

# NUSYM24 – XII<sup>th</sup> International Symposium on Nuclear Symmetry Energy

**A new paradigm in the consistent extraction of  
nuclear symmetry energy and related properties  
using the relativistic application of coherent  
density fluctuation model**

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# Introduction & need for Isospin based study



THAPAR INSTITUTE  
OF ENGINEERING & TECHNOLOGY  
(Deemed to be University)

- ✓ Exploration of nuclear properties close to the  $\beta$ -stability line can be easily performed using the traditional bulk properties.
- ✓ However, moving away from the stability line, the dominance of traditional observables begins to drop, and isospin-dependent quantities come into prominence.

✓ Isospin asymmetry  $\Rightarrow \alpha = \frac{(N-Z)}{N+Z}$

*Phys. Rev. C* **85**, 064319 (2012).

*Phys. Rev. C* **97**, 024322 (2018)

*Chin. Phys. C* **46**, 084101 (2022).

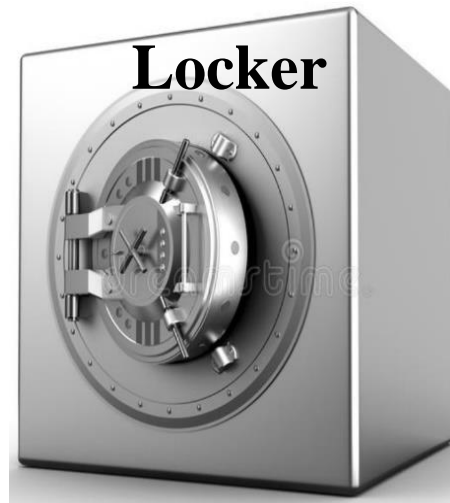
- Symmetry energy  $\rightarrow$  variation of **B.E.** as the neutron to proton ratio of a nuclear system is varied

$$S^{NM}(\rho) = \frac{1}{2} \left( \frac{\partial^2 \varepsilon}{\partial \alpha^2} \right)$$

## ❖ Current aim to study finite nuclei using Symmetry energy

### ○ **Obstacle**

Symmetry energy is not directly measurable



*Phys. Rev. C 97, 024322 (2018)*



- CDFM : Density distribution of nuclear matter fluctuates around average distribution while maintaining spherical symmetry and uniformity.
- The density  $\rho(\mathbf{r}, \mathbf{r}')$  of a finite nucleus can be rewritten as the coherent superposition of infinite number of one-body density matrix (OBDM)  $\rho_x(\mathbf{r}, \mathbf{r}')$  for spherical parts of nuclear matter called **Fluctons**.

*Bulg. J. Phys.* **6**, 151 (1979).

*Z. Phys. A* **304**, 239 (1982).

