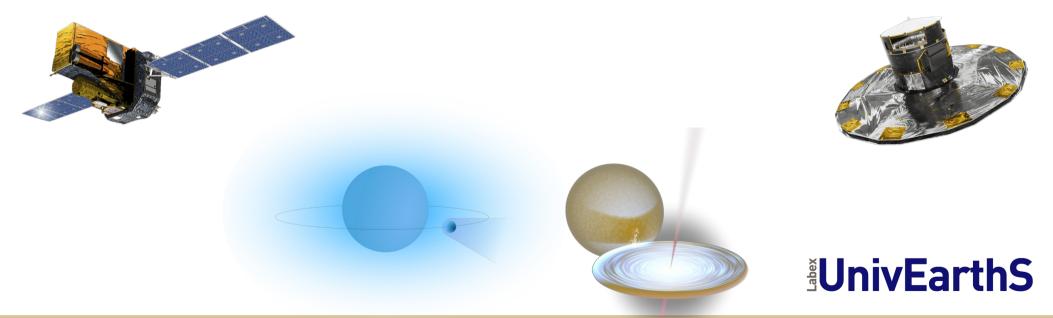






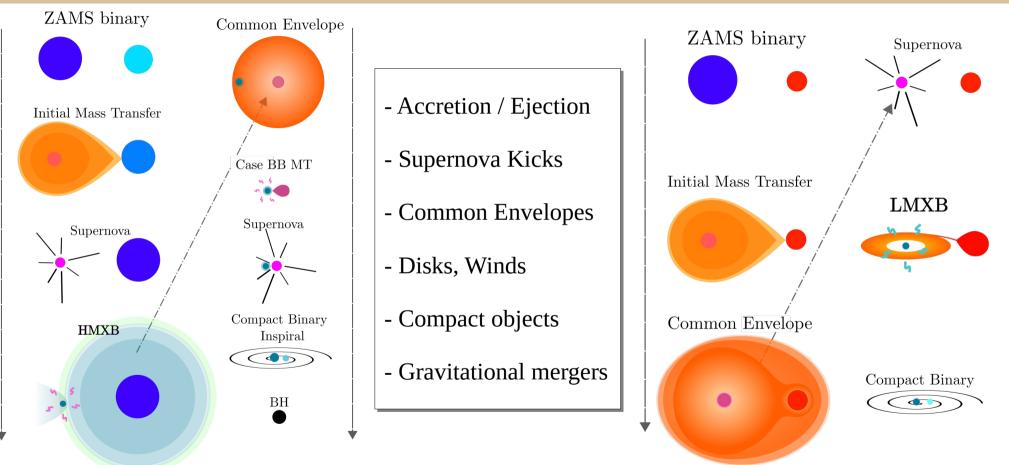
An inventory of Compact objects in X-ray binaries

From the INTEGRAL to the Gaia era



Atelier Ondes Gravitationnelles et Objets Compacts

Evolution of X-ray binaries – Progenitors of GWs ?



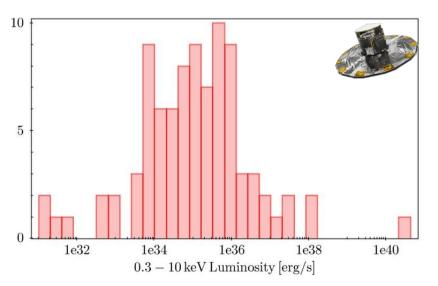
The new catalogue of HMXBs in the Galaxy

Last catalogue of HMXBs : Liu et al. 2006 [N = 114]

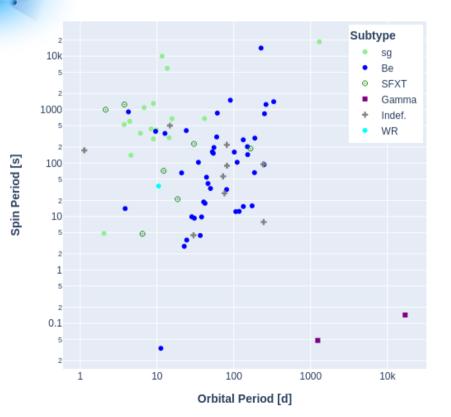
- \rightarrow many new observations since then
- \rightarrow INTEGRAL was just beginning !

