

# Evolution of the pygmy dipole resonance in the Sn mass region studied with the Oslo method

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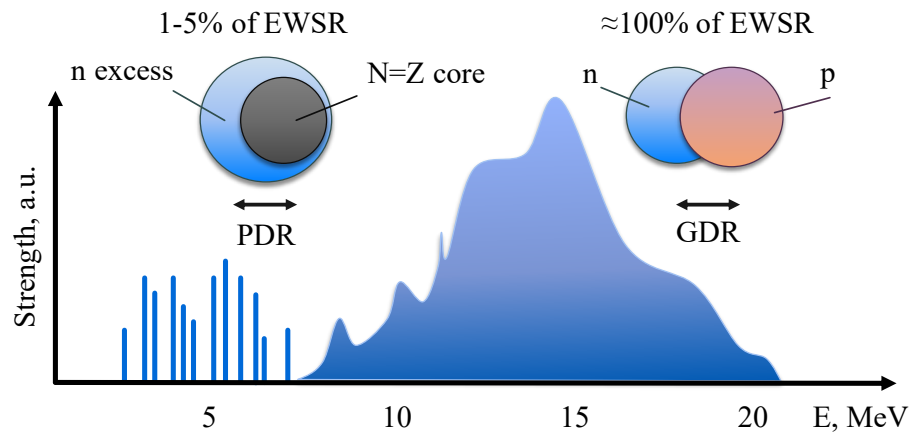
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# The pygmy dipole resonance in nuclei



- ▶ Macroscopic interpretation of the **IVGDR** originating from collective, out-of-phase oscillations of protons against neutrons (enhancement of  $\approx 100\%$  of the EWSR) (**Isovector mode**).
- ▶ The **PDR** is associated with the low-lying  $E1$  strength in the vicinity of the neutron threshold (a few % of the EWSR).
- ▶ Within the macroscopic picture: the PDR is generated by oscillations of a neutron excess (skin) against an isospin saturated core ( $N = Z$ ) (**Isoscalar + isovector modes**).

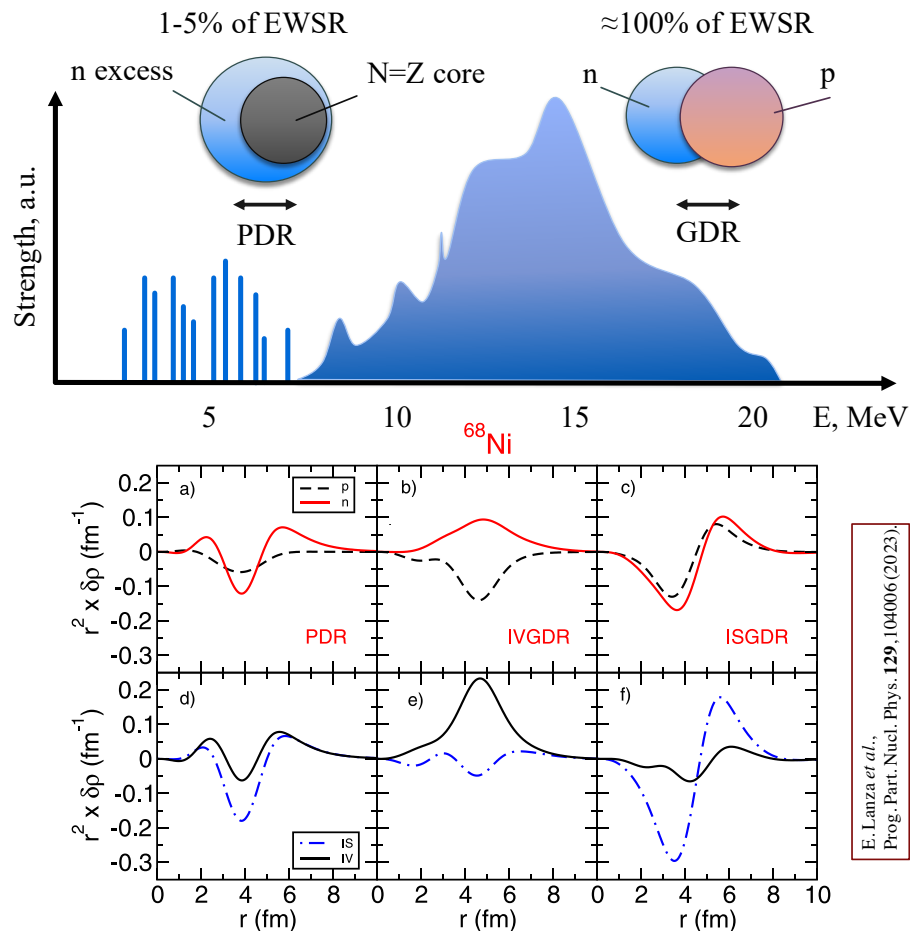


What is the nature of the PDR?

How does it evolve in different isotopic chains?



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Proton, neutron, isoscalar, and isovector transition densities for the dipole states for  $^{68}\text{Ni}$  nucleus (RPA).

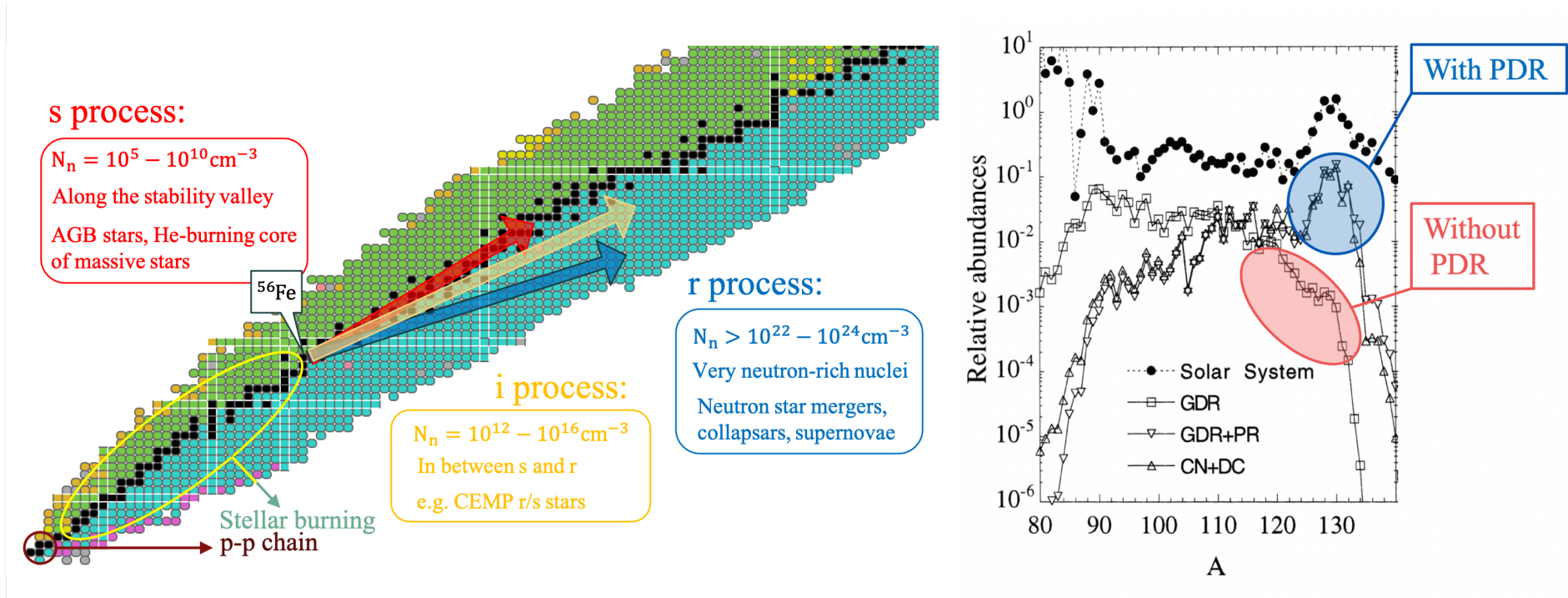
E. Lanza *et al.*,  
Prog. Part. Nucl. Phys. **129**, 104006 (2023).





# Motivation: Why are we interested in the PDR?

- ▶ Relation of the PDR to the **neutron skin thickness** → information on neutron stars?
- ▶ Influence of the PDR on neutron capture rates and resulting **abundances in the r-process**.
- ▶ Appearance of the PDR **increases probability of the  $(n, \gamma)$  reaction**.



