





Caen, France September 25 2025





Update of the Summation Calculations for Reactor Antineutrino Spectra



Outline

- Motivations: Beta Decay study to probe Reactor Antineutrino Anomalies
- Summation Calculations and Nuclear Data
- Updates of the Model and a Few Selection of Results
- Conclusions and Outlooks

Beta Decay Study for Reactor Antineutrinos

• Getting access to the β decay properties and to antineutrino energy spectra

$${}_{Z}^{A}X \rightarrow {}_{Z+1}^{A}Y^{*} + e^{-} + \overline{\nu}_{e}$$

$${}_{Z+1}^{A}Y^{*} \rightarrow {}_{Z+1}^{A}Y + \gamma s$$

- The exploitation of the products of the beta decay is multifold:
 - ☐ The <u>antineutrinos</u> escape and can be detected → essential for fundamental physics but also reactor monitoring and potential non-proliferation tool
 - In laboratory, γ or β measurements of well identified fission products \rightarrow characterize the weak interaction properties, several physics topics in nuclear structure or nuclear astrophysics but also indirect access to antineutrino energy spectra
- Beta decay driven by some selection rules regarding the isospin and the spin-parity between the parent and daughter nuclei
 - Arr Fermi in the 30s: $\lambda = \frac{2\pi}{\bar{h}} |V_{fi}|^2 \rho(E_f)$, $V_{fi} \equiv \langle \psi_f | O_\beta | \psi_i \rangle$
 - ☐ First set of equations for the energy spectrum of electromagnetic particles emitted during a beta decay process
- Over the last 50 years, many computations and improvements of the spectra
- Two methods were revisited in 2011: the conversion and the summation methods.
 - ☐ The HM model used as reference for the neutrino experiments

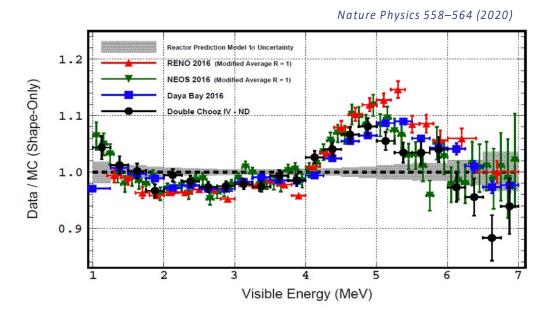
First Observations of the Anomalies of Reactor

Antineutrino Energy Spectra

- ullet Measurement of the $heta_{13}$ oscillation parameter by Double Chooz, Daya Bay, Reno in 2012
 - ☐ Independent evaluation of anti-v energy spectra using NDBs
 - □ 6% deficit in the measured flux at short baseline experiments compared
 - with the best prediction based on ILL data: reactor anomaly (RAA)
 - ☐ Numerous projects in search of the existence of sterile neutrinos

- Y. Abe et al. Phys. Rev. Lett. 108, 131801, (2012) F. P. An et al., Phys. Rev. Lett. 108, 171803 (2012)
- J. K. Ahn et al., Phys. Rev. Lett. 108, 191802 (2012)
- G. Mention et al. Phys. Rev. D83, 073006 (2011)

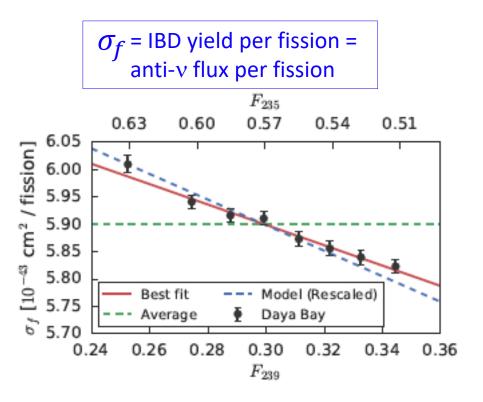
• In 2014, the same three experiments highlighted a spectrum distortion between 4.8-7.3 MeV compared to nuclear models again! (Shape anomaly)



Context by End 2017

In 2017: Daya Bay's new result about the reactor anomaly: pb is in the ²³⁵U spectrum!!!

F. P. An et al. (DB Collaboration), "Evolution of the Reactor Antineutrino Flux and Spectrum at Daya Bay," PRL 118 (2017) 251801. And associated APS Viewpoint by M. Fallot



- Slope not in full agreement with H-M and global deficit of 5.1%
- Deficit in detected antineutrinos compared with predictions depends on the relative fractions of ²³⁵U, ²³⁹Pu, ²³⁸U, and ²⁴¹Pu in the reactor.
- Flux deficit quasi all taken by ²³⁵U —enough of a discrepancy to explain by itself the entire antineutrino anomaly !!!
- In contrast, the discrepancy = almost zero for ²³⁹Pu fissions.