New catalogue of HMXBs : Fortin et al. 2023b [N = 160+]

- \rightarrow automated search for multi-wavelength counterparts
- \rightarrow manual search for spectral types, orbital parameters...







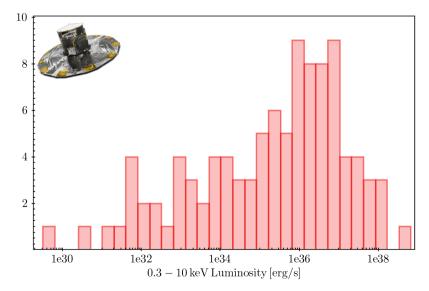
The new catalogue of LMXBs in the Galaxy

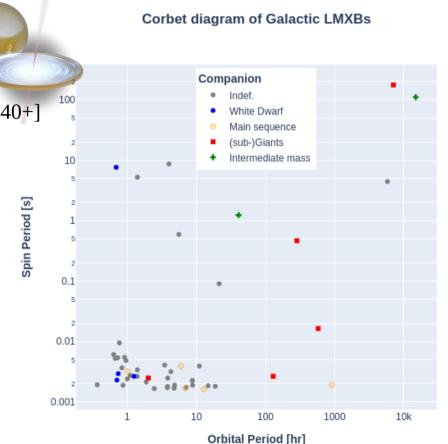
Last catalogue of HMXBs : Liu et al. 2007 [N = 187]

- \rightarrow many new observations since then
- \rightarrow INTEGRAL was just beginning !

New catalogue of HMXBs : Fortin et al. 2023c subm. [N = 340+]

- \rightarrow automated search for multi-wavelength counterparts
- \rightarrow manual search for spectral types, orbital parameters...



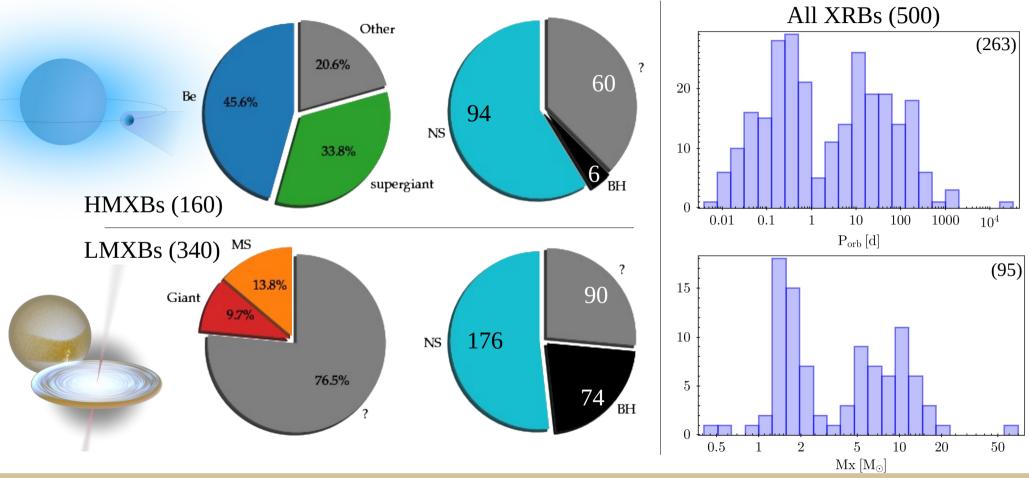


HMXB and LMXB Webcat : participative database

НОМЕ	CATALOGUE NOT	ES DOWNLOADS CI	ONTRIBUTING							AB	BOUT
	HMXB: tip: any identifier know	vn by Simbad will work !							Show: 10 ~		1
	Main ID										
	[Field of View] 1	Spectral type	Class	Right Ascension [®] [J2000]	Declination [J2000] [®]	Error radius 🏨 [mas]	Distance [pc] [®]	Mx [Msun] 14	Mo [Msun]		Pe [d]
		Spectral type BN0.7 lb 2014A&A566A.131G		Ascension 1		radius 🛝				t≱ G	
	[Field of View] IGR J00370+6122	BN0.7 Ib	1 11	Ascension [®] [J2000]	[J2000] ¹¹	radius 🧯 [mas]	[pc] ¹⁰ 3401		[Msun] 22.0		[d

