

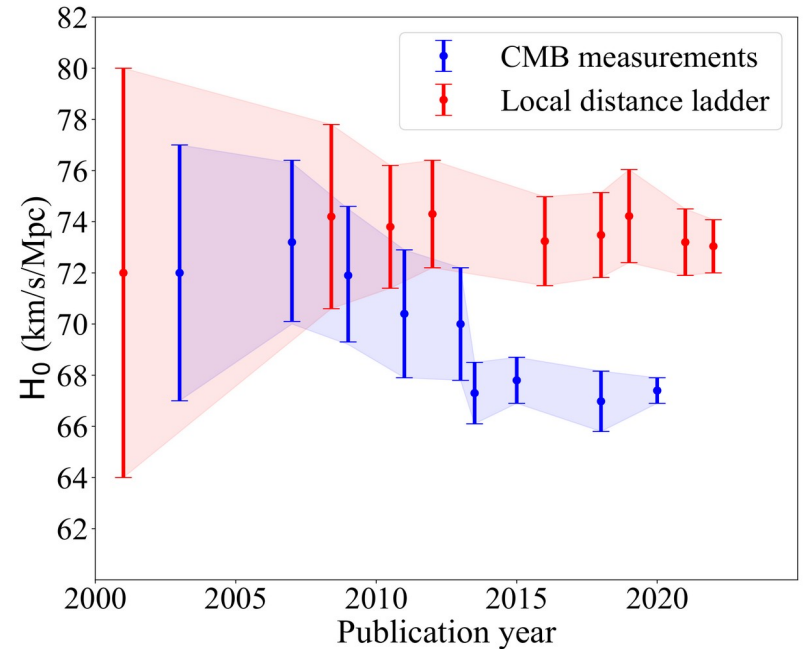
AGN interferometric observations to measure cosmological distances

Pierre Vermot
Atelier API “Ondes gravitationnelles et objets compacts”
Nov 16 – 17, 2023

Context: Hubble tension

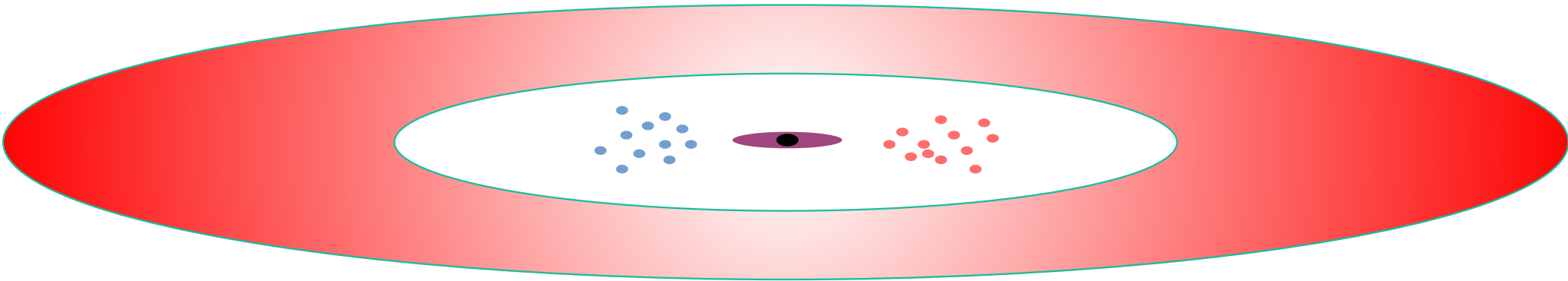
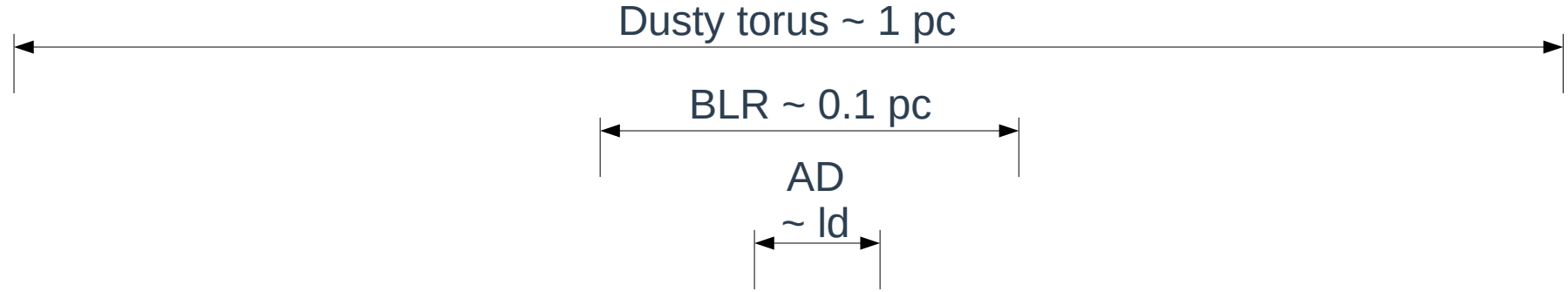
$\sim 5\sigma$ difference between H_0 measurements:

