ATMOSPHERICS ANALYSIS First implementation in MaCh3

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WHAT ARE WE TRYING TO SEE?



• Sensitivity to $\delta_{\rm CP}$ coming from sub-GeV upgoing ν



• Sensitivity to MH coming from few-GeV upgoing ν

МаСн3

MCMC principle

- Sampling likelihood space using Markov Chains
- Stepping in parameter space targeting high-likelihood regions
- Output is a series of steps. Their distribution in parameter space represents the posterior probability



Interpreting the output

- All the variables are sampled simulteanously by Markov Chains
- Full posterior in parameter space available in the end. Allows for any projection through marginalization.



Input sample



INPUT SAMPLE

Sample characteristics

- 300k events in the HD 1x2x6 AV
- Production identical to the large sample request
- Correct fluxes and normalizations applied after the CAF stage $\rightarrow 400 \, \rm kt \, yr$
- Sample split in 24 sub-samples with $\nu_{e,\mu}$ flux, $\nu_{e,\mu,\tau}$ detected and $\nu_{e,\mu}$ reconstructed.

Assumptions

- Perfect flavour determination assumed (as current CVN cannot be used for atmospherics)
- Using the true lepton direction as reco direction
- Using only the ν events (no $\bar{\nu}$)

ENERGY RESOLUTION



Ve

- Reconstructed energy taken as is.
- No correction for energy-scale (center of $1 - E_{u}^{\rm rec} / E_{u}^{\rm truth}$ distribution).
- More details in Henrique's talk.

ANGULAR RESOLUTION



- Using the true lepton direction \rightarrow not too far from reconstructed performances

Oscillations calculation



OSCILLATIONS CALCULATION

Atmospherics oscillations now implemented in MaCh3 for DUNE.

 v_{μ} selec



Oscillations calculated on true variables. Likelihood computed on reco quantities.

 ν_{μ}

Stats-only fit



STATISTICS

Generated 40 chains with 100k steps per chain.

Burn-in

- First steps are biased by the initial values \rightarrow removing first steps of each chain.
- Decide on the burn-in number looking at the steps autocorrelation
- 10k burn-in chosed here.



<u>NB</u>: Not trying to draw big conclusions looking at the following results. Proof of concept of doing an atmospherics oscillation analysis with MaCh3 for DUNE.

LOOKING AT THE CHAINS



• Likelihood in both negative and positive δ_{CP} (sin δ_{CP} vs cos δ_{CP} degeneracy)

LOOKING AT THE CHAINS



- Likelihood in both negative and positive δ_{CP} (sin δ_{CP} vs cos δ_{CP} degeneracy)
- A few steps exploring the IH, not staying there.

1D POSTERIOR PROBABILITY



• Credibility intervals obtained by taking bins with the highest probability until having 68/99/99.73% of probability inside.

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2D POSTERIOR PROBABILITY



Getting the 2D posterior

- Get all the parameters combinations with no additional calculations.
- Marginalization other all the other parameters

2.5

 Δm_{22}^2

2.6

1e-3

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Atmospherics analysis \cdot Stats-only fit $\circ \circ \circ \circ \circ \bullet$

Adding systematics



Adding systematics



SPLINED SYSTEMATICS

Spline-response functions

- Pre-compute the impact of different values of the parameter on "knots" (evaluation points)
- Evaluate looping over the MC sample, for each event.
- Evaluate variations in-between knots by interpolating with piecewise cubic splines.

Nucleon inelastic reaction probability



 \rightarrow Computed using *nusystematics*

Splines are then averaged in E_{ν}^{true} , E_{ν}^{reco} and mode, bins with all the events.

EFFECT OF THE XSEC SYSTEMATICS (PRELIMINARY)



There is an effect of the xsec systematics in degrading the sensitivity \rightarrow good sign

SUMMARY

Summary

- First implementation of DUNE atmospherics analysis within MaCh3 (common framework with LBL analysis).
- Managed to have the first stats-only fits with a limited sample and some assumptions
- First implementation of the LBL xsec systematics

Future work

- Lot of work ongoing to improve the reconstruction for atmospherics (see Henrique's presentation).
- Plan on implementing soon several flux systematics.
- Starting considering non-inclusive xsec systematics for atmospherics (angular dependency).
- Should be able to make use of the future new atmospheric sample.

Backup slides

OSCILLATIONS CALCULATION

Atmospherics oscillations now implemented in MaCh3 for DUNE.



 v_{o} selec

Oscillations calculated on true variables. Likelihood computed on reco quantities.

Vp

IMPLEMENTING SPLINED SYSTEMATICS

LBL xsec systematics

Spline binning

Splines are averaged in E_{ν}^{true} , E_{ν}^{reco} bins with all the events.



USING THE XSEC SYSTEMATICS

Change of GENIE tune

- A new GENIE tune is used for all the new DUNE productions.
- The reference systematics are from the TDR with a different GENIE tune.
- Some changes have to be made to adapt the dials to new tune.
- This developpment is still ongoing for LBL xsec systematics \to I was pointed to the already available list that I'll test soon.

Thinking about atmospherics-specific xsec systematics

- The current LBL xsec systs have no angular dependency. List here
- Need to see with DIRT people what could/should be implemented.
- Significant amount of work to implement.

Adding systematics



FLUX SYSTEMATICS

Implementation in MaCh3

- A new method to set systematics using yaml files was implemented in MaCh3
- Plan to use it very soon to implement first basic flux systematics

Proposed list of flux systematics

- Absolute flux normalization
- Relative normalization of $(\nu_{\mu} + \bar{\nu}_{\mu})/(\nu_{e} + \bar{\nu}_{e})$ ratio
- $\nu/\bar{\nu}$ normalization (flux models differences)
- Zenith angle normalization
- K/π ratio
- Solar activity (impacts geomagnetic field)
- Atmosphere density (impacts production height)

Using physical simulations

Looked at MCEq. They don't implement geomagnetic field which is very important at low E. Is anyone aware of an alternative?



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