

## Colors and spins of asteroids in LSST era



**B. Carry<sup>1</sup>, J. Peloton<sup>2</sup>, R. Le Montagner<sup>2</sup>,  
M. Mahlke<sup>3</sup>, J. Berthier<sup>4</sup>**

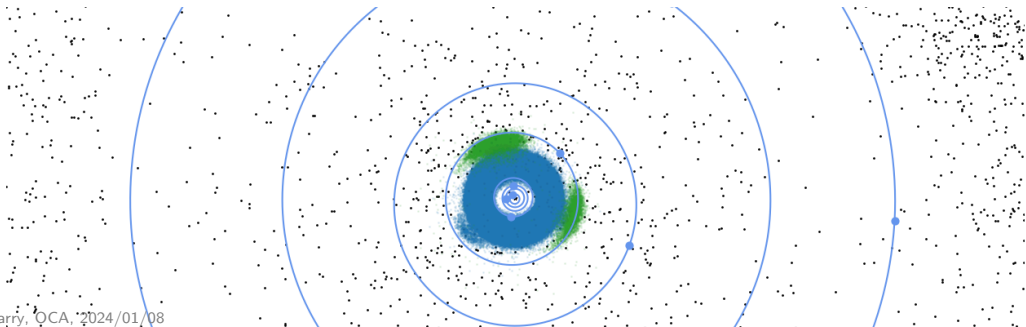
<sup>1</sup>Lagrange, <sup>2</sup>IJCLab, <sup>3</sup>IAS, <sup>4</sup>IMCCE

# Small bodies and planetary formation

- **Leftovers of the early Solar System**
  - Remnants of planet building blocks
  - Limited dynamical & mineralogical evolution
- **Constraints on planetary formation & evolution**
  - Orbital and size distributions
  - Distribution of composition

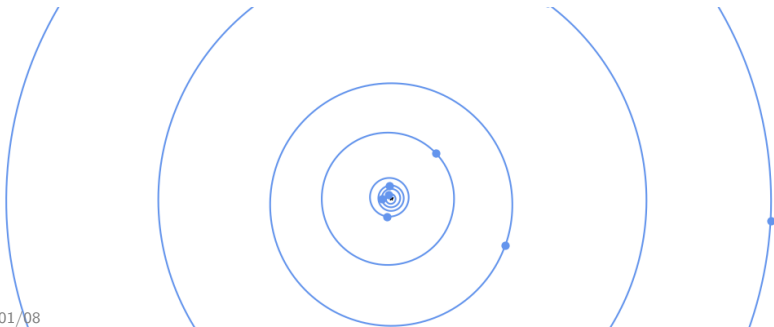
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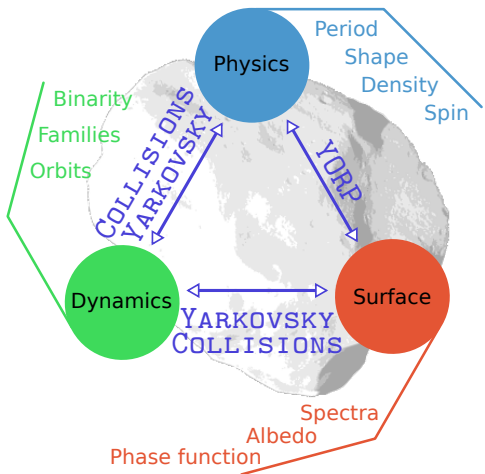


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# What do we need to study?



