

PROBING THE ISOSPIN MIXING IN THE ^{72}Kr COMPOUND NUCLEUS VIA GDR γ DECAY

A. Giaz

INFN Sezione di Milano

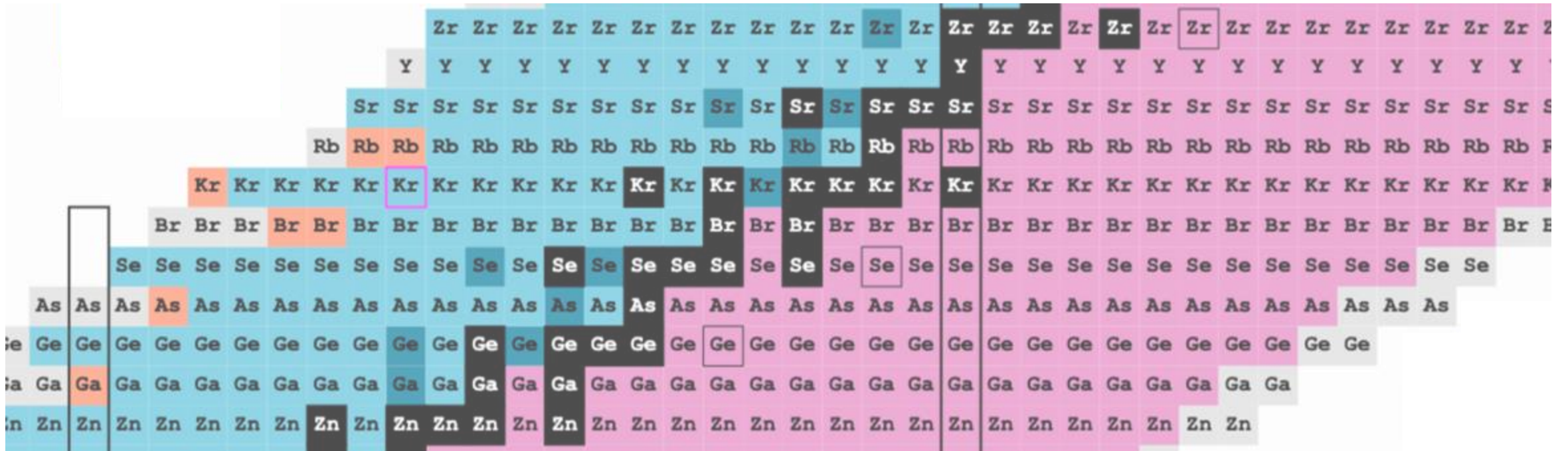


European Nuclear Physics Conference 2025



OUTLINE

- Motivations
 - Isospin mixing
 - Why ^{72}Kr ?
- Experimental Technique & Setup
 - Isospin mixing via GDR
- Results
 - The isospin mixing value α^2 and the correction term δ_c
- Summary & Perspectives



MOTIVATIONS

^{72}Kr AT LOWEST POSSIBLE TEMPERATURE

ISOSPIN SYMMETRY

Isospin symmetry plays a key role in nuclear structure and nuclear reaction.

$$\bullet |p\rangle \equiv |I = \frac{1}{2}, I_z = +\frac{1}{2}\rangle \quad \bullet |n\rangle \equiv |I = \frac{1}{2}, I_z = -\frac{1}{2}\rangle$$

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Isospin mixing beyond nuclear structure: The CKM matrix


$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

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Measuring transition strictly forbidden by isospin conservation.



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$(\Delta J = 0, \Delta I = 0, \pi_i, \pi_f = +)$

- Measurement of the β -particles anisotropy emission
- For β^+ transition \rightarrow initial state, for β^- \rightarrow final state.
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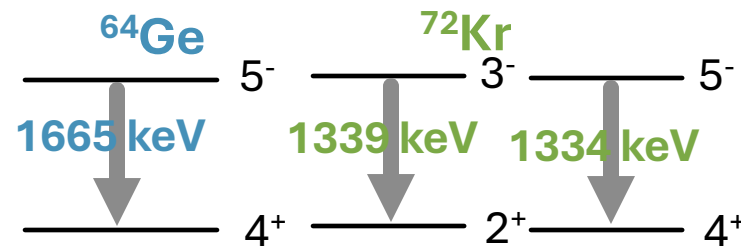
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(polarization & angular distribution)

- Possible only in N=Z nuclei
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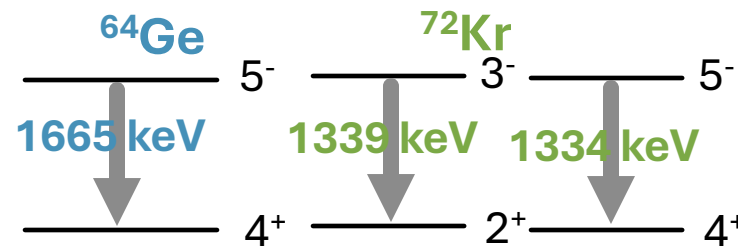
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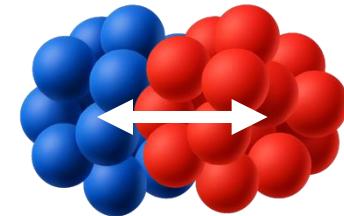


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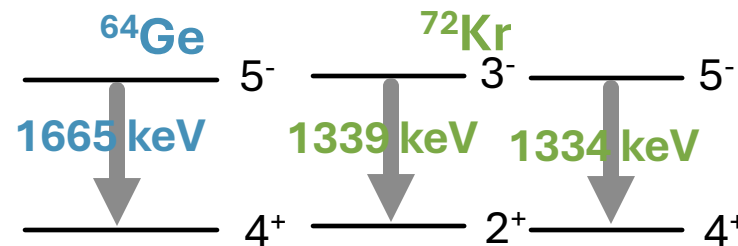
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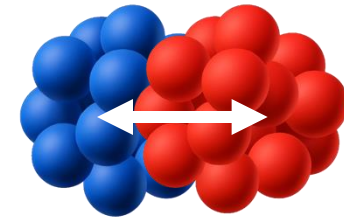


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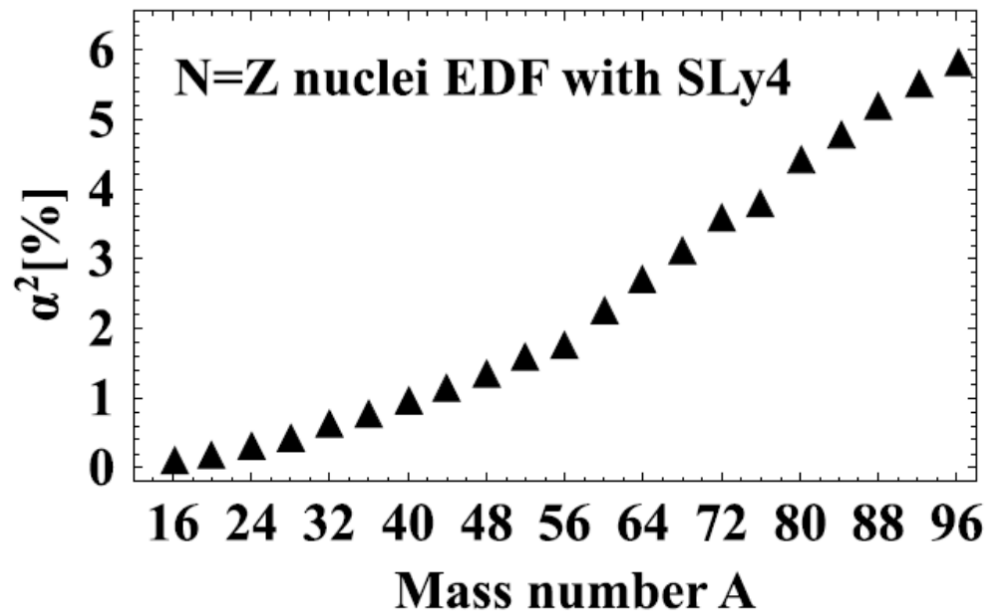
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^{72}Kr this work

ISOSPIN MIXING IN ^{72}Kr

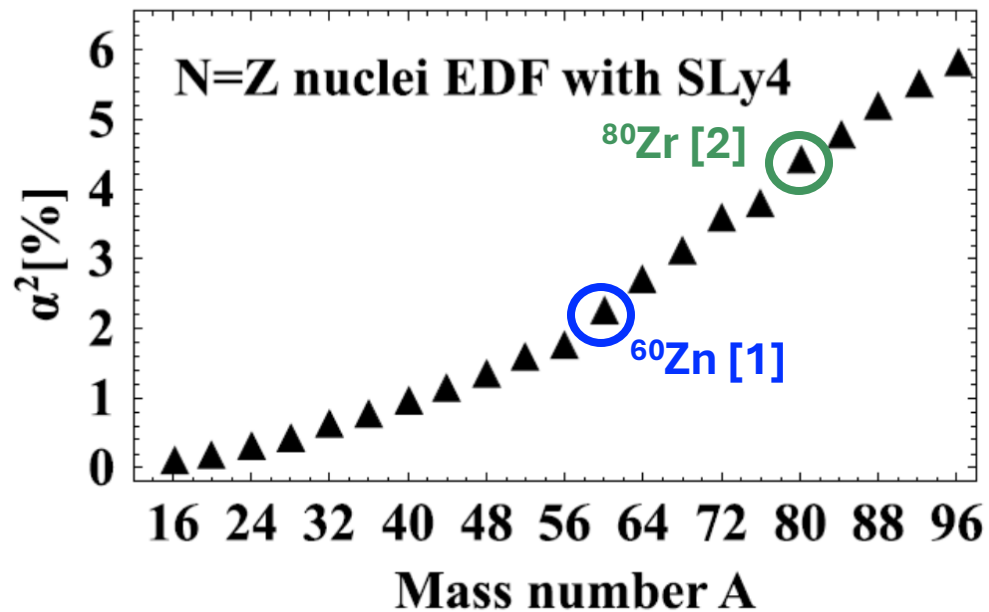
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- It has an intermediate mass between ^{60}Zn and ^{80}Zr .
- It has a E1 discrete transition so it can be measure with two techniques.