- Potential issue in Schreckenbach measurement or H-M model for ²³⁵U?

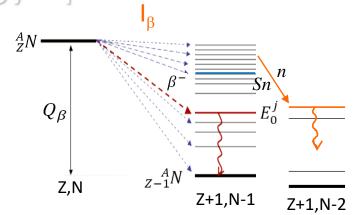
 New DB (ArXiv:1904.07812) and RENO (PRL 122 (2019) 232501) papers re-inforce previous results
 - **➡** Growing interest in summation method to calculate anti-v spectra

Access to the variety Spectral of a FP

• Measurement of well identified fission products:

$${}^{A}_{Z}X \rightarrow {}^{A}_{Z+1}Y^{*} + e^{-} + \overline{\nu}_{e}$$

$${}^{A}_{Z+1}Y^{*} \rightarrow {}^{A}_{Z+1}Y + \gamma s$$



Total energy spectrum of a fission product:

$$S_{fp}(Z,A,p) \propto \sum_{b=1}^{N_b} I_{\beta fp}^{\ \ b} \times S_{fp}^{\ \ b} \left(Z_{fp},A_{fp},E_{0fp}^{\ \ b},E\right)$$

Access to the v Energy Spectra of a EP

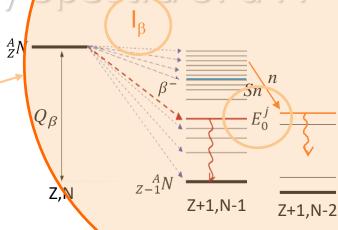
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TAGS measurements

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TAGS measurements

Z+1,N-2

Z+1,N-1

• Energy spectrum of a b branch of a fission product:

$$S_{fp}^b(p) \propto p^2(Q-T_e)^2 F(Z',p) C(Z,p) (1+\delta(Z,A,p))$$
Phase Fermi Shape Subdominant corrections

Access to the 7 Energy Spectra of a FF

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Phase Fermi Shape Subdominant space function factor corrections

A. Alcala et al., accepted to PRL (https://doi.org/10.1103/hyj7-l22h)

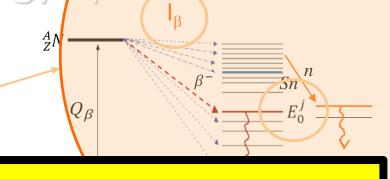
Preprint: <u>arXiv.2505.05929</u>

Access to the V Energy Spectra of a FF

• Measurement of well identified fission products:

$${}^{A}_{Z}X \rightarrow {}^{A}_{Z+1}Y^{*} + e^{-} + \overline{\nu}_{e}$$

$${}^{A}_{Z+1}Y^{*} \rightarrow {}^{A}_{Z+1}Y + \gamma s$$



Caveat:

Summation calculations are strongly dependent on nuclear data!

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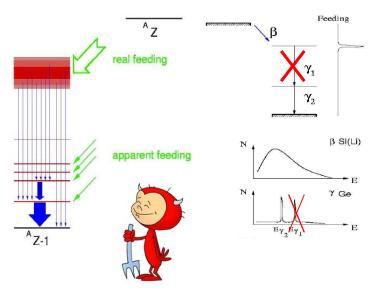
γ Measurement Caveat

- Before the 90s, conventional detection techniques: high resolution γ -ray spectroscopy
 - Excellent resolution but efficiency which strongly decreases at high energy or because of small acceptance
 - Danger of overlooking the existence of β -feeding into the high energy nuclear levels of daugther nuclei
- Incomplete decay schemes: overestimate of the highenergy part of the FP β spectra
- Phenomenon commonly called « pandemonium effect** »

** J.C.Hardy et al., Phys. Lett. B, 71, 307 (1977)

■ Expected distortion of the antineutrino energy spectra computation with SM

Picture credit: A. Algora



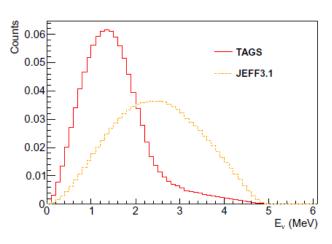


FIG. 1. Illustration of the pandemonium effect on the 105 Mo nucleus anti- ν energy spectrum presents in the JEFF3.1 data base and corrected in the TAS data.