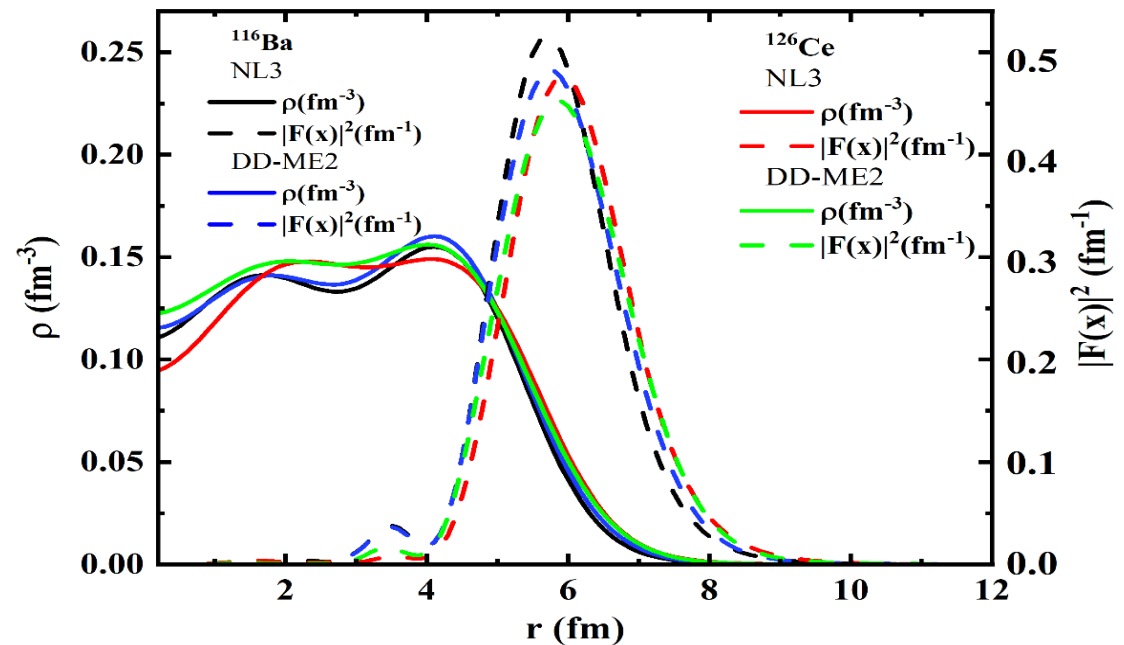
# Methodology of CDFM

- Spherical flucton density:

$$\rho_o(x) = 3A/4\pi x^3$$

- Weight function:

$$|\mathcal{F}(x)|^2 = -\frac{1}{\rho_o(x)} \frac{d\rho(r)}{dr} \Big|_{r=x}$$



*Phys. Rev. C 85, 064319 (2012).*

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Total density distribution  $\rho$  and corresponding weight function  $|\mathcal{F}(x)|^2$  as a function of nuclear distance  $r$

## Symmetry energy:

$$S = \int_0^\infty dx |\mathcal{F}(x)|^2 S^{NM}(\rho(x))$$

$$\kappa = \frac{S_V}{S_S}$$

Volume  
Symmetry Egy

Surface  
Symmetry Egy

$$S_V = S \left( 1 + \frac{1}{\kappa A^{1/3}} \right)$$

$$S_S = \frac{S}{\kappa} \left( 1 + \frac{1}{\kappa A^{1/3}} \right)$$

*Phys. Rev. C 85, 064319 (2012).*

*Phys. Rev. C97, 024322 (2018)*

# Choice of energy density functional



- The original CDFM formalism makes use of Brückner's prescription.
- The studies based on Brückner's prescription have significantly improved our understanding of nuclear matter properties.
- The studies have successfully verified various experimental shell closures across various isotopic & isotonic chains.

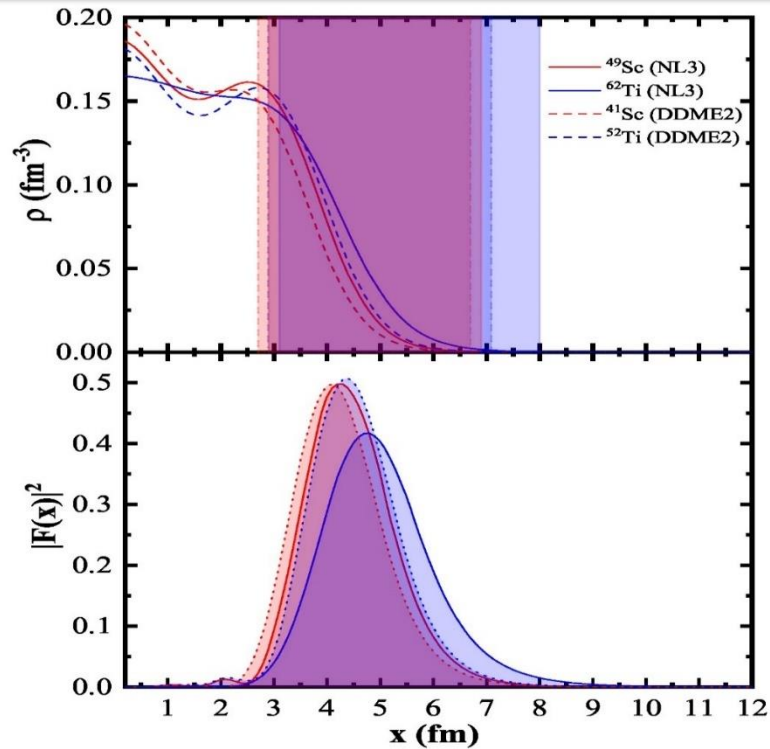
*Phys. Rev. C* **85**, 064319 (2012).