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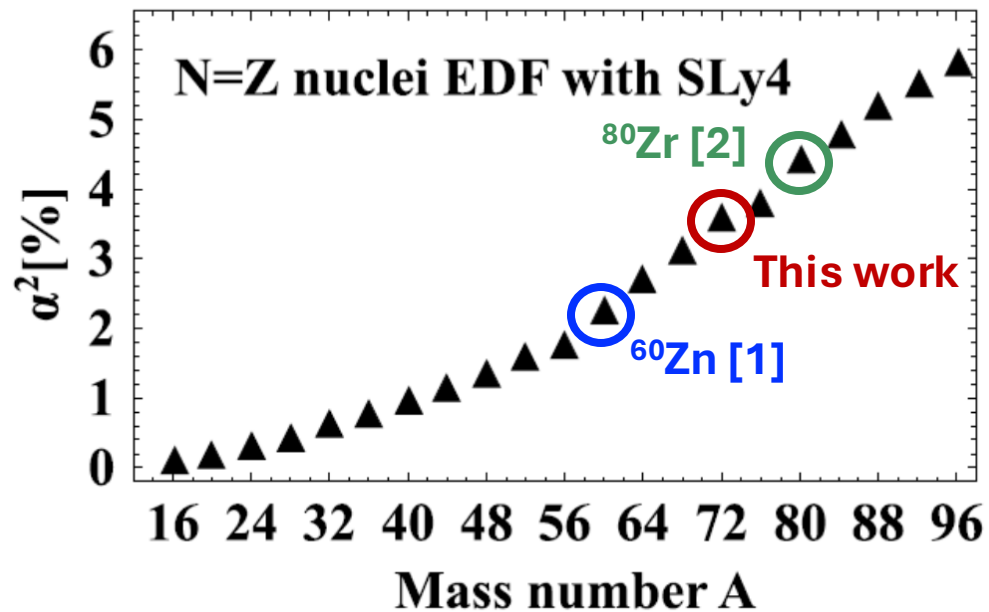
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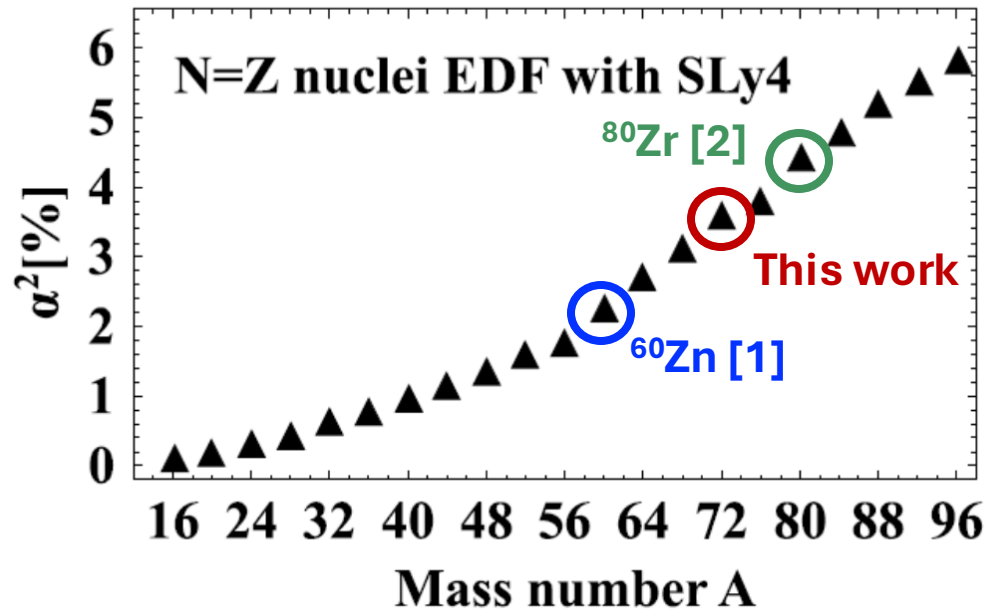
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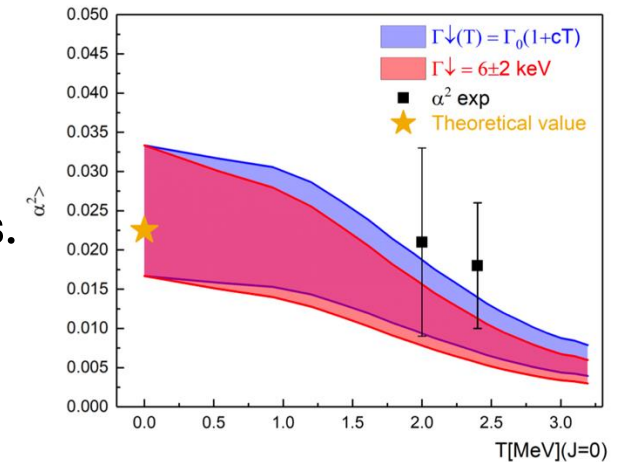
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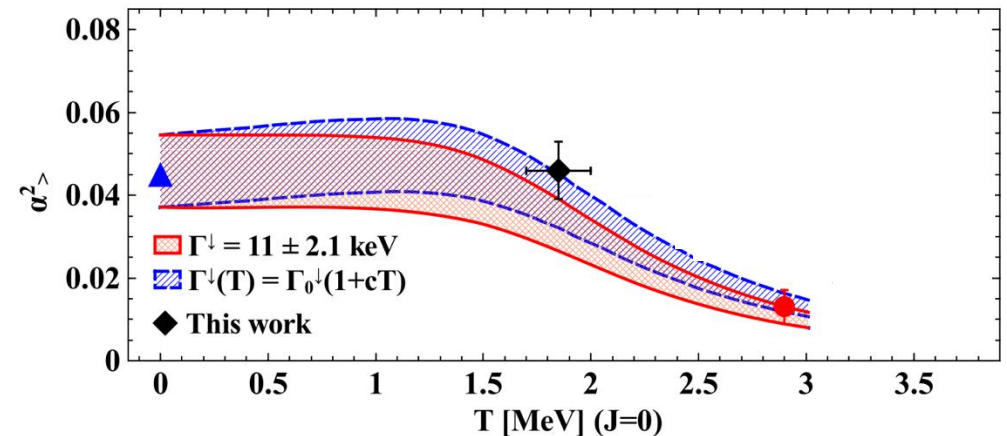
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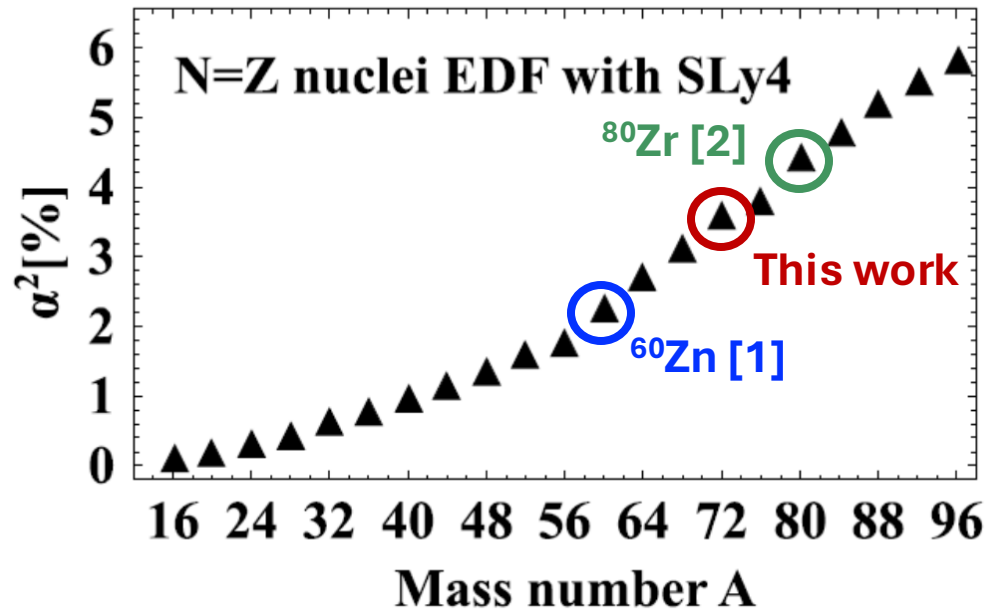
[1] G. Gosta, PhD Thesis, Milano University



[2] S. Ceruti et al., Phys. Rev. Lett., **115**, (2015) 222502

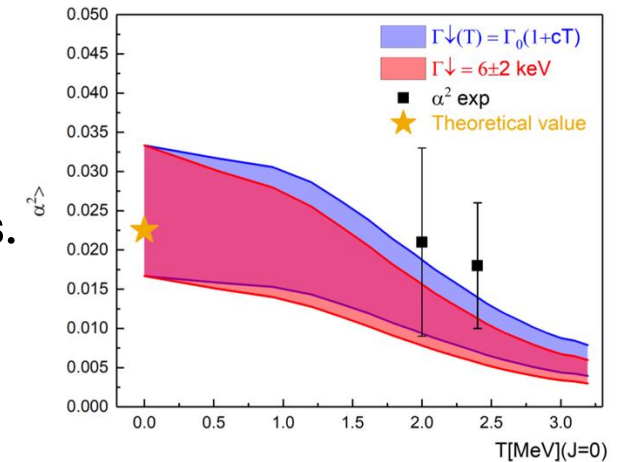
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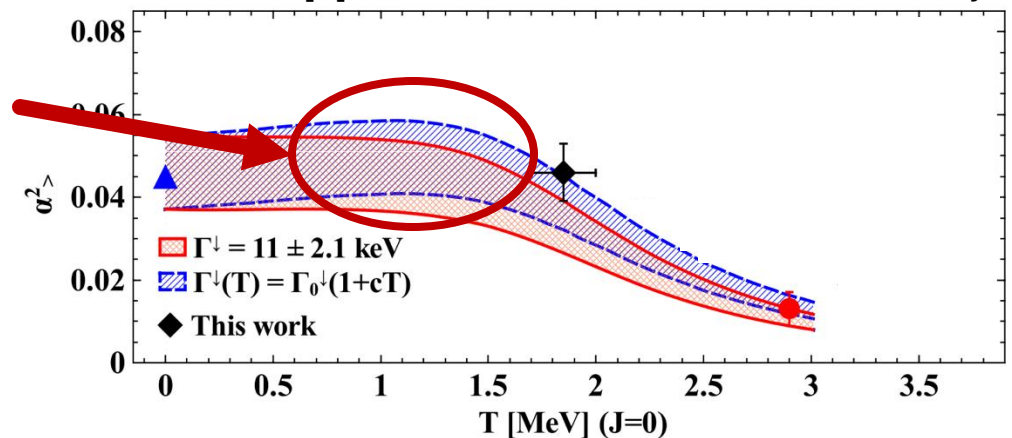


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Lowest possible temperature (low cross-section)



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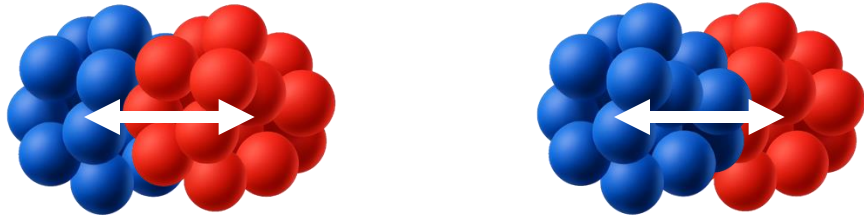
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EXPERIMENTAL TECHNIQUE AND SETUP

ISOSPIN MIXING VIA GDR

EXPERIMENTAL TECHNIQUE: GDR (I)



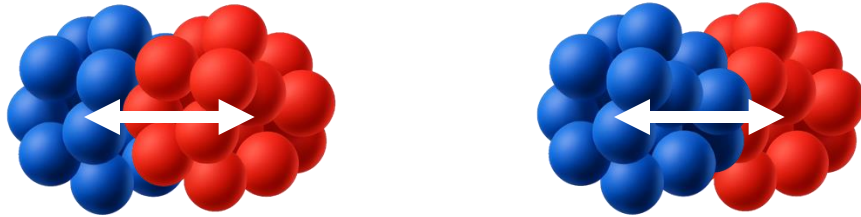
GDR can be described as an oscillation of the protons against the neutrons.

It is a magnetic ($\Delta S = 0$) and isovector ($\Delta I = 1$) transition.

Study of the E1 GDR decay at finite temperature to find the mixing.

To form a compound nucleus at $T > 0$ we used a fusion-evaporation reaction.

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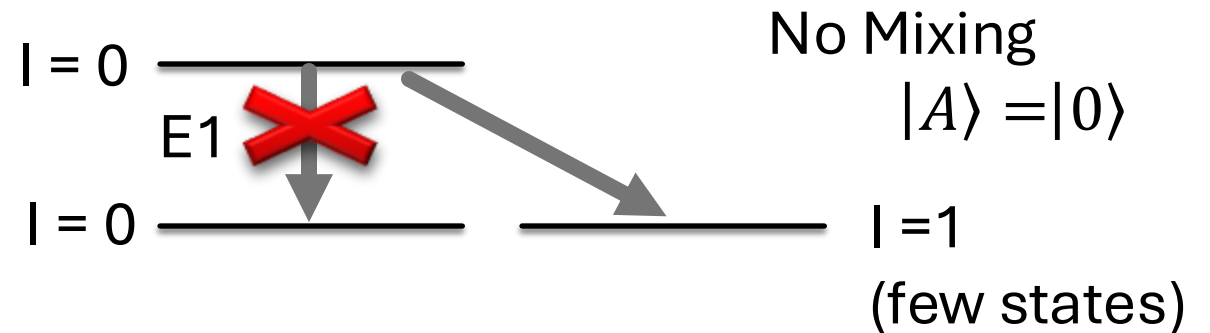


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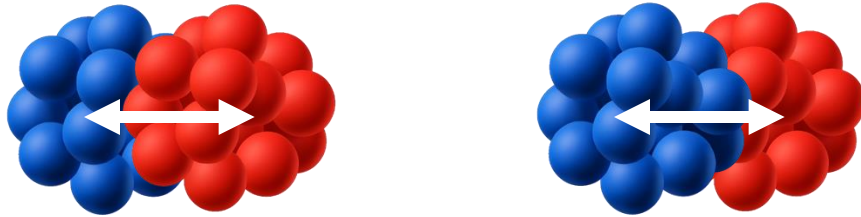
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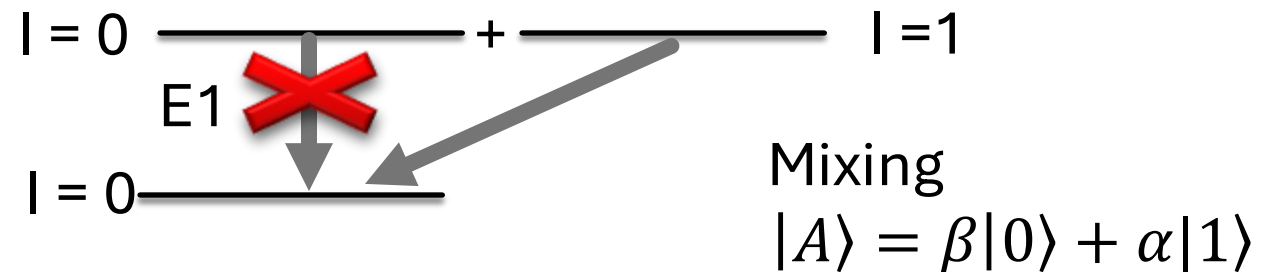
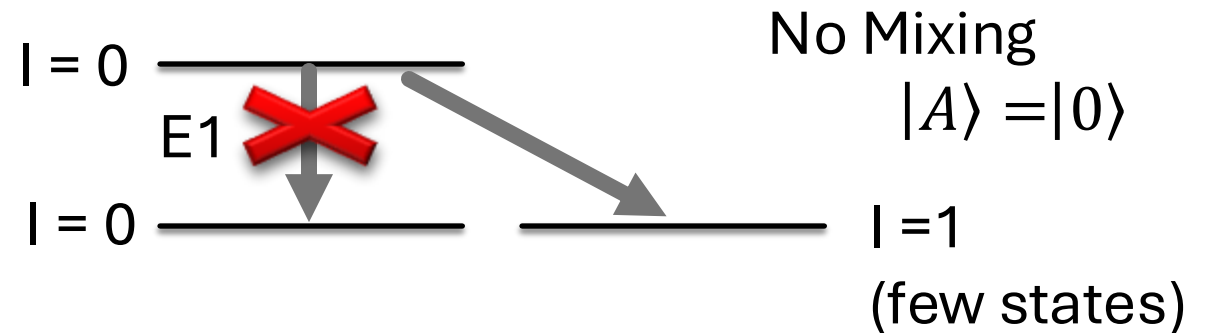