- **Local Universe:**
 $H_0 = 73.04 \pm 1.04 \text{ Km.s}^{-1}.\text{Mpc}^{-1}$
- **CMB:**
 $H_0 = 67.27 \pm 0.60 \text{ Km.s}^{-1}.\text{Mpc}^{-1}$



Hu et al (2023)

Solution: Active Galactic Nuclei as independent distance measurements

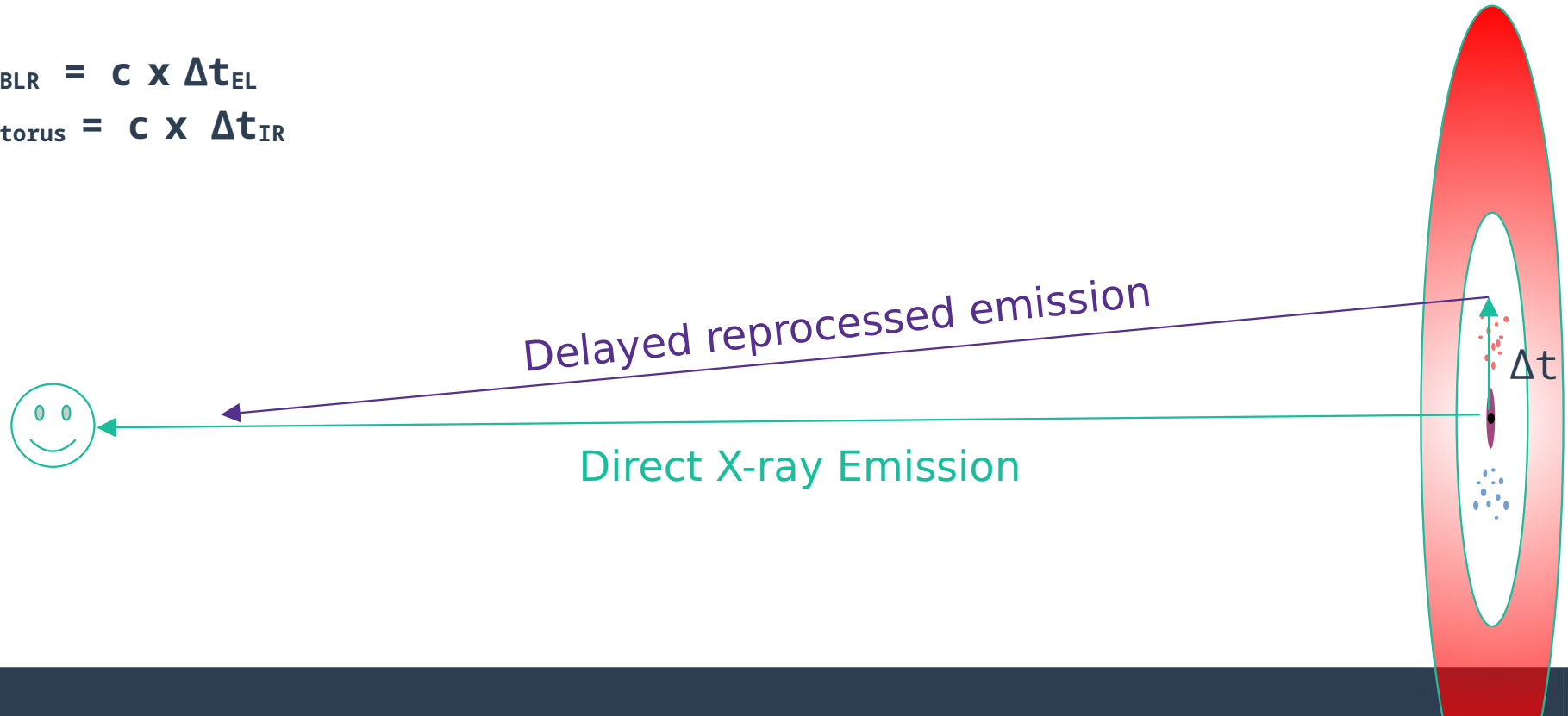


