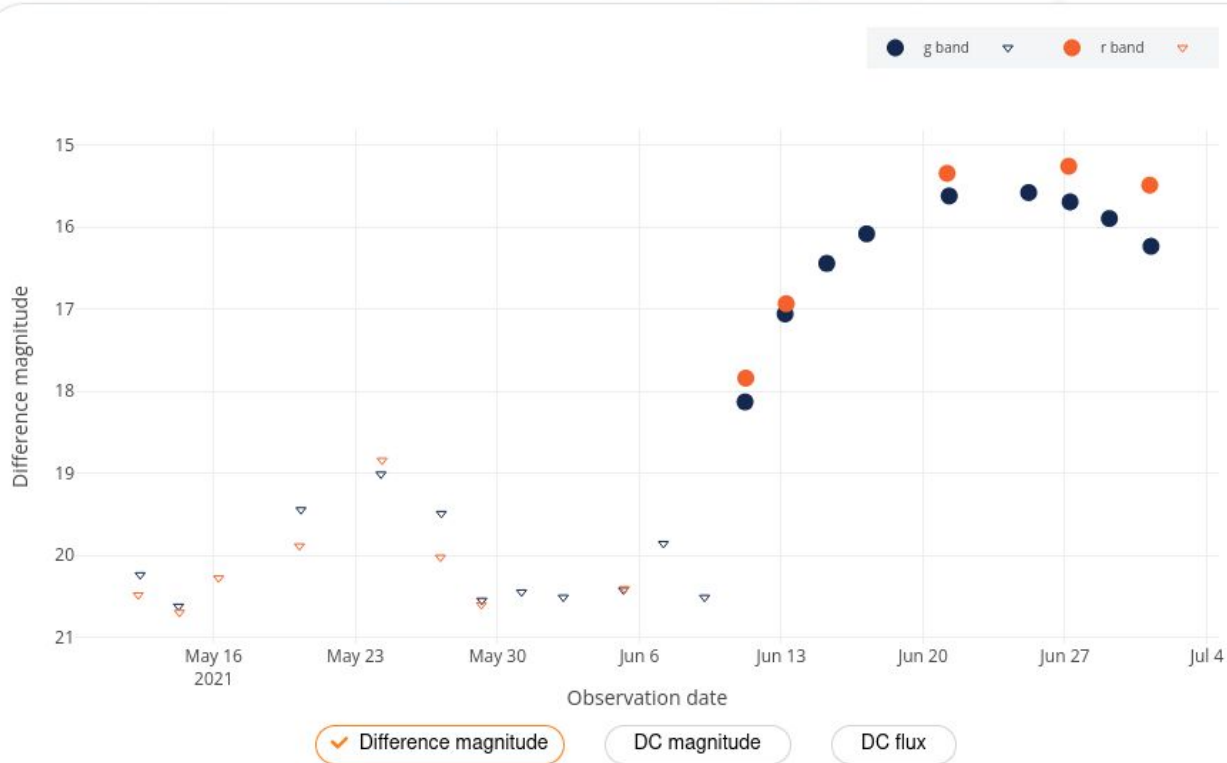


Rainbow

A colorful approach on
multi-passband light curve estimation

E. Russeil, K. L. Malanchev, P. D. Aleo, E. E. O. Ishida, M. V. Pruzhinskaya, E. Gangler, A. D. Lavrukhina, A. A. Volnova, A. Voloshina, T. Semenikhin, S. Sreejith, M. V. Kornilov, V. S. Korolev (The SNAD team)

Light curve characterisation

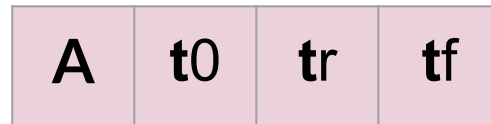
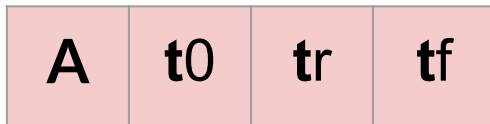
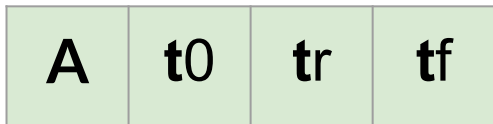
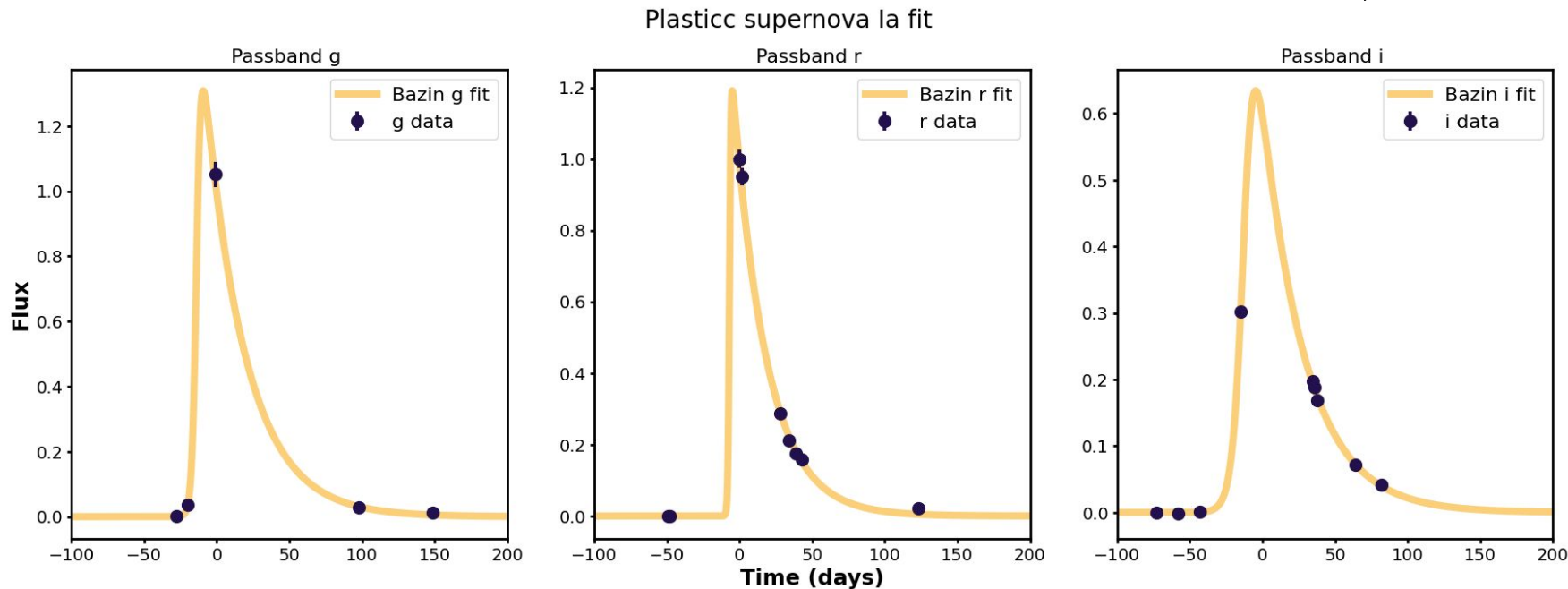


Why produce a fit ?