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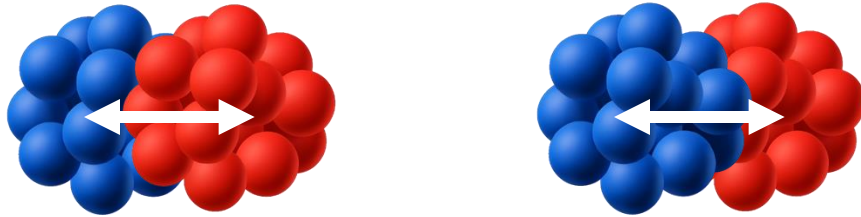
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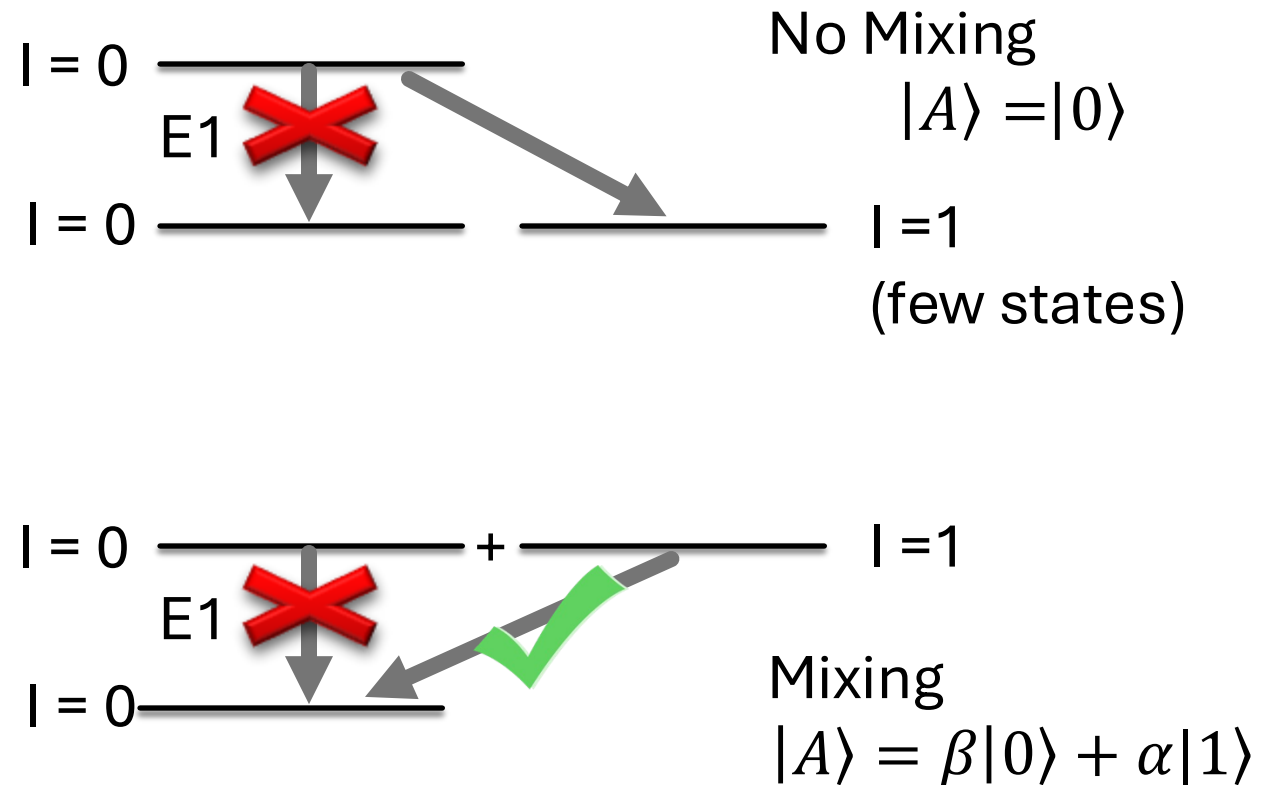


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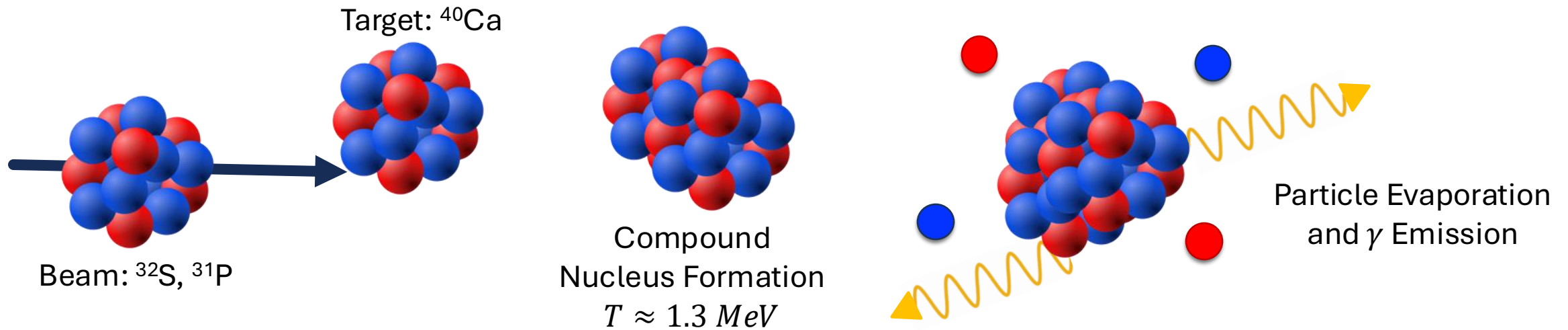
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$I = 0$ compound nucleus reaction:



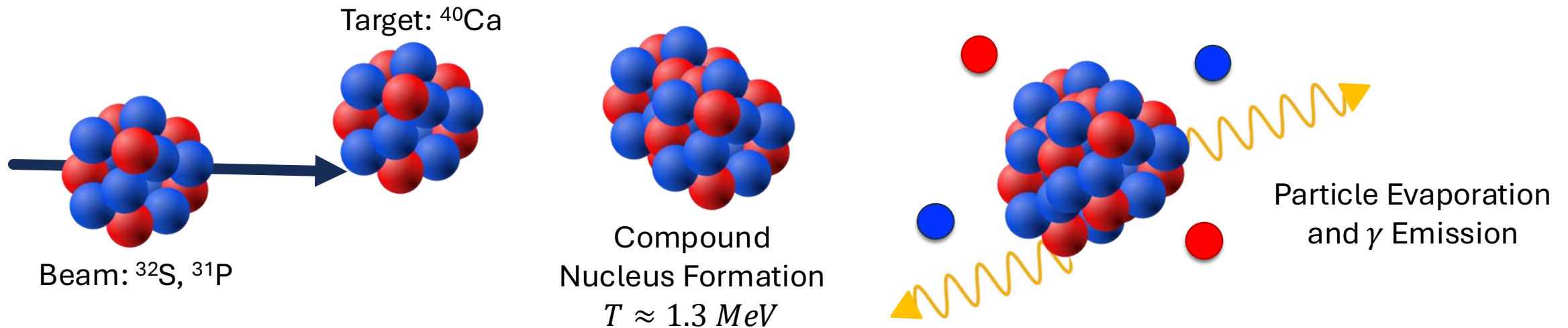
The observed E1 strength is a signature of the mixing.

$I \neq 0$ compound nucleus reaction:



Full mixing. No E1 strength reduction is expected.

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GDR manifest itself as E1 Strength \rightarrow GDR is a probe of the isospin mixing

PREVIOUS EXPERIMENTS

Nucleus	Temperature	Setup
^{80}Zr	2.8 MeV	HECTOR – GARFIELD [1]
^{80}Zr	1.8 MeV	AGATA – HECTOR ⁺ [2]
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The new experiment: Isospin Mixing in ^{72}Kr at the lowest possible temperature.

- In the flat part of the isospin mixing as a function of T
- Low cross-section

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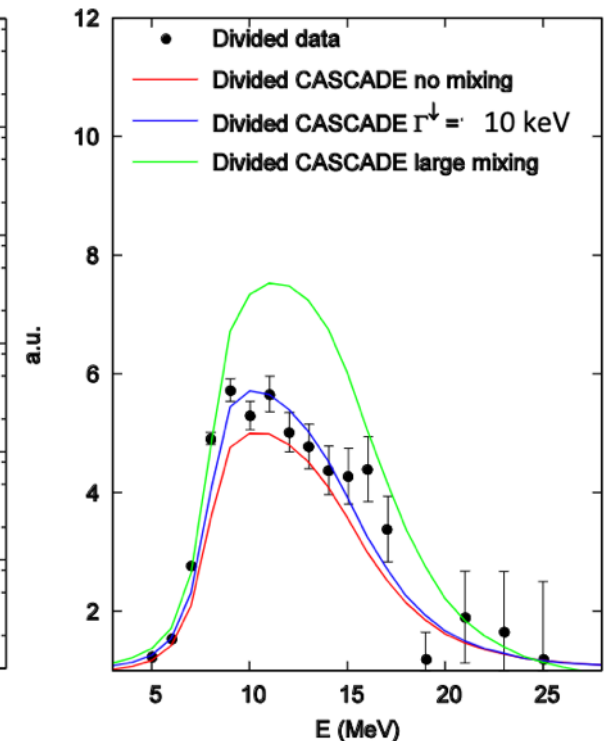
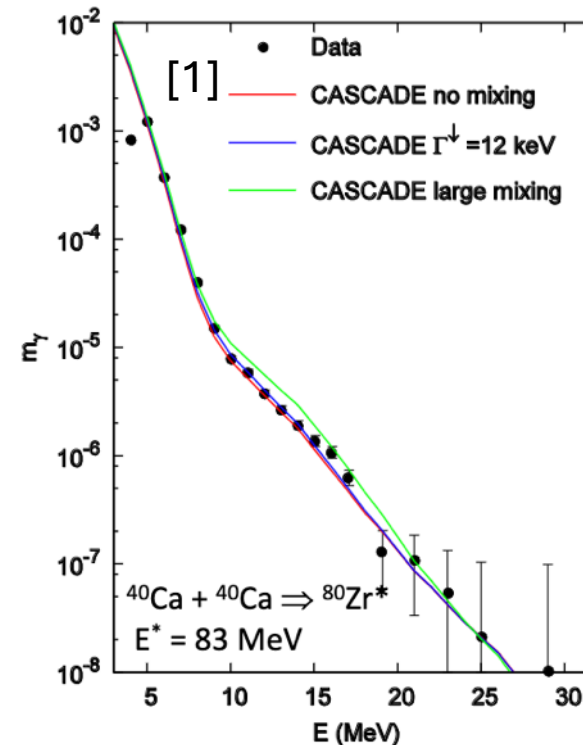
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EXPERIMENTAL SETUP

ELIFANT-GG Array:

High-energy γ rays:

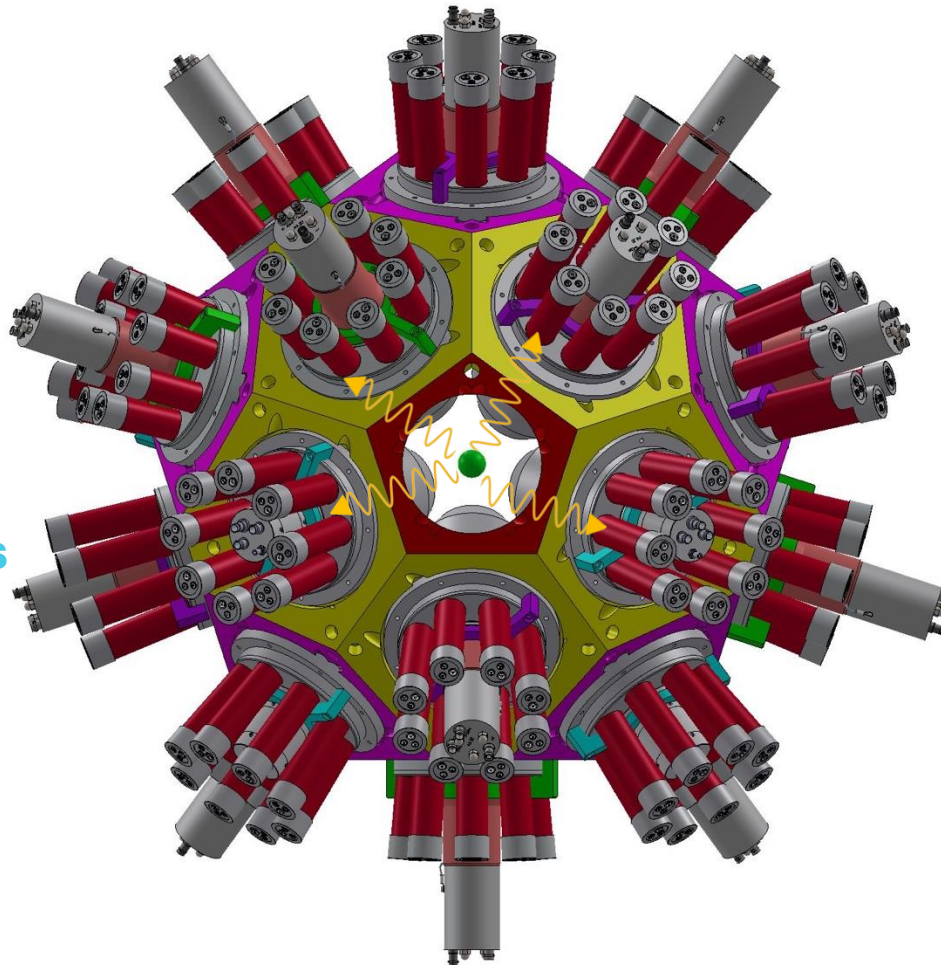
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Detector Array: 11 3''x3''

