

# **Accounting for the beams in the parametric component separation** CMB-France 04/12/2023

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- where:
	- $\mathbf{A} = \text{component}$  *Mixing Matrix*
	- $s$  = true value signals for each component
	- $\mathbf{n}$  = instrumental noise (assumed Gaussian distributed)

# **Component separation**

### **The problem**







Credits: J. Errard

 $\mathbf{d}$  = data vector of the measured signal for all the  $n_{\!f}$  frequencies and  $n_{_S}$  Stokes parameters

### **The model**

## $d = As + n$

**unknown**





where:  $\mathbf{d}$  = data vector of the measured signal for all the  $n_f$  frequencies and  $n_s$  Stokes parameters  $\mathbf{d} =$  data vector of the measured signal for all the  $n_{\!f}$  frequencies and  $n_{\!s}$ 

# **Component separation**

### **The problem**



#### astrophysical foregrounds

*component Mixing Matrix* **parametrised by a set of unknown parameters**  $\mathbf{A}(\beta_i) =$  *component Mixing Matrix* parametrised by a set of <mark>unknown</mark> parameters  $\beta_i$ 

### **The model**

true value signals for each component  $s =$ 

 $n =$  instrumental noise (assumed Gaussian distributed)

**Parametric**

**component** 

**separation** 

## $d = As + n$

**unknown**





## **Parametric maximum likelihood based component separation The solution - maximum likelihood principle**

Full data likelihood:

*Two possibilities:* 

- Characterising  $\mathscr{L}_{data}$  numerically: mapping the likelihood by MCMC sampling
- Maximising  $\mathscr{L}_{data}$ :

■ 2 step approach: [Stompor et al. 2009](http://arxiv.org/abs/0804.2645) **Step 1** recover the spectral parameter estimates by maximising **Step 2** recover the sky components  $-2 \ln \mathcal{L}_{spec}(\beta_i) = \text{const} - (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{d})^T (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{A})^{-1} (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{d})$  $\mathbf{s} = (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{A})^{-1} \mathbf{A}^T \mathbf{N}^{-1} \mathbf{d}$ 

 (sampling at the same time signal and spectral parameters) [De la Hoz et al. 2020](https://arxiv.org/abs/2002.12206)

$$
-2\ln\mathcal{L}_{data}(\mathbf{s},\beta_i)=\text{const}+(\mathbf{d}-\mathbf{A}\mathbf{s})^T\mathbf{N}^{-1}(\mathbf{d}-\mathbf{A}\mathbf{s})
$$

[Eriksen et al. 2006](https://arxiv.org/pdf/astro-ph/0508268.pdf)

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### **Beams in component separation The problem:** *Why do we need to account for the beams in the comp sep?*

 $\mathbf{d} = \mathbf{A}\mathbf{s} + \mathbf{n}$  where  $\mathbf{d} = [d_{\nu_0 \text{ fwhm}_0}, \dots, d_{\nu_n \text{ fwhm}_n}]$ (different frequency channels have different beams)

 $-2 \ln \mathcal{L}_{spec}(\beta) = \text{const} - (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{d})^T (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{A})^{-1} (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{d})$ 

**the map-based spectral likelihood is biased**

**the recovered CMB is biased (much higher CMB residuals)**





# **Beams in component separation**  $d = (B)As + n$ **Extended data model**







#### where **B** true beams of the input frequency maps



# **Beams in component separation**  $d = (B)As + n$ **Extended data model**

where **B** true beams of the input frequency maps

Can we perform the commutation of  $\bf{B}$  and  $\bf{A}$  ?

 $\checkmark$  if  $\bf{B}$  is the same at each frequency band

 $\checkmark$  if there is no spatial variability in  $\hat{A}$  (or spatial variability on scales bigger than the beam size)

Otherwise  $B$  and  $A$  do not commute, unless it is defined an effective Mixing Matrix  $\tilde{A}$  with more complex scaling laws depending on the beams:  $\mathbf{BA} = \tilde{\mathbf{A}}\hat{\mathbf{B}}$ ̂





**component maps frequency maps smoothed at a common beam**



**each smoothed at its beam value**



**component maps smoothed at a common beam**





## **1st solution:** *before* **component separation Beams in parametric component separation**

#### **1st solution:**

dealing with the beams *before* component separation

> **Parametric comp sep**





#### **1st solution:**



## **1st solution:** *before* **component separation Beams in parametric component separation**

**each smoothed at its beam value**



#### **Modify input frequency maps and back to usual component separation:**

Input frequency maps (each convolved at its own beam  $\mathbf{B_{true}}$ ):

