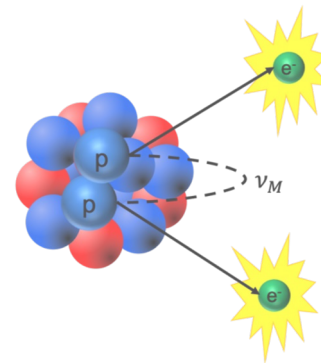




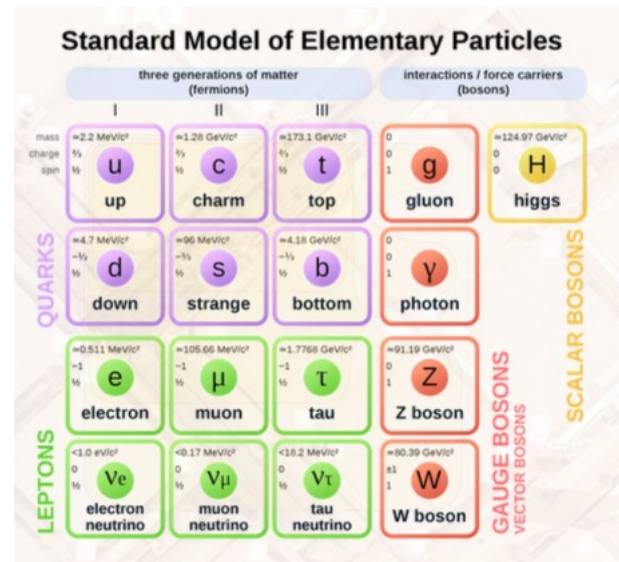
Neutrinoless Double Beta Decay: where we stand and what lies ahead

Claudia Nones
CEA/IRFU



The context: ghostly neutrinos

- The most poorly understood particle within the SM
- First postulated by Pauli in 1930 to explain missing energy in beta decay
- Detected by Reines & Cowan in 1956 (nuclear reactor experiment)
- According to SM: electrically neutral, massless, weakly interacting



- Neutrino oscillations: evidence that neutrinos have mass
- **Physics beyond the Standard Model**

The Majorana Neutrino

- In the **Standard Model**, all the fermions are so-called **Dirac particles** and have an anti-fermion partner.
- In 1937 E. Majorana proposed that in the case of a neutral particle, the fermions and the anti-fermions could coincide. **The neutrino could therefore be a « Majorana particle ».**
- **A Majorana neutrino mass term could then be added to the Standard Model's Lagrangian**, and the « See-Saw » mechanism could explain the smallness of the mass eigenstate by predicting a sterile « right-handed » and heavy neutrino.

Majorana neutrino

Violation of the lepton number conservation

With see-saw mechanism the small mass is explained



$$\nu = \nu^C$$



$$\nu \neq \nu^C$$

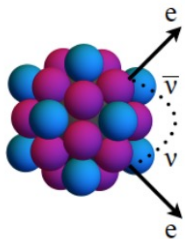
Dirac neutrino

Lepton number conservation

But why the mass is that small ?

The question is still pending if the neutrino is a Dirac or a Majorana neutrino
Answering it is crucial for our understanding of particle physics.

Nuclear Double Beta Decay in a nutshell

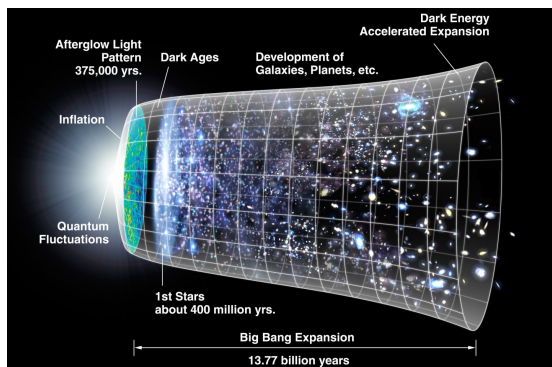


Double Beta Decay is the **rarest nuclear weak process**

It takes place between **two even-even isobars**

$$(A, Z) \rightarrow (A, Z+2) + 2e^- + (...)$$

If only electrons and nothing else:



$\times 10^{15}$

- Need to find single events in a ton of isotope x year(s) of exposure!
- 3×10^{-14} Bq/g
- We go to extreme length to limit ubiquitous radioactivity

Half-life larger $> 10^{25}$ y - 10^{26} y

15 Bq /
banana



More details on double beta decay

Double beta decay is a nuclear transition involving two even-even isobars

① $(A,Z) \rightarrow (A,Z+2) + 2e^- + 2\nu_e$



2ν Double Beta Decay
allowed by the Standard Model
already observed – $\tau \sim 10^{19} - 10^{21} \text{ y}$

② $(A,Z) \rightarrow (A,Z+2) + 2e^-$



$2n \rightarrow 2p + 2e^-$



Neutrinoless
Double Beta Decay ($0\nu 2\beta$)
never observed $\tau > 10^{24} - 10^{26} \text{ y}$