LaBr₃:Ce + 10 CeBr₃ (3''x3'')
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Fusion-evaporation residues (low-energy γ rays):

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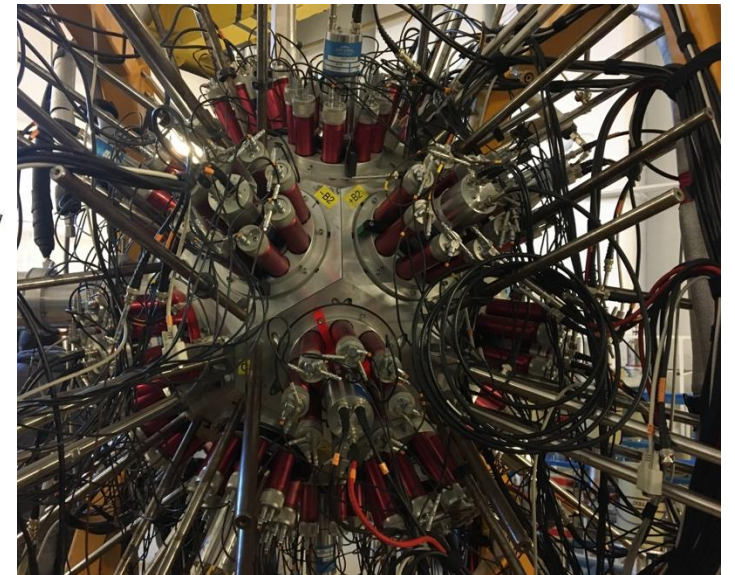
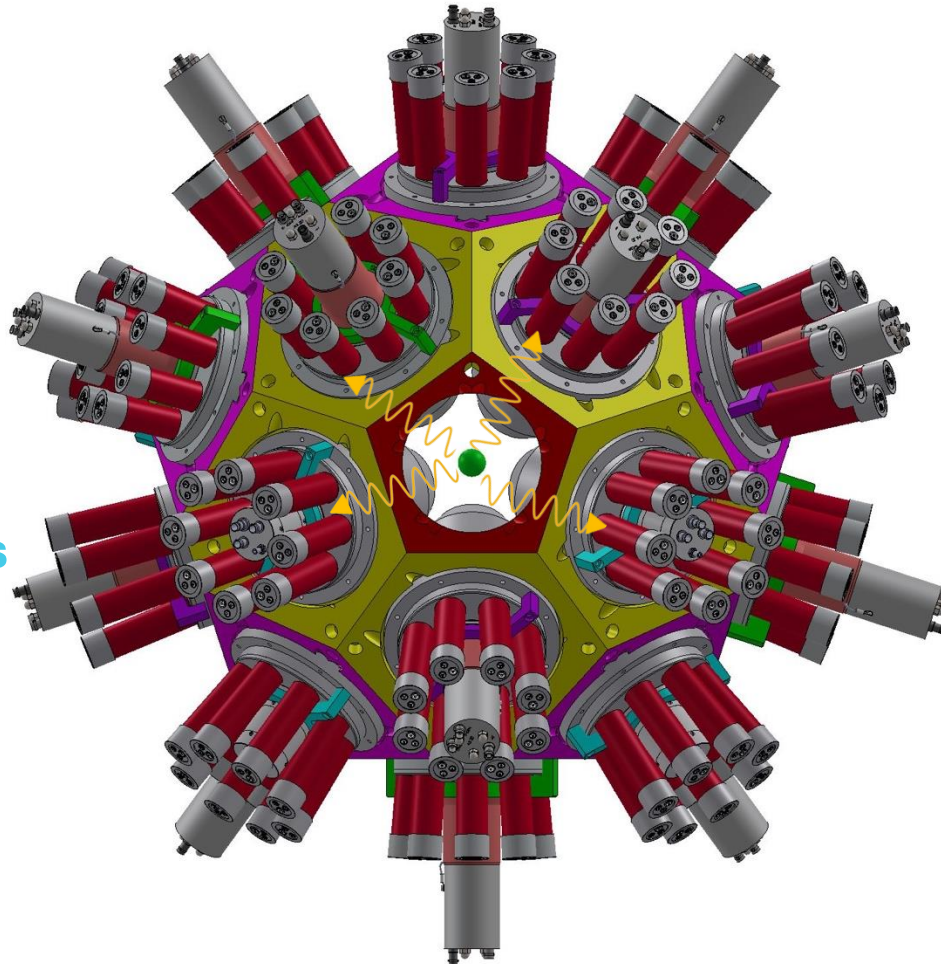
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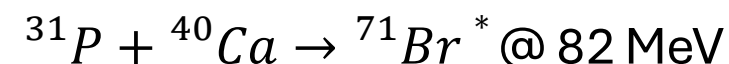
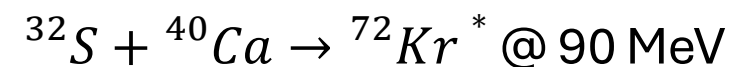
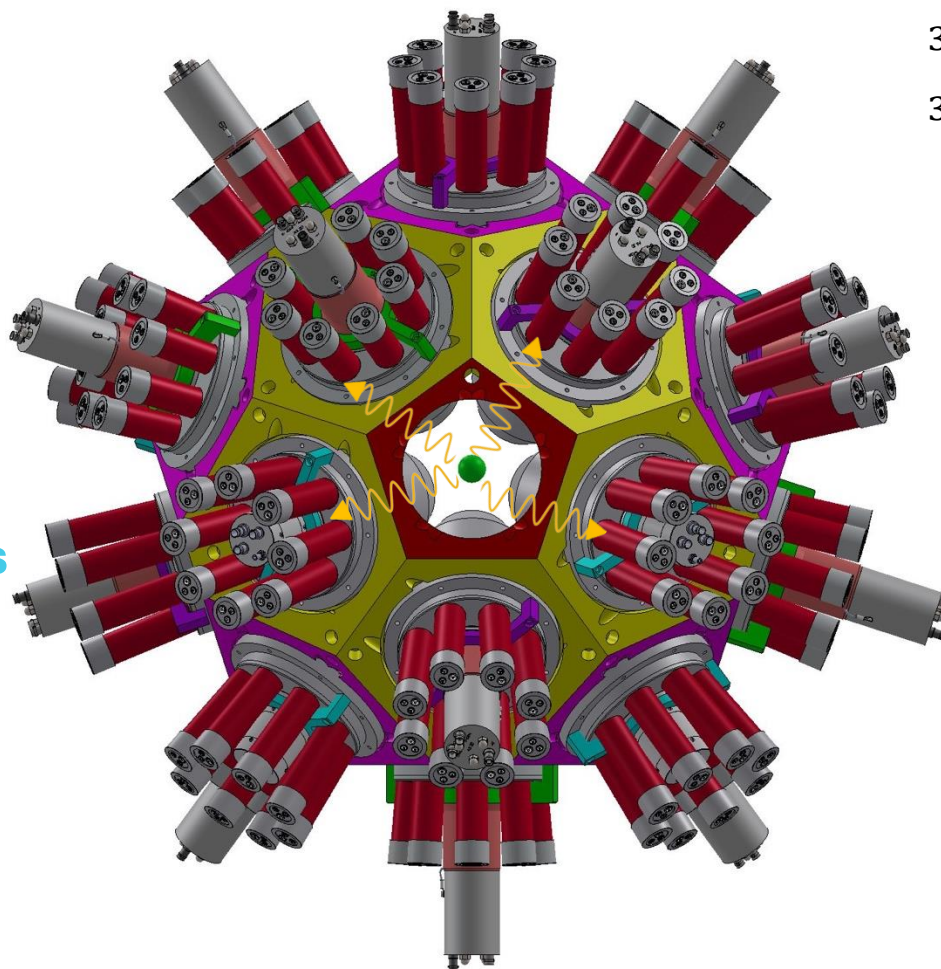
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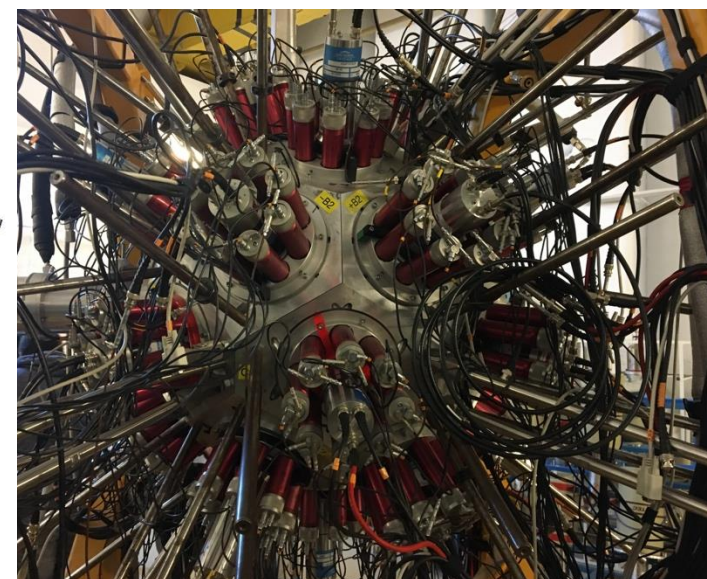
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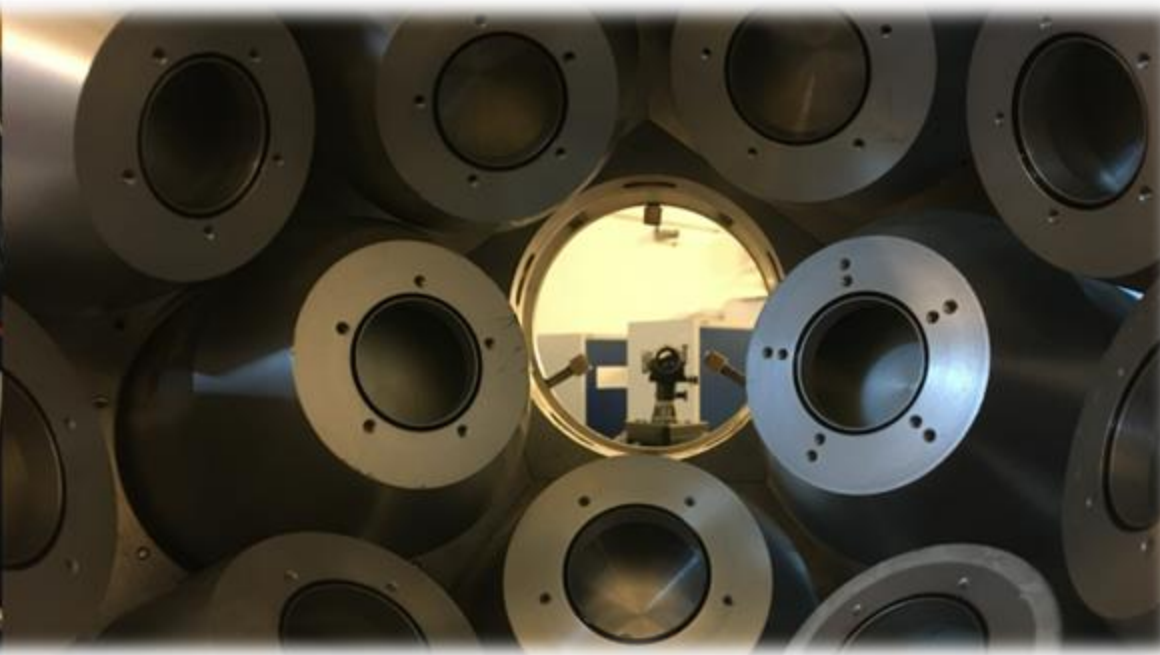
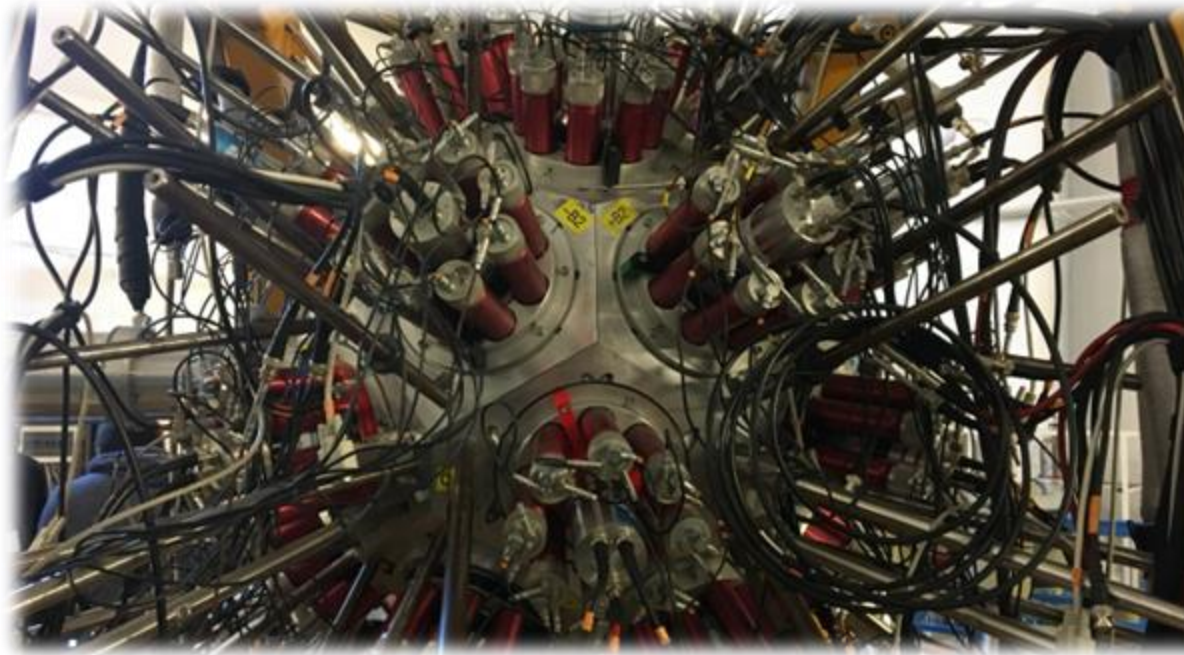


