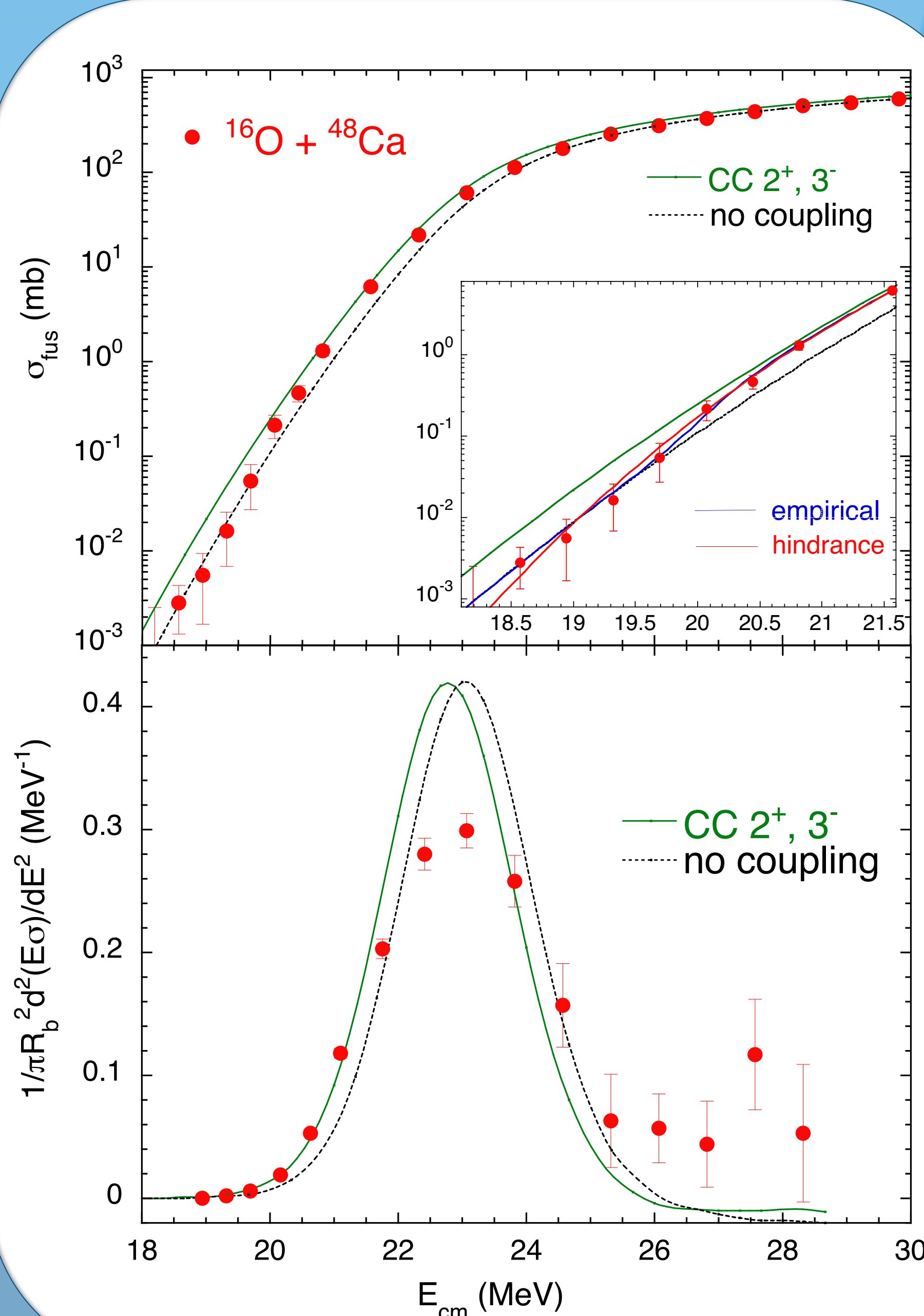


# Low-Energy Fusion of Magic Nuclei: the Case of $^{16}\text{O} + ^{48}\text{Ca}$



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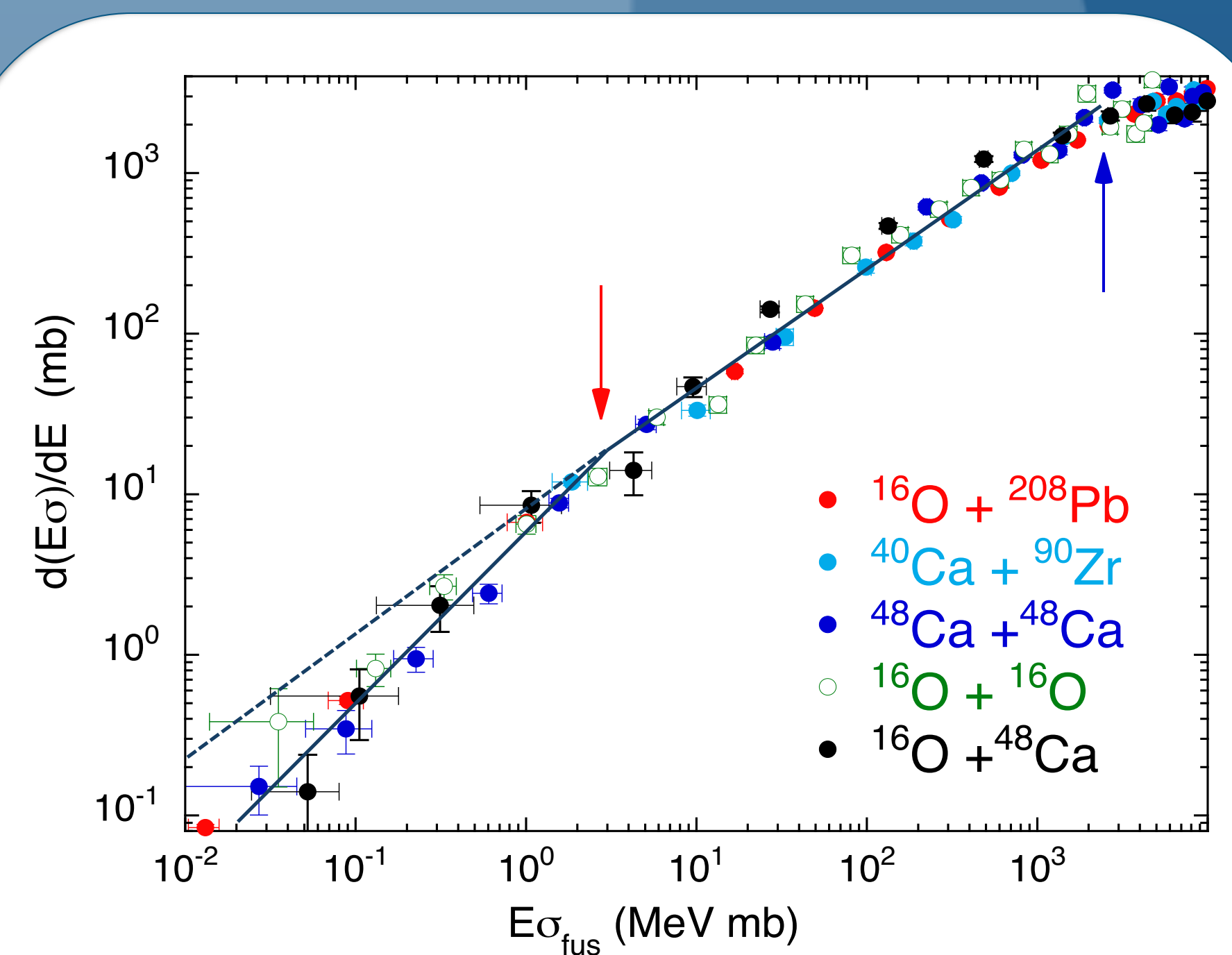
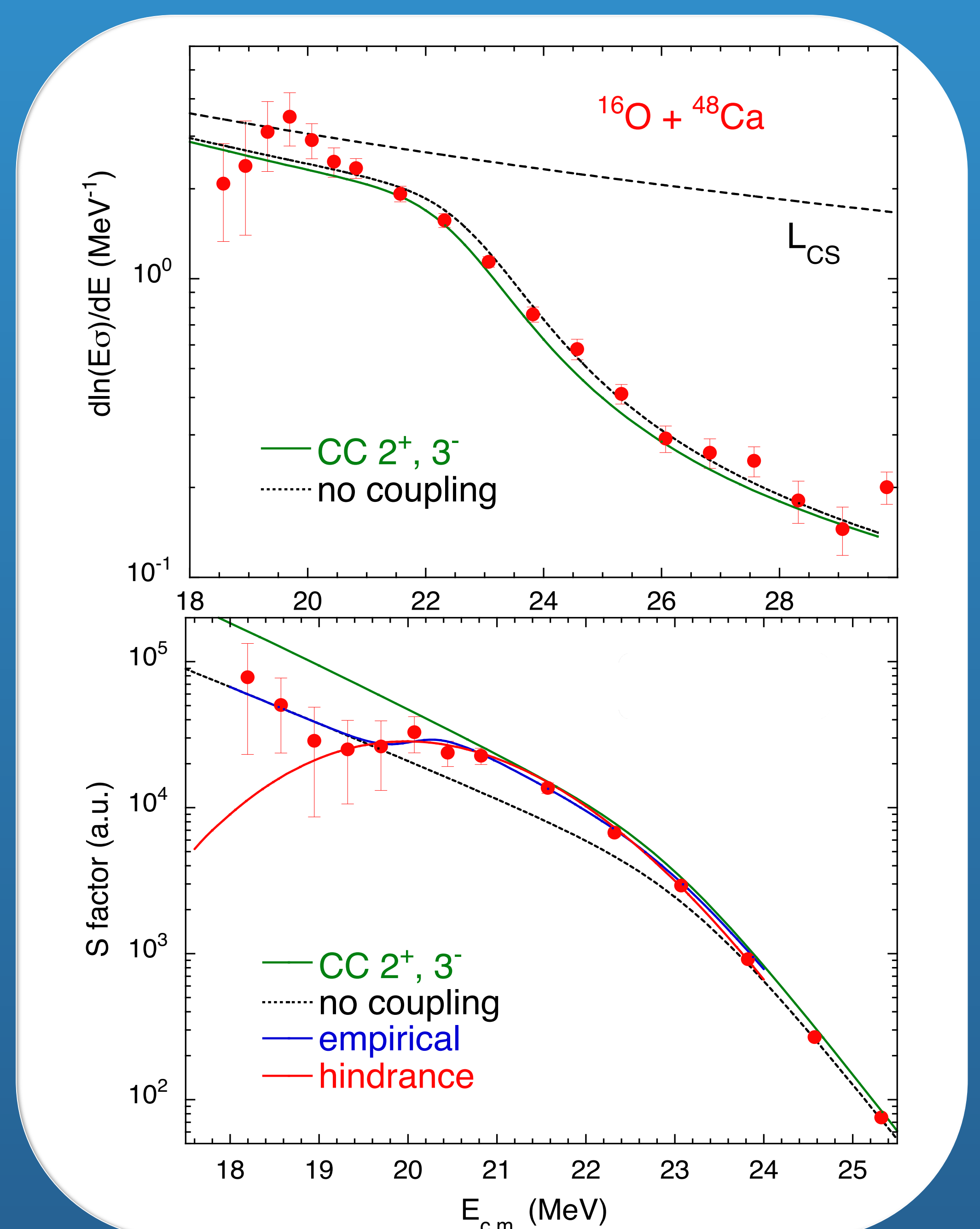
