):**<sup>1</sup> performed on a **calibration set**  $(\hat{\gamma}_i, \kappa_i)_{i=1}^n$ . For each pixel  $k$ :

1. compute a prediction score on each calibration example:

$$e_i[k] := \max\{\hat{\kappa}_i^-[k] - \kappa_i[k], \kappa_i[k] - \hat{\kappa}_i^+[k]\};$$

2. get the  $(1 - \alpha)$ -quantile of  $(e_i[k])_{i=1}^n$ , denoted by  $\mathbf{q}_{(1-\alpha)}[k]$ ;
3. adjust the bounds: set  $\hat{\kappa}^- \leftarrow \hat{\kappa}^- - \mathbf{q}_{(1-\alpha)}$  and  $\hat{\kappa}^+ \leftarrow \hat{\kappa}^+ + \mathbf{q}_{(1-\alpha)}$ .

**THEOREM:**<sup>1</sup>  $\alpha - 1/n_{+1} \leq P\{\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]] \mid (\hat{\gamma}_i, \kappa_i)_{i=1}^n\} \leq \alpha$  for any pixel  $k$ .



<sup>1</sup> Y. Romano, E. Patterson, and E. Candès, “Conformalized Quantile Regression,” in *NeurIPS*, 2019.

# Calibration procedure

**Objective (reminder):** given  $\hat{\gamma}$ , estimate  $\hat{\kappa}^-$  and  $\hat{\kappa}^+$  such that

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Conditionally to a specific  
calibration set

<sup>1</sup> Y. Romano, E. Patterson, and E. Candès, “Conformalized Quantile Regression,” in *NeurIPS*, 2019.

# Calibration procedure

**Objective (reminder):** given  $\hat{\gamma}$ , estimate  $\hat{\kappa}^-$  and  $\hat{\kappa}^+$  such that

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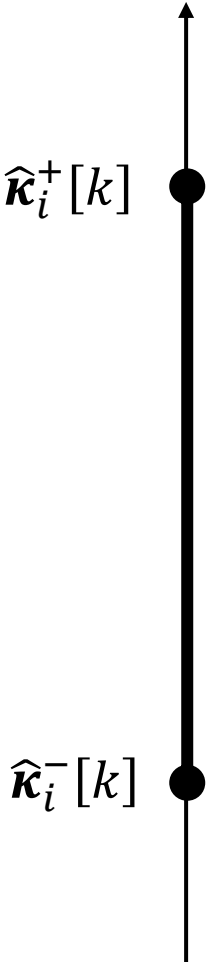
1. compute a prediction score on each calibration example:

$$e_i[k] := \max\{\hat{\kappa}_i^-[k] - \kappa_i[k], \kappa_i[k] - \hat{\kappa}_i^+[k]\};$$

2. get the  $(1 - \alpha)$ -quantile of  $(e_i[k])_{i=1}^n$ , denoted by  $q_{(1-\alpha)}[k]$ ;
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Upper bound: coverage guarantee



<sup>1</sup> Y. Romano, E. Patterson, and E. Candès, “Conformalized Quantile Regression,” in *NeurIPS*, 2019.

# Calibration procedure

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**THEOREM:**<sup>1</sup>  $\alpha - 1/n+1 \leq P\{\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]] \mid (\hat{\gamma}_i, \kappa_i)_{i=1}^n\} \leq \alpha$  for any pixel  $k$ .

Lower bound: prevents  
overconservative prediction bounds

$\hat{\kappa}_i^+[k]$

$\hat{\kappa}_i^-[k]$

<sup>1</sup> Y. Romano, E. Patterson, and E. Candès, “Conformalized Quantile Regression,” in *NeurIPS*, 2019.

# Calibration procedure

**Objective (reminder):** given  $\hat{\gamma}$ , estimate  $\hat{\kappa}^-$  and  $\hat{\kappa}^+$  such that

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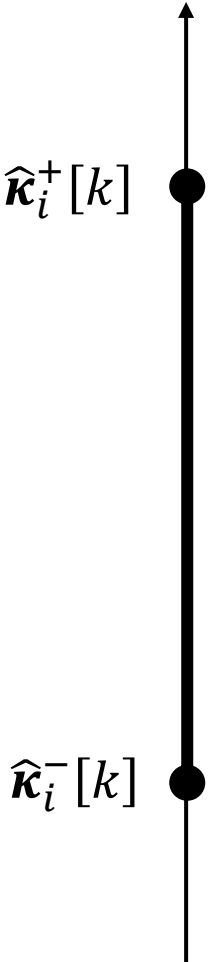
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**Works for any blackbox quantile predictor!**

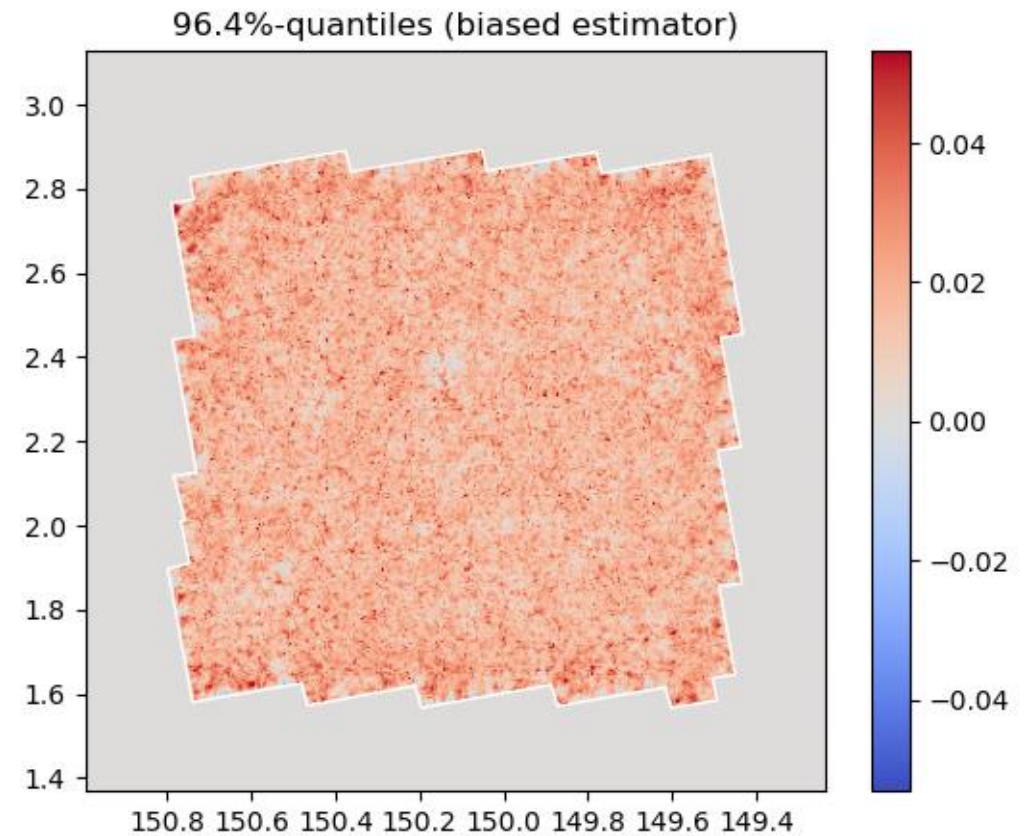
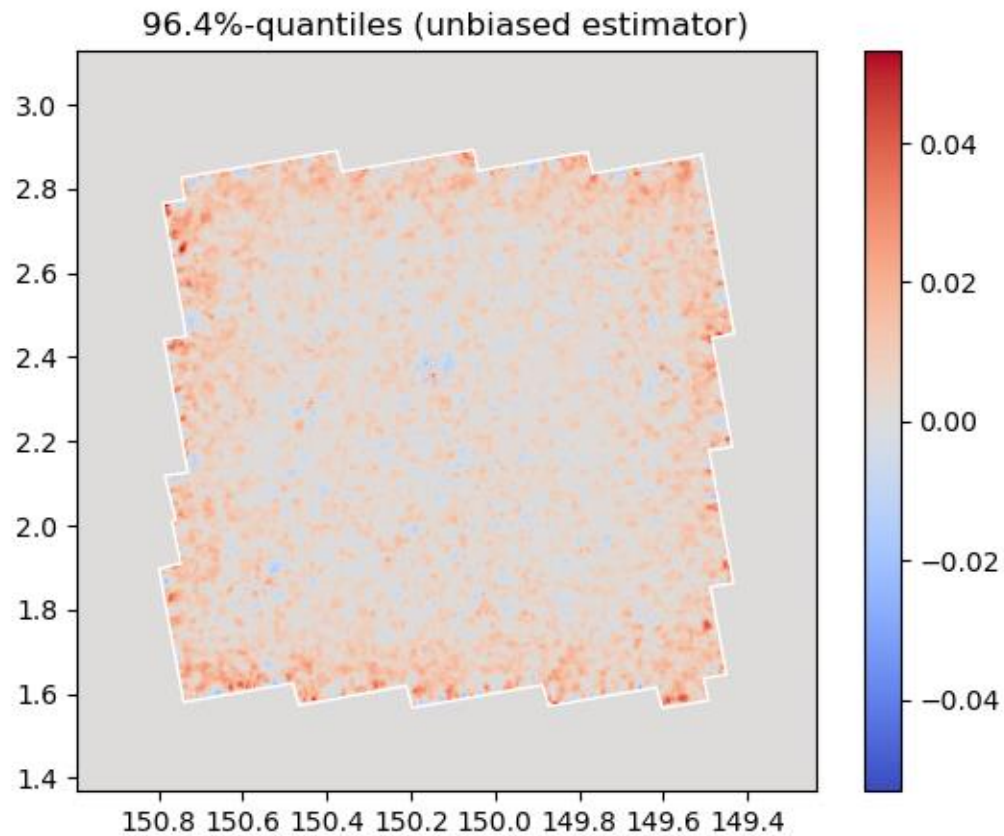


<sup>1</sup> Y. Romano, E. Patterson, and E. Candès, “Conformalized Quantile Regression,” in *NeurIPS*, 2019.