Low cross-section



High efficiency





RESULTS

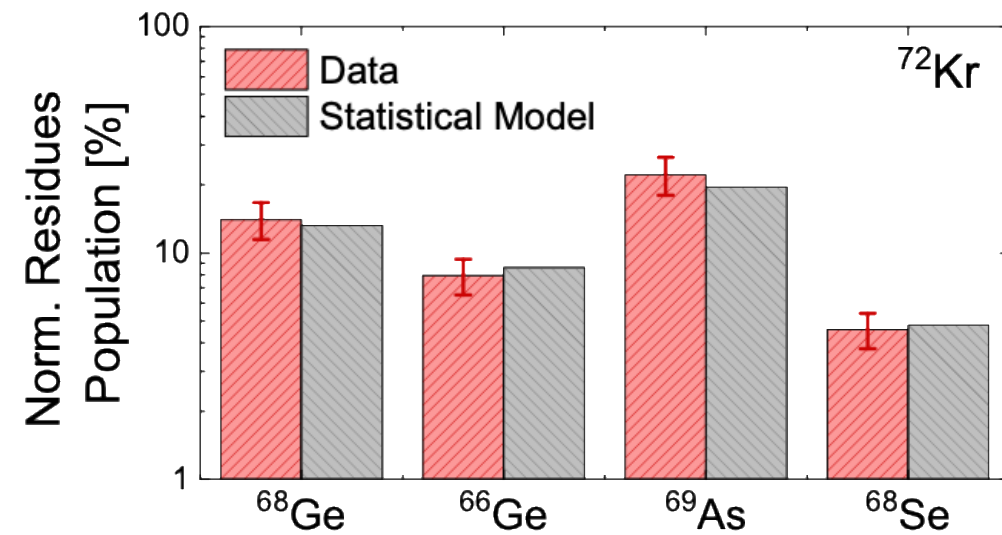
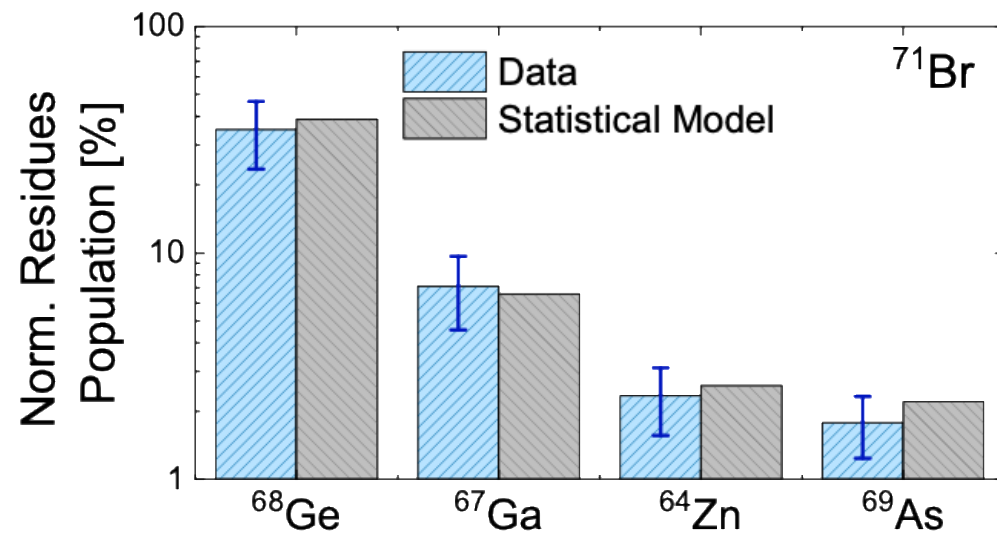
ISOSPIN MIXING PARAMETRES: α^2 & δ_c

RESULTS: EVAPORATION RESIDUES

Fusion-evaporation residues (low-energy γ rays): 4 AC-shield HPGe detectors

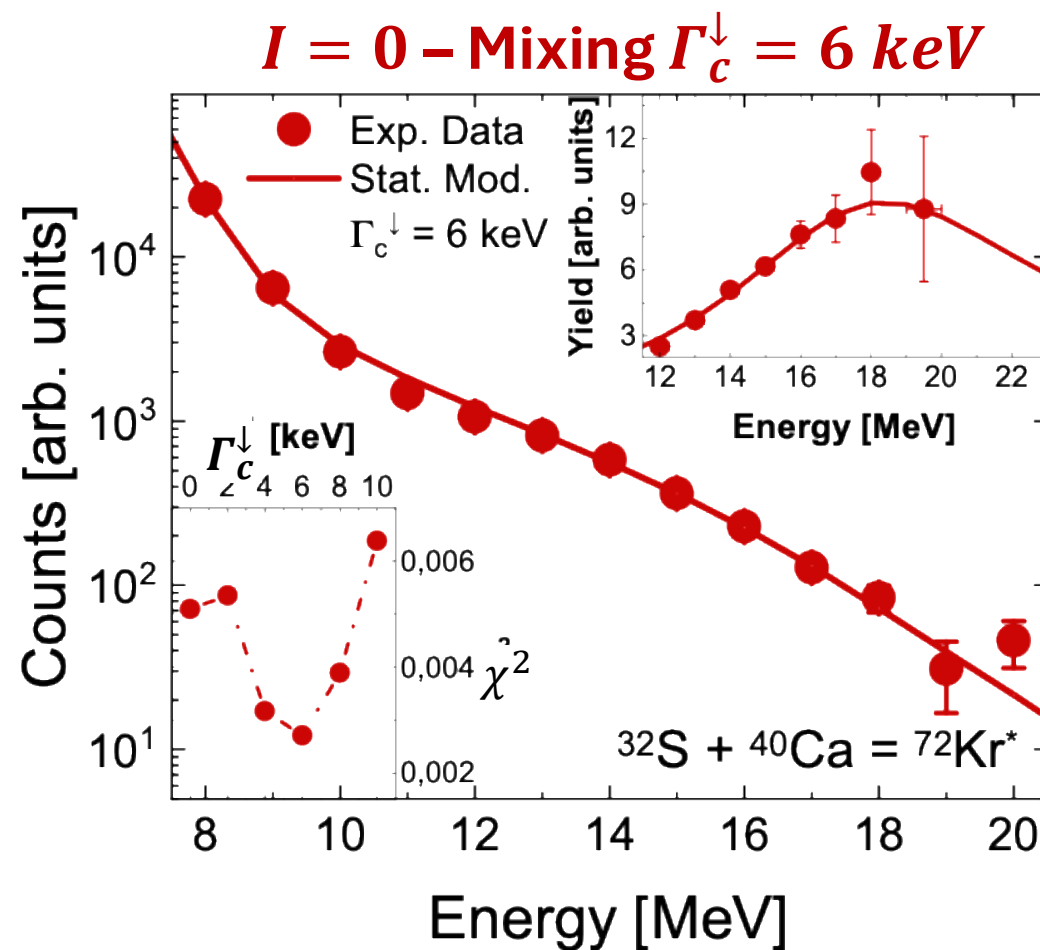
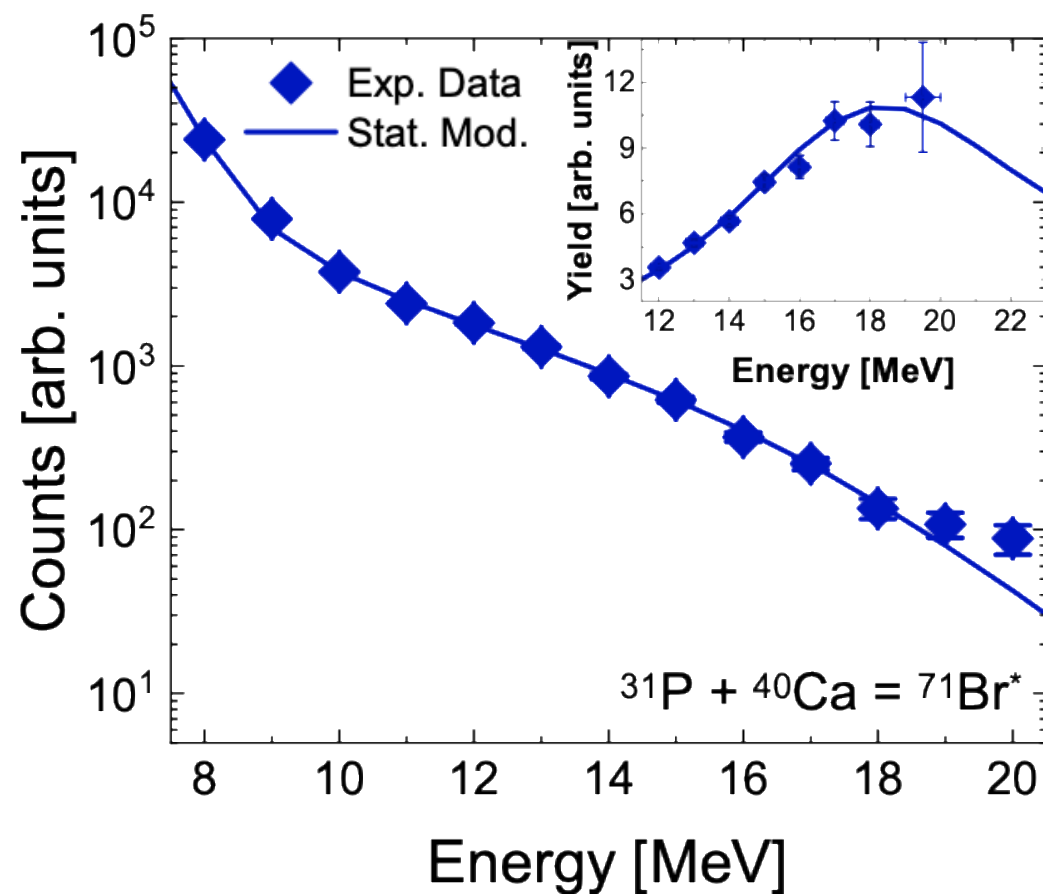
Experimental population of the residues compared with the statistical model prediction.

The most populated residues are reproduced within the error bars.