- Understanding of the object
- Interpolation of missing points
- Prediction of the evolution
- Machine learning analysis

Transient feature extraction :

$$F_{\nu}(t) = A \times \frac{e^{-\frac{(t-t_0)}{t_{fall}}}}{1 + e^{\frac{t-t_0}{t_{rise}}}}$$



If one band is under sampled the feature extraction is impossible



Number of parameters scale with the number of filters of the telescope

Parameters from different bands are correlated



What is Rainbow

$$F_\nu(t, \nu) = \frac{\pi}{\sigma_{SB}} \times \frac{B_\nu(T, \nu)}{T(t)^4} \times F_{bol}(t)$$

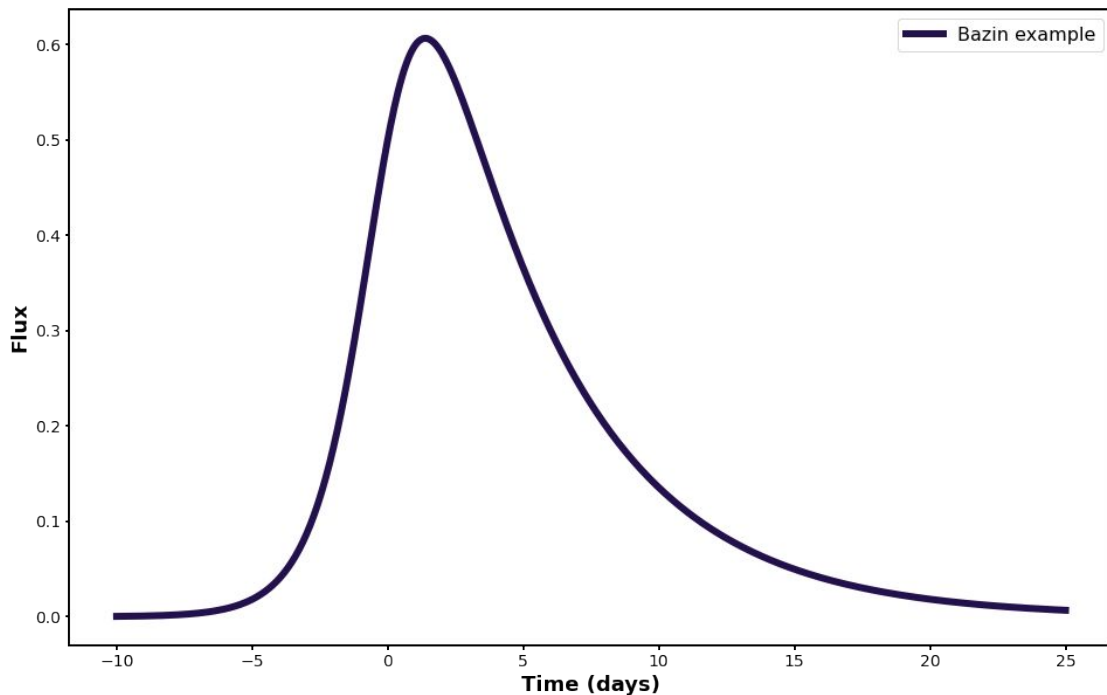
Theoretical solution (assuming that the object is a black body)

$$F_{\nu}(t, \nu) = \frac{\pi}{\sigma_{SB}} \times \frac{B_{\nu}(T, \nu)}{T(t)^4} \times F_{bol}(t)$$

We must decide what to use for
T(t) and **Fbol(t)** ?

Can be adapted to each science case

$$F_\nu(t, \nu) = \frac{\pi}{\sigma_{SB}} \times \frac{B_\nu(T, \nu)}{T(t)^4} \times \boxed{F_{bol}(t)}$$

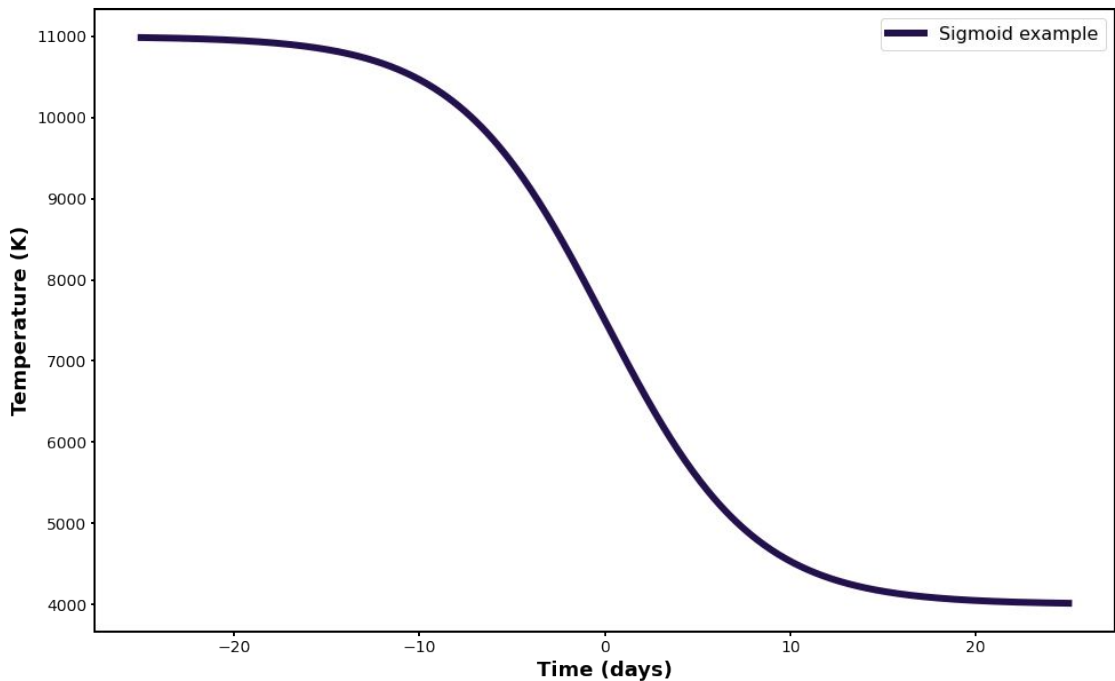


$$F_{bol}(t) = A \times \frac{e^{-\frac{(t-t_0)}{t_{fall}}}}{1 + e^{\frac{t-t_0}{t_{rise}}}}$$

Bazin function : 4 parameters
(here we assume that the baseline is 0)

A = 1
t0 = 0
t_{fall} = 5
t_{rise} = -1

$$F_{\nu}(t, \nu) = \frac{\pi}{\sigma_{SB}} \times \frac{B_{\nu}(T, \nu)}{T(t)^4} \times F_{bol}(t)$$



