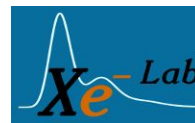


The XENON Group

XENONnT, future experiments (DARWIN/XLZD) and associated R&D (XeLab)

Luca Scotto Lavina – DR, LPNHE
on behalf of the XENON-LPNHE team



Current composition of the group

- 2 permanent researchers

- Luca SCOTTO LAVINA (DR)
- Bernard ANDRIEU (CR)

- 4 permanent ITs (1.1 FTE)

- Romain GAIOR (IR → *chercheur experimentaliste*)
- Nabil GARROUM (IR)
- Olivier DADOUN (IR)
- Yann Orain (AI)

- 1 Postdoc

- Frédéric GIRARD

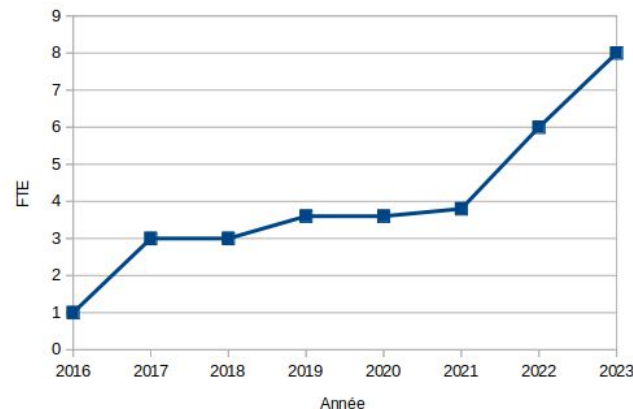
- 3 PhDs

- Layos DANIEL GARCIA (50 % LSL, 50 BA %), exp. defence 11/2024
- Quentin PELLEGRINI (50 % LSL, 50 BA %), exp. defence 09/2025
- Yongyu PAN (70 % LSL, Fei Gao Tsinghua University 30 %), exp. Defence 09/2026

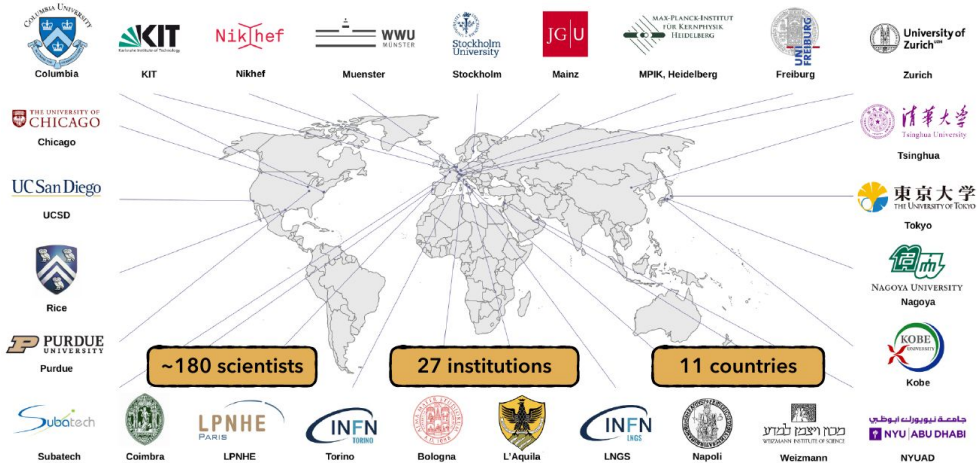
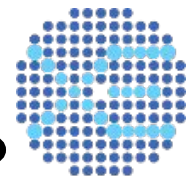
- Former members

Erwann MASSON (Postdoc, 2y, 2021-2023), Sid El Moctar AHMED MAOULOUD (PhD, 3y, 2019-2022), Olivier SPIGA (AI, 9m, 2021-2022), Jean-Philippe ZOPOUNIDIS (PhD, 3y, 2017-2020), Ernesto LOPEZ-FUNE (Postdoc, 2y, 2017-2019)

Group size evolution
since its creation in 2016



The XENON Collaboration



XENON Collaboration Meeting
@ LPNHE, Paris, Sept 2023



Timeline context for the HCERES

4th Nov 2023

HCERES-2023

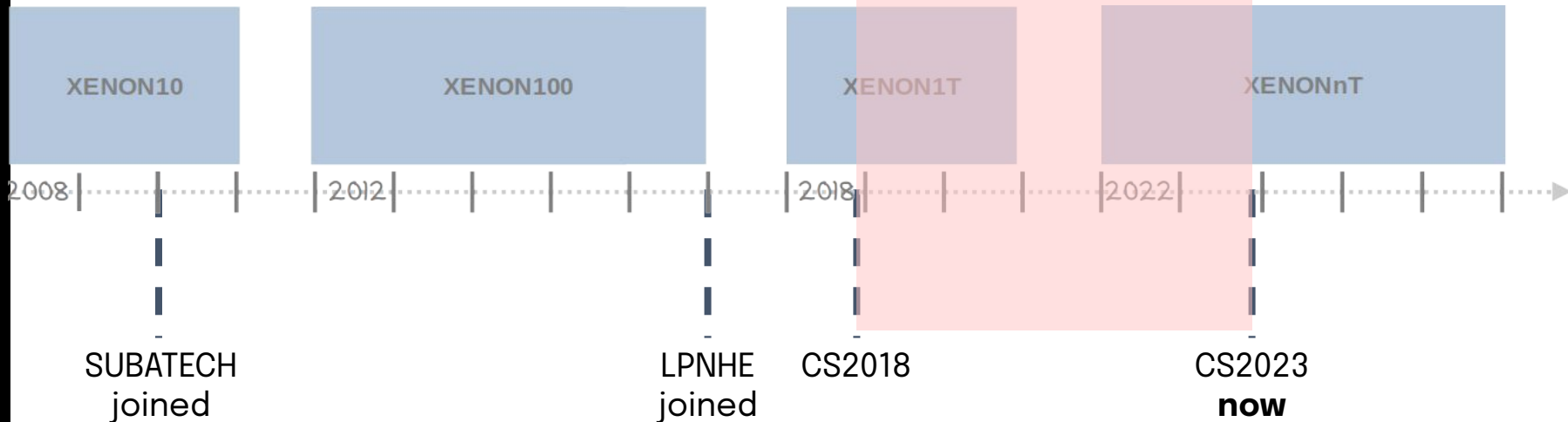
XENON

4



- End of XENON1T data taking and final analyses
- XENONnT construction and commissioning
- XENONnT data taking and analysis