## • Discovery & Dynamics

- Dynamical structure
- Origins & evolution
- ▷ **Astrometry**

## • Composition

- Location & timing of formation
- Compositional structure
- ▷ **Vis-NIR spectro-photometry**

## • Physical properties

- Diameter, Spin, ...
- Main evolutionary drivers
- ▷ **Time serie photometry**

# What LSST will bring to Solar System?



Coming Soon

We are working on website very hard. Estimated remaining time is:

00

days

00

hours

00

minutes

00

seconds

Mail me when site is ready:

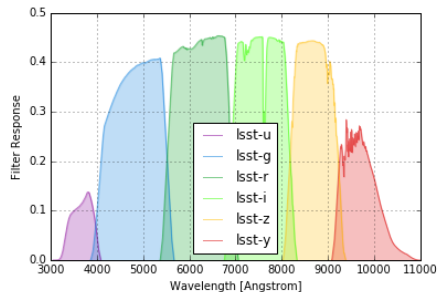
Name:

E-mail address:

# What LSST will bring to Solar System?

	Currently Known	LSST Discoveries	Typical number of observations
Near Earth Objects (NEOs)	~20,000	200,000	(D>250m) 60
Main Belt Asteroids (MBAs)	~650,000	6,000,000	(D>500m) 200
Jupiter Trojans	~7000	280,000	(D>2km) 300
TransNeptunian Objects (TNOs) + Scattered Disk Objects (SDOs)	~3000	40,000	(D>200km) 450
Comets	~3000	10,000	?
Interstellar Objects (ISOs)	2	10	?

LSST SSSC



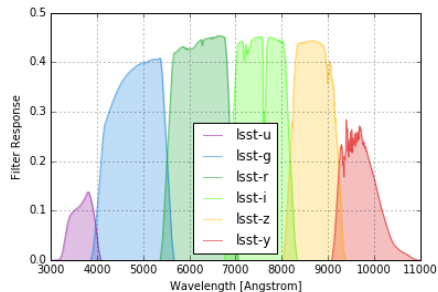
LSST



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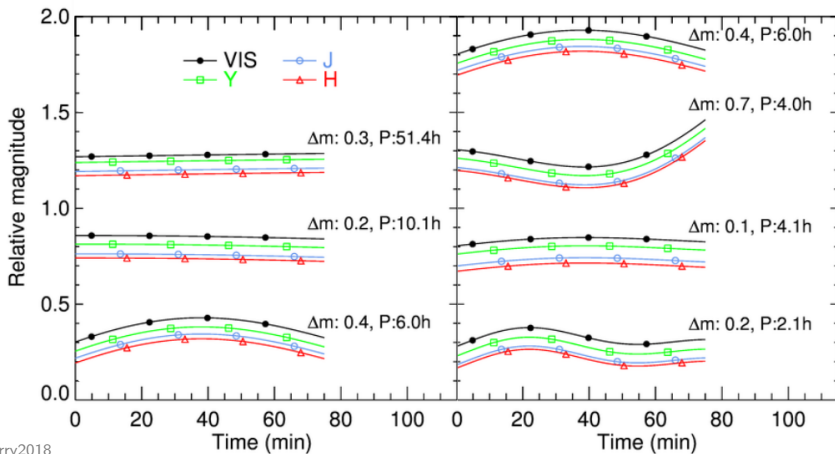
LSST SSSC



LSST

⇒ **Discovery, astrometry, colors, time series**

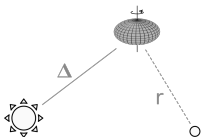
# The challenge of SSO variability



Carry2018

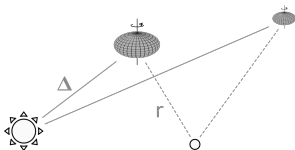
# Modeling the photometry of SSOs

- **Distance**



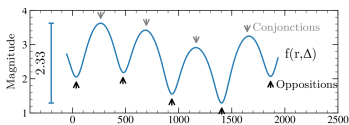
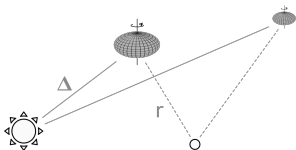
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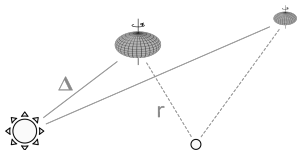