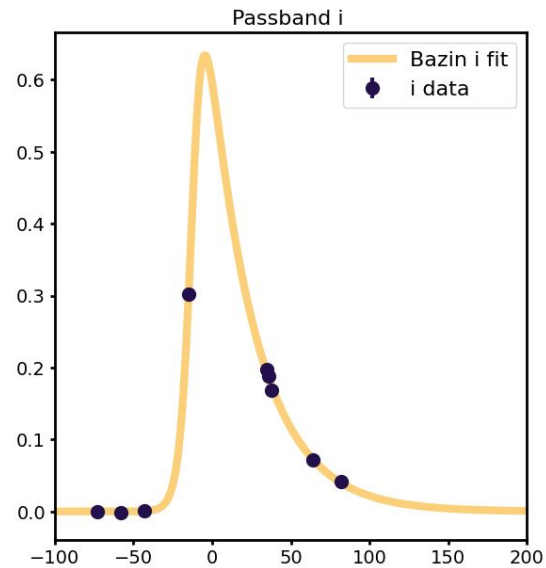
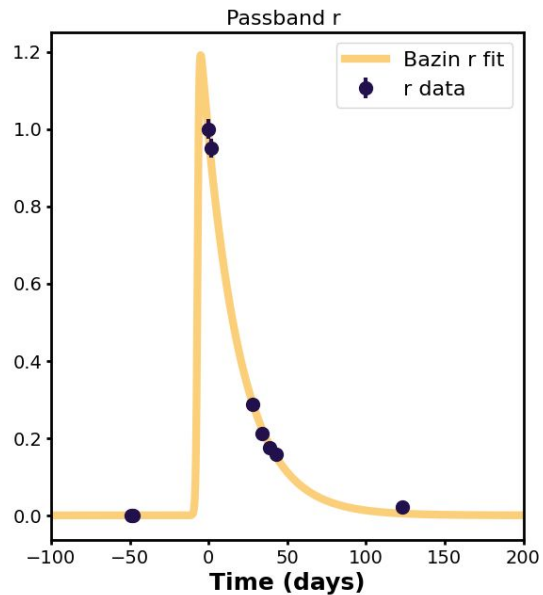
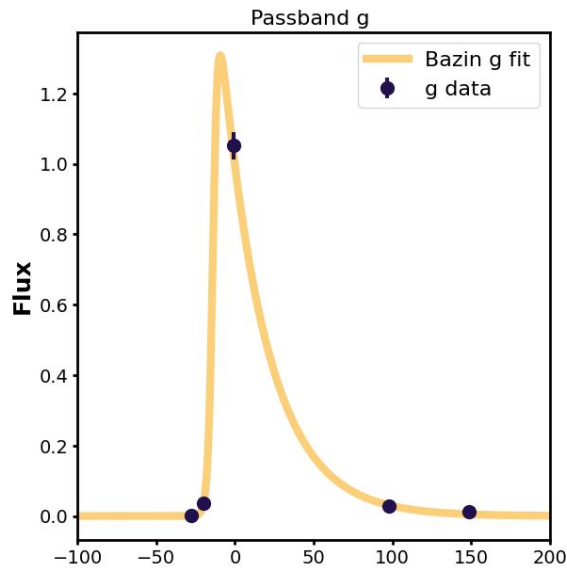
$$T(t) = T_{min} + \frac{T_{var}}{1 + e^{\frac{t-t_0}{k_{sig}}}}$$

Sigmoid function : 3 parameters
(t0 is common with Bazin)

- Tmin = 4000**
- Tvar = 7000**
- ksig = 4**
- (t0 = 0)**

Rainbow fit examples

Plasticc supernova la fit



A

t0

tr

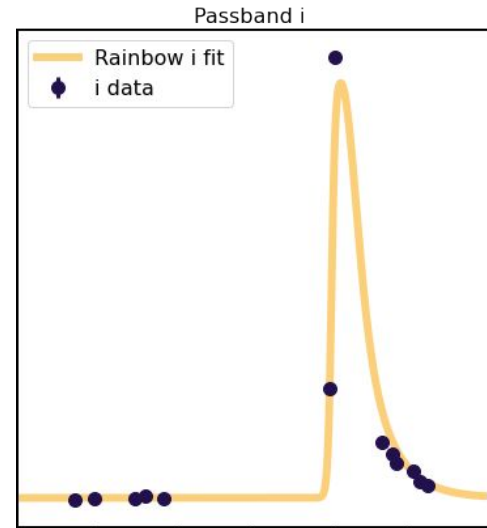
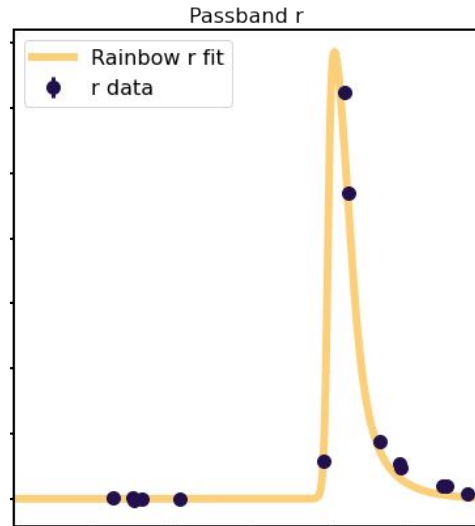
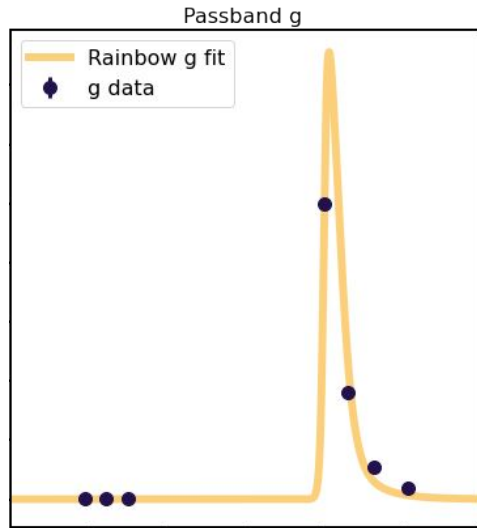
tf