 \rightarrow GitHub/HMXBwebcat

Some statistics on Galactic XRBs



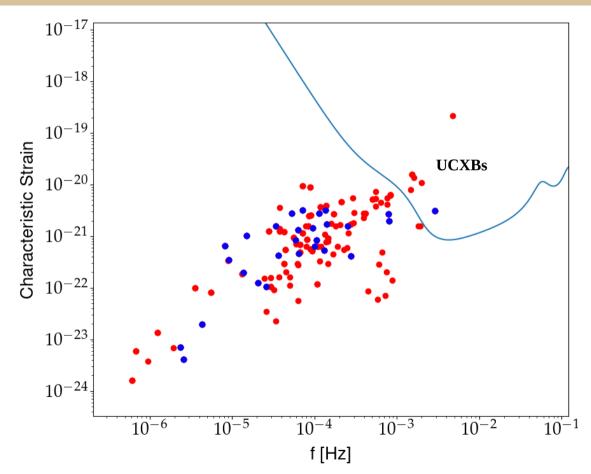
Francis Fortin - IRAP

LMXBs as GW emitters : LISA

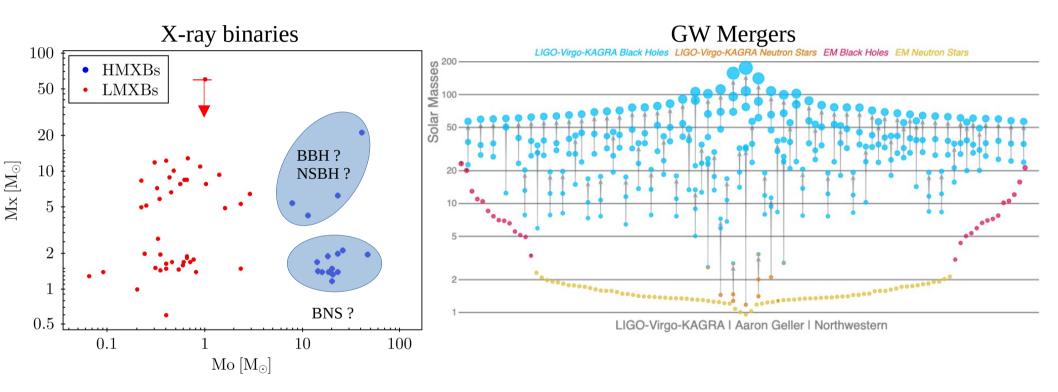
Strain = f(Mx, Mo, Porb, D, Tobs)

Red: filled-in parameters (i.e. 1.4 Msun NS, Galactic Bulge...)

Blue : all parameters available



XRB systems vs. LVK GW progenitors



- → Cyg X-1 : likely NSBH (7% survival after kick), maybe BBH in Hubble time (Neijssel+2021)
- \rightarrow Galactic XRBs vs. far away GW mergers : impact of metallicity ?

Take away messages & questions

- \rightarrow The number of known XRBs in the Galaxy is **500+**.
- → BH in LMXBs : likely biased towards lower BH masses (Jonker+2021). → "hidden" population of BH LMXBs with masses akin to GW BBHs ?
- → GW BBH selection effects + low number of known BH HMXBs : maybe compatible (Fishbach+2022)
- → Different galaxies = different histories: low metallicity allows for Chemically Homogeneous Evolution. → up to 3/4 of the GW BBHs coming from isolated binary evolution (Riley+2021)?

Selection effects ? Use the characteristics of XRBs in the Milky Way to calibrate pop synth models ? Synergy with binary evolution simulations ? How does the MW compares to other galaxies ? Can we probe older populations of massive stars in binaries ?

