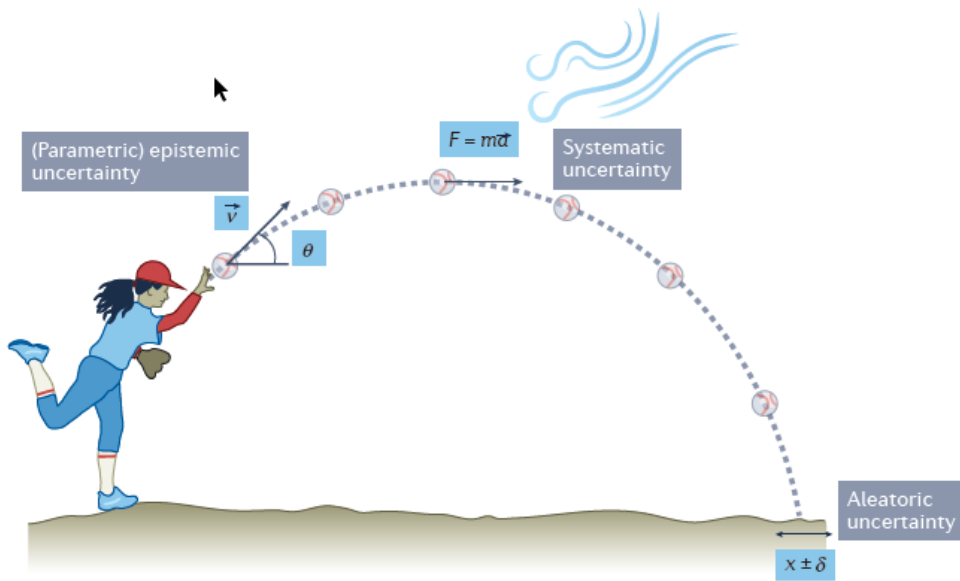


# An introduction to simulation-based inference

AI and the uncertainty challenge in fundamental physics  
November 28, 2023

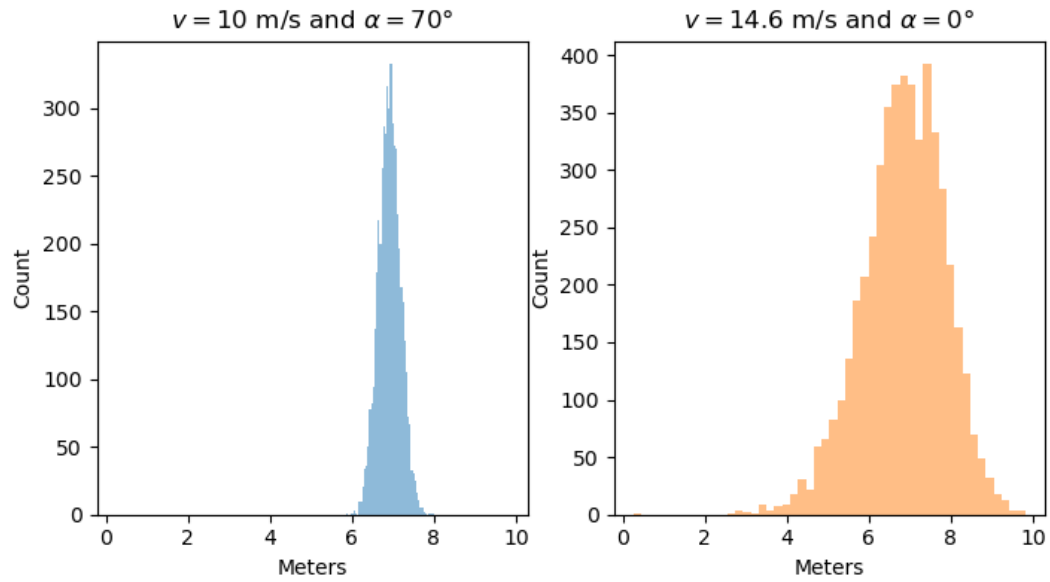
Gilles Louppe  
g.louppe@uliege.be  
@glouppe



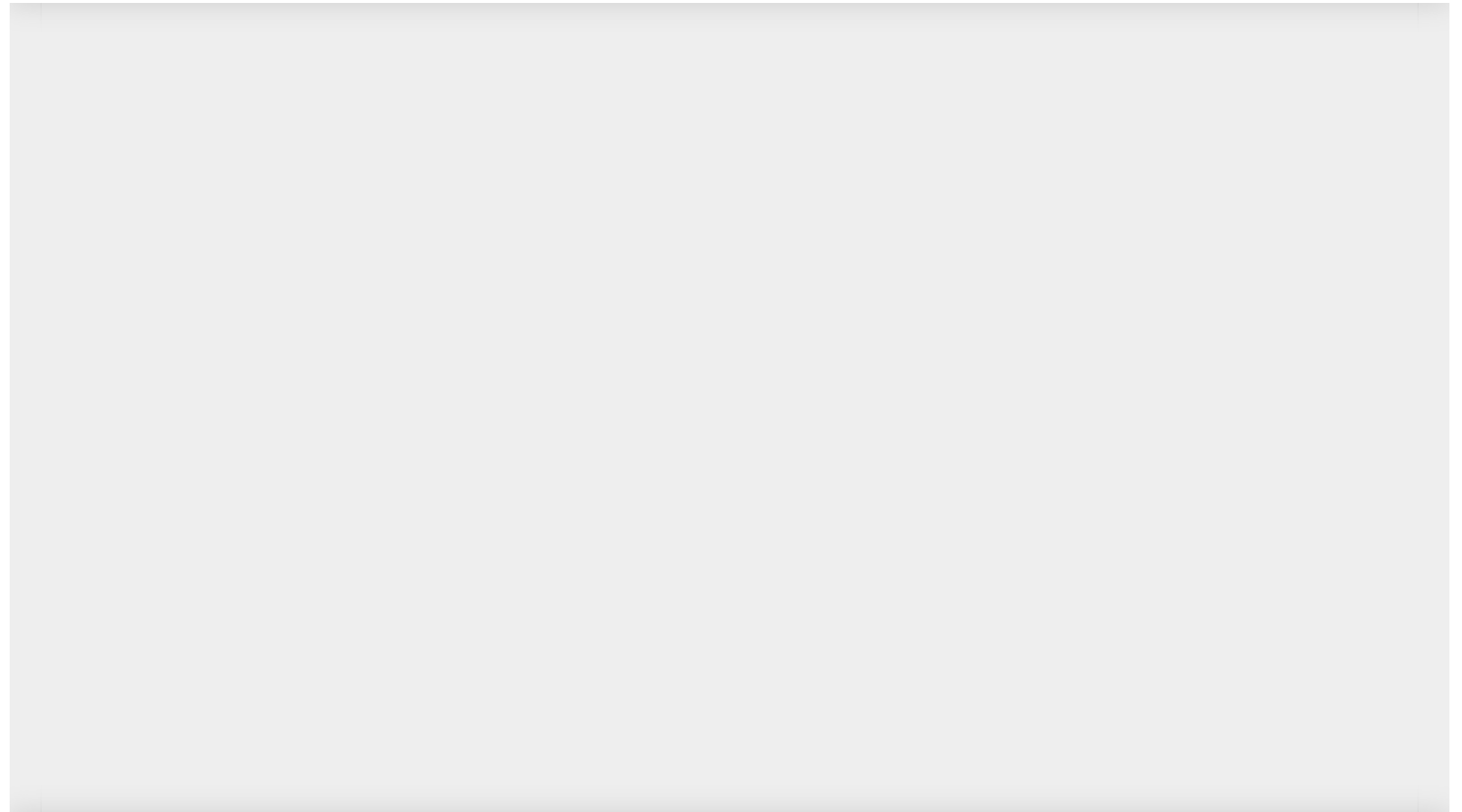


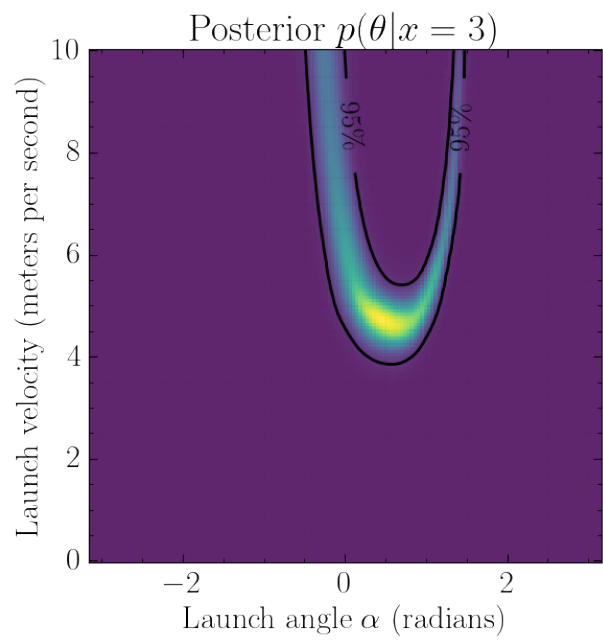
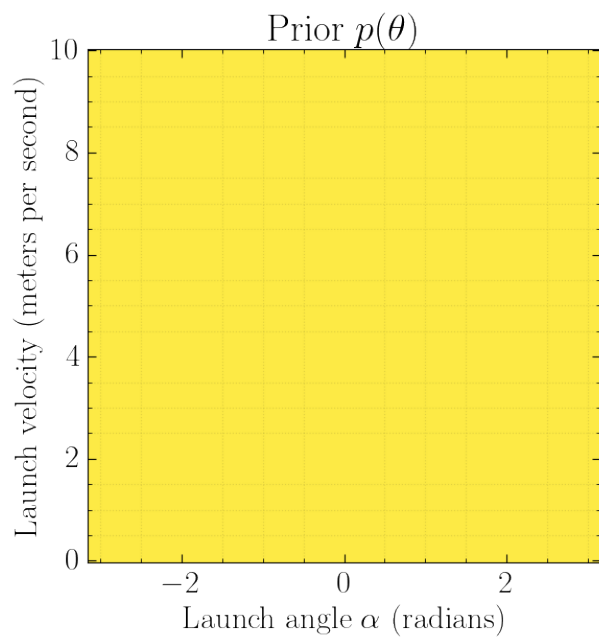
$$v_x = v \cos(\alpha), \quad v_y = v \sin(\alpha),$$
$$\frac{dx}{dt} = v_x, \quad \frac{dy}{dt} = v_y, \quad \frac{dv_y}{dt} = -G.$$

```
def simulate(v, alpha, dt=0.001):  
    v_x = v * np.cos(alpha) # x velocity m/s  
    v_y = v * np.sin(alpha) # y velocity m/s  
    y = 1.1 + 0.3 * random.normal()  
    x = 0.0  
  
    while y > 0: # simulate until ball hits floor  
        v_y += dt * -G # acceleration due to gravity  
        x += dt * v_x  
        y += dt * v_y  
  
    return x + 0.25 * random.normal()
```



What parameter values  $\theta$  are the most plausible?



