

Accounting for the beams in the parametric component separation CMB-France 04/12/2023

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Component separation

The problem





The model

$\mathbf{d} = \mathbf{As} + \mathbf{n}$

- where:
 - A = component Mixing Matrix
 - s = true value signals for each component
 - $\mathbf{n} = instrumental noise$ (assumed Gaussian distributed)

 $\mathbf{d} = \text{data vector of the measured signal for all the <math>n_f$ frequencies and n_s Stokes parameters

Unknown







Component separation

The problem

astrophysical foregrounds



The model

sepc

$\mathbf{d} = \mathbf{As} + \mathbf{n}$

Parametric $A(\beta_i) = component Mixing Matrix parametrised by a set of unknown parameters <math>\beta_i$ s = true value signals for cost to β_i $\mathbf{d} = \text{data vector of the measured signal for all the <math>n_f$ frequencies and n_s Stokes parameters

 $\mathbf{n} = \text{instrumental noise}$ (assumed Gaussian distributed)

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Parametric maximum likelihood based component separation The solution - maximum likelihood principle

Full data likelihood:

$$-2\ln \mathscr{L}_{data}(\mathbf{s},\beta_i) = \text{const} + (\mathbf{d} - \mathbf{A}\mathbf{s})^T \mathbf{N}^{-1} (\mathbf{d} - \mathbf{A}\mathbf{s})$$

Two possibilities:

Characterising \mathscr{L}_{data} numerically: mapping the likelihood by MCMC sampling

• Maximising
$$\mathscr{L}_{data}$$
:

2 step approach: <u>Stompor et al. 2009</u> **Step 1** recover the spectral parameter estimates by maximising $-2\ln \mathscr{L}_{spec}(\beta_i) = \operatorname{const} - (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{d})^T (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{A})^{-1} (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{d})$ **Step 2** recover the sky components $\mathbf{s} = (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{A})^{-1} \mathbf{A}^T \mathbf{N}^{-1} \mathbf{d}$

De la Hoz et al. 2020 (sampling at the same time signal and spectral parameters)

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Eriksen et al. 2006

Beams in component separation The problem: Why do we need to account for the beams in the comp sep?

where $\mathbf{d} = [d_{\nu_0 \text{ fwhm}_0}, \dots, d_{\nu_n \text{ fwhm}_n}]$ $\mathbf{d} = \mathbf{A}\mathbf{s} + \mathbf{n}$ (different frequency channels have different beams)

 $-2\ln \mathscr{L}_{spec}(\beta) = \operatorname{const} - (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{d})^T (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{A})^{-1} (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{d})$

the map-based spectral likelihood is biased

the recovered CMB is biased (much higher CMB residuals)



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Beams in component separation Extended data model $\mathbf{d} = \mathbf{B}\mathbf{A}\mathbf{S} + \mathbf{n}$





where **B** true beams of the input frequency maps



Beams in component separation Extended data model $\mathbf{d} = \mathbf{B}\mathbf{A}\mathbf{s} + \mathbf{n}$

where **B** true beams of the input frequency maps

Can we perform the commutation of ${f B}$ and ${f A}$?

 \checkmark if **B** is the same at each frequency band

 \checkmark if there is no spatial variability in A (or spatial variability on scales bigger than the beam size)

Otherwise **B** and **A** do not commute, unless it is defined an effective Mixing Matrix $ilde{\mathbf{A}}$ with more complex scaling laws depending on the beams: $\mathbf{B}\mathbf{A} = \tilde{\mathbf{A}}\hat{\mathbf{B}}$







frequency maps each smoothed at its beam value

component maps smoothed at a common beam







1st solution:

dealing with the beams **before**

Parametric comp sep

> component maps smoothed at a common beam







frequency maps each smoothed at its beam value

1st solution:

dealing with the beams **before**

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Modify input frequency maps and back to usual component separation:

Input frequency maps (each convolved at its own beam \mathbf{B}_{true}):

