

A night sky filled with stars and the Milky Way galaxy. In the foreground, a large, multi-tiered astronomical observatory building is silhouetted against the dark sky. The building has some lights on, including a prominent green light. The observatory is situated on a dark, rocky mountain peak. In the distance, a city or town is visible with its lights glowing on the horizon.

# New developments in CCD detectors for astronomical observatories

Semi-conductors as photodetectors– 3 juin 2024

Claire Juramy-Gilles



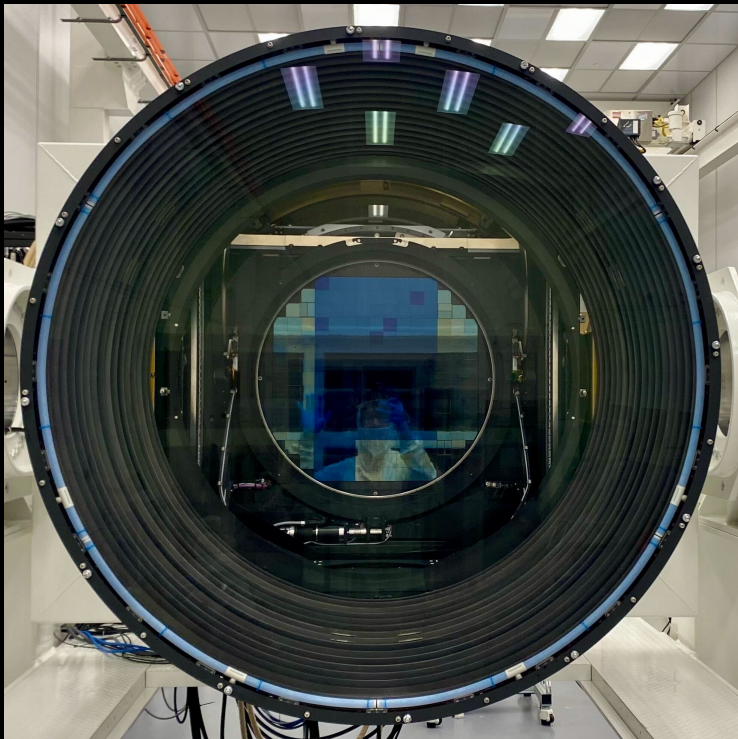
# A CCD for every occasion?

- Traditional CCDs
- CCDs vs CMOS
- Photon counting: EMCCDs
- Skipper CCDs
- Multi-Amplifier Sensing CCDs
- SiSeRO design for readout
- ... and more innovations coming up

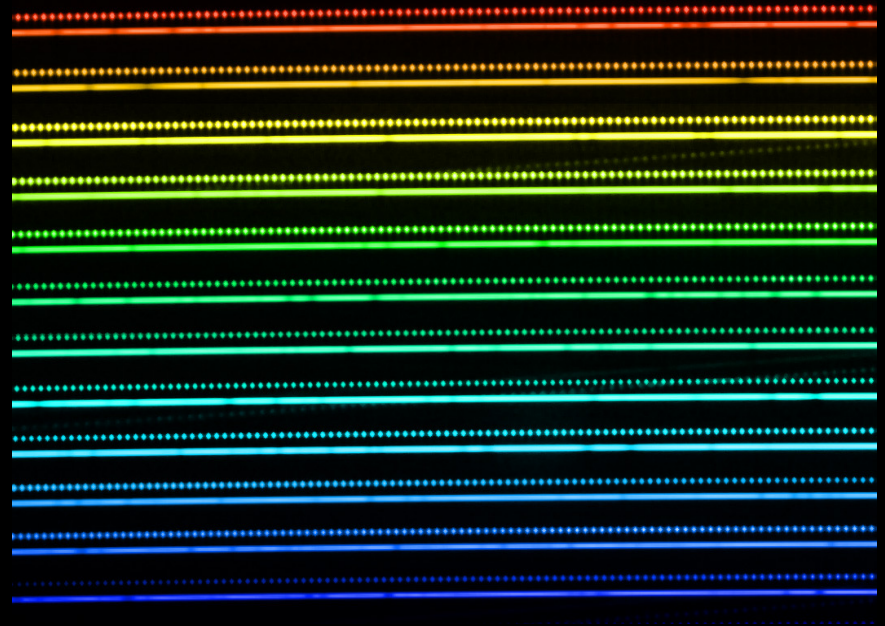


# Traditional CCDs

- Multiple applications in astronomy and astrophysics: UVOIR, X-rays
- Varied requirements: low/high signals, background, radiation hardness



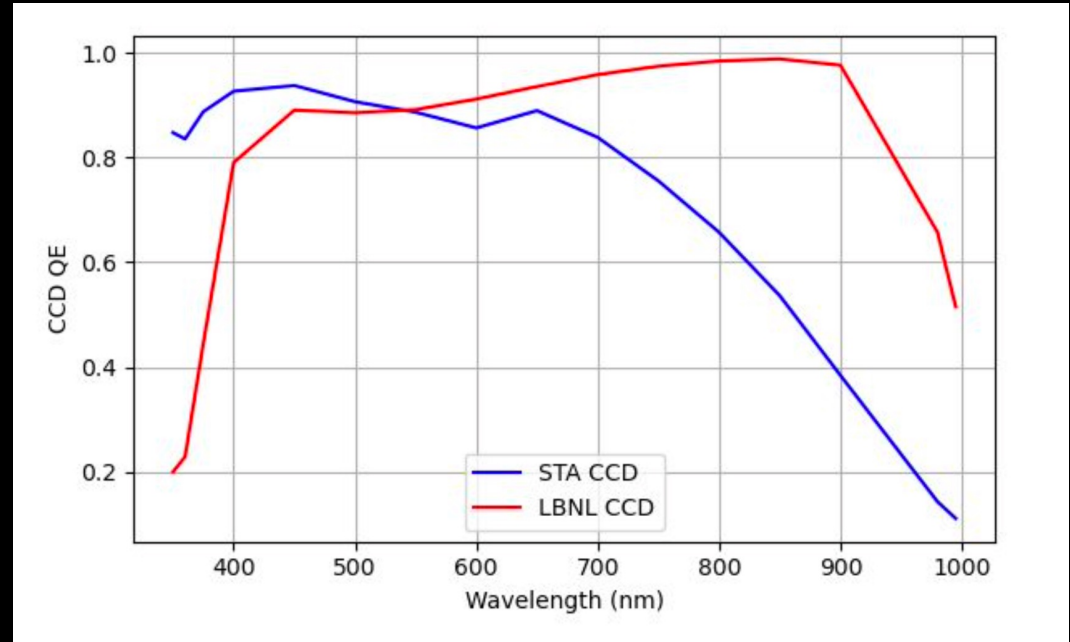
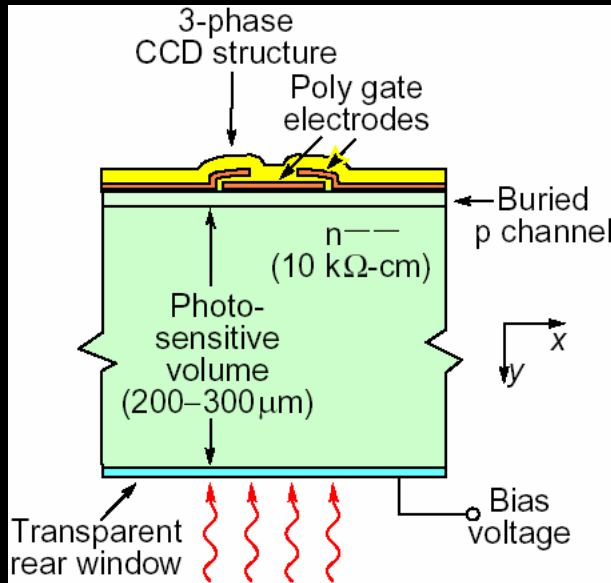
LSST



