



# Focal plane commissioning

LSST-France, 11 June 2024  
Parallel session



U.S. DEPARTMENT OF  
**ENERGY**

**SLAC**

CHARLES AND LISA SIMONYI FUND  
••• FOR ARTS AND SCIENCES •••

**LSST**  
CORPORATION

# Ongoing activities

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- Data replication

- LSSTCam / ISR

CCD configuration optimization

Data quality / investigations of  
puzzling effects

Bias correction strategy

PTC (Brighter-Fatter and gains)

Calibration validation

- CCOB

- CBP

- Photometry

# Data replication (Fabio+Thibault)

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- Goal: to replicate in-dome calibration images from USDF to FrDF
  - To use FrDF's capacity for studying those data
  - Recorded in [PREOPS-4853](#)
- For LSSTCam focal plane tests data we used some components of the replication machinery we plan to use in operations
  - It worked reasonably well, even if we found some aspects to improve
  - Ingestion into FrDF's local butler was done by an ad-hoc solution: the long term solution is being developed and tested, first at SLAC then at EU DFs
- Replicating ComCam data would use the same tools we are currently testing for the operations phase
  - Our goal is to replicate that data to FrDF, after the 30 days-long embargo period during the commissioning phase
  - From our experience replicating Run6 data, the photodiodes files are needed and they were not part of the data automatically replicated

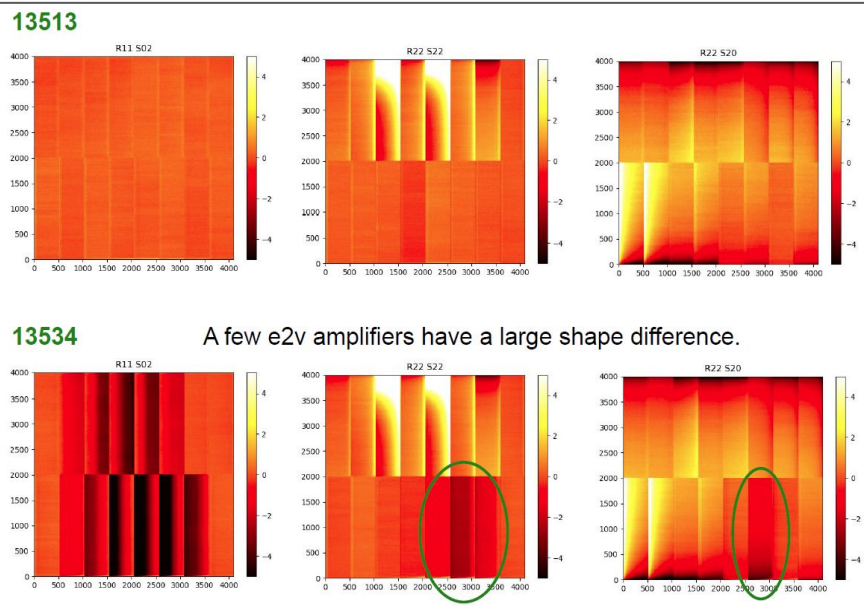
# CCD configuration optimization (Claire+Pierre An.+Pierre As.+Thibault)

Ex: test of voltage configurations in Run 6

## New voltages for e2v

	Pre Run5	Run5	Run 6a	new	
<b>pclkHigh</b>	3.6	3.3	3.0	3.3	Tearing / Persistence / P4
<b>pclkLow</b>	-6.0	-6.0	-6.0	-6.0	Constraint from HV
<b>sclkHigh</b>	3.9	3.9	3.9	3.6	
<b>sclkLow</b>	-5.4	-5.4	-5.4	-5.7	
<b>rgHigh</b>	6.1	6.1	6.1	5.8	
<b>rgLow</b>	-4.0	-4.0	-4.0	-4.3	
<b>rd</b>	11.6	11.6	11.6	11.3	
<b>od</b>	23.4	23.4	23.4	23.1	
<b>og</b>	-3.4	-3.4	-3.4	-3.7	
<b>gd</b>	26.0	26.0	26.0	26.0	
<b>Expectation</b>		Mitigation of Divisadero tearing	Mitigation of Persistence	Mitigation of incomplete reset?	