TAGS: a Solution to the Pandemonium Effect

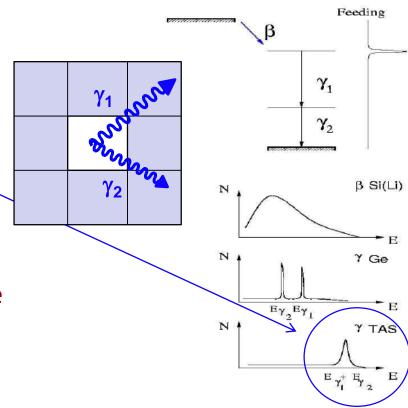
• Total absorption γ -ray spectroscopy (TAGS)

■ A TAS is a calorimeter

 \Box It contains big crystals covering 4π

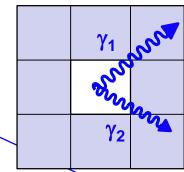
Instead of detecting the individual gamma rays, absorbs the full gamma energy released by the gamma cascades in the β-decay process

 First TAS developed in the 70's but too small detectors to be efficient. Development of the TAGS method efficient and systematic since the 90's (Greenwood & al.)

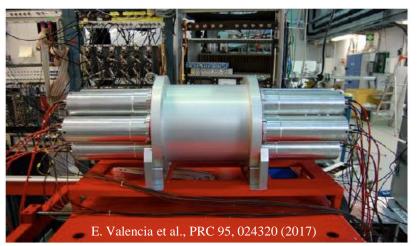


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Rocinante

- \triangleright 12 BaF₂ covering 4π
- Detection efficiency of γ ray cascade >80% (up to 10 MeV)

See J. Pépin, Monday 22 – 2:45pm (NSSD) See A. Porta, Tuesday 23 – 3:05pm (NSSD) See M. Fallot, Thursday 25 – 2:20pm (AI)

DTAS

- > 18 NaI(Tl) crystals of 15cm×15cm×25 cm
- ➤ Individual crystal resolutions: 7-8%
- ➤ Total efficiency: 80-90%



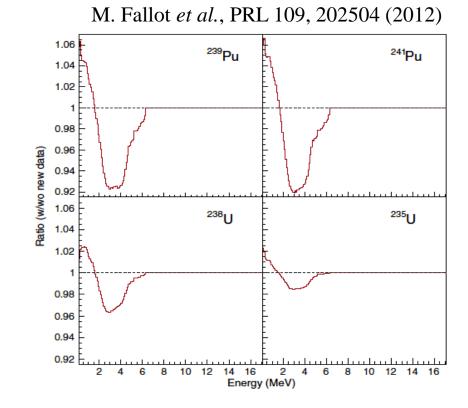
J.L. Tain et al., NIMA 803 (2015) 36 V. Guadilla et al., NIM A 854 (2017) 134

New Summation Model in 2012 – SM-2012

- All fission products in the JEFF3.1.1 fission yields databases taken into account
- To reproduce the experimental conditions of Schrekenbach et al., we simulated irradiation times with MURE: 12 h for ²³⁵U, 1.5 d for ^{239;241}Pu, and 450 d for ²³⁸U.
- We highlighted the necessity:
 - To use nuclear databases Pandemonium free and proposed a cocktail of nuclear databases to follow to reduce as much as possible experimental biases
 - To complete the full list of FP involved in the calculation using models or the Qbeta approximation at minimum.
- Taking into consideration the TAS data of the ^{102;104–}
 ¹⁰⁷Tc, ¹⁰⁵Mo, and ¹⁰¹Nb isotopes measured in 2007 @

 Jyväskylä

 Algora et al., PRL 105, 202501 (2010).
 - □ ~850 nuclei included
 - Noticeable deviation from unity (1.5 to 8% decrease)



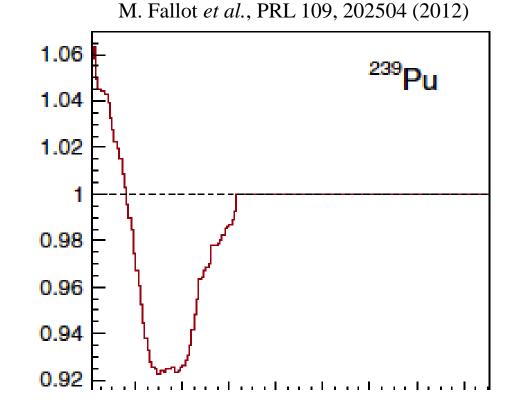
Relative Effects of the 2007 TAS data (published 2010) on the Antineutrino Spectra: typical from Pandemonium: increases the spectrum before 2-3 MeV and decreases it above => strong impact on the IBD yield!