Tvar

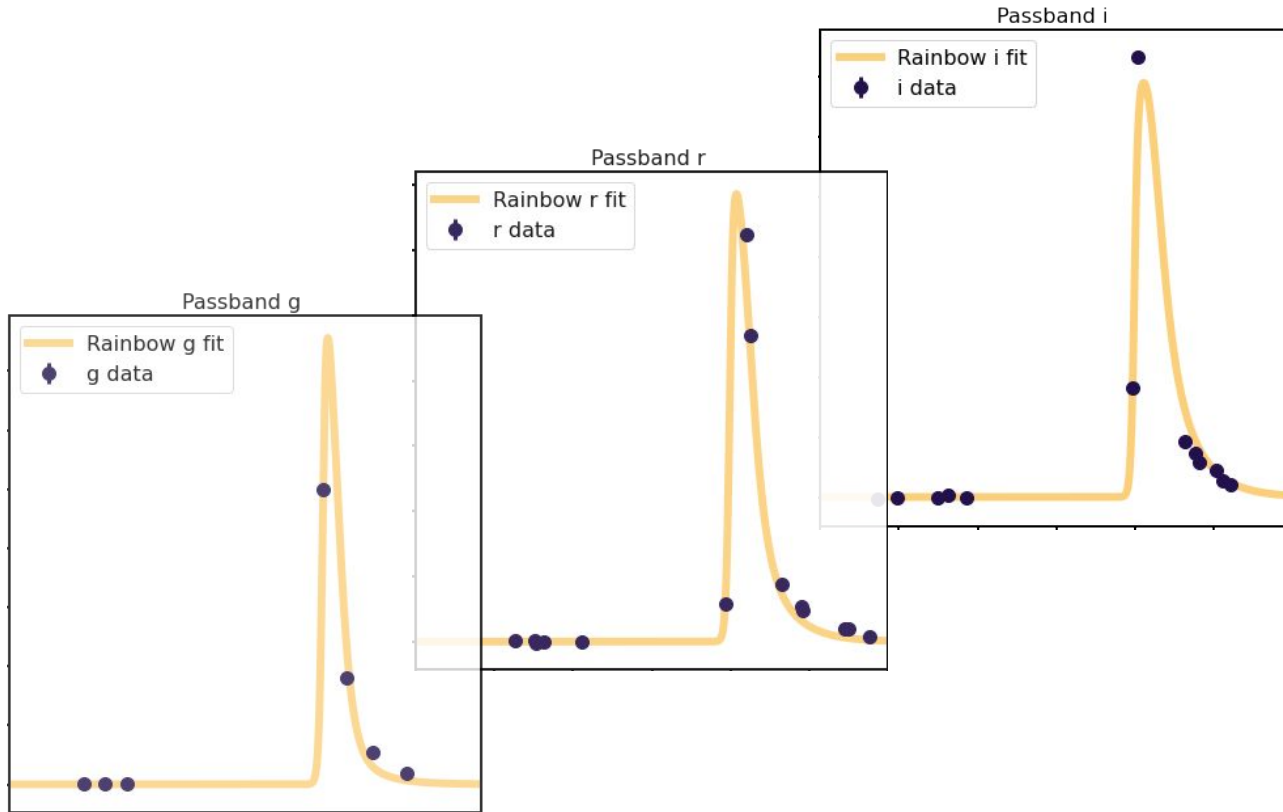
Tmin

ksig

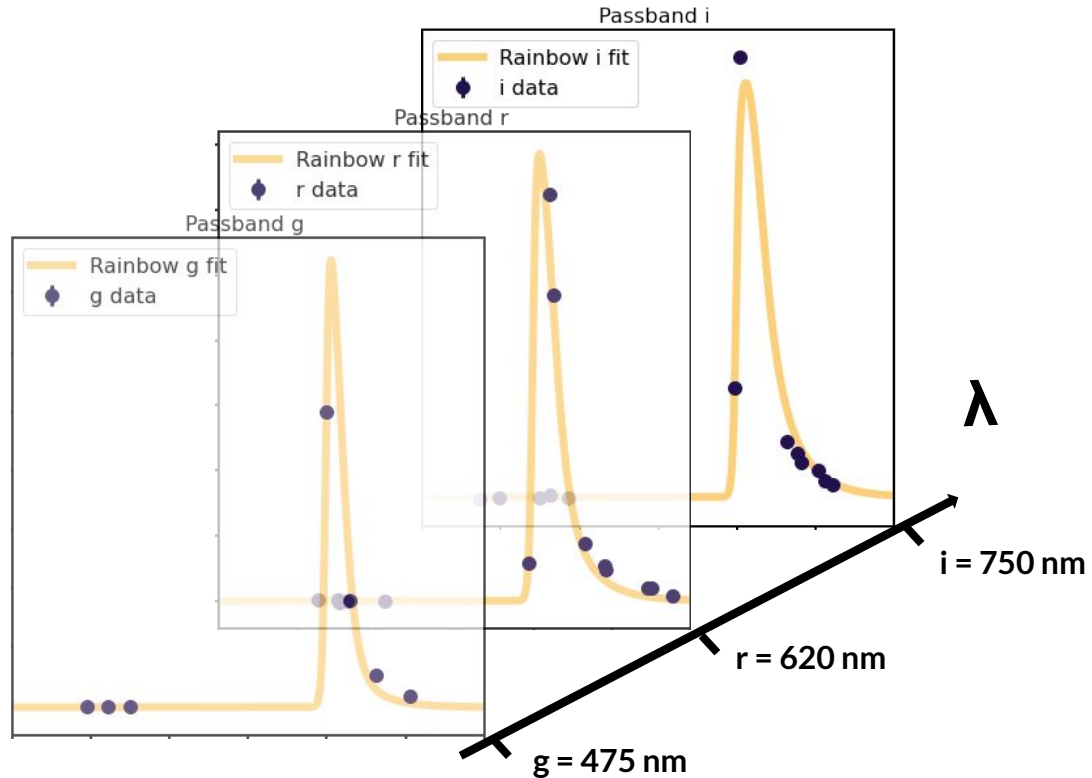
3D visualisation



3D visualisation

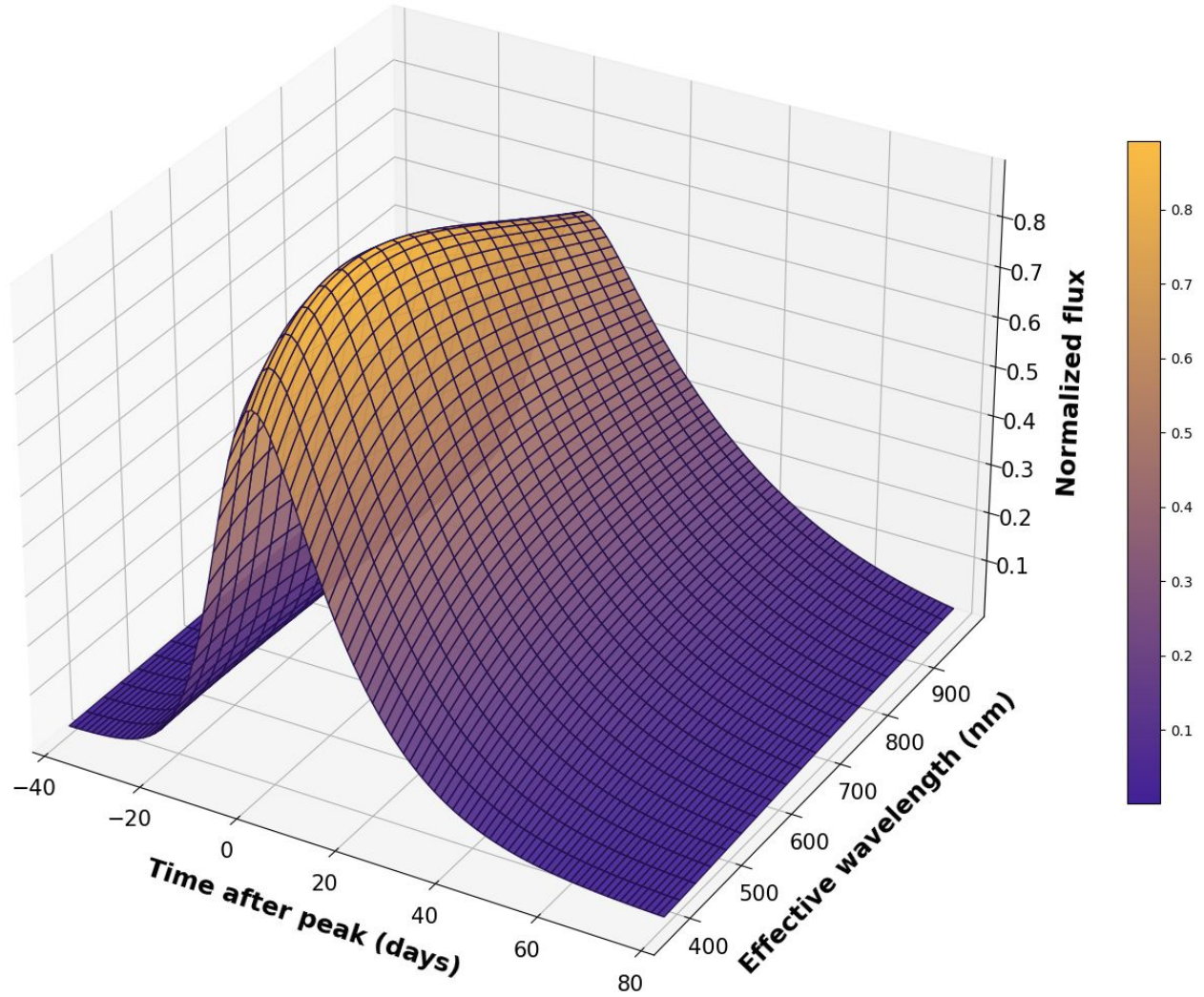


3D visualisation



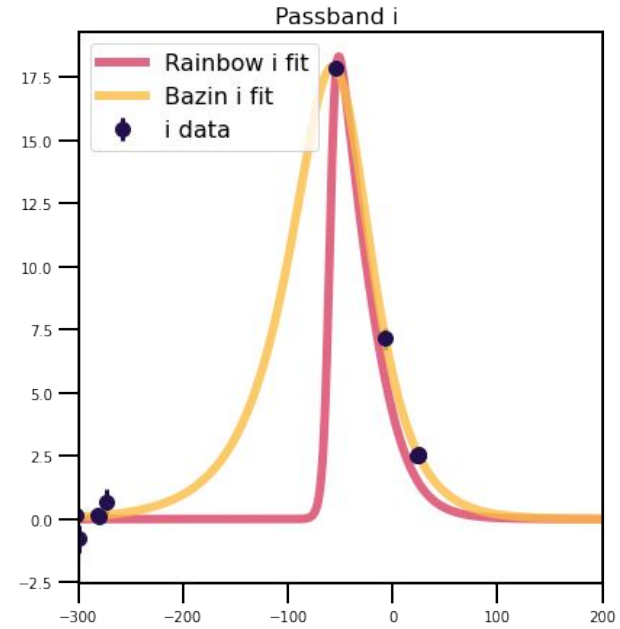
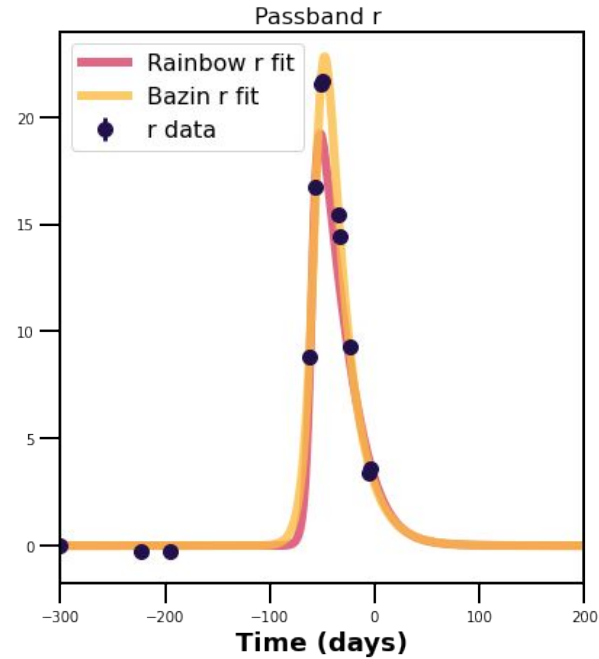
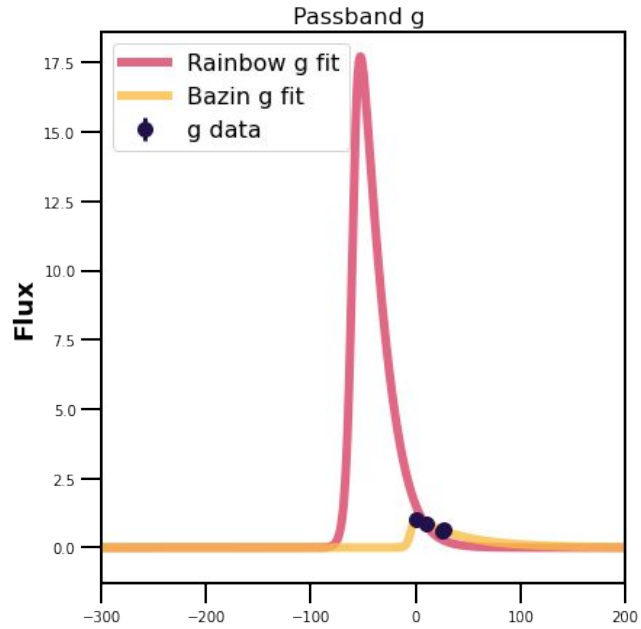
3D visualisation

RAINBOW:



Typical behavior

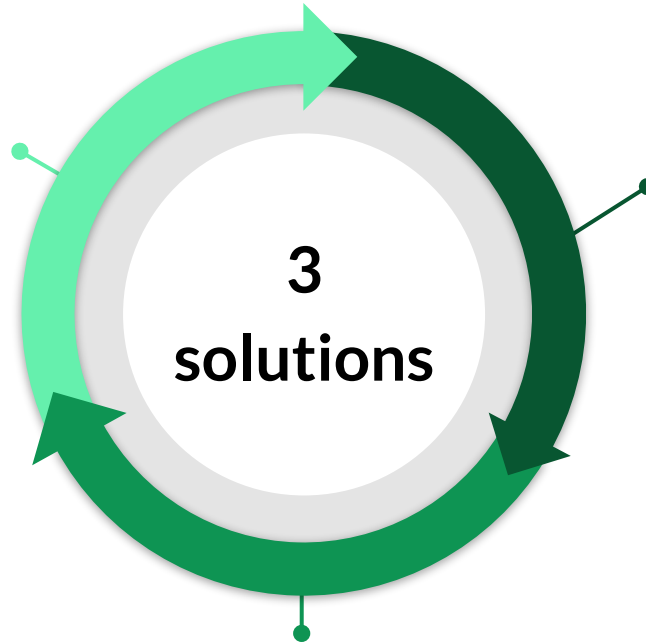