SR0 SR1



Here we will just tell the latest

5 years

Blue bands = science data taking

Outline

Data analysis :

- **Completed XENON1T studies (leptophilic dark matter)**
- **Ongoing XENONnT studies (light dark matter, solar neutrinos, supernova neutrinos)**
- **Simulation studies on XENONnT proportional scintillation**

Technical contributions for XENONnT :

- **Installation and commissioning of the Xenon Storage and Recovery system (ReStoX2)**
- **Leading computing**
- **Improving Geant4 simulations on electrodes and light propagation**
- **Data Quality Monitoring**
- **User Management Tools**

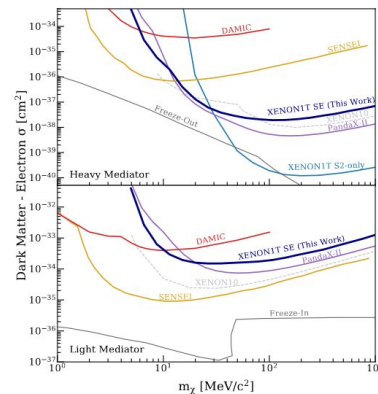
LPNHE XENON1T data analyses

Phys. Rev. D 106, 022001 (2022), [arXiv:2112.12116](https://arxiv.org/abs/2112.12116)

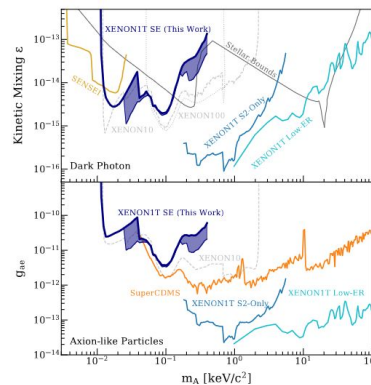
- **Different dark matter models that can be probed:**
 - Low-E Nuclear Recoils (NR)
 - SI elastic scattering
 - SD elastic scattering (LXe-specific)
 - WIMP-pion coupling
 - Effective Field Theory on WIMPs (+iDM) (LXe-specific)
 - Mirror DM
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 - **Dark Photons**
 - Bosonic SuperWIMPs, Magnetic dark matter
 - **Solar axions and Axion-like Particles**
 - Luminous DM
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 - Neutrino magnetic moment
 - Rare events
 - Double electron capture (LXe-specific)
 - Gravitational waves
 - Supernova neutrinos
 - New particles
 - **Solar Dark Photons**

DM-electron scattering

fermion or scalar boson DM candidate scatters off an electron bound in a xenon atom

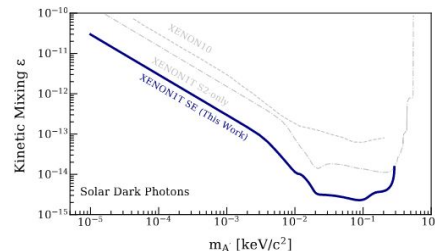


Vector-boson DM \rightarrow dark photons
Pseudo-scalar DM \rightarrow axion-like particles (ALPs)



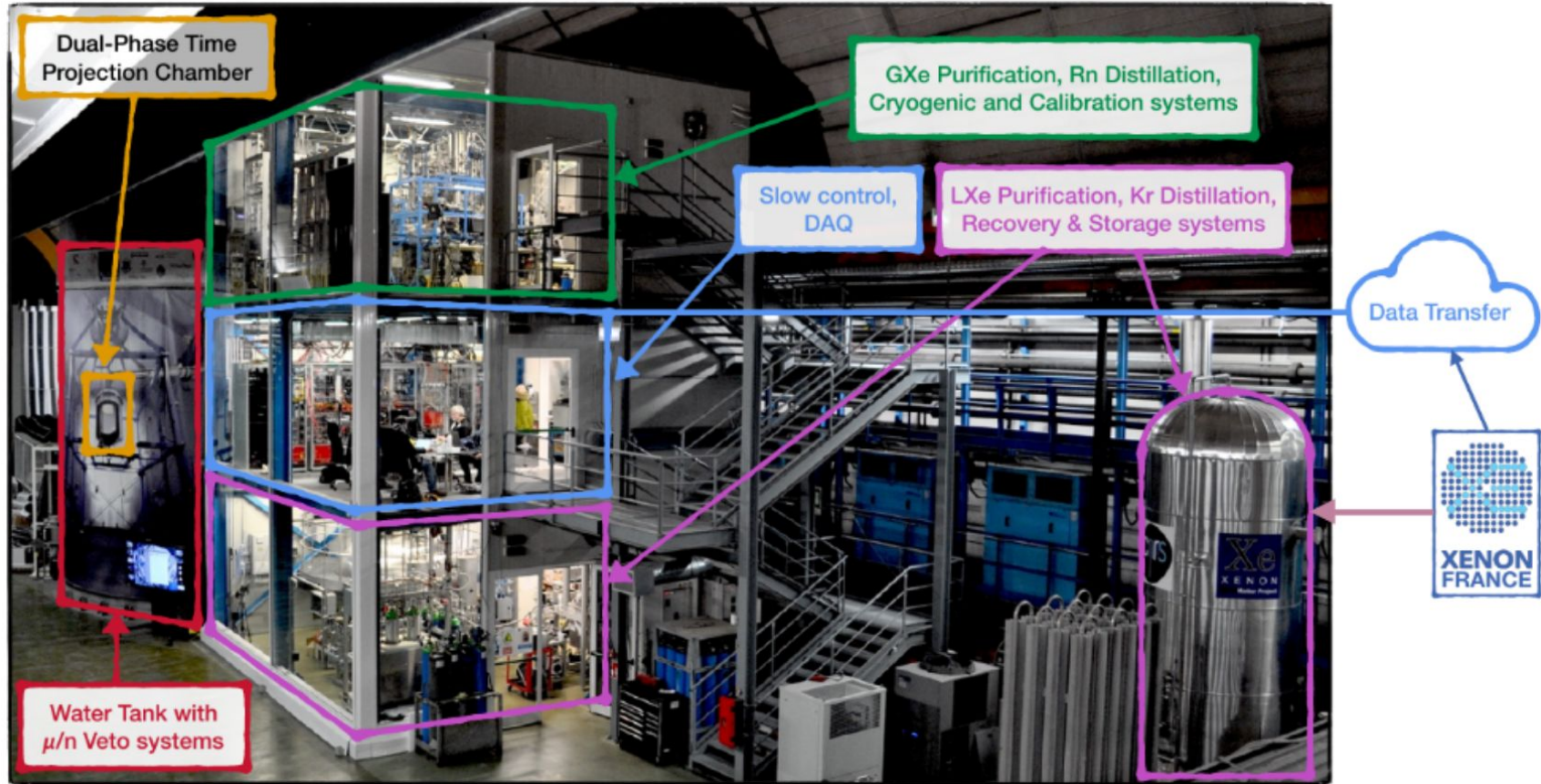
Solar Dark Photons

Higher kinetic energy wrt relic Dark Photons \rightarrow boosted the 2-5 electrons spectrum

