

# Exploring the Isomeric Decay of <sup>180m</sup>Ta with the MAJORANA DEMONSTRATOR:

## New Insights from the Second Year of Data

Ralph Massarczyk (LANL)

LA-UR-24-31756





#### A bit of (ancient) history...



In greek mythology **Tantalus** offended the gods...

... so he was punished to be **trapped** in a pond under a fruit tree.

He could **not** reach **up** to eat.

He could **not** lean **down** to drink.

Tantalus trapped as punishment.

Illustration from www.symmetrymagazine.org/article/majorana-demonstrator-finds-tantalizing-new-purpose

#### A bit of (modern) history...



Level scheme of <sup>180m</sup>Ta

Ta used in the experiment

For nuclear physics **Tantalum** (named 1802) is one of the rarest elements and has two isotopes...

... one of them (<sup>180m</sup>Ta) is **trapped** in an isomeric state while the ground state decays.

It can **not** go to a **higher** state due to energy.

It can **not** go down to a **lower** state due to spins

- The origin of Tantalum in the universe :
  - Study helps to understand the observed abundance of <sup>180m</sup>Ta within a wider nucleosynthesis framework
  - Understand which candidate processes are strong enough to produce Ta (*v*-interactions, thermal excitation in early universe)
- Longest lived metastable state never observed to decay
  - Most extreme case to study nuclear structure spin traps
  - Theory varies on predictions for half-life
  - Variety of transitions possible:
    β-decay, electron capture (EC), internal conversion,
    γ-transition, α-decay
  - Ground-state <sup>180</sup>Ta is unstable ( $T_{1/2} \sim 8$  hours)



Decay scheme of <sup>180m</sup>Ta with possible decay channels (red) and detection signatures (blue)

4

- The origin of Tantalum in the universe :
  - Study helps to understand the observed abundance of <sup>180m</sup>Ta within a wider nucleosynthesis framework
  - Understand which candidate processes are strong enough to produce Ta (*v*-interactions, thermal excitation in early universe)



- <u>The origin of Tantalum in the universe</u> :
  - Study helps to understand the observed abundance of <sup>180m</sup>Ta within a wider nucleosynthesis framework
  - Understand which candidate processes are strong enough to produce Ta (*v*-interactions, thermal excitation in early universe)
- Longest lived metastable state never observed to decay
  - Most extreme case to study nuclear structure spin traps
  - Theory varies on predictions for half-life
  - Variety of transitions possible:
    β-decay, electron capture (EC), internal conversion,
    γ-transition, α-decay
  - Ground-state <sup>180</sup>Ta is unstable ( $T_{1/2} \sim 8$  hours)



Decay scheme of <sup>180m</sup>Ta with possible decay channels (red) and detection signatures (blue)

- <u>The origin of Tantalum in the universe</u> :
  - Study helps to understand the observed abundance of <sup>180m</sup>Ta within a wider nucleosynthesis framework
  - Understand which candidate processes are strong enough to produce Ta (*v*-interactions, thermal excitation in early universe)
- Longest lived metastable state never observed to decay
  - Most extreme case to study nuclear structure spin traps
  - Theory varies on predictions for half-life
  - Variety of transitions possible:
    β-decay, electron capture (EC), internal conversion,
    γ-transition, α-decay
  - Ground-state <sup>180</sup>Ta is unstable ( $T_{1/2} \sim 8$  hours)



Decay scheme of <sup>180m</sup>Ta with possible decay channels (red) and detection signatures (blue)

- The origin of Tantalum in the universe :
  - Study helps to understand the observed abundance of <sup>180m</sup>Ta within a wider nucleosynthesis framework
  - Understand which candidate processes are strong enough to produce Ta (*v*-interactions, thermal excitation in early universe)
- Longest lived metastable state never observed to decay
  - Most extreme case to study nuclear structure spin traps
  - Theory varies on predictions for half-life
  - Variety of transitions possible:  $\beta$ -decay, electron capture (EC), internal conversion, y-transition,  $\alpha$ -decay
  - Ground-state <sup>180</sup>Ta is unstable ( $T_{1/2} \sim 8$  hours)
- <u>Search for Dark Matter interaction</u>
  - Additional energy from the isomer allows reaction with particles that would not interact otherwise
  - Candidates: Strongly Interacting DM, Inelastic DM



Decay scheme of <sup>180m</sup>Ta with possible decay channels (red) and detection signatures (blue)



#### What is needed for a measurement...





History of Tantalum decay measurements with predictions (dashed lines), from arxiv 2305.17238

- Large exposure (material and time)
- Detector with excellent energy resolution
- If possible multiple detectors, that can detect coincidences
- A clean, ultra low-background system and environment

## Perfect use of MJD facility after enriched detector removal

#### Tantalum – a rare element (isotope)

- 1 2 ppm of earth's crust is Ta
- 99.98% is <sup>181</sup>Ta
- Best previous measurement used ~1kg of <sup>nat</sup>Ta (~0.2 g of <sup>180m</sup>Ta)
- All the <sup>180</sup>Ta is metastable:

the only naturally occurring long-lived isomer





#### MAJORANA DEMONSTRATOR 🚳





Searching for neutrinoless double-beta decay of <sup>76</sup>Ge in HPGe detectors, probing additional physics beyond the standard model, and informing the design of the next-generation LEGEND experiment