Plasticc supernova Ia fit





Rainbow efficiency

Consider all points available at once. An undersample passband doesn't matter



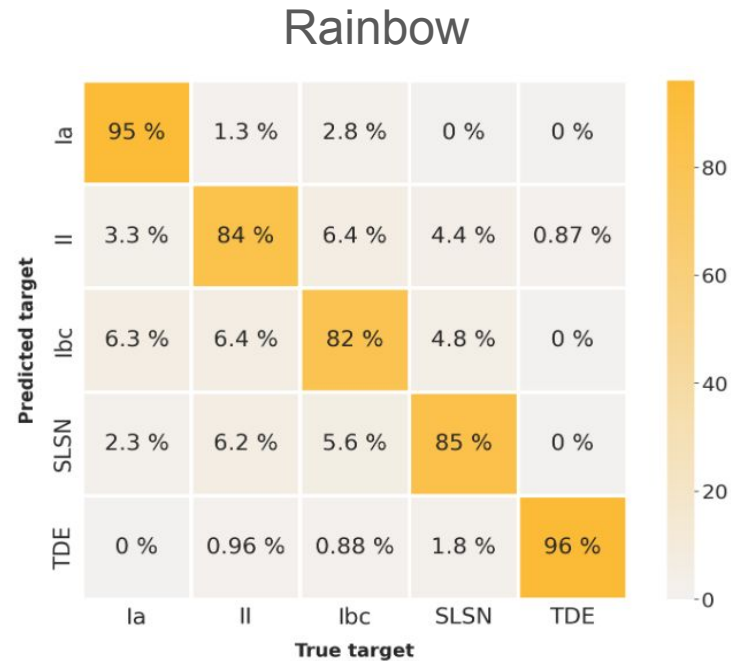
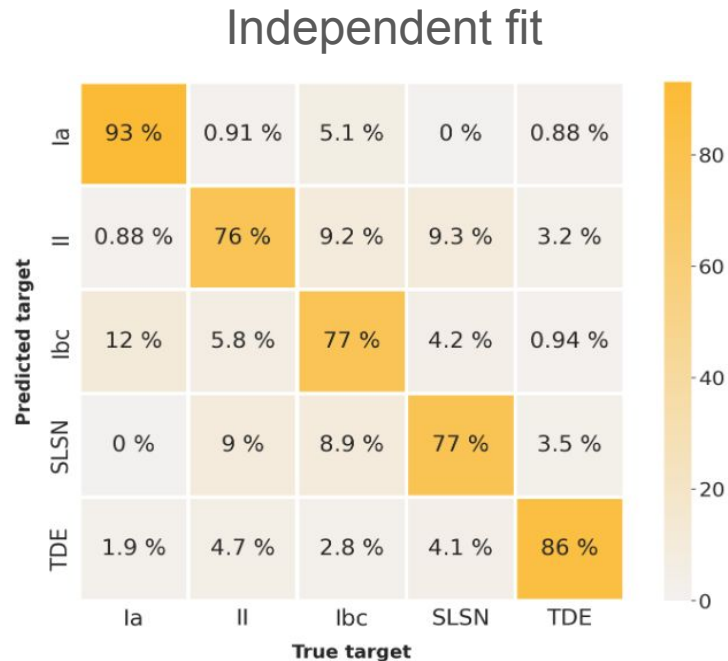
Number of parameters is constant independently of the number of filters

We obtain only one set of parameters : get rid of the correlations

Paper conclusion : simpler and better !