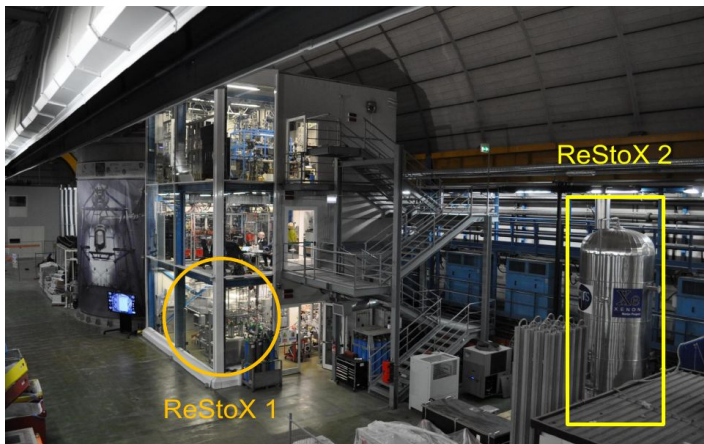


**LPNHE main contribution,
corresponding author J.P. Zopounidis (PhD thesis)**

From XENON1T to XENONnT, the advantages of quick updates



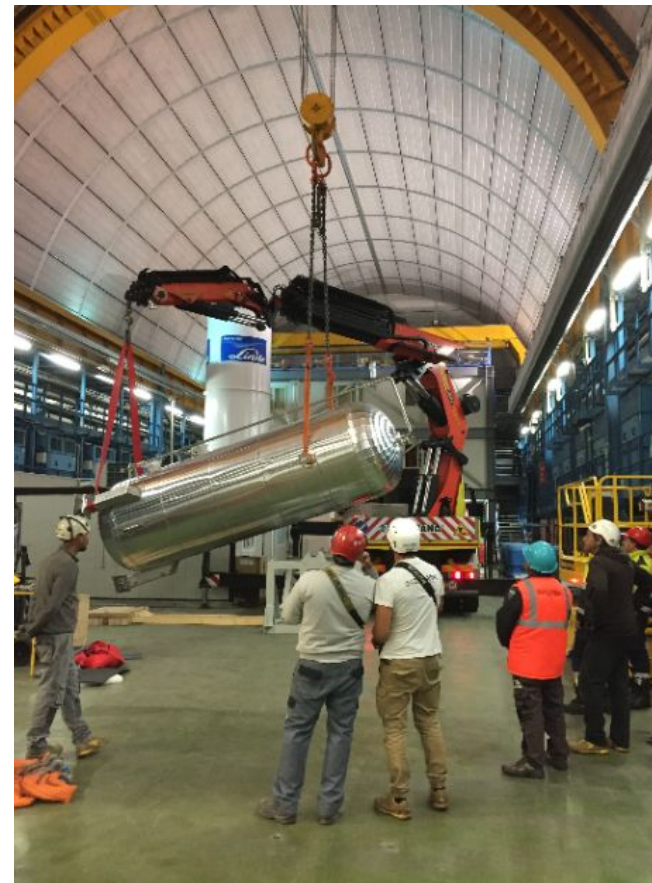
The Recovery and Storage System of XENONnT (ReStoX2)



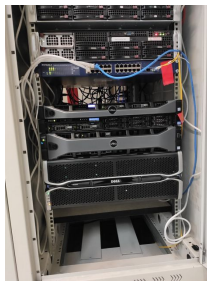
ReStoX1 : Columbia, Subatech and Mainz

ReStoX2 : 100% contribution of XENON-France (Subatech, LPNHE, LAL). **LPNHE contributed with the inner heat exchanger (DATE)**, SUBATECH with the vessel. Funded by IN2P3 and the two regions : *Pays de la Loire* and *Île-de-France* (DIM-ACAV+)

Reached a fast recovery with a rate of 1 tonne / hour !



LPNHE : Leading Computing for XENONnT



Design, installation and commissioning of the computing machines at LNGS

Development of the data transfer software (aDMIX) :

<https://github.com/XENONnT/admix>

Development of the Offline Data Quality Monitoring (XOM) :

<https://github.com/XENONnT/xom>

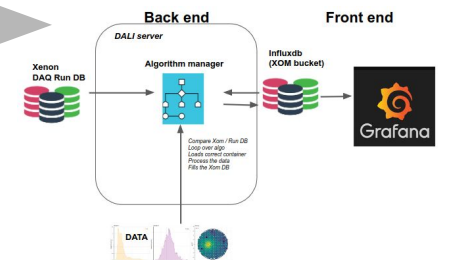
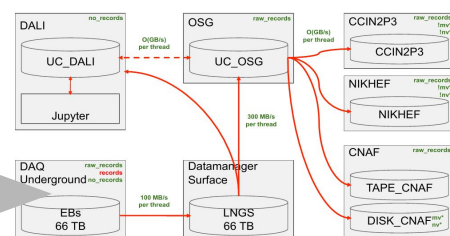
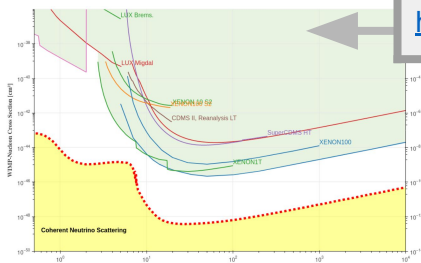
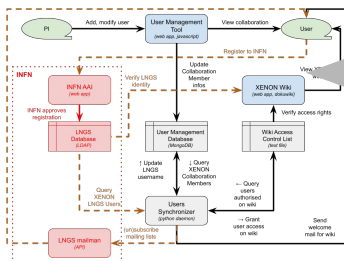
Development of the User Management System

GRID Administration

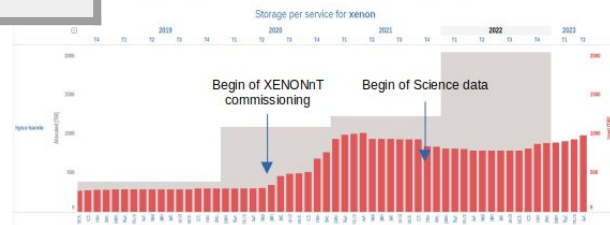
CC-IN2P3 resources (2.5PB of HPSS, 4M HS06.Hour)