A) Reverberation Mapping

Time delay between X-ray variations vs BLR and torus IR variations

$$R_{\text{BLR}} = c \times \Delta t_{\text{EL}}$$

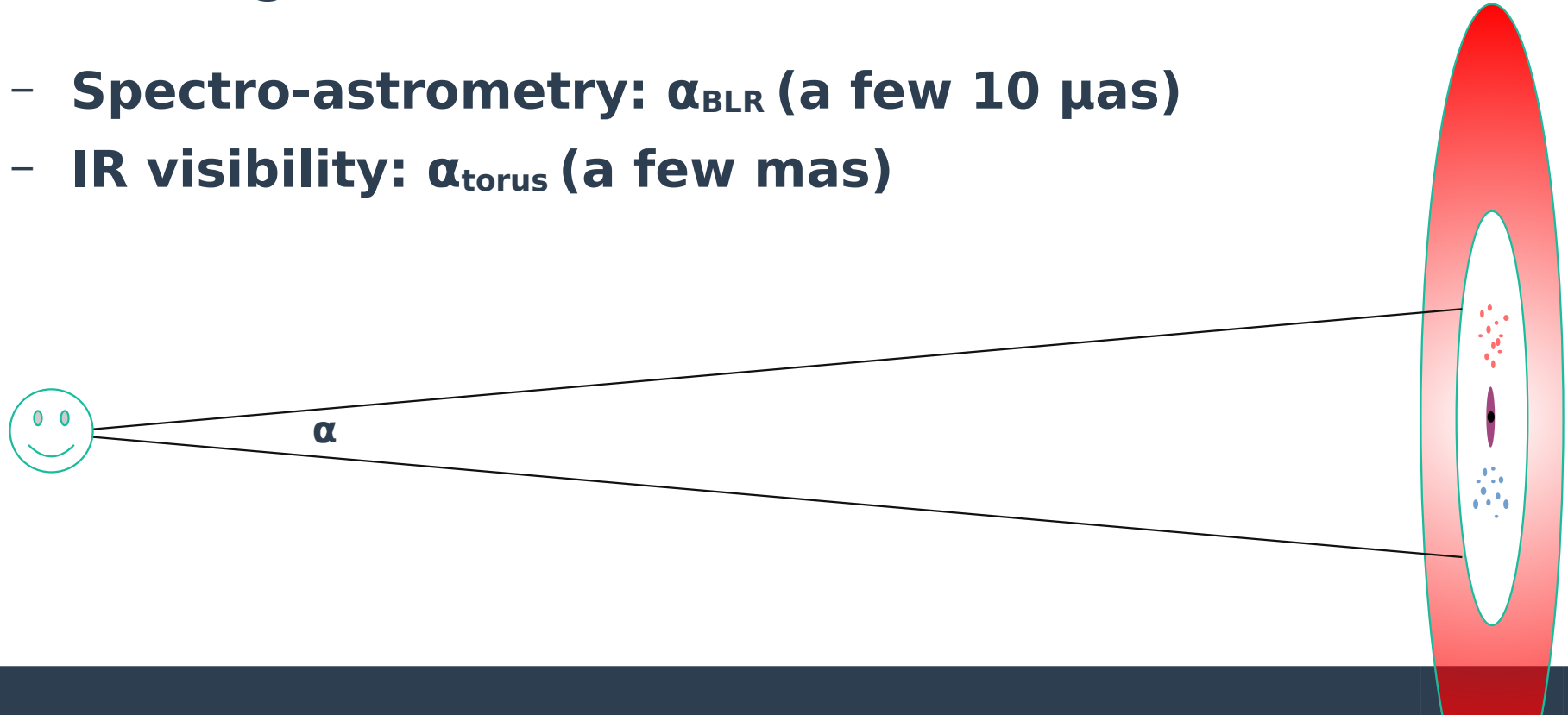
$$R_{\text{torus}} = c \times \Delta t_{\text{IR}}$$



A) Interferometry

Direct angular sizes with GRAVITY/VLTI:

