

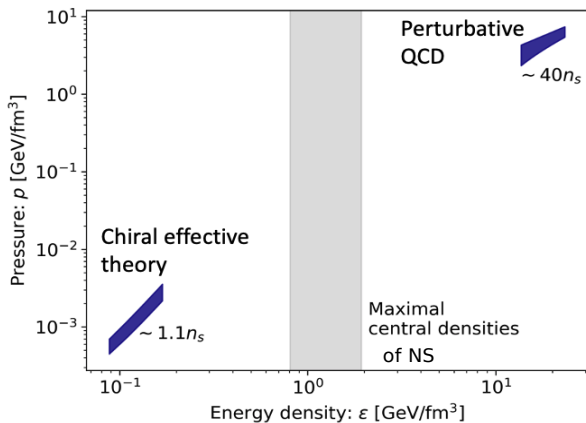
CALIBRATING THE GLOBAL BEHAVIOUR OF EOS IN A MACHINE LEARNING APPROACH

Adil Imam

Based on : **arXiv:2407.08553**

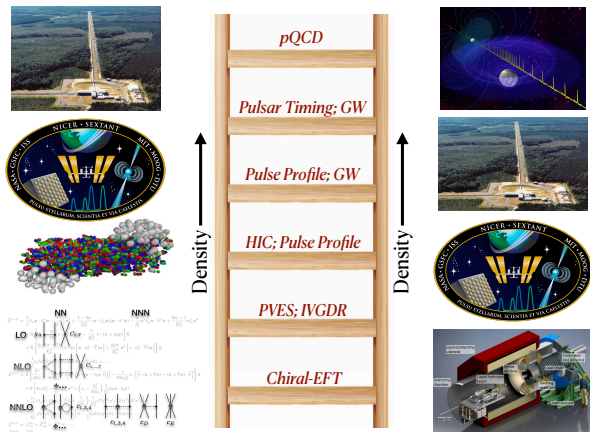


INTRODUCTION : SIMPLISTIC VIEW OF PHASE DIAGRAM



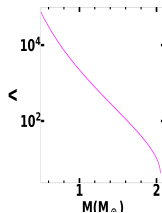
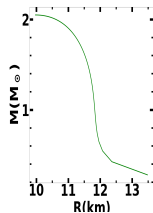
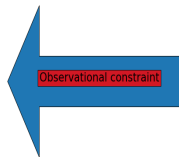
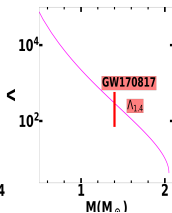
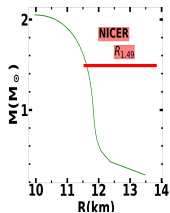
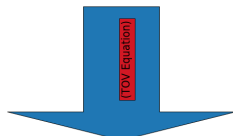
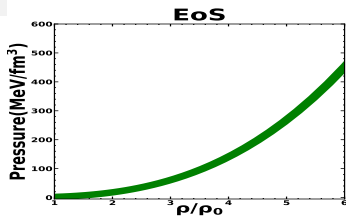
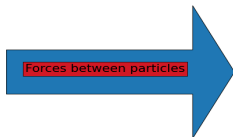
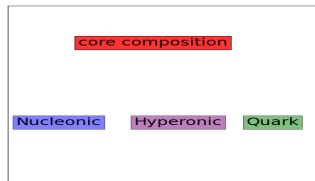
arXiv:2211.11414v1

THE DENSITY LADDER



arXiv:2304.05441v1

MICRO-MACRO COLLISION



THE EOS FOR NS MATTER :

NUCLEONIC CORE COMPOSITION

- The NS core is composed of neutron (n), proton(p), electron(e) and muon(μ)

THE EOS FOR NS MATTER :

NUCLEONIC CORE COMPOSITION

- The NS core is composed of neutron (n), proton(p), electron(e) and muon(μ)
- The EoS is determined by the features of nucleon-nucleon (N-N) interaction

THE EOS FOR NS MATTER

THE PARABOLIC APPROXIMATION :