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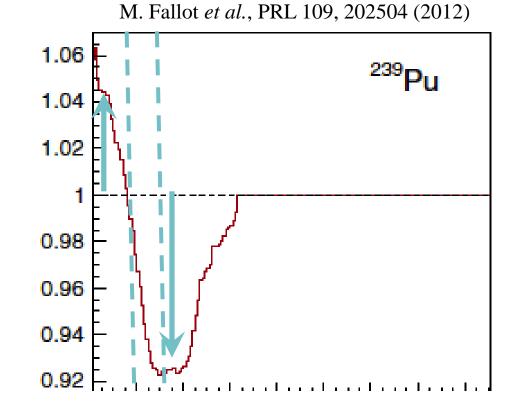
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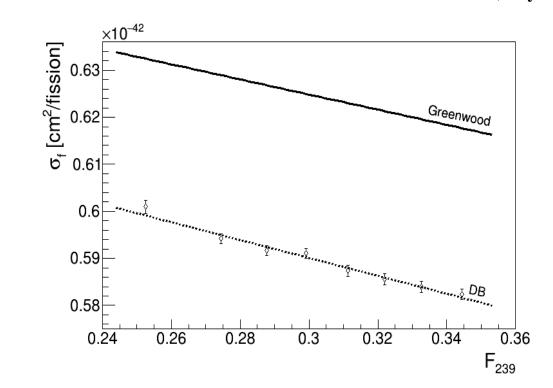


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SM-2018 – Impact of a Decade of TAGS data

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^{102;104–107} Tc, ¹⁰⁵ Mo & ¹⁰¹ Nb	SM-2012 M. Fallot et al. PRL 109, 202504 (2012)	A. Algora et al. PRL 105, 202501 (2010), D. Jordan et al. PRC 87, (2013) 044318
+ ⁹² Rb	SM-2015	A.A. Zakari-Issoufou et al. PRL 115, 102503 (2015)
+ ^{87,88} Br and ⁹⁴ Rb + ⁸⁶ Br and ⁹¹ Rb	SM-2017	E. Valencia et al., PRC 95, 024320 (2017) S. Rice et al. PRC 96 (2017) 014320
+ ^{100,100m,102,102m} Nb	SM-2018 (PRL quoted below)	V. Guadilla et al. PRL 122, (2019) 042502

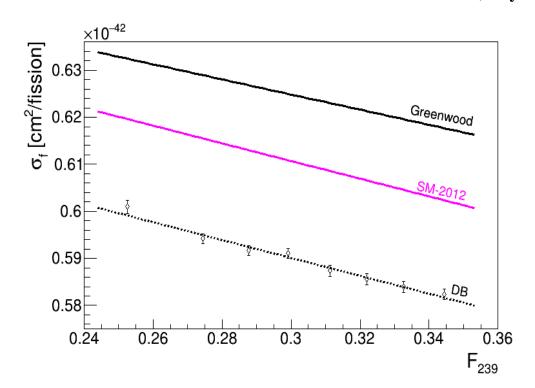
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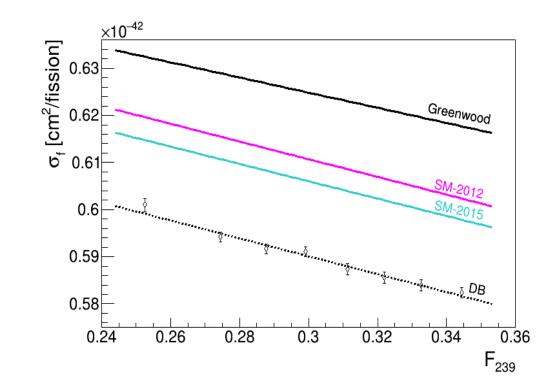
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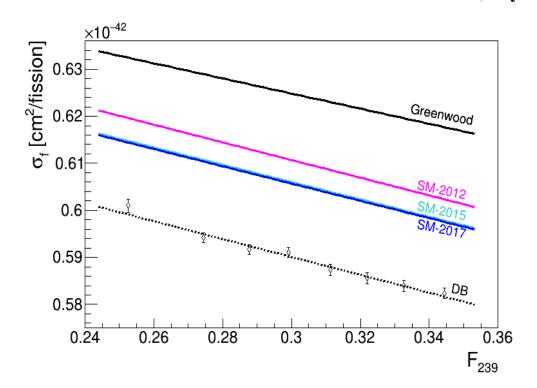
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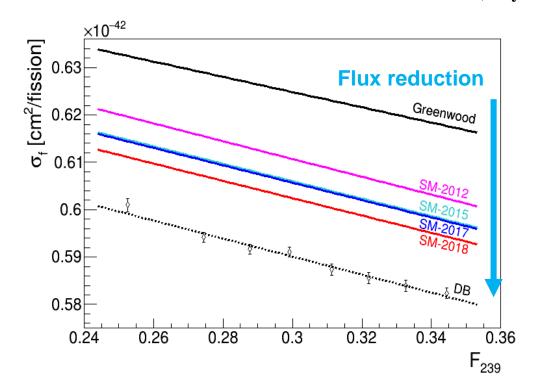
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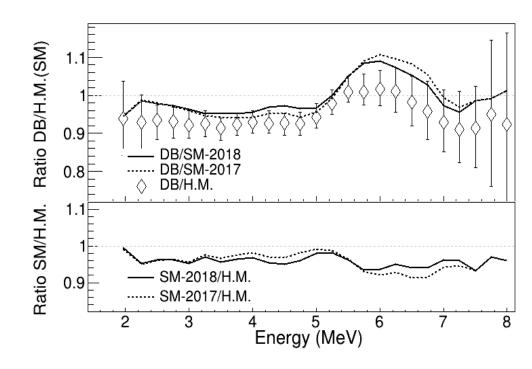
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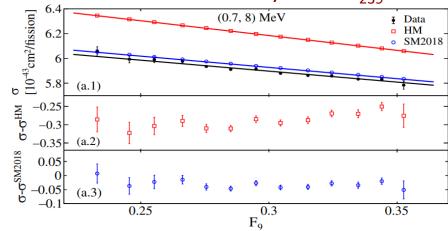
First full spectrum prediction without any renormalization.