*Phys. Rev. C* **97**, 024322 (2018)

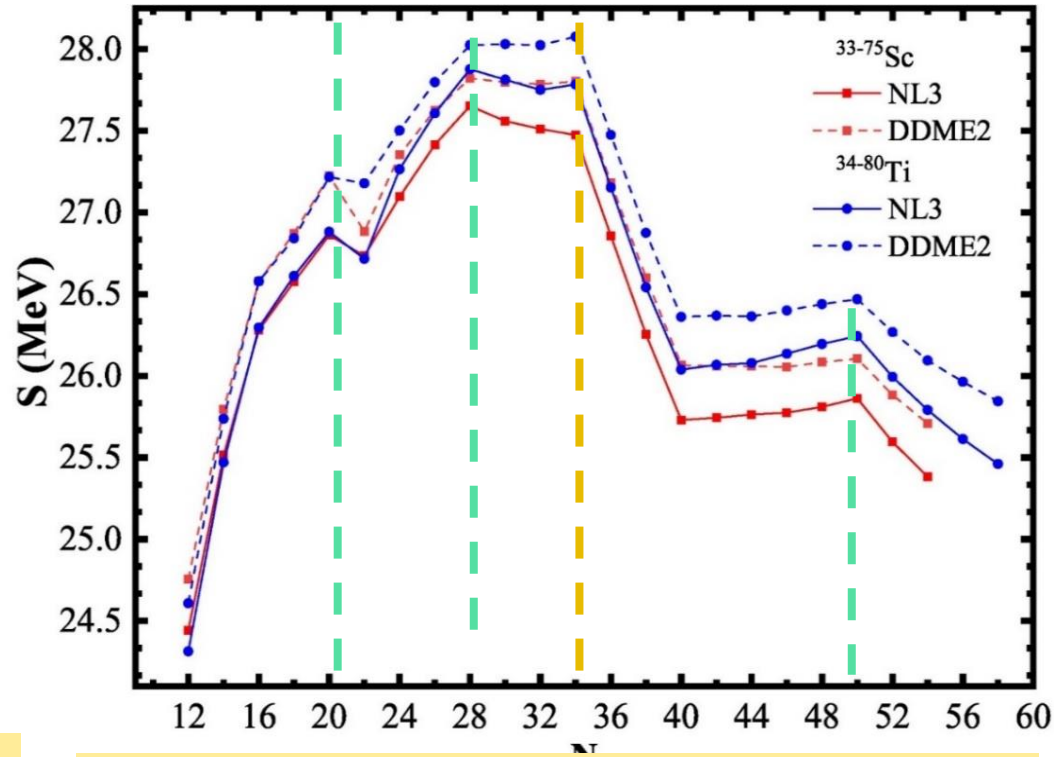
*Chin. Phys. C* **46**, 084101 (2022).

# Using CDFM with Brueckner EDF

Verified Exp. Shell closure at  $N = 20, 28$  &  $50$ .  
 Predicted novel shell closure at  $N = 34$ .



Total density distribution  $\rho$  (upper panel) and corresponding weight function  $|F(x)|^2$  (lower panel) as a function of nuclear distance  $x$