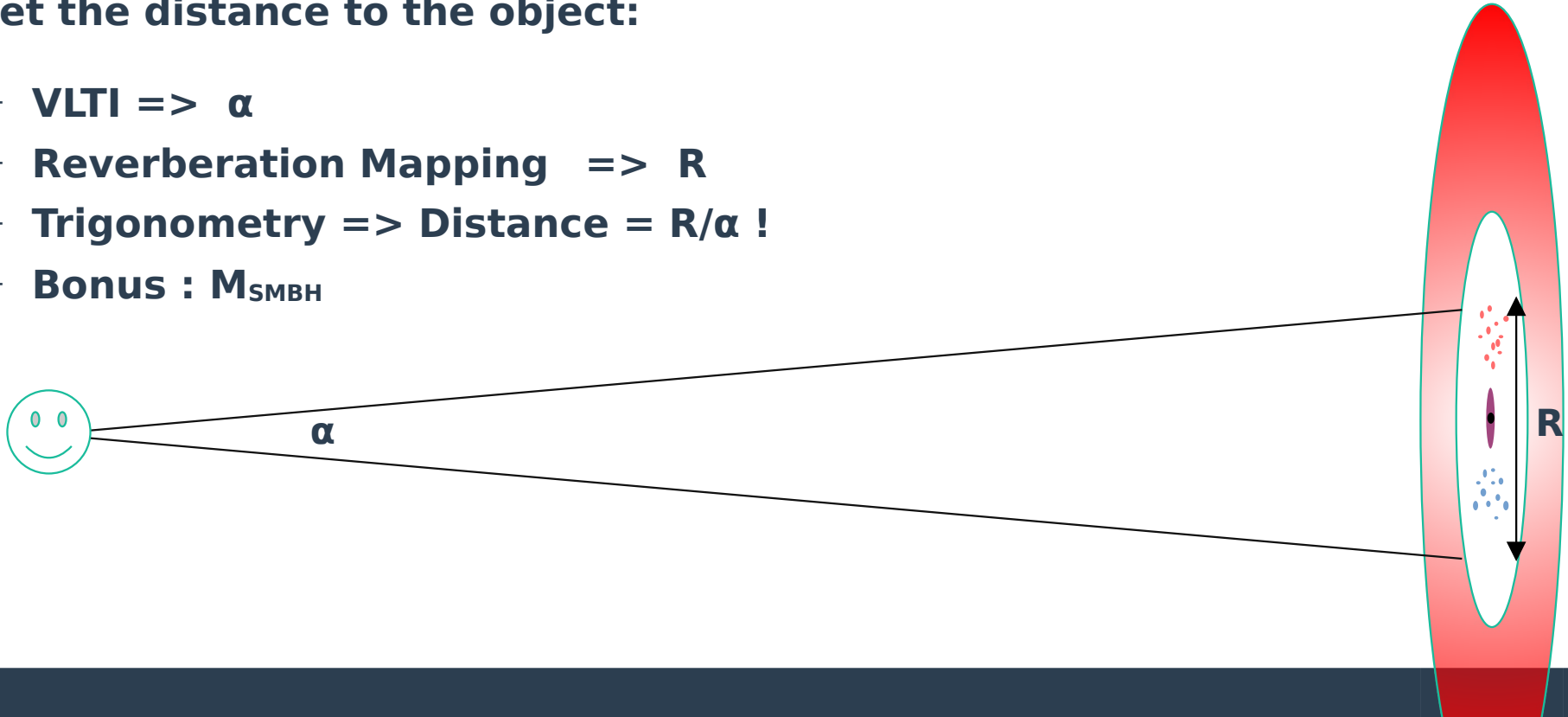
- Spectro-astrometry: α_{BLR} (a few $10 \mu\text{as}$)
- IR visibility: α_{torus} (a few mas)



Interferometry + Reverberation mapping

Knowing the angular size and the physical size we can get the distance to the object:

- VLTi => α
- Reverberation Mapping => R
- Trigonometry => Distance = R/α !
- Bonus : M_{SMBH}



Problems

- **While independent from the distance ladder and the Λ CDM model, there are some problems:**
 - Dependent on the model of the torus/BLR
 - Degeneracy between parameters, in particular i and M_{SMBH}
 - Low number of targets
- **Projet AGN_MELBA**
- **My contribution:**
 - Improve the model of the dusty torus
 - Improve interferometric image reconstructions

