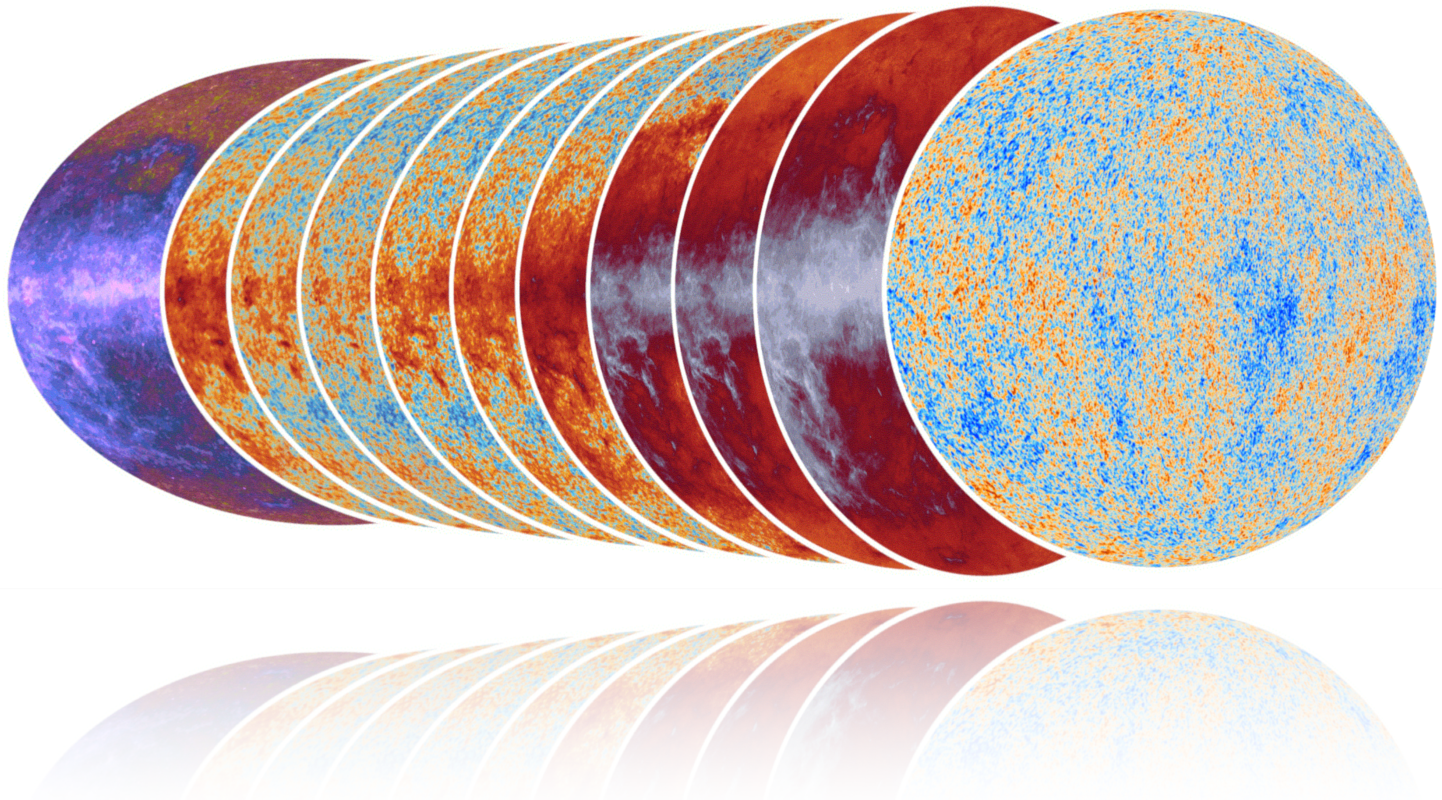

Cosmological parameters derived from the final Planck (PR4) data release

M. Tristram

CMB-France #5, Dec 2023

PLANCK NPIPE Team

Cosmological parameters derived from the final (PR4) *Planck* data release

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[Tristram et al. A&A (2023)]
astro-ph/2309.10034

PLANCK PR4 likelihoods

[Tristram+ (2023)]

Planck likelihoods are splits in **two parts** due to different **statistical assomptions**

large scales (low ℓ)

- **lowT: Commander**
  [Planck Collaboration V 2020]

Bayesian posterior Gibbs sampling that combines astrophysical component separation and likelihood estimation

- **lowE(B): Lollipop**
  [Tristram et al. 2022]

H&L likelihood based on cross-spectra between CMB clean maps on 50% of the sky

ℓ = 2-30

small scales (high ℓ)

- **Hillipop: TT, TE, EE, TTTEEE**
  [Tristram et al. 2023]

Gaussian likelihood based on cross-spectra from frequency maps on 75% of the sky, including models for the foreground residuals

ℓ = 30-2500

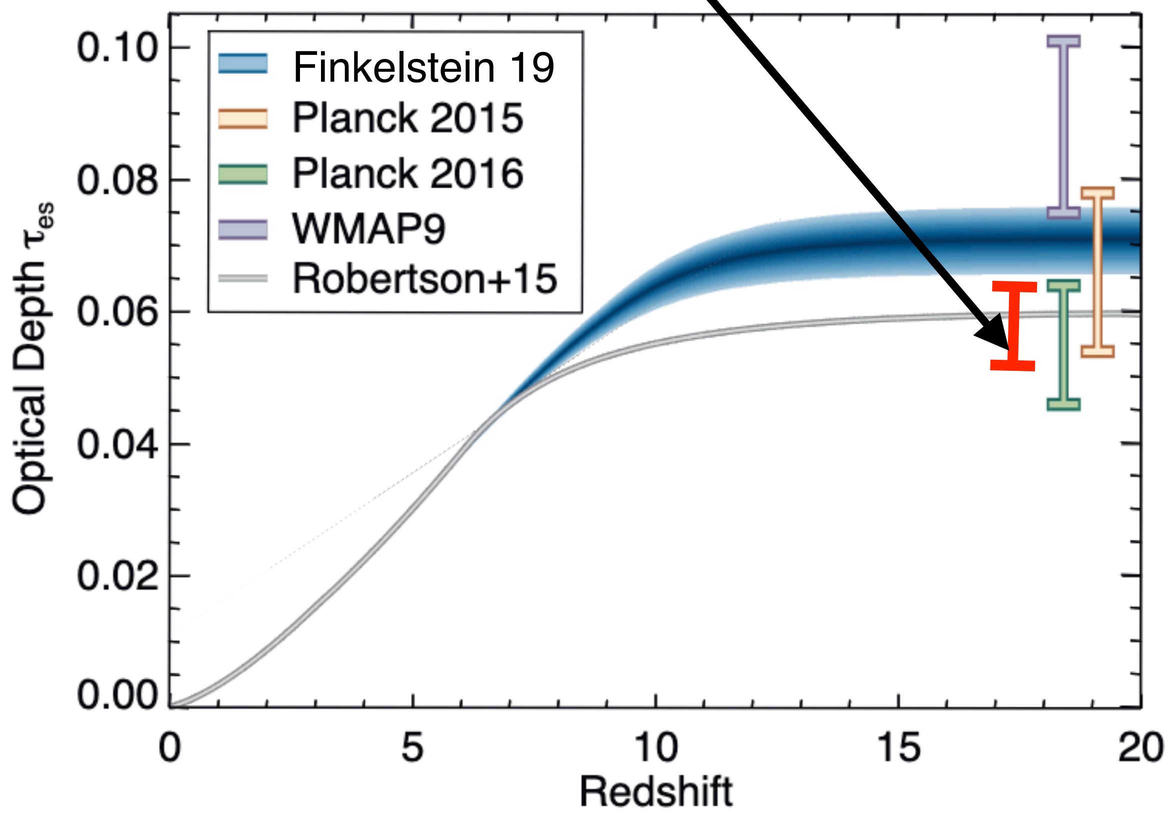
Lollipop

Tensor-to-scalar ratio & Reionization

[Tristram et al. A&A 647, A128 (2021)]
[Tristram et al. PRD 105, 083524 (2022)]