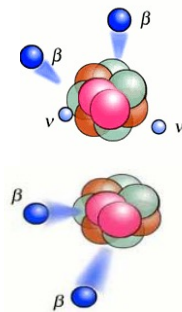
Creation of matter without antimatter partners

Beyond Standard Model

Lepton Number Violation (LNV)



The only currently viable experimental approach to
probe the Majorana nature of neutrino



More details on double beta decay

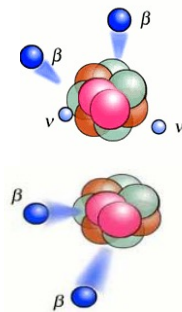
Double beta decay is a nuclear transition involving two even-even isobars



2ν Double Beta Decay
allowed by the Standard Model
already observed – $\tau \sim 10^{19} - 10^{21} \text{ y}$



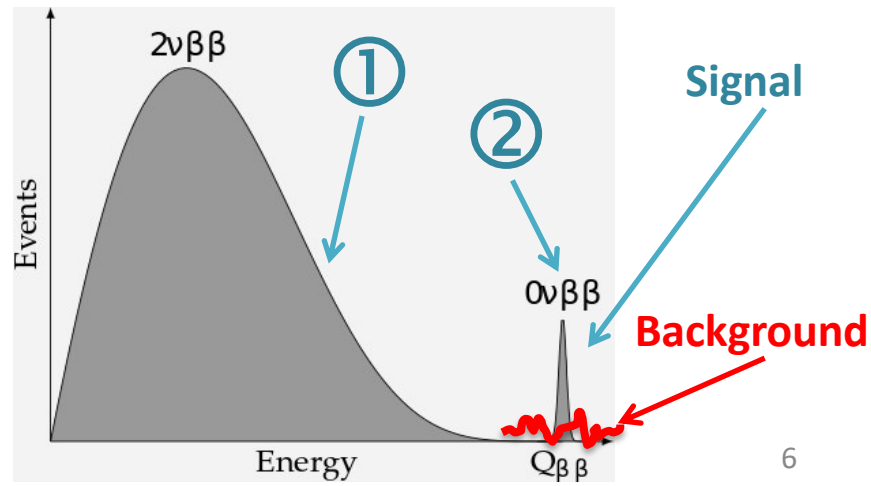
Neutrinoless
Double Beta Decay (0ν2β)
never observed $\tau > 10^{24} - 10^{26} \text{ y}$



Experimental signatures based on the
Sum energy spectrum of the two electrons

$$Q_{\beta\beta} \sim 2\text{-}3 \text{ MeV}$$

for the most promising candidates



Why neutrinoless Double Beta Decay is important

- Majorana nature of neutrino (irrespectively of the mechanism)
- See-saw mechanism \Rightarrow naturalness of small neutrino masses
- Leptogenesis and matter-antimatter asymmetry in the Universe
- Neutrino mass scale and hierarchy

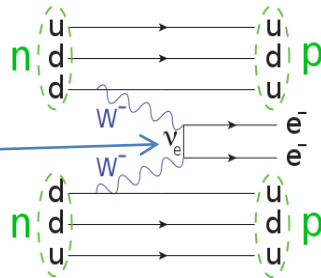
Standard mechanism: **neutrino physics**

$0\nu 2\beta$ is mediated by

light massive Majorana neutrinos

(exactly those which oscillate)

Sometimes defined “**mass mechanism**”



Non-standard mechanisms:

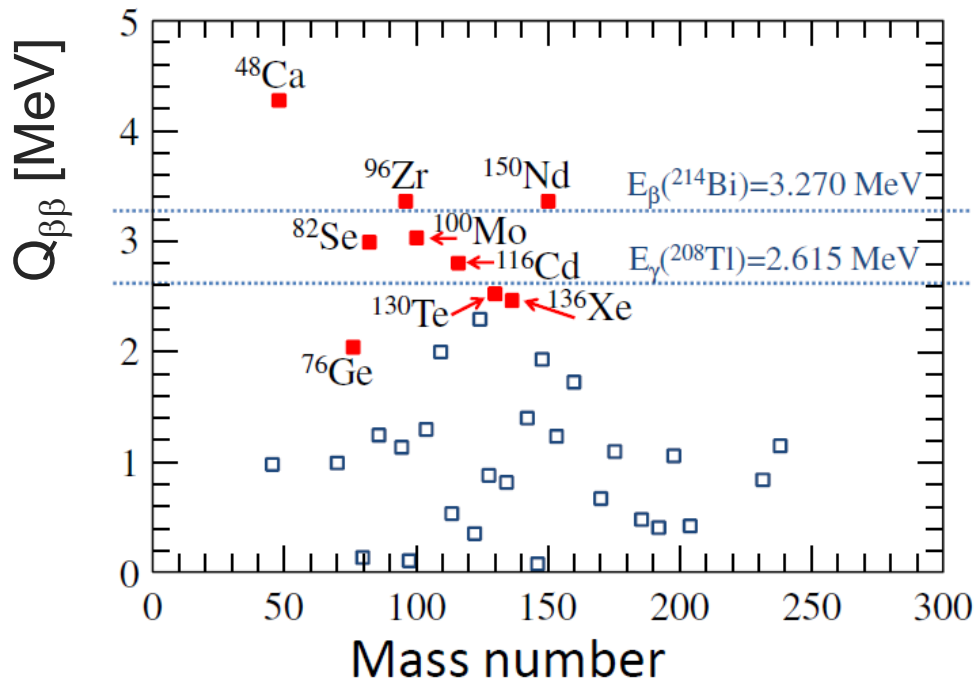
Sterile ν , LNV,...