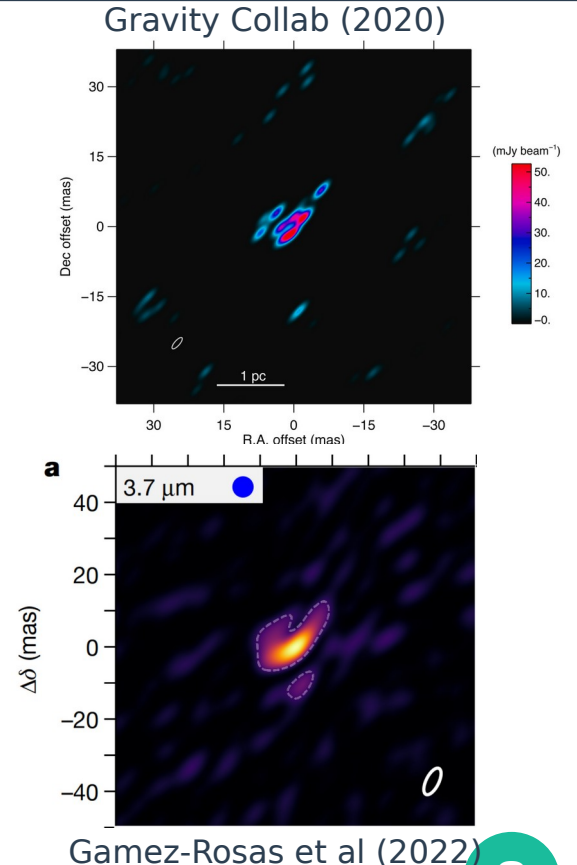
Interferometric image reconstruction

- **Problem:**

- Interferometers give access to very high angular resolution, but do not provide images
- They provide an incomplete and irregular sampling of the Fourier Transform of the images
- Reconstructing the images is a highly degenerate inverse problem with an infinite number of solutions

- **Solution:**

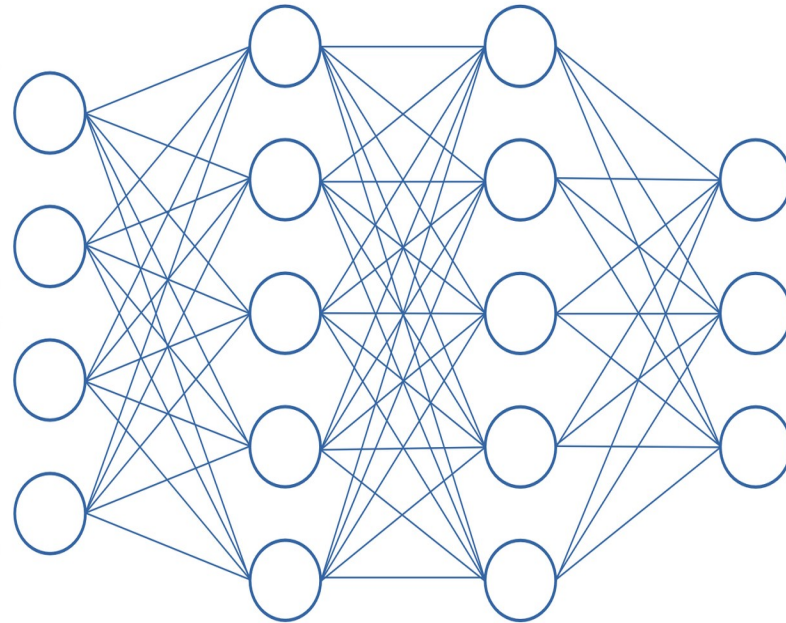
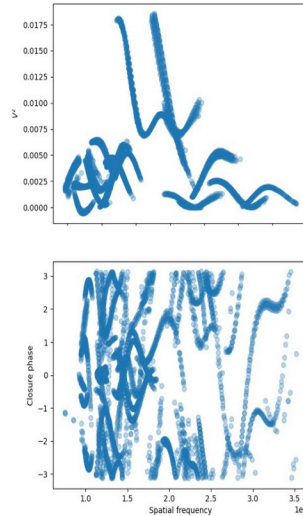
- Include *a priori* knowledge about the geometry of the source, i.e. regularization
 - Find the best compromise between fit to the data and adequation to this regularization term
 - My new method :)



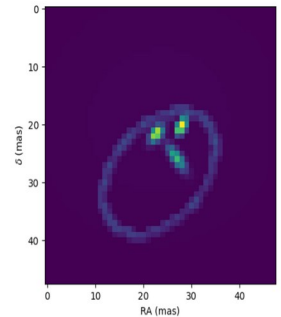
Neural network based image reconstructions