Reionization optical depth (scattering of photons by free electrons)

$$\tau = 0.0580 \pm 0.0062$$



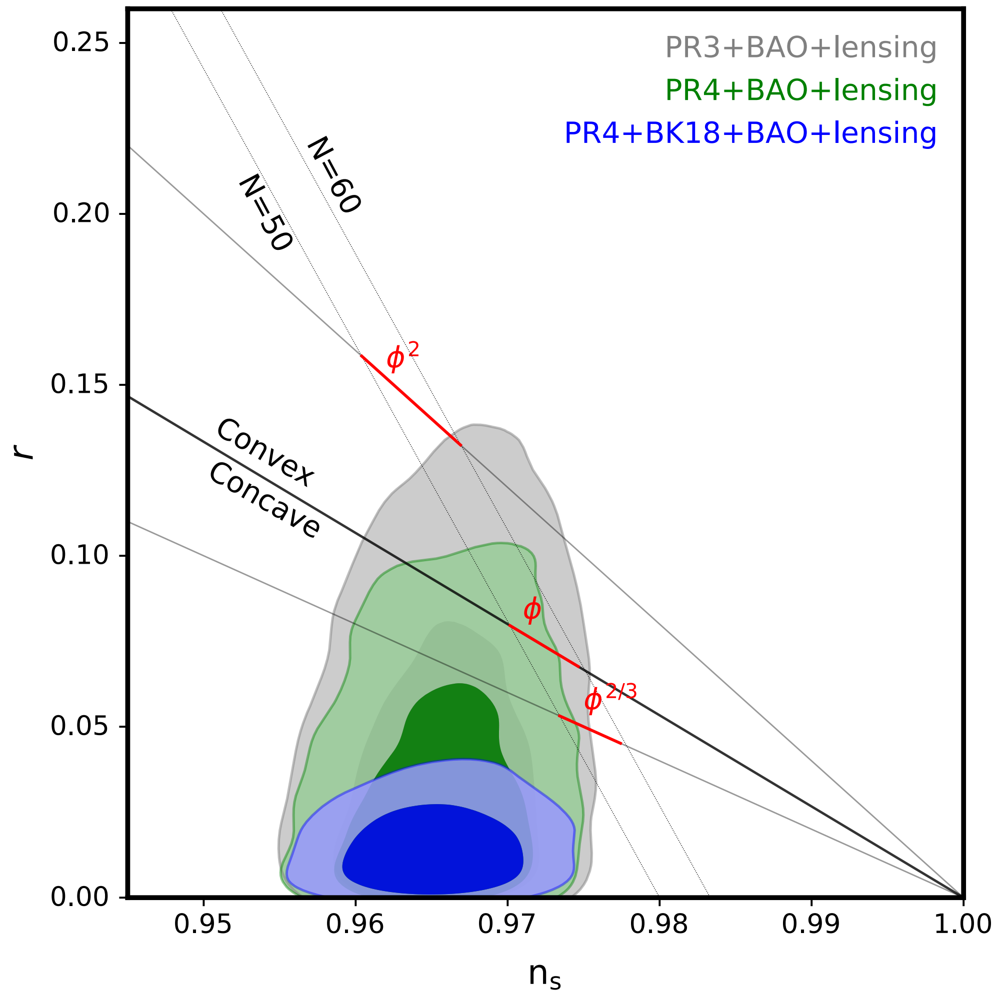
Galaxies become more efficient producers of ionizing photons at higher redshifts and fainter magnitudes

Faintest galaxies (MUV > –15) dominate the ionizing emissivity

$r_{0.05} < 0.042$	BICEP2/Keck 2018	1% of the sky
$r_{0.05} < 0.069$	Planck EB (2020)	50% of the sky

[Tristram et al. PRD 105, 083524 (2022)]

$r_{0.05} < 0.032$ (Planck + BK18)