○ HG Bowell1989

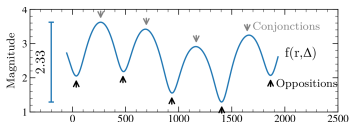
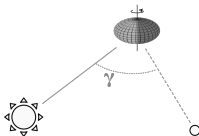
$$H = m - f(r, \Delta)$$

# Modeling the photometry of SSOs

## • Distance



## • Phase

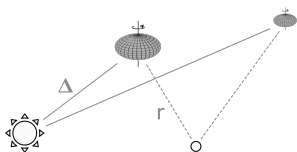


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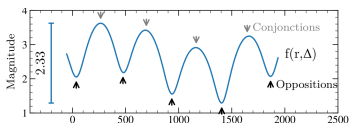
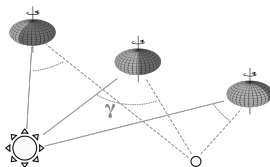
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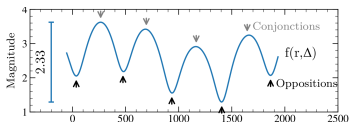
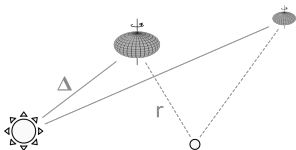


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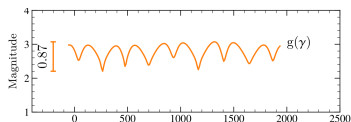
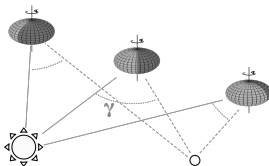
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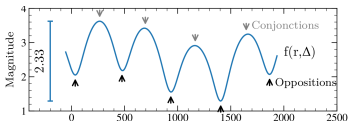
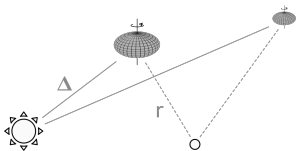
○ HG<sub>1</sub>G<sub>2</sub> Muinenon+2010

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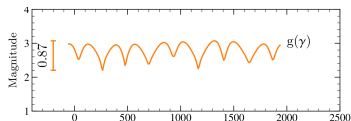
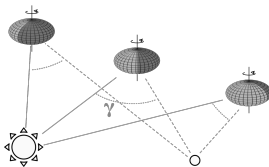
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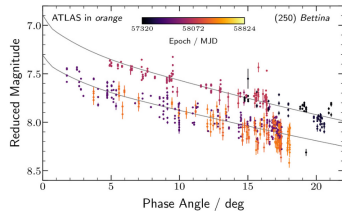


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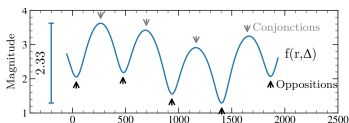
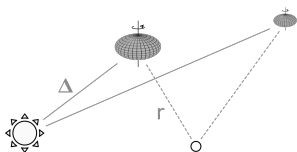
## • Aspect!



Mahlke+2021

# Modeling the photometry of SSOs

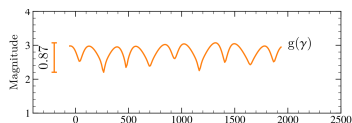
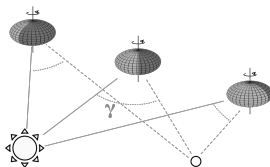
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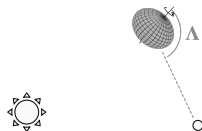


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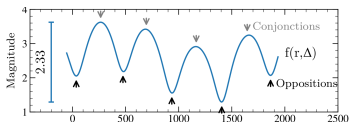
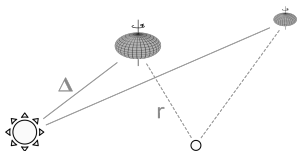
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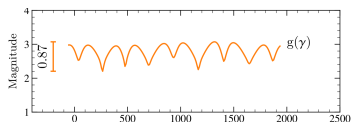
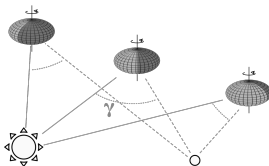
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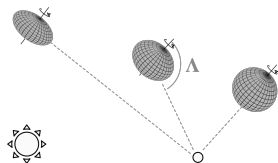