For the GDR DUPhy : development of a Dark Matter plotter :

<https://github.com/odadoun/DarkPlotter>



Evolution over time using the CC-IN2P3 web app



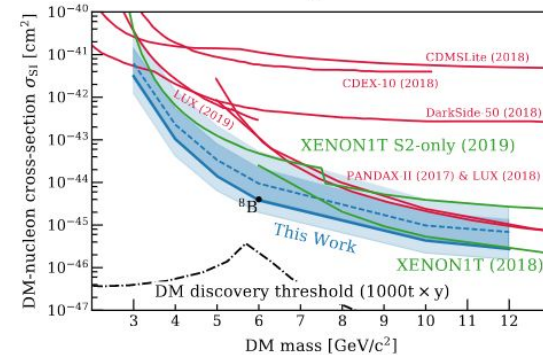
Ongoing LPNHE XENONnT data analyses

- **Different dark matter models that can be probed:**
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 - SI elastic scattering
 - SD elastic scattering (LXe-specific)
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 - Neutrino magnetic moment
 - Rare events
 - Double electron capture (LXe-specific)
 - Gravitational waves
 - Supernova neutrinos
 - New particles
 - Solar Dark Photons

Solar ^8B neutrinos

Goals:

- Develop new techniques to **improve sensitivity** near the threshold
- Quantify the ^8B **neutrinos** component in our background (in XENON1T: 6 events observed in the ROI, 5.38 background expected, whose 2.11 from CEvNS)
- Improve our DM limit at low masses



Challenges (aka, technical contributions):

- Improve S2 simulation
- Study of the biases of the reconstruction algorithms

Q. Pellegrini PhD thesis (ongoing)

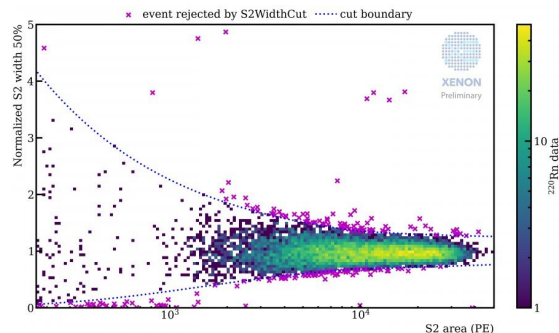
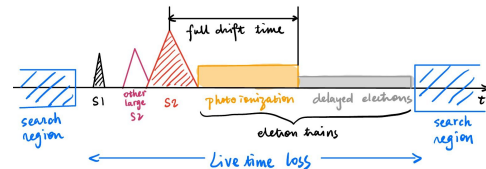
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 - Supernova neutrinos
 - New particles
 - Solar Dark Photons

Low mass Dark Matter

Goals:

- Study tiny signals induced by light dark matter
- Single electrons' contamination (develop rejection cuts and background model)
- Improve our DM limit at very low masses (few electrons signal)



Challenges (aka, technical contributions):

- Improve the knowledge on secondary S1s and S2s
- Develop new background rejection cuts

S. el M. Ahmed Maouloud PhD thesis (completed)
Y. Pan PhD thesis (ongoing)

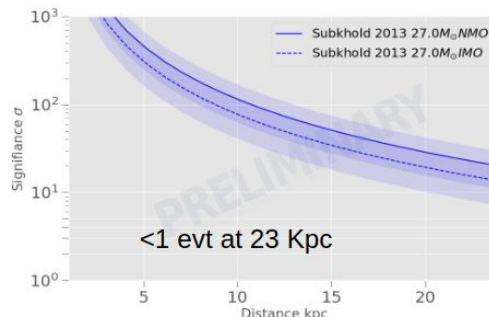
Ongoing LPNHE XENONnT data analyses

- **Different dark matter models that can be probed:**
 - Low-E Nuclear Recoils (NR)
 - SI elastic scattering
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 - WIMP-pion coupling
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 - Neutrinoless double-beta decay (LXe-specific)
 - Neutrino magnetic moment
 - Rare events
 - Double electron capture (LXe-specific)
 - Gravitational waves
 - **Supernova neutrinos**
 - New particles
 - Solar Dark Photons

Detection of supernova neutrinos

Goals:

- Make XENONnT capable to detect promptly neutrinos coming from a supernova via Coherent scattering in the TPC
- Include also the two (water tank based) veto systems signal via Inverse Beta Decay



Challenges (aka, technical contributions):

- Study all SN models
- Background study and sensitivity curves as a function of SN distance
- Developing a reliable MC simulation for vetoes
- Validate MC with real neutrons (AmBe)

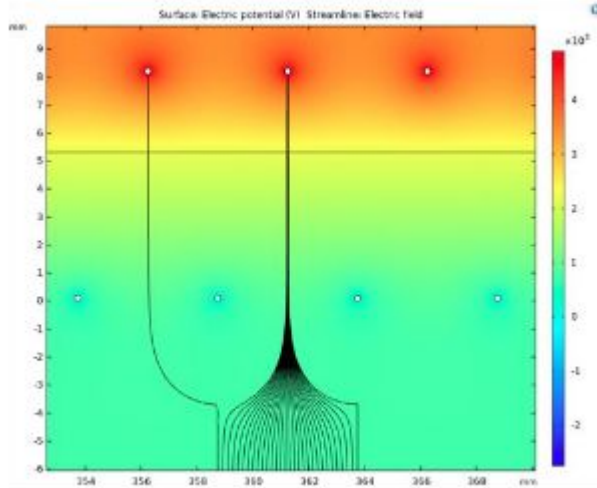
L. Daniel Garcia PhD thesis

Electroluminescence and new cathode in MC Geant4 simulation

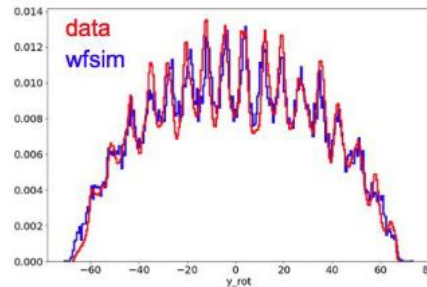
Goals:

- Improving MC simulation on secondary scintillation signal (high impact on any XENON analysis)