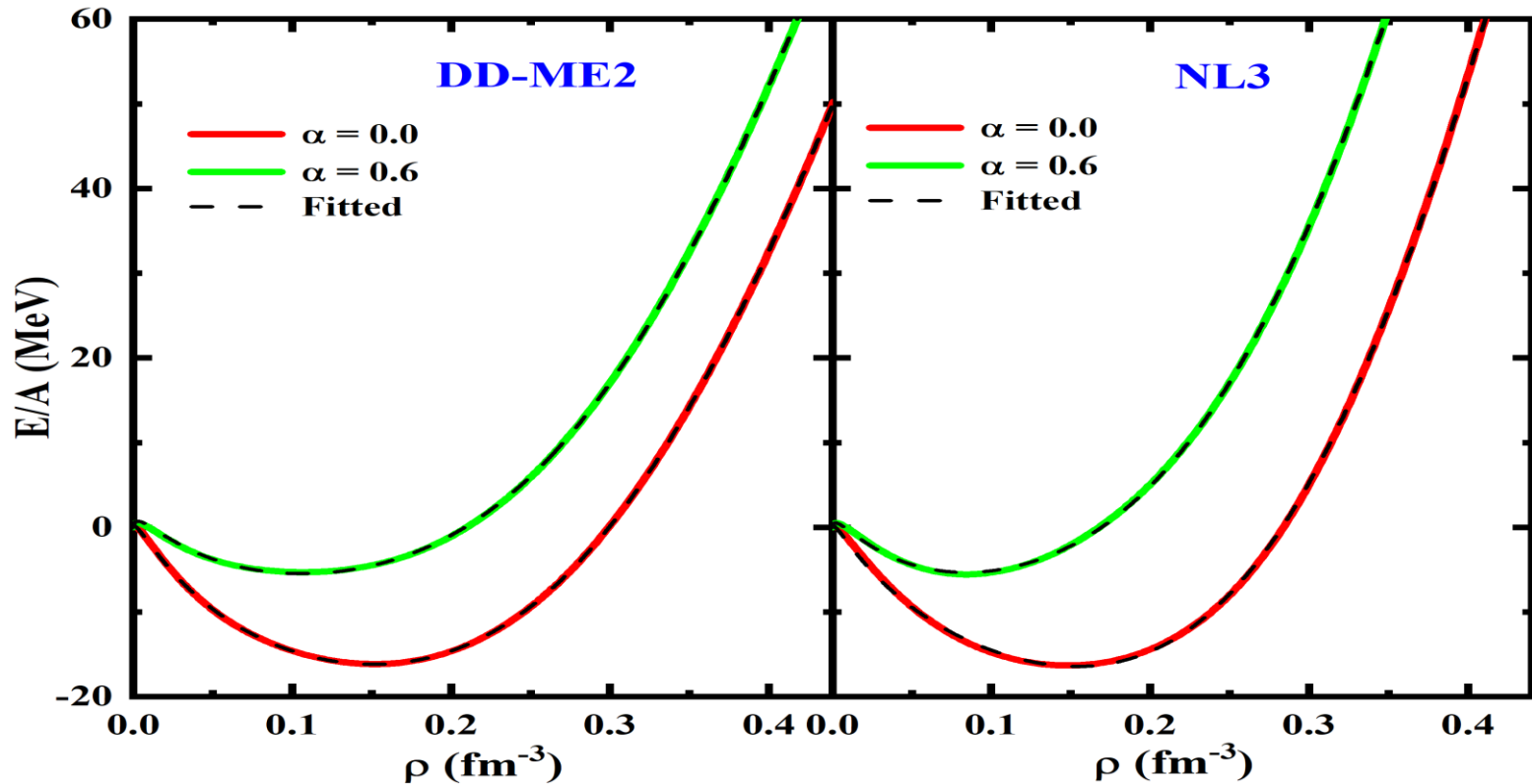
Variation of Symmetry Energy  $S$  as a function of neutron number  $N$

*Chin. Phys. C* **46**, 084101 (2022).



- The next step is to include relativistic prescription within the CDFM, i.e. make use of Relativistic energy density functional (R-EDF).
- Recently, we have introduced a parameterization of R-EDF in Ref: [*Europhys. Lett.*, **146**, 14001 (2024).]
- The newly fitted R-EDF with CDFM [*Europhys. Lett.*, **146**, 14001 (2024)] inherently counters the Coester-band problem.
- This novel parametrization based on R-EDF provides precise results in the study of nuclear structure at drip-line & possible correlations of isospin properties with the nuclear bulk properties.