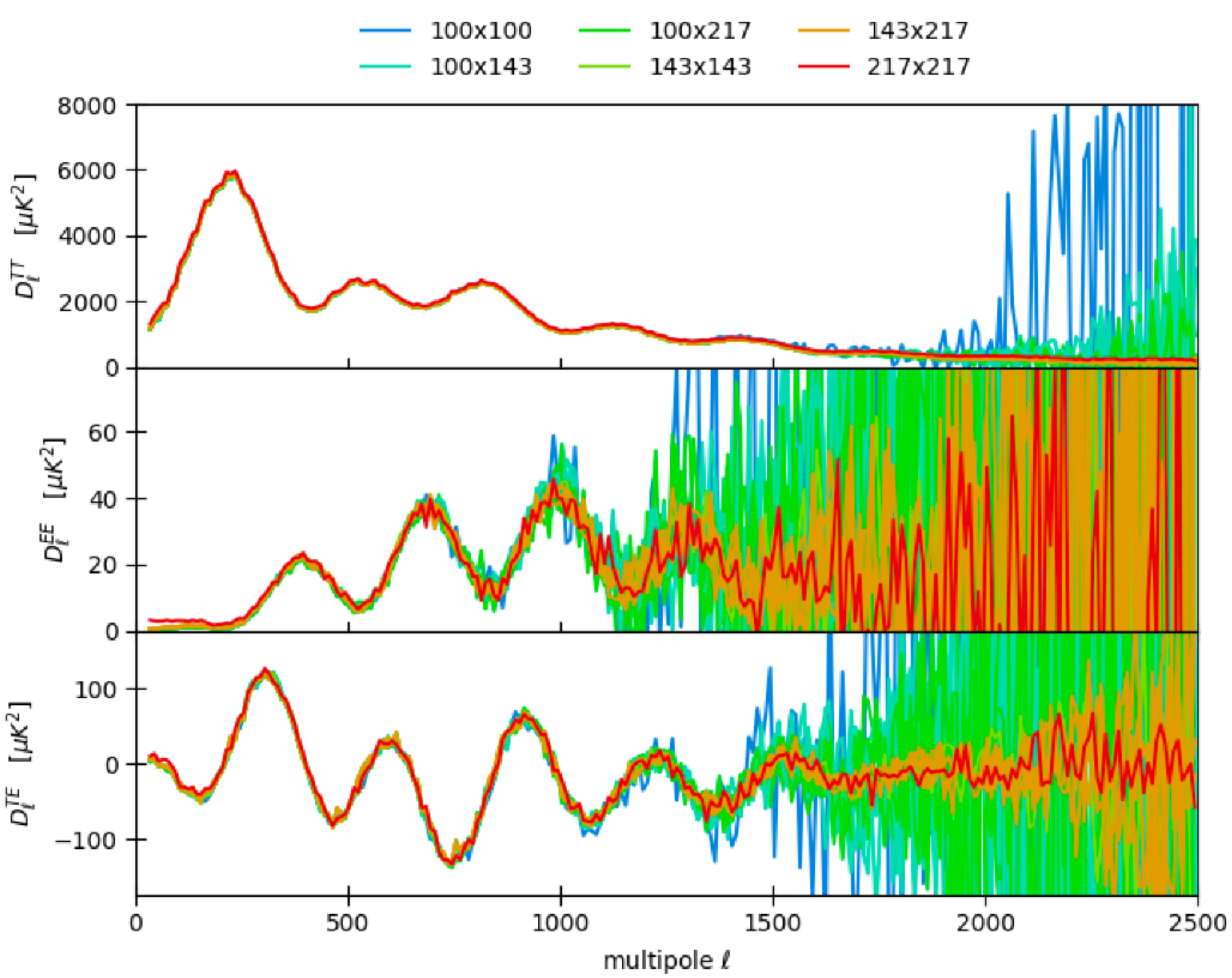


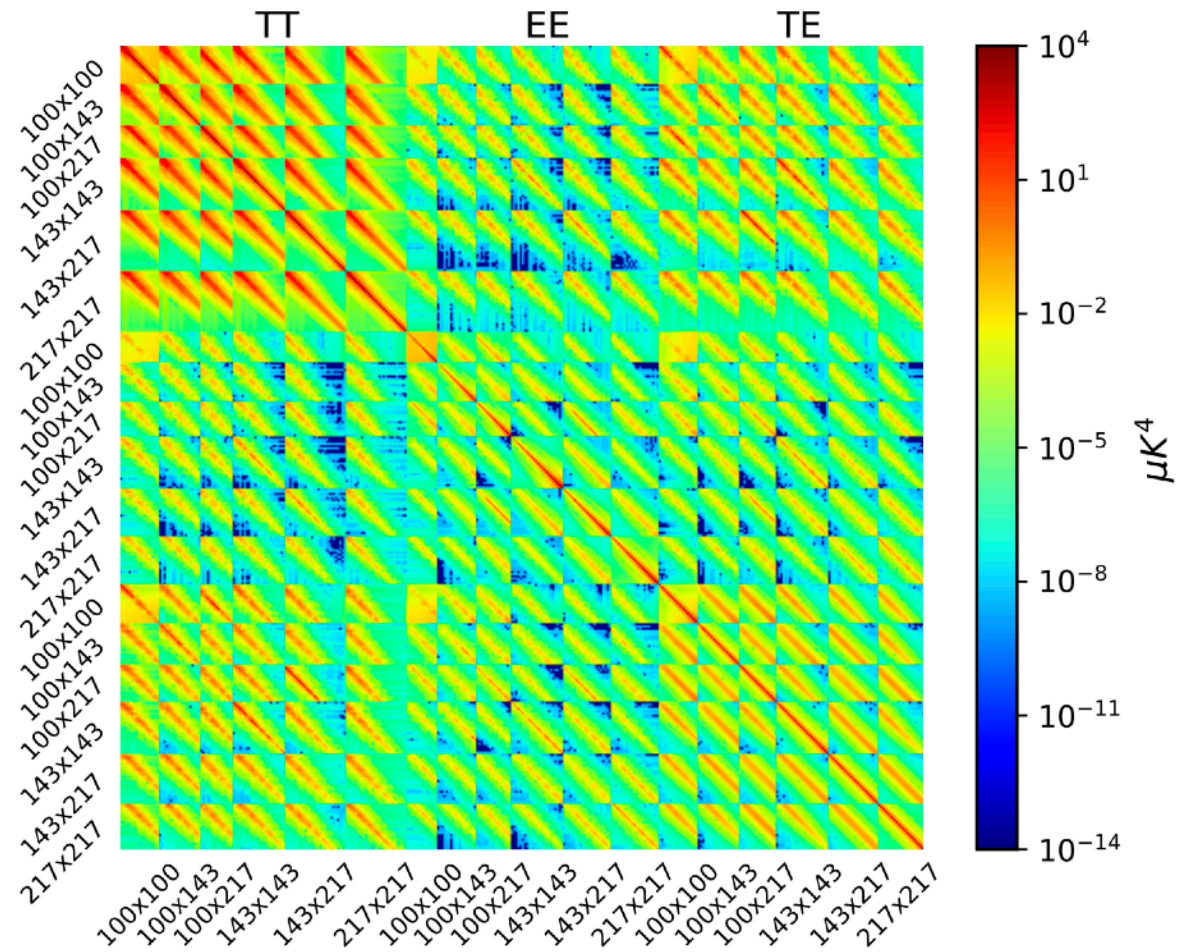
Hillipop

PR4 power-spectra

[Tristram+ (2023)]

2 maps per frequencies at 100, 143 and 217 GHz
15 cross-spectra at 6 cross-frequencies





xpol

[https://gitlab.in2p3.fr/tristram/xpol]

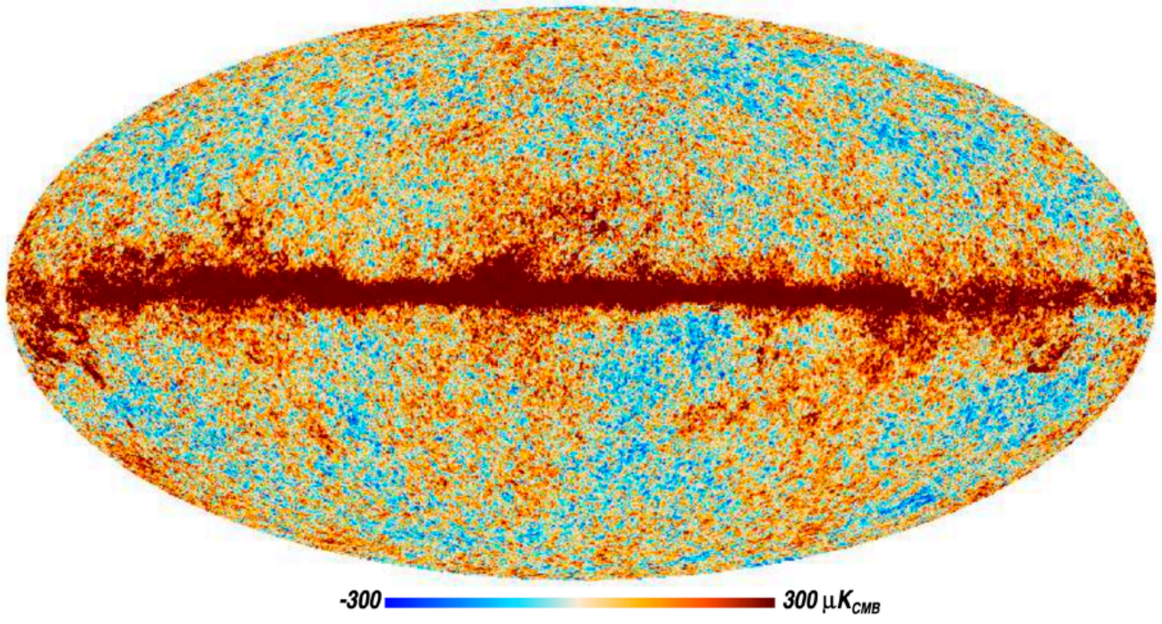
Hillipop

PR4 TT-TE-EE likelihood

[Tristram+ (2023)]