 $\mathbf{d} = \mathbf{B}_{true}(\mathbf{As}) + \mathbf{n}$

Applying deconvolution and convolution on input frequency maps through:

$$\mathbf{B}_{eff}\mathbf{d} = \mathbf{B}_{eff}(\mathbf{B}_{true}(\mathbf{As}) + \mathbf{n}) = \mathbf{B}_{final}(\mathbf{As}) + \mathbf{B}_{eff}(\mathbf{As}) + \mathbf{B}_$$

 $(B_{eff} = \frac{B_{final}}{B_{true}}$ is the effective beam representing the deconvolution by B_{true} and the convolution to a common beam B_{final})

maps $\mathbf{s}' = \mathbf{B}_{final}\mathbf{s}$ beam-smoothed at the final resolution.

$A_{\text{ff}}n = A\hat{B}_{\text{final}}s + B_{\text{eff}}n = As' + B_{\text{eff}}n$

The usual component separation is performed on the ${f B}_{
m eff}{f d}$ frequency maps, to recover the component

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Modify input frequency maps and back to usual component separation:

Pros

- No need to modify the component separation technique
- Fast

Cons VS

- **B**_{eff} applied to **n**, correlations introduced. The diagonal noise covariance matrix doesn't describe the data anymore. We can correct for it, but it remains an approximation.
- Common resolution has to be chosen \geq than the worst input resolution
- Need to be able to smooth the maps in a reliable way (for example needed maps at sufficiently high resolution)

Modify input frequency maps and back to usual component separation:

Example of this approach already used in the literature

Probing Cosmic Inflation with the LiteBIRD Cosmic Microwave Background Polarization Survey (LiteBIRD Collaboration, 2023)



The Simons Observatory: pipeline comparison and validation for large-scale B-modes (Wolz et al, 2023) [fgbuster = pipeline C below]



Fig. 7. Compilation of the mean r with its (16, 84)% credible interval as derived from 500 simulations, applying the three nominal pipelines (plus extensions) to four foreground scenarios of increasing complexity. We assume a fiducial cosmology with r = 0 and $A_{\text{lens}} = 1$, inhomogeneous noise with goal sensitivity and optimistic 1/f noise component (dot markers), and inhomogeneous noise with baseline sensitivity and pessimistic 1/f noise component (cross markers). Note that the NILC results for Gaussian foregrounds are based on a smaller sky mask, see Appendix B.

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frequency maps each smoothed at its beam value

2nd solution:

dealing with the beams *in* the component separation algorithm

> **Parametric** comp sep



component maps

smoothed at a common beam





domain	applying B	spatial variability of A	homogenous noise	ell correlated noise
harmonic	\checkmark	X	X	\checkmark
pixel	X	\checkmark	\checkmark	X
mixed	\checkmark	\checkmark	\checkmark	\checkmark

back and forth pixel/harmonic domain, to apply each object where it naturally lives





Input frequency maps (each convolved at its own beam B_{true}): $d = B_{true}(As) + n$ It can be rewritten as: $\mathbf{d} = \mathbf{B}_{true} \mathbf{B}_{final}^{-1} \mathbf{A} \hat{\mathbf{B}}_{final} \mathbf{s} + \mathbf{n} = \mathbf{A}' \mathbf{s}' + \mathbf{n}$ where: $\mathbf{A}' = \mathbf{B}_{true} \mathbf{B}_{final}^{-1} \mathbf{A}$ (redefinition of the Mixing Matrix)

<u>Comp sep</u>:

➡ 2 step approach

Step 1 recover the spectral parameter estimates by maximising = const - $(\mathbf{A}'^T \mathbf{N}^{-1} \mathbf{d})^T (\mathbf{A}'^T \mathbf{N}^{-1} \mathbf{A}')^{-1} (\mathbf{A}'^T \mathbf{N}^{-1} \mathbf{d})$ **Step 2** recover the sky components $\mathbf{s} = (\mathbf{A}'^T \mathbf{N}^{-1} \mathbf{A}')^{-1} \mathbf{A}'^T \mathbf{N}^{-1} \mathbf{d}$

$$-2\ln \mathscr{L}_{spec}(\beta_i) =$$

- $\mathbf{s}' = \hat{\mathbf{B}}_{\text{final}} \mathbf{s}$ is the recovered components beam-smoothed at the final resolution