GARFIELD++/COMSOL simulation



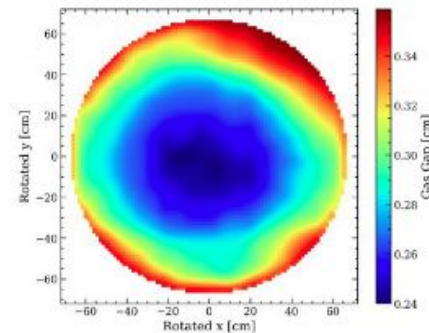
Simulation of shadowing + PMTs
pattern perfectly matches with data



Challenges (aka, technical contributions):

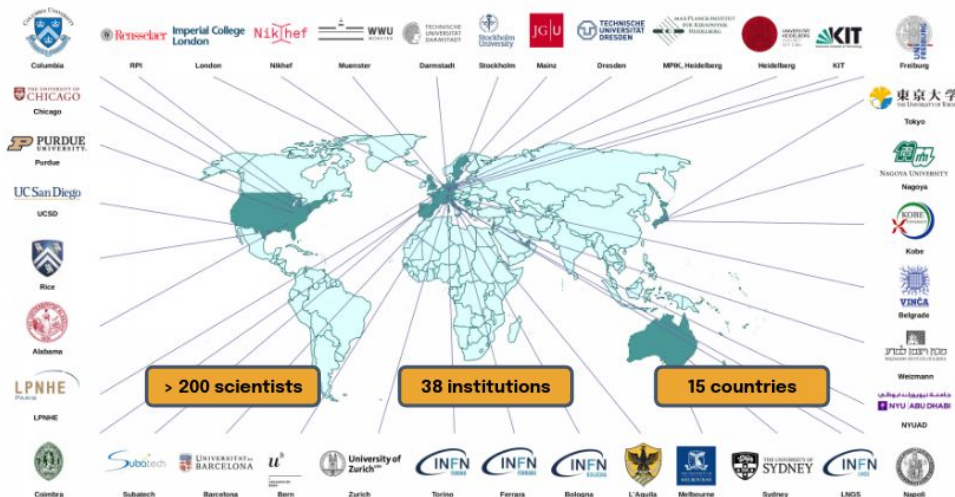
- Studying the shadowing effect
- Simulate the light propagation (each photon separately!)
- Improving the details on electrodes in Geant4 geometry

Electrodes sagging



B. Andrieu (CR), O. Dadoun (IR)

The future : DARWIN → XLZD



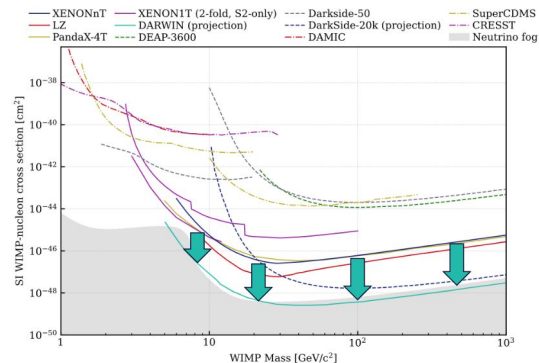
A Next-Generation Liquid Xenon Observatory for Dark Matter and Neutrino Physics

J. Aalbers,^{1,2} K. Abe,^{3,4} V. Aerne,⁵ F. Agostini,⁶ S. Ahmed Maouloud,⁷ D.S. Akerib,^{1,2} D.Yu. Akimov,⁸ J. Akshat,⁹
A.K. Al Musallhi,¹⁰ F. Alder,¹¹ S.K. Alsum,¹² L. Althueser,¹³ C.S. Amarasinghe,¹⁴ F.D. Amaro,¹⁵ A. Ames,^{1,2}
T.J. Anderson,^{1,2} B. Andrieu,⁷ N. Angelides,¹⁶ E. Angelino,¹⁷ J. Angevaere,¹⁸ V.C. Antochi,¹⁹ D. Antón Martín,²⁰
B. Antonenko,^{21,22} F. Anzile,²³ H.M. Araújo,¹⁶ I.F. Assaert,²⁴ F. Arceudo,²⁵ M. Aschauer,¹⁴ D. Auer,²⁶
S. Aubert,²⁷ P. Baai,²⁹ A. Balashov,³¹ M. Balci,³² A. Balci,³³ A. Balci,³⁴ A. Balci,³⁵ A. Balci,³⁶ A. Balci,³⁷ A. Balci,³⁸ A. Balci,³⁹ A. Balci,⁴⁰
W. Bargema,⁴¹ D. Baur,⁴² M. Bauer,⁴³ M. Bauer,⁴⁴ M. Bauer,⁴⁵ M. Bauer,⁴⁶ M. Bauer,⁴⁷ M. Bauer,⁴⁸ M. Bauer,⁴⁹ M. Bauer,⁵⁰
M. Bazky,³⁹ K. Beattie,⁴⁰ J. Behrens,⁴¹ N.F. Bell,³⁵ L. Bellagamba,⁶ P. Beltrame,⁴² M. Benabderrahmane,²⁵
E.P. Bernard,^{43,40} G.F. Bertone,¹⁸ P. Bhattacharjee,⁴⁴ A. Bhatti,²⁴ A. Biekert,^{43,40} T.P. Biesiadzinski,^{1,2}
A.R. Binay,⁹ R. Biordi,⁴⁵ Y. Biordi,⁴⁶ H.J. Blachnik,¹⁴ F. Bleckmann,¹⁴ F. Bleckmann,¹⁴ F. Bleckmann,¹⁴ F. Bleckmann,¹⁴

<https://arxiv.org/abs/2203.02309>

DARWIN

- 2.6 m diameter x 2.6 m height
- 40 t LXe active target
- Two arrays of photosensors (1910 3" PMTs)
- 24 PTFE reflector walls
- Passive and active muon and neutron vetos
- Located at LNGS