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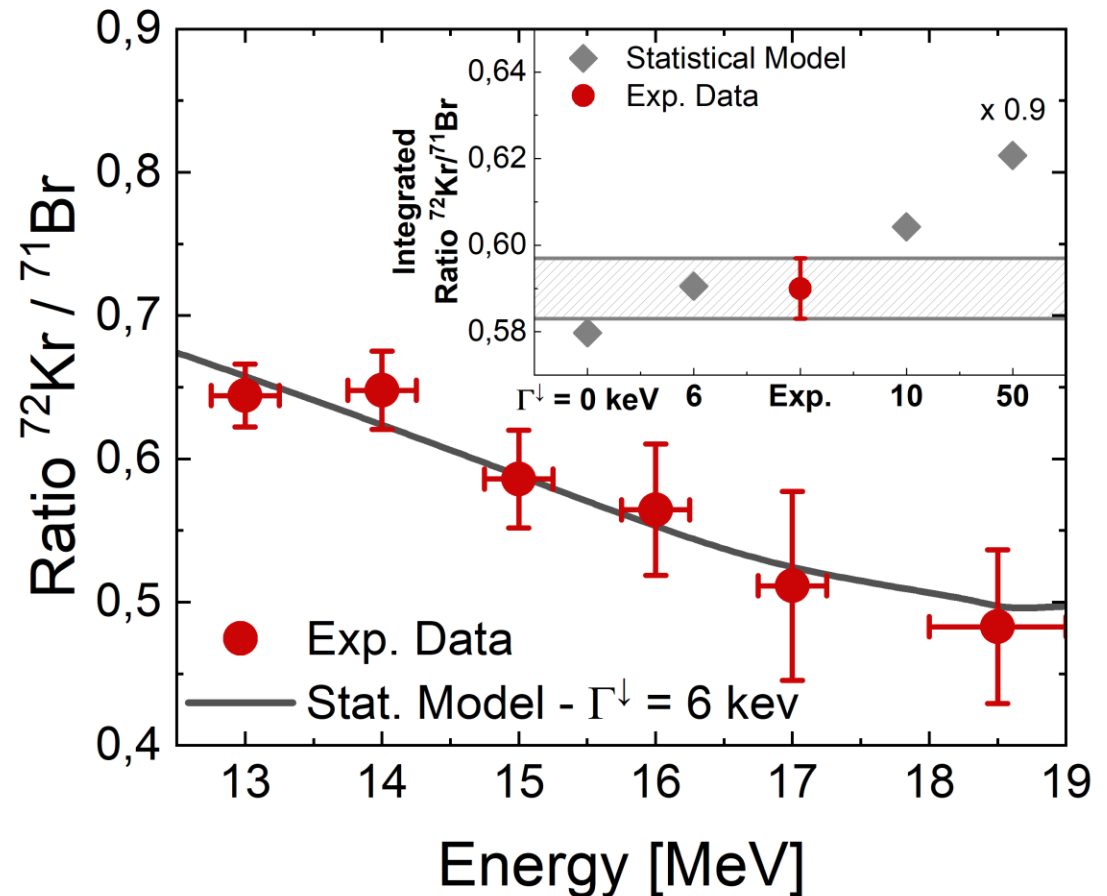
RESULTS: HIGH-ENERGY γ -RAY SPECTRA



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RESULTS: MIXING ESTIMATION

Statistical model with different values of Γ_c^\downarrow : 0 (no mixing), 6, 10 and 50 keV (full mixing).



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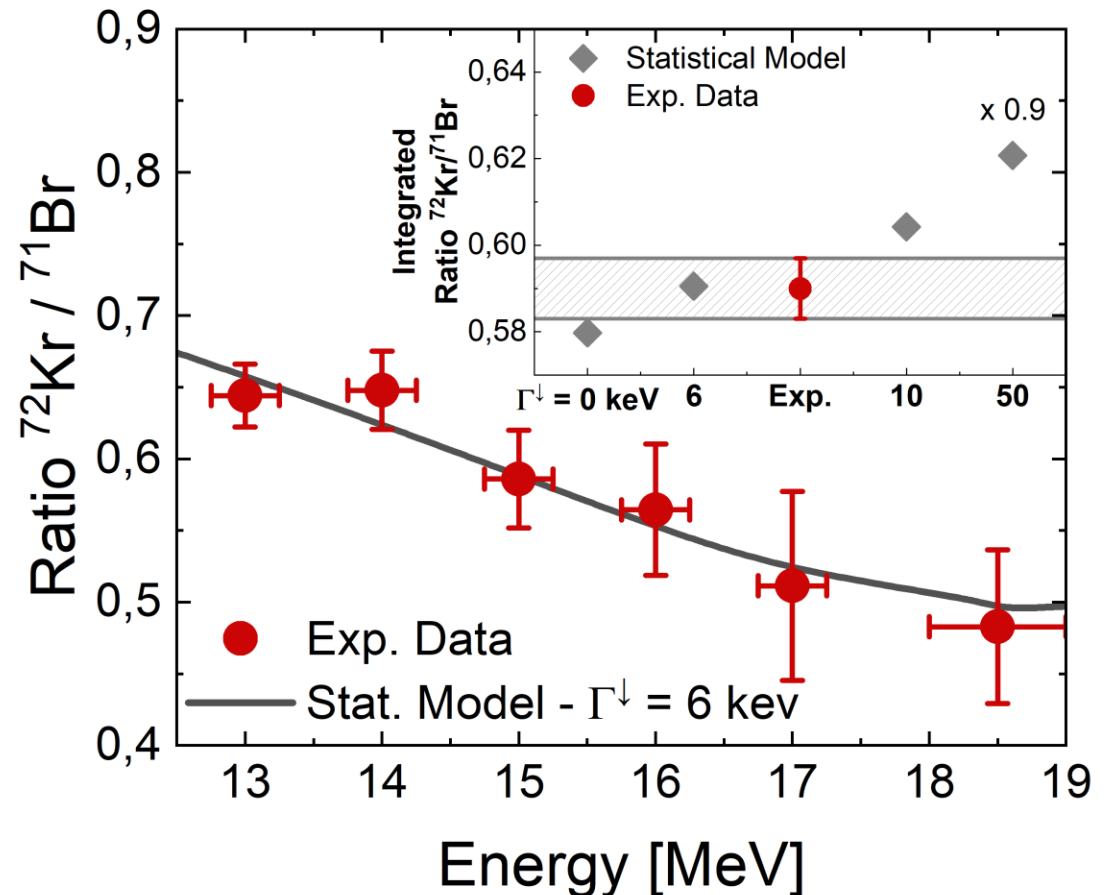
Statistical model with different values of Γ_c^\downarrow : 0 (no mixing), 6, 10 and 50 keV (full mixing).

Integrated Ratio ^{72}Kr and ^{71}Br from 13 MeV to 18 MeV:

- Experimental data: **0.590 ± 0.007**
- Statistical model $\Gamma_c^\downarrow = 6 \text{ keV}$:

0.591

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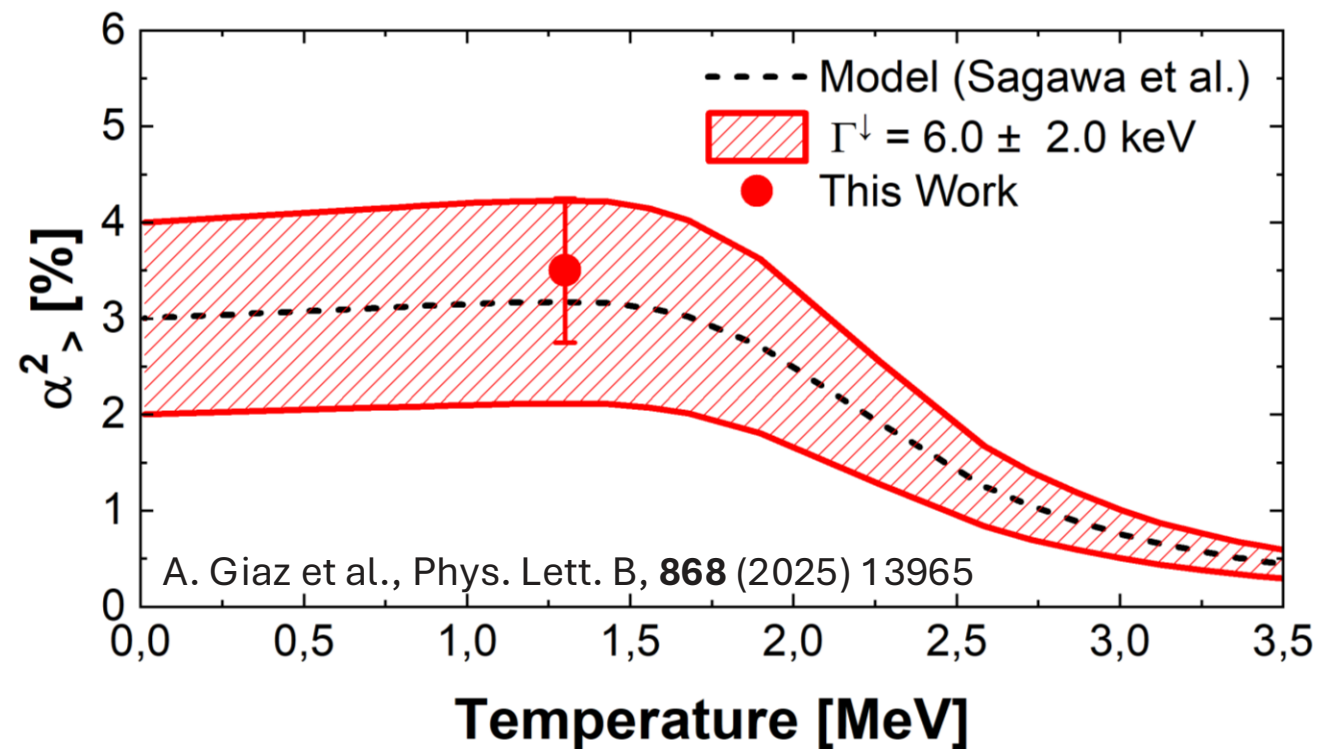


RESULTS: α^2 MIXING PARAMETER (I)

From $T > 0$ to $T = 0$

α^2 trend as a function of temperature:

$$\alpha^2(T) = \frac{1}{I_0 + 1} \frac{\Gamma_{IAS}^\downarrow}{\Gamma_{CN}(T) + \Gamma_{IVM}^{(IAS)}}$$



[1] G.C.H. Sagawa, P.F. Bortignon, Phys. Lett. B, 444, (1998), 1-6

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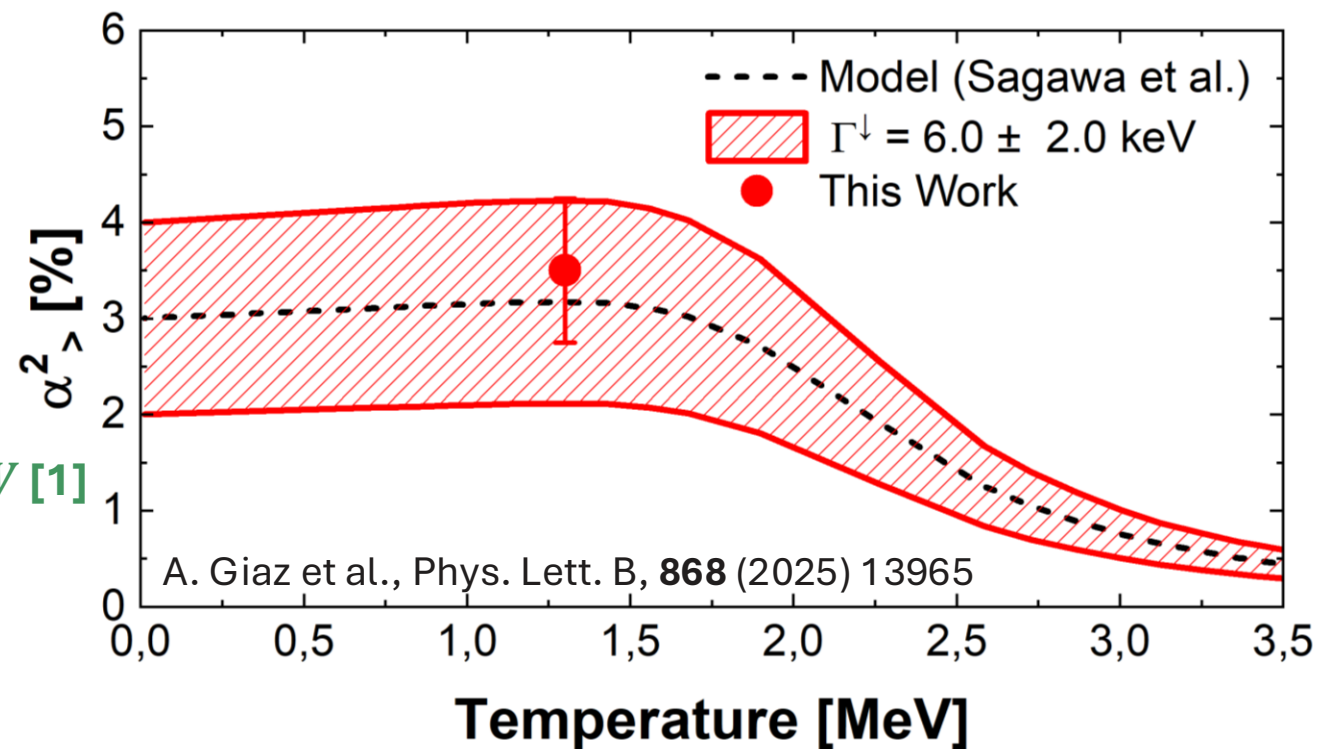
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- $\Gamma_{IVM}^{(IAS)}$ width of IVM resonance at excitation energy of IAS

200 keV [1]