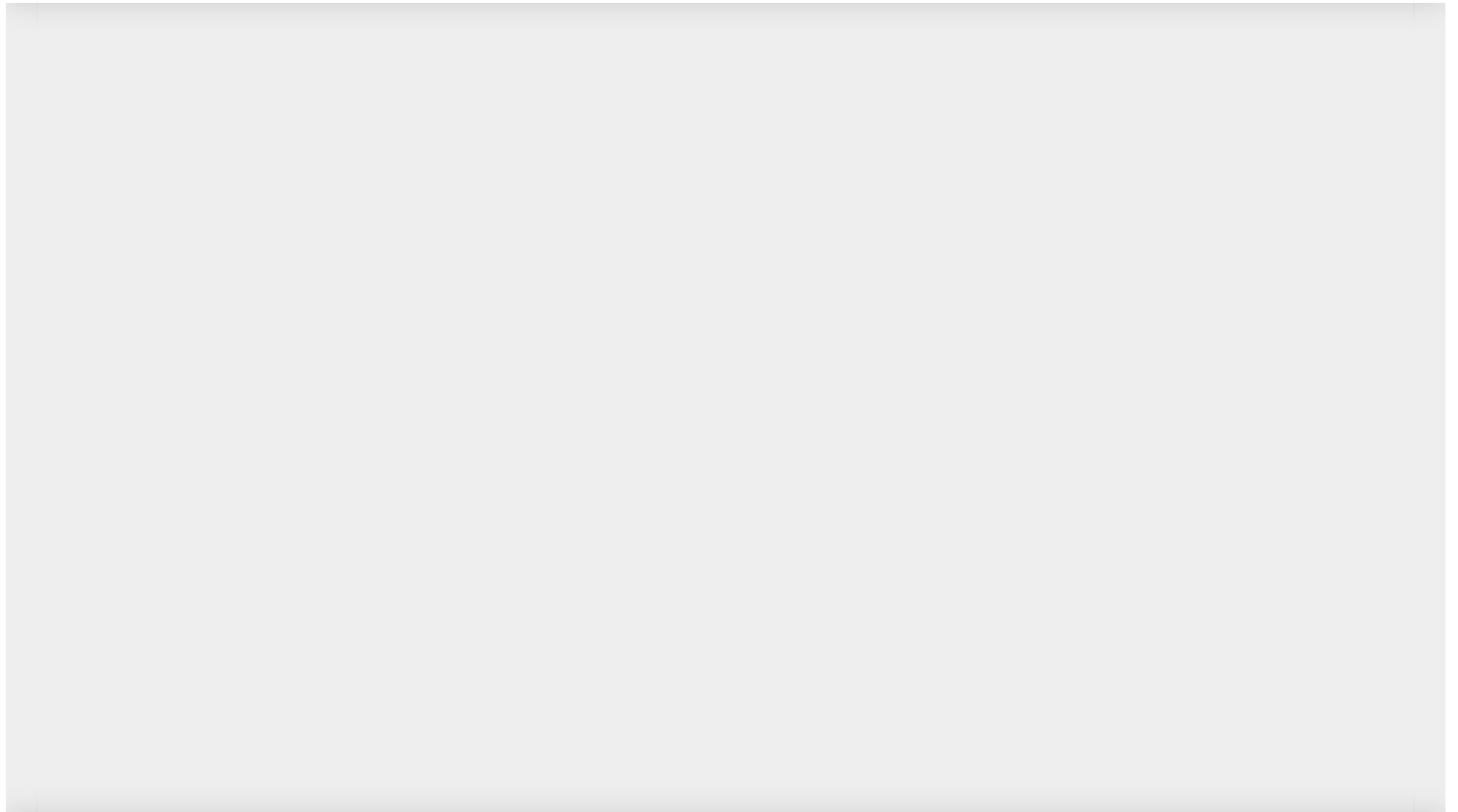


# Outline

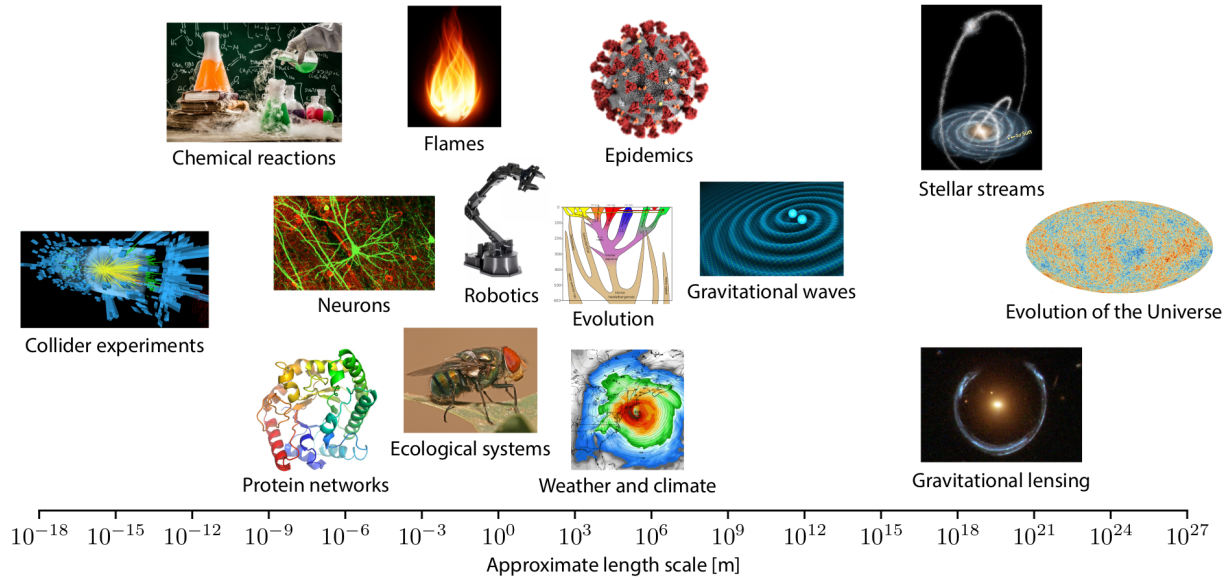
1. Simulation-based inference
2. Algorithms
  - Neural ratio estimation
  - Neural posterior estimation
  - Neural score estimation
3. Diagnostics

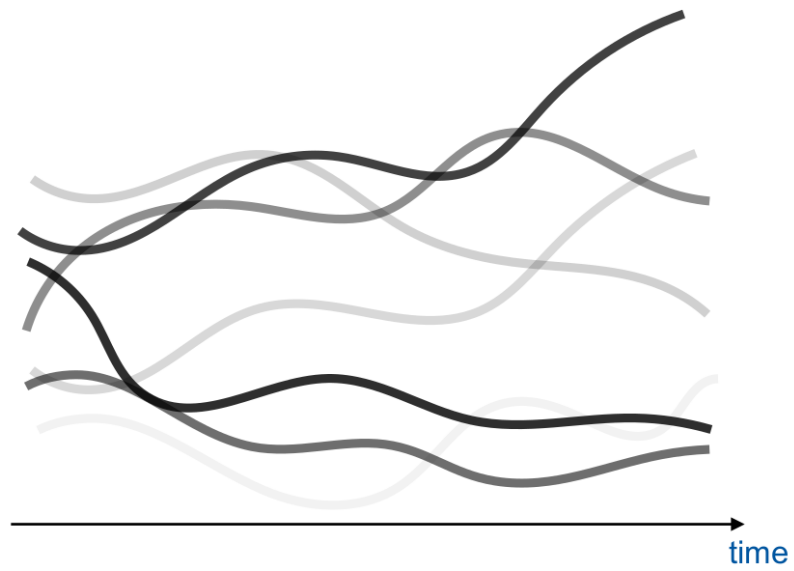


# Simulation-based inference

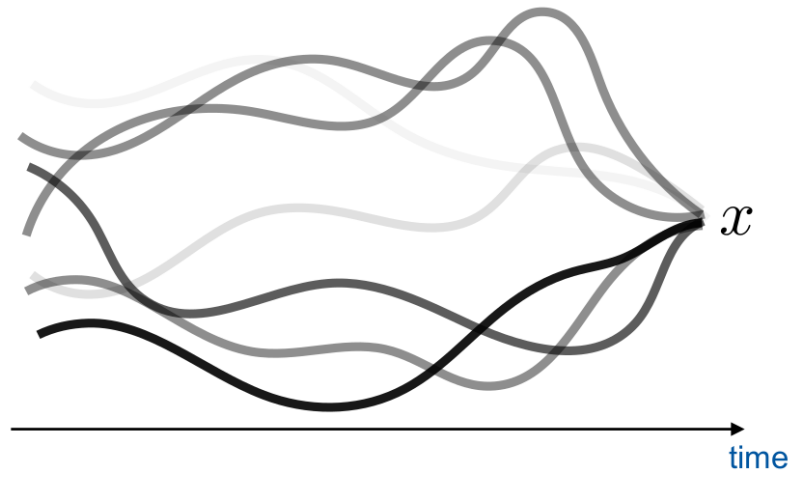


# Scientific simulators





$$\theta, z, x \sim p(\theta, z, x)$$



$$\theta, z \sim p(\theta, z|x)$$

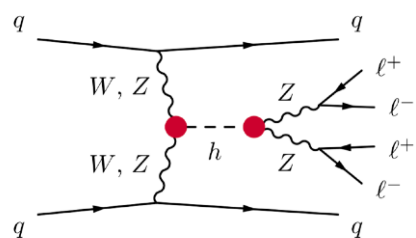
Latent variables

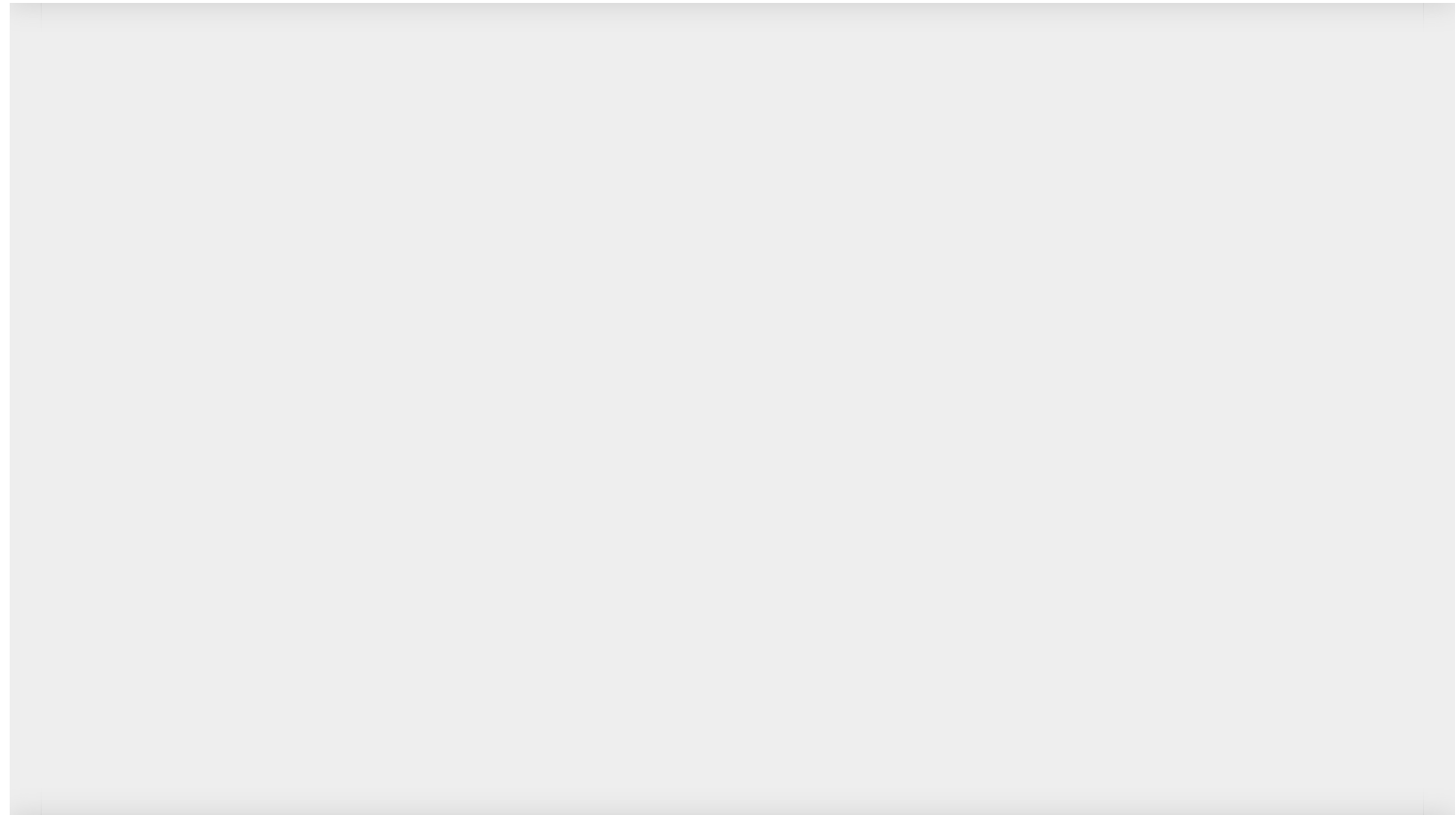
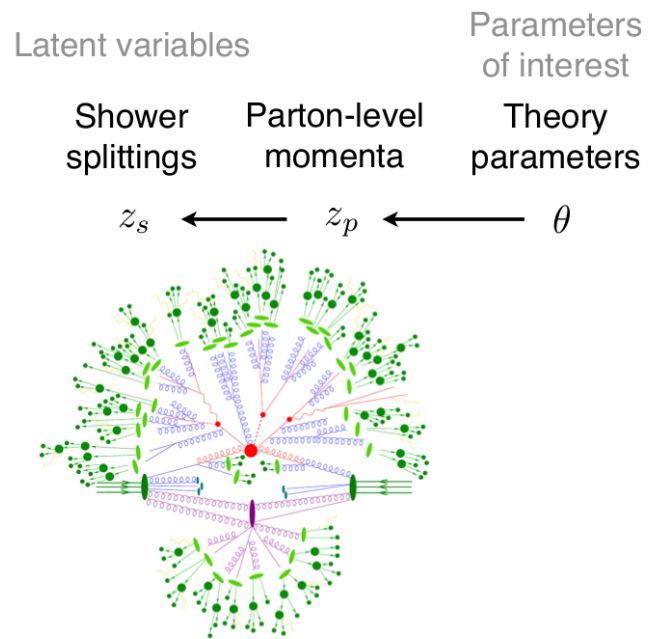
Parameters of interest