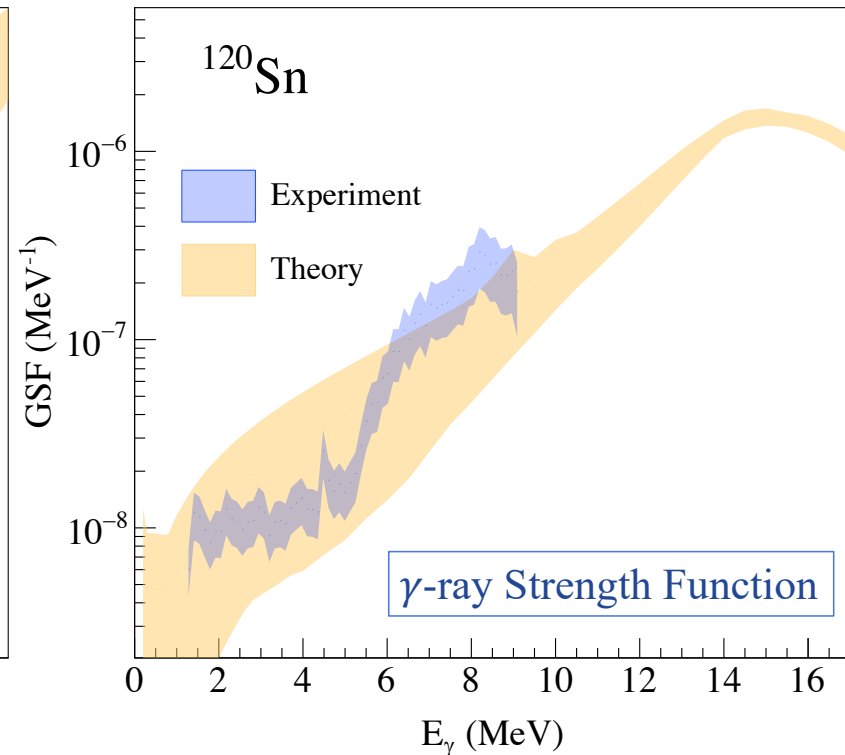
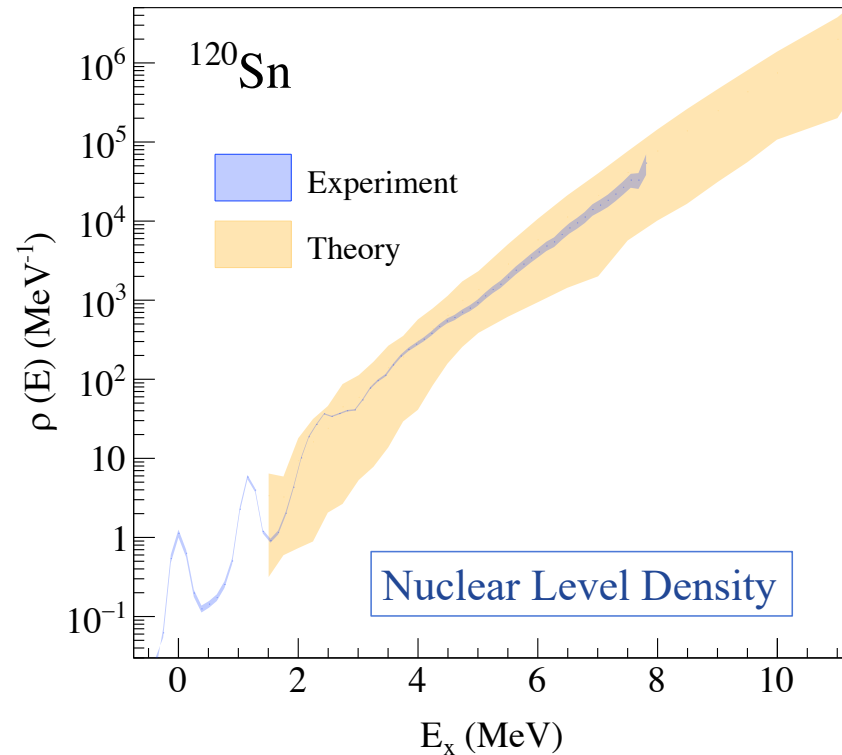
**Left part:** Probable pathways of nucleosynthesis process. **Right part:** Abundances of elements produced in the r-process  
S. Goriely, Phys. Lett. B 436 (1998).





# GSF and NLD: The Oslo method as a method of choice

- ▶ Both the **NLD** and the **GSF** are important ingredients of statistical model calculations.
- ▶ The **Oslo method** is commonly used for a simultaneous extraction of these nuclear characteristics.
- ▶ There is a plethora of theoretical approaches, and systematic experimental constraints are highly demanded.

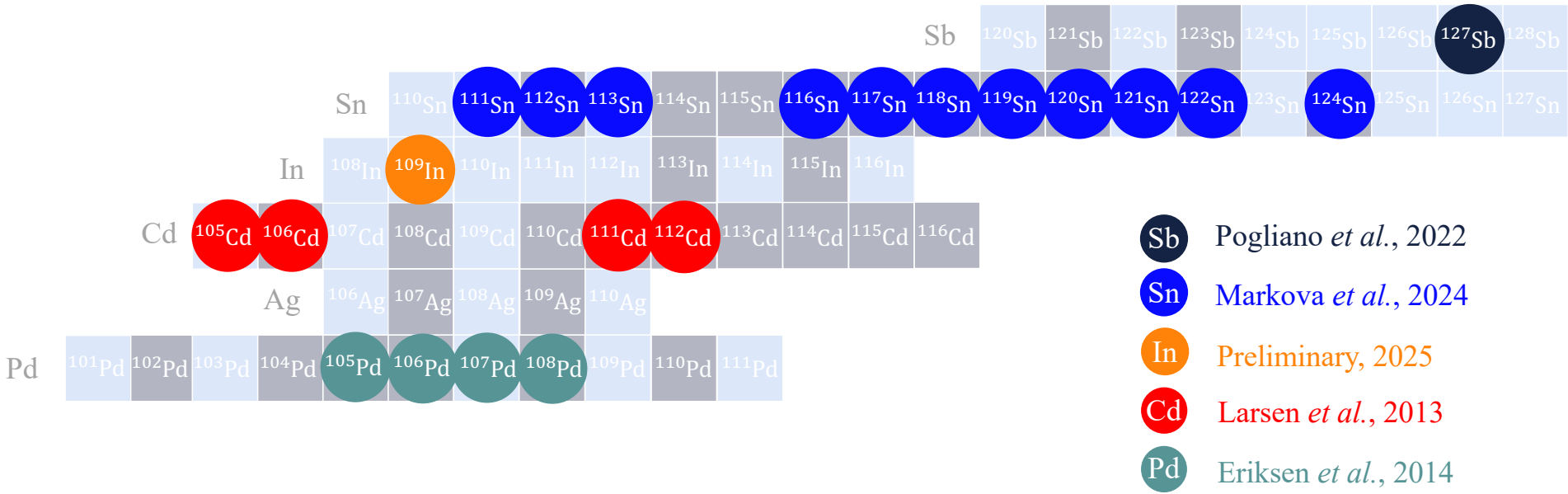


# Isotopes in the Sn region studied at the OCL

											Sb											
											<sup>120</sup> Sb	<sup>121</sup> Sb	<sup>122</sup> Sb	<sup>123</sup> Sb	<sup>124</sup> Sb	<sup>125</sup> Sb	<sup>126</sup> Sb	<sup>127</sup> Sb	<sup>128</sup> Sb			
				Sn	<sup>110</sup> Sn	<sup>111</sup> Sn	<sup>112</sup> Sn	<sup>113</sup> Sn	<sup>114</sup> Sn	<sup>115</sup> Sn	<sup>116</sup> Sn	<sup>117</sup> Sn	<sup>118</sup> Sn	<sup>119</sup> Sn	<sup>120</sup> Sn	<sup>121</sup> Sn	<sup>122</sup> Sn	<sup>123</sup> Sn	<sup>124</sup> Sn	<sup>125</sup> Sn	<sup>126</sup> Sn	<sup>127</sup> Sn
				In	<sup>108</sup> In	<sup>109</sup> In	<sup>110</sup> In	<sup>111</sup> In	<sup>112</sup> In	<sup>113</sup> In	<sup>114</sup> In	<sup>115</sup> In	<sup>116</sup> In									
		Cd	<sup>105</sup> Cd	<sup>106</sup> Cd	<sup>107</sup> Cd	<sup>108</sup> Cd	<sup>109</sup> Cd	<sup>110</sup> Cd	<sup>111</sup> Cd	<sup>112</sup> Cd	<sup>113</sup> Cd	<sup>114</sup> Cd	<sup>115</sup> Cd	<sup>116</sup> Cd								
					Ag	<sup>106</sup> Ag	<sup>107</sup> Ag	<sup>108</sup> Ag	<sup>109</sup> Ag	<sup>110</sup> Ag												
Pd	<sup>101</sup> Pd	<sup>102</sup> Pd	<sup>103</sup> Pd	<sup>104</sup> Pd	<sup>105</sup> Pd	<sup>106</sup> Pd	<sup>107</sup> Pd	<sup>108</sup> Pd	<sup>109</sup> Pd	<sup>110</sup> Pd	<sup>111</sup> Pd											

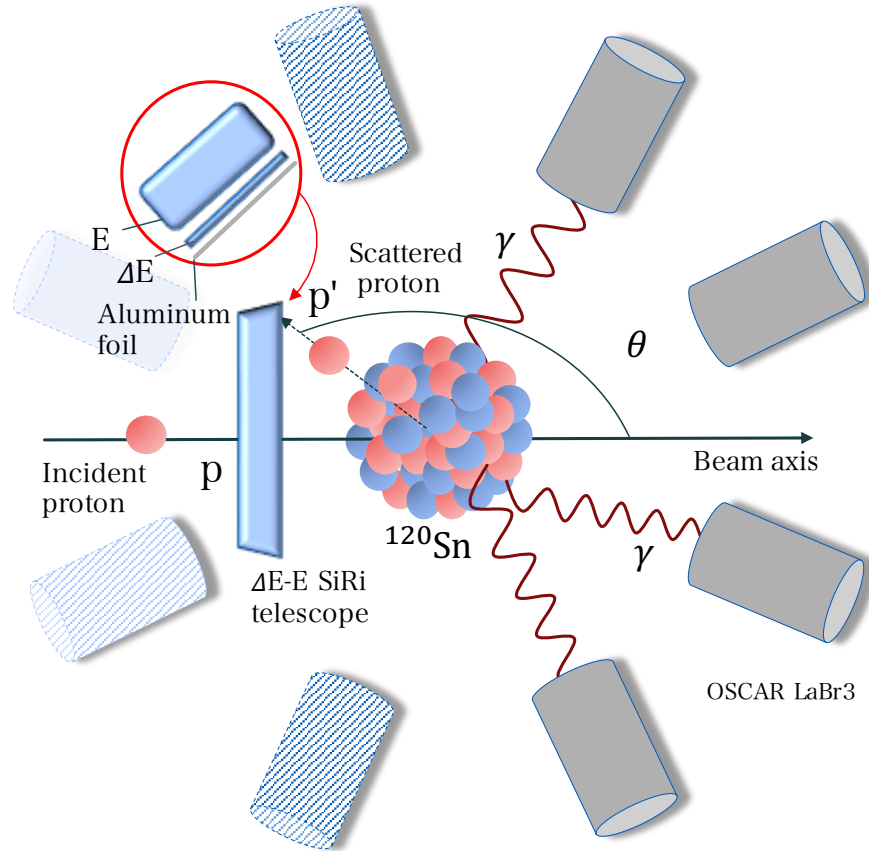


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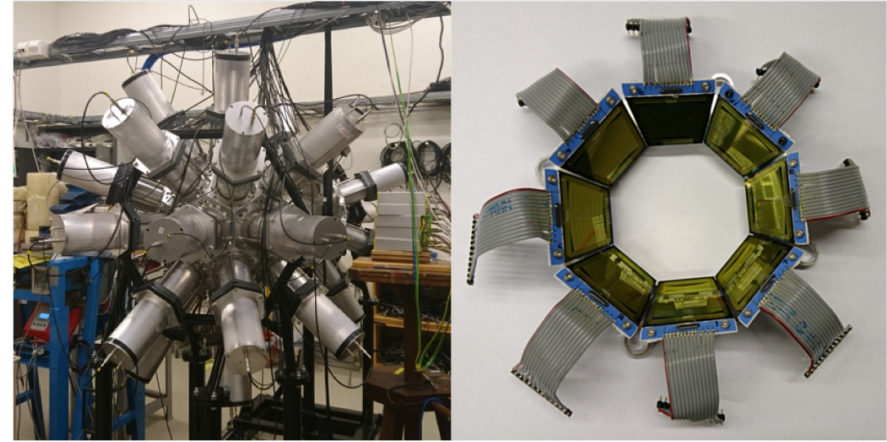