## Bias shapes

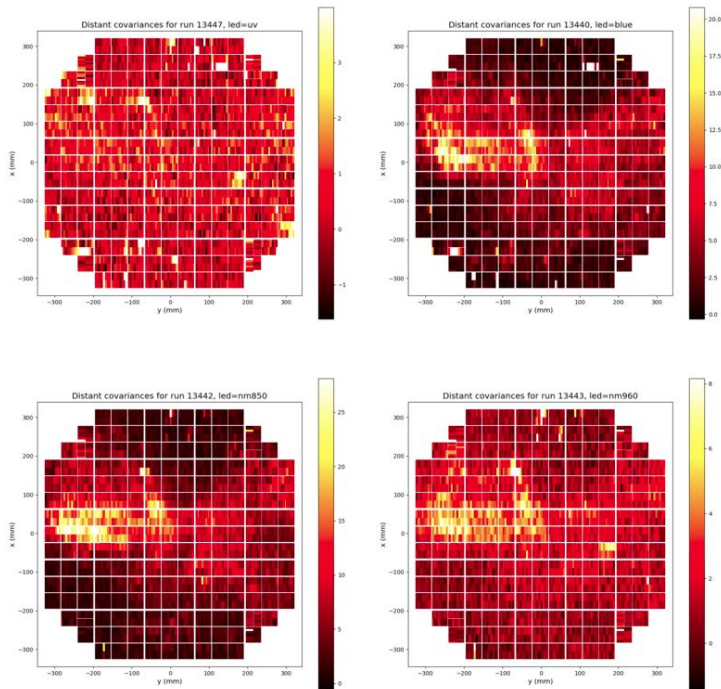


From Yousuke

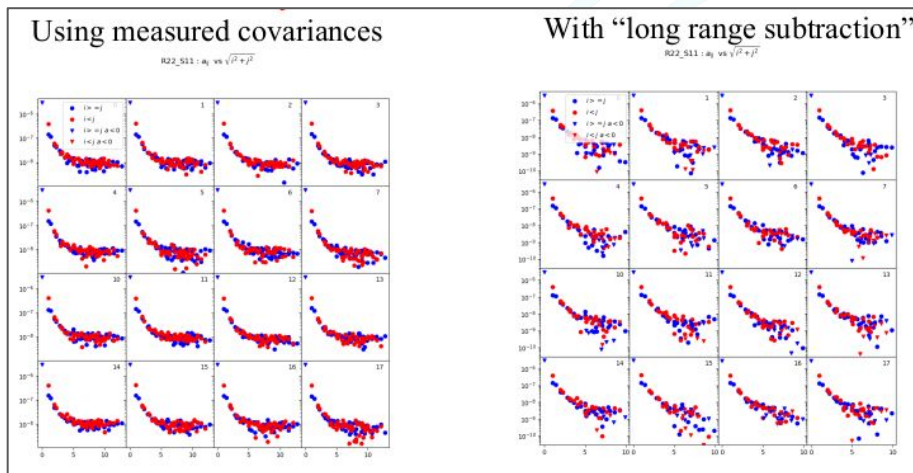
Noise, persistence, stability, etc.

# Data quality / investigations of puzzling effects (Claire+Pierre An.+Pierre As.+Thibault)

Ex: long range covariances in Run 6 flats



- Effect depends on LEDs (pattern and intensity)
- Needed to model this effect for PTC analysis

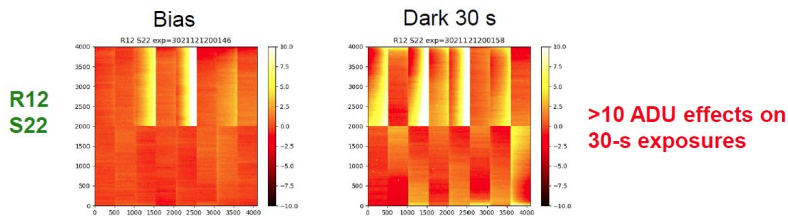


# Bias correction strategy (Pierre An.+ Thibault)

## Main bias issues

e2v

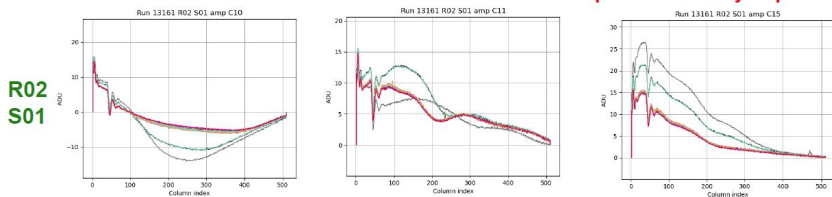
Impact of the exposure time on the 2D bias shape