An accurate masking

- our Galaxy
- point sources
- nearby extended galaxies (e.g. M31)

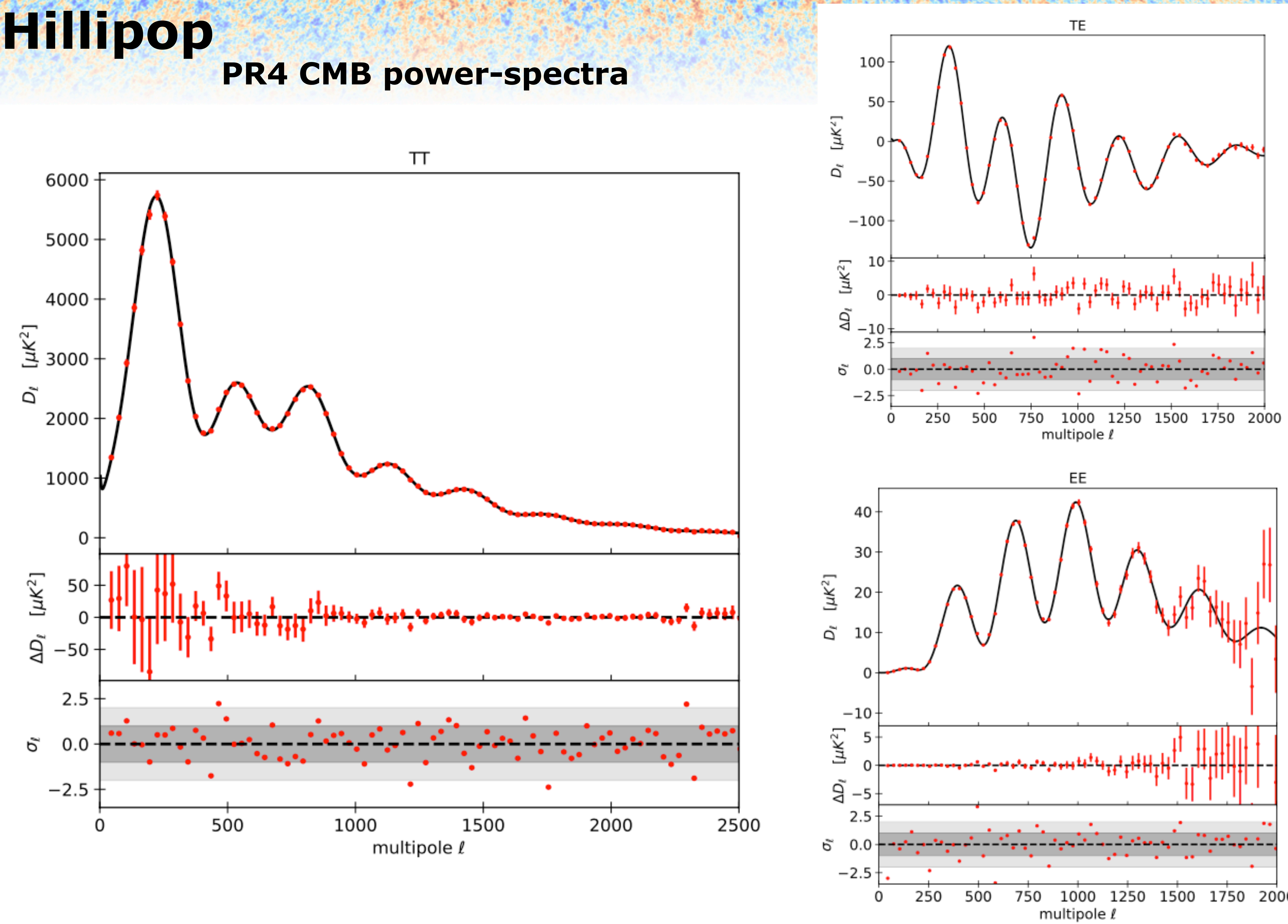


An accurate foreground model

- Galactic dust
- cosmic infrared background (CIB)
- thermal (tSZ) and kinetic (kSZ) Sunyaev-Zeldovich components
- Poisson-distributed point sources from radio and infrared star-forming galaxies
- the correlation between CIB and the tSZ effect (tSZ×CIB)

Hillipop

PR4 CMB power-spectra



ΛCDM cosmology
model

- **6 parameters**
  - 3 for the primordial matter spectra

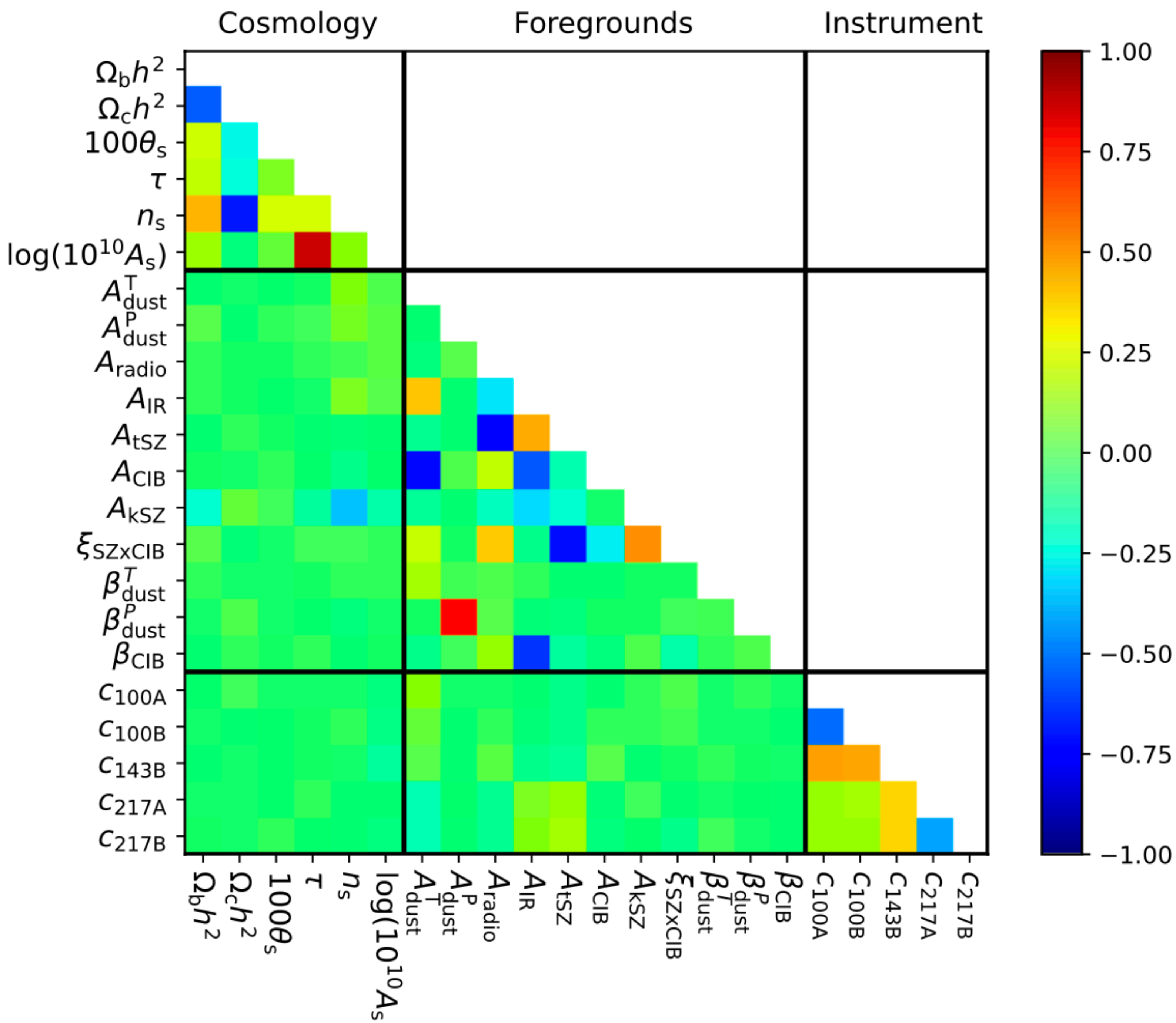
$$\mathscr{P}_s(k) = A_s \left(\frac{k}{k_0}\right)^{n_s - 1}$$