Source & Detector:

U.S. DEPARTMENT OF

- Array of p-type, point contact detectors 30 kg of 88% enriched <sup>76</sup>Ge crystals
- Included 6.7 kg of <sup>76</sup>Ge inverted coaxial, point contact detectors in final run
- Enriched detectors removed in 2021 for LEGEND
- 14 kg of natural Ge crystals

Office of

Science

- Excellent Energy Resolution: 2.5 keV FWHM @ 2039 keV
- Low Analysis Threshold: 1 keV
- Low Background: 2 modules within a compact graded shield and active muon veto using ultra-clean materials
- Final Result, (PRL 130, 062501, 2023)
  - 65 kg-yr exposure
  - Median  $T_{1/2}$  Sensitivity: 8.1 × 10<sup>25</sup> yr (90% C.I.)
  - Limit: T<sub>1/2</sub> > 8.3 × 10<sup>25</sup> yr (90% C.I)



#### **Reconfiguring of the DEMONSTRATOR**

• Detectors and Ta arranged to maximize efficiency







#### **Reconfiguring of the DEMONSTRATOR**

- Detectors and Ta arranged to maximize efficiency
- **17.4 kg installed** ~ 2 g <sup>180m</sup>Ta, (*x10 more than best previous measurement*)
- 23 active detectors (before only one or two HpGe setups)
- Operating since May 2022





(left) cleaning and installation in the MJD strings

(right) schematic arrangement of detectors, green, and Ta, grey, and photograph of the full detector array





#### Installation











#### **Data Overview and Analysis**

- Data Set of 859 days (84.0% live)
- Background contributions from:
  - natural radioactivity within the Tantalum disks (< 0.5 mBq/kg<sub>Ta</sub>)
  - surface activation in Ta
    - <sup>182</sup>Ta (T<sub>1/2</sub> = 114 days)
    - <sup>175</sup>Hf (T<sub>1/2</sub> = 70 days)
- Background improving over time



#### **Data Overview and Analysis**

- Data Set of 859 days (84.0% live)
- Background contributions from:
  - natural radioactivity within the Tantalum disks (< 0.5 mBq/kg<sub>Ta</sub>)
  - surface activation in Ta
    - <sup>182</sup>Ta (T<sub>1/2</sub> = 114 days)
    - $^{175}$ Hf (T<sub>1/2</sub> = 70 days)
- Background improving over time

n: 101 /2-) 5.2 m	<sup>176</sup> W	<sup>177</sup> W	<sup>178</sup> W	<sup>179</sup> W	<sup>180</sup> W	<sup>181</sup> W	<sup>182</sup> W	<sup>183</sup> W	<sup>184</sup> W
	<sup>175</sup> Ta z: 73 n: 102 Jπ 7/2+ T <sub>1/2</sub> :10.5 h	<sup>176</sup> Ta	<sup>177</sup> Ta	<sup>178</sup> Ta	<sup>179</sup> Ta	<sup>180</sup> Ta	<sup>181</sup> Ta	<sup>182</sup> Ta	<sup>183</sup> Ta
i.	<sup>192</sup> Hf	<sup>175</sup> Hf z: 72 n: 103 Jπ 5/2(-) T <sub>1/2</sub> :70 d 2	<sup>176</sup> Hf	<sup>177</sup> Hf	<sup>178</sup> Hf	<sup>179</sup> Hf	<sup>180</sup> Hf	<sup>181</sup> Hf	<sup>182</sup> Hf
J	<sup>173</sup> Lu	<sup>174</sup> Lu	<sup>175</sup> Hf Hafnium Jn T <sub>1/2</sub> or <u>F</u> Delta (keV) Bind/A (keV)	n 103 z 72 5/2(-) 70 d 2 -54481.763 23 8060 7625 13	<sup>77</sup> Lu	<sup>178</sup> Lu	<sup>179</sup> Lu	<sup>180</sup> Lu	<sup>181</sup> LU
) I	<sup>172</sup> Yb	<sup>173</sup> Yb	Mass (μAMU) Qα (keV) Qβ (keV) Qec (keV)	174941511.424 2 2400.1392 23 -2073.1103 28 683.915 26	25 <sup>76</sup> Yb	<sup>177</sup> Yb	<sup>178</sup> Yb	<sup>179</sup> Yb	I80Yb
n	<sup>171</sup> Tm	<sup>172</sup> Tm	Sn (keV) Sp (keV) Decay Major radiatio Type keV β+	6200.4421 22 ec 100% ns %	<sup>75</sup> Tm	<sup>176</sup> Tm	<sup>177</sup> Tm	<sup>178</sup> Tm	
2	<sup>170</sup> Er	<sup>171</sup> Er	γ 343.40 s 54.070 s	84 46.5	<sup>74</sup> Er	<sup>175</sup> Er	<sup>176</sup> Er		



#### A look in a few region of interests – full data



♦ U/Th decay chain

#### **Multiplicity analysis**



- Replaces search of individual signatures
- Efficiency lower
- Background much lower !