# Parameterization of EDF



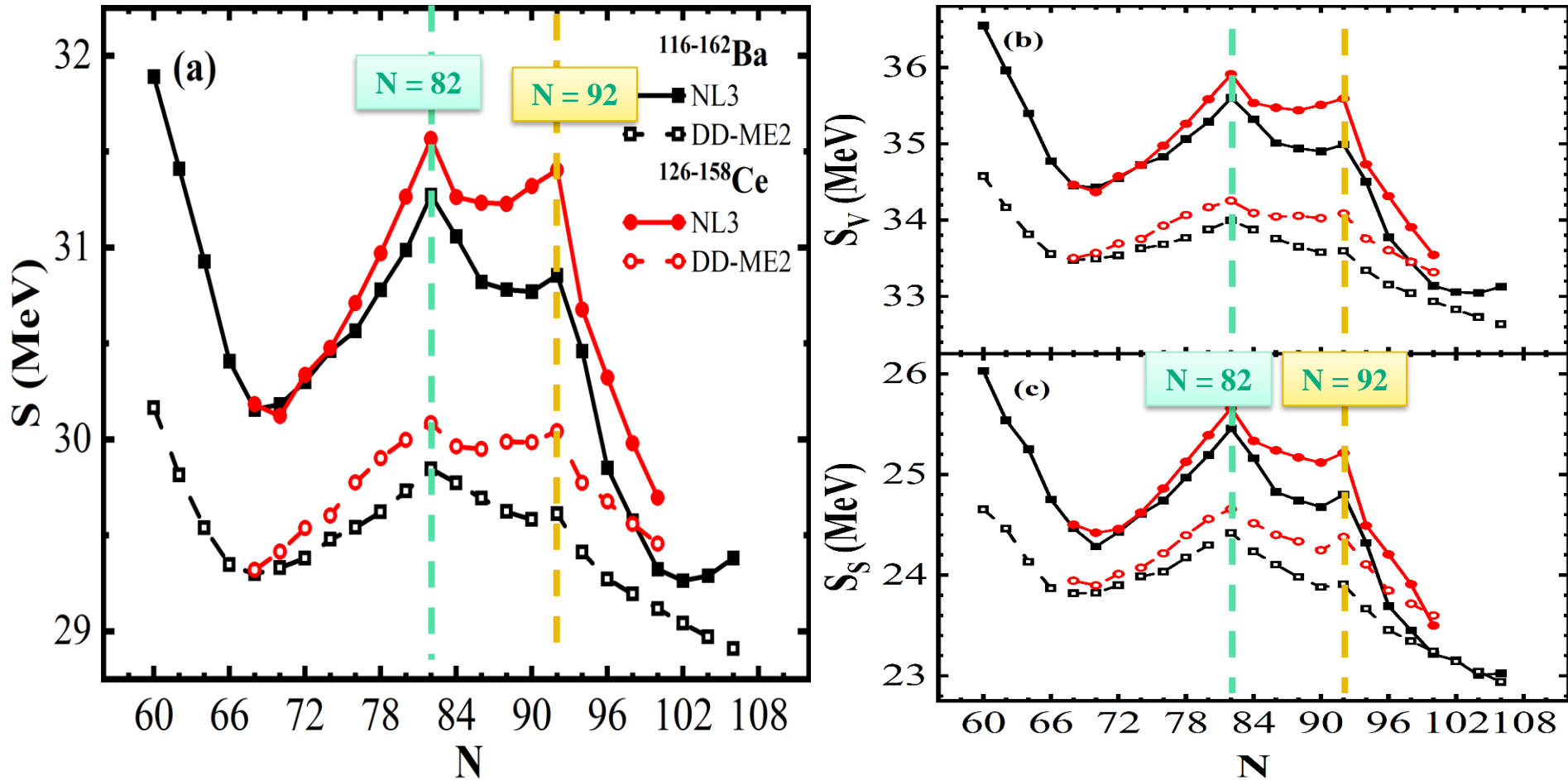
The energy per nucleon ( $E/A$ ) as a function of total number density ( $\rho$ ) of DD-ME2 and NL3 parameter sets for different asymmetry

- ✓ The parametrization resulted in the root-mean-square deviation of  $E/A$  for DD-ME2 and NL3 as 0.07704 and 0.19146 MeV, respectively.