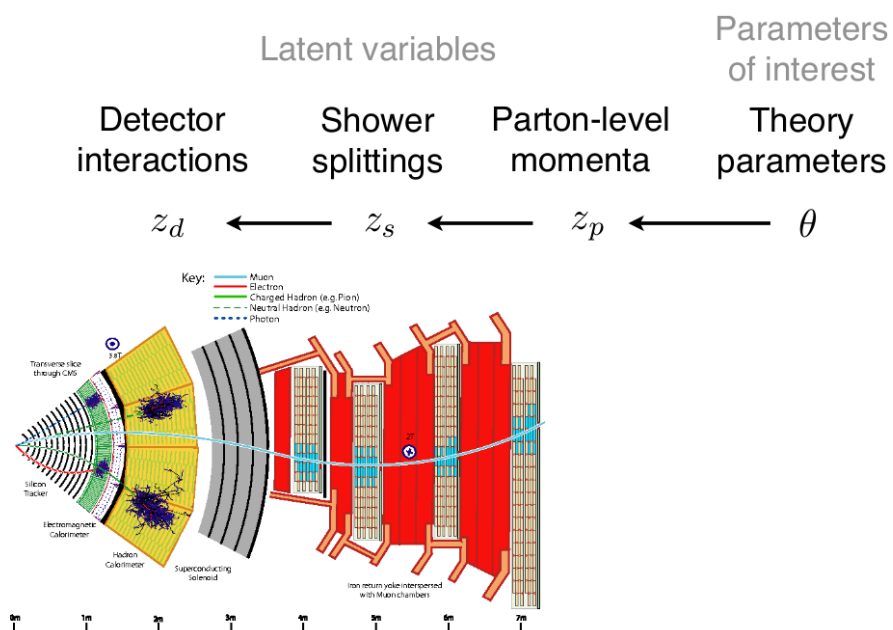
Parton-level momenta

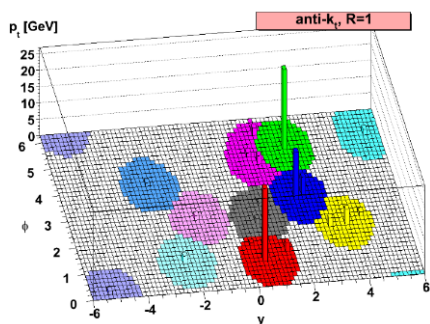
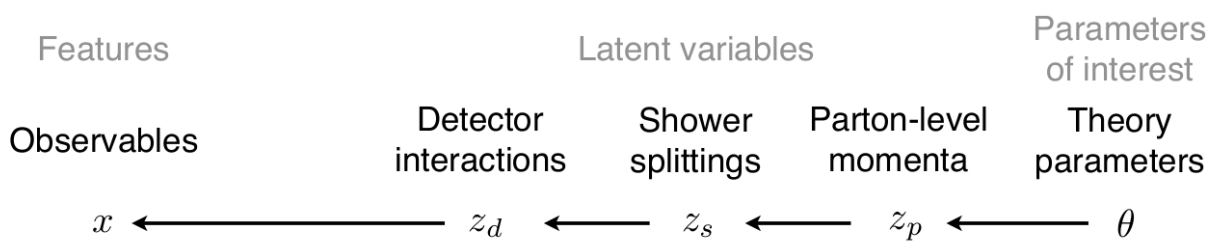
Theory parameters

$$z_p \longleftarrow \theta$$









[Image source: M. Cacciari, G. Salam, G. Soyez 0802.1189]



$$p(x|\theta) = \underbrace{\iiint}_{\text{yikes!}} p(z_p|\theta)p(z_s|z_p)p(z_d|z_s)p(x|z_d)dz_pdz_sdz_d$$

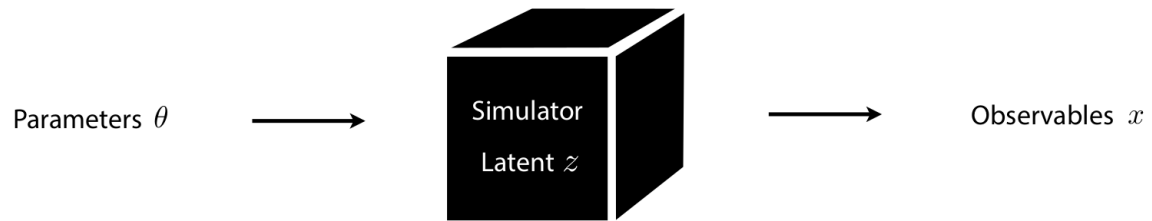
## Bayesian inference

Start with

- a simulator that can generate  $N$  samples  $\mathbf{x}_i \sim p(\mathbf{x}_i|\theta_i)$ ,
- a prior model  $p(\theta)$ ,
- observed data  $\mathbf{x}_{\text{obs}}$ .

Then, estimate the posterior

$$p(\theta|\mathbf{x}_{\text{obs}}) = \frac{p(\mathbf{x}_{\text{obs}}|\theta)p(\theta)}{p(\mathbf{x}_{\text{obs}})}.$$