# Oslo method: Experiments at the Oslo Cyclotron Laboratory



The principal scheme of the experiment.



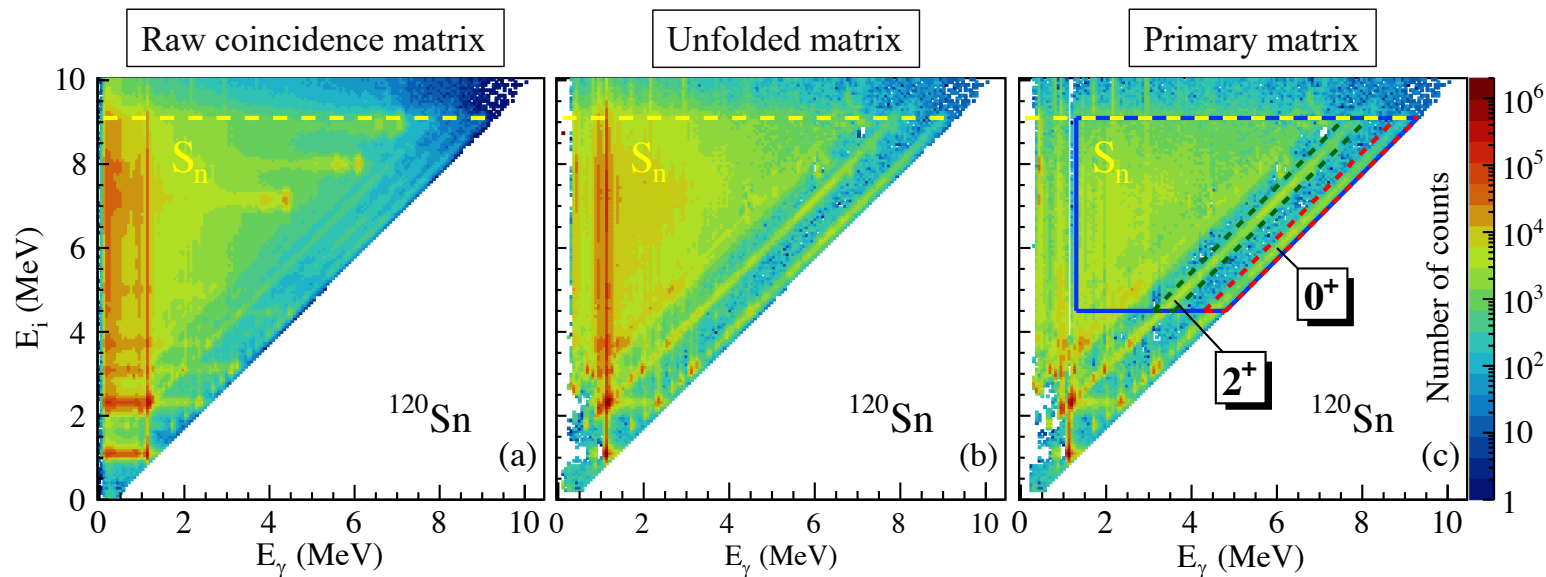
**Left:** OSCAR LaBr<sub>3</sub>:Ce detector array. **Right:** SiRi particle telescope.

<sup>105–108</sup>Pd, <sup>105,106,111,112</sup>Cd, <sup>109</sup>In, <sup>111–113,116–122,124</sup>Sn,  
<sup>127</sup>Sb:

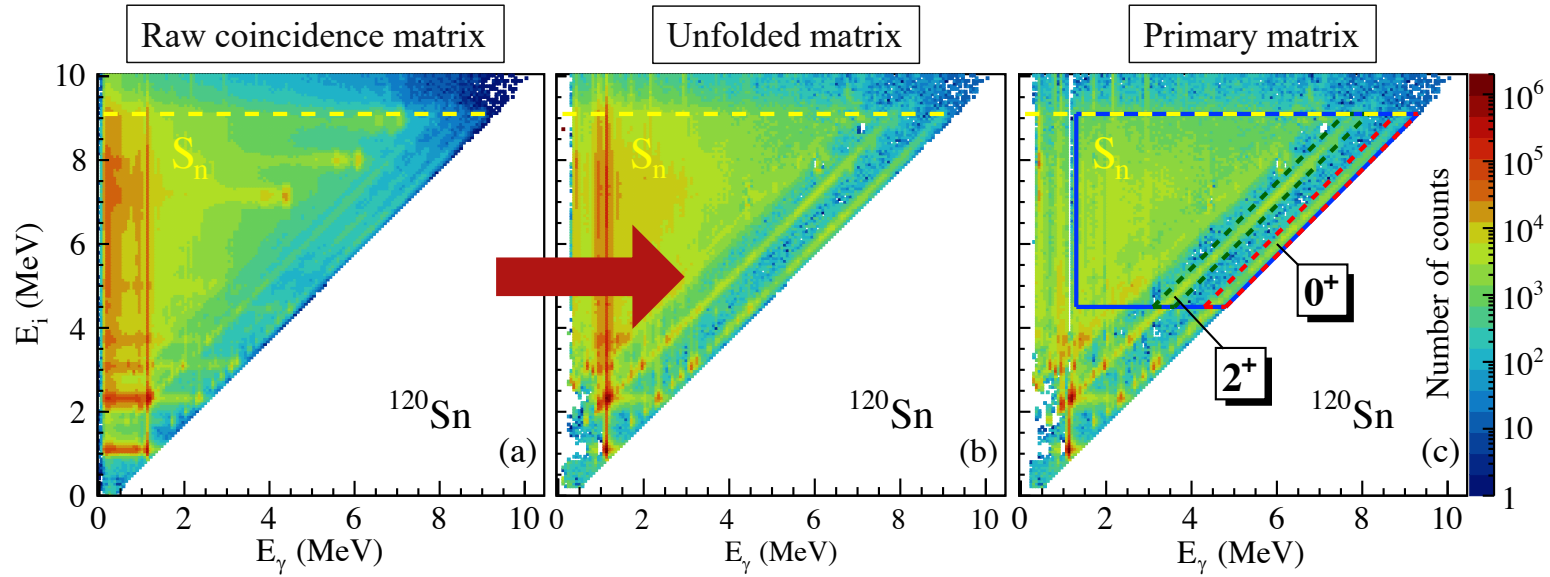
- ▶ SiRi Si particle telescope + NaI(Tl) (older)/ LaBr<sub>3</sub>(Ce) (new) detector array.
- ▶ 126°–140° particle angles are covered.
- ▶ ( $p, p'\gamma$ ), ( $p, d\gamma$ ), ( $d, p\gamma$ ), ( $^3\text{He}, ^3\text{He} \gamma$ ), ( $^3\text{He}, \alpha\gamma$ ), ( $\alpha, p\gamma$ ) reactions with 16 and 25 MeV proton, 11.5 MeV deuteron, 38 MeV  $^3\text{He}$ , and 24 MeV  $\alpha$  beams.
- ▶ particle- $\gamma$  coincidences were extracted.