  - 1 expansion rate H_0 (in practice sound horizon θ_s)
  - 2 parameters for densities $\Omega_b h^2$ $\Omega_c h^2$
  - reionization τ

- **hypothesis**
  - flat Universe $\Omega_k = 0$
  - No running $\partial n_s / \partial \ln k = 0$
  - no tensor $r = 0$
  - 3 neutrinos $N_{\text{eff}} = 3.044$
  - standard neutrinos with low mass $\sum m_\nu = 0.06$ eV

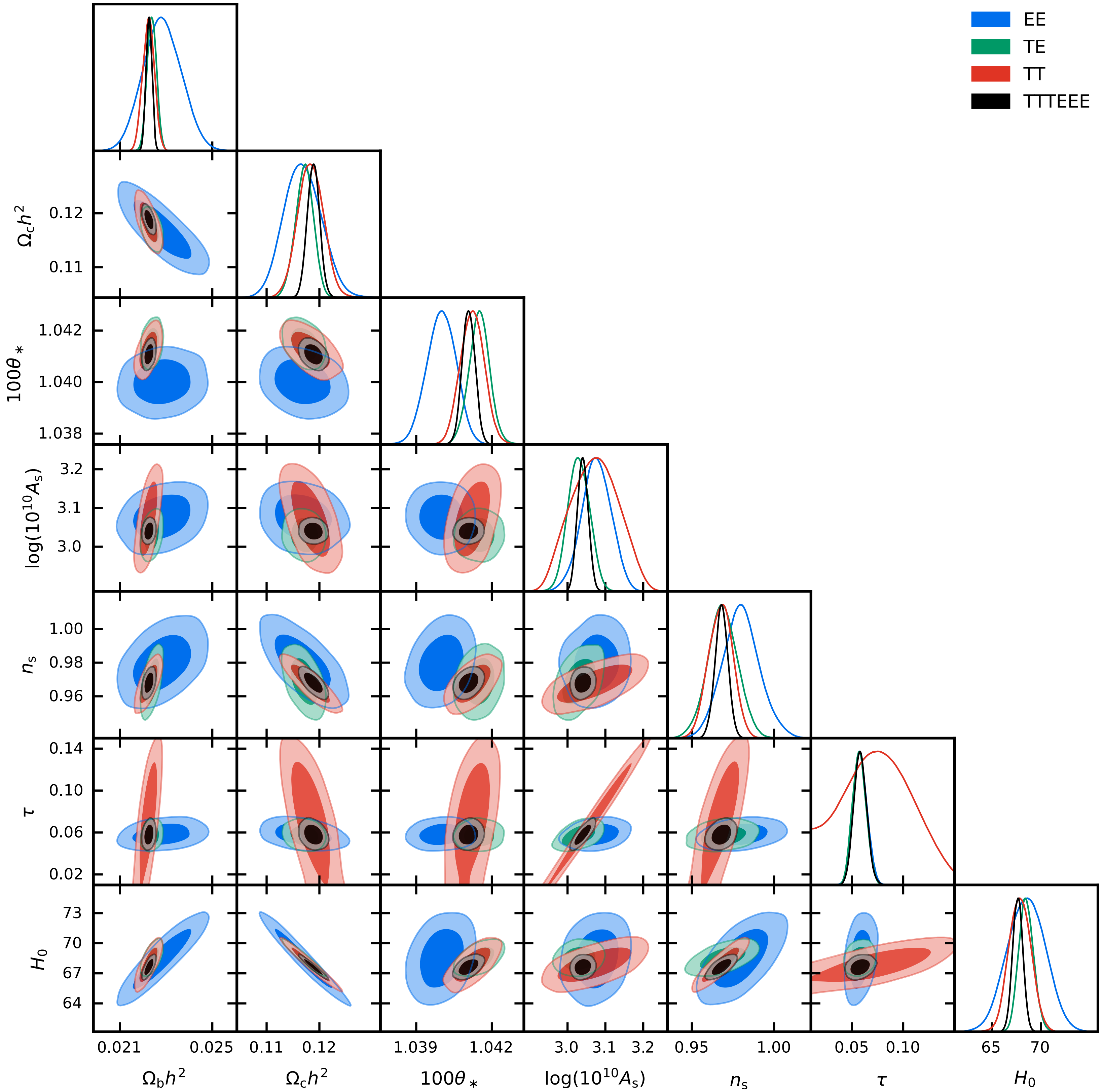
ΛCDM cosmology
parameters

[Tristram+ (2023)]



ΛCDM cosmology
TT, TE, EE

[Tristram+ (2023)]



ΛCDM cosmology

+lensing+BAO

[Tristram+ (2023)]

Parameter	TTTEEE	TTTEEE +lensing	TTTEEE +lensing+BAO	
$\Omega_b h^2$	0.02226 ± 0.00013	0.02226 ± 0.00013	0.02229 ± 0.00012	0.5 %
$\Omega_c h^2$	0.1188 ± 0.0012	0.1190 ± 0.0011	0.1186 ± 0.0009	0.75 %
$100\theta_*$	1.04108 ± 0.00026	1.04107 ± 0.00025	1.04111 ± 0.00024	0.02 %
$\log(10^{10}A_s)$	3.040 ± 0.014	3.045 ± 0.012	3.048 ± 0.012	0.39 %
n_s	0.9681 ± 0.0039	0.9679 ± 0.0038	0.9690 ± 0.0035	0.36 %
τ	0.0580 ± 0.0062	0.0590 ± 0.0061	0.0605 ± 0.0059	9.75 %
H_0	67.64 ± 0.52	67.66 ± 0.49	67.81 ± 0.38	0.56 %
σ_8	0.8070 ± 0.0065	0.8113 ± 0.0050	0.8118 ± 0.0050	0.61 %
S_8	0.819 ± 0.014	0.824 ± 0.011	0.821 ± 0.009	1.09 %
Ω_m	0.3092 ± 0.0070	0.3092 ± 0.0066	0.3071 ± 0.0051	1.66 %