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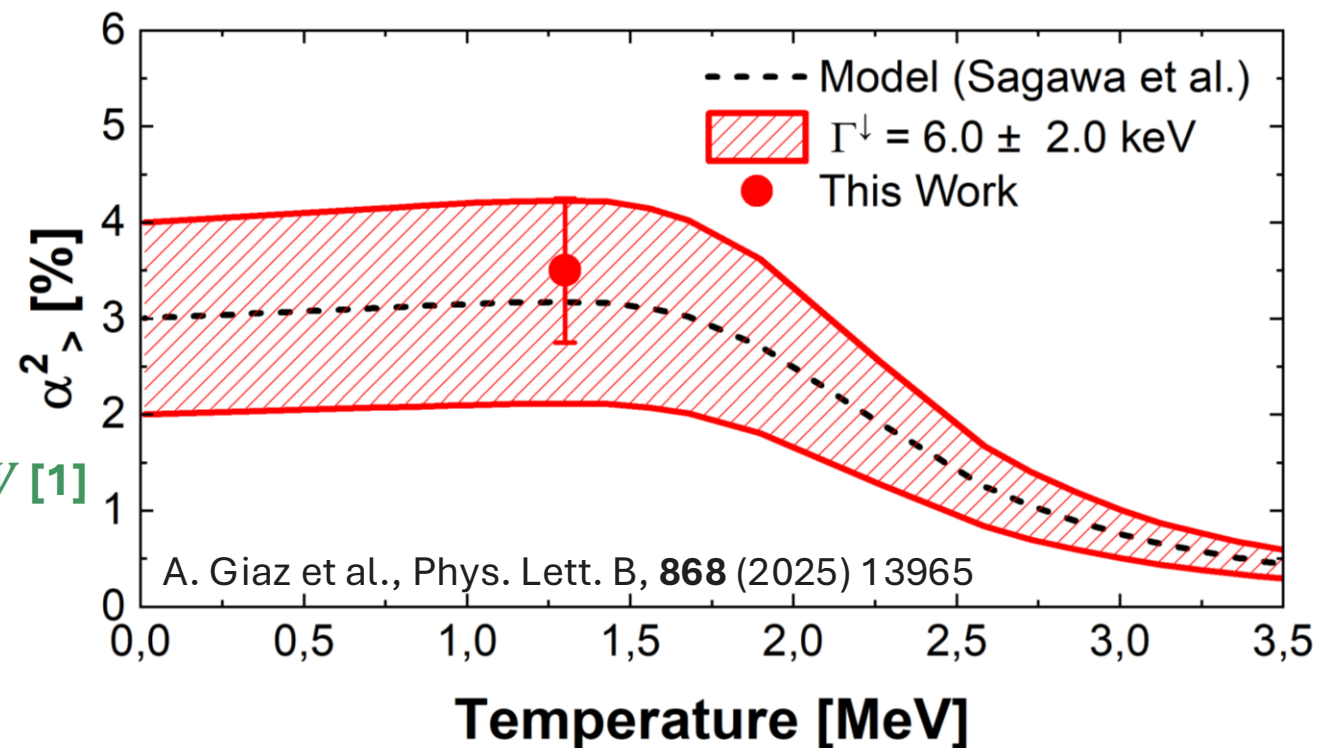
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From Stat. Model

- $\Gamma_{IVM}^{(IAS)}$ width of IVM resonance at excitation energy of IAS
- $\Gamma_{CN}(T)$ is the compound nucleus decay width

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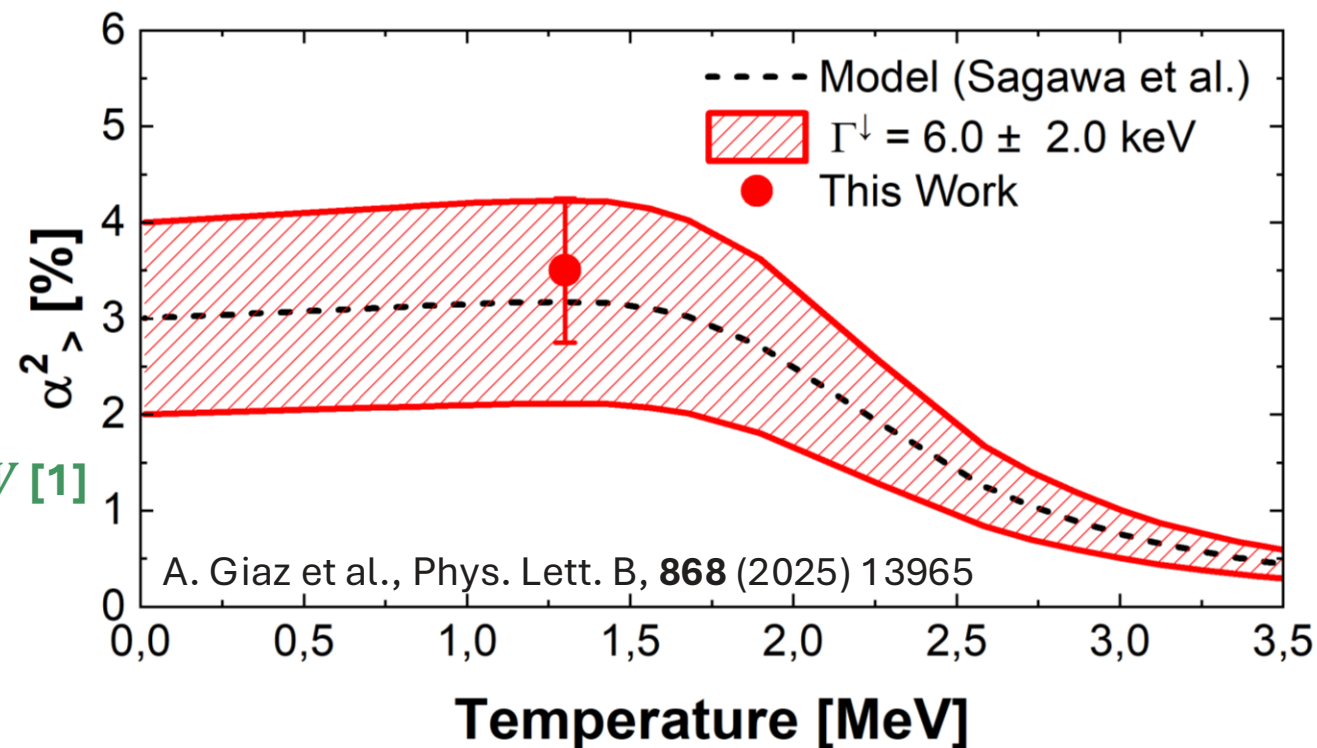
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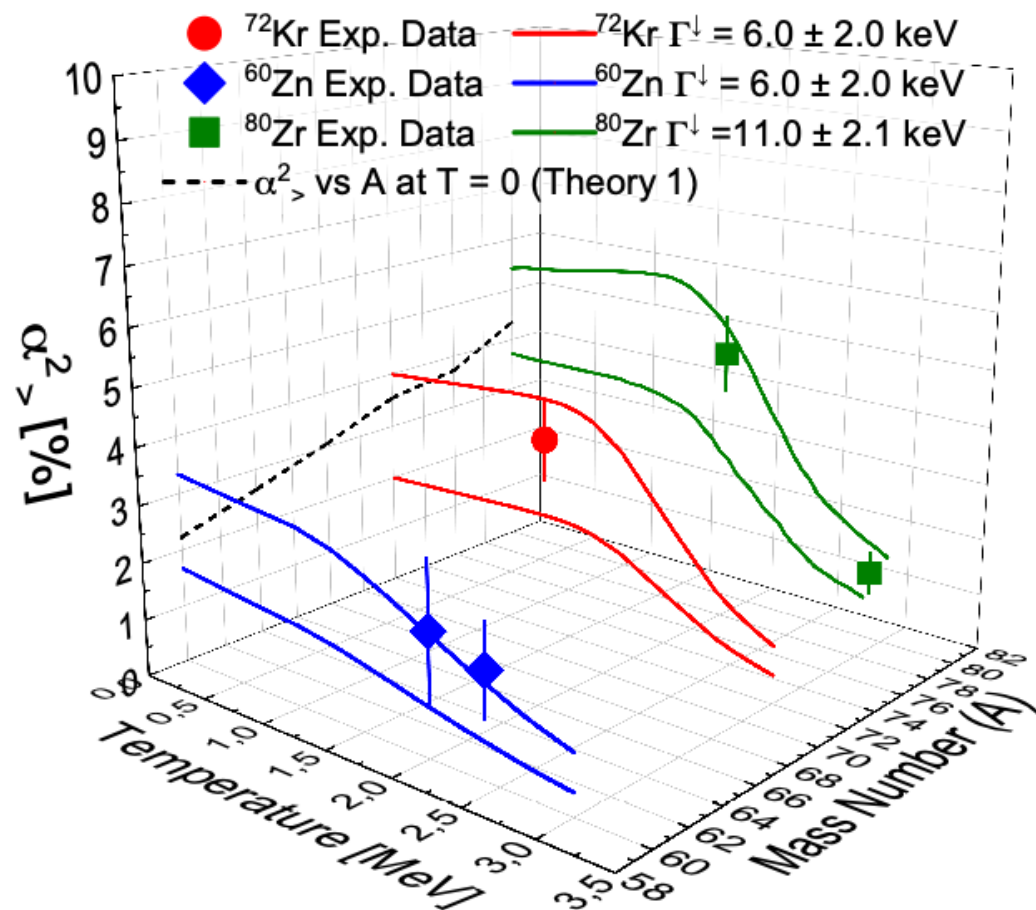
6.0 \pm 2.0 keV (pointing to Γ_{IAS}^\downarrow)
From Stat. Model (pointing to $\Gamma_{CN}(T)$)
200 keV [1] (pointing to $\Gamma_{IVM}^{(IAS)}$)

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RESULTS: α^2 MIXING PARAMETER (II)

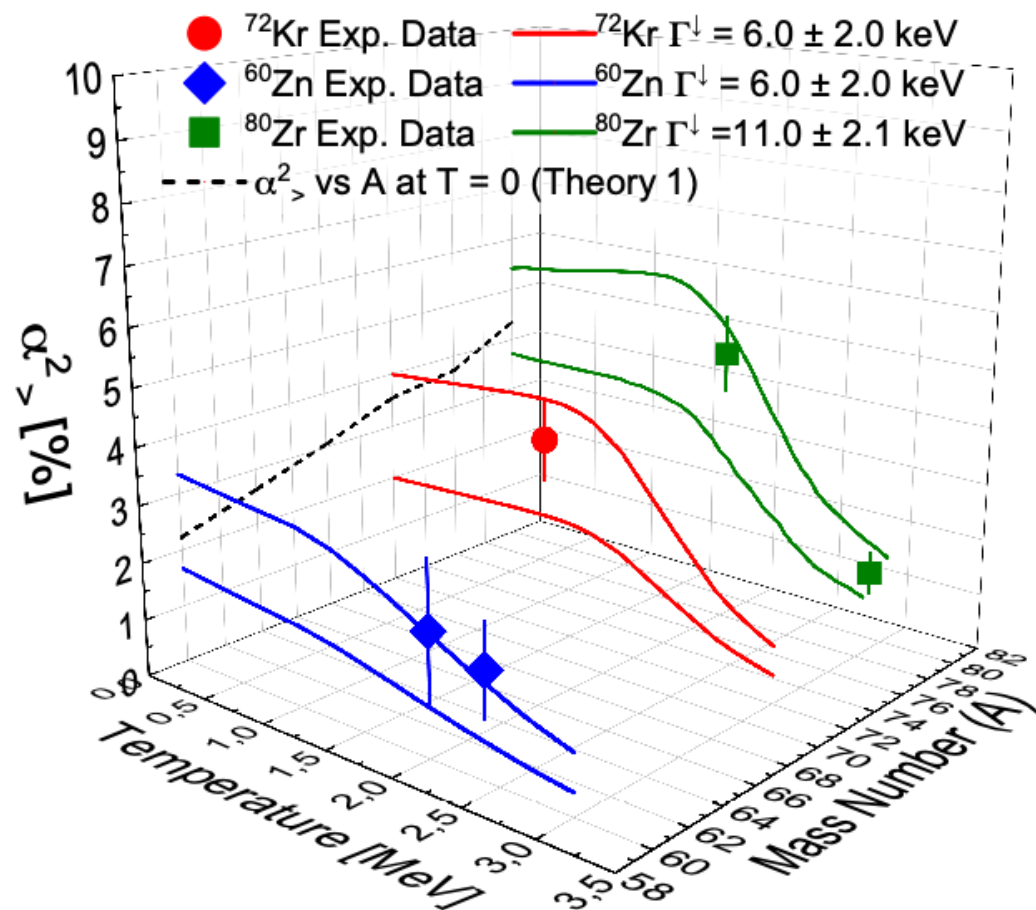


From finite temperature to zero temperature:

- ^{60}Zn : $(2.5 \pm 0.8) \% @ T = 0$ [1]
- ^{80}Zr : $(4.6 \pm 0.9) \% @ T = 0$ [2]
- ^{72}Kr : $(3.0 \pm 0.8) \% @ T = 0$ and $(3.5 \pm 0.8) \% @ T = 1.3 \text{ MeV}$ [3]

- [1] G. Gosta et al., Phys. Rev. C **103**, (2021) L041302
 [2] S. Ceruti et al., Phys. Rev. Lett. **115**, (2015) 222502
 [3] A. Giaz et al., Phys. Lett. B, **868** (2025) 13965

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Mixing Probability α^2 follows the expected trend both as a function of **Mass Number (A)** and **Temperature**