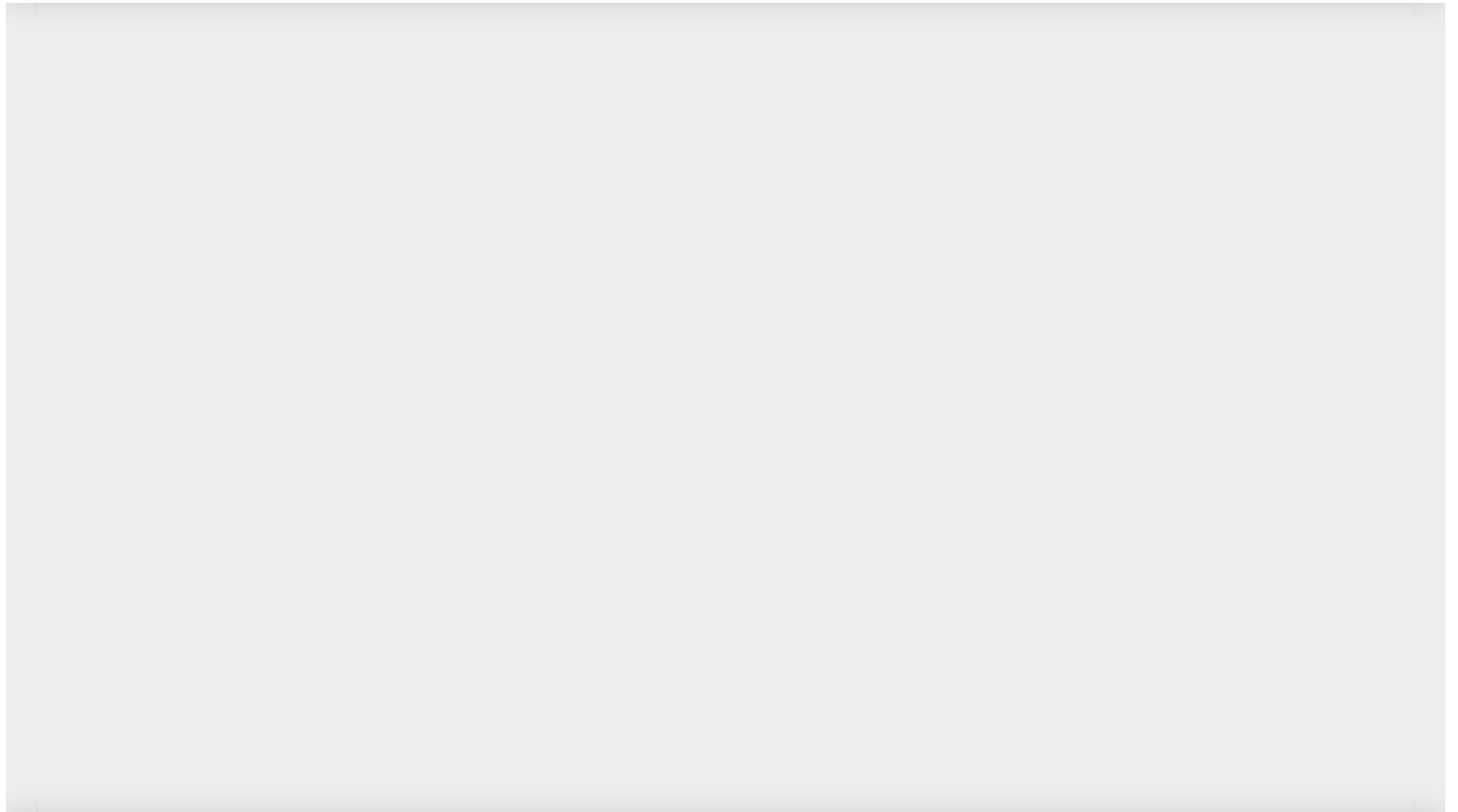


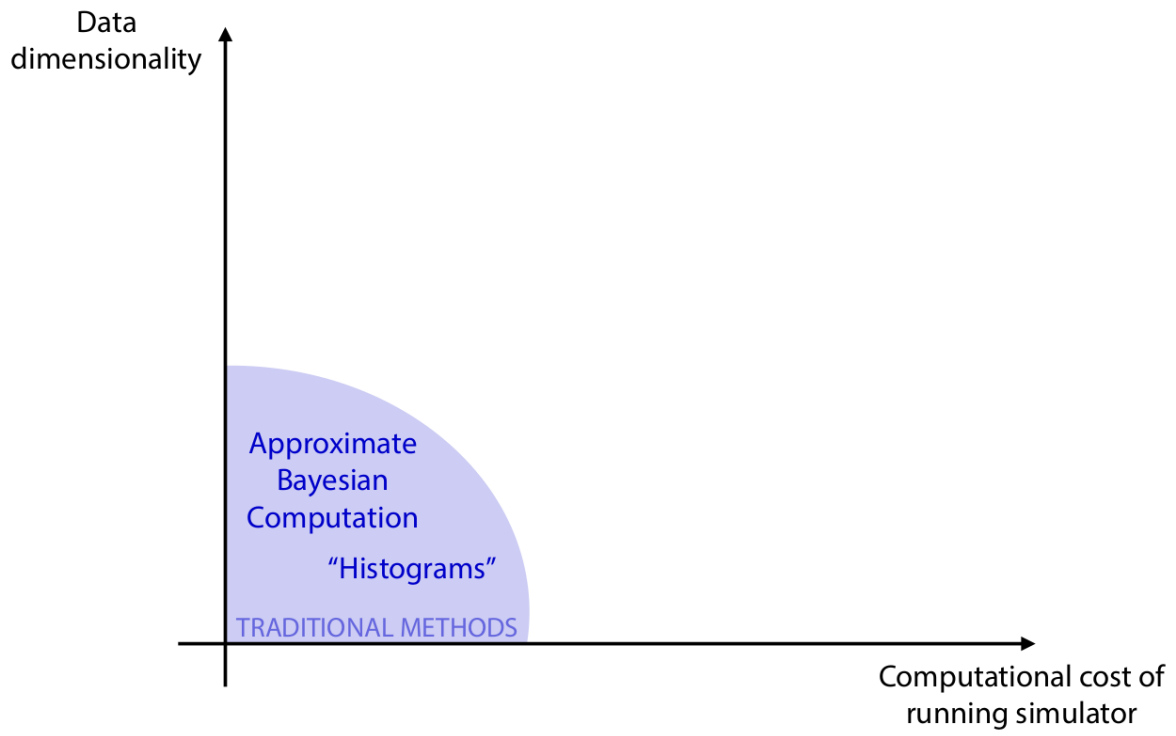
- Prediction:
- Well-motivated mechanistic, causal model
  - Simulator can generate samples  $x \sim p(x|\theta)$



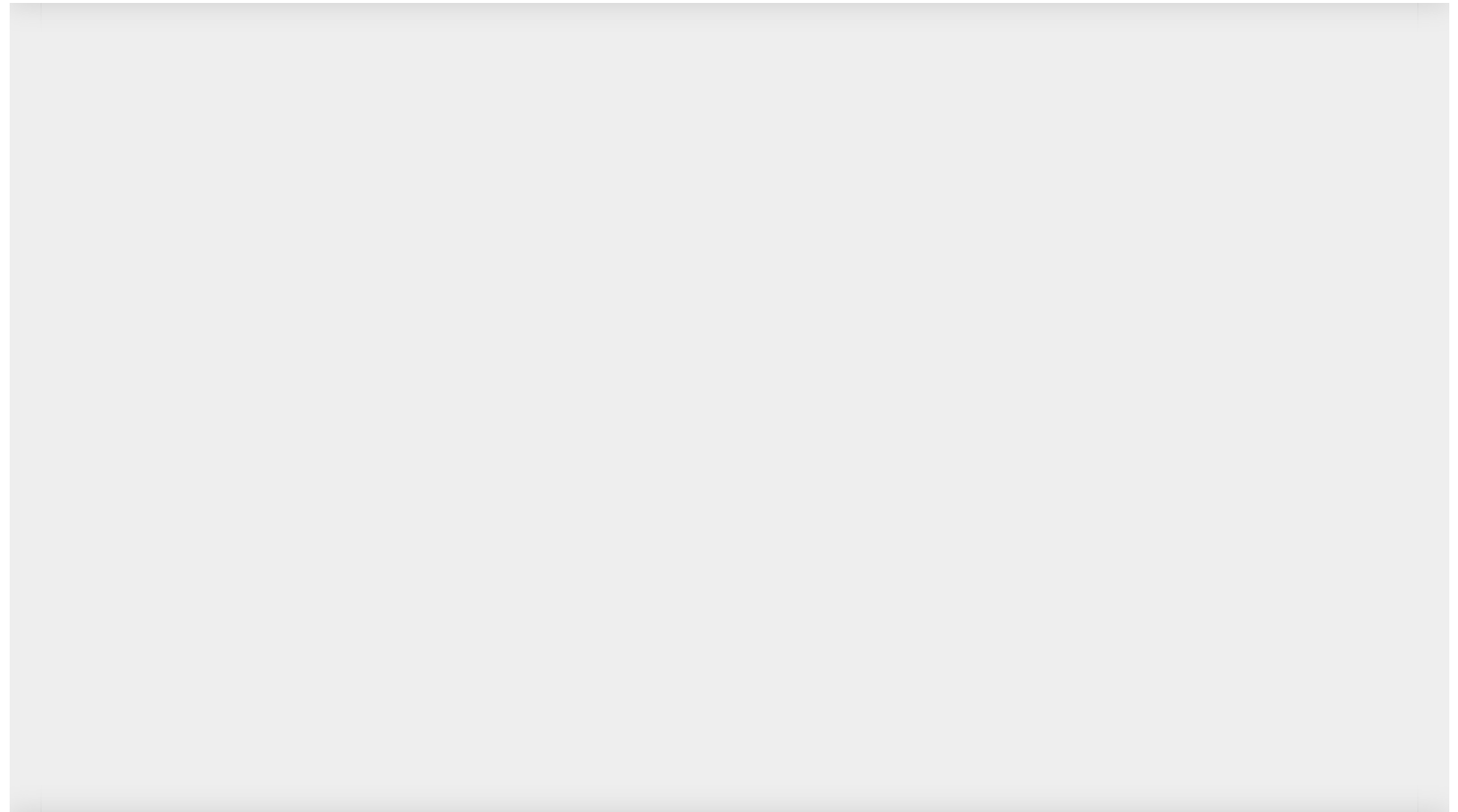
- Inference:
- Interactions between low-level components lead to challenging inverse problems
  - Likelihood  $p(x|\theta) = \int dz p(x, z|\theta)$  is intractable

# Algorithms

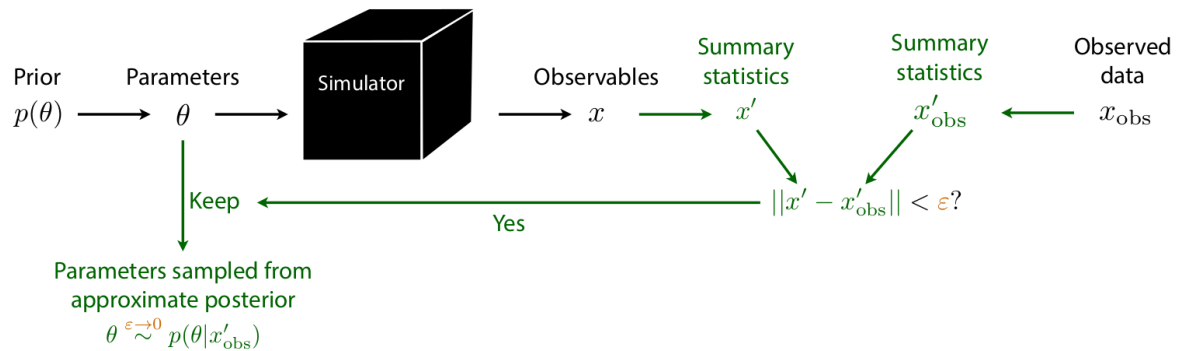




Credits: [Cranmer, Brehmer and Louppe](#), 2020.