ITL

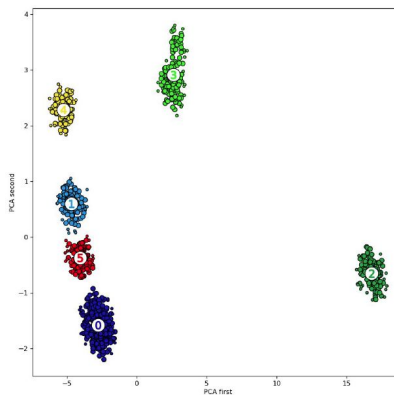
Multimodal biases (a.k.a. ITL bias jumps)

Up to ~10 ADU jumps

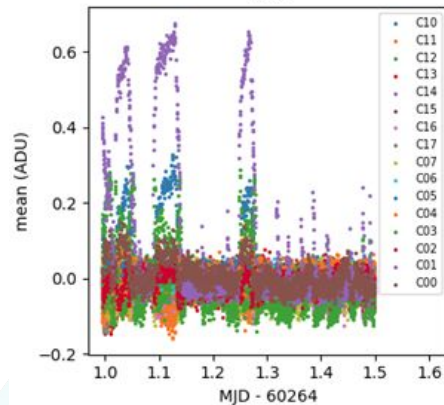


eo\_pipe studies  
of OpSim runs

run 13390 , RAFT R02 CCD S01  
pre14pre2-pre0 per event for amplifier C10\_C11\_C12\_C13\_C14\_C15\_C05\_C04\_C03\_C02  
DBSCAN clustering with eps=0.300000



R22, MD15\_1D  
S02



Identification of  
ITL families

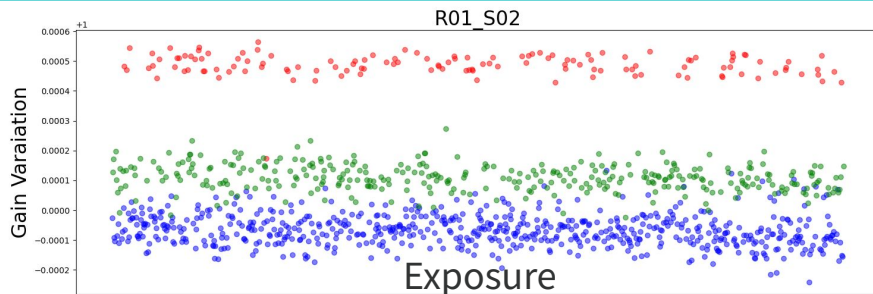
# Brighter-Fatter (Pierre As.)

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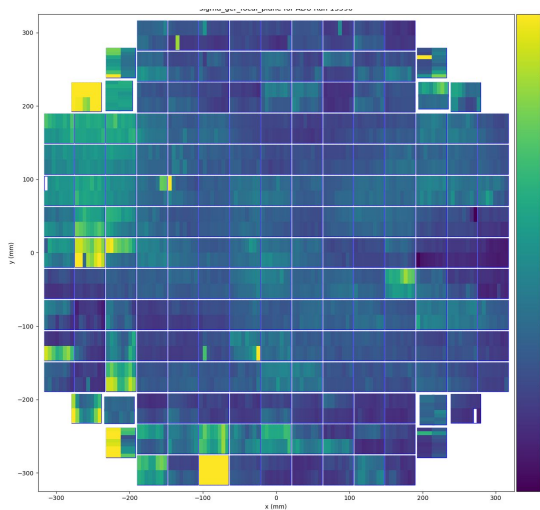
- On Auxtel, because of the fine image sampling (FWHM  $\sim 10$  pixels) the BF effect is probably far too small to provide an interesting test on the sky.
- Comcam is the next opportunity. There are flat series already available, but we could rely on the in-dome flat system. For non-linearity, the illumination monitoring system should deliver a proper reference, once properly configured.
- Regarding DM code, the situation is confusing:
  - The non-linearity fit is implemented but not available yet (nor its performance characterized, AFAIK)
  - For image BF-correction, there are competing options, and I don't know how decisions will be made.
  - I would advocate that comcam will be an excellent sandbox.