ESO, MPQ: frequency comb

# CCD performances and processes

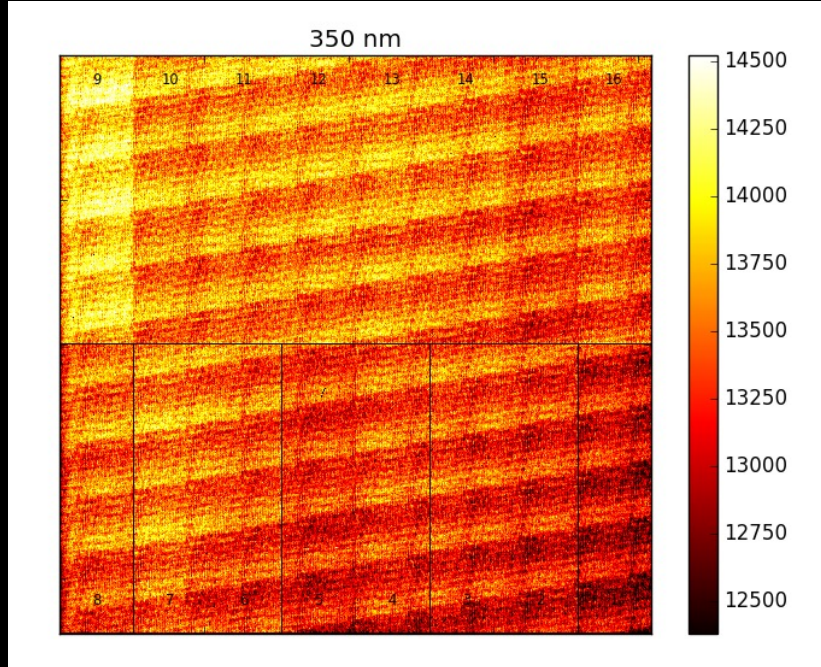
- Quantum efficiency in NIR: thickness
- n-channel (Teledyne-e2v, Semiconductor Technologies Associates Inc) and p-channel (LBNL, FermiLab, Hamamatsu)
- Radiation hardness



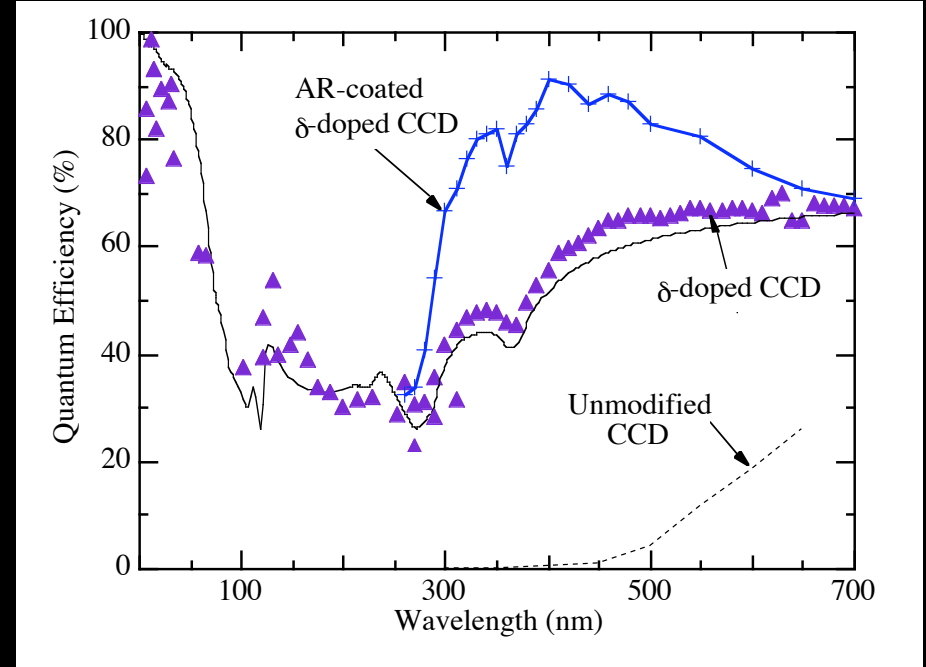


# CCD performances and processes

- Back-illuminated, thinned
- Quantum Efficiency in UV and blue: coating
- Development of 2D delta-doped CCDs with Molecular Beam Epitaxy