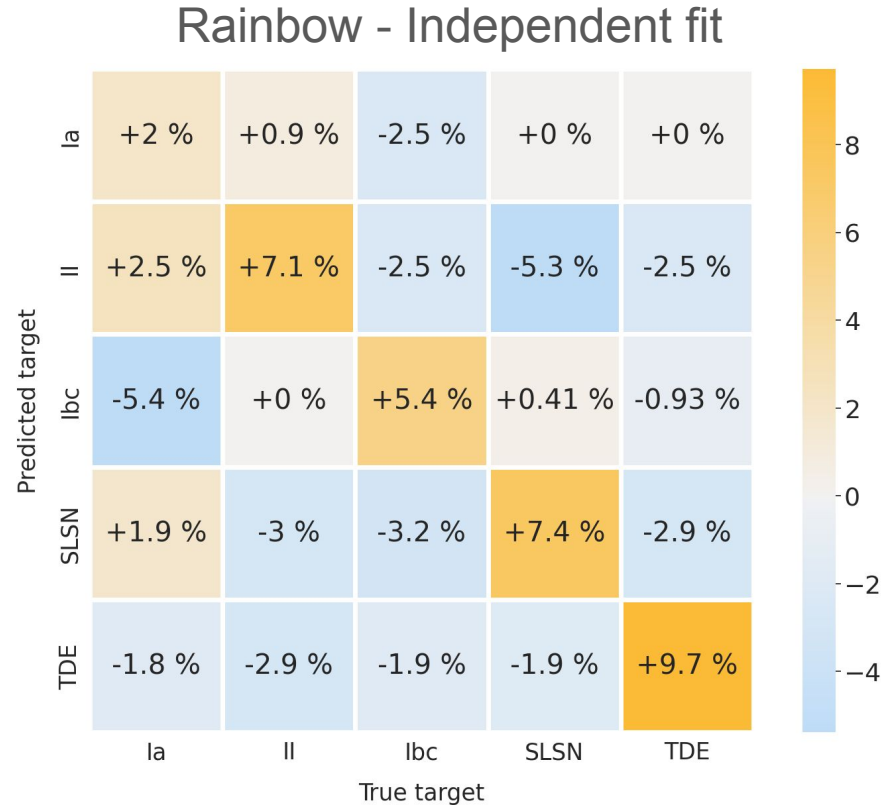


Paper teaser: PLAsTiCC transient classification



Balanced dataset (300 objects per class)

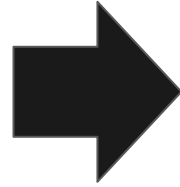
Paper teaser: PLAsTiCC transient classification





Fink applications

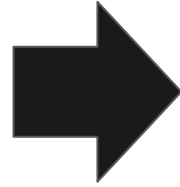
Science modules



- **Active galactic nuclei**
- **Anomaly detection**
- **Early SNIa**
- **Super Luminous Supernovae**
- **Tidal disruption event**
- *More to come ?*

All require a parametric fit at some point

Science modules



- Active galactic nuclei
- Anomaly detection
- Early SNIa
- Super Luminous Supernovae
- Tidal disruption event
- *More to come ?*

All require a parametric fit at some point

**Need to adapt from 2 to 6 passbands with a low
lower cadence for LSST !**

Current work using Rainbow

Early SNIa

- Fit only rising light curves
- Details to be presented by Emille

Early TDE

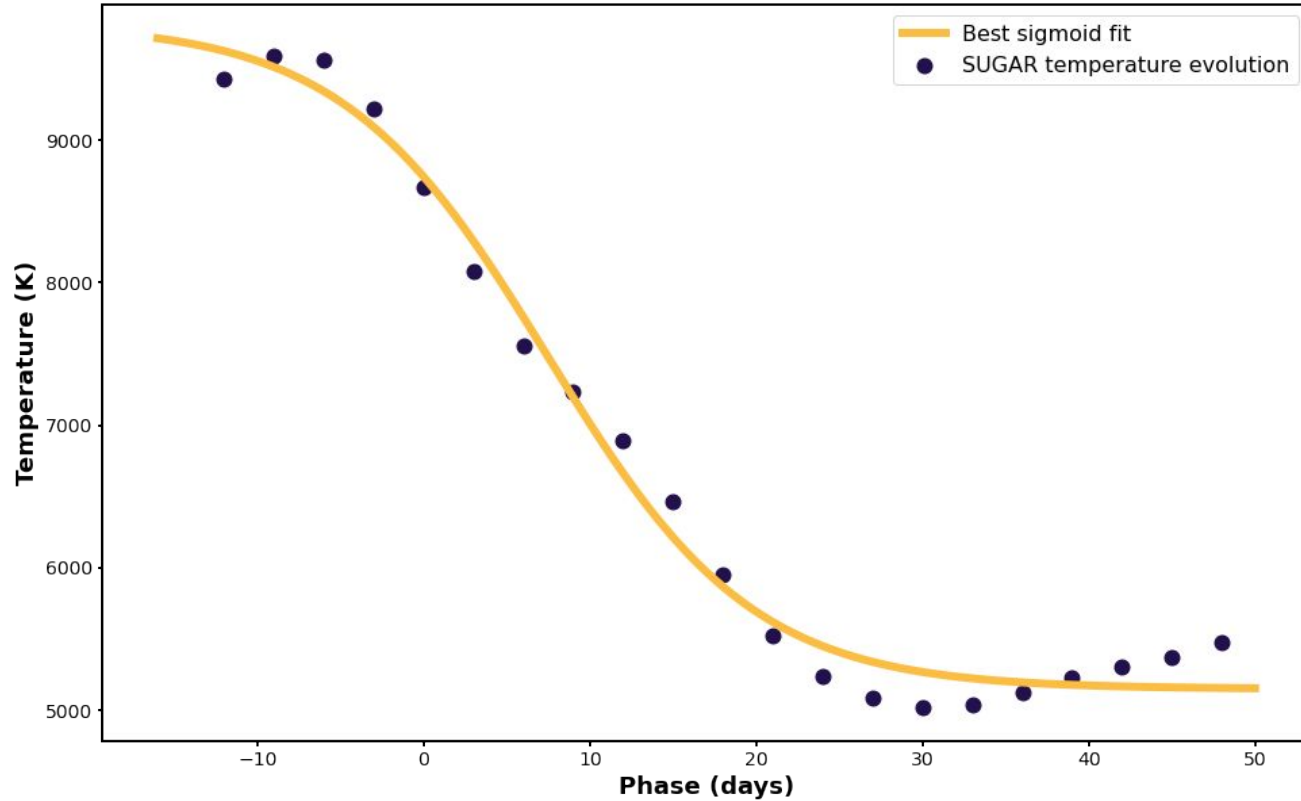
- Fit only rising light curves
- Use a constant temperature
- Work in progress