- Energy per nucleon for neutron star matter at a given nucleon density ρ and asymmetry $\delta = (\frac{\rho_n}{\rho} - \frac{\rho_p}{\rho})$ can be written as,

$$\varepsilon(\rho, \delta) = \sum \frac{\partial^n}{\partial \delta^n} \varepsilon(\rho, \delta) |_{\delta=0} = \mathbf{e}(\rho, \mathbf{0}) + \mathbf{e}_{sym}(\rho) \delta^2 + \dots,$$

$\mathbf{e}(\rho, \mathbf{0})$: Energy for SNM and

$\mathbf{e}_{sym}(\rho) = \varepsilon(\rho, 1) - \varepsilon(\rho, 0)$: Density dependent symmetry energy

THE EOS FOR NS MATTER

THE PARABOLIC APPROXIMATION :

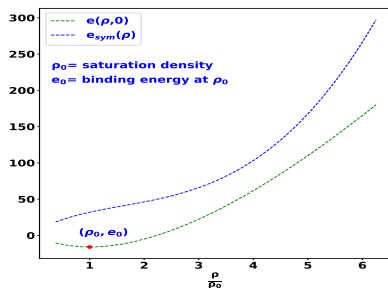


FIGURE: Energy per nucleon for Symmetric Nuclear Matter (SNM) (green curve) and symmetry energy (blue curve)

■ Nuclear Matter Parameters (NMPs) $\propto \frac{\partial^n}{\partial \rho^n} e(\rho, 0)$ or $\propto \frac{\partial^n}{\partial \rho^n} e_{sym}(\rho)$

THE EOS OF NS MATTER

NUCLEAR MATTER PARAMETERS (NMPS) :

- Taylor expansion : $\varepsilon(\rho, \delta) = \sum_{n=0}^4 (a_n + b_n \delta^2) \left(\frac{\rho - \rho_0}{3\rho_0} \right)^n$

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- Taylor expansion : $\varepsilon(\rho, \delta) = \sum_{n=0}^4 (a_n + b_n \delta^2) \left(\frac{\rho - \rho_0}{3\rho_0}\right)^n$
- **SNM parameters (at ρ_0) :**
- $e_0 \equiv$ Binding energy = $e(\rho, 0)|_{\rho=\rho_0} = a_0$
- $K_0 \equiv$ Incompressibility coefficient $\propto \frac{\partial^2 e(\rho, 0)}{\partial \rho^2} |_{\rho=\rho_0} = a_2$
- $Q_0(Z_0) \equiv$ Third(Fourth)order derivative = $a_3(a_4)$
- **Symmetry Energy parameters (at ρ_0) :**
- $J_0 \equiv$ Symmetry energy = $e_{sym}(\rho)|_{\rho=\rho_0} = b_0$
- $L_0 \equiv$ Slope of symmetry energy $\propto \frac{\partial e_{sym}(\rho)}{\partial \rho} |_{\rho=\rho_0} = b_1$
- $K_{sym,0} \equiv$ Curvature of symmetry energy $\propto \frac{\partial^2 e_{sym}(\rho)}{\partial \rho^2} |_{\rho=\rho_0} = b_2$
- $Q_{sym,0}(Z_{sym,0}) \equiv$ Third(Fourth)order derivative = $b_3(b_4)$

THE EOS OF NS MATTER

THE $\frac{n}{3}$ EXPANSION MODEL :

- $\frac{n}{3}$ Expansion: $\varepsilon(\rho, \delta) = \sum_{n=2}^6 (c_n + d_n \delta^2) \left(\frac{\rho}{\rho_0}\right)^{n/3}$