Pandemonium correction contributes to reduce the bump but it is still there.

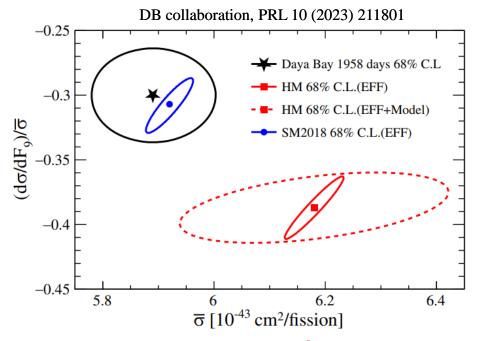


2023 Results from Daya Bay

Evolution of the total IBD yield vs F₂₃₉



- Using of 1958-day data sample (~6 years of measurement 2011-2017), study of the total IBD yield.
- Spent nuclear-fuel contribution and out-ofequilibrium effect taken into account as a function of antineutrino detector and time.
- Data sorted into 13 groups of ²³⁹Pu fission fraction.



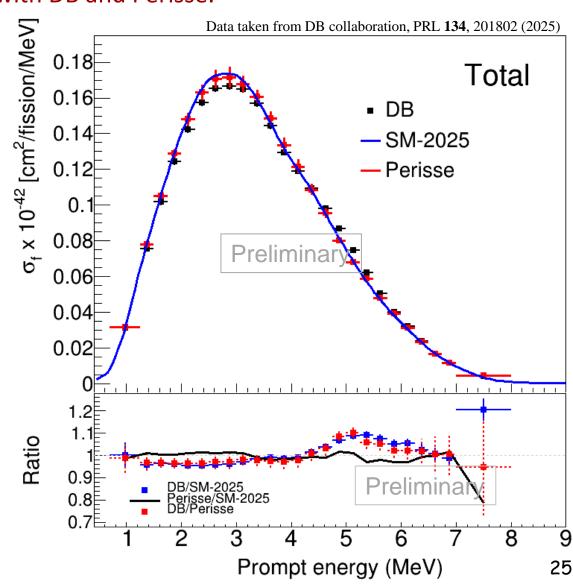
- H-M predictions for $\bar{\sigma}$ and $[\frac{d\sigma}{dF_9}]/\bar{\sigma}$ rejected at 3.6 and 3 standard deviations.
- SM18 consistent with DB
 RAA relieved by SM-2018 but not the spectrum shape
- Issue with the ²³⁵U measurement from Schrekenbach et al. Confirmed