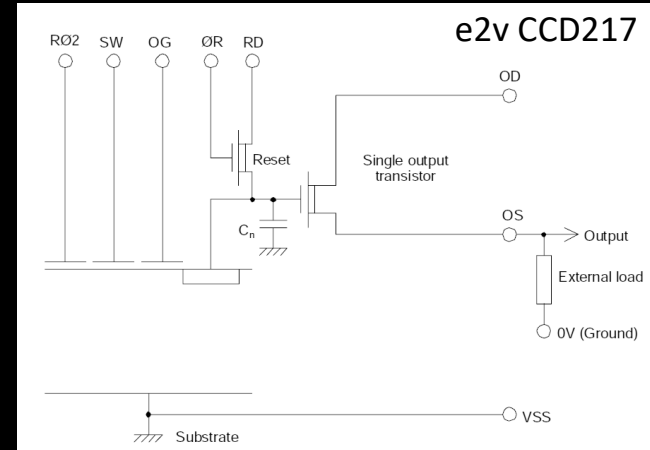
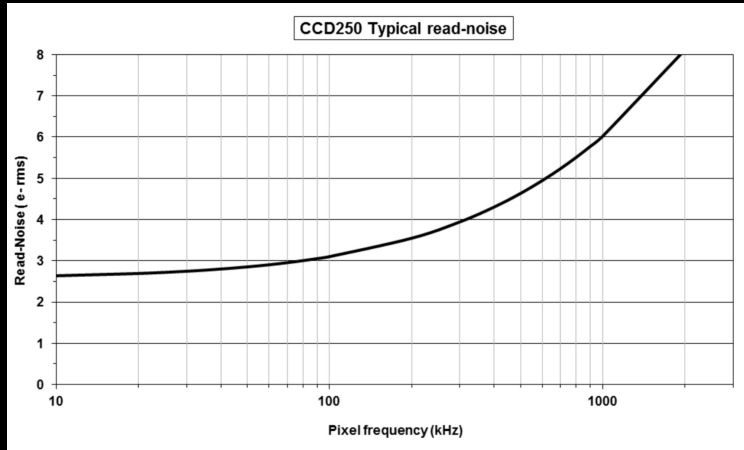
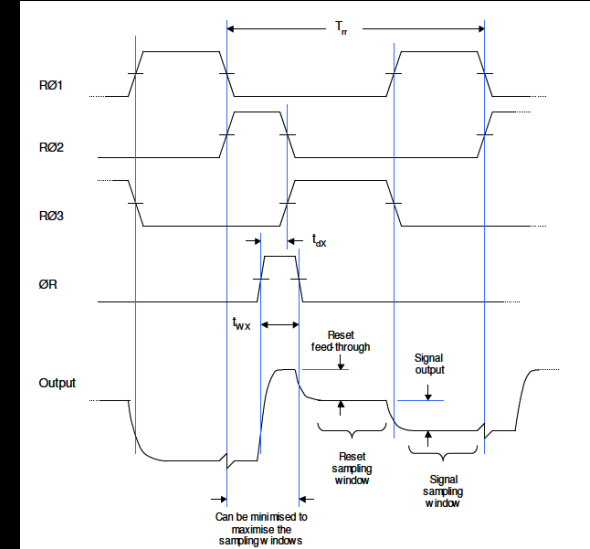
T-e2v CCD250, LSST



JPL, Caltech

# Pros and cons of traditional CCDs

- ⊕ Linearity, full well  $\sim 100s$  ke (depending on pixel size)
- ⊕ Low dark current:  $\sim e/pix/hour$  (depending on temperature)
- ⊖ Serialized readout: efficiency of the charge transfer (CTE), readout time
- ⊖ Noise corner of the MOSFET output transistor
- ⊖ Pixel reset:  $kTC$  noise subtraction

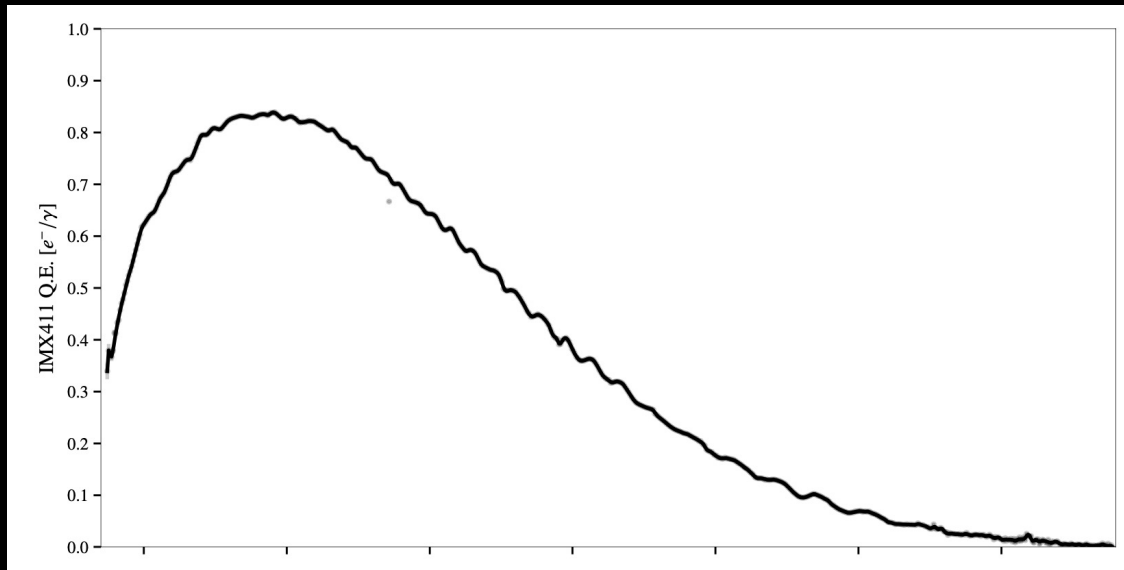




# The competition from CMOS sensors

... for astronomy/astrophysics UVOIR imaging

- “CMOS sensors” = 1 amplifier/pixel, multiplexed readout
- Pixel size, Quantum Efficiency: development of dedicated back-illuminated, high-resistivity CMOS sensors



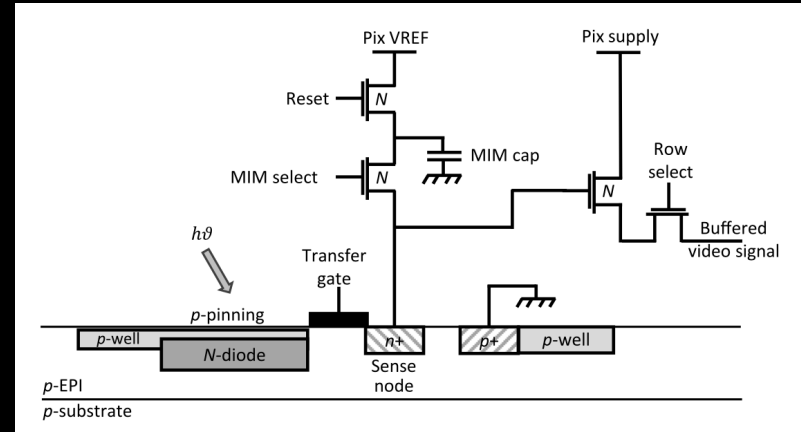
M. Betoule, Sony IMX411ALR in QHY411M (151 Mpix)