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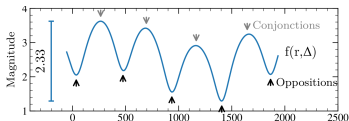
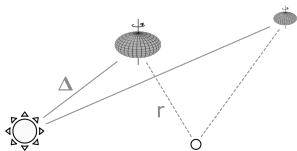
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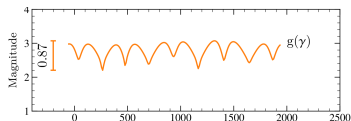
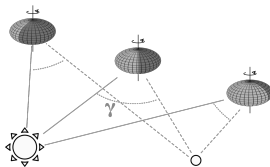
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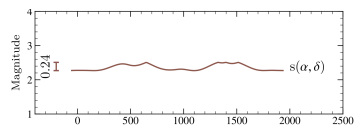
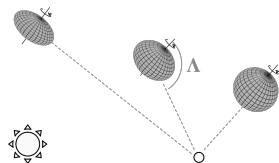


○ HG Bowell1989

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$$H = m - f(r, \Delta) - g(\gamma)$$

## • Aspect



○ sHG<sub>1</sub>G<sub>2</sub> Carry+2024

$$H = m - f(r, \Delta) - g(\gamma) - s(\alpha, \delta)$$

# Data

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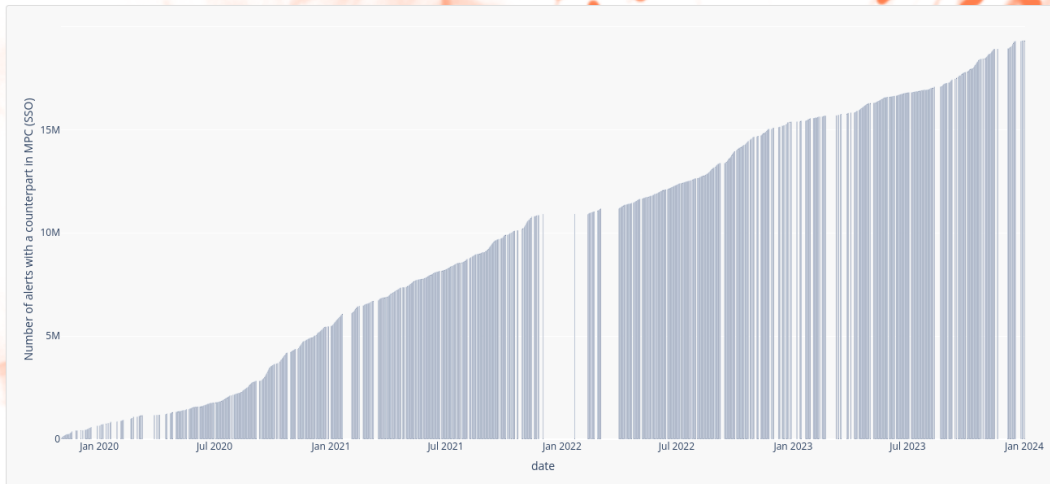
# Data: FINK!