Summation Model Updates - 2025

• Ingredients of our model core calculation (SM-2023) : ☐ 1keV energy bins Collaboration with L. Hayen: Screening corrections: Rose replaced by Salvat (L. Hayen, N. Severijns et al. Rev. Mod. Phys. 90, 015008 (2018)) \square Nubase 2020 for Q_{β} approximation □ 65 nuclei from Rudstam / Tengblad et al. (elimination of those in agreement with JEFF/ENDFB8, and a few « odd » ones) \Rightarrow Small change in the global flux (~+0.25%) 2014 TAGS campaign: quantification of the impact of 7 new nuclei: □ ⁹⁵Rb et ¹³⁷I: 2 nuclei from V. Guadilla et al. **Phys. Rev. C 100, 044305 (2019)** □ ^{96gs}Y and ^{96m}Y (Pandemonium): 2 nuclei from V. Guadilla et al. Phys. Rev. C 106, 014306 (2022) □ ⁹⁹Y, ¹⁴²Cs and ¹³⁸I: 3 Pandemonium nuclei from L. Le Meur et al., **in preparation**

Total Prompt Energy Spectra Comparison

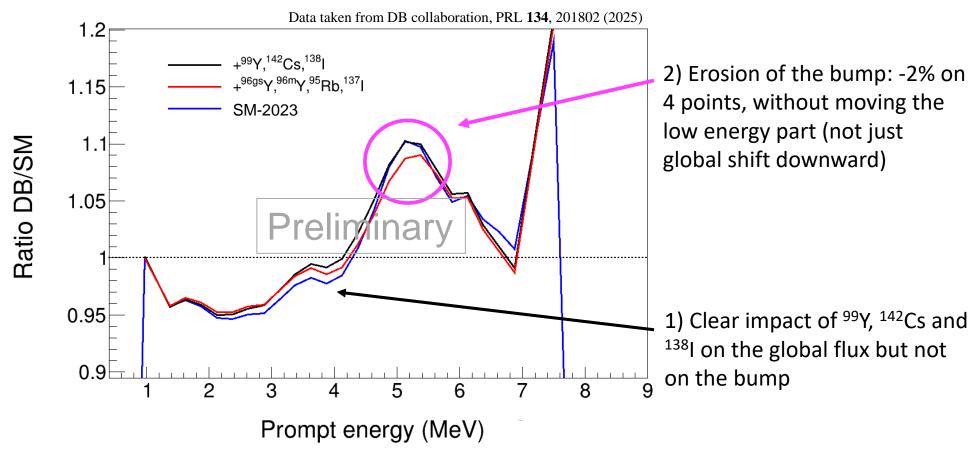
- SM-2025 total prompt energy spectrum compared with DB and Perissé.
 - ☐ Convolution of our summation calculations with DB response matrix
 - As already seen, the bump is still there and we are below DB between 1.5 and 3 MeV in the dominant IBD cross section region by 4 to 5%. We reach 2% difference between 3 and 4 MeV before the bump.
 - ☐ With our updated model, good agreement with Perissé et al.
 - ☐ Discrepancies +1% in the 2 to 3 MeV region and -2-3% between 5.5 and 6.5 MeV.
 - => Two SM compatible within Perissé et al.'s uncertainties but we have more TAGS Decay Data included



Impact of the « TAGS » Nuclei Added in the Model

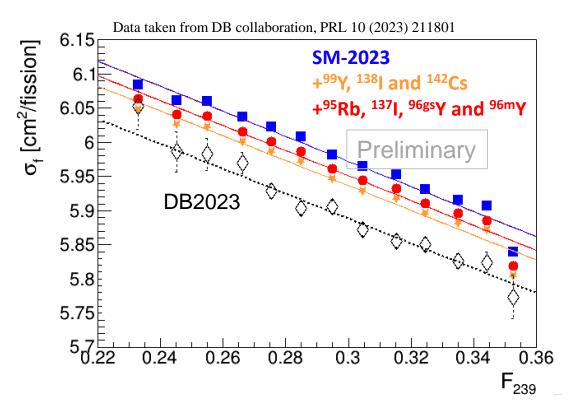
Ratio Daya Bay Spectrum 2025 over SM-2025

- ☐ Comparison of the SM-2023 and SM-2025 to DB in prompt energy of the total spectrum.
- ☐ Out of equilibrium effects corrected in both cases.