# CQR calibration

Target:  $\alpha \approx 4,6\%$  ( $2\sigma$ -confidence)

$n = 100$  (size of the calibration set)

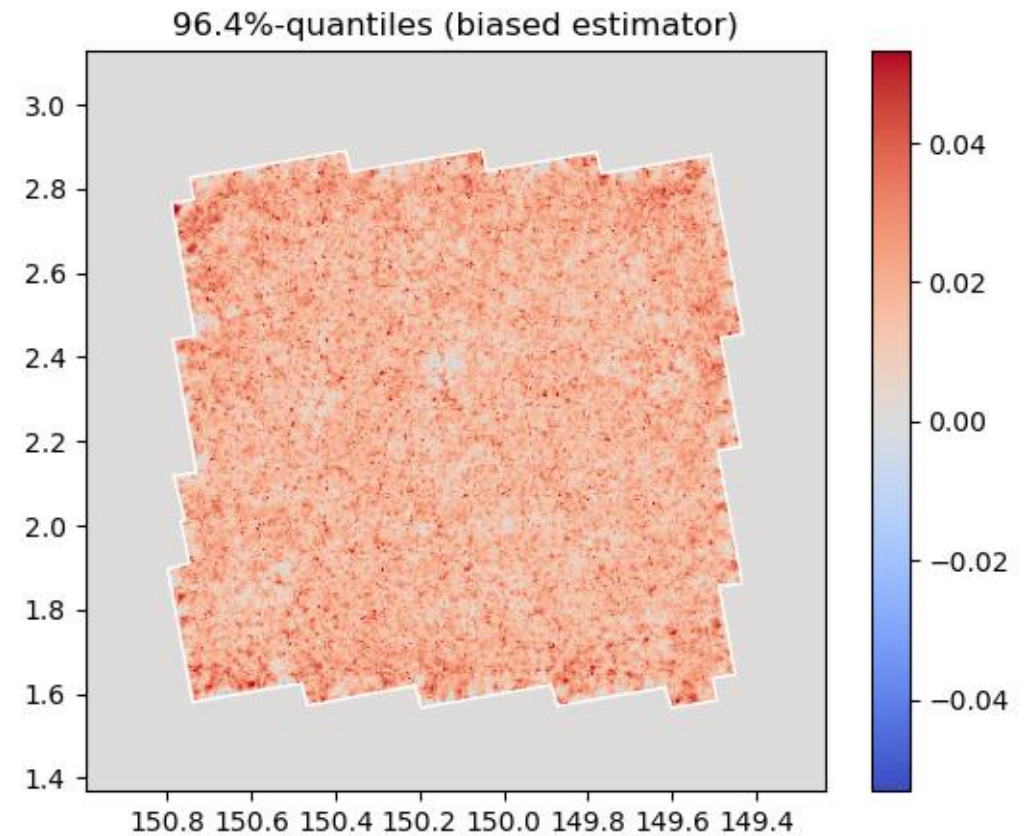
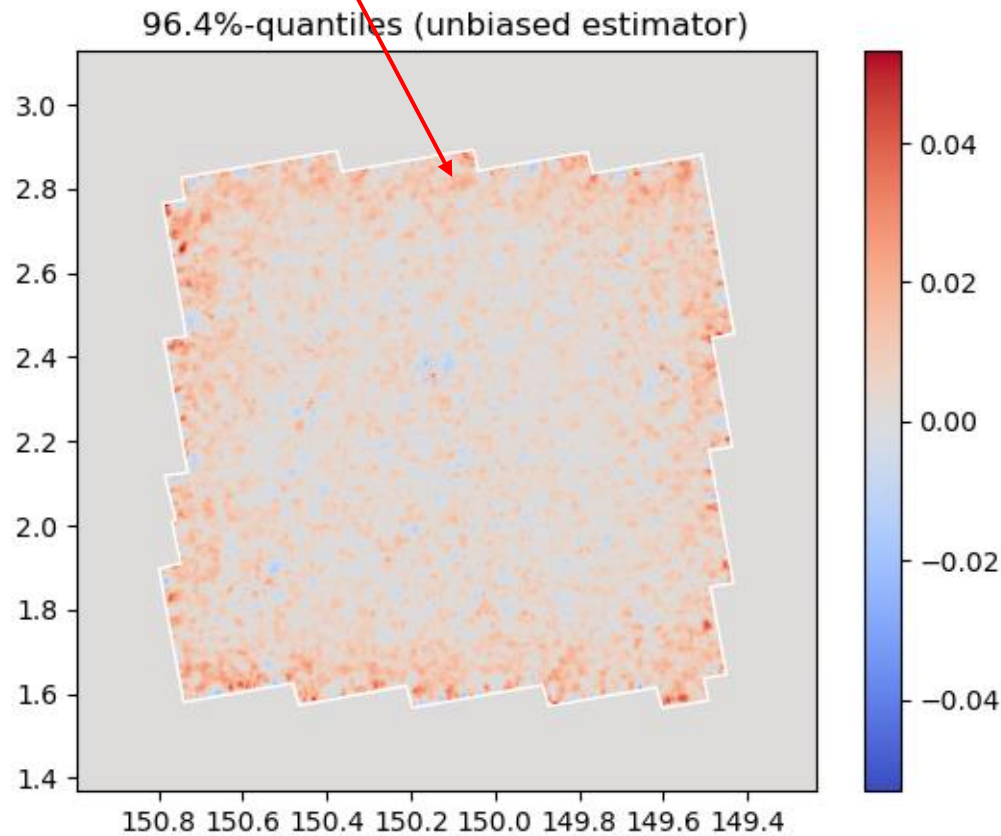


# CQR calibration

Target:  $\alpha \approx 4,6\%$  ( $2\sigma$ -confidence)

Undercoverage near the edges

$n = 100$  (size of the calibration set)

