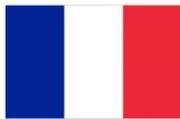
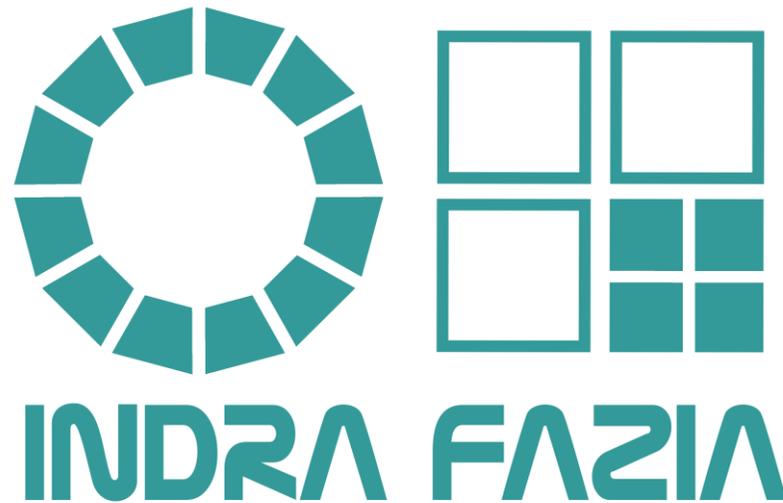


Plans for symmetry energy research within the INDRA-FAZIA collaborations



<http://fazia.in2p3.fr>

*Nicolas Le Neindre* for the INDRA & FAZIA Collaborations

Université de Caen Normandie, ENSICAEN, CNRS/IN2P3, LPC Caen UMR6534, F-14000 Caen, France

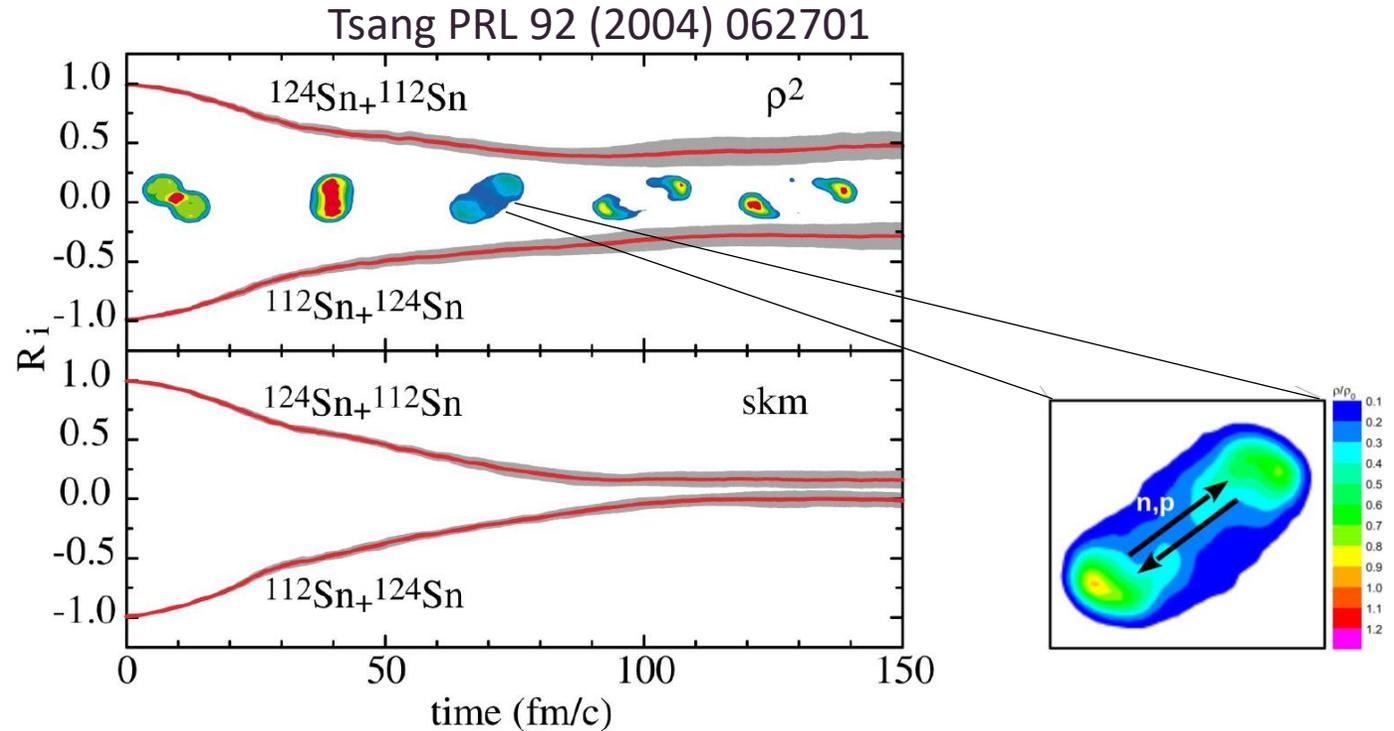
**NUSYM 2024, XII<sup>th</sup> International Symposium on Nuclear Symmetry Energy, September 9<sup>th</sup>-13<sup>th</sup> 2024, Caen**

## INTRODUCTION

In heavy ions collisions, during peripheral reactions, projectile and target interact and exchange nucleons. Isospin equilibration: projectile and target with different neutron to proton ratio equilibrate their  $N/Z$  over time. Different interactions, leading to different EoS, produce different equilibration path.

Any experimental measurements of the isospin equilibration rate would constraint the EoS  
-> See C. Ciampi, A. Jedye, S. Mallik, R. Bougault, A. Le Fèvre and many other talks

Critical role of clustering at low density



*From Diego Gruyer IWM 2024*

## What is FAZIA?



### **FAZIA at GANIL:**

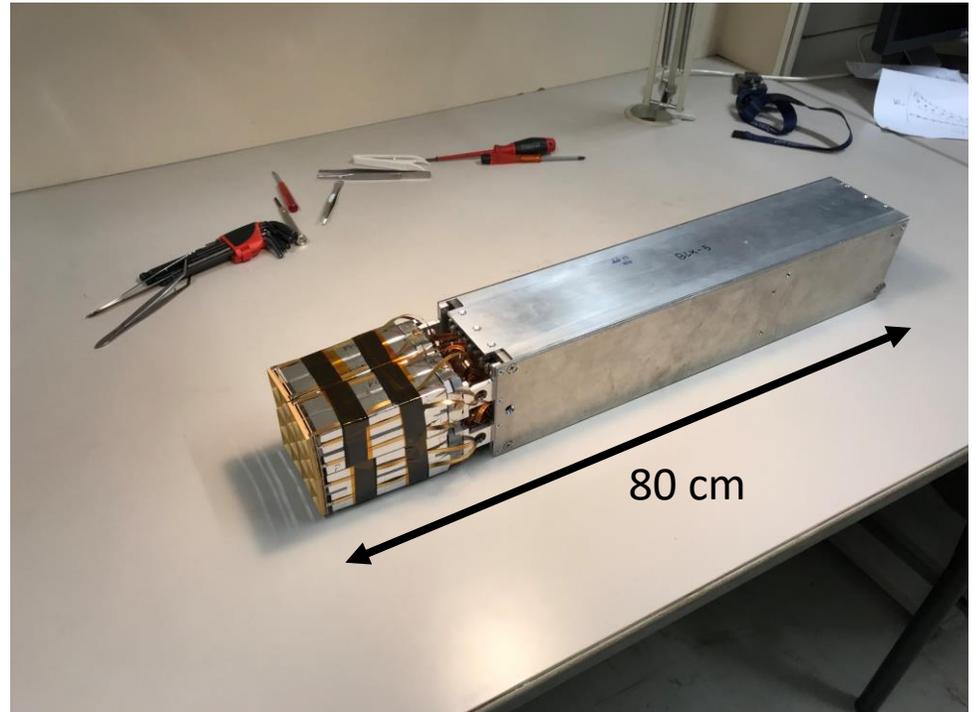
12 blocks  $1,8^{\circ}$ - $13,5^{\circ}$

192 telescopes Si-Si-CsI(Tl)