Impact on the IBD yield

Ratio Daya Bay 2023 IBD flux vs SM-2025

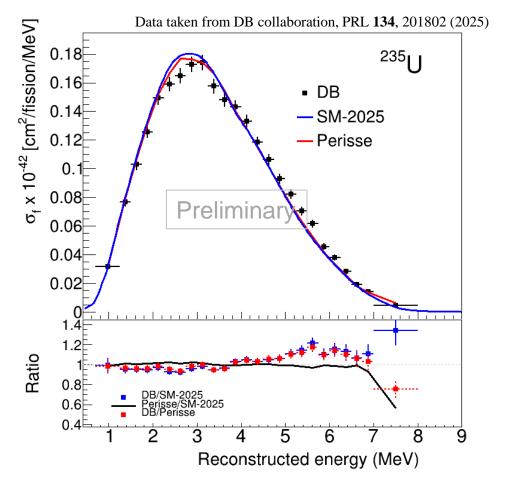


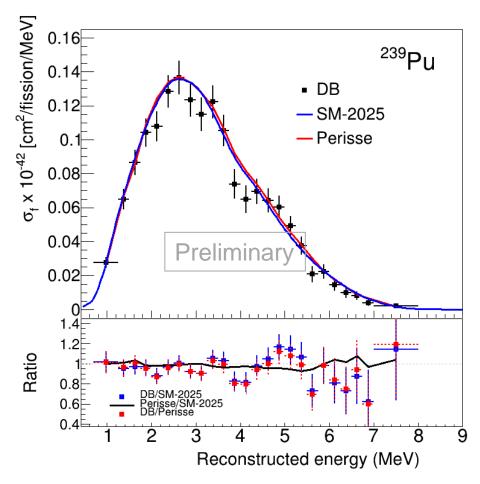
- SM-2025 is ~1.5% above DB2023:
- +0.25% increase due to update of the model (SM-2023)
- -0.6% due to ¹⁴²Cs, ¹³⁸I and ⁹⁹Y
- +0.25% due to ^{96, 96m}Y, ¹³⁷I and ⁹⁵Rb
- +0.25% due to new Daya Bay points
- ☐ Pandemonium correction still decreases the discrepancy but some nuclei are corrected from other systematic effect (i.e. w.r.t. Rudstam or new data for isomer)

- Still systematic trend reducing the flux including pandemonium free data
- SM-2025 is the most up-to-date in terms of TAGS decay data (most dominant uncertainty)
- More to come with new TAGS campaign

²³⁵U and ²³⁹Pu Prompt Energy Spectra Comparisons

SM-2025 prompt energy spectrum compared with DB and Perissé.





- ☐ More pronounced bump in the case of ²³⁵U wrt ²³⁹Pu but too large DB uncertainties to quantify the deviation in the ²³⁹Pu case.
- ☐ SM models in good agreement in both cases. Largest discrepancy for ²³⁹Pu in the range 4.5 to 5.5 MeV

Conclusions & Perspectives

- TAGS data (Pandemonium free) correct NDB from Pandemonium effect
- SM-2018 updated to SM-2025 :
 - ☐ Internal ingredients changed or updated to newer databases
 - ☐ Seven TAGS nuclei added either Pandemonium free or different from previous database taken as reference
- Pandemonium systematic still observed and improvement in both flux (1.7%) and shape (-2%) in the bump region wrt to DB
- Good agreement in flux with Daya Bay and Perissé
- Still some discrepancies in shape with DB and good agreement with Perissé et al. except in an energy range in the bump region for ²³⁹Pu
- SM-2025 most up-to-date Pandemonium free data
- More to come with new TAGS data and eShape data to take into account forbidden non unique corrections

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Thank you!

The TAS Collaboration



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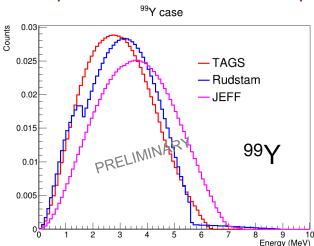
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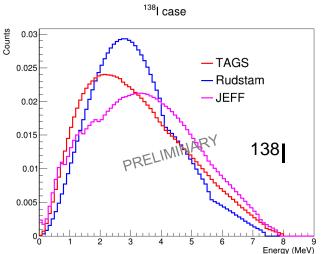
HUN-REN ATOMKI: D.Sohler, G.Kiss, I.Kuti

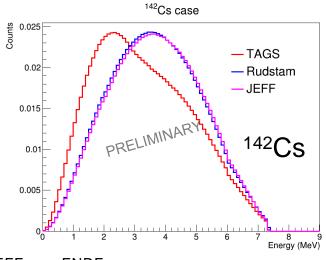
Extra

Individual Anti-v Energy Spectra: 99Y, 1381 and 142Cs

- Comparison of the individual antineutrino energy spectra between DTAS and the preferred nuclear database that was used for our previous calculation (Rudstam).
 - \square Rudstam β spectra converted
 - ☐ Non pandemonium free data in JEFF 3.1.1
 - ☐ Shift vs low energy in TAS: apparent biases in Rudstam measurement and large error bars
 - ☐ Impact the total antineutrino spectrum