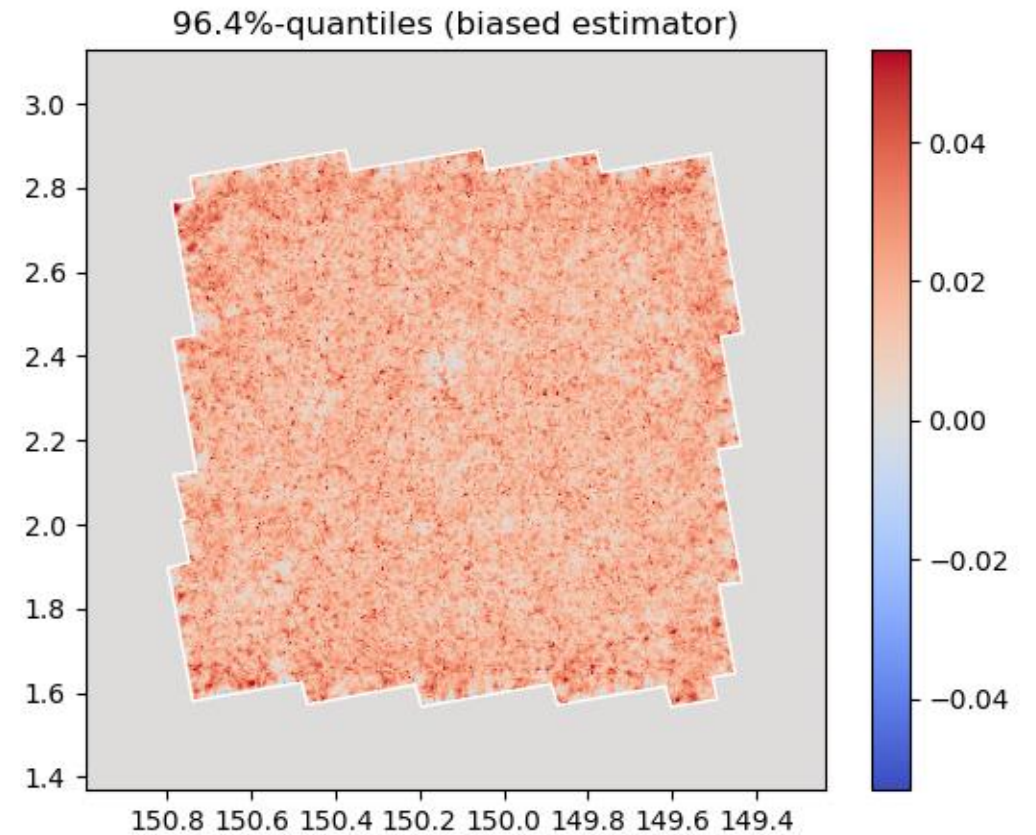
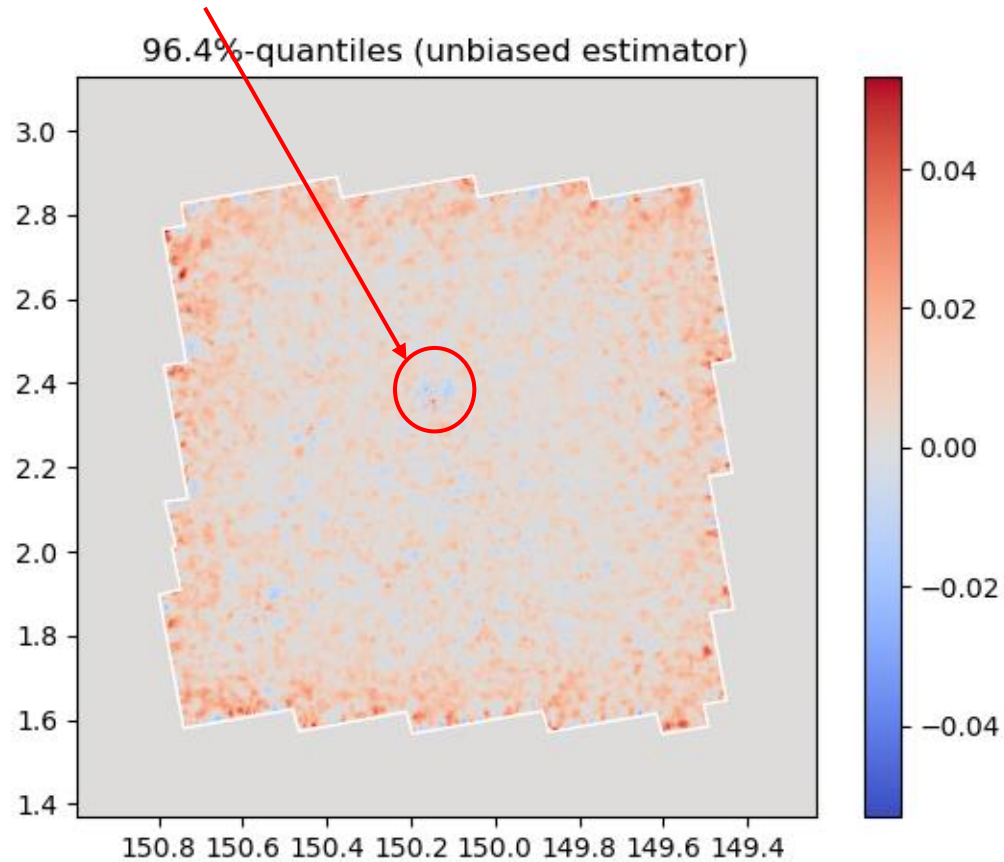


# CQR calibration

Target:  $\alpha \approx 4,6\%$  ( $2\sigma$ -confidence)

$n = 100$  (size of the calibration set)

Overcoverage near the  
masked regions

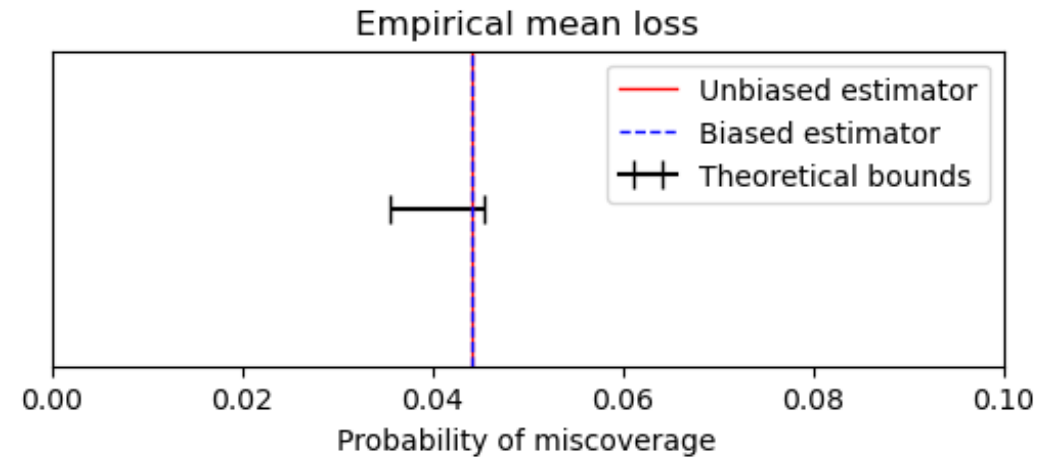
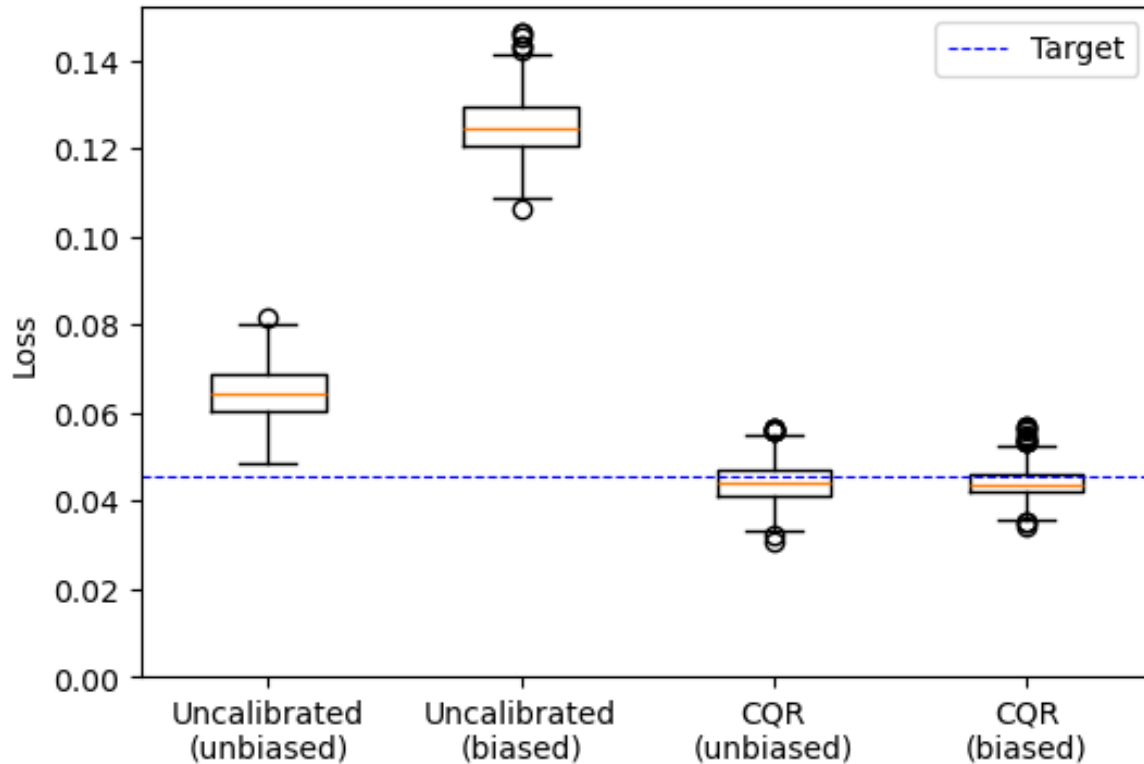




# CQR calibration – losses

Target:  $\alpha \approx 4,6\%$  ( $2\sigma$ -confidence)

$n = 100$  (size of the calibration set)