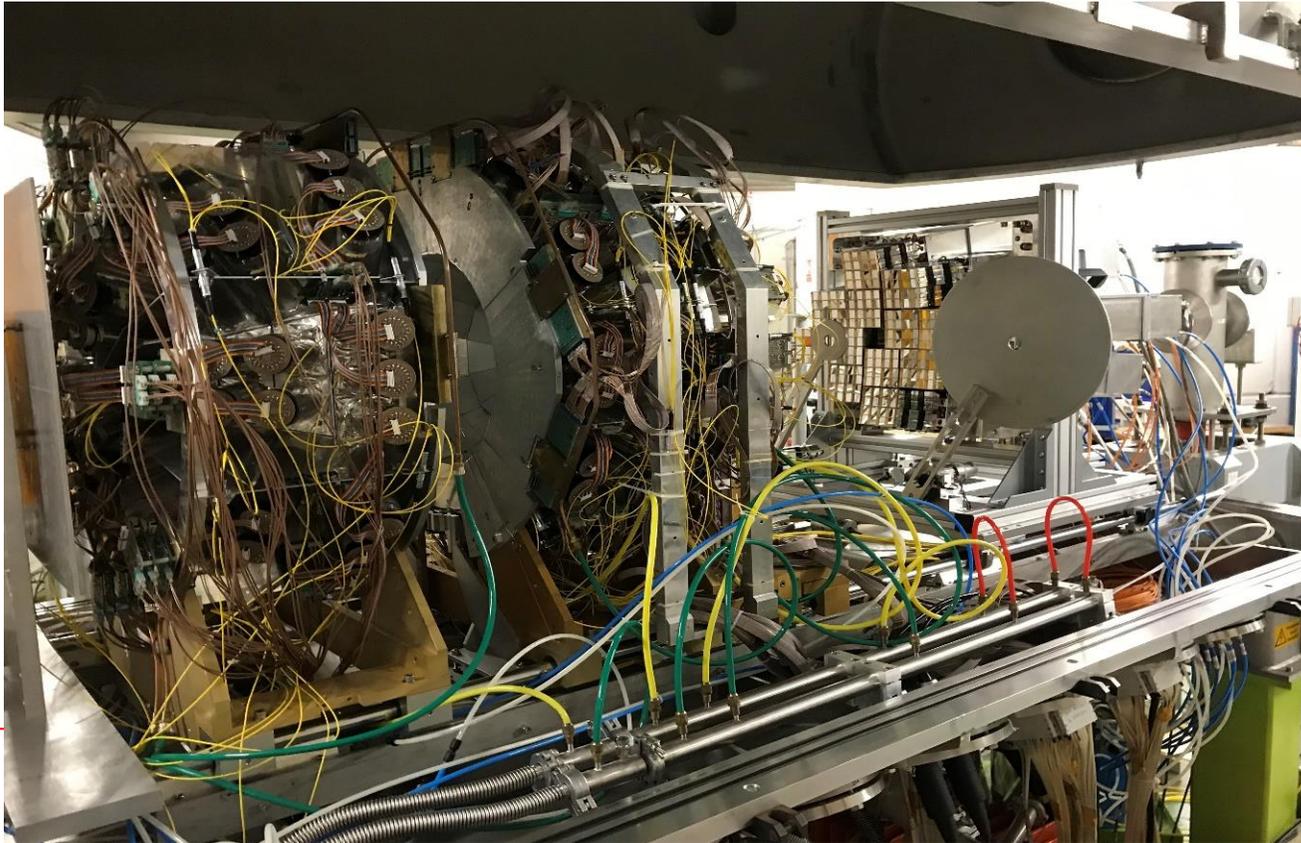
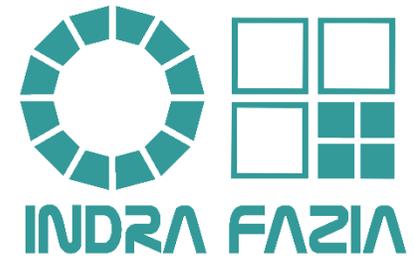
Z identification 1-54

A resolution  $Z \sim 20$  PSA

&  $Z \sim 25$   $\Delta E$ -E



# INDRA & FAZIA at GANIL in 2024



## INDRA:

12 rings  $14^{\circ}$ - $176^{\circ}$

240 CsI(Tl)

96 Si detectors

Z identification 1-54

A resolution  $Z=1-4$  CsI,  $Z=1-6$  Si-CsI

## FAZIA:

12 blocks  $1,8^{\circ}$ - $13,5^{\circ}$

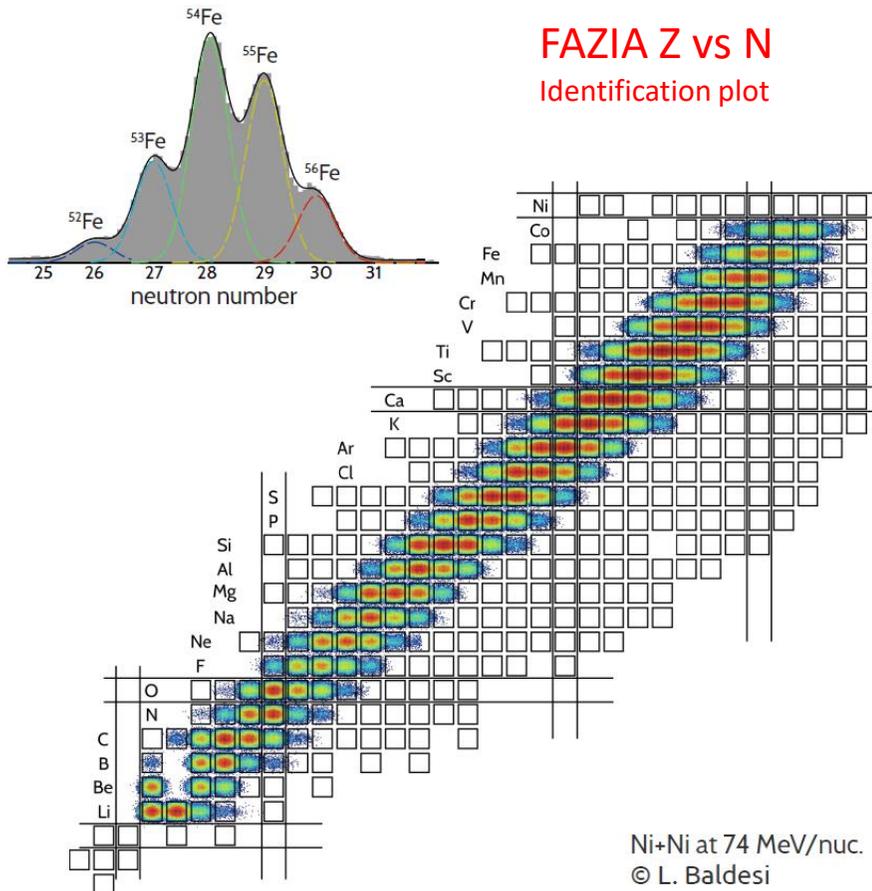
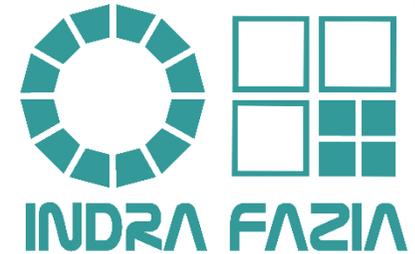
192 telescopes Si-Si-CsI(Tl)

Z identification 1-54

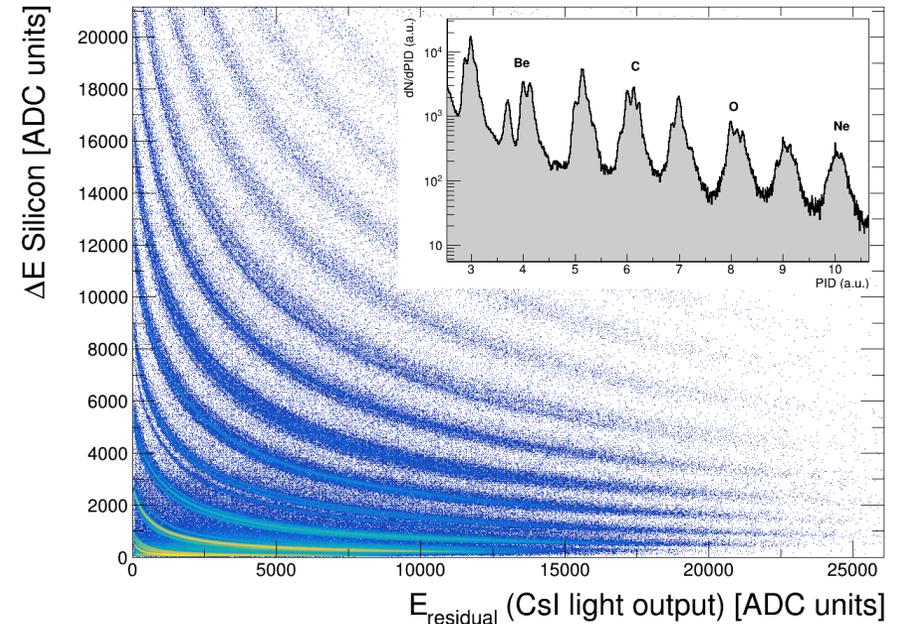
A resolution  $Z\sim 20$  PSA  $Z\sim 25$   $\Delta E-E$

# Why INDRA and FAZIA for symmetry energy research?

- Very good isotopic identification on a large scale
- Large angular isotopic resolution with INDRA upgrade  $\rightarrow 45^\circ$
- Good acceptance and low thresholds  $\rightarrow$  good event characterization
- 192 telescopes for FAZIA and 240 for INDRA  $\rightarrow$  high multiplicity
- Full digital electronics, low dead time and high acquisition rate