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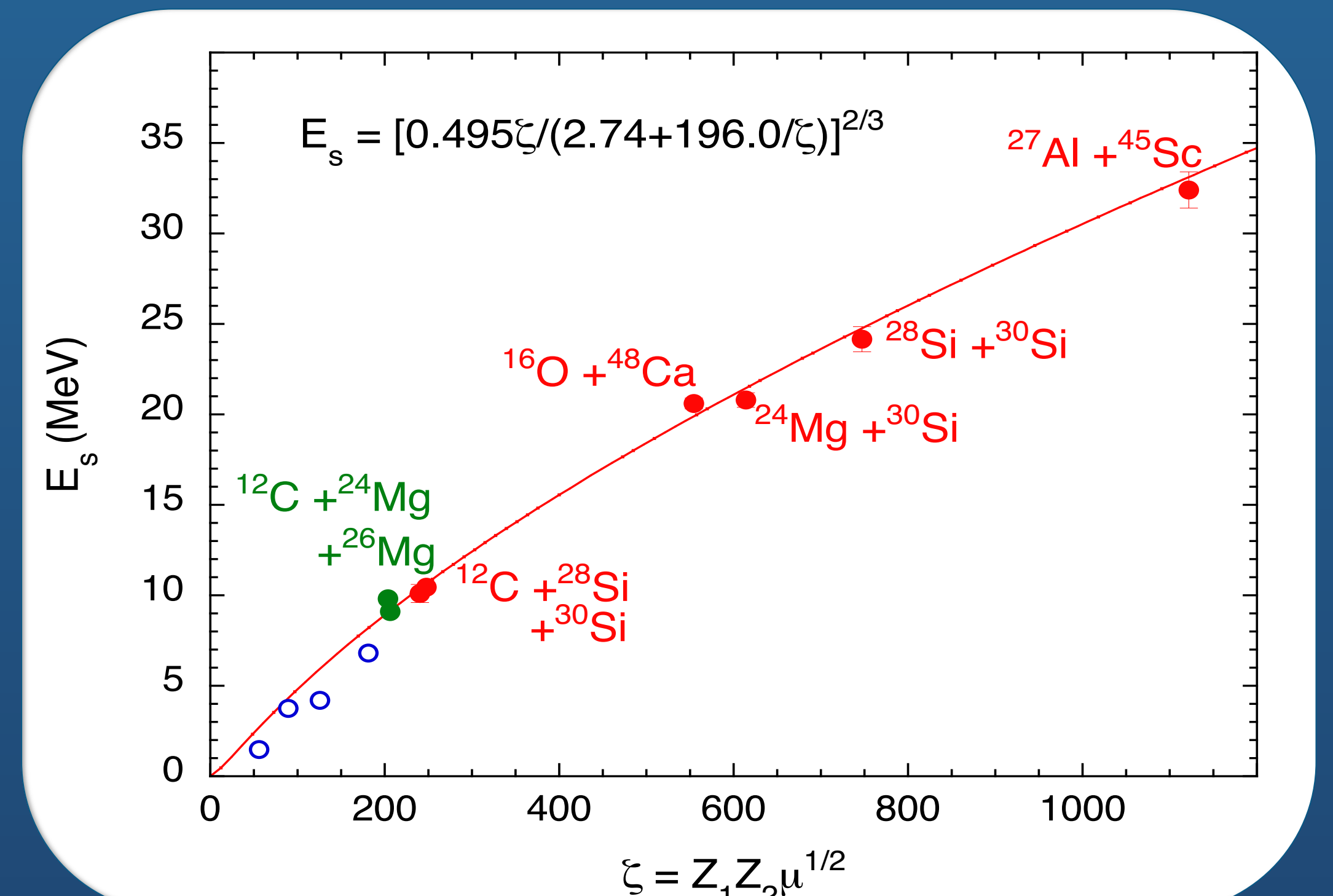
Motivated by the doubly magic nature of the system  $^{16}\text{O} + ^{48}\text{Ca}$ , we have measured [1] its fusion excitation function from above to far below the barrier at the Laboratori Nazionali di Legnaro of INFN. We have used the  $^{16}\text{O}$  beams from the XTU Tandem accelerator. The fusion cross sections were measured down to a few  $\mu\text{b}$  by identifying the evaporation residues in a detector telescope, downstream of an electrostatic beam deflector.