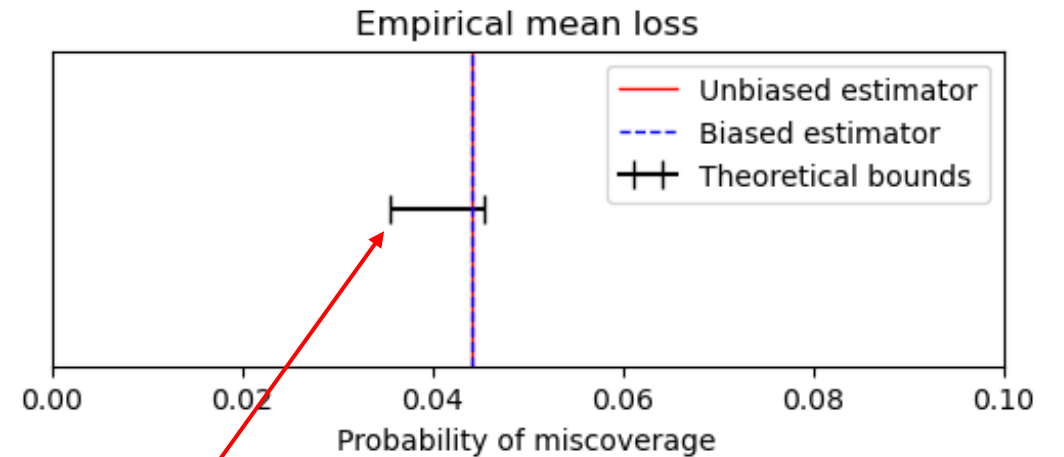
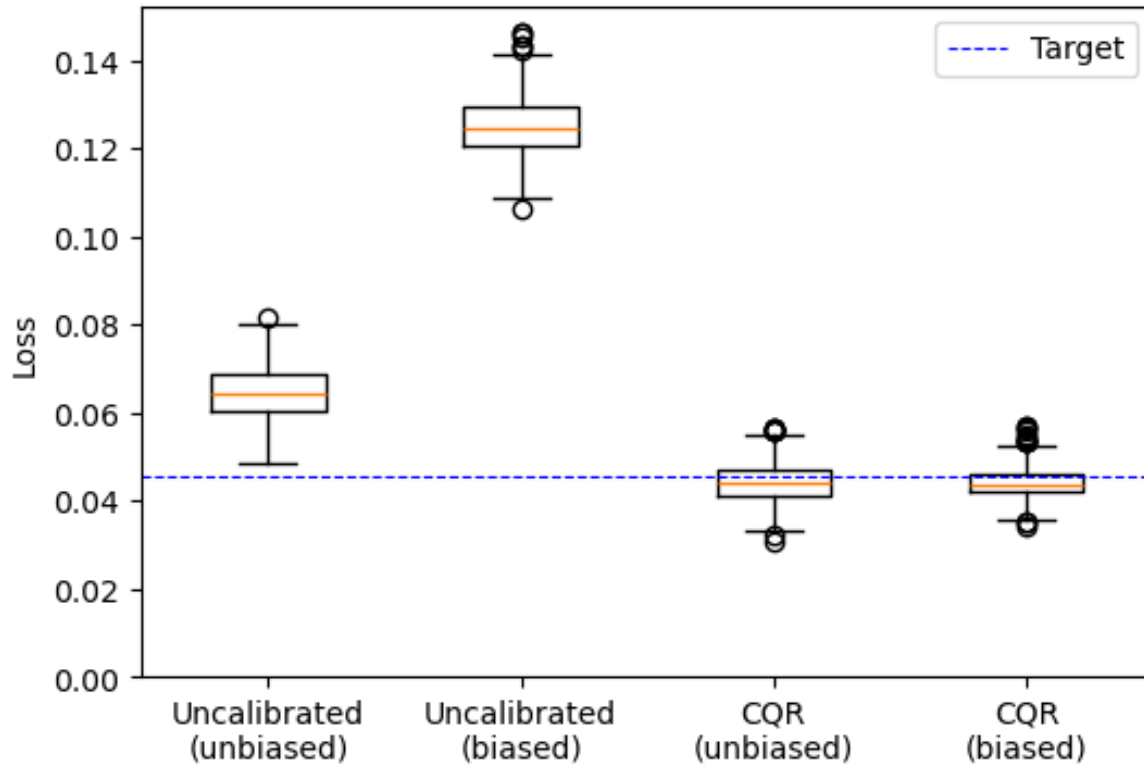


$$\alpha - \frac{1}{n+1} \leq P\{\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]] \mid (\hat{\mathcal{Y}}_i, \kappa_i)_{i=1}^n\} \leq \alpha$$

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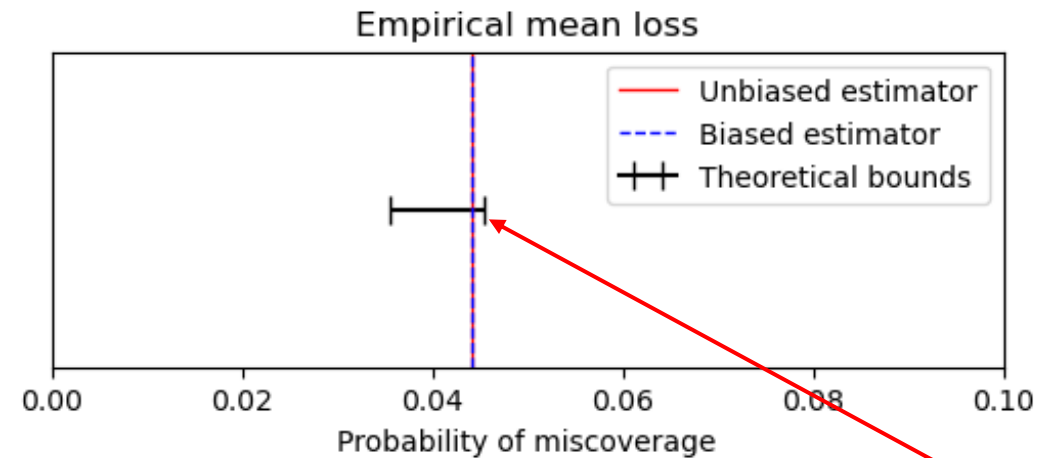
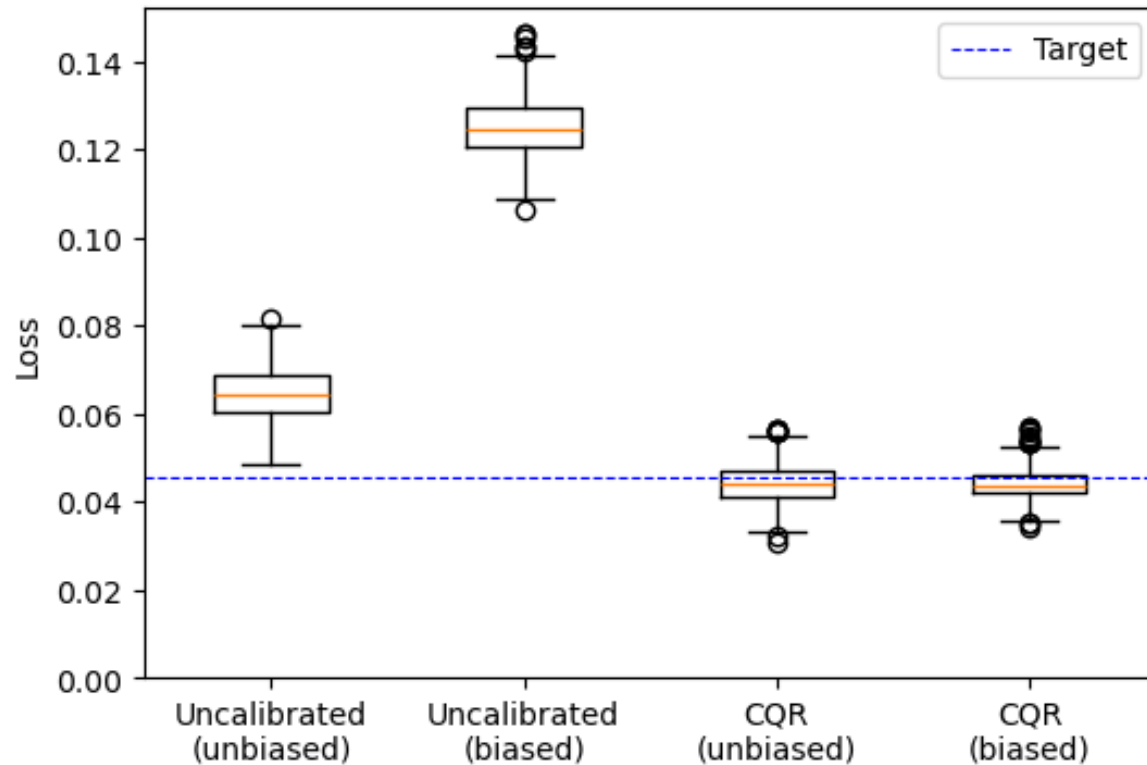


$$\alpha - \frac{1}{n+1} \leq P\{\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]] | (\hat{\gamma}_i, \kappa_i)_{i=1}^n\} \leq \alpha$$