## INDRA upgrade



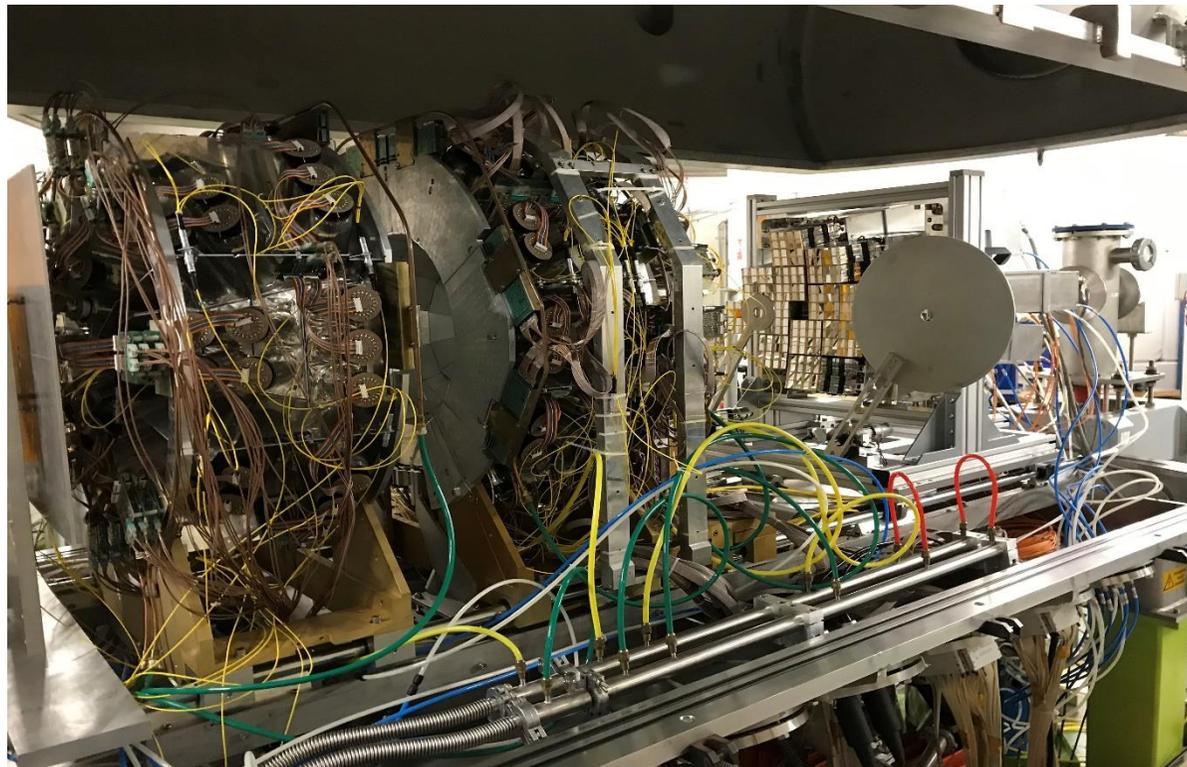
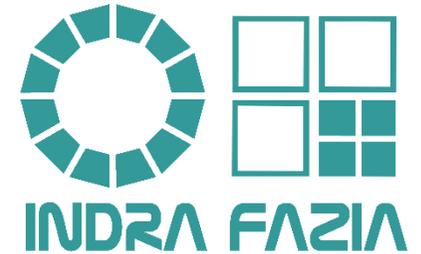
So far, what have been done?

INDRA-FAZIA at GANIL

- E789 – 2019 :  $^{58-64}\text{Ni} + ^{58-64}\text{Ni}$  @ 32 & 52 A MeV
- E818 - 2022 :  $^{36}\text{Ar} + ^{58}\text{Ni}$  @ 74 A MeV &  $^{58}\text{Ni} + ^{58}\text{Ni}$  @ 74 A MeV

To come soon (2025)

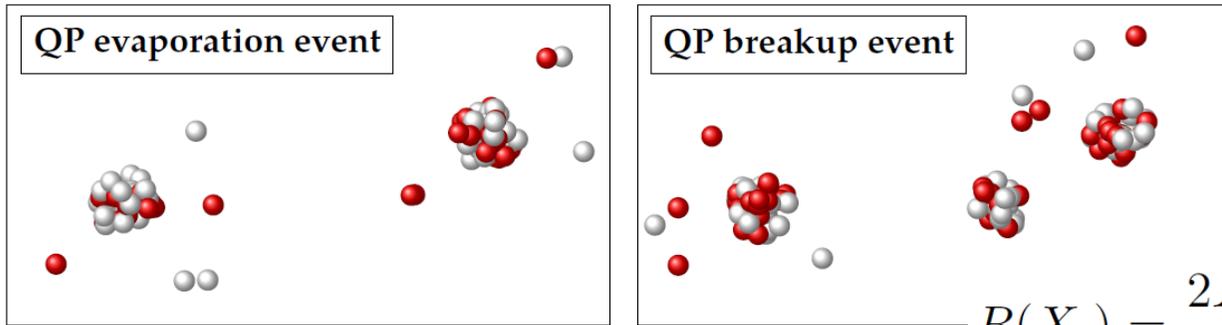
- E884 - 2025 :  $^{70}\text{Zn} + ^{27}\text{Al}, ^{70}\text{Zn}, ^{208}\text{Pb}$  @ 35 A MeV



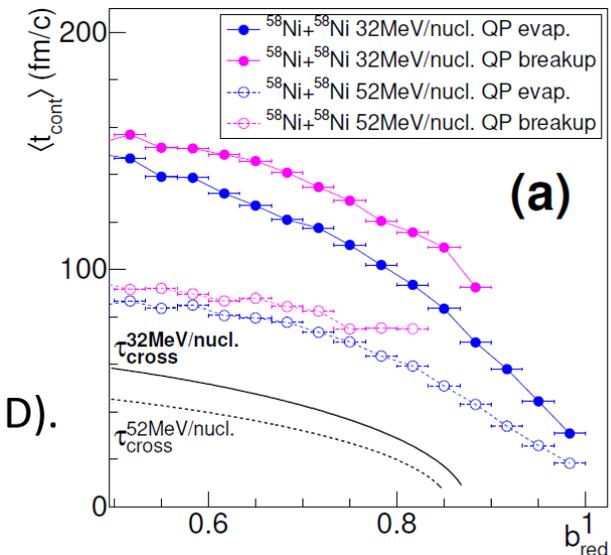
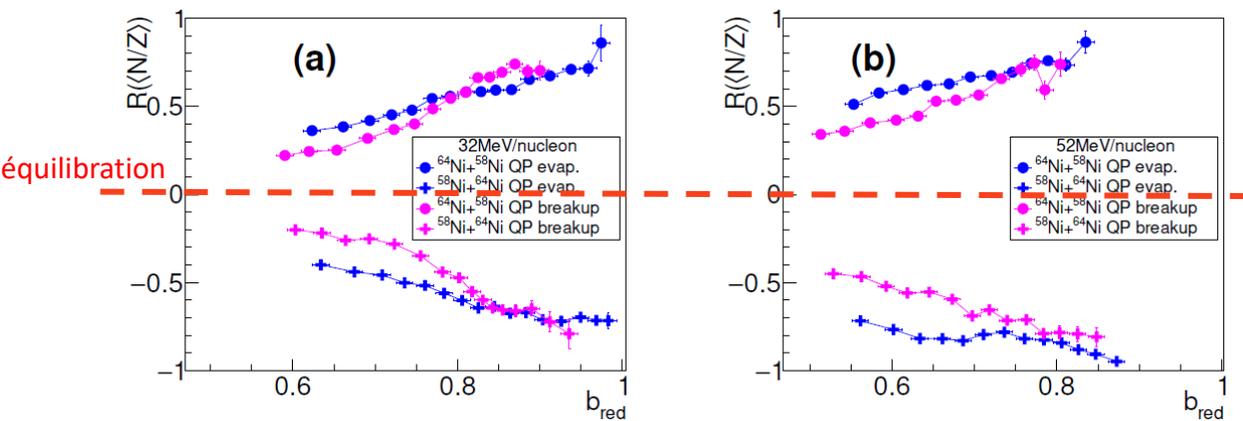
# E789: Studies on isospin transport ratio with $^{64-58}\text{Ni}+^{64-58}\text{Ni}$ @ 32 & 52 A MeV



- C. Ciampi et al. Physical Review C 106 (2022) 024603
- C. Ciampi et al. Physical Review C 108 (2023) 054611