But still:

- **low H0** compared to SNIa

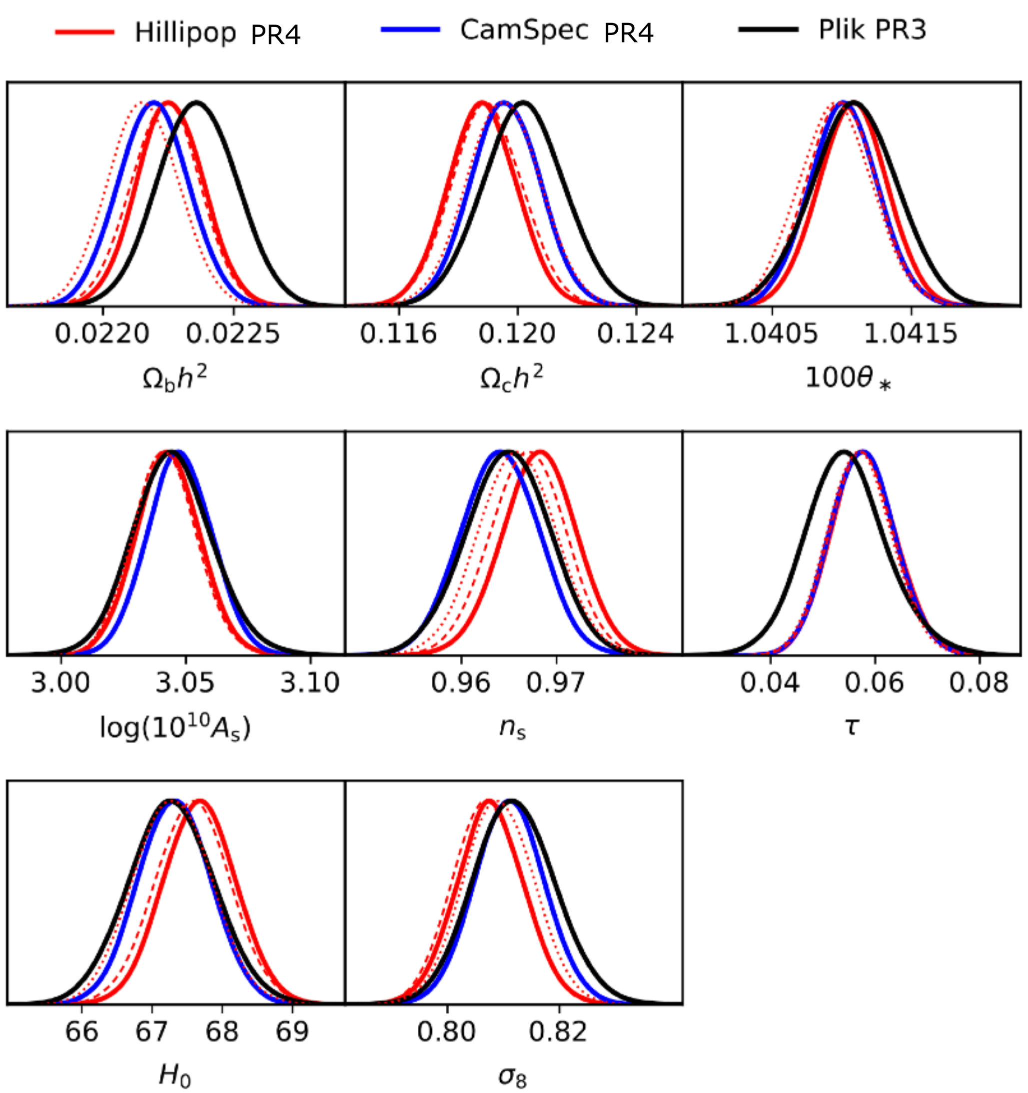
$$H_0 = 67.64 \pm 0.52 \text{ km s}^{-1} \text{ Mpc}^{-1} \quad \text{(TTTEEE)}$$
$$H_0 = 67.81 \pm 0.38 \text{ km s}^{-1} \text{ Mpc}^{-1} \quad \text{(TTTEEE+lensing+BAO)}$$

ΛCDM cosmology

Comparison with PR3 and CamSpec

[Planck 2018 Results. VI. (2020)]
[Rosenberg, Gratton, Efstathiou, MNRAS, 517, 4620 (2022)]
[Tristram+ (2023)]

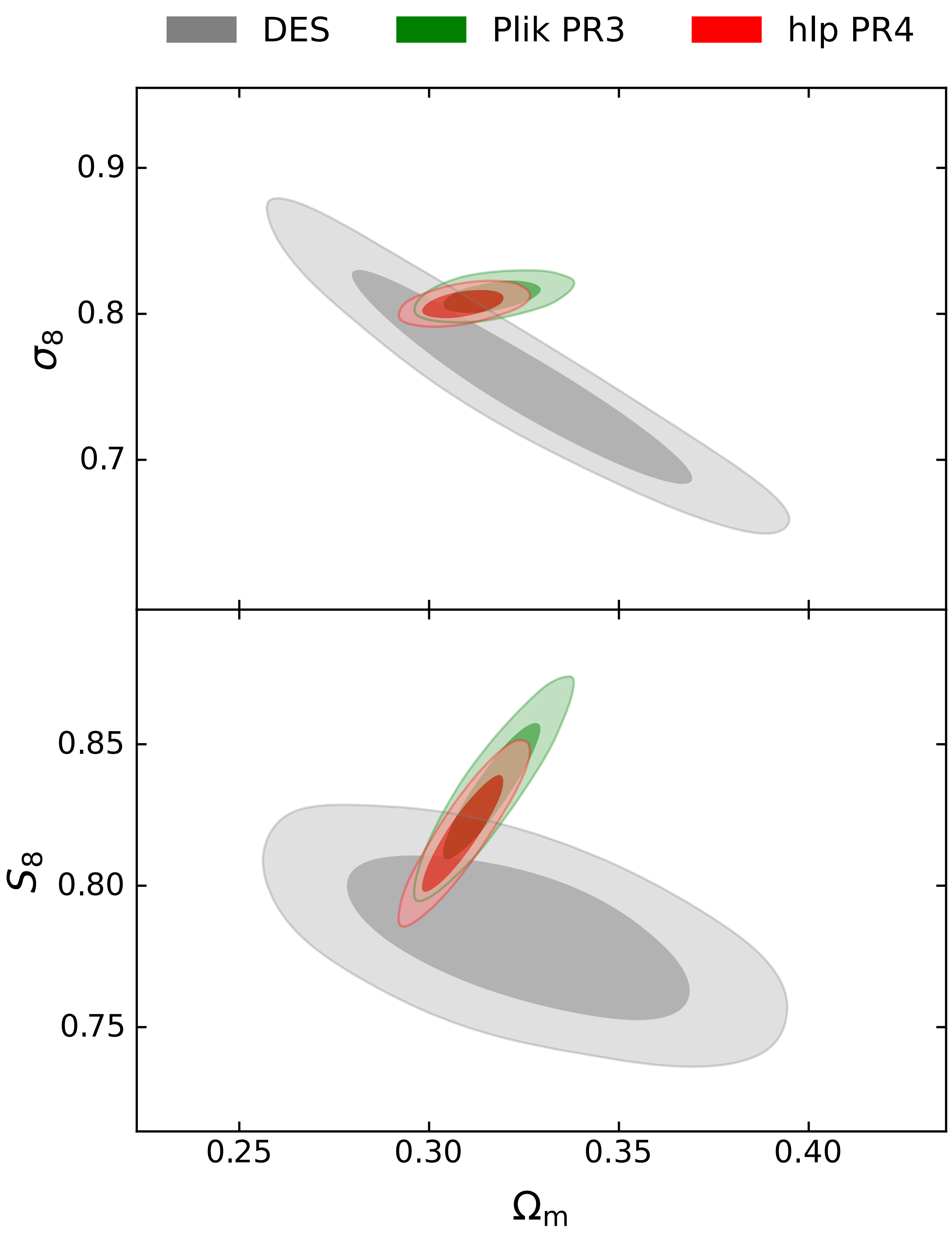
- **Good consistency between the PR4 and PR3 power spectra**, which translates to very good agreement on cosmological parameters as well.
- Lower noise of the NPIPE maps + improvement in polarization signal provides **tighter parameter constraints**, with more than **10% improvement** for ΛCDM parameters in TTTEEE