# Gain correction (Pierre An.+Yassine)

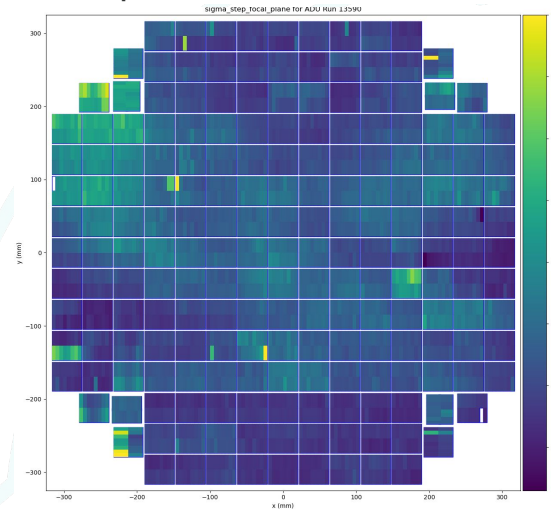
Some itl ccds present gain variations as seen in R01\_S02 for run 13590 (one color per gain value organized per “family”):



If we take those gain values per family into consideration we can reduce it’s raft dispersion and bring them to e2v level :



On the left general dispersion, on the right dispersion with family identification





# Calibration validation (Thibault)

Mainly just following this activity recently...  
Critical activity with a LOT of things to validate  
Calibration Rehearsal 27-29 May

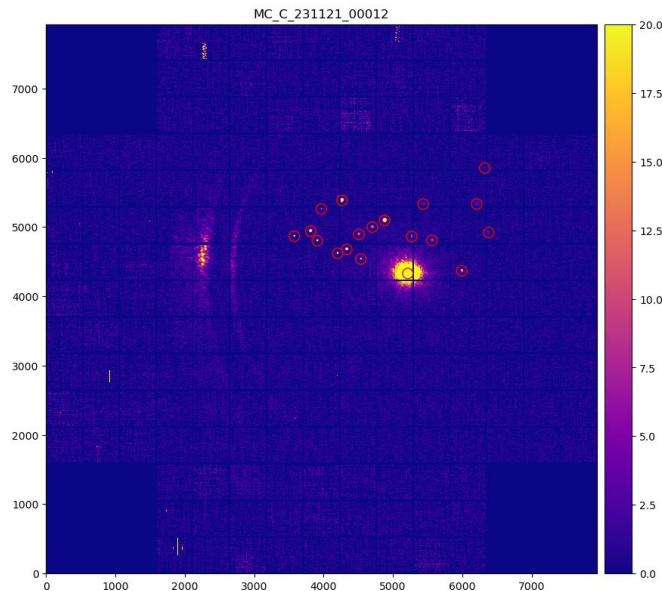
## Validation chain

cp\_pipe → calibration collections  
→ cp\_verify → statistics results  
→ analysis\_tools → metrics/plots  
→ web report for TAXICAB

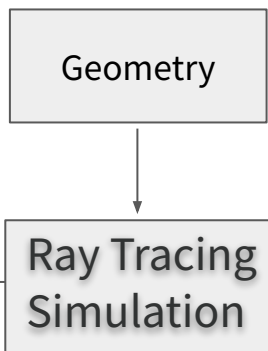
## Web report (Chris Waters)

- Use butler to find all relevant input datasets:
  - analysis\_tools plots and metrics.
  - cp\_verify focal plane residual mosaics and final result table.
  - cp\_pipe focal plane calibration mosaics.
- Copy input datasets to output `src` directory, with:
  - FITS mosaics converted to PNG plots.
  - analysis\_tools metric bundles converted to tables.
  - cp\_verify results converted to tables.
- Iterate over calibration type (bias, dark, etc), generating:
  - An index page containing focal plane results and metric bundle tables.
  - An exposure page containing per-exposure plots and metrics, organized by exposure.
  - A detector page containing per-detector plots and metrics, organized by detector.
- Clearly label everything with the butler collection, `dataset_type`, `dataId`.
- The full camera will need navigation helpers.

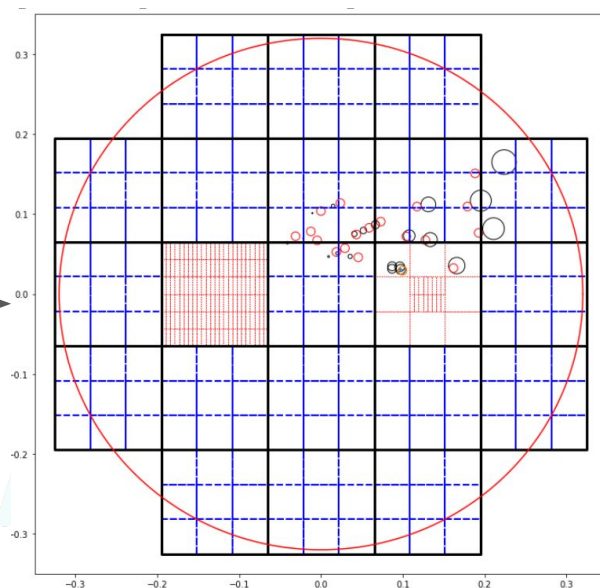
## Camera Optics alignment and transmission measurement