*Europhys. Lett.*, **146**, 14001 (2024)

# Results using Relativistic-EDF within CDFM

## Verified existing & predicted novel shell closure



The calculated symmetry energy ( $S$ ), volume ( $S_V$ ) & surface symmetry ( $S_S$ ) energy using the parameterized Relativistic-EDF within the CDFM formalism.

# Summary / Take away

- CDFM formalism can be successfully used to study finite nuclei across nuclear chart.
- Rather than using Brueckner-EDF, with the incorporation of Relativistic prescription (Relativistic-EDF) we can avoid the Coester band problem.
- This method can successfully verify existing experimental shell closure & predict novel ones.

*Thank  
you*



# Extra Sildes

Following the literature, as detailed in Ref. [Z Phys A **297**, 257–260 (1980), JINR-E-2-11282, and Z Physik **304**, 239–243 (1982)] we can derive the expression of the weight function.

The mixed density, whose Fourier transformation is connected with the momentum distribution, can be written quite generally in the form:

$$\rho(\mathbf{r}, \mathbf{r}') = \int_0^\infty |\mathcal{F}(x)|^2 \rho_x(\mathbf{r}, \mathbf{r}') dx. \quad (1)$$

Here,  $|\mathcal{F}(x)|^2$  is the weight function for the different uniform distributions in the average density distribution. Moreover,  $\rho_x$  is the density matrix for  $A$  nucleons uniformly distributed in the sphere with radius  $x$  and spherical flucton density  $\rho_o(x) = 3A/4\pi x^3$ , expressed as:

$$\rho_x(\mathbf{r}, \mathbf{r}') = 3\rho_o(x) \frac{J_1(k_F(x)|\mathbf{r}-\mathbf{r}'|)}{(k_F(x)|\mathbf{r}-\mathbf{r}'|)}. \quad (2)$$

Here,  $J_1$  is a Bessel function of first-order, and  $k_F(x)$  is known as the Fermi momentum of the nucleons given as:

$$k_F(x) = \left( \frac{3\pi^2}{2} \rho_o(x) \right)^{1/3} = \left( \frac{9\pi A}{8} \right)^{1/3}. \quad (3)$$

It is important to note that Eq. 1 relates to a general CDFM assertion that the density distribution of nuclear matter fluctuates around average distribution while maintaining spherical symmetry and uniformity. Following the definition of density matrix, one-particle density  $\rho(\mathbf{r})$  is given by its diagonal elements as:

$$\rho(\mathbf{r}) = \rho(\mathbf{r}, \mathbf{r}') \Big|_{\mathbf{r}' \rightarrow \mathbf{r}}. \quad (4)$$

Using Eqs. 1 and 4 gives

$$\rho(\mathbf{r}) = \int_0^\infty |\mathcal{F}(x)|^2 \rho_x(\mathbf{r}) dx, \quad (5)$$

which is a decomposition of  $\rho$  in terms of  $\Theta$ -like densities as:

$$\rho_x(\mathbf{r}) = \rho_o(x) \theta(x - r), \quad \theta(y) = \begin{cases} 1, & y \geq 0 \\ 0, & y < 0 \end{cases}. \quad (6)$$

Consequently, after the differentiation of Eq. 5 results in:

$$|\mathcal{F}(x)|^2 = - \frac{1}{\rho_o(x)} \frac{d\rho(r)}{dr} \Big|_{r=x}, \quad (7)$$

# Why RMF?

- Main advantage: Takes account of spin-orbit interactions automatically.
- Coester band problem (Nuclear saturation & binding energy) is solved by relativistic models.
- Empirical saturation point of symmetric nuclear matter that is  $E/A \approx -16$  MeV at  $\rho \approx 0.2$  fm<sup>-3</sup> instead of  $\rho \approx 0.15$  fm<sup>-3</sup>, commonly known as the Coaster-Band problem.
- Provides reasonably good description of finite nuclei & infinite nuclear matter

Prog. Part. Nucl. Phys. **37**, 193 (1996), Phys. Rev. C **97**, 024322 (2018).