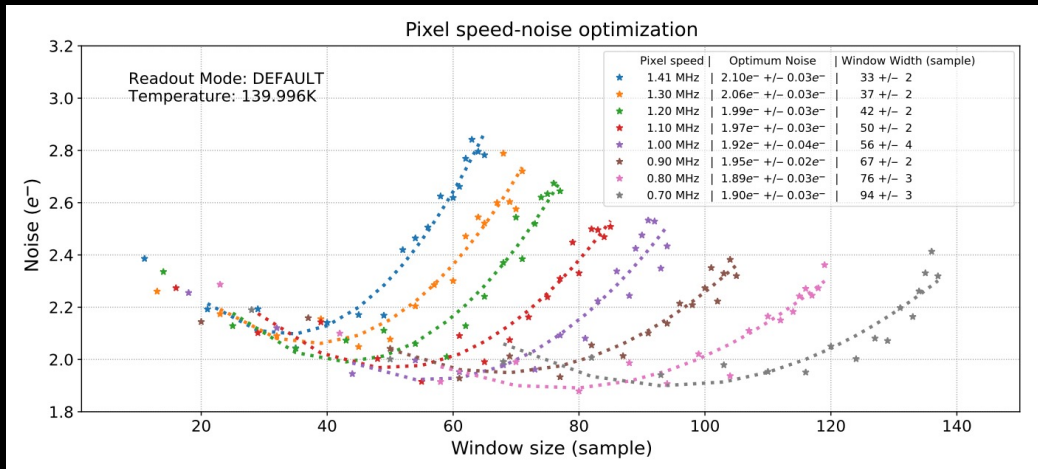
T-e2v COSMOS-66 (64Mpix) 7

# The competition from CMOS sensors

- Complexities (and opportunities) in the readout :
  - Electronic shutter (rolling/global)
  - High Dynamic Range (readout with high/low gain)
- Energy consumption, self-heating, dark current, inter-pixel capacitance



J. Janesick, 5T pixel with pinned photodiode

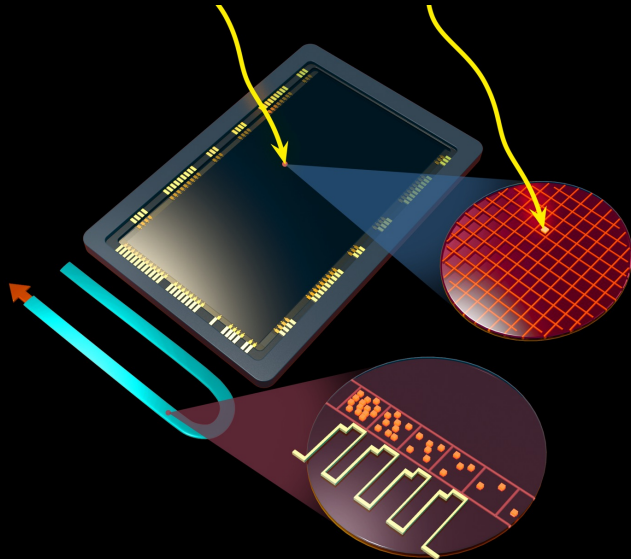


T. Greffe, SRI International 4kx2k prototype for UVEX, 2D delta-doped, Archon + FPGA 100 MS/s readout

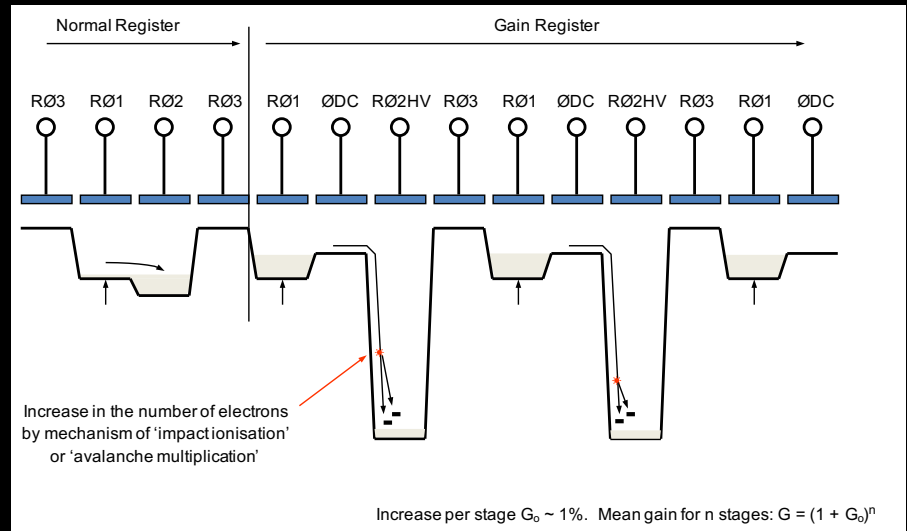


# Photon counting: Electron-Multiplying CCDs

- Also in use: Silicon APDs (avalanche photodiodes) arrays
- EMCCD: suppress read noise, also amplify dark current and clock-induced charges, additional noise from stochastic process
- T-e2v CCD311: reduce cosmic rays overspill through use of dump gate to remove high signals

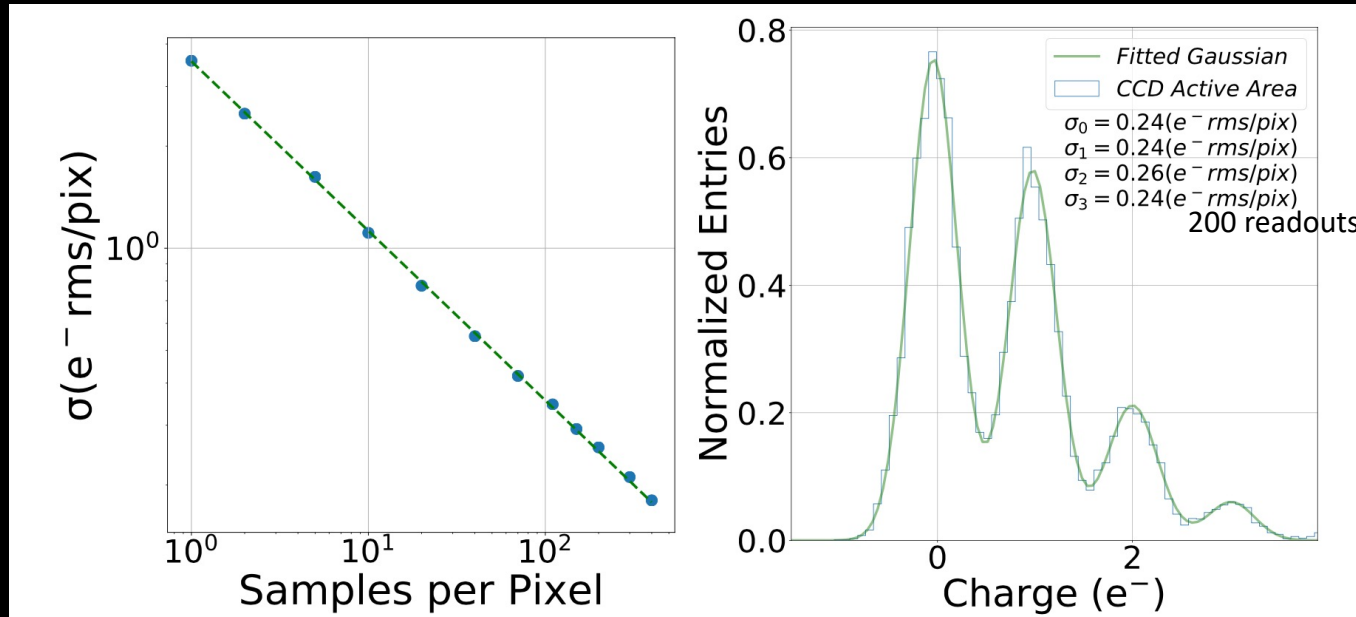
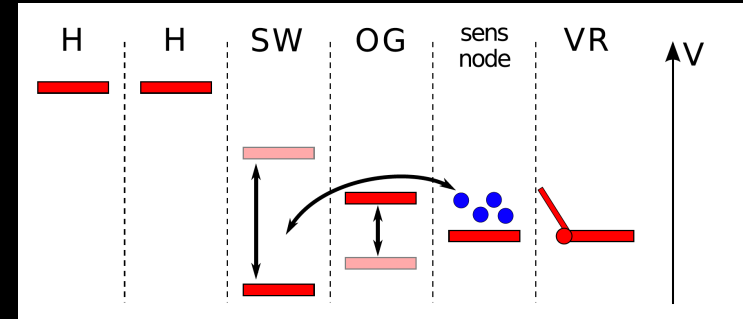