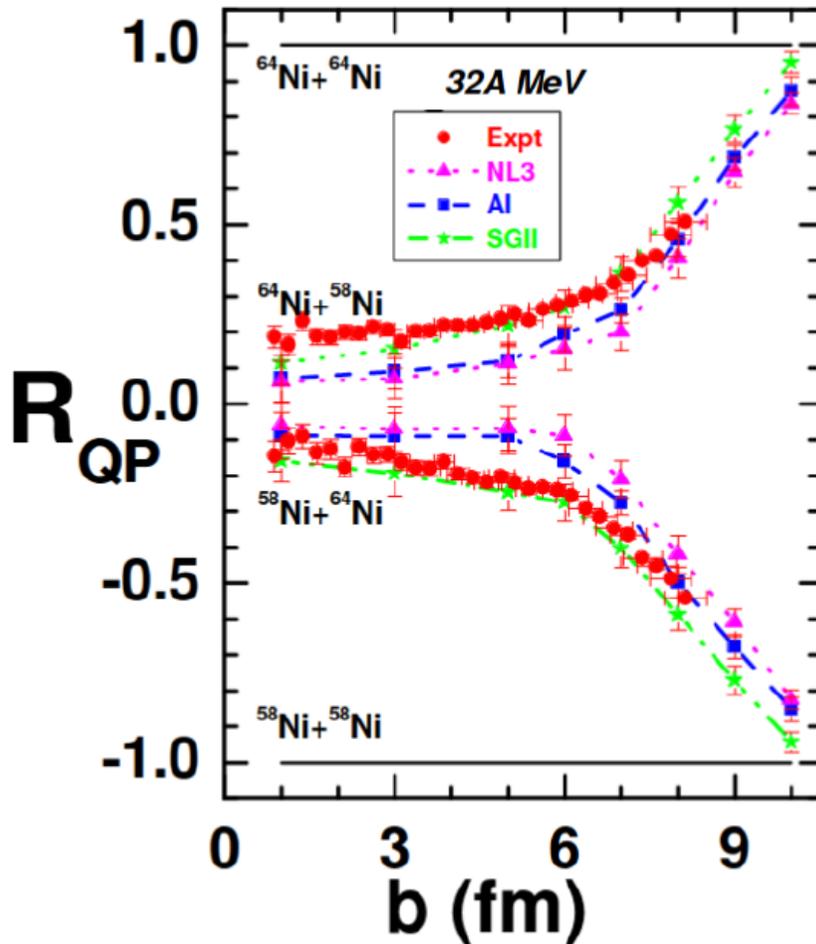


$$R(X_i) = \frac{2X_i - X_{6464} - X_{5858}}{X_{6464} - X_{5858}}$$



Contact time extraction via comparison with transport models (AMD).  
 Longer contact time for break-up respect to QP evaporation.  
 +time = more isospin equilibration

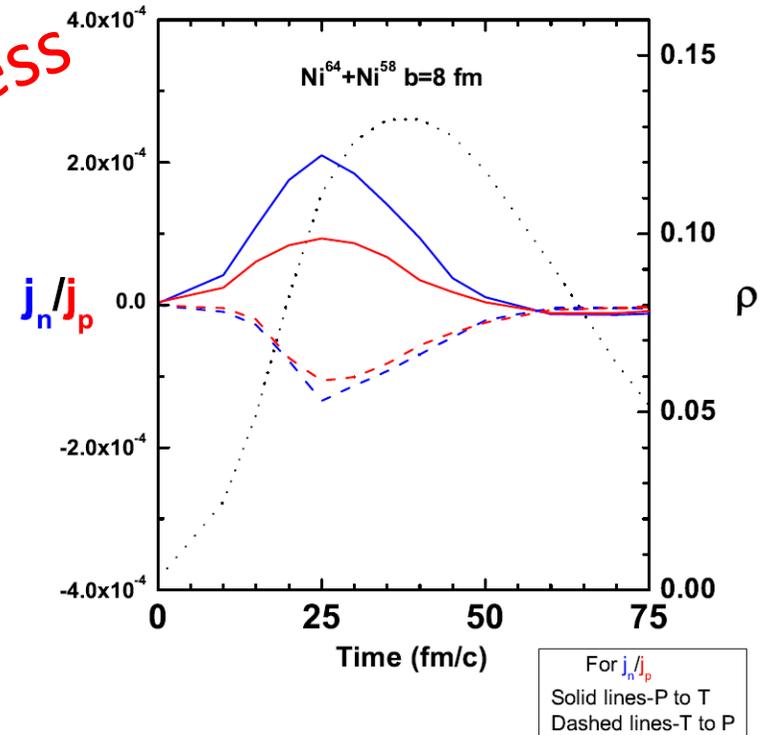
Constraining the parameters of the nuclear EoS



© Caterina Ciampi

In progress

Coll. FAZIA+ F. Gulminelli & Swagata Mallik (Kolkata)  
 BUU calculations with different equation of state vs Exp.



© Swagata Mallik

Next INDRA-FAZIA campaign at GANIL

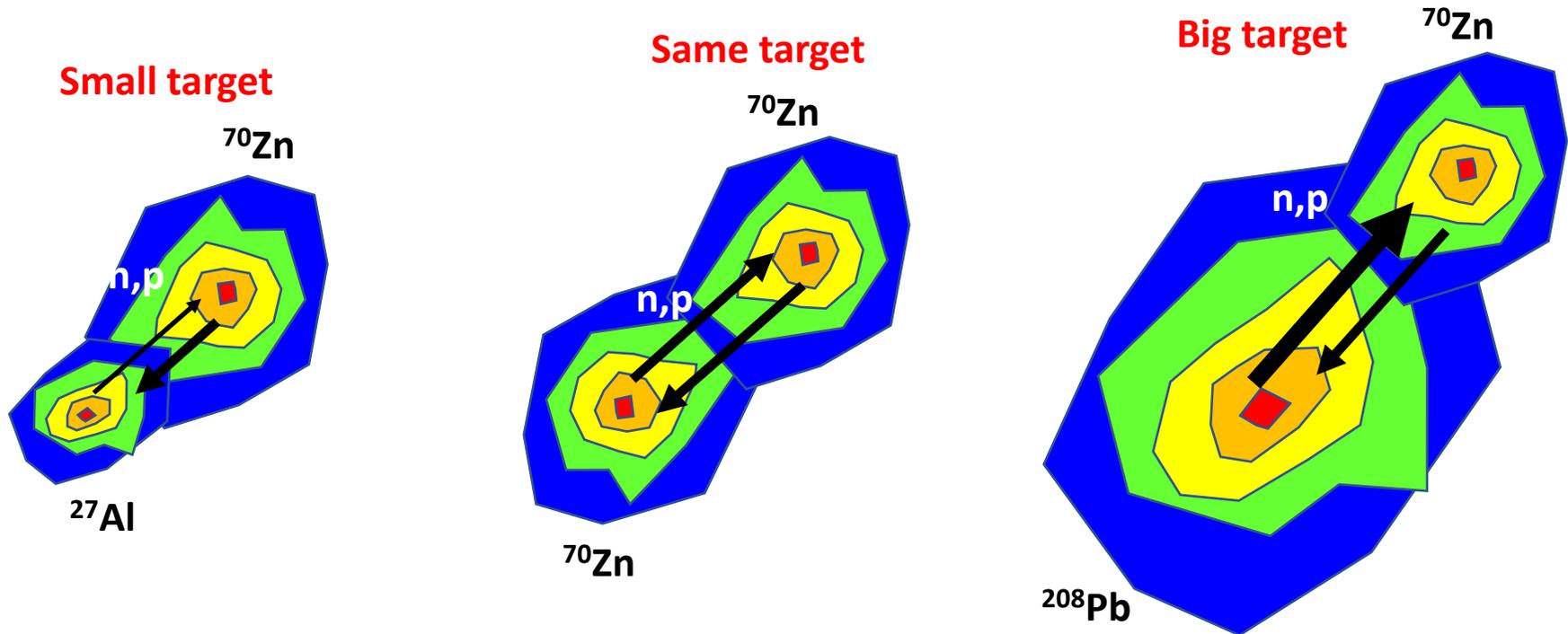
E884 - 2025 :  $^{70}\text{Zn} + ^{27}\text{Al}, ^{70}\text{Zn}, ^{208}\text{Pb}$  @ 35 A MeV

*"Impact of projectile-target size asymmetry on the isospin equilibration rate extracted From quasi projectile break-up reactions."*

We have seen in previous analyses how isospin transport evolves according to QP decay (evaporation vs break-up) see C. Ciampi et al.

What is the role of the target?

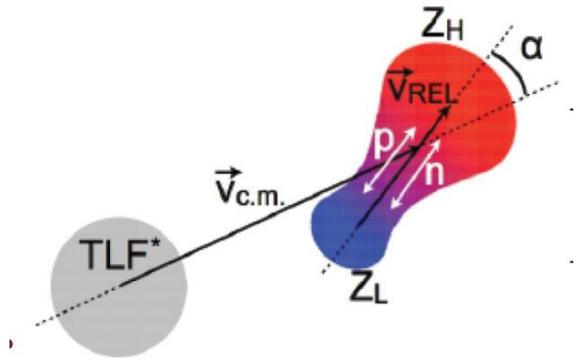
Target as an isospin reservoir and size influence on QP evolution.