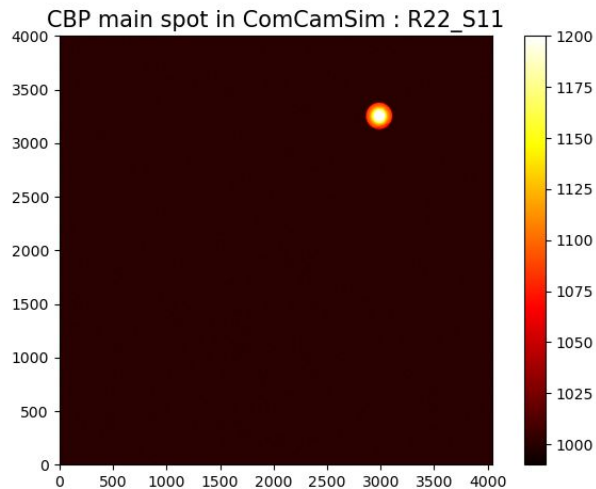
Narrow beam Run 6 data



## Ghosts Match ?



## CBP spot in ComCam Image simulation with Batoid + ImSim



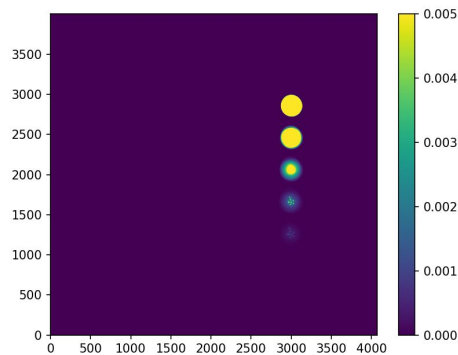
Goal : photometry with this spot

Ongoing : Normalize spot flux depending on exptime

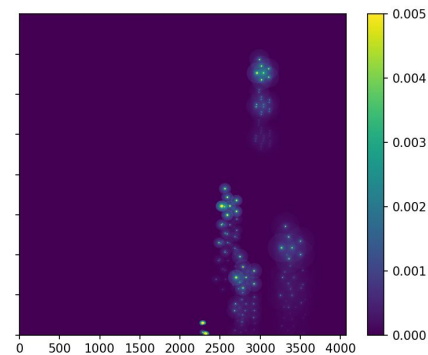
→ In collaboration with Harvard student

## Ghosts simulation with Batoid

Princ. CCD

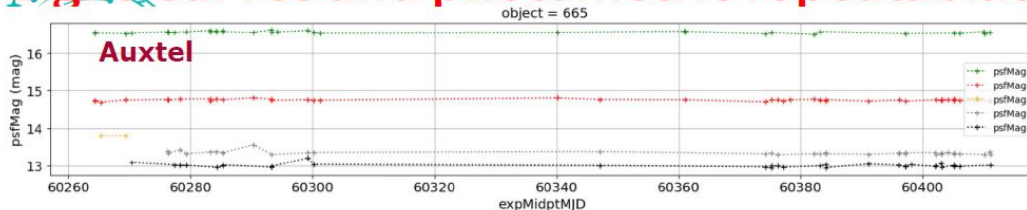


CCD below

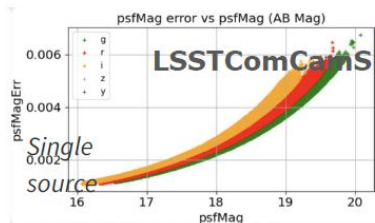
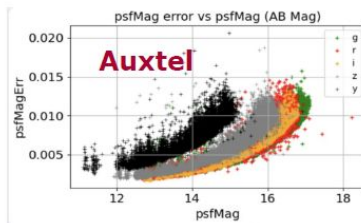