Solved with a PCG solver: $\mathbf{A}^{T}\mathbf{N}^{-1}\mathbf{A}^{T}\mathbf{x} = \mathbf{A}^{T}\mathbf{N}^{-1}\mathbf{A}$

Step 1

Step 2

recovered components: **X**

Step 1 recover the spectral parameter estimates by maximising $-2\ln \mathscr{L}_{spec}(\beta_i) = \operatorname{const} - (\mathbf{A}^{T}\mathbf{N}^{-1}\mathbf{d})^{T}(\mathbf{A}^{T}\mathbf{N}^{-1}\mathbf{A}^{T})^{-1}(\mathbf{A}^{T}\mathbf{N}^{-1}\mathbf{d})$

- $\mathscr{L}_{spect} = -(\mathbf{d}^T \mathbf{N}^{-1} \mathbf{A}') \mathbf{X} \quad \text{(PCG at each step of the maximisation)}$



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Adding the beam operator in the spectral likelihood

Pros

- More rigorous treatment
- The recovered component maps resolutions are not limited by the worst input map resolution

vs Cons

- Edge effects due to cut sky, where the convolution doesn't perform well
- Potentially more computationally involved











LiteBIRD-like bands Full Sky fgs models d0s0

- CMB residuals (input fwhm=LB-like, final fwhm=17arcmin)
- CMB residuals (input fwhm=80arcmin)
- dust
- synch
- -- D_{ℓ}^{EE} ref





Conclusion and prospects

The input frequency maps are often convolved to different resolutions, we need a strategy to deal with this:

- Ist solution: dealing with the beams before component separation
- 2nd solution: we add the beam operator in the component separation method

Apply the beam in the harmonic domain, but stay in pixel domain as much as possible (to be able to address the spatial variability of the foregrounds)

<u>Next steps</u>:

- study the effect of **more complex beam** in the component separation, exploiting more advanced softwares to perform the deconvolution ex.
- use implementation with PCG to deal with correlated noise
- going to **TODs domain** (to include pixel dependent beams)







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Thanks





Parametric maximum likelihood based component separation The solution - two step procedure

2 step procedure: we do not directly maximised the full data likelihood, but rather the spectral likelihood (computationally easier!) "true" parameters estimated parameters



<u>Stompor et al. 2009</u>

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Beams in parametric component separation

d = BAs + n where B are the true beams (could be the same for all the frequencies or not)

Commutation of the beam operator and the Mixing Matrix:

- If A doesn't depend on the pixel and B is the same for all frequency channels: $BA = A\hat{B}$ The data model becomes: $\mathbf{d} = \mathbf{A}\mathbf{s}' + \mathbf{n}$ where $\mathbf{s}' = \hat{\mathbf{B}}\mathbf{s}$ are beam-smoothed component maps. (This can be extended to \mathbf{A} with spatial variability as long as those are at scale bigger than the beam size.)
- If A does depend on the pixel: $BA \neq AB$. We can still have: $BA = A'\hat{B}$, where A' is a modified ("effective") Mixing Matrix. The data model becomes: $\mathbf{d} = \mathbf{A}'\mathbf{s}' + \mathbf{n}$

The problem: How to recover component maps from beam convolved frequency maps?







frequency maps each smoothed at its beam value component maps smoothed at a common beam

Performing the comp sep multiple times, each time with less maps, but better resolution.

The resolution is limited by the resolution of the larger fwhm, therefore two regimes explored. The final result is a combination of those.

	final (common) resolution	frequency band used	reconstructed components
regime 1	80 arcmin	all bands	CMB, dust, synch
regime 2	40 arcmin	excluded lower freq bands	CMB, dust