# Next INDRA-FAZIA campaign at GANIL

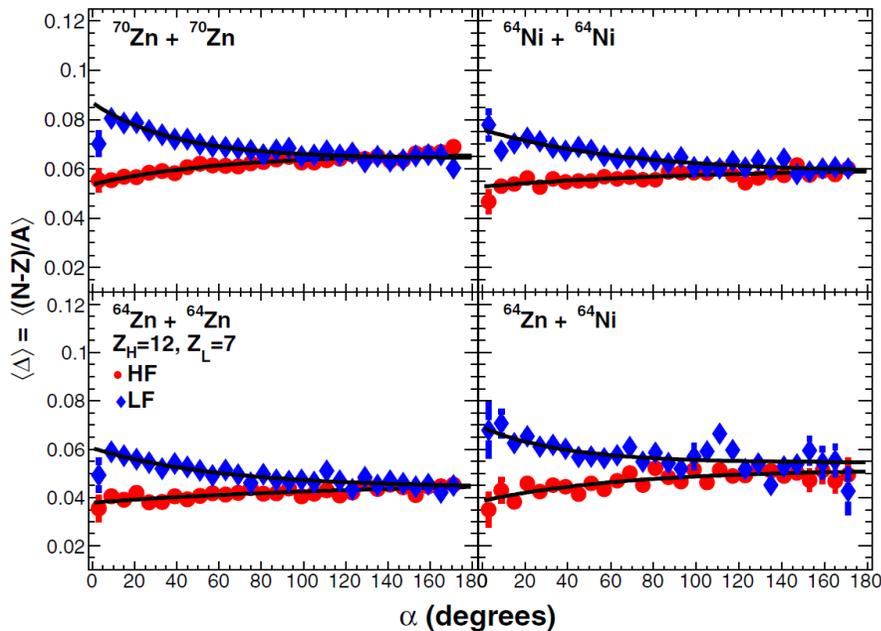
E884 - 2025 :  $^{70}\text{Zn} + ^{27}\text{Al}$ ,  $^{70}\text{Zn}$ ,  $^{208}\text{Pb}$  @ 35 A MeV

"Impact of projectile-target size asymmetry on the isospin equilibration rate extracted From quasi projectile break-up reactions."

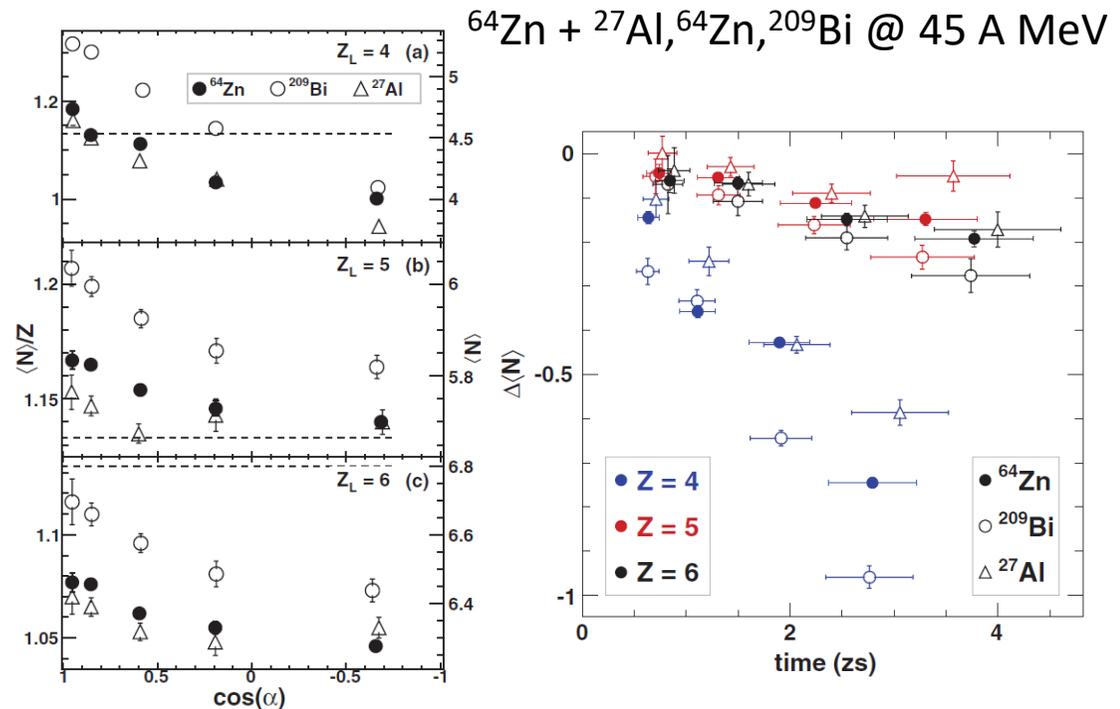


$\alpha$  angle between the QP-QT separation axis and the break-up axis can be used as a « clock » for the breakup time scale.

- Such interpretation is still debated
- As well the target role is contrasted in literature.



A. Rodriguez-Manso et al. PRC 95 (2017) 044604



K. Brown et al. PRC 87 (2013) 061601

# Experimental campaigns at GANIL: INDRA+FAZIA

Measurement of the  $^{12}\text{C}$  Hoyle state radius via double-excitation inelastic scattering

Diego Gruyer\*

*Université de Caen Normandie, ENSICAEN, CNRS/IN2P3, LPC Caen UMR6534, F-14000 Caen, France*

E881

Daniele Dell'Aquila

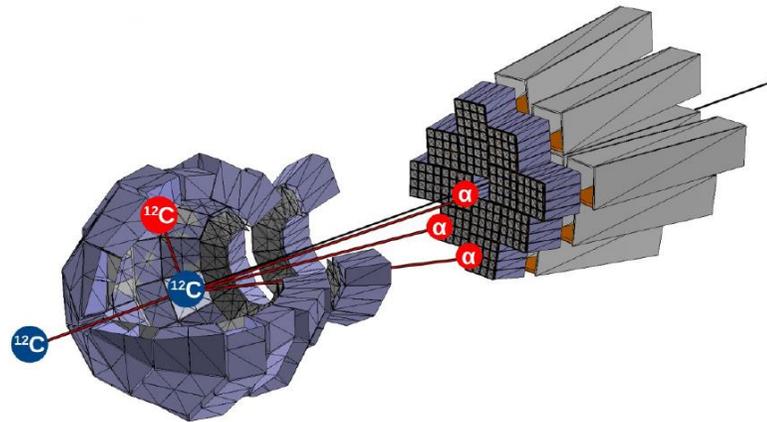
*Dipartimento di Fisica, Università degli Studi di Napoli 'Federico II', Naples, Italy*

(INDRA and FAZIA Collaborations)

The detailed properties of the  $^{12}\text{C}$  second  $0^+$  excited state, known as the Hoyle state, are both a challenge for nuclear structure theory and have a key role in the synthesis of the elements. We propose to measure the mean matter radius of this state by analyzing the diffraction structure of single- and double-excitation in  $^{12}\text{C}+^{12}\text{C}$  inelastic scattering at 105 MeV beam energy. The experimental setup will consist in the FAZIA detector.

**In addition to symmetry  
energy studies**

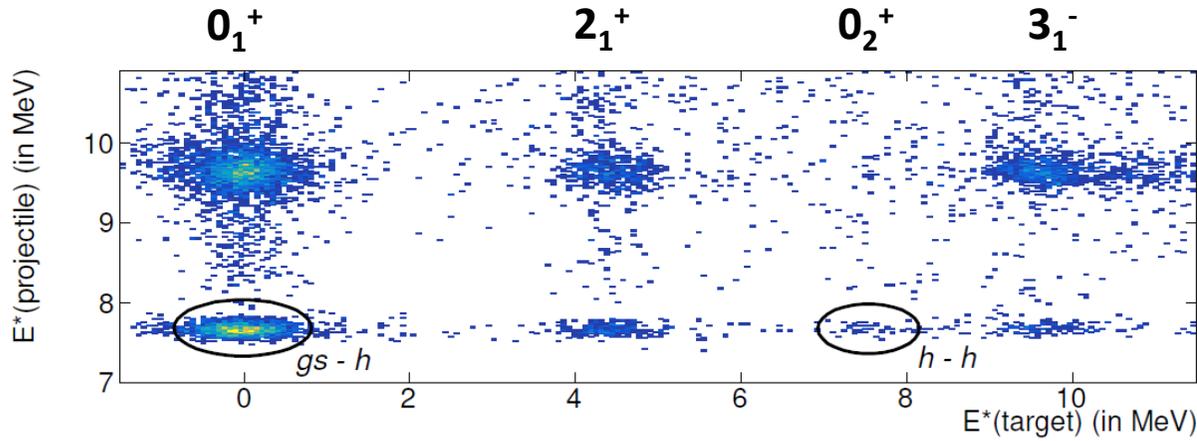
In 2022 during E818 campaign, we had calibration beams  $^{12}\text{C}$  at 8,75 and 13,75 A MeV. After having completed calibrations, during 6 hours we put a C target . For fun...



