

Analysis of the CCOB-NB Run6 data

Camera Calibration Optical Bench – Narrow Beam

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- Commissioning of the integrated camera (LCA-283-F)
 - precise measurement of the **optical throughput**
 - determination of the optics alignment/tilt
 - mapping of the **baffle** between L1 and L2
- The Camera Calibration Optical Bench Narrow Beam (CCOB-NB) is a calibrated and $(X,Y)+(\Theta, \Phi)$ positionable pencil beam monochromatic light source.
 - allows the illumination of the focal plane through the full optical system from a variety of incident angles in the 6 spectral bands
 - 2.5 mm wide monochromatic beam ($\delta\lambda\sim$ 1 nm)
 - from 300 nm to 1100 nm









In-operation picture

Designed and built at LPSC, used at SLAC for Run 6 in 2023





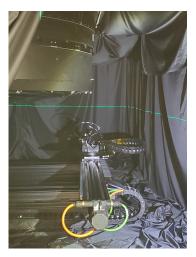






In-operation picture

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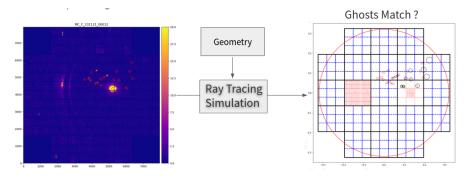




Main idea



- Run6 data taken in november 2024 in 27 different configurations
- Ray tracing simulation package based on batoid

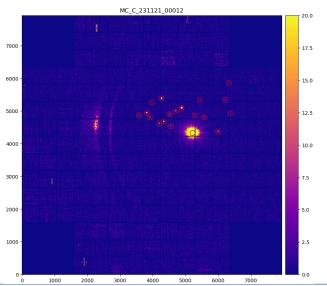






Full focal plane image with ghosts









Rubin

Observatory





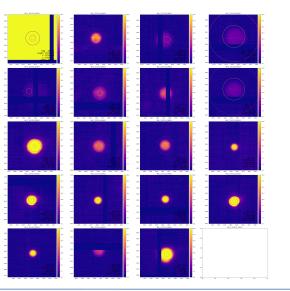
- Run the ISR on each CCD, relying on eo_pipe
- Make a mosaic image : assembling CCDs into one focal plane image
- Point and click on each ghosts
- Fit each ghost position with a 2D gaussian to get a precise position (and a rough flux estimation)
- Coordinate transform to get the CCD in which each ghost is
- Fit each ghost position again on the post–ISR CCD images: to get a correct coordinate transform to absolute focal plane position (a.k.a DVCS)







Ghosts fits





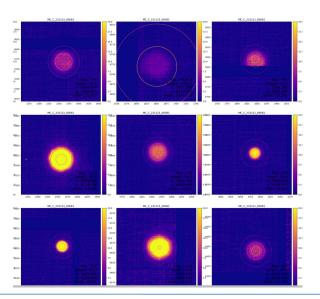


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Ghosts fits



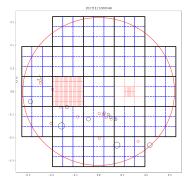


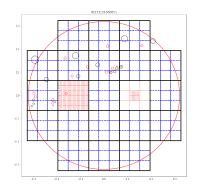






Match data and simulations





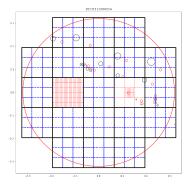


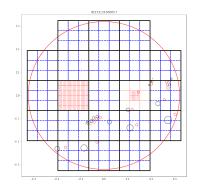


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Match data and simulations







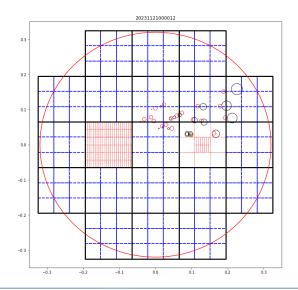


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Match data and simulations













ghost spot $g_{r,i}$ has parameters $[x_{r,i}, dx_{r,i}, y_{r,i}, dy_{r,i}, rad_{r,i}, drad_{r,i}, p_{r,i}]$ for position in x and y with uncertainties dx and dy, radius rad and uncertainty drad and intensity.

the 2D Euclidean distance between 2 ghosts spots is defined as:

$$d(g_{r,j},g_{s,i}) = \sqrt{(x_{s,i}-x_{r,j})^2 + (y_{s,j}-y_{r,j})^2}$$

and the associated error is:

$$\sigma_d(g_{r,j},g_{s,i}) = \sqrt{(dx_{s,i}^2 + dy_{s,i}^2) + (dx_{r,j}^2 + dy_{r,j}^2)}$$

with the distances spot to spot between the 2 lists $(g_{r,j1..m}, g_{s,i1..n})$, one can them associate each spot in the S_s set with the closest spot in the S_r set.

we note g_{r,k_i} the closest ghost spot in S_r to the ghost spot $g_{s,i}$ in S_s .

the reduced distance between 2 sets of ghosts spots may then be defined as follow:

$$L = rac{\sqrt{\sum_{i=1}^{n} rac{d(g_{s,i},g_{r,k_i})^2}{\sigma_d(g_{s,i},g_{r,k_i})^2}}}{n}$$

if 2 sets of beam spots are the same L=0, and if they are really close then L should be small.

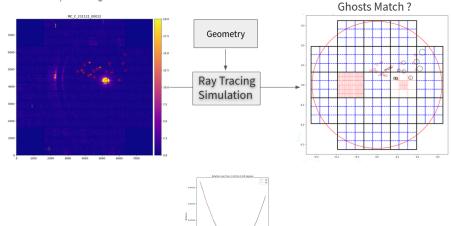






Fit principle





nin ols sis rotation angle degrees









- Some complexities
 - 5 parameters × (3 lenses + focal plane)
 - the beam position is not well known: need to fit it as well (not implemented yet)
 - ghosts sometimes fall in between CCDs (error on position)
 - some ghosts are really dim
 - running one simulation for one geometry configuration is quite fast O(1) s... but that's atually slow when you need to run O(10⁵) for 27 configurations
- Fit strategy
 - fit iteratively one parameter for each lens
 - fit iteratively one lens at a time
 - fit all lenses together
 - fit using multiple configurations at a time
- \Rightarrow No results yet...









- CCOB–NB designed and built at LPSC, sent and received at SLAC in Summer 2021 (after a long journey)
- Used at SLAC on the LSST Camera for Run6 in 2023: lots of high quality data
- Data analysis for optical alignment underway!
 - [github:lsst-camera-dh/ghosts]
 - [github:lsst-camera-dh/ghosts_nb_analysis/]
- New data taking will occur on the summit this summer both for transmission and alignment purposes