# The Oslo method: Step-by-step

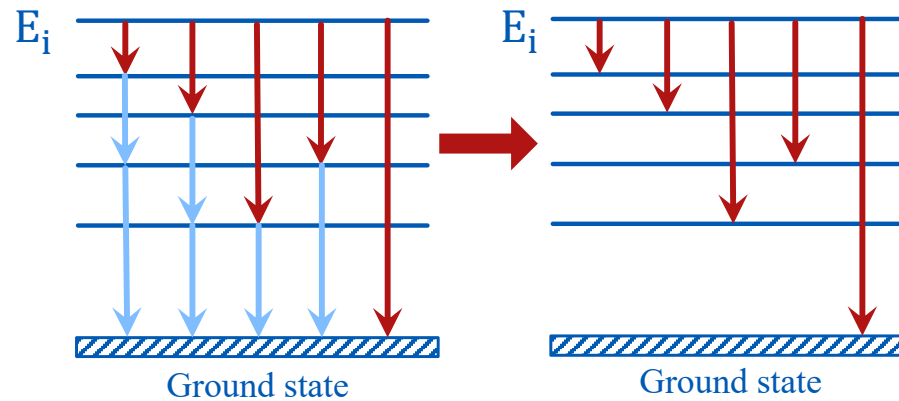
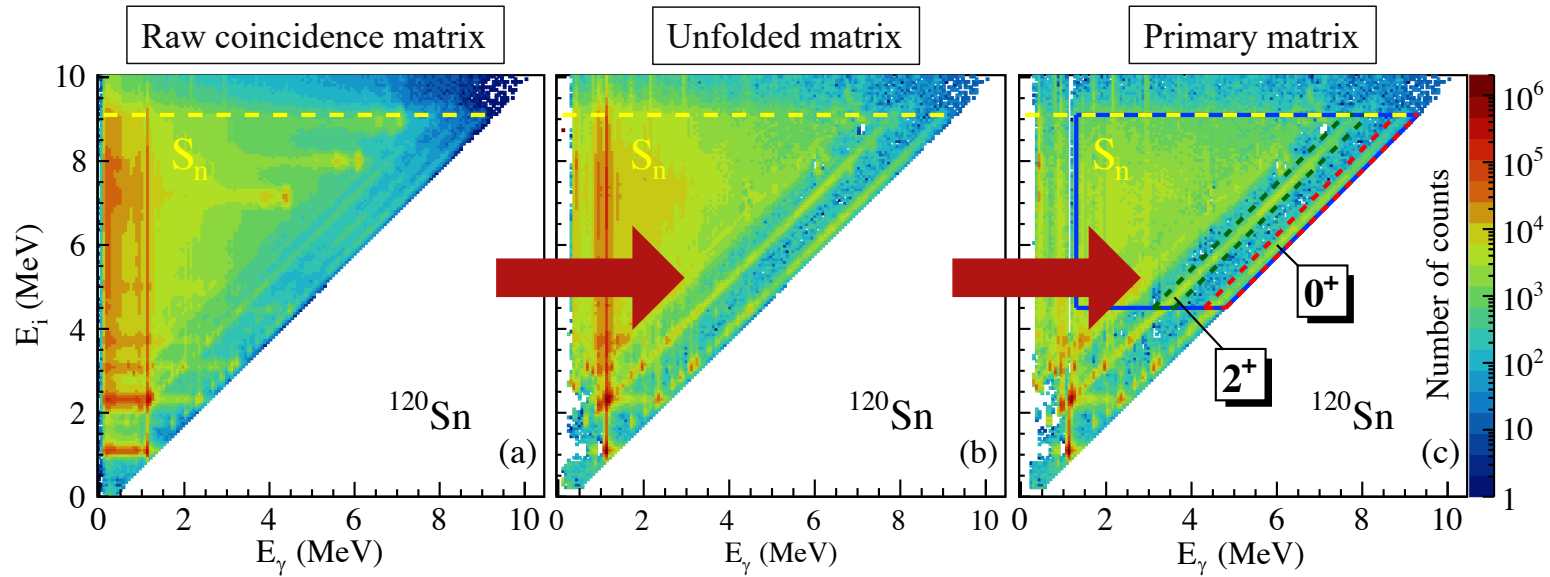


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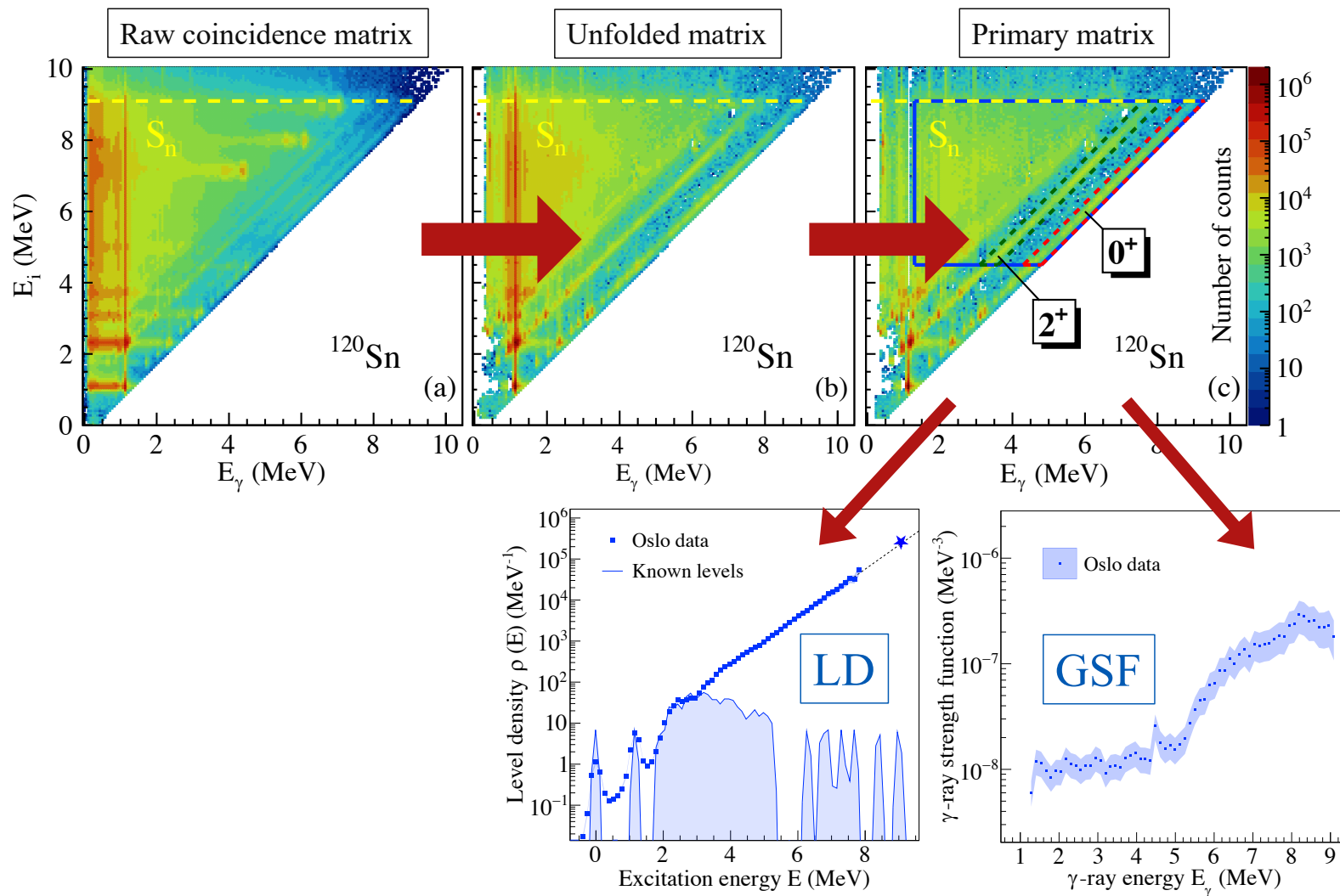




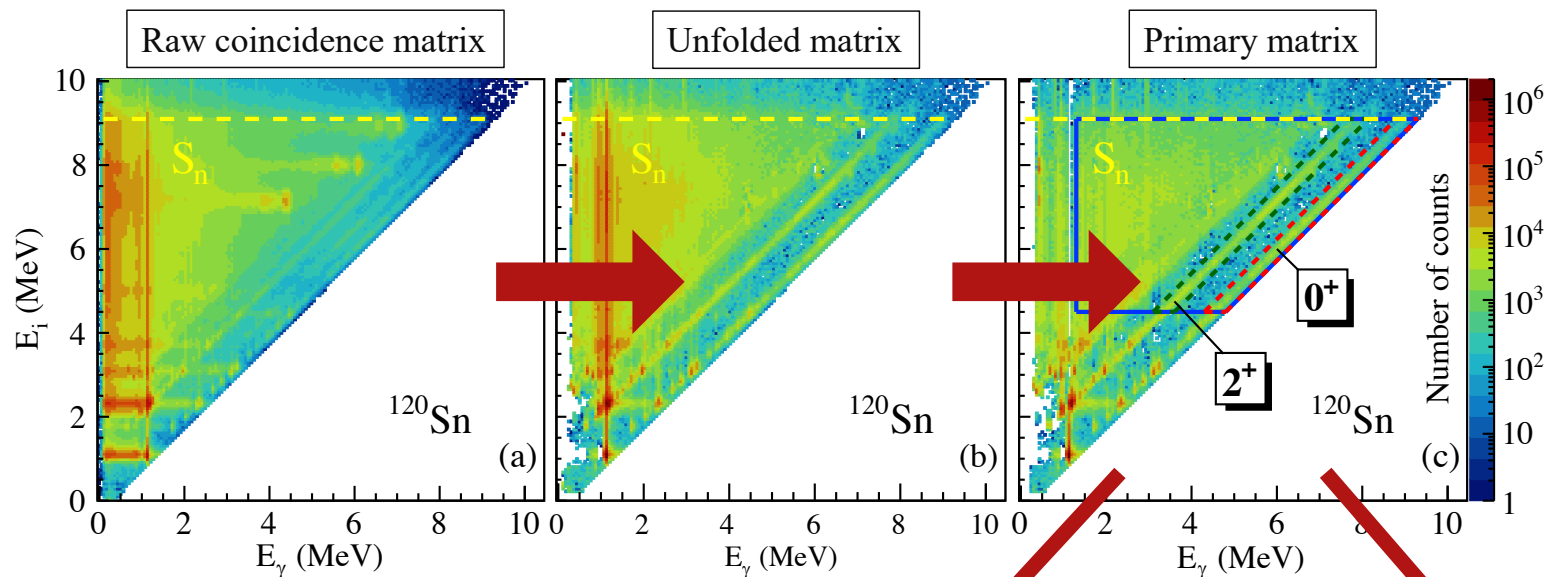
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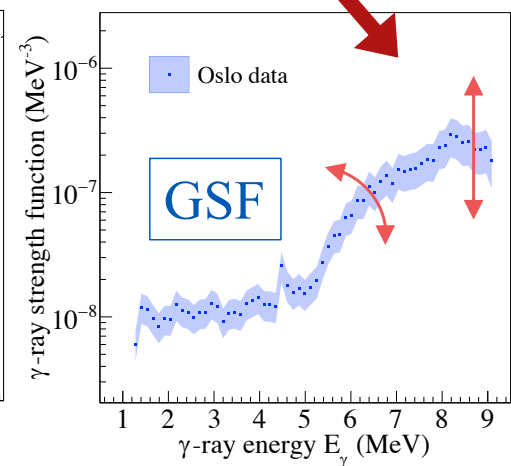
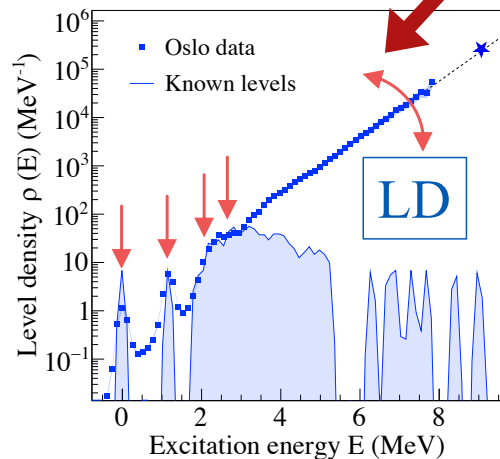