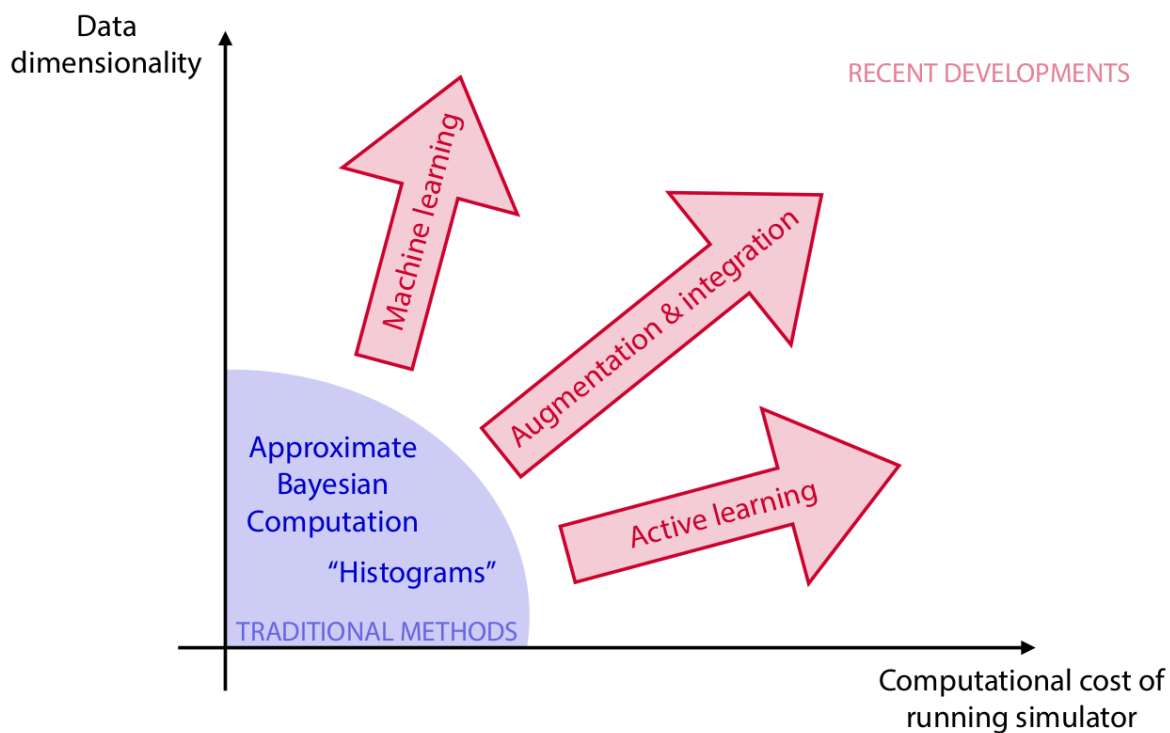


## Approximate Bayesian Computation (ABC)



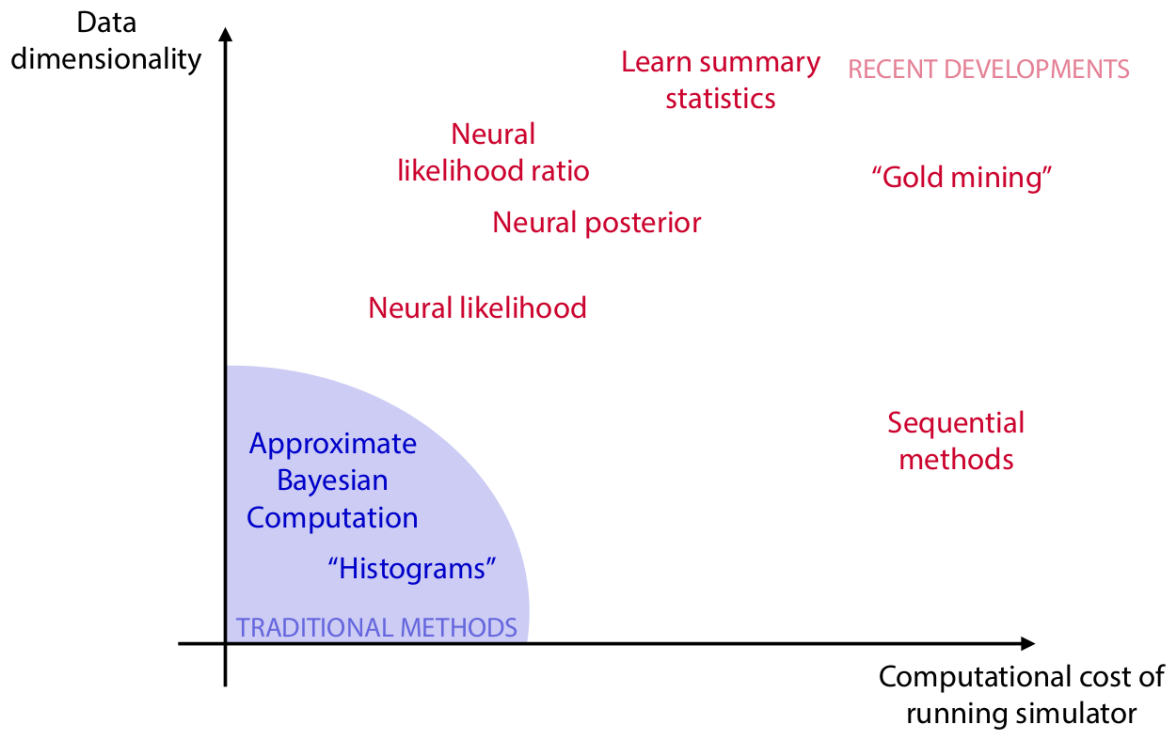
*Issues:*

- How to choose  $x'$ ?  $\epsilon$ ?  $\| \cdot \|$ ?
- No tractable posterior.
- Need to run new simulations for new data or new prior.



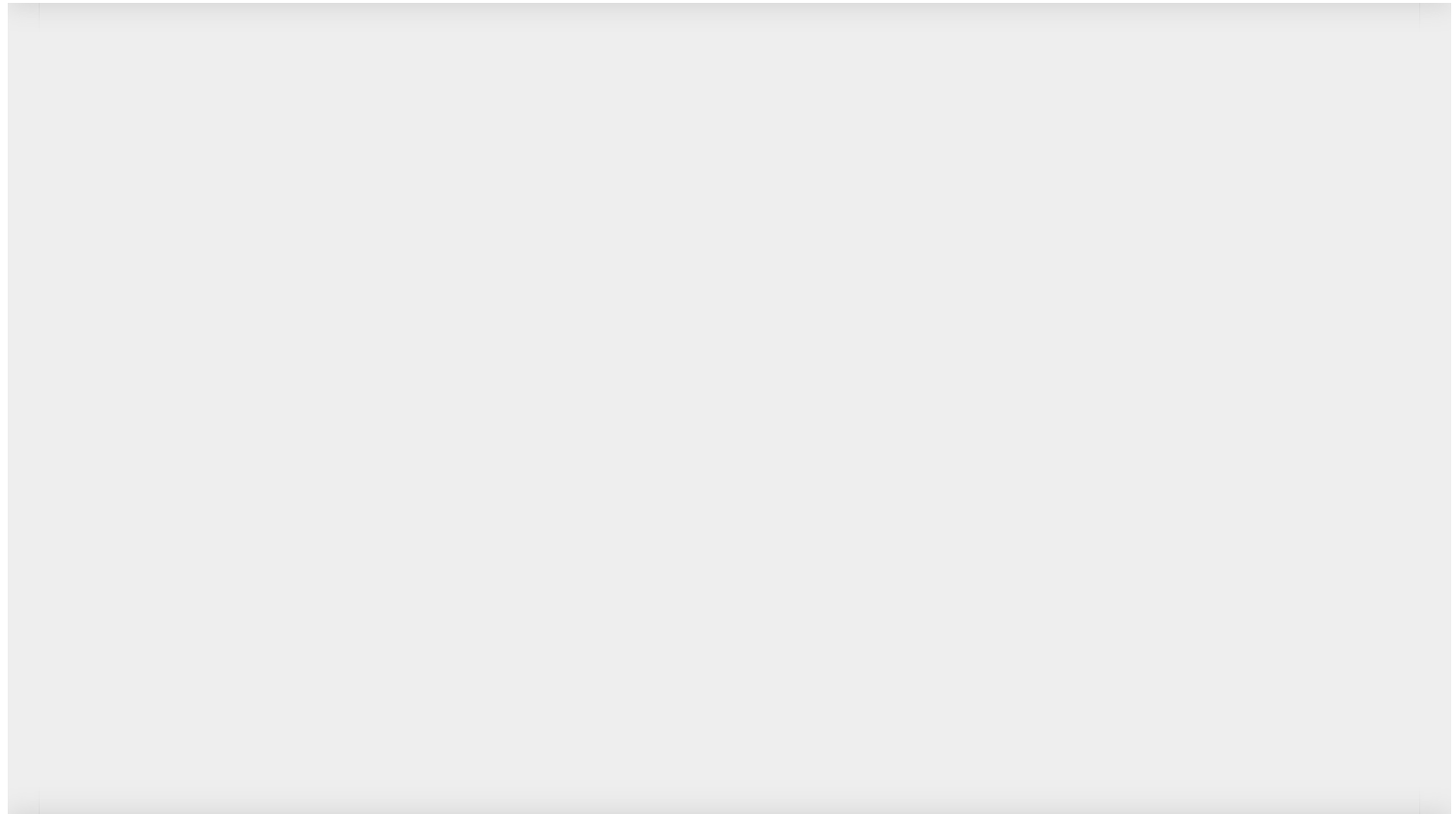
Credits: Cranmer, Brehmer and Louppe, 2020.

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Credits: [Cranmer, Brehmer and Louppe, 2020](#).

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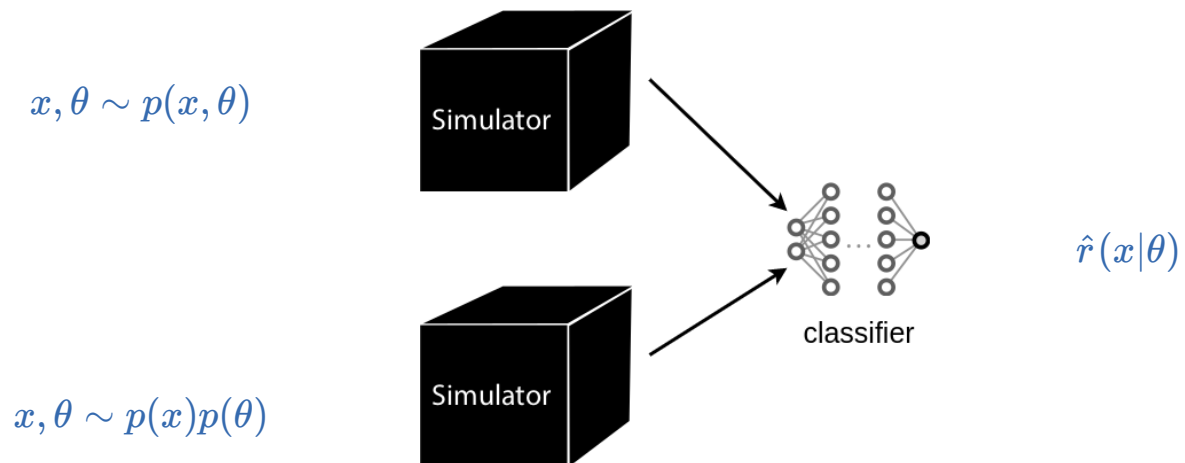




# Neural ratio estimation



The likelihood-to-evidence  $r(x|\theta) = \frac{p(x|\theta)}{p(x)} = \frac{p(x,\theta)}{p(x)p(\theta)}$  ratio can be learned, even if neither the likelihood nor the evidence can be evaluated:



—  
Credits: [Cranmer et al, 2015](#); [Hermans et al, 2020](#).

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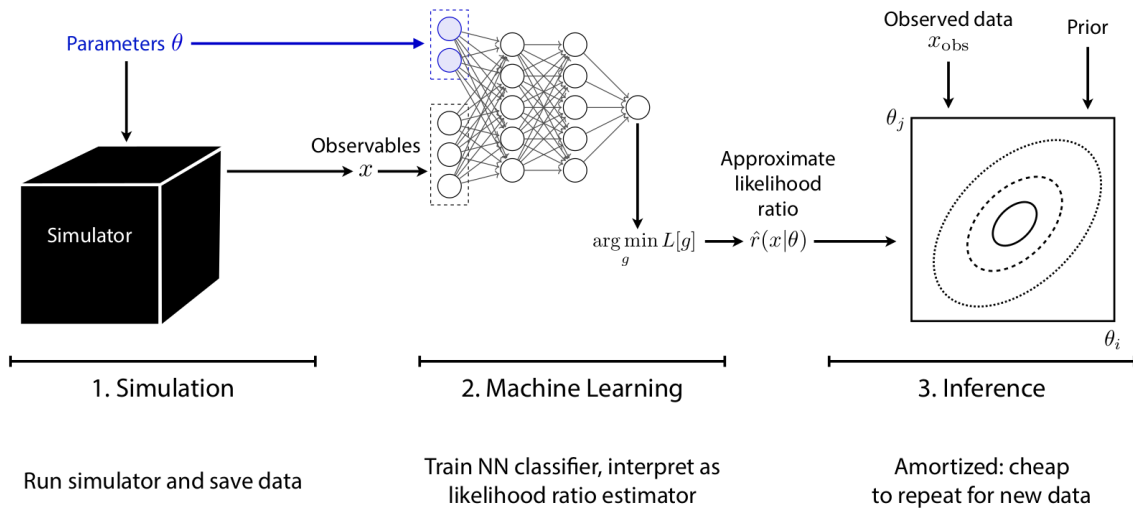


The solution  $d$  found after training approximates the optimal classifier

$$d(x, \theta) \approx d^*(x, \theta) = \frac{p(x, \theta)}{p(x, \theta) + p(x)p(\theta)}.$$

Therefore,

$$r(x|\theta) = \frac{p(x|\theta)}{p(x)} = \frac{p(x, \theta)}{p(x)p(\theta)} \approx \frac{d(x, \theta)}{1 - d(x, \theta)} = \hat{r}(x|\theta).$$



$$p(\theta|x) \approx \hat{r}(x|\theta)p(\theta)$$

# Constraining dark matter with stellar streams



**Palomar 5 (Pal5) stream**  
 Pal5 was discovered in 2001 as the first thin stream formed from a globular cluster. Its current orbit takes it far over the galactic center.