improvement wrt Planck 2018

Parameter	$\Delta\sigma$
$\Omega_b h^2$	-13.7 %
$\Omega_c h^2$	-15.2 %
$100\theta_*$	-16.1 %
$\log(10^{10}A_s)$	-12.0 %
n_s	-11.0 %
τ	-21.4 %
H_0	-13.7 %
σ_8	-11.5 %
S_8	-14.2 %
Ω_m	-16.1 %

ΛCDM cosmology
growth of structures



- **DES**

$$S_8 = 0.782 \pm 0.019 \quad \text{(DES-Y3)}$$

- **Planck**

$$S_8 = 0.834 \pm 0.016 \quad \text{(PR3 TTTEEE)}$$

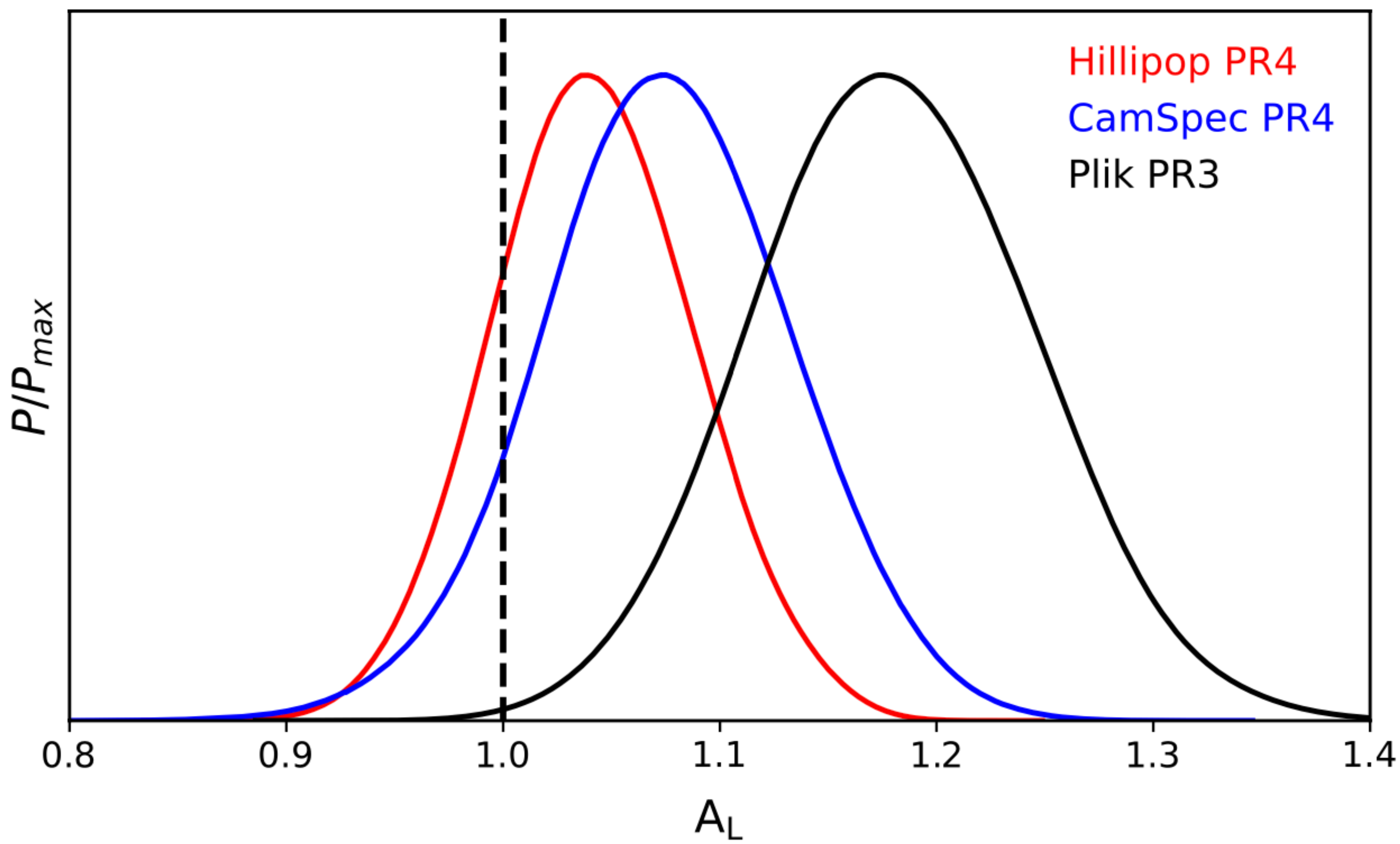
$$S_8 = 0.819 \pm 0.014 \quad \text{(PR4 TTTEEE)}$$