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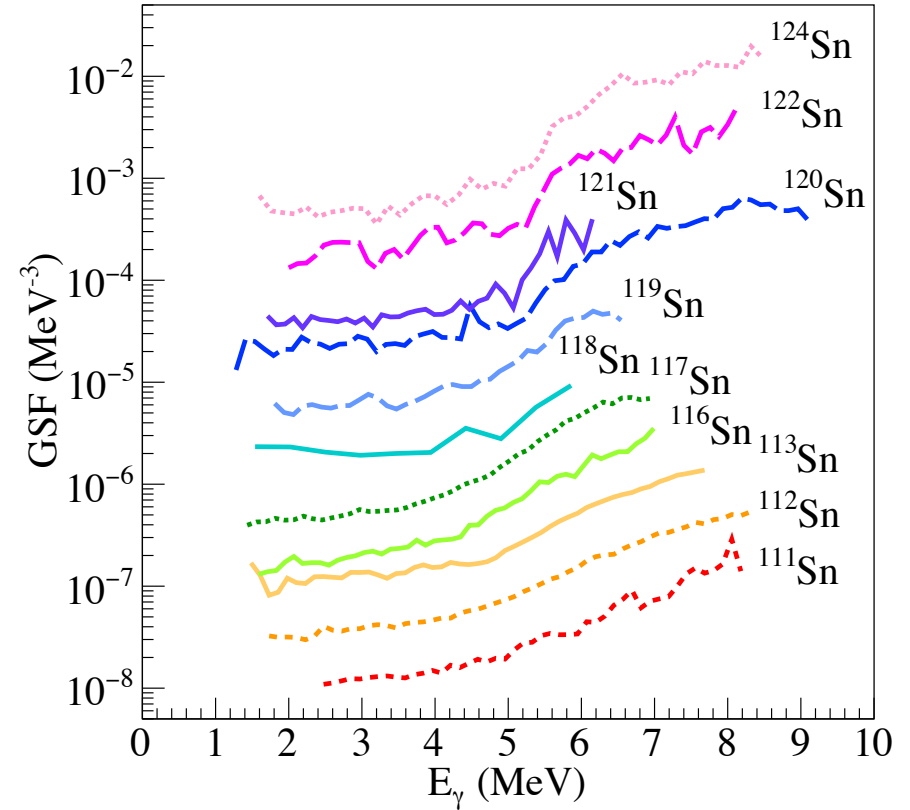
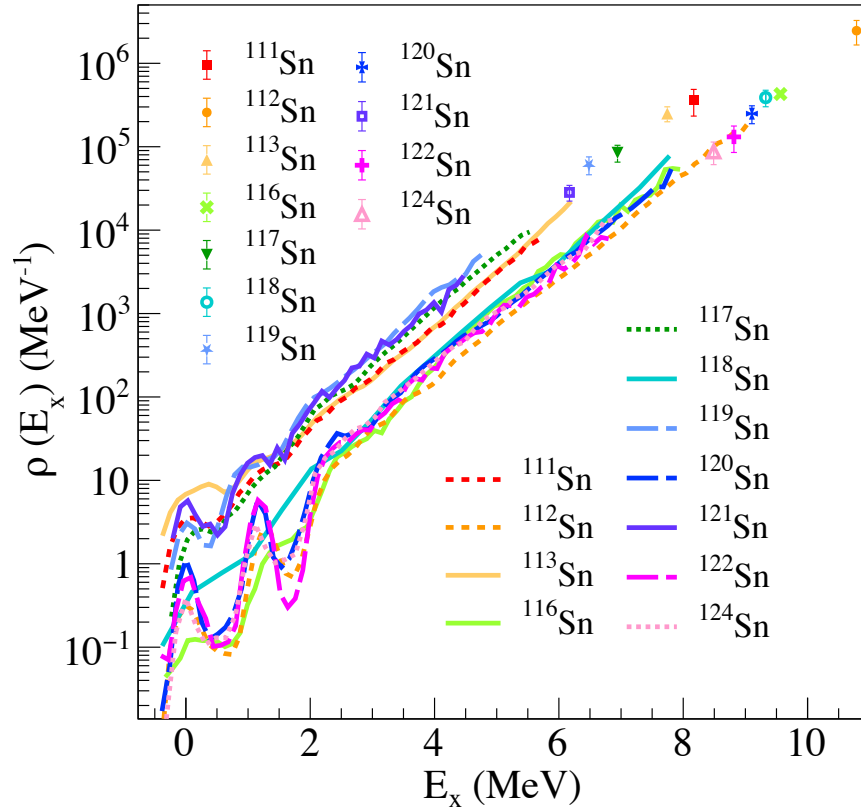
## Normalization:

- **Scaling** the NLD to discrete low-lying states.
- Extracting the NLD and GSF **slope** from neutron resonance data.
- **Scaling** the GSF to the neutron resonance data.