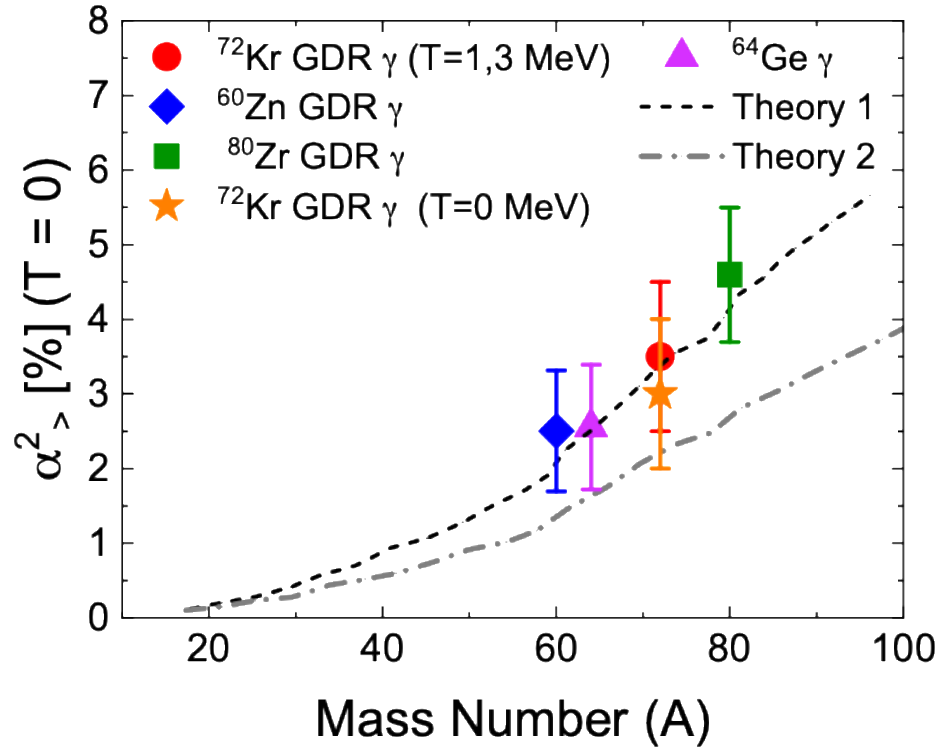
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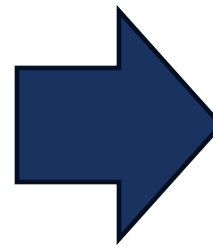
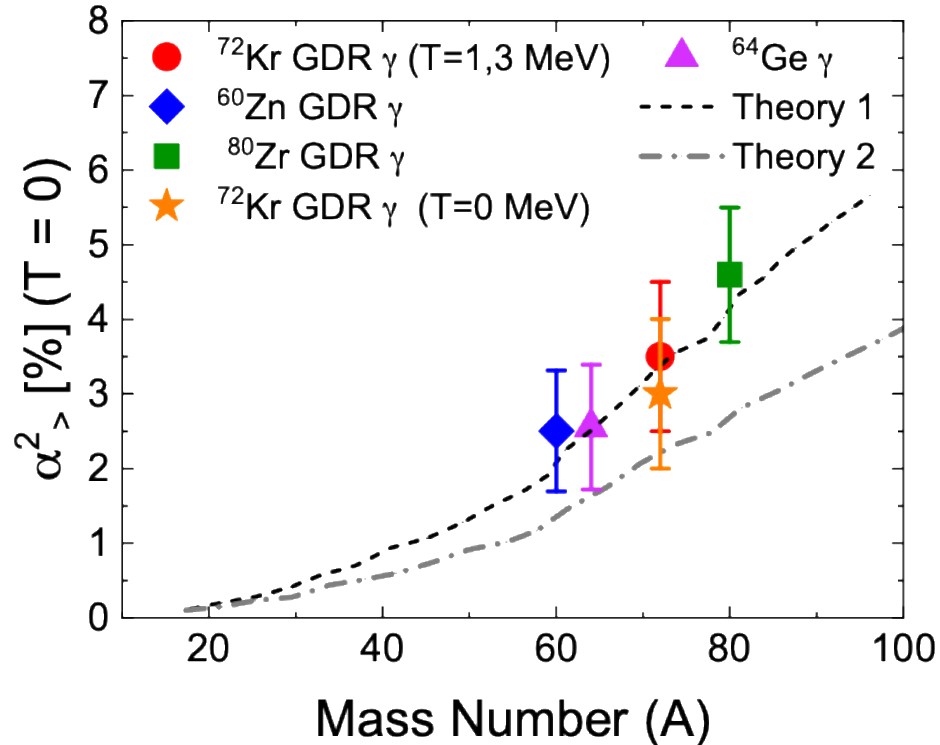
RESULTS: CORRECTION TERM δ_c

From α^2



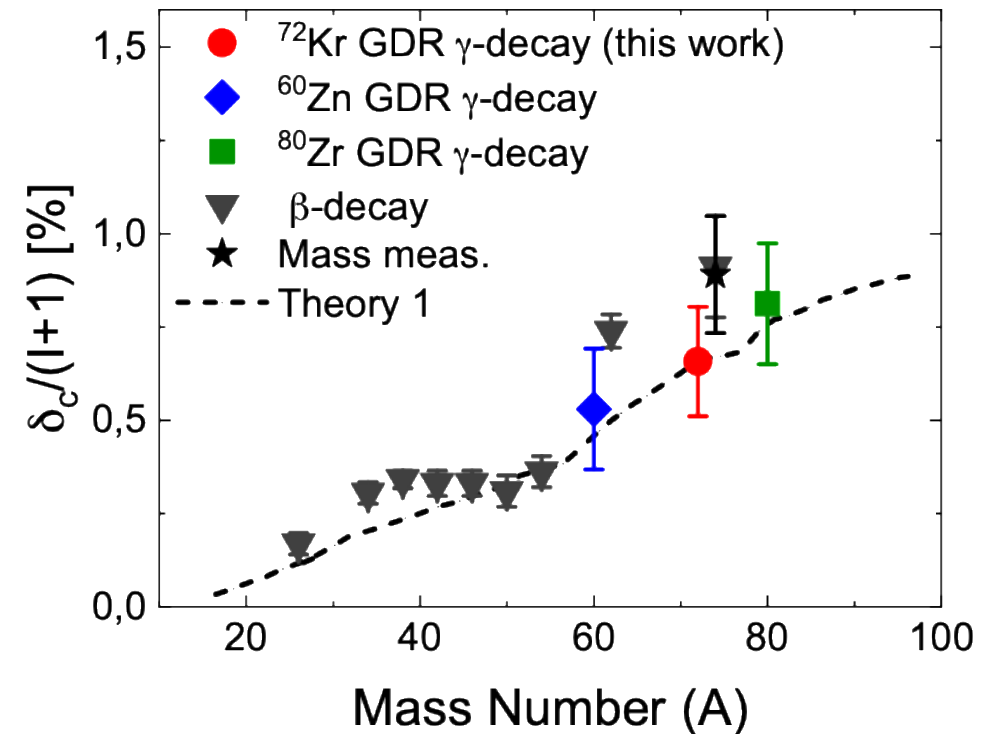
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$$\text{To } \delta_c = 4(I + 1) \frac{V_1}{41\xi A^{2/3}} \alpha^2$$

$$\delta_c = (0.66 \pm 0.14)\%$$



SUMMARY AND PRESPECTIVES

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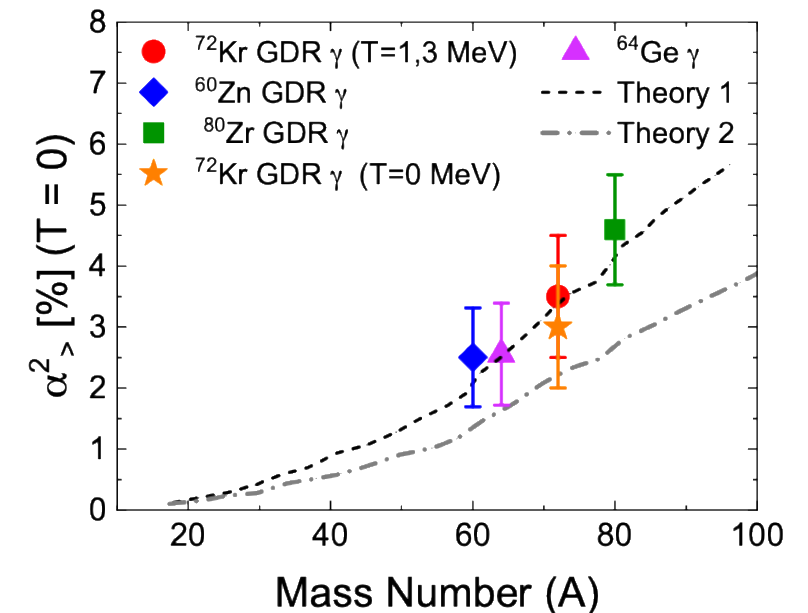
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PRESPECTIVES

- Exotic beam to reach $N=Z$ nuclei with $A > 80$
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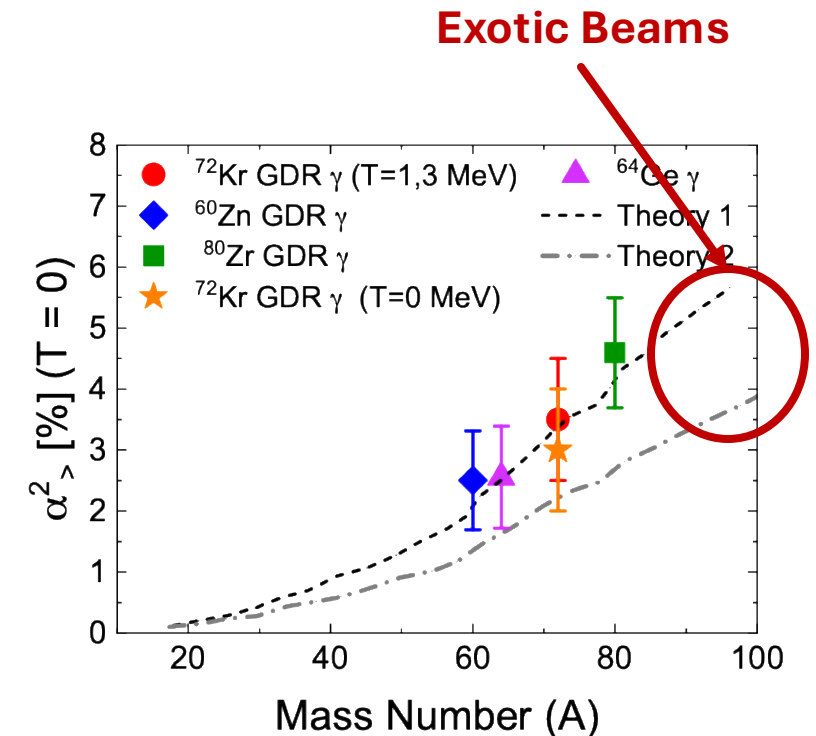
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COLLABORATORS

Agnese Giaz^a, Oliver Wieland^a, Angela Bracco^{b,a}, Franco Camera^{b,a}, Soichiro Aogaki^c, Dimiter L. Balabanskic, Maria Brezeanu^c, Ruxandra Borcea^d, Marian Boromiza^d, Ion Burducea^e, Stefana Calinescu^d, Adina Coman^d, Paul Constantin^c, Cristian Costache^d, Michał Ciemata^f, Gheorghe Ciocan^d, Cristina Clisu^d, Fabio C.L. Crespi^{b,a}, Mihai Cuciuc^c, Anukul Dhal^c, Irina Dinescu^d, Nikolay Djourellov^c, Nicoleta M. Florea^d, Ioana Gheorghe^d, Decebal Iancu^e, Violeta Iancu^c, David M. Kahl^c, Maria Kmiecik^f, Asli Kuşoğlu^{c,g}, Razvan Lica^d, Adam Maj^f, Nicolae M. Mărginean^d, Raluca Mărginean^d, Constantin Mihai^d, Radu E. Mihai^{d,h}, Bénédicte Million^a, Catalin Neacsu^d, Dragos Nichita^c, Cristina R. Nită^d, Haridas Pai^c, Alfio D. Pappalardo^c, Teodora Petruse^c, Adrian Rotaru^c, Andreea B. Serban^c, Pär-Anders Söderström^c, Cristophe O. Sotty^d, Anamaria Spătaru^c, Lucian Stan^d, Alexandru N. State^c, Alexandru Stoica^d, Dmitry A. Testov^c, Sebastian Toma^d, Tatiana Tozar^c, Andrei Turturică^d, Gabriel V. Turturică^c, Sorin Ujeniuc^d, Vlad Vasilca^c, Yi Xu^c, Fan Zhu^c

^aINFN, sezione di Milano, Via Celoria, 16, Milano, 20133, MI, Italy

^bUniversità degli studi di Milano and INFN, sezione di Milano, Via Celoria, 16, Milano, 20133, MI, Italy

^cExtreme Light Infrastructure-Nuclear Physics, Str. Reactorului 30, Bucharest-Măgurele 077125, Romania

^dDepartment of Nuclear Physics, Str. Reactorului 30, Bucharest-Măgurele, 077125, Romania

^eApplied Nuclear Physics Department, Str. Reactorului 30, Bucharest-Măgurele, 077125, Romania

^fThe Henryk Niewodniczanski Institute of Nuclear Physics, Polish Academy of Sciences, ul. Radzikowskiego 152, Krakow, 31-342, Poland

^gDepartment of Physics, Faculty of Science, Istanbul University, Vezneciler/Fatih, Istanbul, 34134, Turkey

^hInstitute of Experimental and Applied Physics, Czech Technical University in Prague, Husova 240/5, Prague, 110 00, Czech Republic

Thank you for the attention!