T-e2v



# Below 1 electron noise : skipper CCDs

- ‘Skipper’: pixel charge moved back and forth to readout node
- N readouts of the same charge: statistic reduction of noise in  $\sqrt{N}$
- Bonus: gain measurement



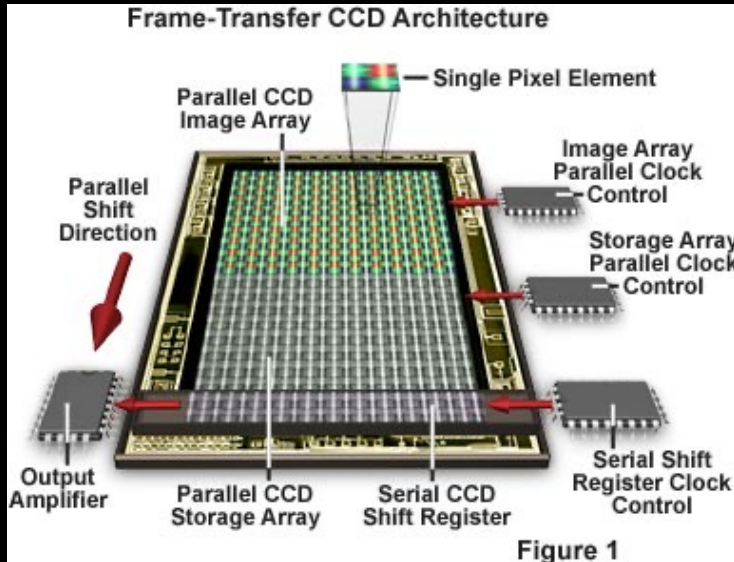
J. Tiffenberg  
2017

A. Drlica-Wagner,  
SPIE2020



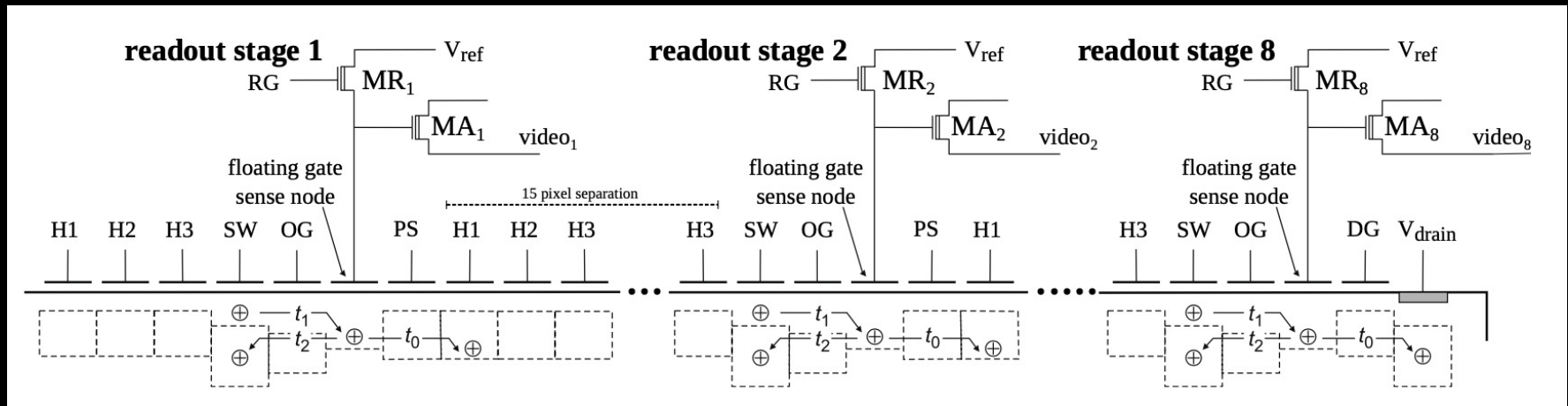
# Below 1 electron noise : skipper CCDs

- Readout time
- Frame-transfer: readout during exposure
- Use of skipper readout only in regions of interest