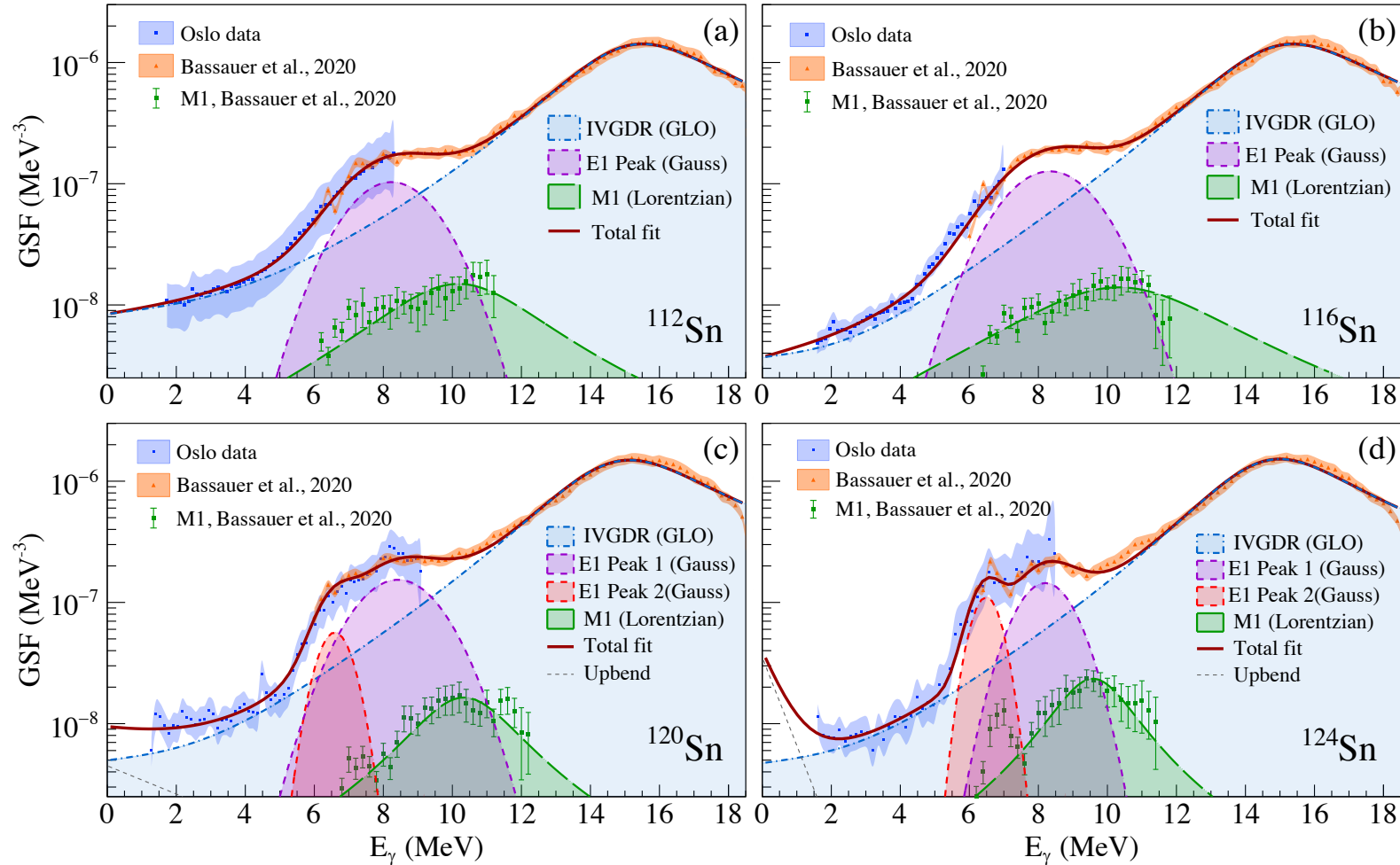




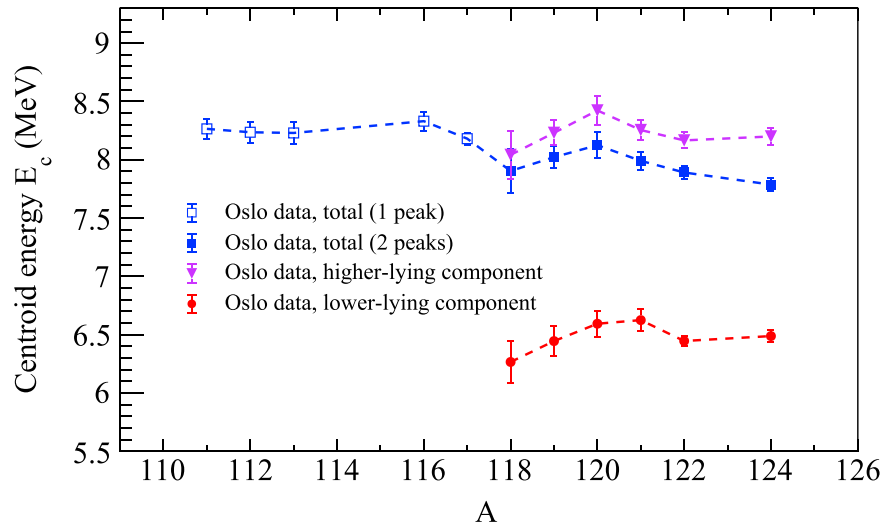
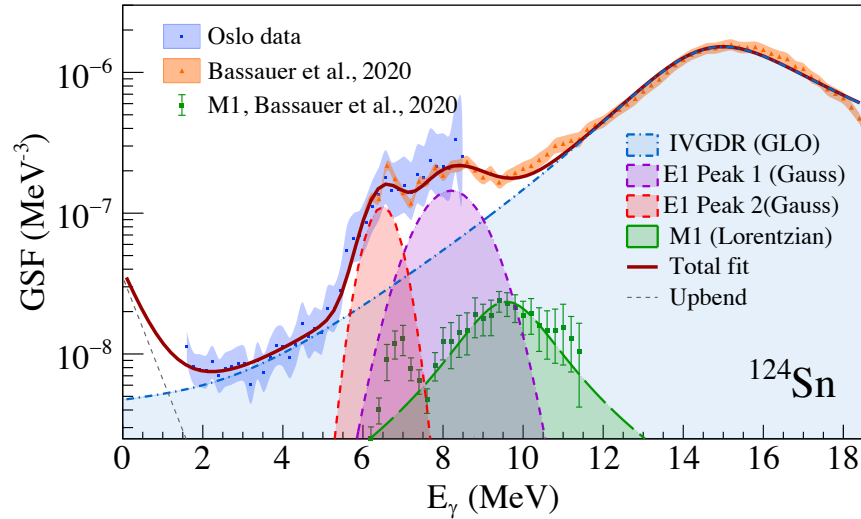
# NLDs and GSFs of Sn isotopes



# Evolution of the low-lying $E1$ strength

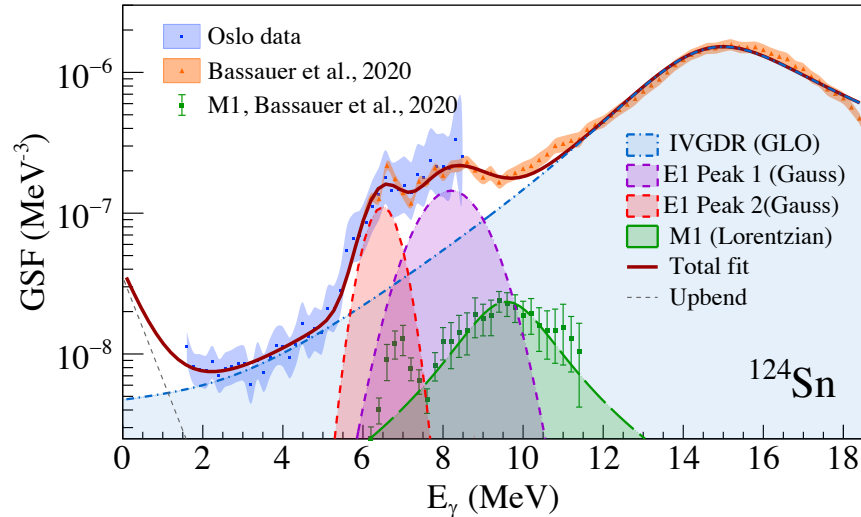


# Main results: Evolution of the low-lying strength



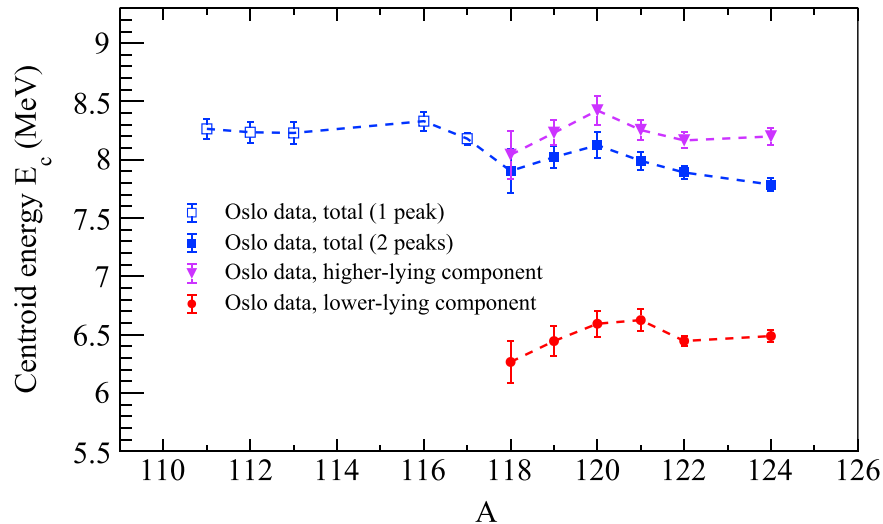


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## Conclusion 1:

The low-lying strength is centered at  $\approx 8.4$  MeV in all studied Sn isotopes.

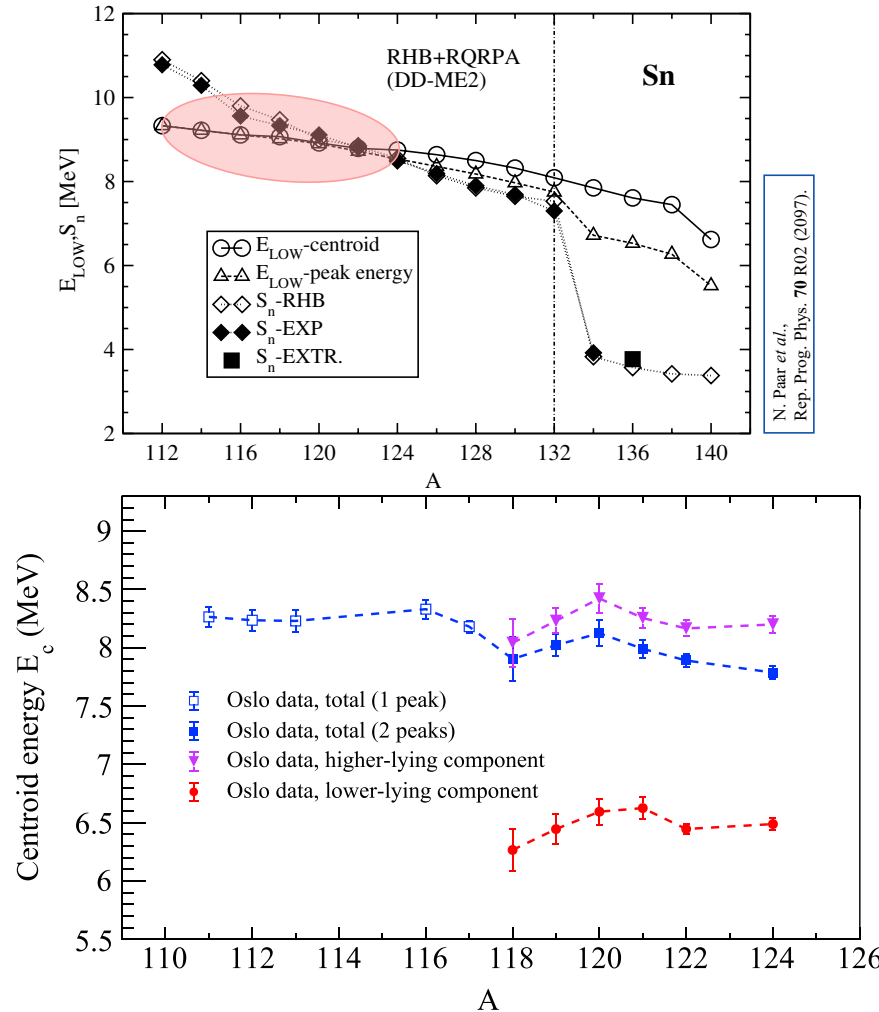


## Conclusion 2:

The low-lying component remains at  $\approx 6.5$  MeV in heaviest tins.



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N. Paar *et al.*,  
Rep. Prog. Phys. **70** R02 (2007).