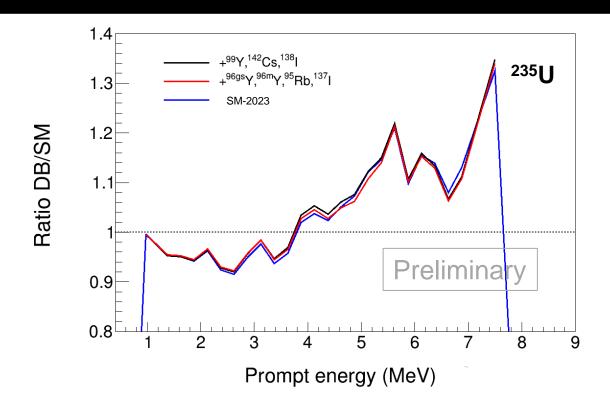


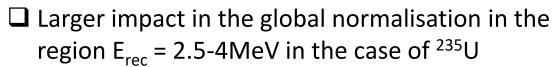


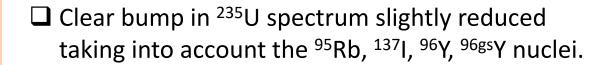
PhD Thesis work:				
Loïc Le Meur (Subatech, Nantes)				
Publication in preparation				

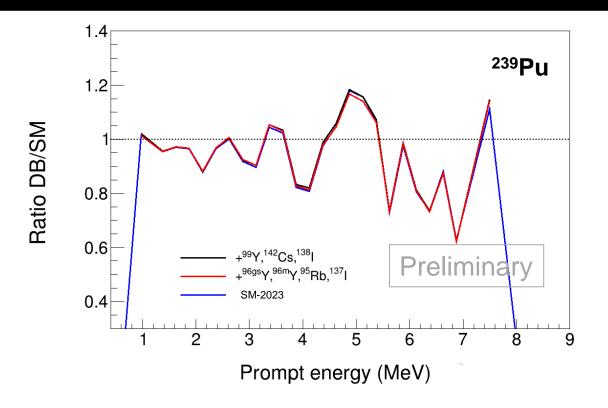
Decay	Ε	DTAS	JEFF	ENDF
		[keV]	[keV]	[keV]
¹⁴² Cs	u	$1526.3^{+\ 83}_{\neq\ 54}$	676.48	952.37
	_	2535.0 ^{+ 25} _{≠ 39}	2919(178)	2919(179)
99 Y	u	1584 ^{+ 46} _{≠ 31}	917	1006
	_	$2379^{+\ 15}_{\neq\ 22}$	2949(146)	2931(208)
¹³⁸	u	$2005^{+\ 106}_{\neq\ 99}$	1325	1420
•	_	$2475^{+64}_{\neq 33}$	2721(125)	3068(290)

Impact as a Function of the Isotope ²³⁵U or ²³⁹Pu









- □ ²³⁹Pu prompt spectrum less impacted in shape and flux than the ²³⁵U spectrum (strong fluctuations in DB measurements)
- ☐ Small bump in the ²³⁹Pu spectrum slightly affected as well by correcting the databases of these nuclei

Propagation of Yield Uncertainties

	DB (x10 ⁻⁴³ cm ² /fission)	SM-2025
σ_{f}	5.84+/-0.07	5.93 +/-0.13
σ_5	6.16+/-0.04(stat)+/-0.08(syst)+/-0.18(model)	6.22
σ_9	4.16+/-0.07(stat)+/-0.09(syst)+/-0.18(model)	4.34
$d\sigma_f/dF_{239}$	-1.96+/-0.11(stat)+/-0.07 (syst)	-1.83

Consistent wrt SM-2025 (1% deficit) 4.3% déficit wrt SM-2025

- Using the (improved) GEF covariance matrices of the yields and an analytical error propagation
- Study of the dependence on beta decay uncertainties which are underestimated in nuclear BDD in many cases: assuming
 4 uncertainties, all the same for all nuclei: 5, 10, 15, 20%

$\Delta_{DD} = 5\%$	Δ_{DD} = 10%	Δ_{DD} = 15%	Δ _{DD} = 20%
$\Delta\sigma_{\rm f}$ = +-1.00%	$\Delta\sigma_{\rm f}$ = +-1.33%	$\Delta\sigma_{\rm f}$ = +-1.70%	$\Delta \sigma_{\rm f}$ = +-2.2%

- Note that this calculation was performed a long time ago...
 - ☐ Yields are very correlated
 - ☐ Final uncertainty is quite small, even when assuming 20% on all decay data
- ⇒ On-Going work
- ⇒ Truth is some BDD should have less than 20% and some more (100%)