Number of alerts with a counterpart in MPC (SSO)

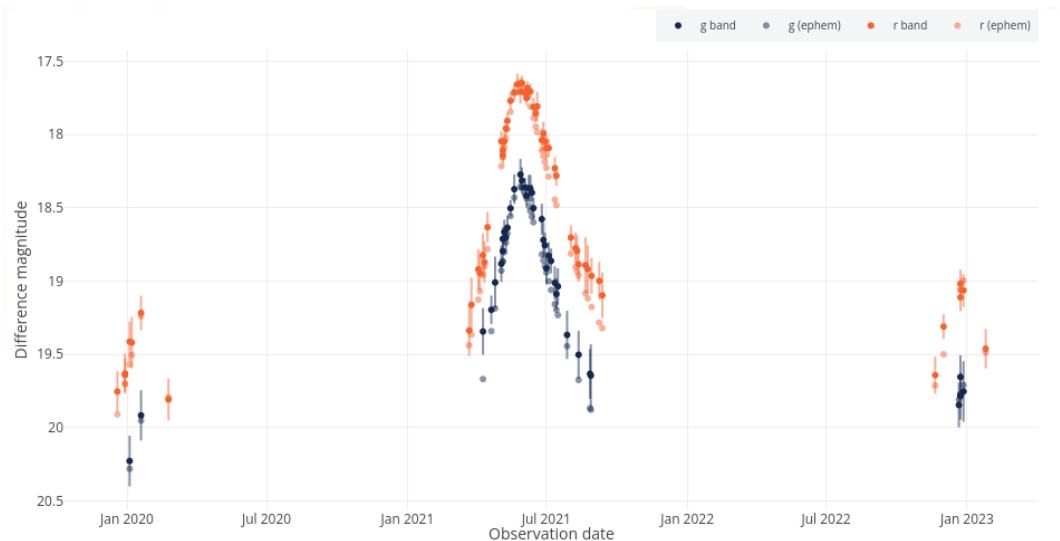


Cumulative

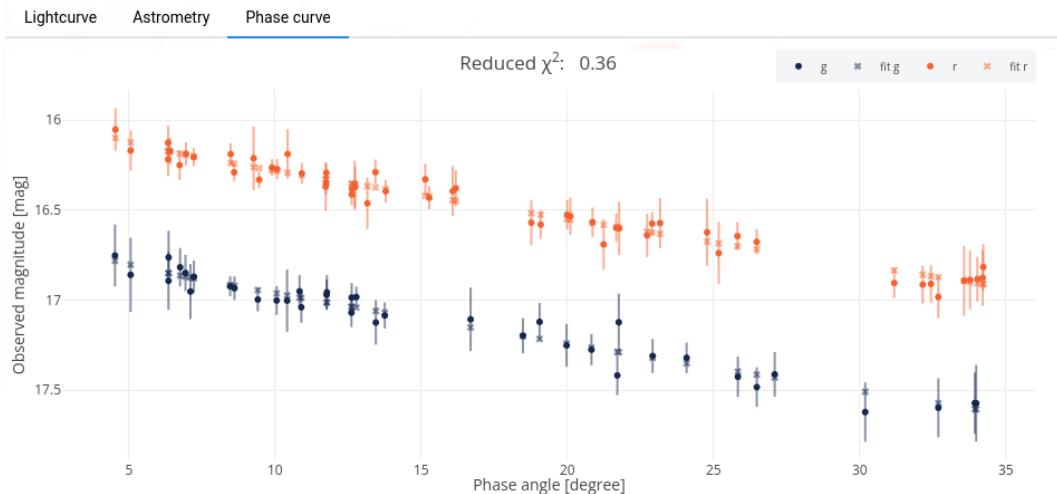
Percentage