xlzd

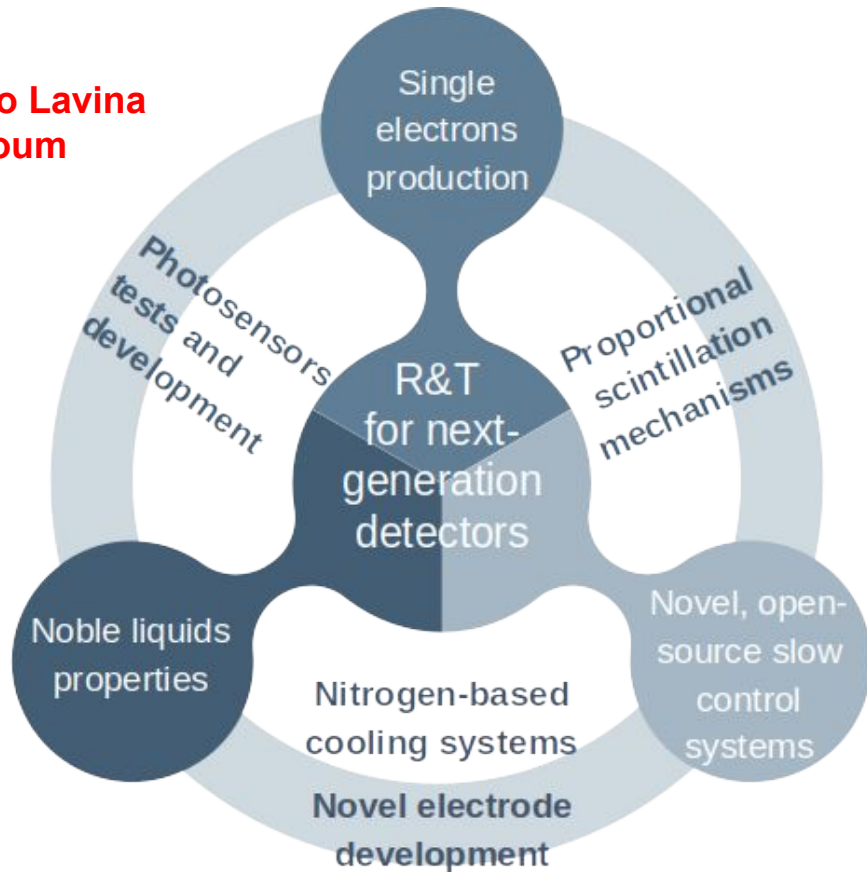
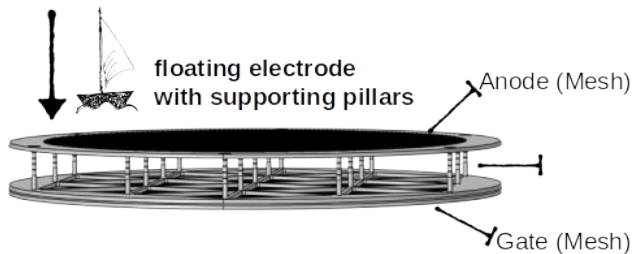
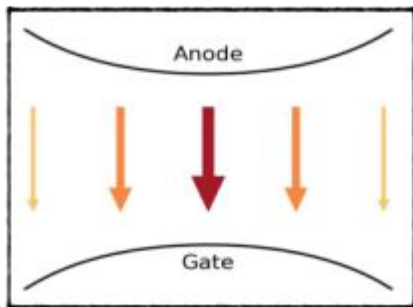
The IN2P3 XeLab R&T Project



IN2P3
Les deux infinis

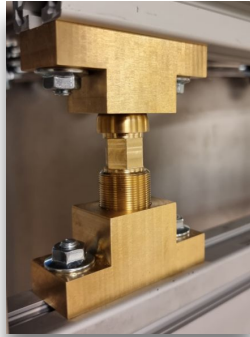


Scientific leader : L. Scotto Lavina
Technical leader : N. Garroum



The R&D in France: XeLab and its cryogenic system

See our
poster on
XeLab!



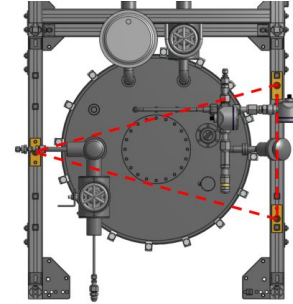
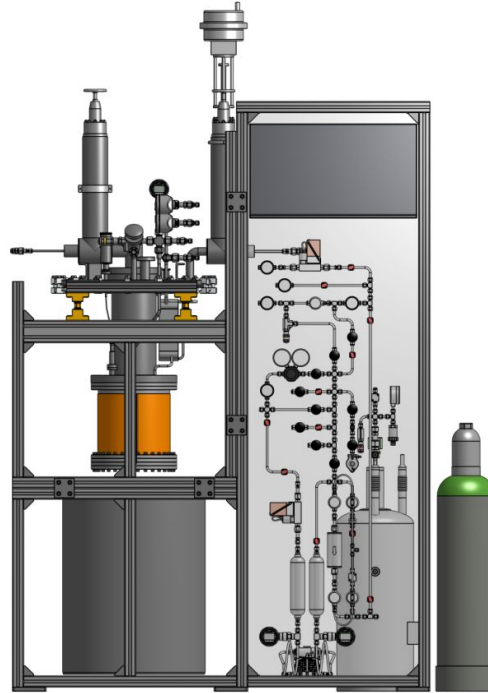
R&D
Levelling
system



R&D
Three-way
heat
exchanger

R&D
Cryostat
LN2-cooling
with copper
belt

Cryogenics and design : F. Girard
Slow Control : R. Gaior, Y. Pan
Technical support : Y. Orain