# Multi-Amplifier Sensing (MAS) CCDs

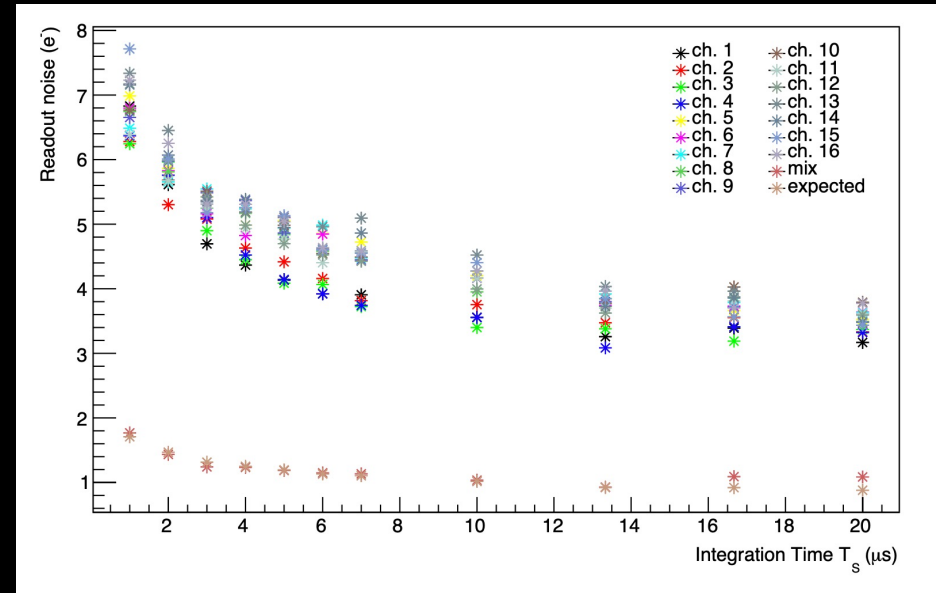
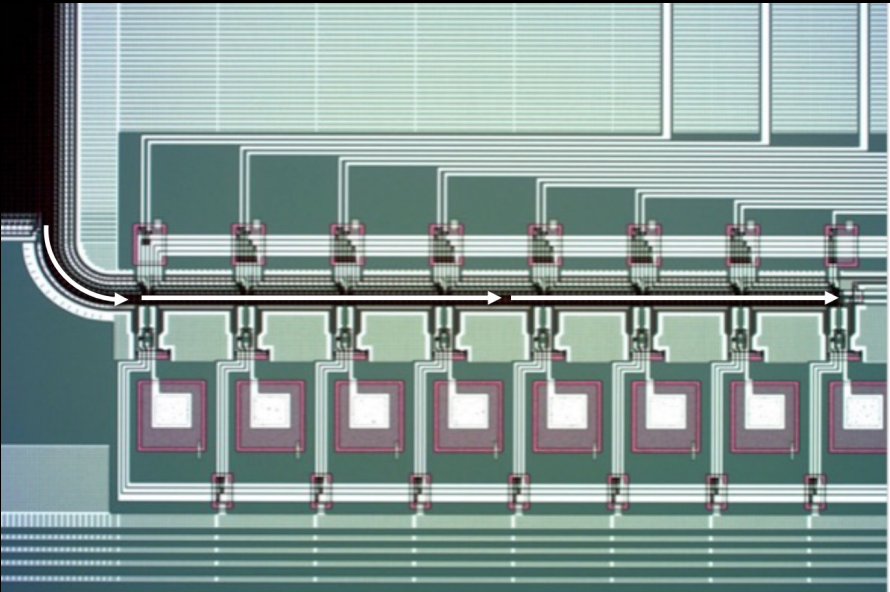
- Repeated readout of the same charge... in sequential amplifiers
- Average (or weighted average) on 8, 16, 32 channels
- Readout system
- Correlated noise suppression: read the same value at different times



G. Fernandez Moroni,  
FermiLab

# MAS CCDs Implementations

- FermiLab / Lawrence Berkeley National Laboratory
- Skipper CCDs for dark matter, tests for astronomy, MAS prototypes



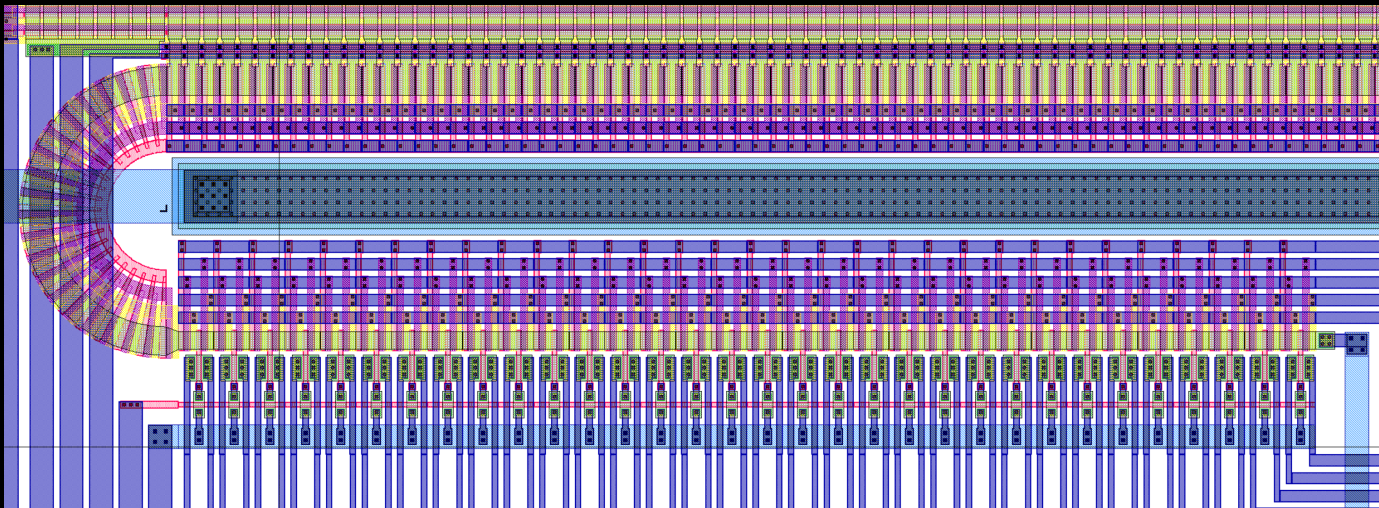
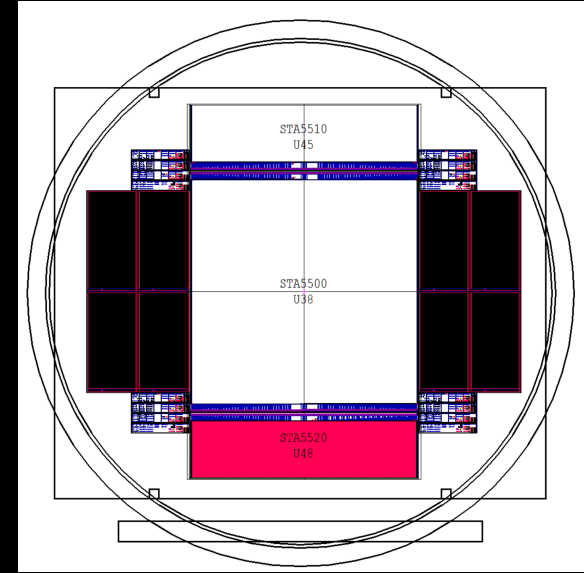
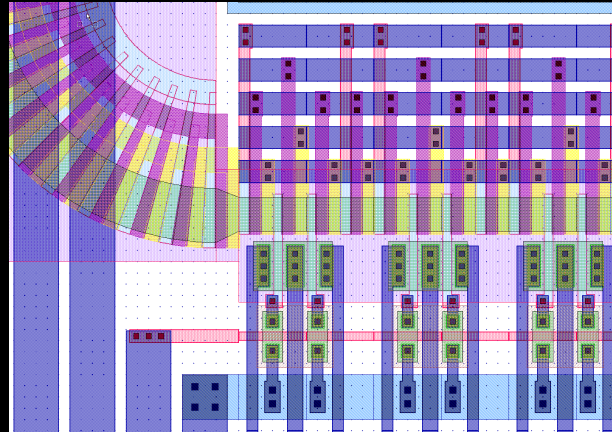
K.Lin