© Diego Gruyer

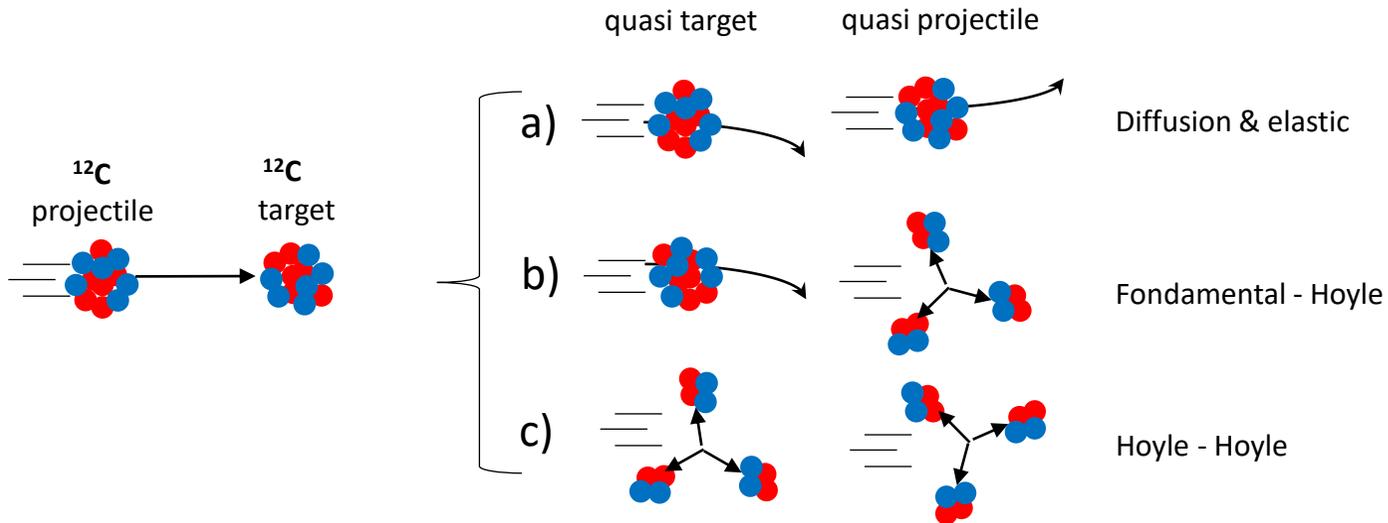
# Experimental campaigns at GANIL: INDRA+FAZIA

*In addition to symmetry energy studies*



*experimental data  
6 hours of beam test*

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# High energies - high density aspects of nuclear EoS

## Dense Nuclear Matter Equation of State from Heavy-Ion Collisions

Agnieszka Sorensen<sup>1</sup>, Kshitij Agarwal<sup>2</sup>, Kyle W. Brown<sup>3,4</sup>, Zbigniew Chajęcki<sup>5</sup>, Paweł Danielewicz<sup>2,6</sup>, Christian Drischler<sup>7</sup>, Stefano Gandolfi<sup>8</sup>, Jeremy W. Holt<sup>9,10</sup>, Matthias Kaminski<sup>11</sup>, Che-Ming Ko<sup>9,10</sup>, Rohit Kumar<sup>3</sup>, Bao-An Li<sup>12</sup>, William G. Lynch<sup>3,6</sup>, Alan B. McIntosh<sup>10</sup>, William G. Newton<sup>12</sup>, Scott Pratt<sup>3,6</sup>, Oleh Savchuk<sup>3,13</sup>, Maria Stefanik<sup>14</sup>, Ingo Tews<sup>8</sup>, ManYee Betty Tsang<sup>3,6</sup>, Ramona Vogt<sup>15,16</sup>, Hermann Wolter<sup>17</sup>, Hanna Zdrodzczyk<sup>18</sup>

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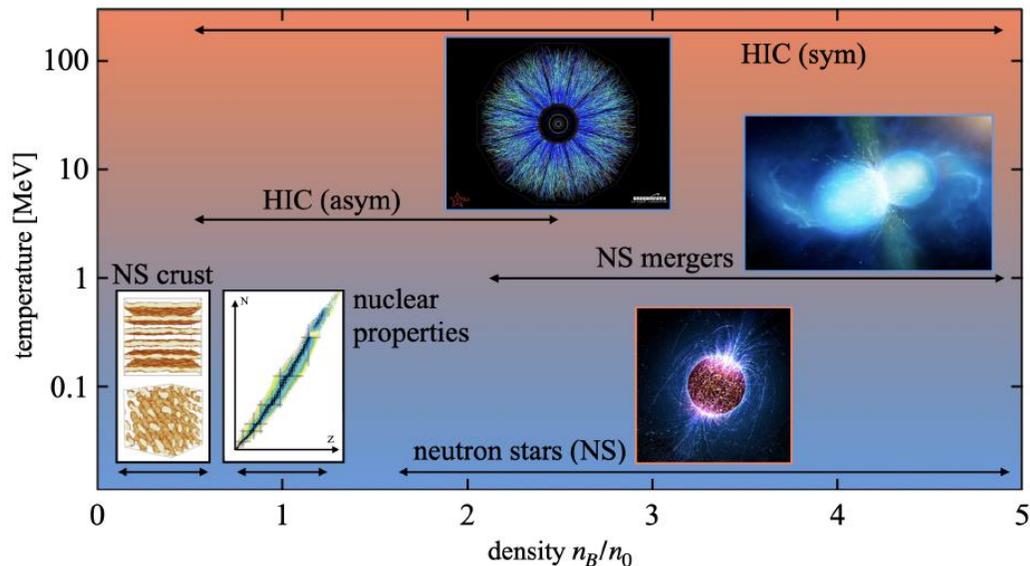
theoretical and experimental understanding of the properties of mergers. With advances in experiments probing nuclear matter astronomy, the next decade

strong interactions remain between a hadronic gas and a consistent with a crossover the finite-density region of The isospin-dependence of neutron-rich nuclei and the duration density. Moreover, the EOS in neutron-rich maturation density, leading to light (breaking the conformal or not known, but it is also early-symmetric nuclear maturation density, leading to opportunities for studying

heavy nuclei at intermediate about 25 GeV/nucleon baryon density and temperature about 5 times the nuclear a hundred MeV, respectively devoted to uncovering the studies at FRIB where collisions of rare isotopes. field measurements that temperature, and isospin

means of interpreting as at intermediate beam effort aimed at uncovering ends on the continued development the imminent results is an urgent need for a accelerate the development EOS from experimental realize the potential of matter EOS.

## Long Range Plan: Experimental side



High E radioactive beams, for larger  $\rho$ , T & isospin

### VI. Exploratory directions

- A. Dense nuclear matter EOS meeting extreme gravity and dark matter in supermassive neutron stars
- B. Nuclear EOS with reduced spatial dimensions

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# High energies - high density aspects of nuclear EoS

## Long Range Plan: Dense matter theory for heavy-ion collisions

### and neutron stars

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arXiv:2211.02224v2 [nucl-th] 8 Nov 2022

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## Long Range Plan: Theory side