 $d = B_{true}(As) + n$ 

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Applying deconvolution and convolution on input frequency maps through:

#### $(As) + B_{eff}n = AB_{final}s + B_{eff}n = As' + B_{eff}n$ ̂

The usual component separation is performed on the  $\mathbf{B}_{\text{eff}}\mathbf{d}$  frequency maps, to recover the component

*maps*  $\mathbf{s}' = \mathbf{B}_{\text{final}}\mathbf{s}$  *beam-smoothed at the final resolution.* ̂

### **1st solution:** *before* **component separation Beams in parametric component separation**

$$
B_{eff}d = B_{eff}(B_{true}(As) + n) = B_{final}(As) + B_{el}
$$

 $\beta_{\rm eff}$  =  $\frac{1000}{\rm R}$  is the effective beam representing the deconvolution by  ${\rm B_{true}}$  and the convolution to a common beam  ${\rm B_{final}}$  $\mathbf{B}_{\text{final}}$  $\mathbf{B}_{\text{true}}$  $\mathbf{B_{true}}$  and the convolution to a common beam  $\mathbf{B_{final}}$ 

#### **Modify input frequency maps and back to usual component separation:**

- No need to modify the component separation technique
- Fast

### Pros vs Cons

- B<sub>eff</sub> applied to **n**, correlations introduced. The diagonal noise covariance matrix doesn't describe the data anymore. We can correct for it, but it remains an approximation.
- Common resolution has to be chosen  $\geq$  than the worst input resolution
- Need to be able to smooth the maps in a reliable way (for example needed maps at sufficiently high resolution)

### **1st solution:** *before* **component separation Beams in parametric component separation**

#### **Modify input frequency maps and back to usual component separation:**

### Example of this approach already used in the literature



Probing Cosmic Inflation with the LiteBIRD Cosmic Microwave Background Polarization Survey (LiteBIRD Collaboration, 2023)



### **1st solution:** *before* **component separation Beams in parametric component separation**

The Simons Observatory: pipeline comparison and validation for large-scale B-modes (Wolz et al, 2023) **[fgbuster = pipeline C below**]



Fig. 7. Compilation of the mean r with its  $(16, 84)$ % credible interval as derived from 500 simulations, applying the three nominal pipelines (plus extensions) to four foreground scenarios of increasing complexity. We assume a fiducial cosmology with  $r = 0$  and  $A_{\text{lens}} = 1$ , inhomogeneous noise with goal sensitivity and optimistic  $1/f$  noise component (dot markers), and inhomogeneous noise with baseline sensitivity and pessimistic  $1/f$  noise component (cross markers). Note that the NILC results for Gaussian foregrounds are based on a smaller sky mask, see Appendix B.









**component maps frequency maps smoothed at a common beam**

**each smoothed at its beam value**



dealing with the beams *in* the component separation algorithm

## **2nd solution:** *in the* **component separation Beams in parametric component separation**

**Parametric comp sep**







### **2nd solution:** *in the* **component separation Beams in parametric component separation**



back and forth pixel/harmonic domain, to apply each object where it naturally lives



Input frequency maps (each convolved at its own beam  $\mathbf{B}_{\text{true}}$ ):  $\quad \mathbf{d} = \mathbf{B}_{\text{true}}(\mathbf{A}\mathbf{s}) + \mathbf{n}$ It can be rewritten as:  $\mathbf{d} = \mathbf{B}_{true} \mathbf{B}_{final}$ where:  $\mathbf{s}' = \mathbf{B}_{\text{final}} \mathbf{s}$  is the recovered components beam-smoothed at the final resolution  $A' = B_{true}B_{final}^{-1}A$  (redefinition of the Mixing Matrix) −1 ̂ ̂ −1 **A**

> **Step 1** recover the spectral parameter estimates by maximising **Step 2** recover the sky components −2 ln ℒ*spec*(*β<sup>i</sup>*  $\mathbf{D} = \text{const} - (\mathbf{A}'^T \mathbf{N}^{-1} \mathbf{d})^T (\mathbf{A}'^T \mathbf{N}^{-1} \mathbf{A}')^{-1} (\mathbf{A}'^T \mathbf{N}^{-1} \mathbf{d})$

$$
-2\ln\mathcal{L}_{spec}(\beta_i) =
$$

 $\mathbf{s} = (\mathbf{A}'^T \mathbf{N}^{-1} \mathbf{A}')^{-1} \mathbf{A}'^T \mathbf{N}^{-1} \mathbf{d}$ 