Goal : use Ghosts positions for  
photometry + Rubin-CBP alignment

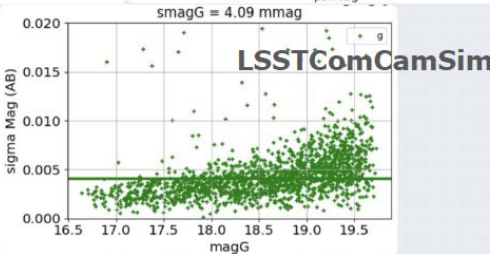
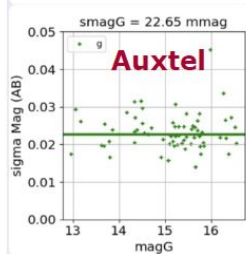
## Light-Curves and photometric repeatability



band	AUXTEL	LSST
g	23 mmag	4.8mmag
r	21mmag	4.7mmag
i		4.8mmag
z	23mmag	
y	36 mmag	



Single source

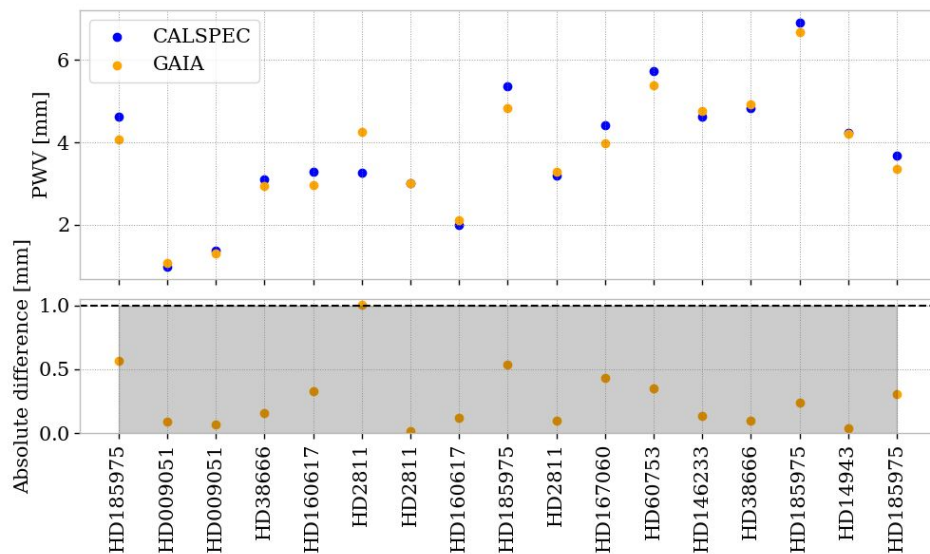


Photometric repeatability:  
all sources a given object

Auxtel does not follow requirements on photometric repeatability

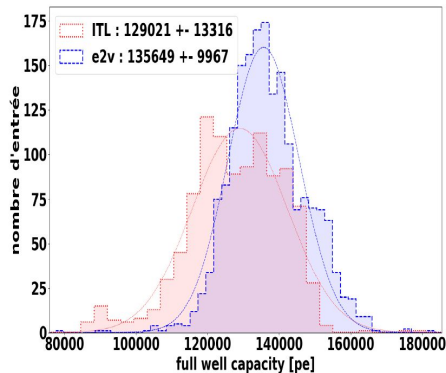
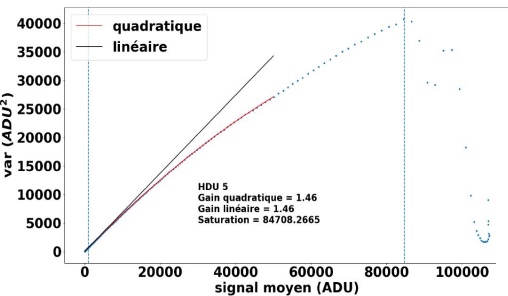
- Standard spectra used in AuxTel from CALSPEC and no coverage of the full LSST sky, nor the DDF.
- **First test:** using GAIA spectra as standards in the spectroscopic extraction of AuxTel and comparison with CALSPEC
- Conclusive feasibility study, still a lot of work on GAIA catalog and spectra
- See presentation in AuxTel parallel

Precipitable water vapor measurement by AuxTel with CALSPEC and GAIA spectra



# Ccd saturation and SNe Ia (Julie+Philippe)

- ccd saturation -> impact on low-z ( $z \lesssim 0.05$ ) SNe Ia
- Estimate gain/ccd full well using run5



$$Flux = flux[pe/seconde] \times temps\ de\ pose \times f(seeing)$$