reduced from 2.1σ to **1.5σ**

ΛCDM extensions

A_{lens}

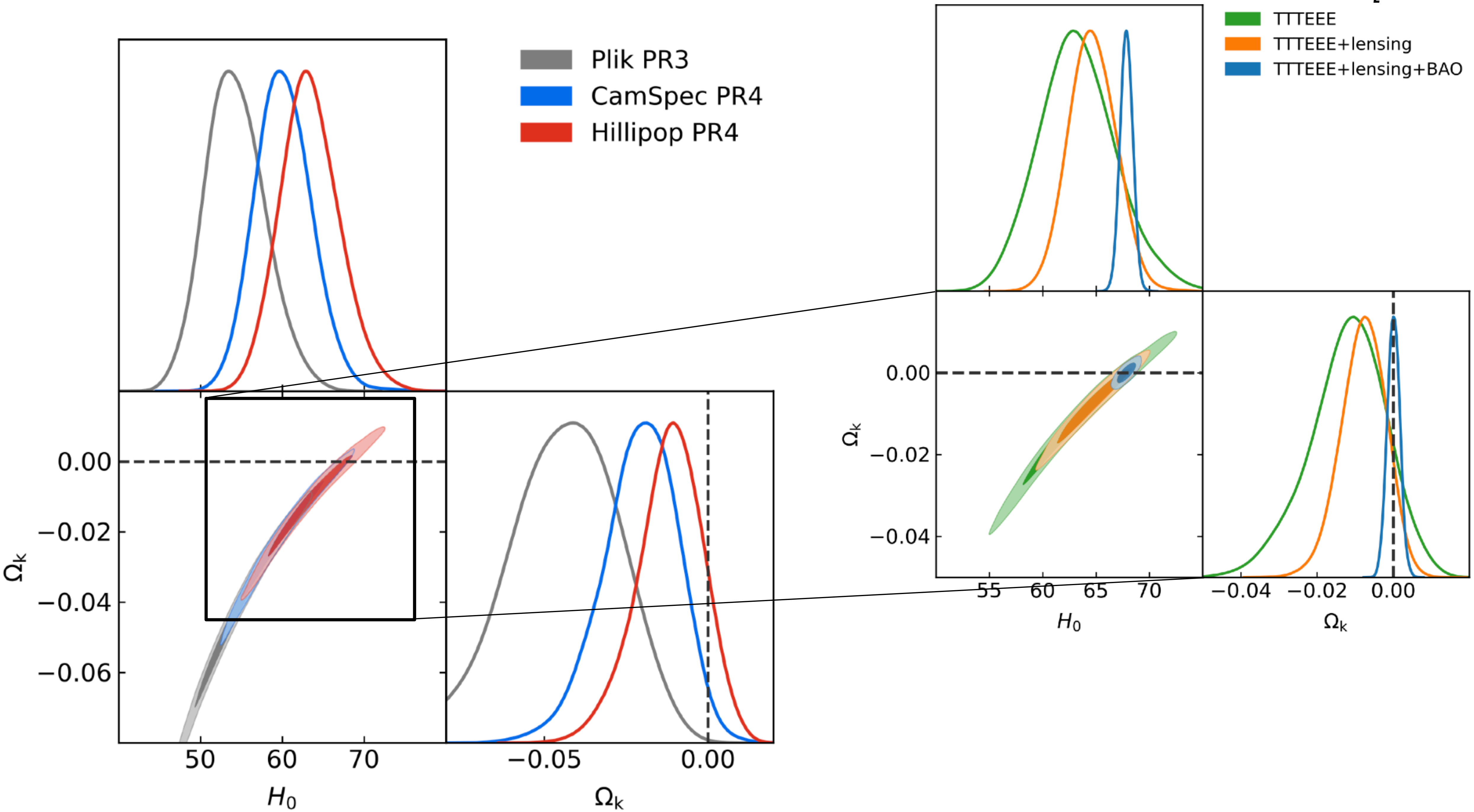
[Tristram+ (2023)]



$$A_{\rm lens} = 1.039 \pm 0.052 \quad \text{(TTTEEE)}$$
$$A_{\rm lens} = 1.037 \pm 0.037 \quad \text{(TTTEEE+lensing)}$$

ΛCDM extensions
curvature Ω_K

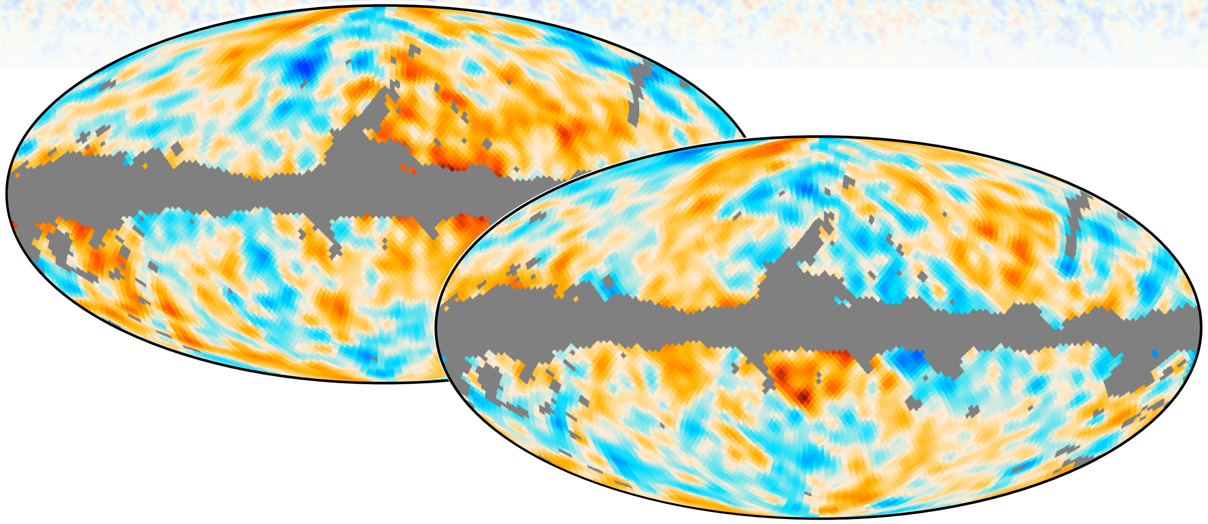
[Tristram+ (2023)]



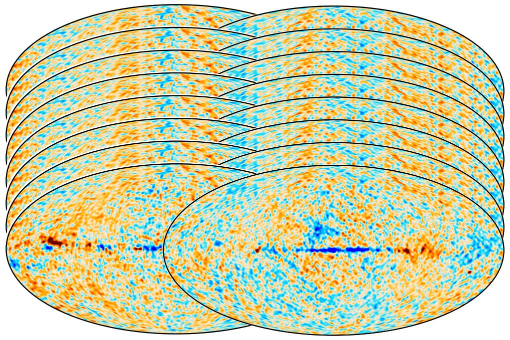
$$\Omega_K = -0.012 \pm 0.010 \quad \text{(TTTEEE)}$$

$$\Omega_K = 0.0000 \pm 0.0016 \quad \text{(TTTEEE+lensing+BAO)}$$

Conclusions



- **PR4 final PLANCK maps**
  - **cleaner** (less systematics)
  - **more sensitive** (less noisy)
  - split-maps not correlated
  - sims consistent with the data, include uncertainties from systematics (both instrumental and astrophysical)



- **CMB likelihoods (Lollipop & Hillipop)**
  - Cosmology consistent with the PR3 and with CamSpec
  - about **10% improvement** in most of ΛCDM parameters
  - give the tightest constraints from Planck CMB today
  - **no deviation from standard ΛCDM**

$$A_{\mathrm{lens}} = 1.039 \pm 0.052$$

$$\Omega_K = -0.012 \pm 0.010$$

available on github
https://github.com/planck-npipe
for Cobaya and MontePython