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■

$$\begin{pmatrix} e_0 \\ 0 \\ K_0 \\ Q_0 \\ Z_0 \end{pmatrix} = \begin{pmatrix} 1 & 1 & 1 & 1 & 1 \\ 2 & 3 & 4 & 5 & 6 \\ -2 & 0 & 4 & 10 & 18 \\ 8 & 0 & -8 & -10 & 0 \\ -56 & 0 & 40 & 40 & 0 \end{pmatrix} \begin{pmatrix} c_0 \\ c_1 \\ c_2 \\ c_3 \\ c_4 \end{pmatrix}$$
$$\begin{pmatrix} J_0 \\ L_0 \\ K_{\text{sym},0} \\ Q_{\text{sym},0} \\ Z_{\text{sym},0} \end{pmatrix} = \begin{pmatrix} 1 & 1 & 1 & 1 & 1 \\ 2 & 3 & 4 & 5 & 6 \\ -2 & 0 & 4 & 10 & 18 \\ 8 & 0 & -8 & -10 & 0 \\ -56 & 0 & 40 & 40 & 0 \end{pmatrix} \begin{pmatrix} d_0 \\ d_1 \\ d_2 \\ d_3 \\ d_4 \end{pmatrix} .$$

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- Symmetry energy : $e_{sym}(\rho) > 0$
- $M_{TOV} \geq$ mass of observed massive NS (non-rotating)
(M_{TOV} = maximum mass in the stable branch)

KNOWLEDGE ABOUT THE NMPS

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- The parameters which are known within about 10 percent: J_0, K_0
- The parameters which are known within about 50 percent: L_0
- The parameters which are almost unknown:
 $Q_0, Z_0, K_{sym,0}, Q_{sym,0}, Z_{sym,0}$

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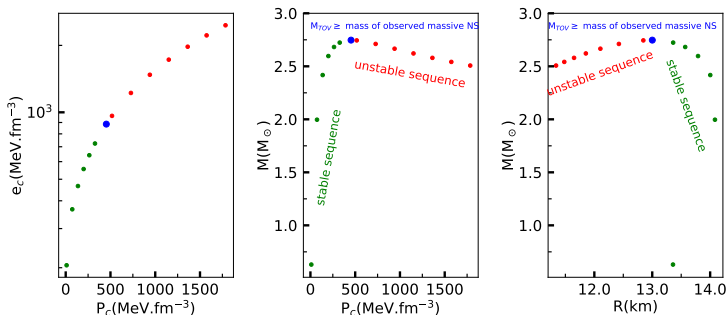


CALIBRATING THE GLOBAL BEHAVIOUR OF EoS

■ NMPs \rightarrow EoS \rightarrow TOV \rightarrow NS properties

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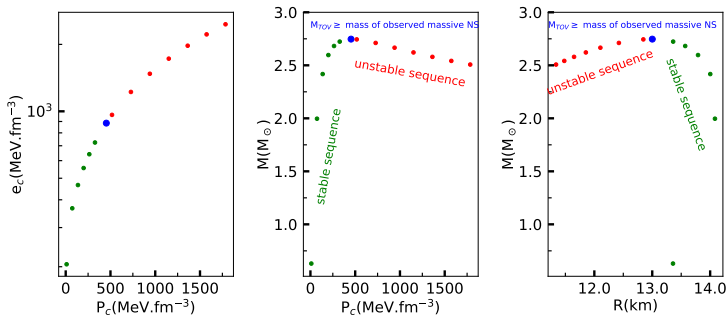
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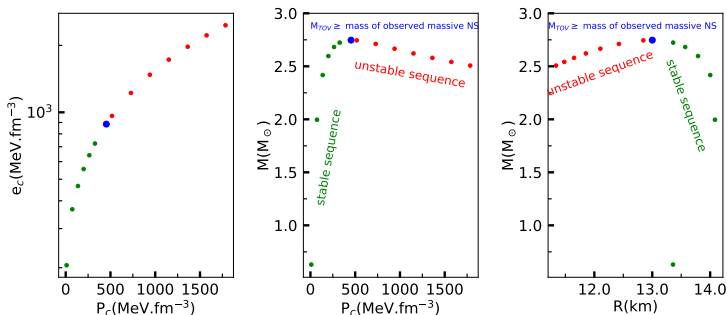


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■ NMPs \rightarrow NS properties

■ We use a machine learning approach : Symbolic Regression

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- A K-fold validation procedure will give multiple equations in each of the k fold
- To choose the best equation one has to perform a goodness of fit