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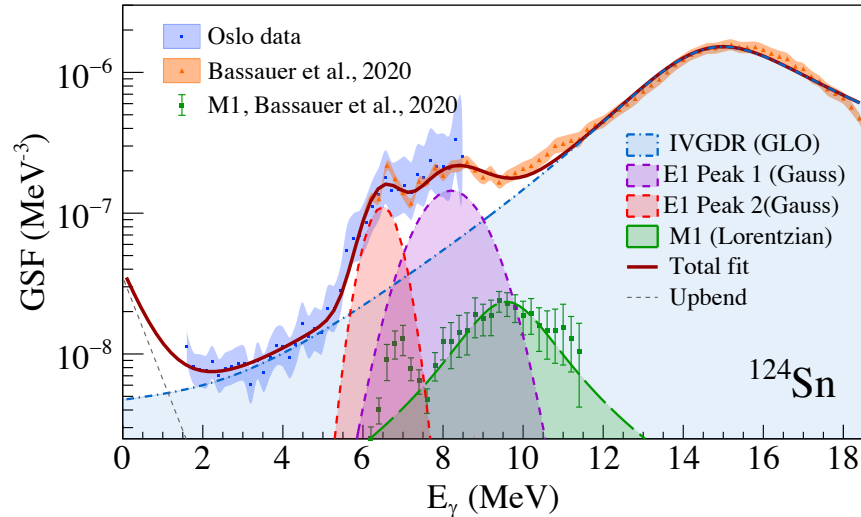
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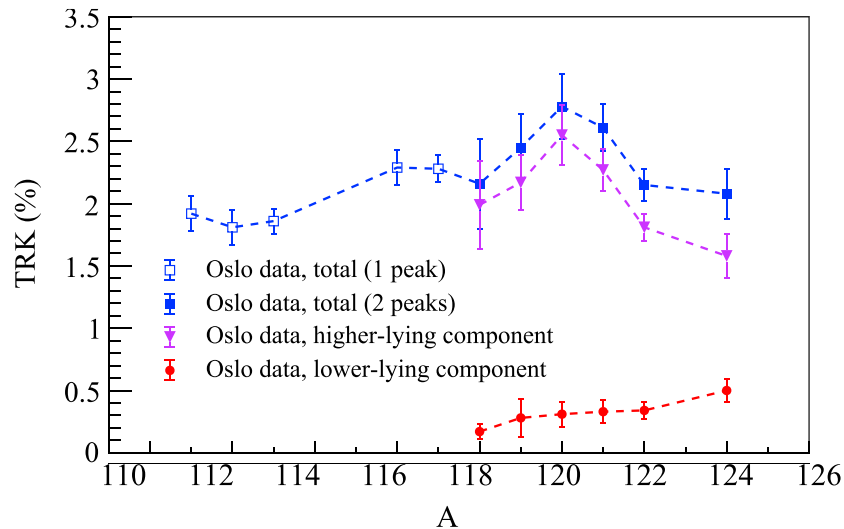


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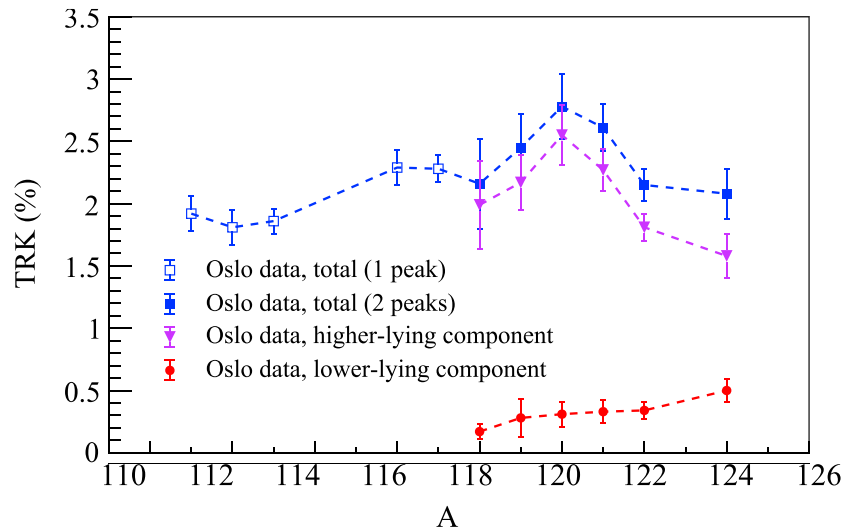
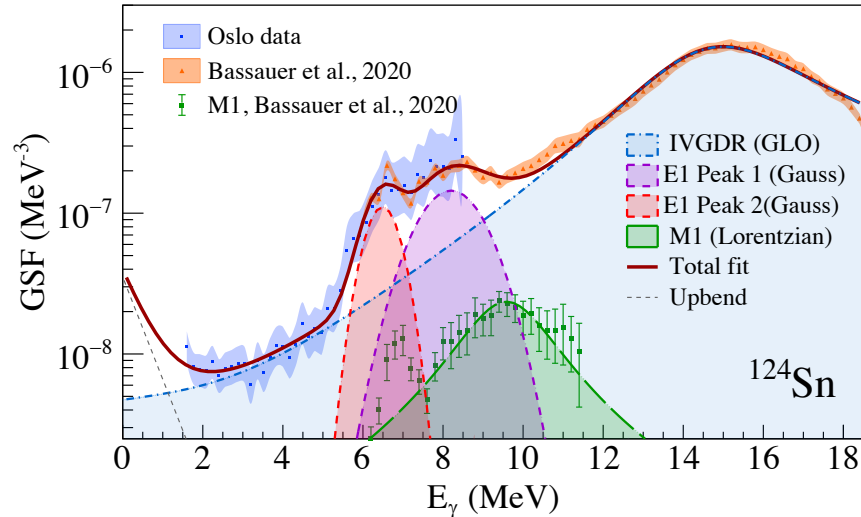


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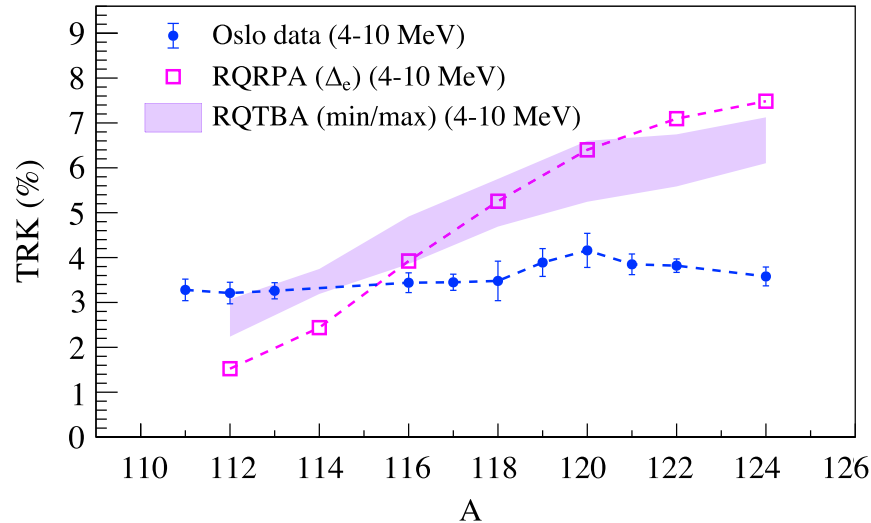
Ranges from 2% to 3%, the largest strength in  $^{120}\text{Sn}$ .

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The low-lying component increases in strength with  $N$ , the IV component of PDR?