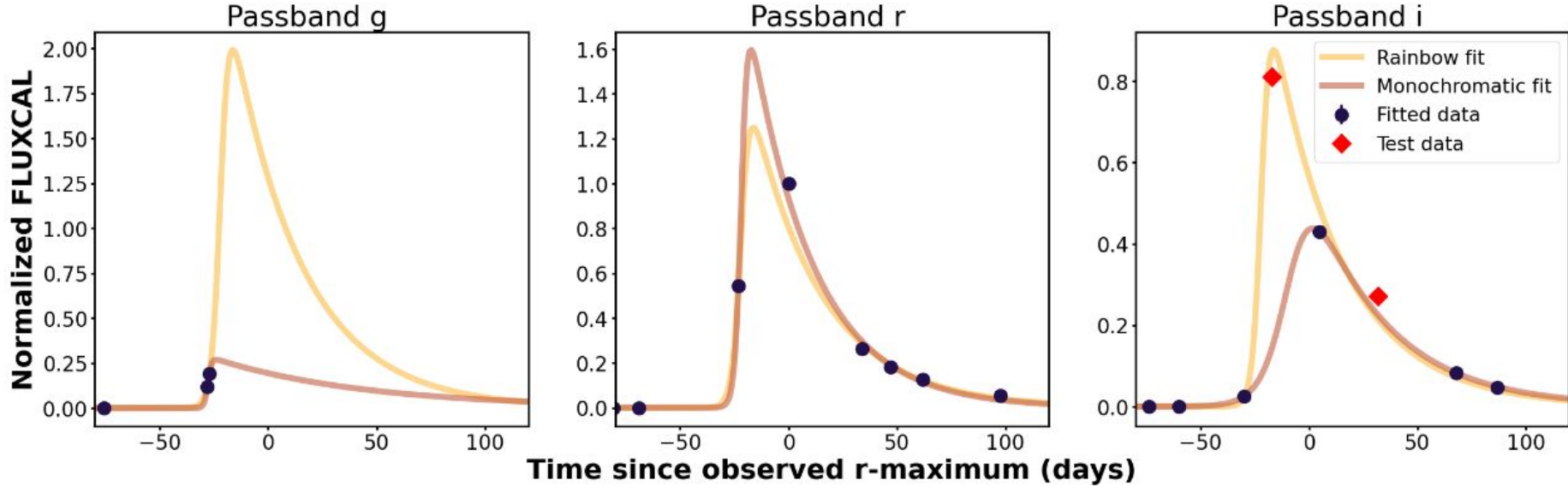
Conclusion

- Rainbow is displaying excellent results, see [paper](#)
- Offers a good transition from ZTF to LSST
- An implementation is available for easy use here:
<https://github.com/light-curve/light-curve-python>
- Adaptable to future science cases ! Need to choose bolometric flux and temperature evolution
- If you are interested in using Rainbow don't hesitate to contact me

Why sigmoid ?

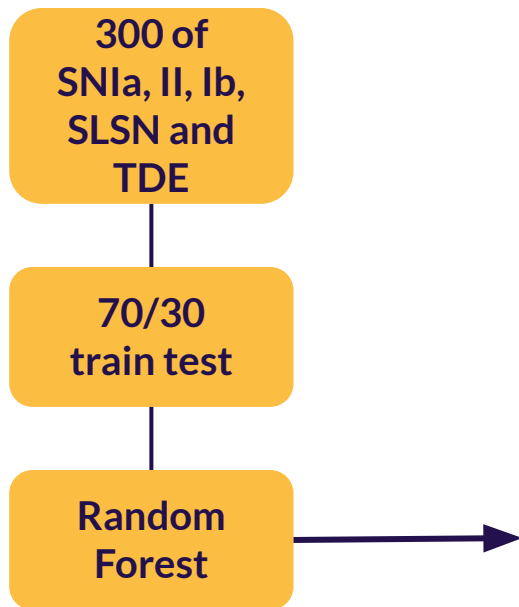


Goodness of fit : random points removal

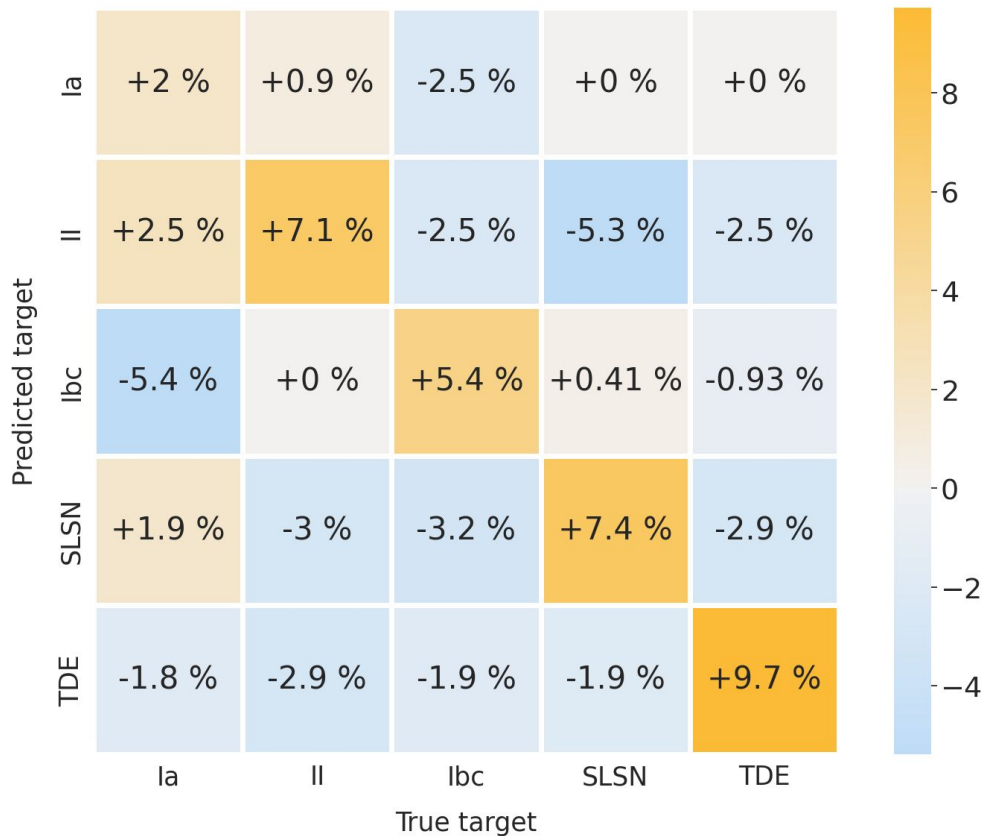


$$\text{nRMSEo} = \sqrt{\frac{1}{m} \sum_i \left[\frac{(y_i - \mu(t_i))^2}{2\epsilon_i^2} \right]}$$