Three
levelling
systems to
define a plane



R&D
Storage and
recovery
system

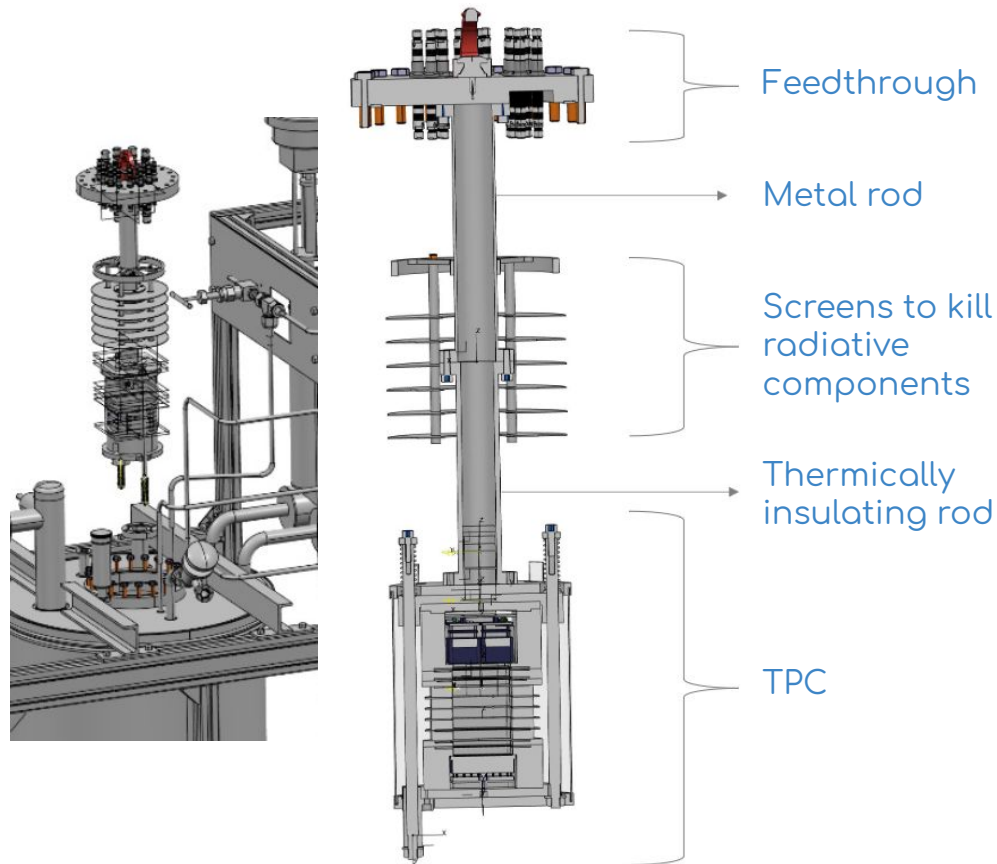
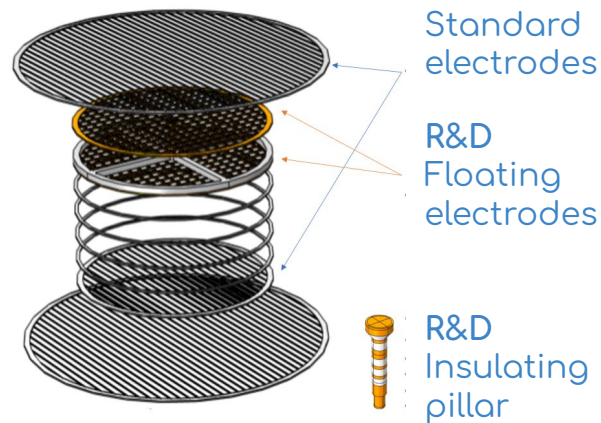
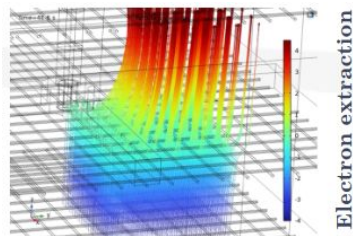
R&D
Slow Control
with RevPI



The R&D in France: XeLab and its TPC

See our
poster on
XeLab!

Studies and simulations : N. Garroum
MC simulation : E. Masson, O. Dadoun
Feedthrough : R. Gaïor



R&D in Collaboration with ...

Germany

International Research Laboratory (IRL) CNRS / Helmholtz Foundation (DMLab).

Three axes of common work for DARWIN:

- Liquid xenon technology
- Electrodes
- Computing



Australia

Submitted an International Research Project (IRP) proposal (LPNHE, Subatech, Sydney and Melbourne)

- DARWIN simulations
- Joint PhDs
- R&D with XeLab (students coming to Paris for short stay)



Conclusions and outlook

XENON1T:

- Physics production completed. LPNHE major contribution to **analysis** : leptophilic dark matter studies

XENONnT:

- LPNHE : strong hardware contribution (**Xenon Storage Systems** and **Computing**)
- LPNHE : **leadership in computing**
- LPNHE : Data analysis on **four topics** : low mass DM, solar neutrinos, SN neutrinos, S2 simulation

DARWIN/XLZD:

- Data taking not before 2032, multi-stage strategy before
- LPNHE : leadership in computing (>2023) and liquid xenon technology (<2023)

XeLab as R&T project towards next-generation detectors:

- **R&T** on: liquid xenon technology (LN2 cooling systems, storage and recovery), electrodes, slow control

Very ambitious program, but already successfully ongoing and compatible with the increasing size of the group (plus, aiming for a MdC Sorbonne experienced on hardware and DM data analysis)

Enlightening the Dark

Thanks for your attention

4th Nov 2023

HCERES-2023

XENON

20



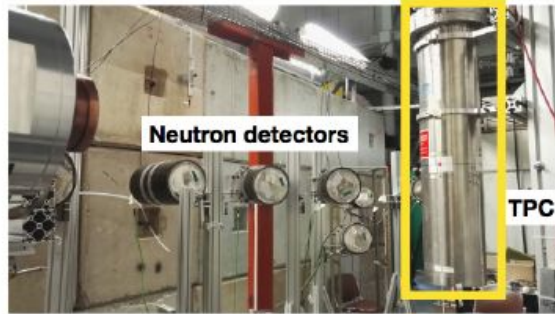
The DarkSide group

DarkSide-50 detector

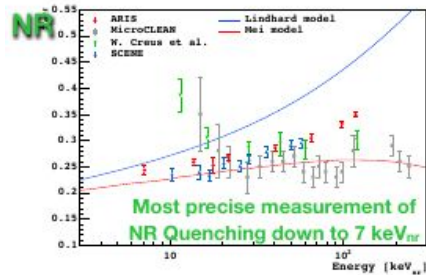
Claudio Giganti – CR, LPNHE

- The DarkSide activity in the lab **started in 2014**
- Originally formed by two researchers (Sandro De Cecco and Claudio Giganti)
- **Sandro is currently on leave** → Professor at University of Rome where he is working full time on DarkSide
- **Claudio is mostly working on the neutrino physics** → T2K and Hyper-K
- Two PhD students:
 - Anyssa Navrer-Agasson (2016-2019) → defence 09/2019
 - ARIS experiment
 - DarkSide-50 low mass analysis
 - Julie Rode (co-direction with Davide Franco, APC) → defence 09/2022
- **Participation to the DS-50 analysis** (calibration, background model, limits for axions and sterile neutrinos)
- Development of DS-20k reconstruction tools

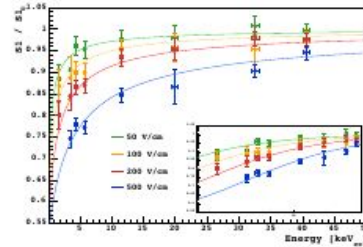
The ARIS experiment



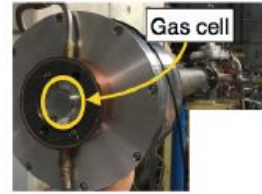
12 days of data taking
at ALTO@IPNO in 2016