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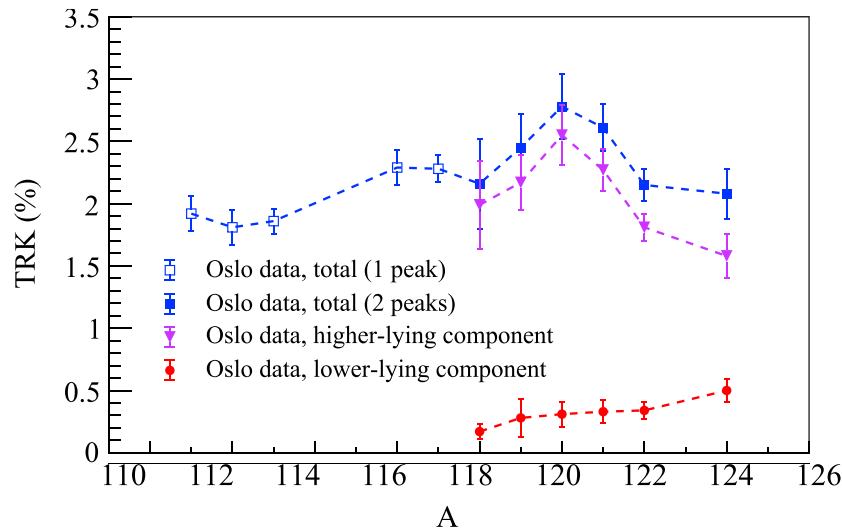
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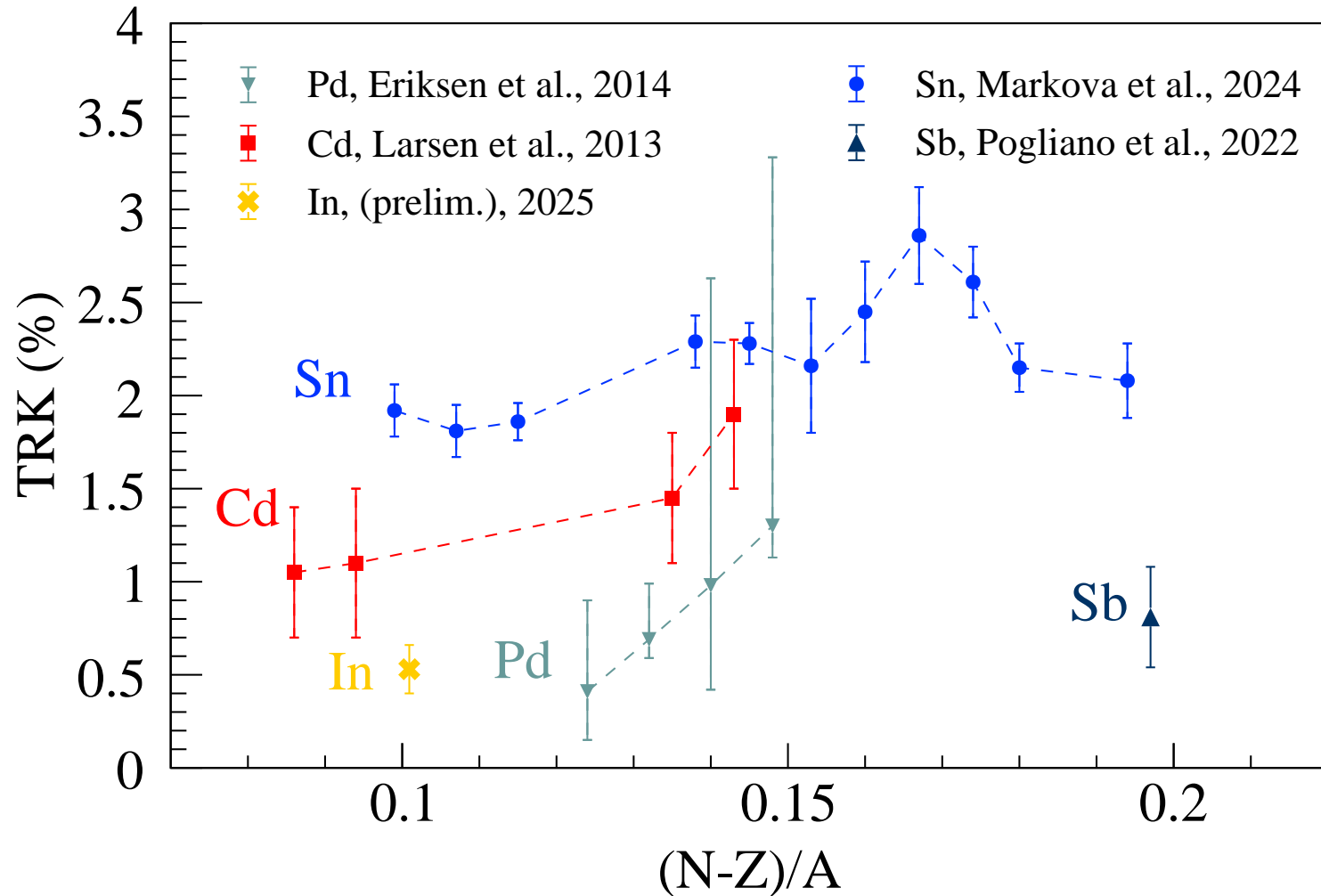
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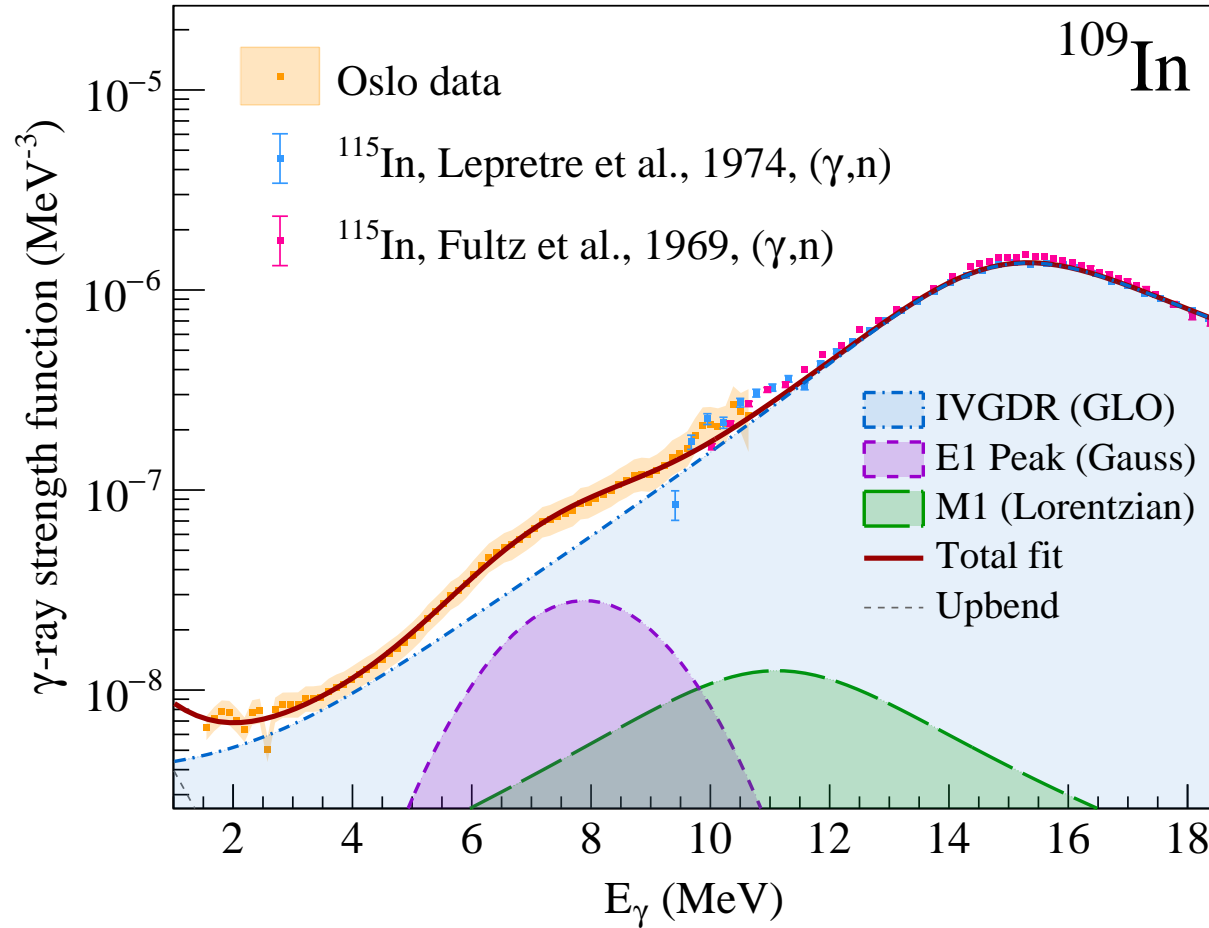




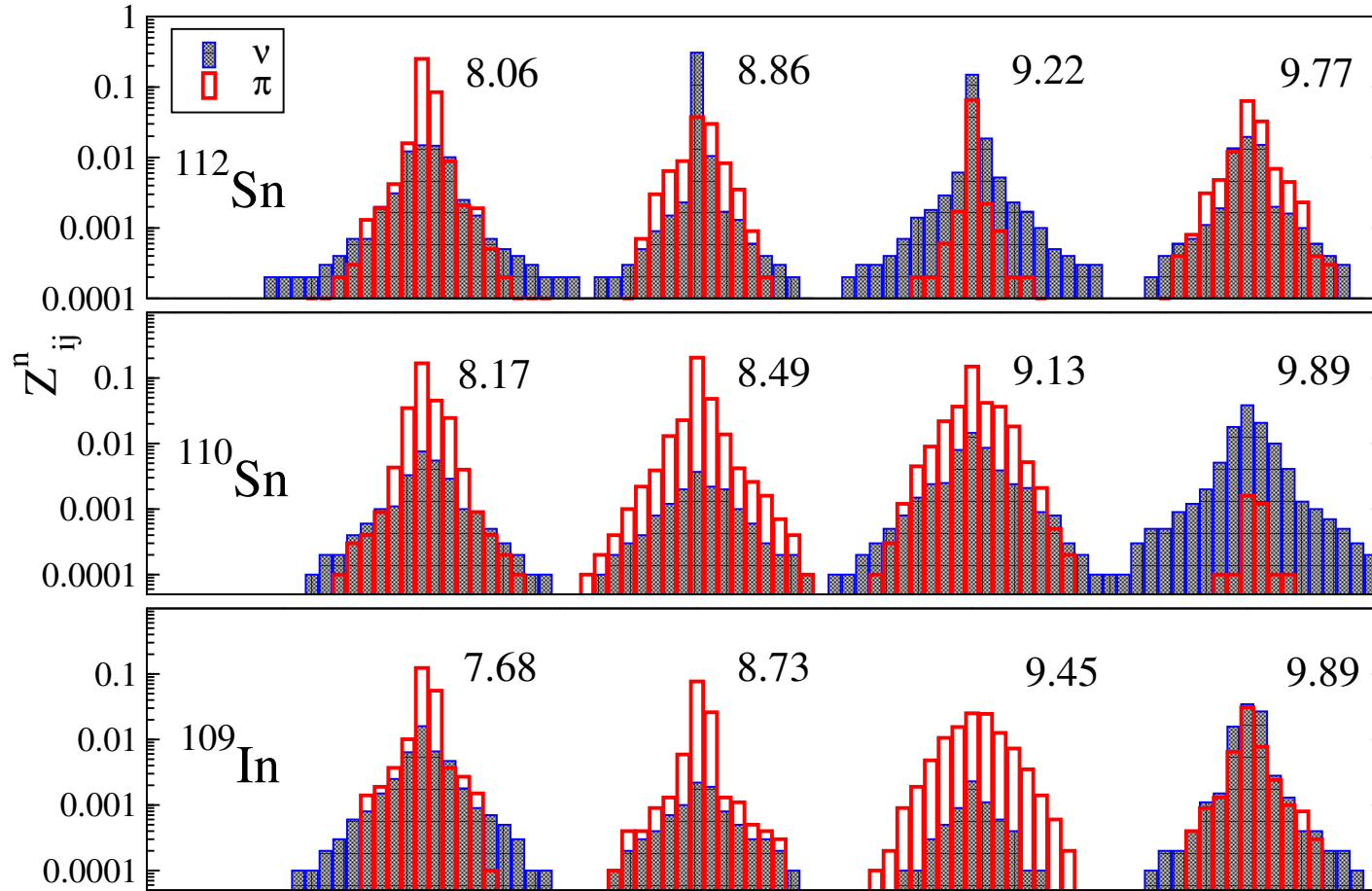
# PDR in the Sn region



# PDR in the Sn region: $^{109}\text{In}$



# PDR in the Sn region: $^{109}\text{In}$



# Summary and conclusions

- ▶ The nuclear level densities and  $\gamma$ -ray strength functions of Pd, Cd, In, Sn, Sb have been extracted in a model-consistent way with the Oslo method.
- ▶ The low-lying electric dipole strength in stable Pd, Cd, Sn isotopes is located at  $\approx 8$  MeV and exhausts  $\approx 0.5 - 3\%$  of the TRK sum rule.
- ▶ No systematic increase of the strength with  $N$  was observed in Sn isotopes.
- ▶ RQRPA and RQTBA calculations were performed to interpret the evolution of the low-lying  $E1$  strength.

## Thank you for your attention!

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