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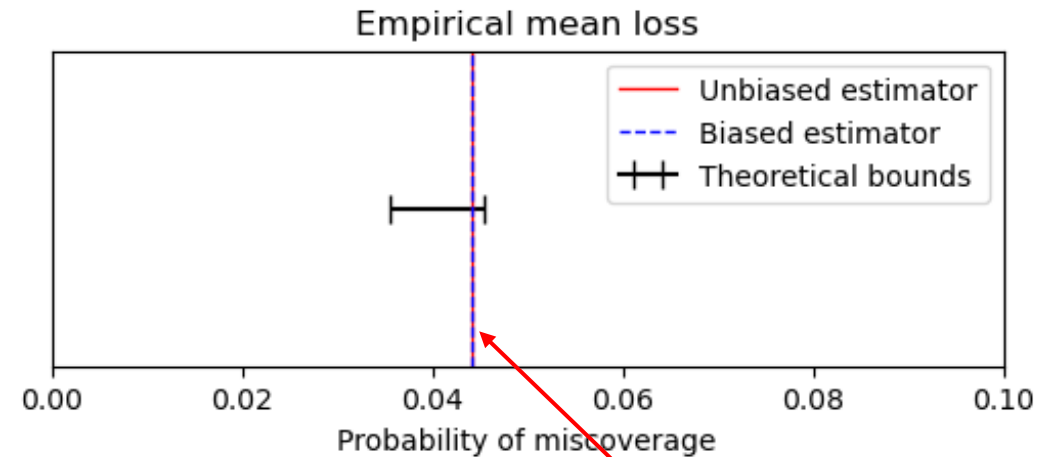
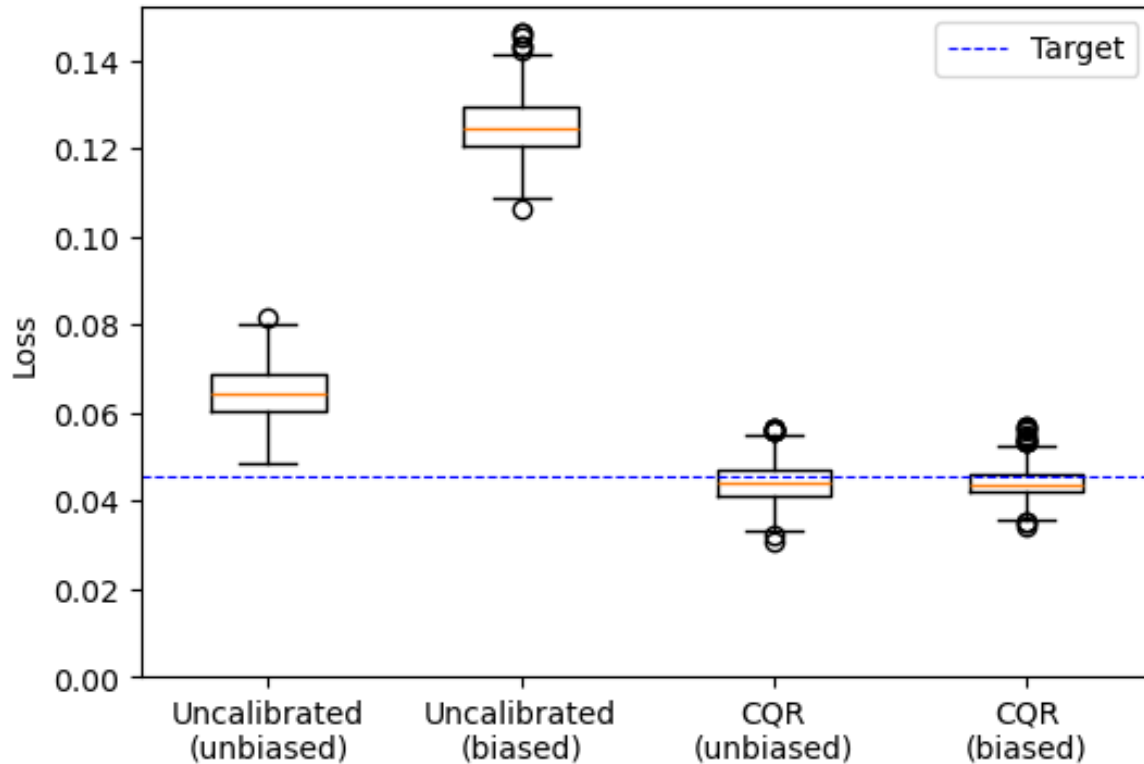


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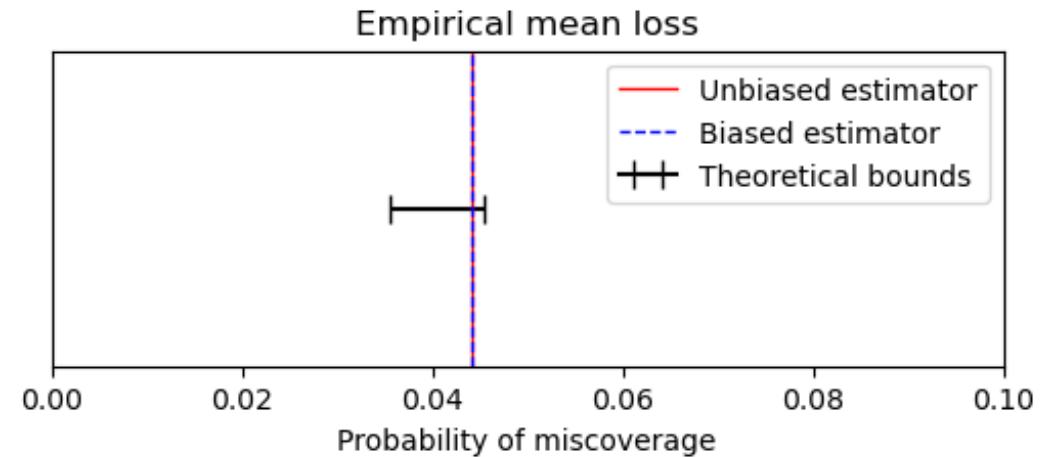
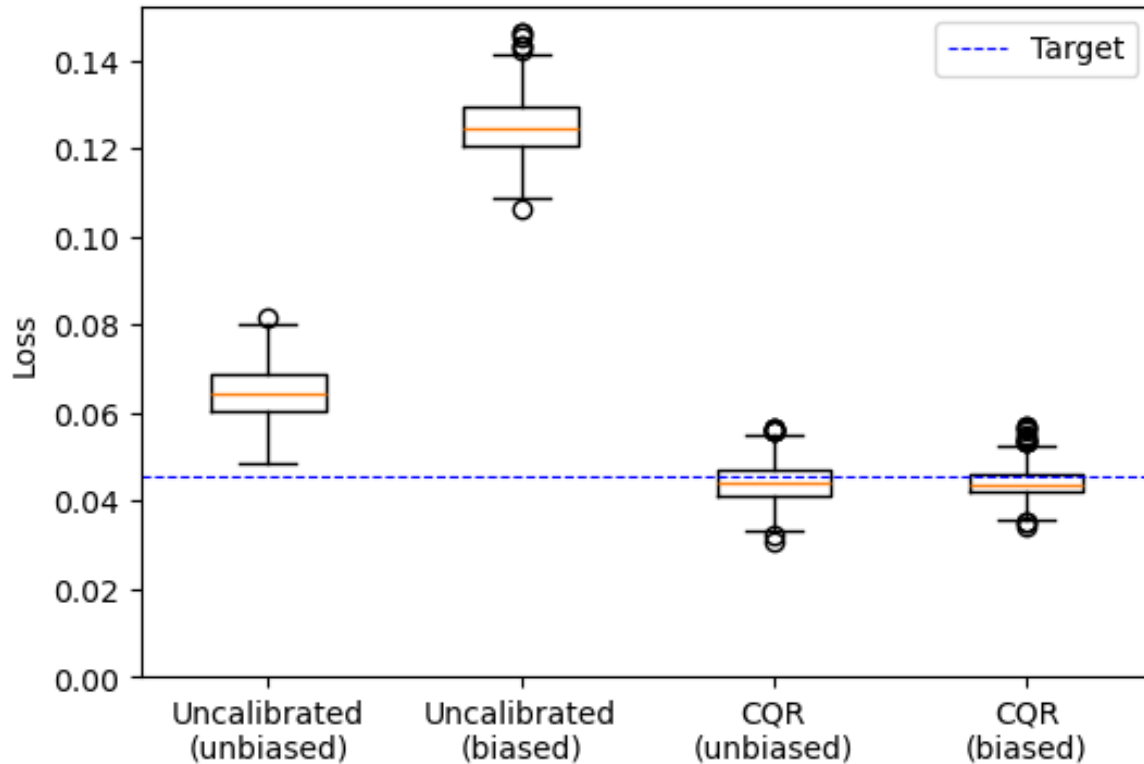
$n = 100$  (size of the calibration set)



$$\alpha - \frac{1}{n+1} \leq P\{\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]] | (\hat{\gamma}_i, \kappa_i)_{i=1}^n\} \leq \alpha$$