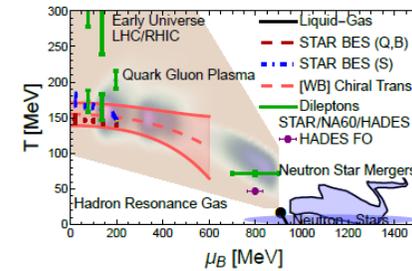


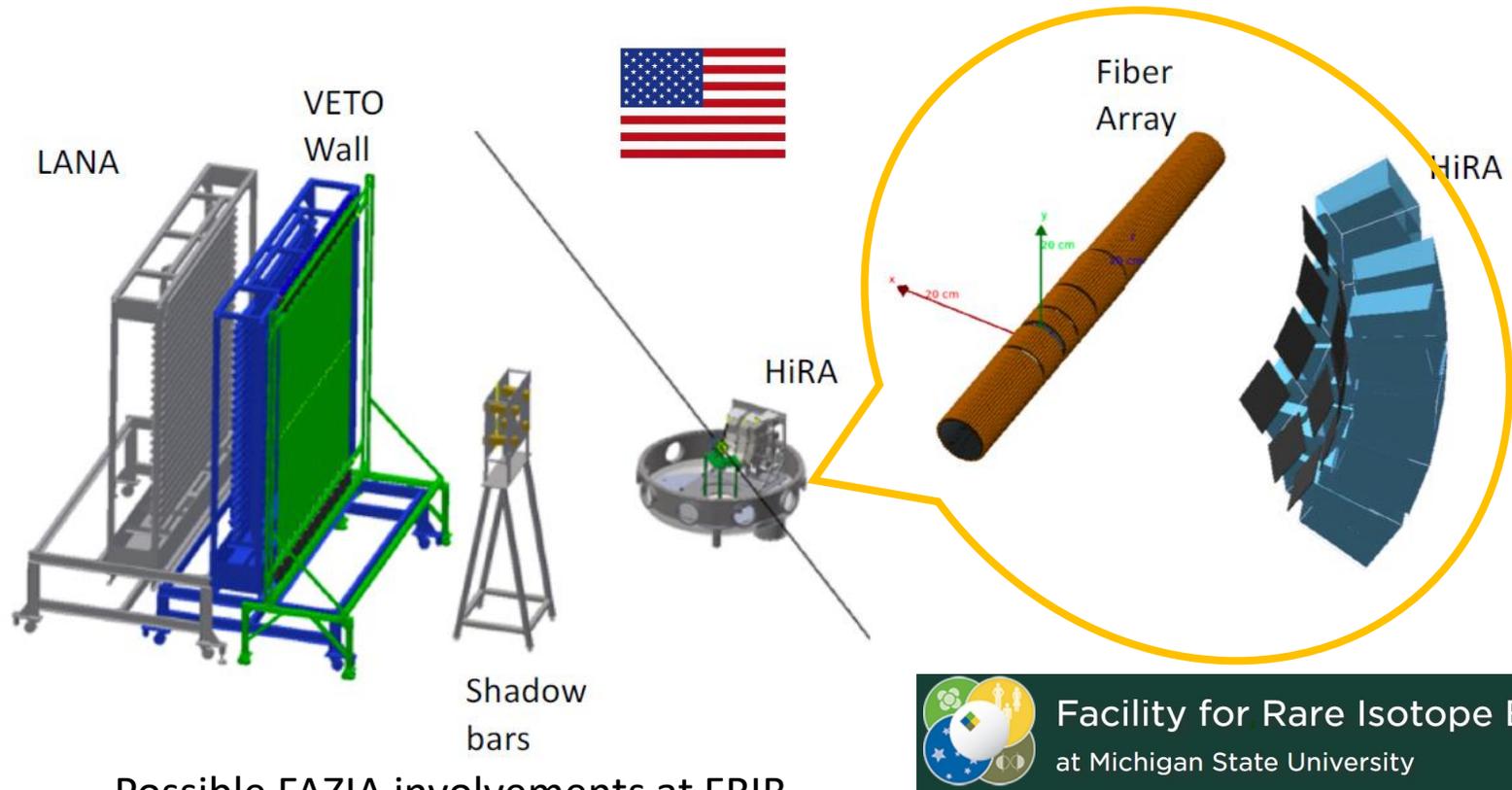
FIG. 1. QCD phase diagram with the latest interpretation of experimental data and dynamical simulations of regions probed by various systems. The zero baryon chemical potential axis follows the trajectory of the early universe and coincides with the Large Hadron Collider (LHC), RHIC, through the STAR BESII and fixed target (FXT) programs, explores the brown shaded region as inferred from relativistic viscous hydrodynamic simulations [46] for  $\sqrt{s_{NN}} = 3, 7.7, 27$  GeV. Estimates for neutron star mergers comes from numerical relativity simulations [47] and the  $T = 0$  neutron star range comes from various EOS estimates. Net-proton, K,  $\pi$  data from STAR were used to extract  $T, \mu_B$  of light particles at freeze-out (red) and net-K,  $\Lambda$  for strange particles (blue) from [48]. Dilepton measurements of temperatures within the quark-gluon plasma phase from STAR, NA60 [49], and HADES [6] are shown in green. Thermal model fits to HADES particle yields provide a freeze-out estimate for  $\sqrt{s_{NN}} = 2.4$  GeV [50] (some ambiguity still exists [51]), shown in maroon. The nuclear liquid-gas phase transition is based on experimental data from [52] and the  $\mu_B$  estimate from [53]. The chiral transition (light red) comes from lattice QCD calculations [54]. The possible QCD critical point and the associated 1st-order phase transition line are not shown due to uncertainty, which STAR BESII and FXT aim to reduce.

### I. OVERVIEW

Understanding the behavior of dense baryonic matter is one of the central problems in nuclear physics. Dense matter in this context is any nuclear system that contains a net

# American set-up foreseen so far for an approved experiment at FRIB (end 2025?)

$^{56,70}\text{Ni} + ^{58,64}\text{Ni}$  @ 175 A MeV



## Possible FAZIA involvements at FRIB

- Bringing some FAZIA blocks to join/complete the set-up?
  - Imagine our own set-up? Of course with local support.
  - Lol for PAC3 at FRIB?
- => Everything is open, let's discuss and collaborate.

Building additional FAZIA blocks for RAON 2024-....



Two production methods: Separation on line (ISOL) and in flight fragmentation (IF):

- ISOL:  $^{238}\text{U}$  fission by proton beam at 70 MeV
- IF:  $^{238}\text{U}$  beam at 200 A MeV (8,3p  $\mu\text{A}$ )

Final project associates a large variety of radioactive beams with a big range of beam energies.

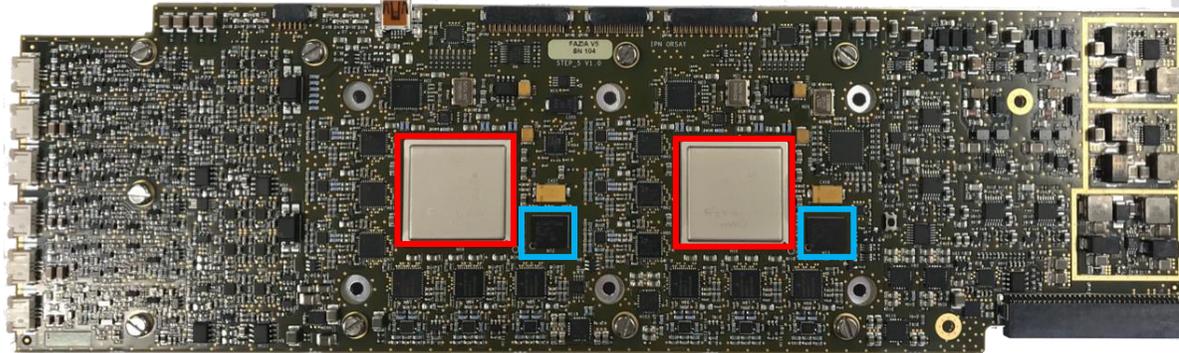
For example  $^{132}\text{Sn}$  at 250 A MeV with  $10^9$  pps!

But not before 2030 after building SCL2!

# New developments for the future

## FAZIA Front End Electronics update in South Korea

Old/current FAZIA FEE card (2 FPGA Virtex 5)



Analog stage

Digital stage

Converters stage



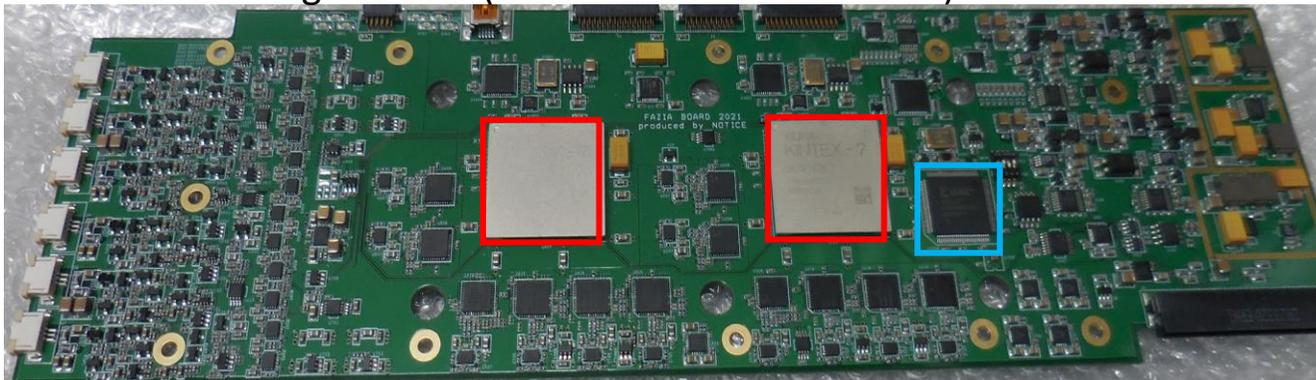
IJCLab Orsay  
Naples

*Courtesy of Minjung Kweon,  
Jiyong Kim & Simone Valdré*

New prototypes FAZIA FEE card (FPGA Virtex 5->Kintex 7)

One Complex Programmable Logic Device chip (VHDL) makes two FPGAs

New clock generator (old one no more available) 250->500 MHz



NOTICE Co  
Korea

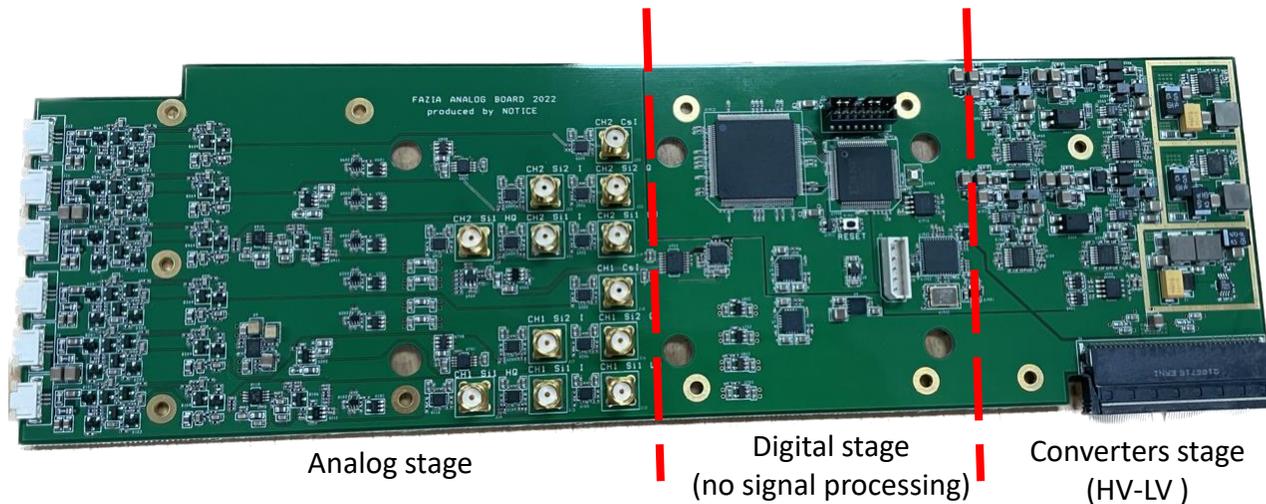
Update of the components

After a series of tests during summer 2022 and 2023: the two new prototypes were validated!