CALIBRATING THE GLOBAL BEHAVIOUR OF EOS

SYMBOLIC REGRESSION :



$$R^2 = 1 - \frac{\sum (y_i - \hat{y}_i)^2}{\sum (y_i - \bar{y})^2}$$

y_i : observed or actual value

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$$RMSE = \sqrt{\frac{\sum(y_i - \hat{y}_i)^2}{n}}$$

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$$\begin{aligned}M_{\max} &= 0.14\hat{K}_0 - 0.02\hat{L}_0 + 0.27\hat{Q}_0 + 0.16\hat{Z}_0 + 2.33 \\R_{1.4} &= -0.31(\hat{J}_0 - \hat{Q}_0 - \hat{Q}_{\text{sym}0}) - 0.62\hat{K}_{\text{sym}0} + 1.75\hat{L}_0 + 14.22 \\ \Lambda_{1.4} &= -54.02(\hat{J}_0 - \hat{K}_0 - \hat{Q}_0 - \hat{Q}_{\text{sym}0}) + 76.09\hat{K}_{\text{sym}0} \\ &+ 178.12\hat{L}_0 + 737.93\end{aligned}$$

CALIBRATING THE GLOBAL BEHAVIOUR OF EOS

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- GW170817, mass-radius posterior data from NICER

MAPPING NS PROPERTIES TO NMPS

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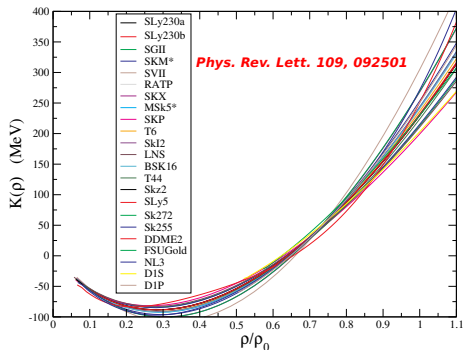


FIGURE: Variation of incompressibility coefficient, K with density.

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TABLE: S_0 : All the data except E_{pnm} and M_c ; S_{ALL} : All the data.

	Scenario	K_0	Q_0	J_0	L_0	K_{sym0}
ref	-	266	-90	33.17	67	-47
\mathcal{T}_{SRM}	S_0	256^{+46}_{-56}	-148^{+255}_{-266}	$32.92^{+1.36}_{-1.33}$	65^{+14}_{-14}	-27^{+81}_{-79}
	S_{ALL}	262^{+21}_{-23}	-167^{+175}_{-194}	$32.97^{+1.29}_{-1.26}$	65^{+14}_{-14}	-26^{+79}_{-80}
\mathcal{T}_{TOV}	S_{ALL}	262^{+20}_{-20}	-203^{+182}_{-176}	$32.74^{+1.24}_{-1.24}$	63^{+14}_{-13}	-34^{+76}_{-74}

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- Bayesian inferences are performed with realistic data including E_{pnm} and M_c

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TOV vs SYMBOLIC REGRESSION MODELS :

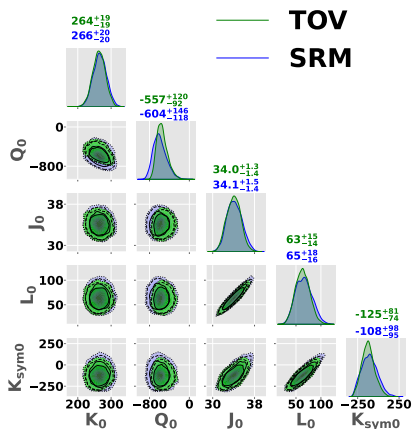


FIGURE: Posterior obtained using symbolic regression models and TOV solutions

SUMMARY

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- Symbolic regression based equations are constructed to replace EoS computation and solutions of TOV equations
- Using TOV and SRMs in Bayesian inference give comparable results while SRMs provides results within 20 minutes which is 100 times faster than the traditional TOV approach

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3. Dr. Naresh K. Patra

Thank You