# CQR calibration – losses

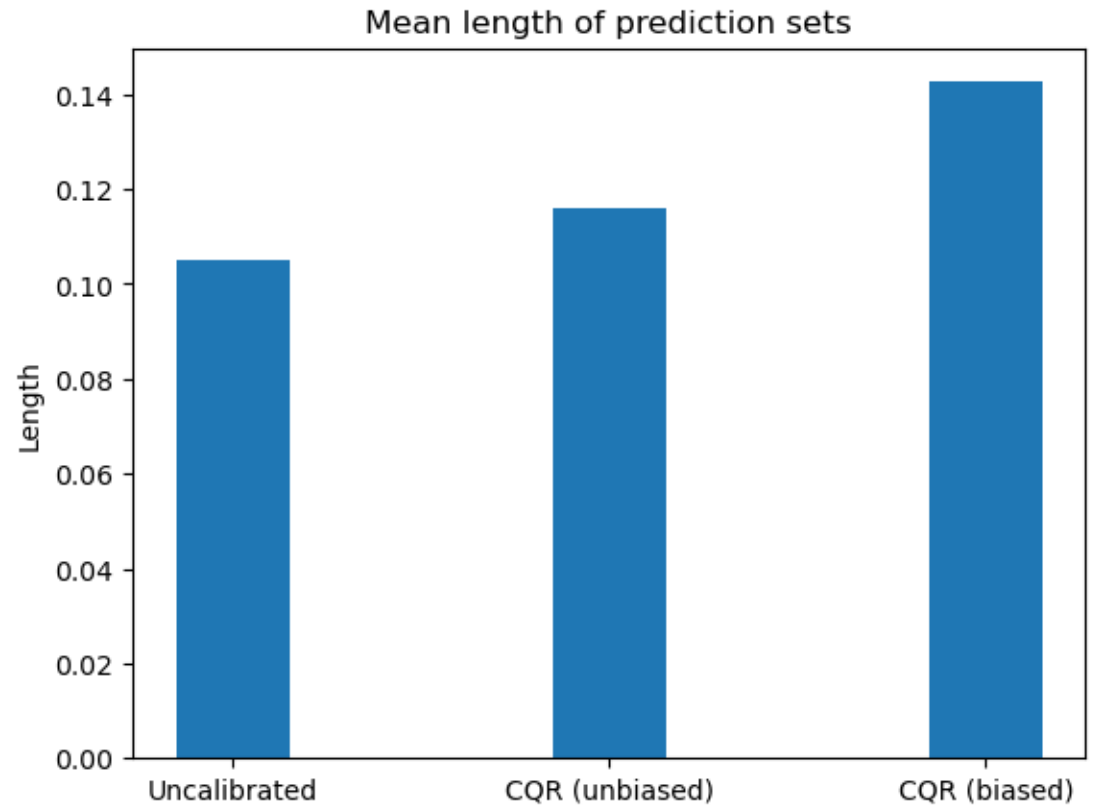
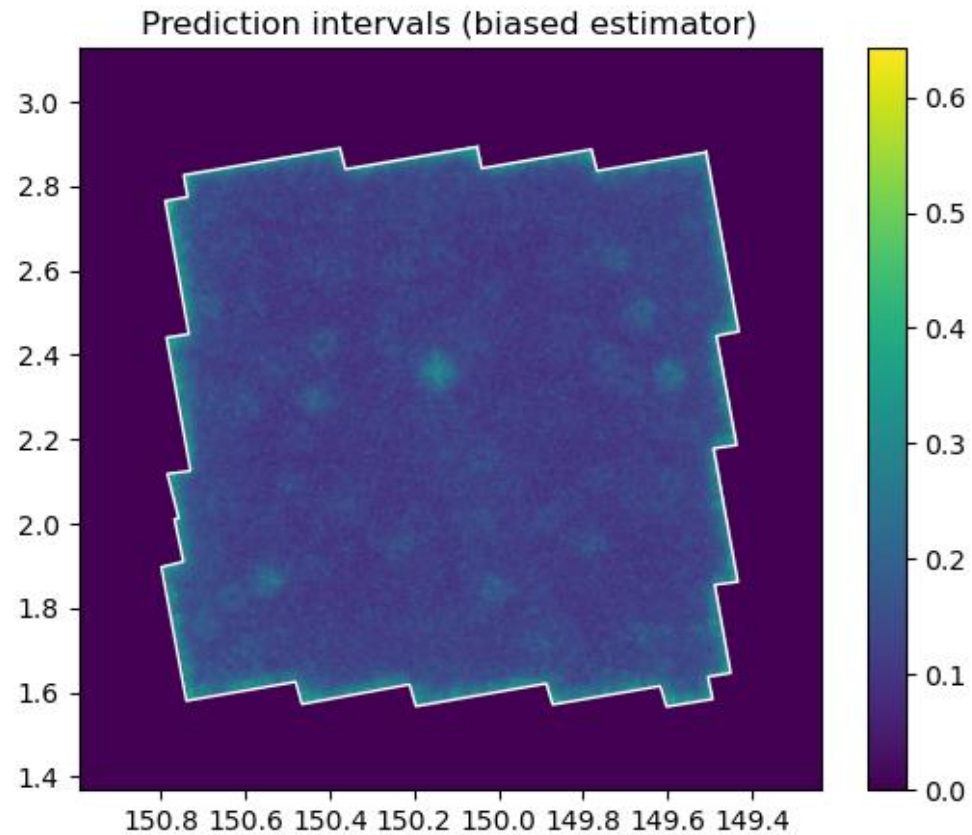
Target:  $\alpha \approx 4,6\%$  ( $2\sigma$ -confidence)  
 $n = 100$  (size of the calibration set)



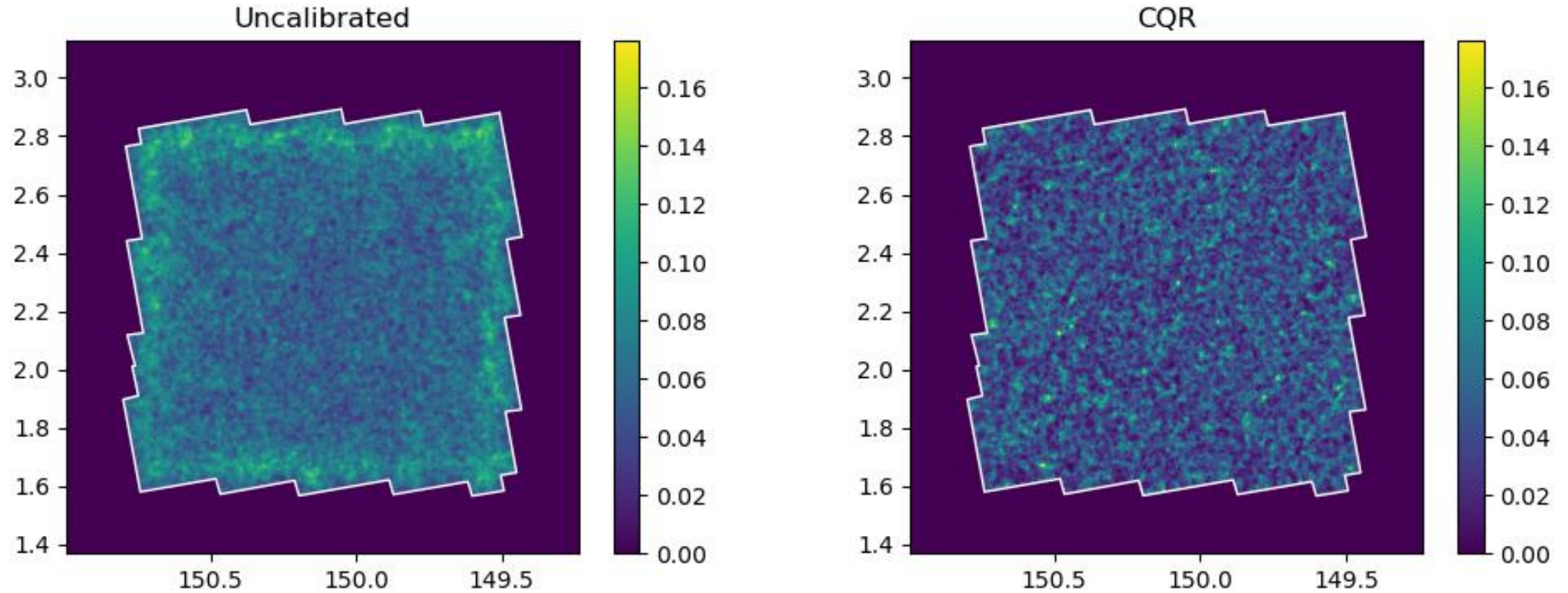
$$\alpha - \frac{1}{n+1} \leq P\{\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]] \mid (\hat{\gamma}_i, \kappa_i)_{i=1}^n\} \leq \alpha$$

→ These experiments are in line with the theoretical results.