G. Fernandez Moroni



# MAS CCDs Implementations

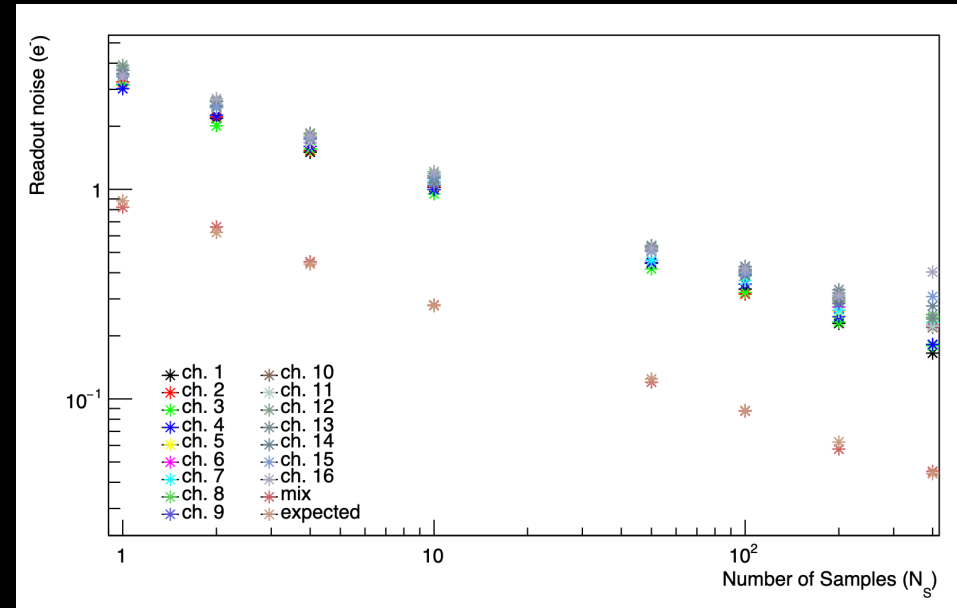
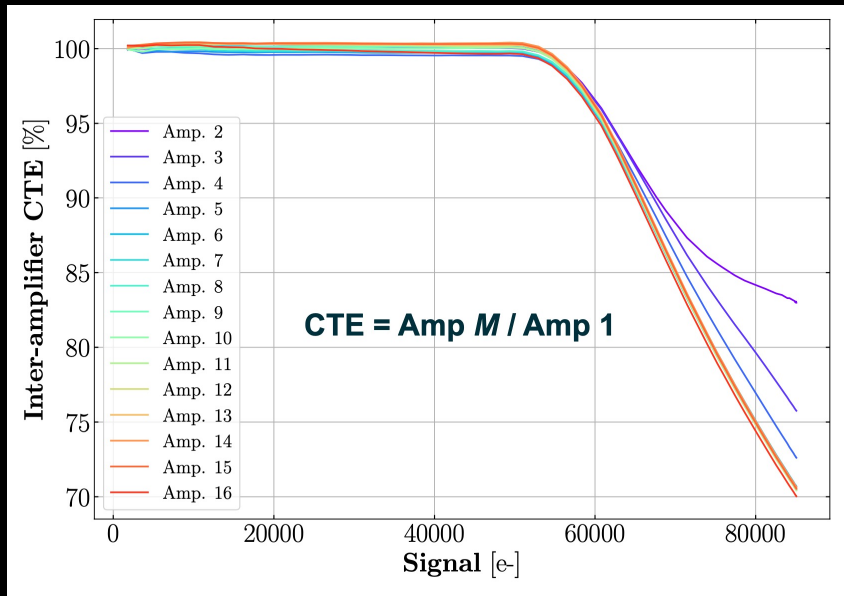
- STA5500
- Differential outputs
- Turn-around





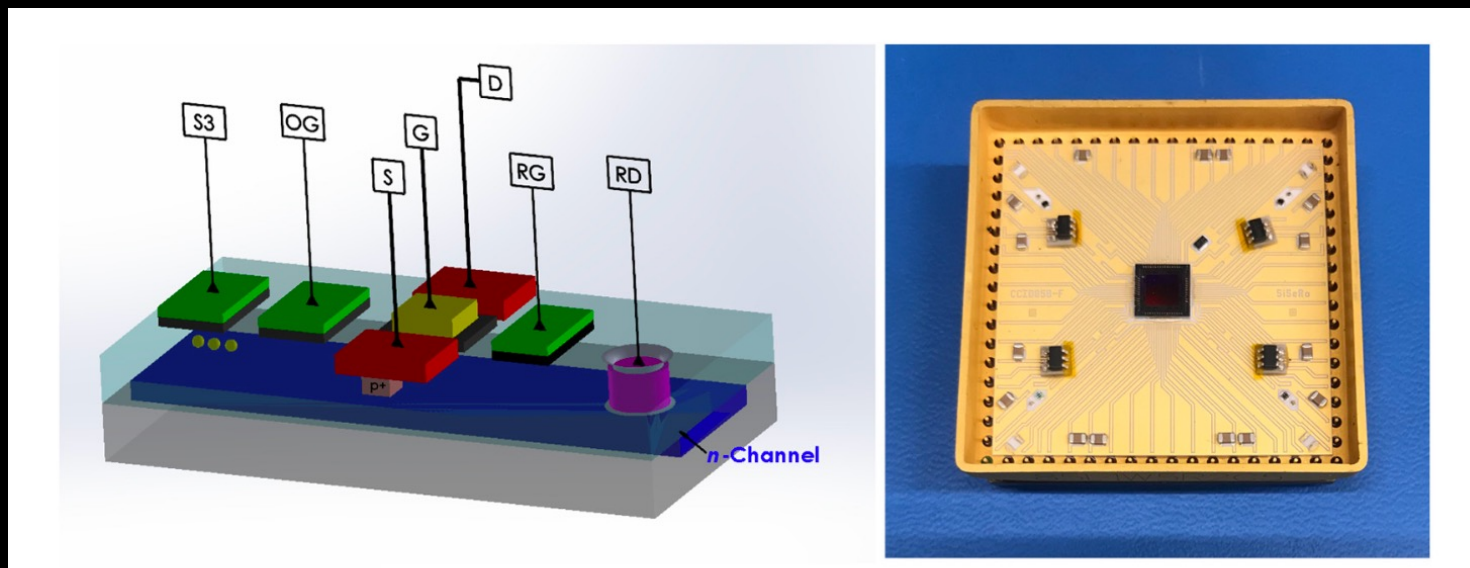
# Multi-Amplifier Sensing (MAS) CCDs

- Challenges: clock-induced charges, transfer efficiency at high flux
- Optimization of clock voltages: noise vs full well
- Can be combined with skipper mode