# New developments for the future

FAZIA Korean initiative: a simpler version of the FEE  
Development of a “small” card with mainly the analog part (PreAmp)  
=> Coaxial outputs

*Courtesy of Minjung Kweon  
& Jiyong Kim*



NOTICE Co  
Korea

This initiative brings new developments towards simpler FAZIA block

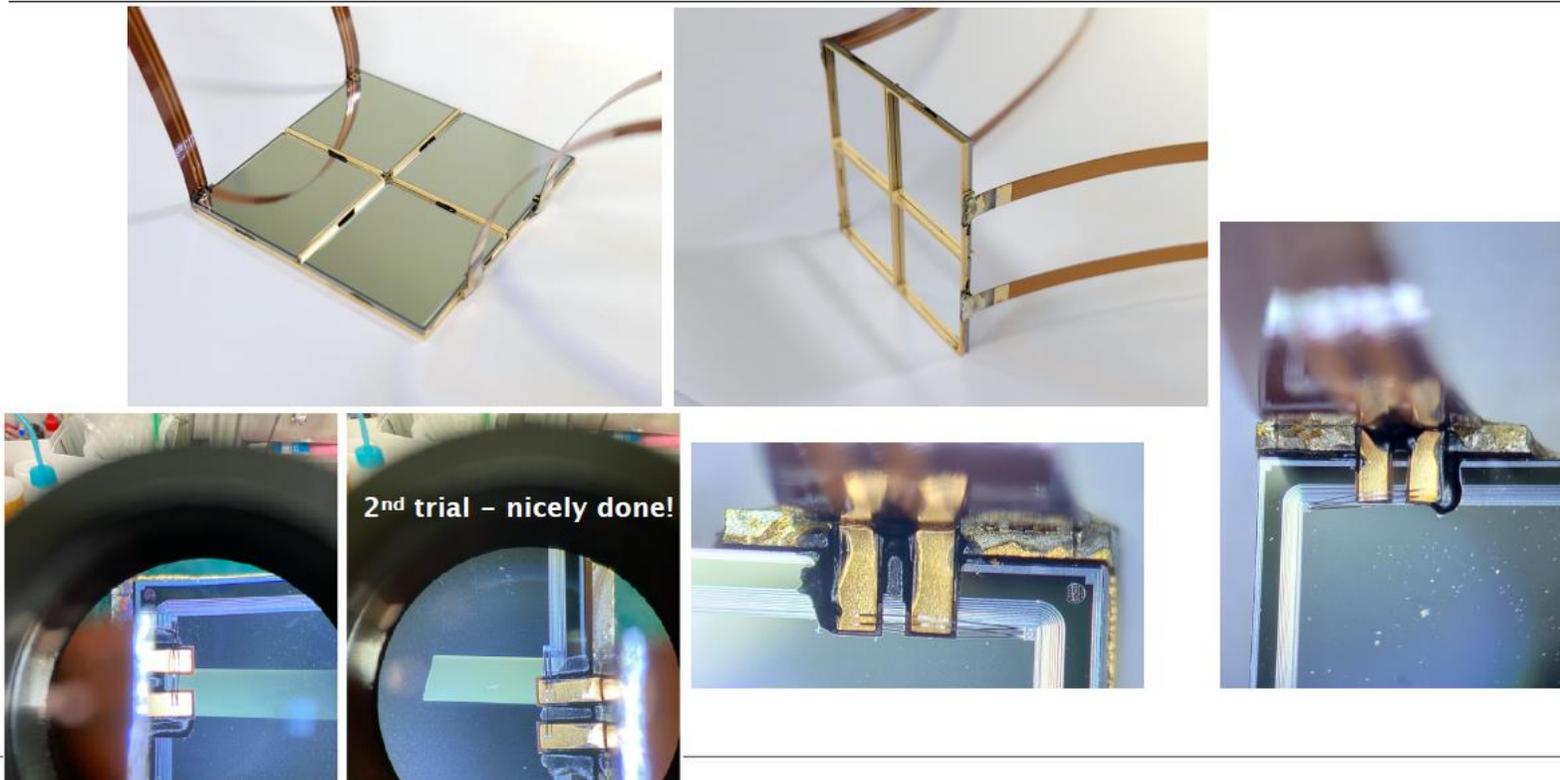
- Increasing the angular coverage especially at larger angle (mid velocity) or backward
- More versatile for experiments with other groups (LISE, ACTAR, FRIB...)

## New developments for the future

- Korean colleagues already delivered 500 and 750  $\mu\text{m}$  thick for FAZIA at GANIL
- New silicon chip detectors for FAZIA developed in Korea as well (100 to 1000  $\mu\text{m}$ )
- Better partnership between them and the detector companies too

### Quartetto produced by MEMSPACK (chip mounting & wire-bonding)

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*Courtesy of Minjung Kweon*

## First step: The low energy RI beams (GANIL SPIRAL 2, SPES, RAON)

For low energy experiments (radioactive ion beams SPES, Spiral 2, FRAISE...) we must lower the identification thresholds => Thin silicon prototypes (20-30  $\mu\text{m}$ ).

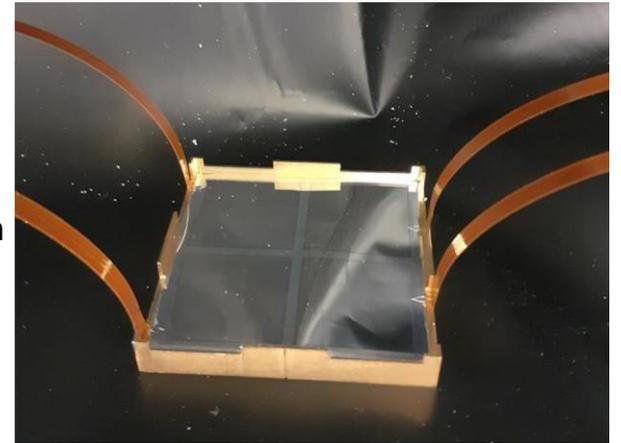
INFN budget 20 k€ in 2023 + IN2P3

⇒ Ordering 4 protos at Micron semi-conductor.

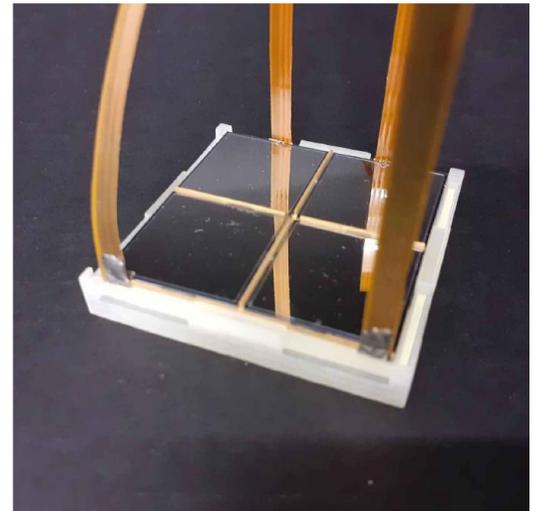
⇒ Beam test at LNS Legnaro 2024? GANIL 2025?

⇒ Interesting developments to lower the dead zone for other silicon detectors.

New quartet:  
1 silicium/4 20  $\mu\text{m}$



Quartet standard:  
4 siliciums 500  $\mu\text{m}$



# Conclusions

## Plans for symmetry energy researches.

### **Now**

Experimental program at GANIL with INDRA & FAZIA (in parallel with other thematic)

- Complete maintenance + improvements on electronics and detectors.

### **Future**

Development of additional FAZIA blocks in Korea for RAON.

- New updated electronics FEE boards.
- More simple electronic cards for better versatility.
- New silicon detectors of various thicknesses according to beam energies and studies.

Participation to experiments at FRIB.

- First as partner, with some spare blocks added to existing devices.
- As a whole device??