**Globular clusters**  
 These hives typically hold 100,000 stars or fewer and give rise to long, thin streams.

Gap

Sun

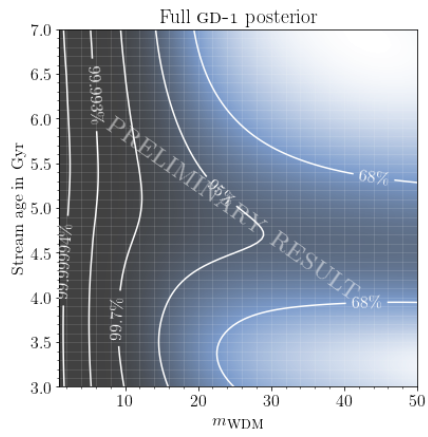
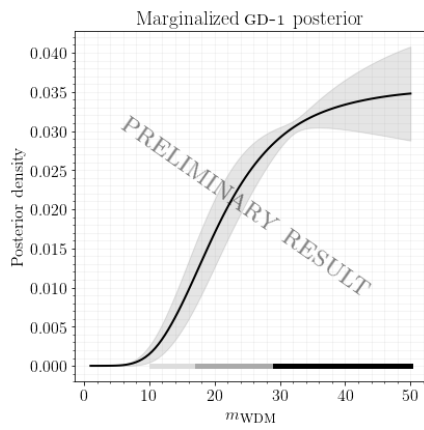
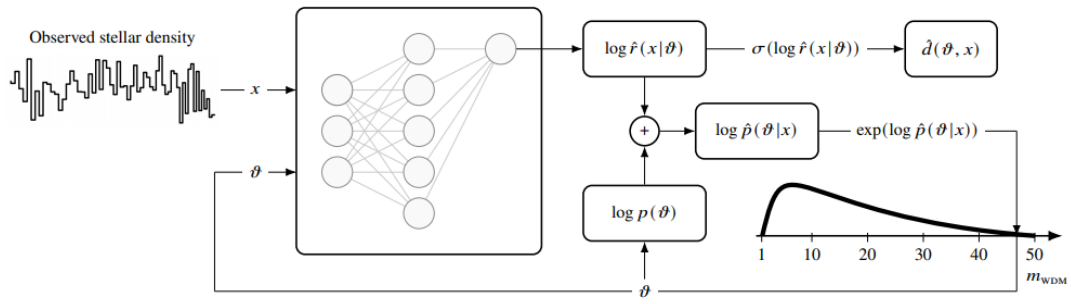
Milky Way

**GD1 stream**  
 Discovered in 2006, GD1 is the longest known thin stream, stretching across more than half the northern sky. It contains a gap that could be the scar of a dark matter collision 500 million years ago.

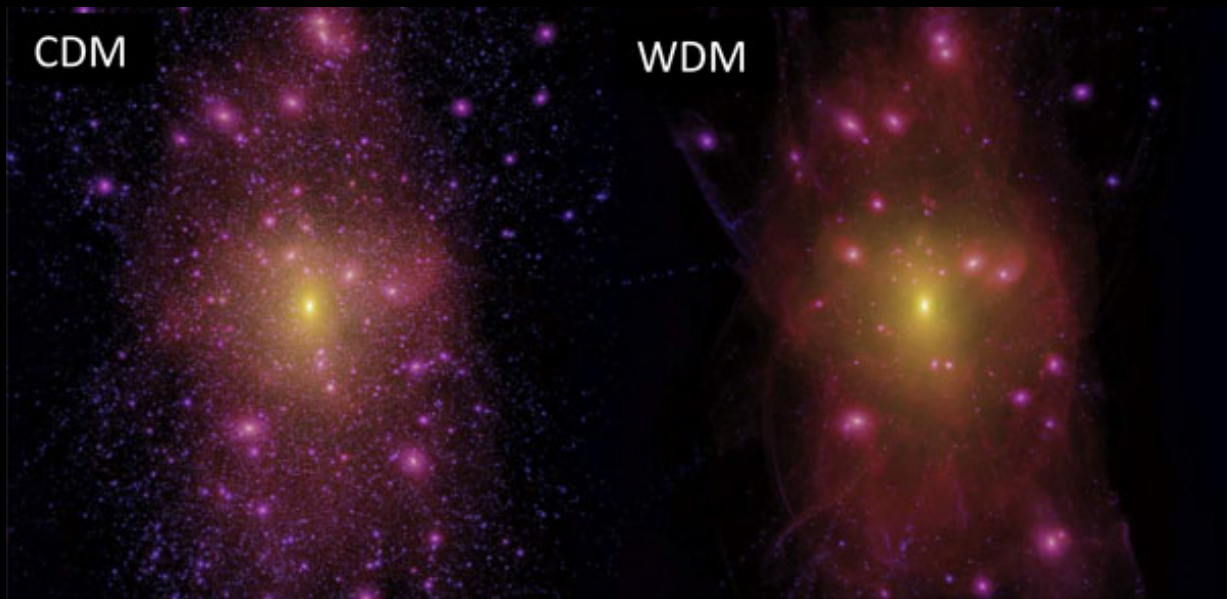
Interaction of Pal 5 with two dark ...



Image credits: C. McKee, R. S. Stencel, D. Lyra

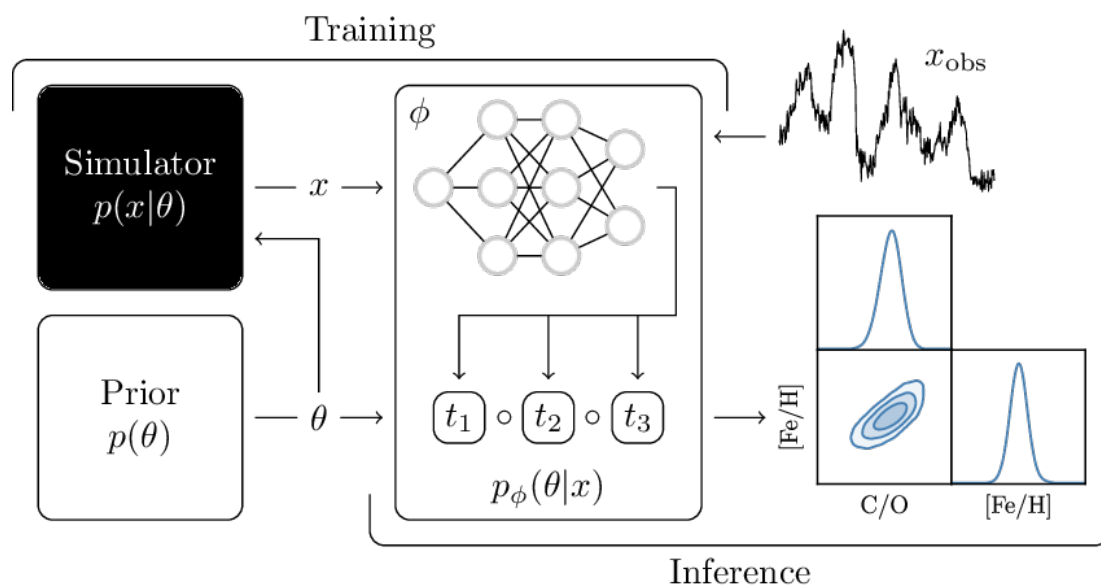


Credits: Hermans et al, 2021.



Preliminary results for GD-1 suggest a **preference for CDM over WDM.**

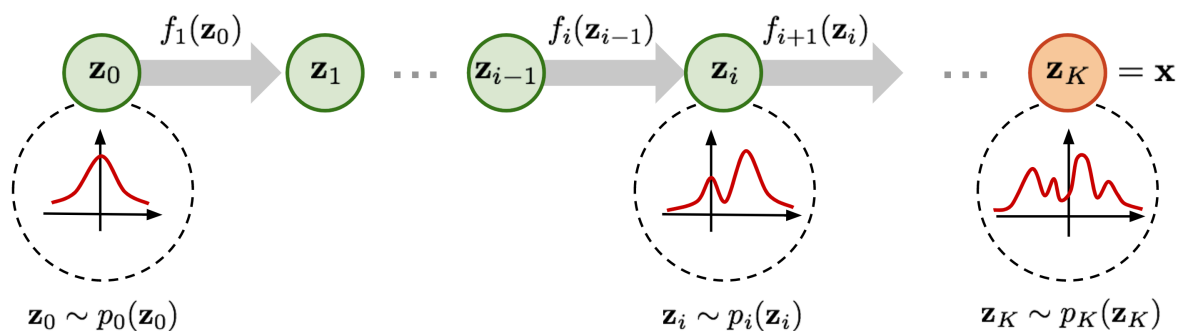
# Neural Posterior Estimation



$$\min_{q_\phi} \mathbb{E}_{p(x)} [\text{KL}(p(\theta|x) || q_\phi(\theta|x))]$$

## Normalizing flows

A normalizing flow is a sequence of invertible transformations  $f_k$  that map a simple distribution  $p_0$  to a more complex distribution  $p_K$ :

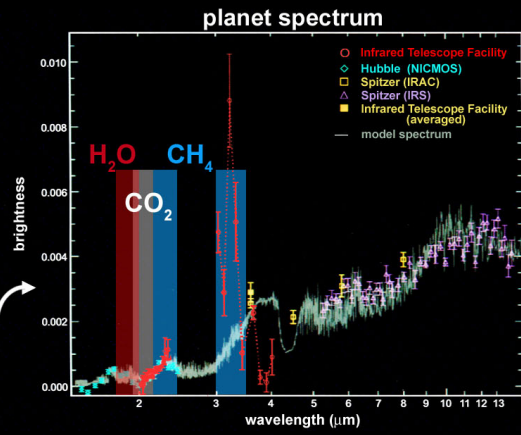
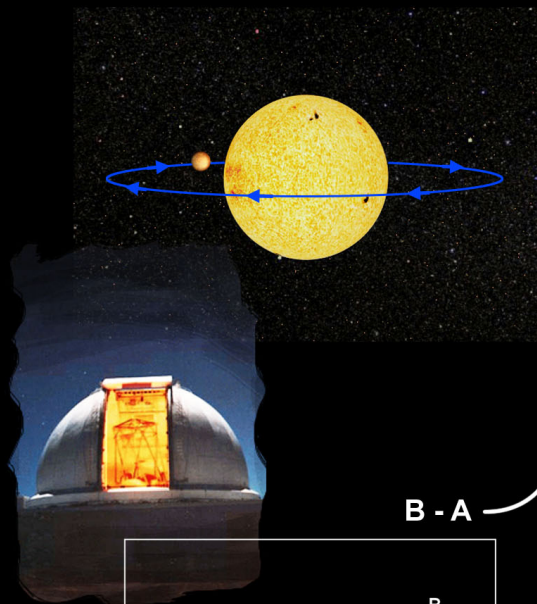


By the change of variables formula, the log-likelihood of a sample  $\mathbf{x}$  is given by

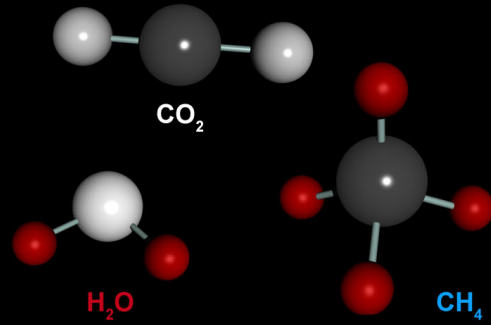
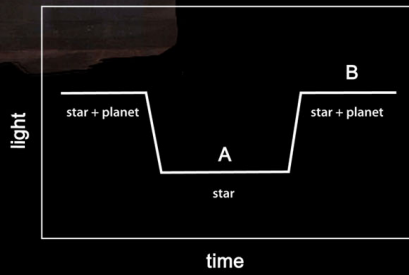
$$\log p(\mathbf{x}) = \log p(z_0) - \sum_{k=1}^K \log |\det J_{f_k}(z_{k-1})|.$$