# SiSeRO CCDs

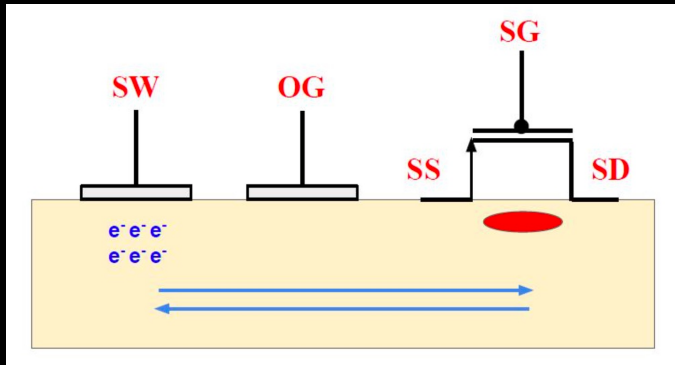
- = Single electron Sensitive ReadOut
- Pixel charge modulates current in readout transistor
- Faster readout (X-ray), high conversion gain, no kTC, compact
- Expect  $1e$  noise at 1 MHz for 1500 pA/e



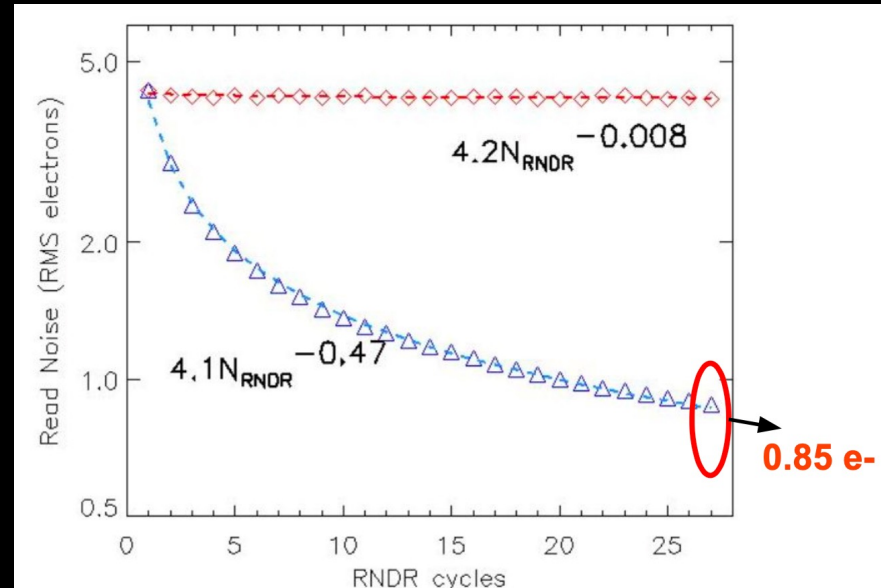
MIT Lincoln Lab CCID85F prototype

# SiSeRO CCD Implementations

- Prototyping: moved to buried channel SiSeRO, 1/f noise filtering
- Tested Repetitive Non-Destructive Readout (RNDR)
- Also demonstrated by FNAL/LBL with NMOS FET on p-CCD
- Combined with multi-amplifier architecture



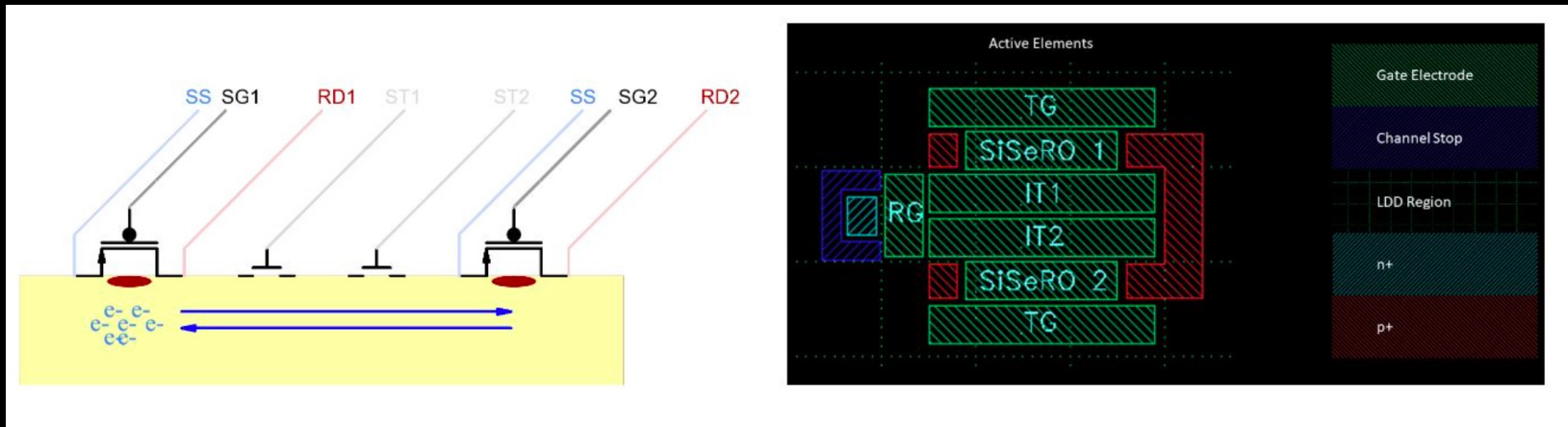
T. Chattopadhyay



# More innovations on the way

Other sensors benefitting from CCD R&D:

- Skipper CMOS
- Active Pixel Sensor with two SiSeROs per pixel, with alternating measurements of baseline and pixel values on two channels on the same output



T. Chattopadhyay



# Some takeaways

- Innovations on all fronts for CCDs: sub-e noise, readout speed, output amplifier
- Convergent evolution of CCDs and CMOS sensors: outputs, noise reduction strategies
- Dedicated sensors for astronomical observations: back-illuminated, high-resistivity, pixel size, new designs
- Development costs