#### **Updated results**

- Current improvements
  - Efficiency (x 2-3) Ο
  - Mass (x 12) 0
  - Background 0
- multiplicity analysis allows high sensitivity search

$$\lambda_{total} = \lambda_{EC} + \lambda_{\beta^-} + \lambda_{\gamma} + \lambda_{IC} + \lambda_{\alpha} + \lambda_{DM}$$

	EC	β <sup>-</sup>	¥	IC	α
Previous Limits	> 1.6 x 10 <sup>18</sup>	> 1.1 x 10 <sup>18</sup>	> 4.5 x 10 <sup>14</sup>	> 4.5 x 10 <sup>14</sup>	-
MJD - 2023	> 1.3 x 10 <sup>19 **</sup>	> 1.5 x 10 <sup>19 **</sup>	> 6.0 x 10 <sup>17</sup>	> 2.9 x 10 <sup>17</sup>	> 1.1 x 10 <sup>19 **</sup>
Full data (preliminary)	3.7 x 10 <sup>19</sup>	4.3 x 10 <sup>19</sup>	1.3 x 10 <sup>18</sup>	5.9 x 10 <sup>17</sup>	3.0 x 10 <sup>19</sup>
Theory	10 <sup>23</sup>	10 <sup>20</sup>	10 <sup>31</sup>	10 <sup>18</sup>	10 <sup>25</sup>

Previous limits:

MAJORANA

#### **Recent updates**

- Before 1D spectral analysis
- Inclusion of correlated peak fits



#### **Recent updates**

- Before 1D spectral analysis
- Inclusion of correlated peak fits
- Investigation in <sup>175</sup>Hf and other short lived
- Optimize sensitivity by using 2D fits in energy and time
  - Background decaying component
  - Background constant component
  - Signal constant in time for Ta, or decaying for <sup>175</sup>Hf



2D fit of constant signal on background in presence of other peaks

#### **Recent updates**

- Before 1D spectral analysis
- Inclusion of correlated peak fits
- Investigation in <sup>175</sup>Hf and other short lived
- Optimize sensitivity by using 2D fits in energy and time
  - Background decaying component
  - Background constant component
  - Signal constant in time for Ta, or decaying for <sup>175</sup>Hf
- 20-30 % improvement over 1D fits



#### Makes use of all data

- Early data good background determination and statistics
- Later data: better signal to background ratio

#### **Dark matter induced deexcitation**

- No observation of <sup>180m</sup>Ta decay  $\rightarrow$  no DM-induced decay
- Improved sensitivities to strongly interacting DM (siDM)
- Additional sensitivities to more complex DM with multiple states
- and/or particles via inelastic scattering





#### Similar studies in other isomers

- Different approach
  - Short half life
  - High activity using a LANL in-house source
- Portable Ge detector, short measurement time (15m)
- Search for new transitions next to known transitions



#### Similar studies in other isomers

- Different approach
  - Short half life
  - High activity using a LANL in-house source
- Portable Ge detector, short measurement time (15m)
- Search for new transitions next to known transitions

#### $\chi + {}^{178m}\text{Hf} \rightarrow \chi^* + {}^{178}\text{Hf}_j$

PHYSICAL REVIEW LETTERS 131, 141801 (2023)

Editors' Suggestion

#### Dark Matter Constraints from Isomeric <sup>178m</sup>Hf

D. S. M. Alves<sup>1</sup>, S. R. Elliotte<sup>1,\*</sup> R. Massarczyk<sup>0,1</sup> S. J. Meijer<sup>0,1</sup> and H. Ramani<sup>2</sup> <sup>1</sup>Los Alamos National Laboratory, Los Alamos, New Mexico 87545, USA <sup>2</sup>Stanford Institute for Theoretical Physics, Stanford University, Stanford, California 94305, USA

(Received 7 June 2023; revised 10 August 2023; accepted 11 September 2023; published 5 October 2023)

We describe a first measurement of the radiation from a  $^{178m}$ Hf sample to search for dark matter. The  $\gamma$  flux from this sample, possessed by Lee Alemoe National Laboratory puckas chamietry, use measured





## Summary

- Most sensitive search for half-life measurements in isomers world-wide
- Data improved previous measurements by 1-2 orders of magnitude
- Background continues to improve
- Expect final results at the end of project Spring 2025













#### Inelastic dark matter motivation

- What if the DM species evade direct detection ?
- Production in colliders possible, but very low
- No reason that there is no a whole DM zoo of particles or composite DM
- Based on Tucker-Smith & Weiner (2001) or Alves et al (2010)
- Requires additional "energy" in the system
  - Sufficiently fast DM interacts with standard detectors measuring nuclear recoils OR



Recoil energy as a function of incoming DM velocity for non-isomers