- 1) Create fake astronomical images
- 2) Compute synthetic interferometric observations for each
- 3) Relentlessly train a neural network to reconstruct the original image based on the mock observation
- 4) Apply to the real observation

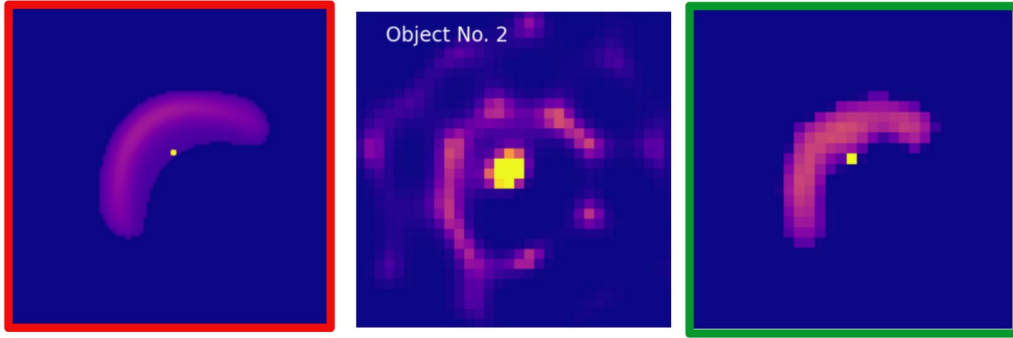
Input:
Interferometric
observables



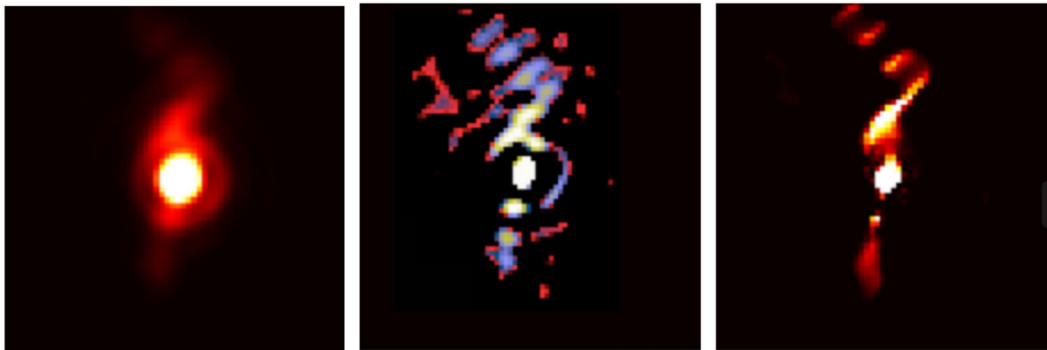
Output:
Images



Promising results

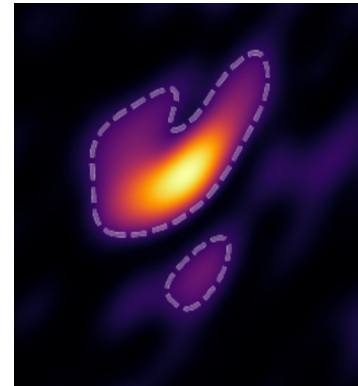


Optical interferometry imaging contest IX, Sanchez et al (2022)

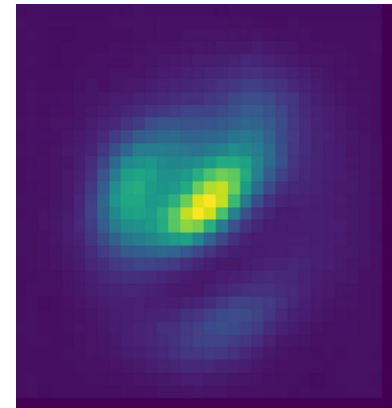


Bonus: Deconvolution of Adaptive Optics images

MATISSE/VLTI image of NGC 1068
(nearby AGN)



Classical image
reconstruction
Gamez-Rosas et al
(2022)



Neural Network
based image
reconstruction

Conclusions

- 1) Improve interferometric image reconstructions
- 2) Use it to improve our knowledge of AGN torii
- 3) Use it to improve SARM measurements of cosmological distances (and SMBH masses)
- 4) Get a precise new H_0 estimate, independent from distance ladder and Λ CDM model