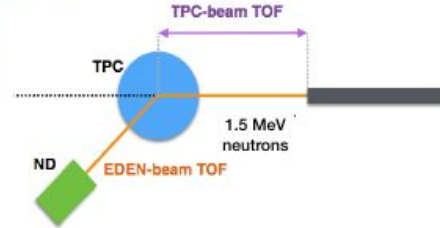
No field



Phys. Rev. D 97 (2018) no.11, 112005



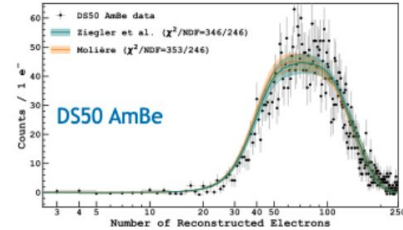
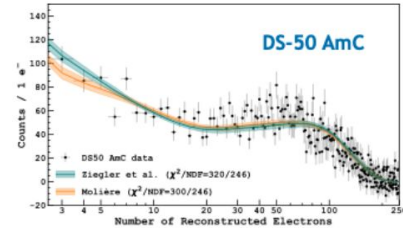
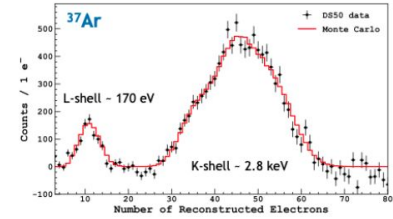
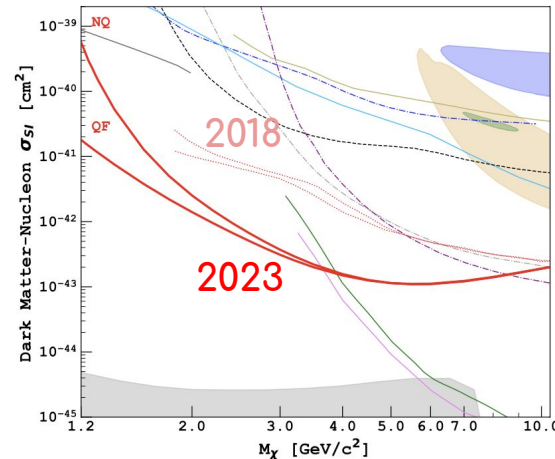
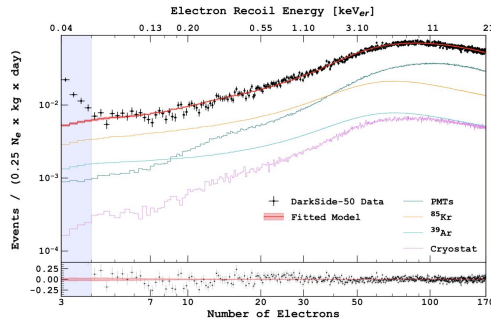
Phys. Rev. D 97 (2018) no.11, 112005



- *Experiment proposed by DS French groups
- *France/Italy/US collaboration ~ 30 people
- *Measure scintillation yield for ER and NR
- *Measure quenching (Leff) in the [7,120] keV_{nr} range
- *Full description of recombination processes for ER and NR

Low mass WIMP searches (2018) and reanalysis (2023)

- End of DS-50 data taking → work on a re-analysis of S2 only data
- Goal : understand the observed excess between 4 and 7 electrons
 - New calibration, new selection, new background model
 - Improve by one order of magnitude previous limits with (almost) same data-set



Conclusions

- The LPNHE DarkSide group joined the experiment in 2014 and **contributed to the low-mass searches of DS-50**
- The collaboration is now building the next generation, DarkSide-20k, detector
- Unfortunately the small size of the DS group at LPNHE and other commitments of the only permanent researcher involved on DS makes it very difficult to reasonably contribute to this experiment
- **We decided not to take any commitment in the construction of DS-20k**

Material for discussions

- **Different dark matter models that can be probed:**

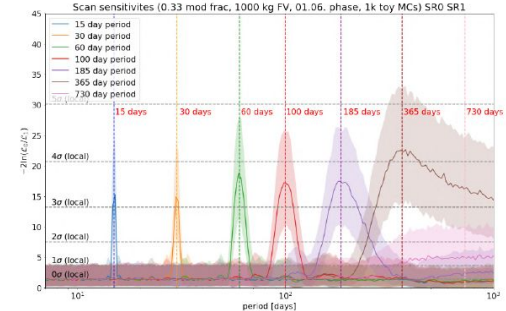
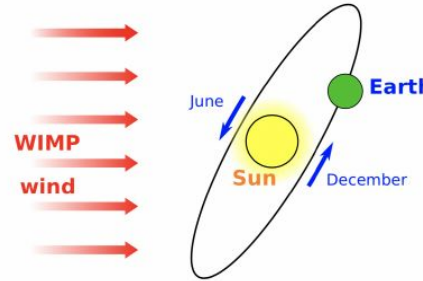
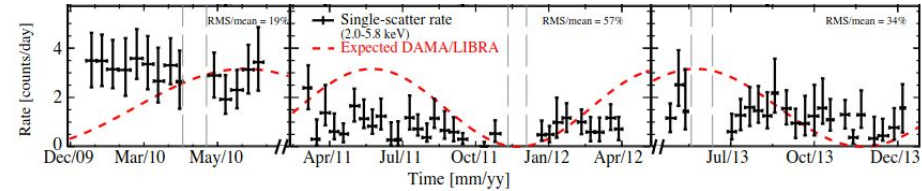
- Low-E Nuclear Recoils (NR)
 - SI elastic scattering
 - SD elastic scattering (LXe-specific)
 - WIMP-pion coupling
 - Effective Field Theory on WIMPs (+iDM) (LXe-specific)
 - Mirror DM
- Electronic Recoils (ER)
 - Dark Photons
 - Bosonic SuperWIMPs, Magnetic dark matter
 - Solar axions and Axion-like Particles
 - Luminous DM
- Both (NR+ER)
 - Inelastic DM
 - **Annual modulation search**
 - Low mass WIMPs ($<10\text{GeV}$)
 - Multiply-Interacting Massive Particles (MIMPs)
 - Migdal Effect and Bremsstrahlung

- **New physics can be scoped:**

- Neutrinos
 - Solar 8B neutrinos ($\text{CEvNS} \rightarrow \text{NR}$)
 - Neutrinoless double-beta decay (LXe-specific)
 - Neutrino magnetic moment
- Rare events
 - Double electron capture (LXe-specific)
 - Gravitational waves
 - Supernova neutrinos
- New particles
 - Solar Dark Photons

Annual modulation

Two publications with XENON100 with 4 years of data:



**XENON1T study done at LPNHE,
 J.P. Zopounidis (PhD thesis)**

The scoped mass domains of Dark Matter in a visual way