- 
- $AB$  **final**<sup> $S$ </sup> + **n** =  $A'S'$  + **n** 
	-
	-

### **2nd solution:** *in the* **component separation Beams in parametric component separation**

*Comp sep:* 

■ 2 step approach

**Step 1** recover the spectral parameter estimates by maximising  $-2 \ln \mathcal{L}_{spec}(\beta_i) = \text{const} - (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{d})^T (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{A}')^{-1} (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{d})$ 

- 
- **Step 1**  $\mathscr{L}_{\text{spect}} = -(\mathbf{d}^T \mathbf{N}^{-1} \mathbf{A}')\mathbf{x}$  (PCG at each step of the maximisation)

### **2nd solution:** *in the* **component separation Beams in parametric component separation**



Solved with a PCG solver:  $A^T N^{-1} A' x = A^T N^{-1} d$ 

 $\mathscr{L}_{\text{spect}} = -(\mathbf{d}^T \mathbf{N}^{-1} \mathbf{A}')\mathbf{x}$ 

**Step 2** recovered components: **x**







### Pros vs Cons

- More rigorous treatment
- The recovered component maps resolutions are not limited by the worst input map resolution
- Edge effects due to cut sky, where the convolution doesn't perform well
- Potentially more computationally involved

### **Adding the beam operator in the spectral likelihood**

### **2nd solution:** *in the* **component separation Beams in parametric component separation**



## **2nd solution:** *in the* **component separation Beams in parametric component separation**







### **2nd solution:** *in the* **component separation Beams in parametric component separation**



Full Sky *fgs models d0s0*

- CMB residuals (input fwhm=LB-like, final fwhm=17arcmin)
- CMB residuals (input fwhm=80arcmin)
- dust
- synch
- $D_{\ell}^{EE}$  ref



The input frequency maps are often convolved to different resolutions, we need a strategy to deal with this:

- ‣ 1st solution: dealing with the beams *before* component separation
- ‣ 2nd solution: we add the beam operator **in** the component separation method

*Next steps*:

- study the effect of **more complex beam** in the component separation, exploiting more advanced softwares to perform the deconvolution ex.
- use implementation with PCG to deal with correlated noise
- going to **TODs domain** (to include pixel dependent beams)







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# **Conclusion and prospects**

Apply the beam in the harmonic domain, but stay in pixel domain as much as possible (to be able to address the spatial variability of the foregrounds)







➡ 2 step procedure: we do not directly maximised the full data likelihood, but rather the *spectral likelihood* (computationally easier!) "true" parameters estimated parameters

## **Parametric maximum likelihood based component separation The solution - two step procedure**





### [Stompor et al. 2009](http://arxiv.org/abs/0804.2645)





#### **Commutation of the beam operator and the Mixing Matrix:**

- $\blacktriangleright$  If  $\bf{A}$  doesn't depend on the pixel and  $\bf{B}$  is the same for all frequency channels:  $\bf{BA} = \bf{A}\bf{B}$ The data model becomes:  $\mathbf{d} = \mathbf{A}\mathbf{s}' + \mathbf{n}$  where  $\mathbf{s}' = \mathbf{B}\mathbf{s}$  are beam-smoothed component maps. (This can be extended to  ${\bf A}$  with spatial variability as long as those are at scale bigger than the beam size.) ̂ ̂
- $\blacktriangleright$  **If A** does depend on the pixel:  $\mathbf{BA} \neq \mathbf{AB}$ . We can still have:  $\mathbf{BA} = \mathbf{A}'\mathbf{B}$ , where  $\mathbf{A}'$  is a modified ("effective") Mixing Matrix. The data model becomes:  $\mathbf{d} = \mathbf{A}'\mathbf{s}' + \mathbf{n}$ ̂ ̂

# **Beams in parametric component separation**

 $\mathbf{d} = \mathbf{B}\mathbf{A}\mathbf{s} + \mathbf{n}$  where  $\mathbf{B}$  are the true beams (could be the same for all the frequencies or not)

#### **The problem:** *How to recover component maps from beam convolved frequency maps?*



**smoothed at a common beam**



**frequency maps each smoothed at its beam value**

## **Extension to 1st solution:** *before* **component separation Beams in parametric component separation**

#### **Performing the comp sep multiple times, each time with less maps, but better resolution.**







## **Extension to 1st solution:** *before* **component separation Beams in parametric component separation**

The resolution is limited by the resolution of the larger fwhm, therefore two regimes explored. The final result is a combination of those.