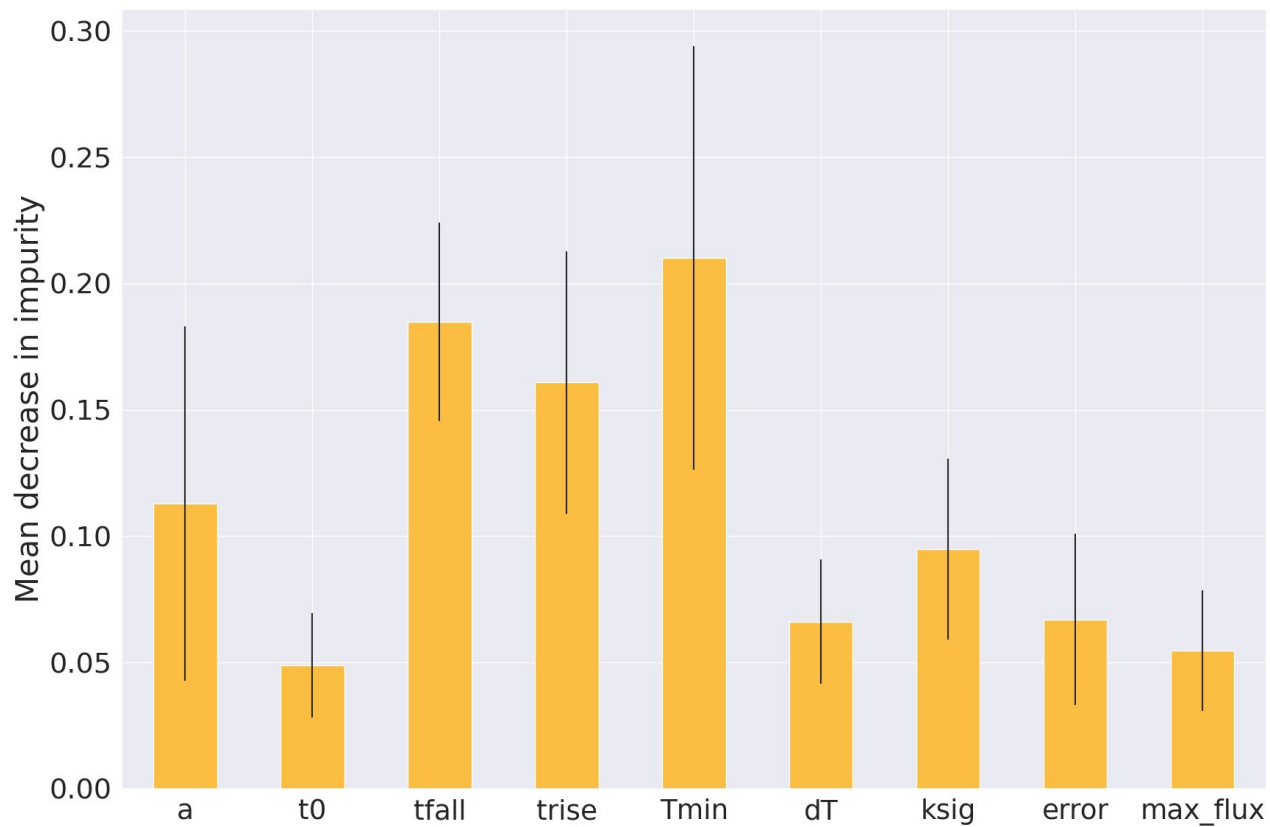
Classification task



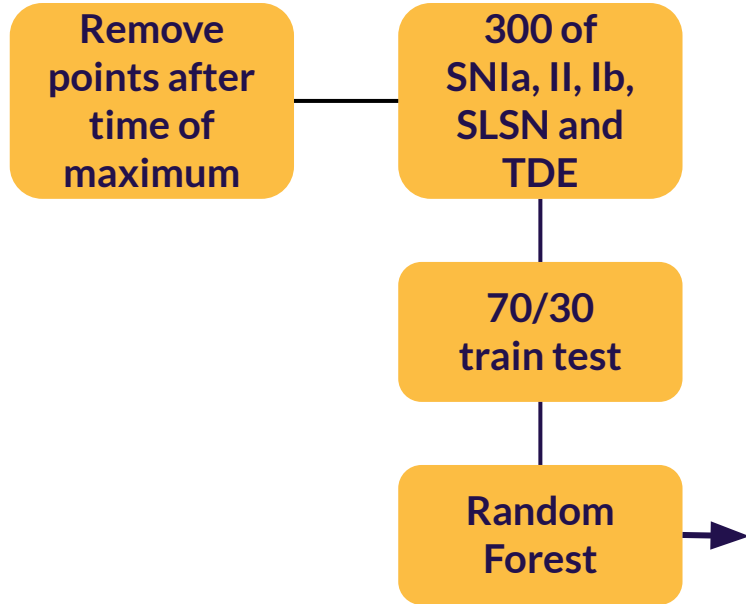
Rainbow - Monochromatic



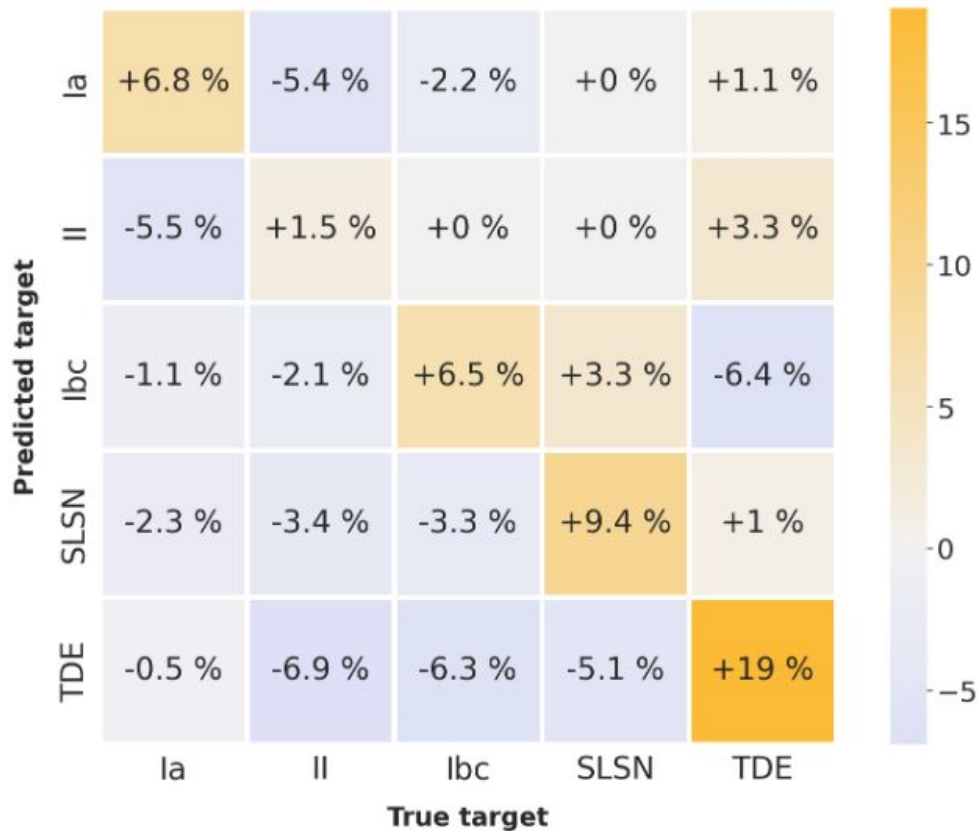
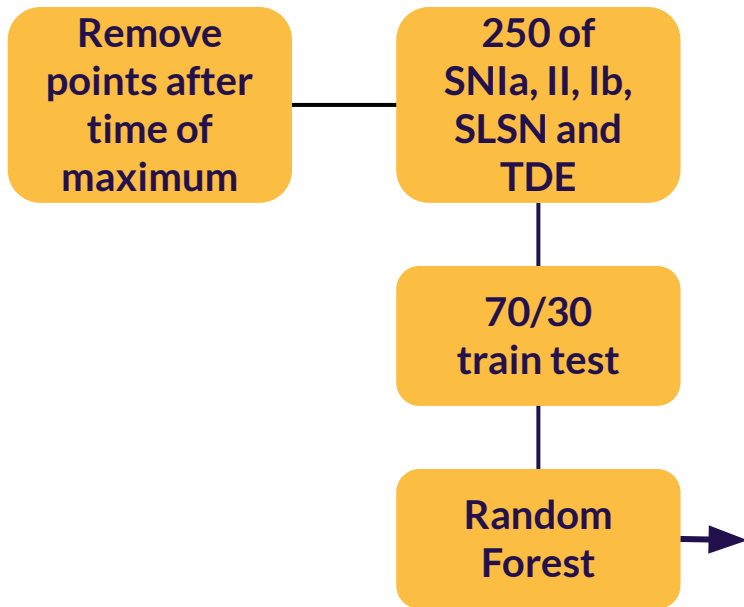
Classification task



Classification rising light curves



Classification rising light curves



$$F_\nu(t, \nu) = \frac{\pi}{\sigma_{SB}} \times \frac{B_\nu(T, \nu)}{T(t)^4} \times F_{bol}(t)$$

Planck's law

$$B_\nu(\nu, T) = \frac{2h\nu^3}{c^2} \times \frac{1}{\exp\left(\frac{h\nu}{k_B T}\right) - 1}$$