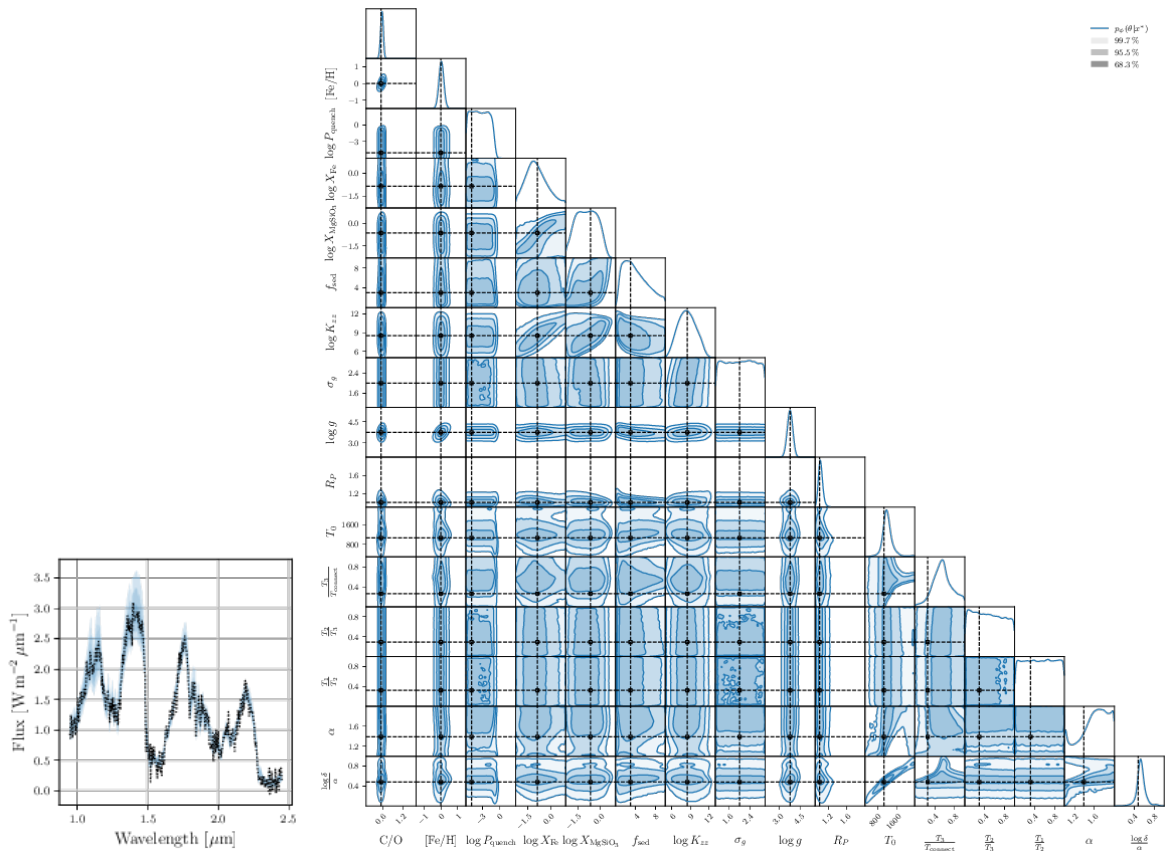
# Exoplanet atmosphere characterization



B - A

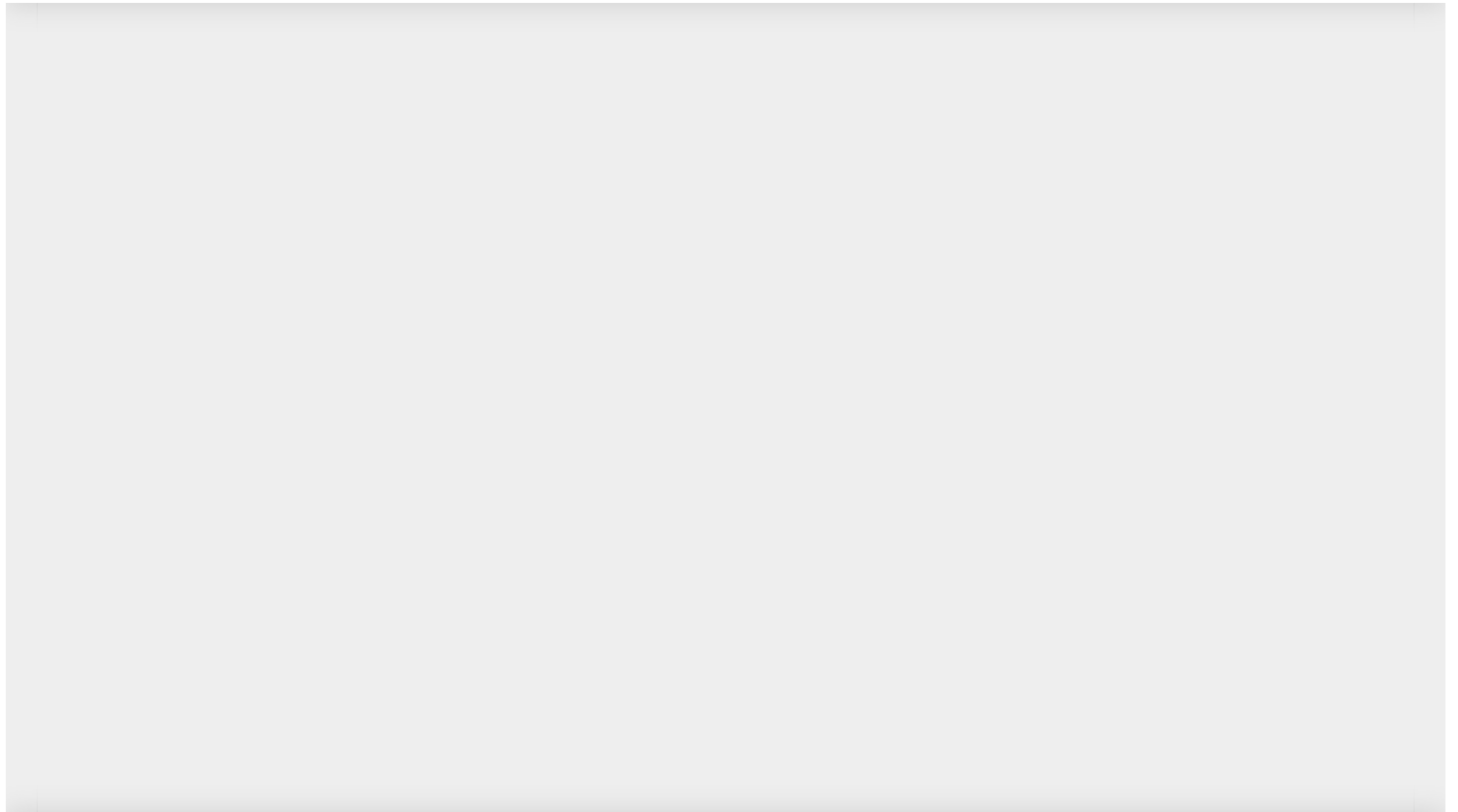


Credits: NSA/JPL-Caltech, 2010.



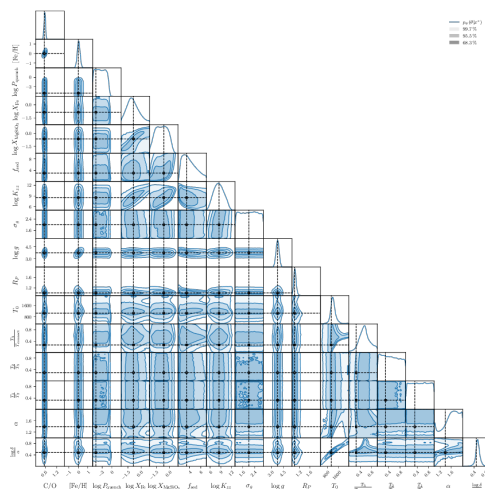
Credits: Vasist et al, 2023.

# Diagnostics



$$\hat{p}(\theta|x) = \text{sbi}(p(x|\theta), p(\theta), x)$$

We must make sure our approximate simulation-based inference algorithms can (at least) actually realize faithful inferences on the (expected) observations.



*How certain are you of your uncertainties?*

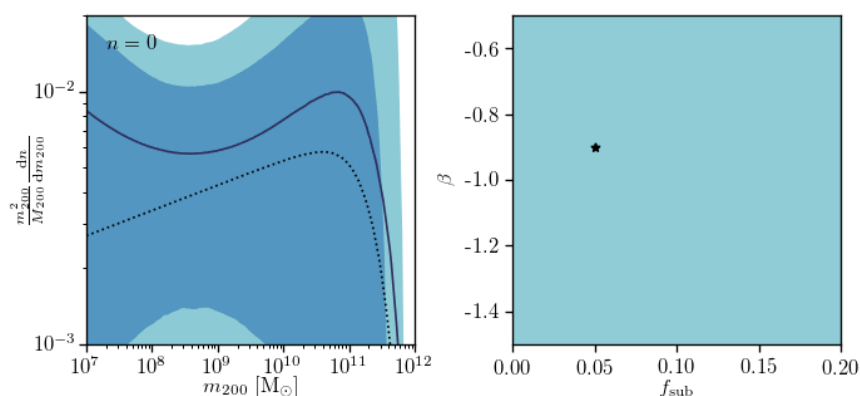


## Mode convergence

The maximum a posteriori estimate converges towards the nominal value  $\theta^*$  for an increasing number of independent and identically distributed observables

$x_i \sim p(x|\theta^*)$ :

$$\begin{aligned} & \lim_{N \rightarrow \infty} \arg \max_{\theta} p(\theta | \{x_i\}_{i=1}^N) \\ &= \lim_{N \rightarrow \infty} \arg \max_{\theta} p(\theta) \prod_{x_i} r(x_i | \theta) = \theta^* \end{aligned}$$



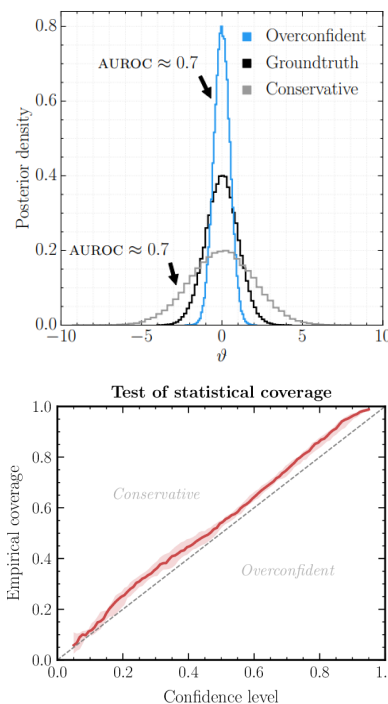
Credits: Brehmer et al, 2019.