# Data: FINK!... Julienpeloton

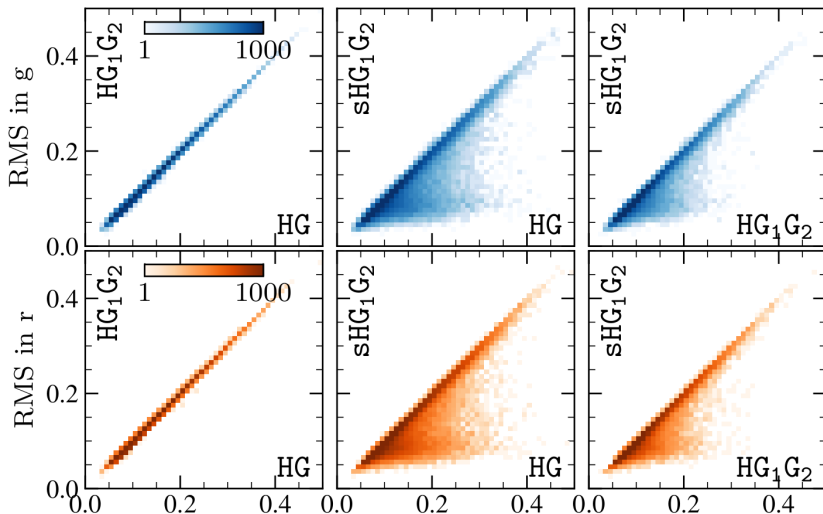


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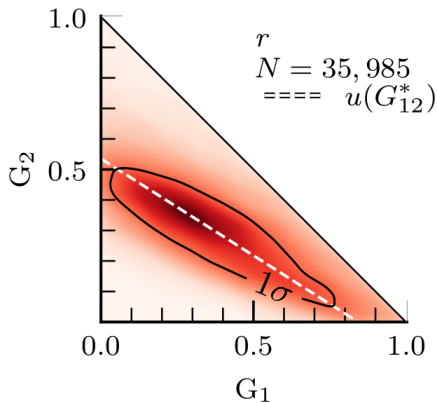
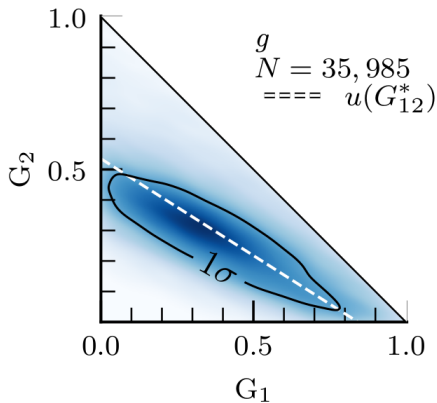