CC calculations with the Akyuz-Winther potential, including the lowest  $2^+$  and  $3^-$  states of  $^{48}\text{Ca}$ , well fit the data down to  $\sigma_{\text{fus}} \approx 0.8 \text{ mb}$ . At lower energies, the hindrance effect shows up. The fusion barrier distribution shows a single main peak.

At lower energies, the data are consistent with pure one-dimensional tunnelling, as observed for  $^{12}\text{C} + ^{24}\text{Mg}$ ,  $^{30}\text{Si}$  [2]. The logarithmic slope reaches the  $L_{\text{CS}}$  value, and the S factor develops a maximum vs energy. The low energy data are well fit by an empirical approach simulating the coupling strength damping (adiabatic model), while the hindrance model fits the S factor maximum but not its increase at the lowest energies.



Doubly-magic systems were previously investigated, and the present case  $^{16}\text{O} + ^{48}\text{Ca}$  does not show significant differences when the various Coulomb barriers are considered. The  $E\sigma$ -range between the red and blue arrows indicates the total width of the barrier distribution.



The phenomenological systematics proposed for heavier stiff systems several years ago [3], have required adjusting the fit parameters (see figure), leading to updated hindrance predictions for the light systems of astrophysical interest.

[1] A.M. Stefanini et al., to be published

[2] G. Montagnoli et al., J. Phys. G 49, 195101 (2022); Phys. Rev. C 97, 024610 (2018)

[3] C.L. Jiang et al., Phys. Rev. C 79, 044601 (2009)