# CQR calibration – length of prediction sets

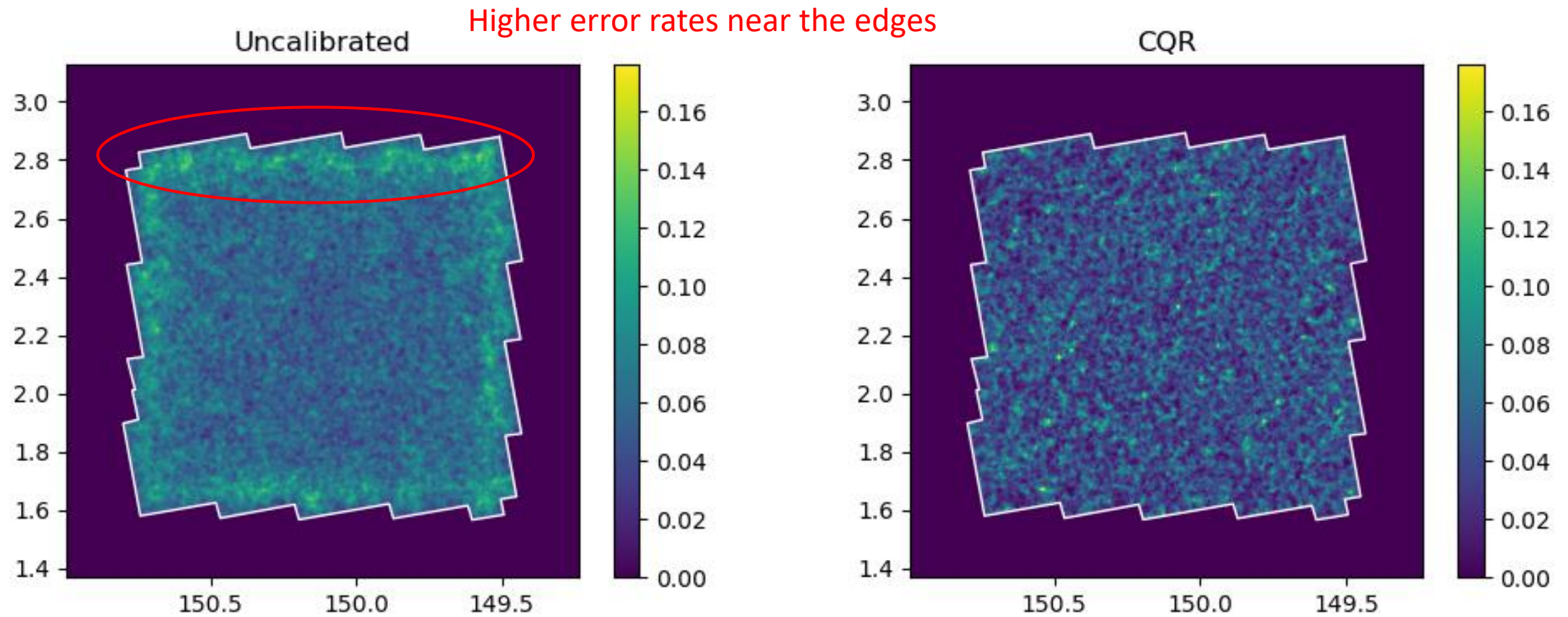


# CQR calibration – error rate per pixel



Ground truth =  $\mathbf{S}\boldsymbol{\kappa}$  (unbiased estimator)

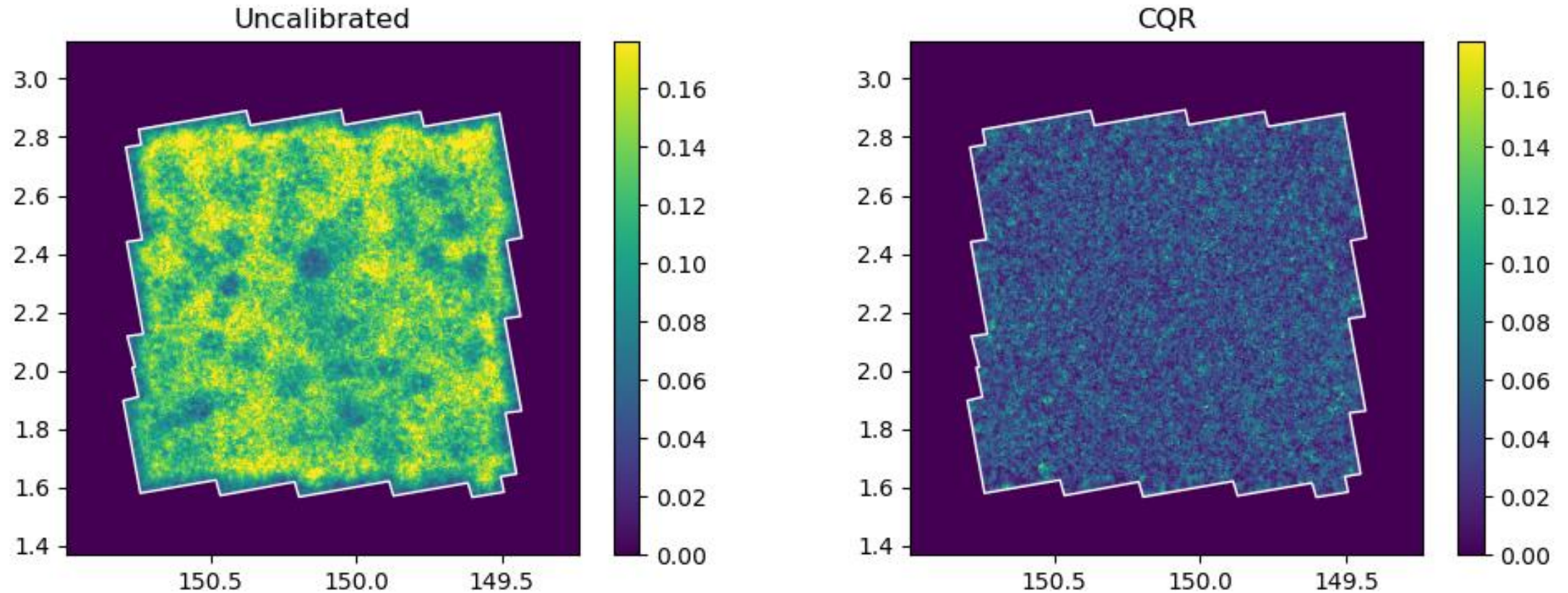
# CQR calibration – error rate per pixel



Ground truth =  $\mathbf{S}\boldsymbol{\kappa}$  (unbiased estimator)



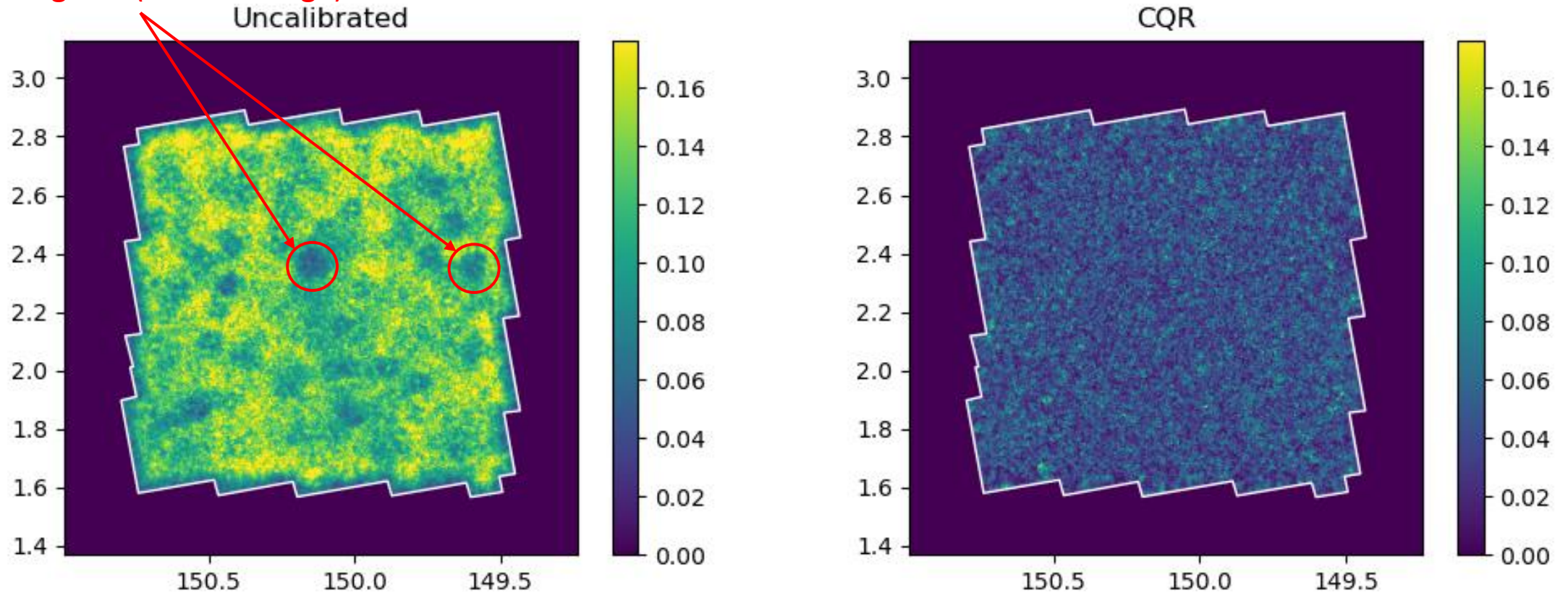
# CQR calibration – error rate per pixel



Ground truth =  $\kappa$  (biased estimator)

# CQR calibration – error rate per pixel

Lower error rates around  
masked regions (overcoverage)



Ground truth =  $\kappa$  (biased estimator)

# CQR calibration – still biased??

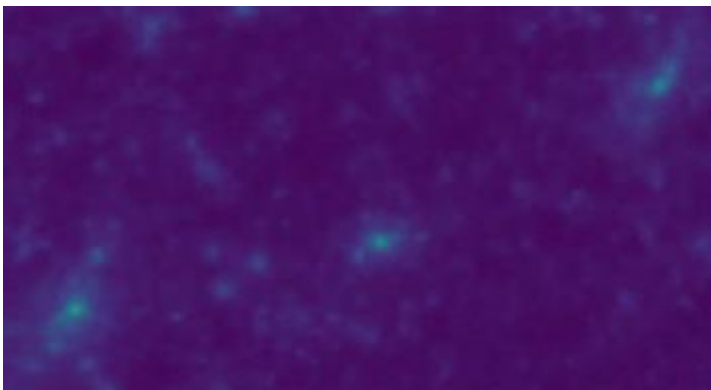
- Back to the theoretical guarantee:

$$\alpha - 1/n_{+1} \leq P\{\boldsymbol{\kappa}[k] \notin [\hat{\boldsymbol{\kappa}}^-[k], \hat{\boldsymbol{\kappa}}^+[k]] \mid (\hat{\boldsymbol{\gamma}}_i, \boldsymbol{\kappa}_i)_{i=1}^n\} \leq \alpha.$$