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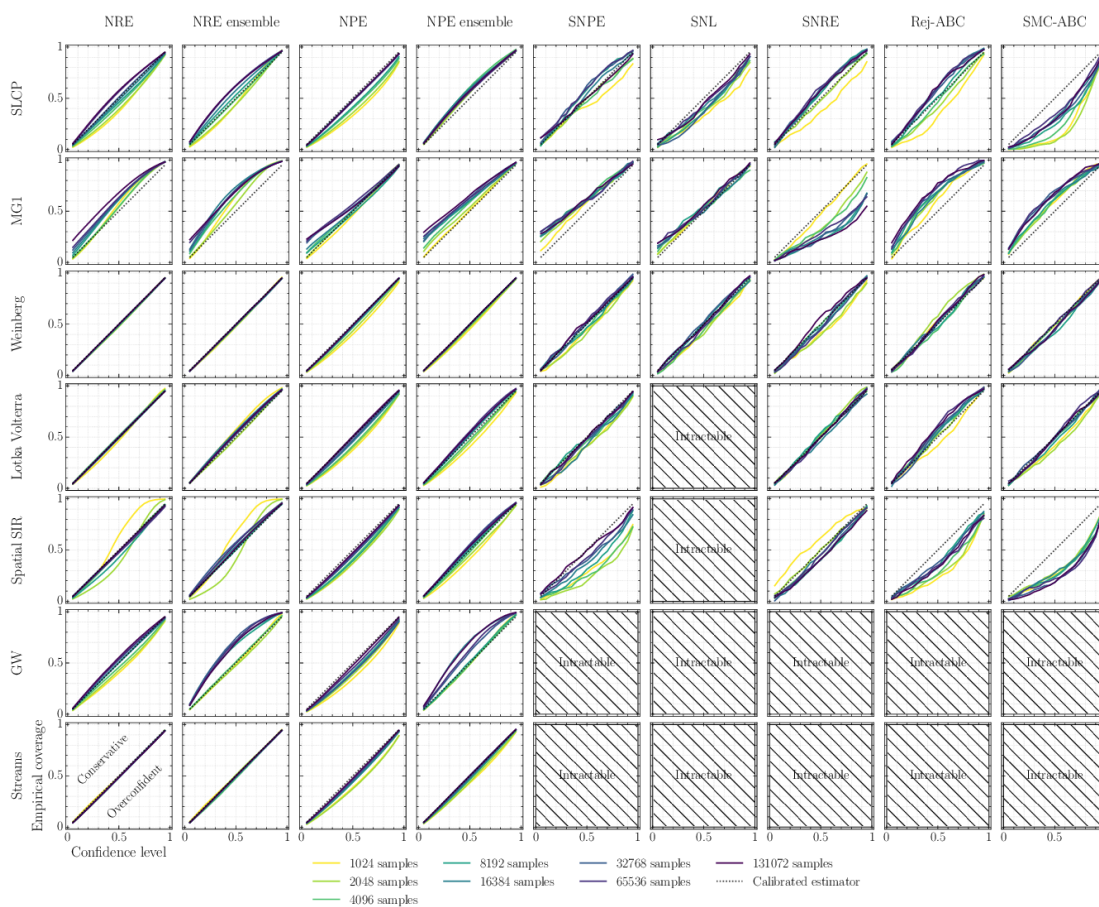
## Coverage diagnostic

- For  $x, \theta \sim p(x, \theta)$ , compute the  $1 - \alpha$  credible interval based on  $\hat{p}(\theta|x)$ .
- If the fraction of samples for which  $\theta$  is contained within the interval is larger than the nominal coverage probability  $1 - \alpha$ , then the approximate posterior  $\hat{p}(\theta|x)$  has coverage.



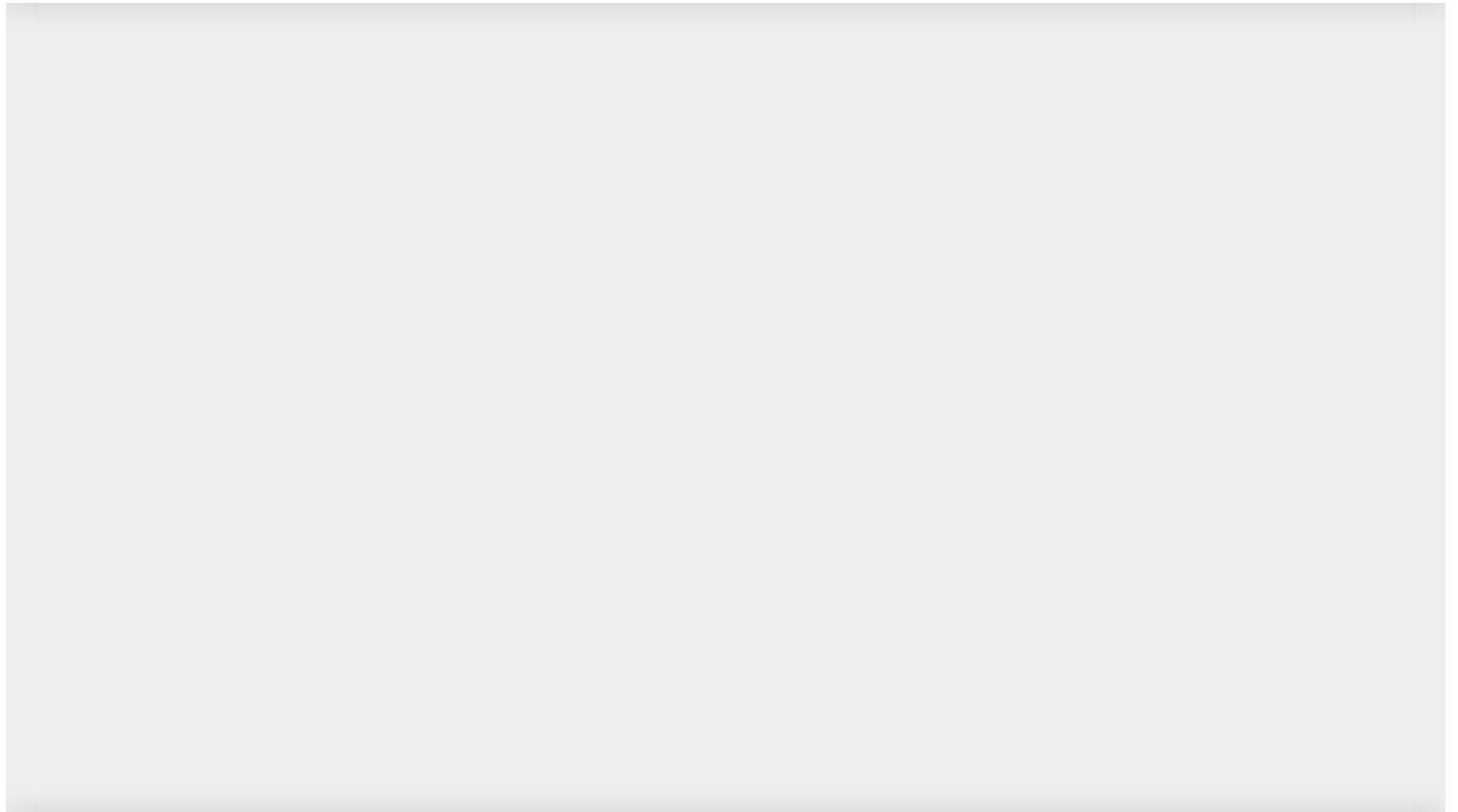
—  
Credits: Hermans et al, 2021; Siddharth Mishra-Sharma, 2021.

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Credits: Hermans et al, 2021.

What if diagnostics fail?



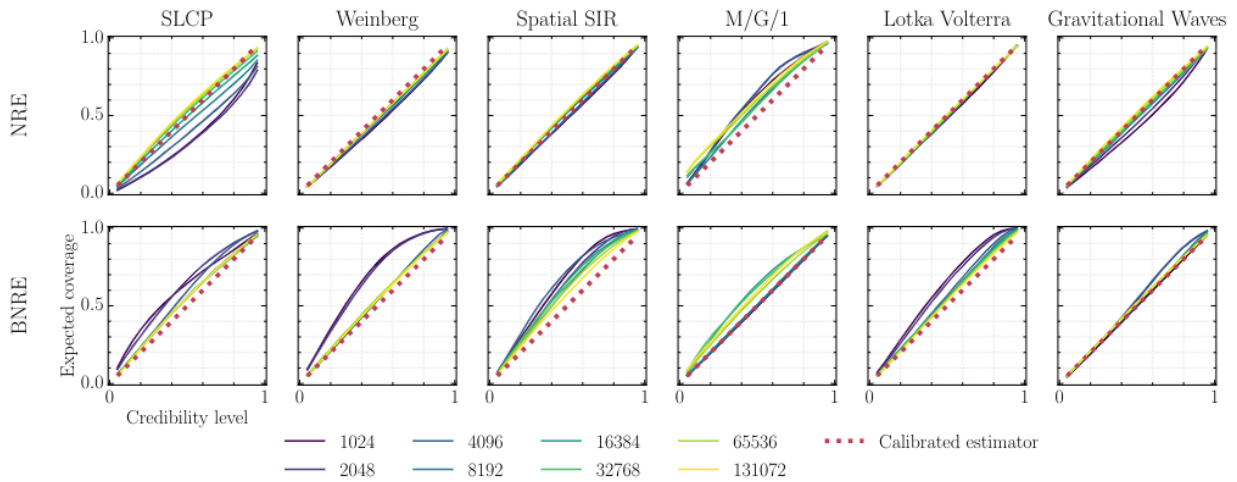


# Balanced NRE

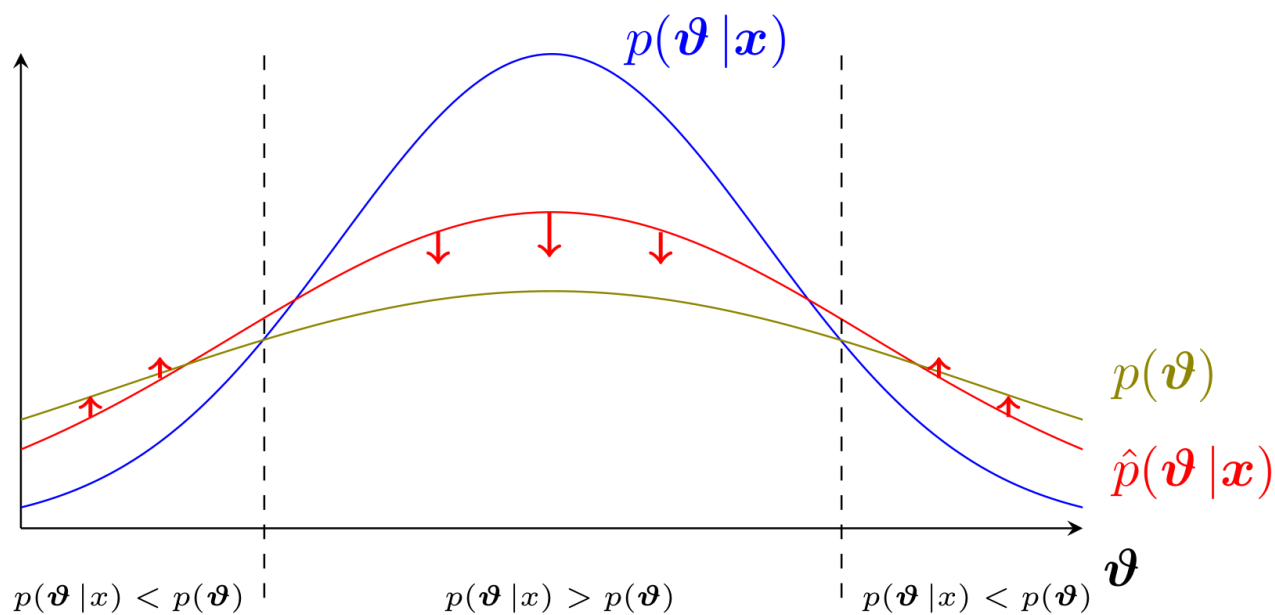


Enforce neural ratio estimation to be **conservative** by using binary classifiers  $\hat{d}$  that are balanced, i.e. such that

$$\mathbb{E}_{p(\theta, x)} \left[ \hat{d}(\theta, x) \right] = \mathbb{E}_{p(\theta)p(x)} \left[ 1 - \hat{d}(\theta, x) \right].$$



Credits: [Delaunoy et al, 2022](#).



Credits: Delaunoy et al, 2022.

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# Summary

Advances in deep learning have enabled new approaches to statistical inference.

This is major evolution in the statistical capabilities for science, as it enables the analysis of complex models and data without simplifying assumptions.

Inference remains approximate and requires careful validation.

Obstacles remain to be overcome, such as the curse of dimensionality and the need for large amounts of data.

The end.