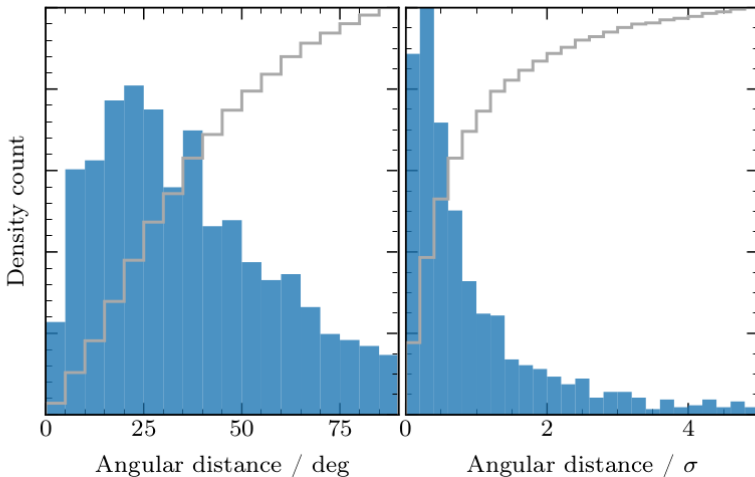
# Validation: Fit



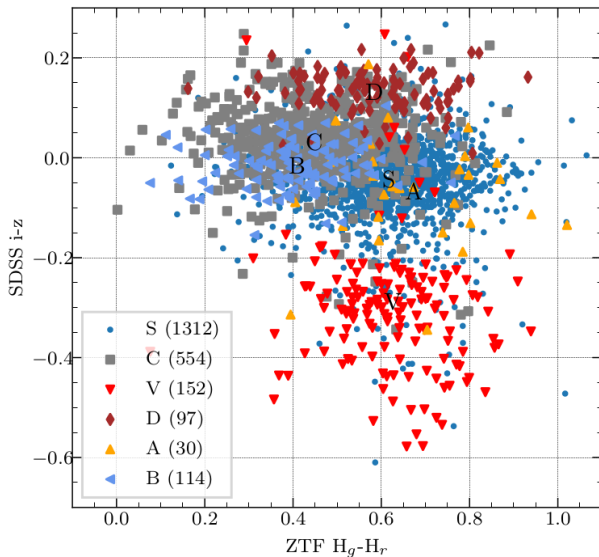
# Validation of phase: $g(\gamma)$



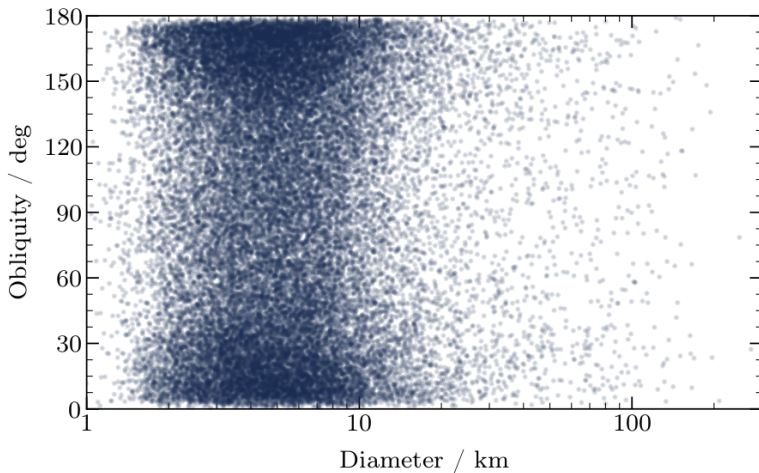
# Validation of spins: $s(\alpha, \delta)$



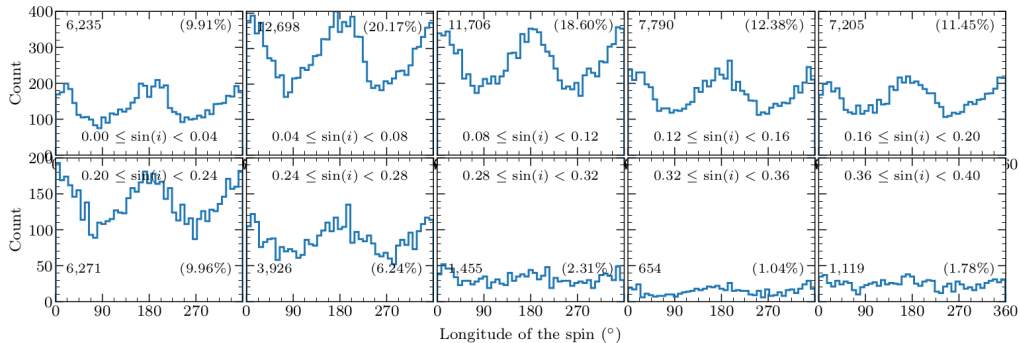
# Validation of colors: sHG<sub>1</sub>G<sub>2</sub>



# Spin orientation



# Spin orientation



# Summary

- **Questions on Solar System & Planetary formation require**

- Discoveries for large statistics on SSOs
- Including multi-filter and time-serie photometry
- ▶ LSST of Vera C. Rubin is **highly** promising

- **Challenges linked with SSO photometry**

- Combination of short- and long-term variability
- ▶ New model put forward here: sHG<sub>1</sub>G<sub>2</sub>

- **sHG<sub>1</sub>G<sub>2</sub> implemented in FINK over 2023**

- Run monthly → Dec.2023 → 115,000 SSOs
- sHG<sub>1</sub>G<sub>2</sub> works and improves over previous models
- Phase & spin parameters & *g-r* color for ≈50,000 SSOs
- ▶ Ready for LSST!