- Conditioned on the calibration set, but NOT on the neighboring pixels. E.g.,

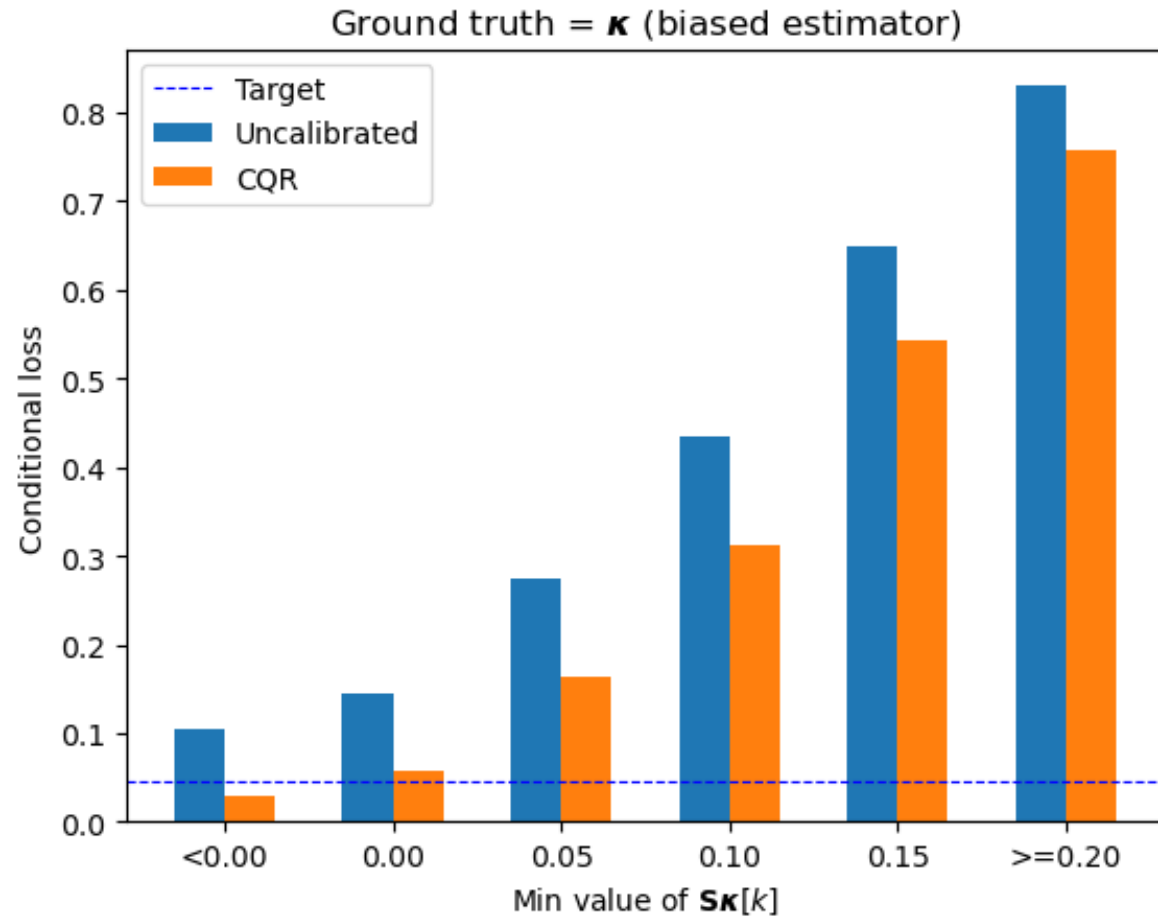
$$\alpha - 1/n_{+1} \leq P\{\boldsymbol{\kappa}[k] \notin [\hat{\boldsymbol{\kappa}}^-[k], \hat{\boldsymbol{\kappa}}^+[k]] \mid \mathbf{S}\boldsymbol{\kappa}[k], (\hat{\boldsymbol{\gamma}}_i, \boldsymbol{\kappa}_i)_{i=1}^n\} \leq \alpha.$$

→ CQR could fail to provide valid coverage guarantees in the regions of interest!



# CQR calibration – still biased??

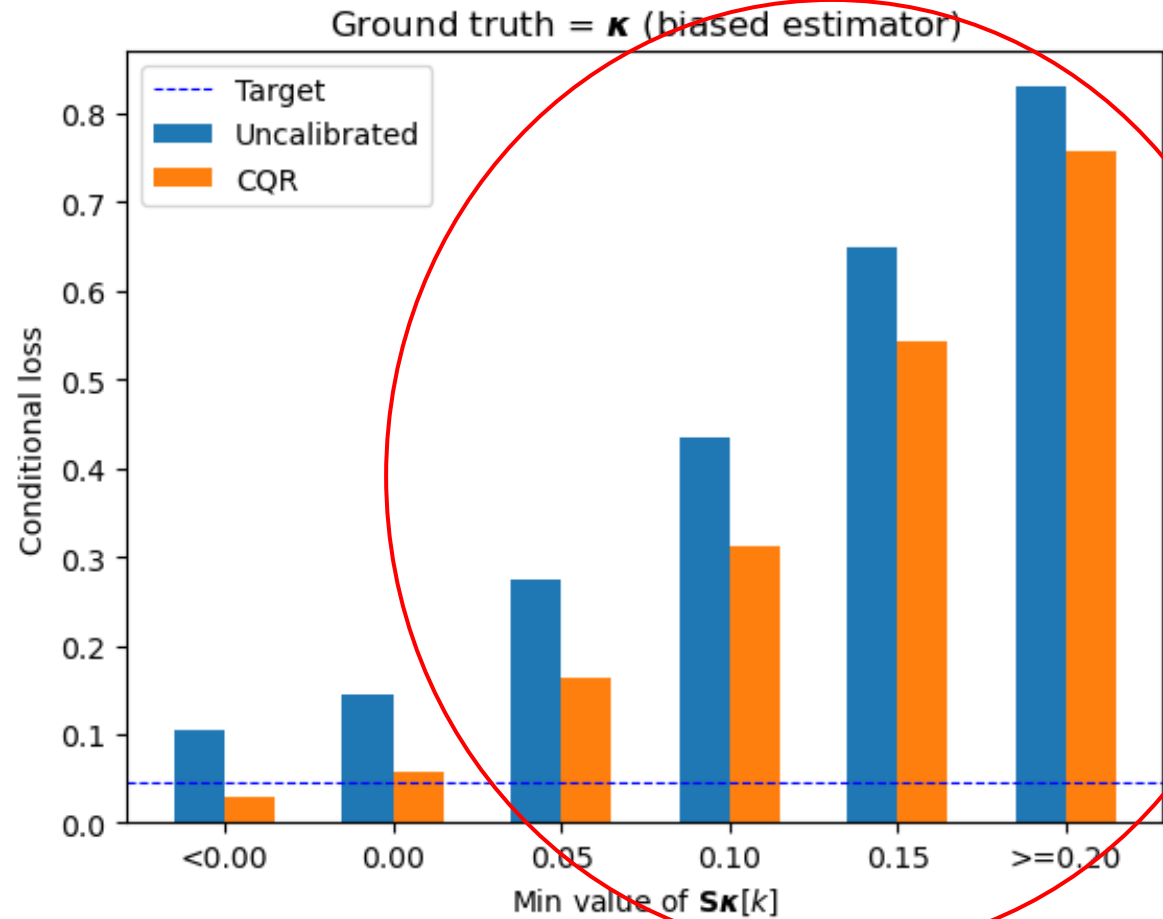
Ratio of pixels falling outside the predicted bounds:  $\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]]$ , conditioned on the value of  $S\kappa[k]$ .



# CQR calibration – still biased??

Ratio of pixels falling outside the predicted bounds:  $\kappa[k] \notin [\hat{\kappa}^-[k], \hat{\kappa}^+[k]]$ , conditioned on the value of  $S\kappa[k]$ .

Way above target for large values of convergence!



# Next steps

- Run similar experiments on more advanced reconstruction methods.
- In particular, apply plug-and-play methods to this framework?
- Adapt the results to simulated galaxy shape catalogs predicting results from SKA (T-RECS).<sup>1</sup>
- Can we get coverage guarantees conditionally to the neighboring pixels? E.g., around peak values.
- Detection of extensive objects: existing work based on prediction masks.<sup>2</sup>
- Other avenue: exploit correlation between pixels?<sup>3</sup>

<sup>1</sup> A. Bonaldi et al., “The Tiered Radio Extragalactic Continuum Simulation (T-RECS),” *Monthly Notices of the Royal Astronomical Society*, vol. 482, no. 1, pp. 2–19, Jan. 2019.

<sup>2</sup> G. Kutiel, R. Cohen, M. Elad, and D. Freedman, “What’s Behind the Mask: Estimating Uncertainty in Image-to-Image Problems.” *arXiv*, Nov. 28, 2022.

<sup>3</sup> O. Belhasin, Y. Romano, D. Freedman, E. Rivlin, and M. Elad, “Principal Uncertainty Quantification with Spatial Correlation for Image Restoration Problems.” *arXiv*, May 17, 2023.