Not necessarily neutrino physics

- Minimal straightforward extension of the Standard Model
- Metric to compare experiments and technologies

Which and how many nuclei?

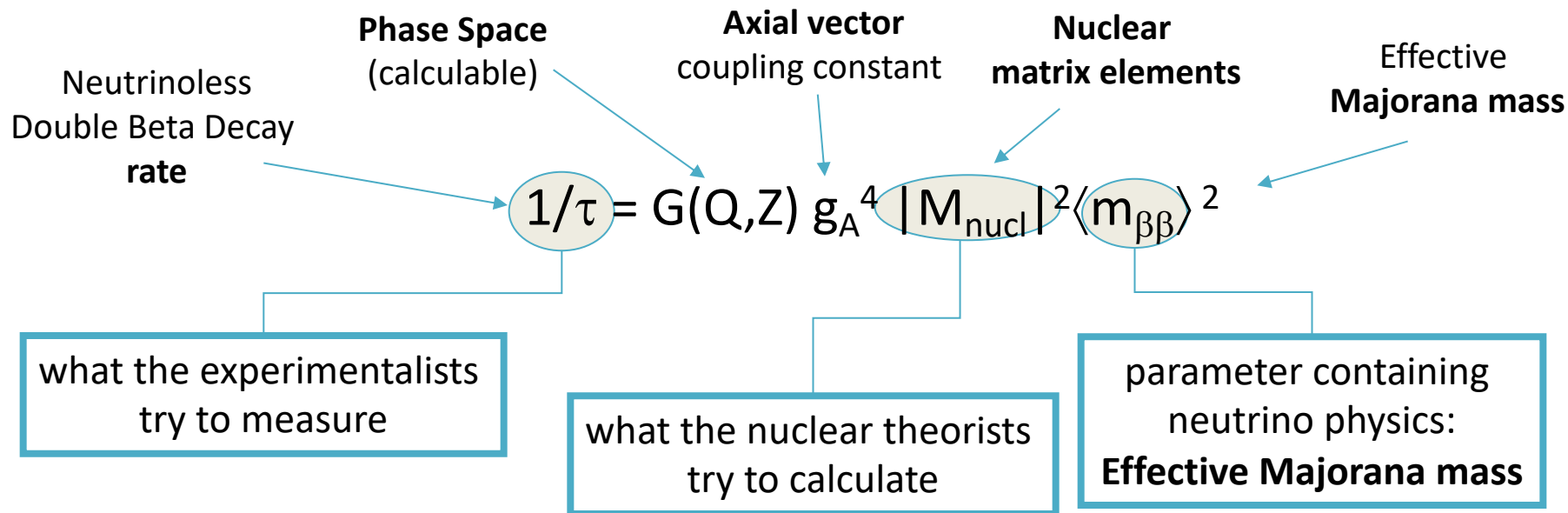
Double Beta Decay is the main decay channel for 35 nuclei, with a large span of $Q_{\beta\beta} \rightarrow$ energy available for the decay products



**Most promising
candidates**
High $Q_{\beta\beta}$

Rate in case of mass mechanism

how **0ν-DBD** is connected to **neutrino mixing matrix** and **masses**
in case of process induced by light ν exchange (**mass mechanism**)



$$\langle m_{\beta\beta} \rangle = \left| |U_{e1}|^2 m_1 + e^{i\alpha_1} |U_{e2}|^2 m_2 + e^{i\alpha_2} |U_{e3}|^2 m_3 \right|$$

Majorana phases

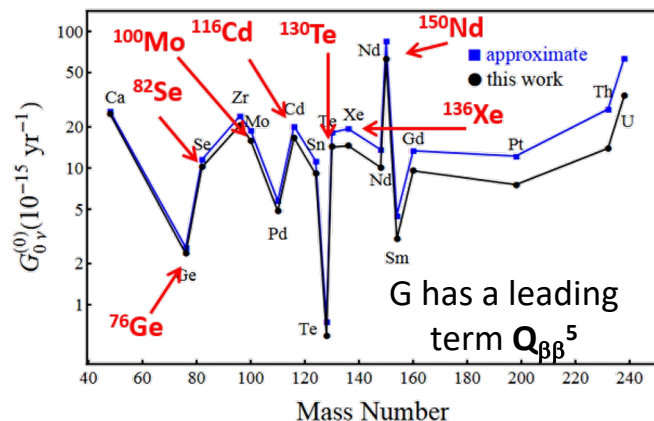
Phase space and nuclear matrix elements



WE WANT YOU

$$1/\tau = G(Q,Z) g_A^4 |M_{\text{nuc}}|^2 \langle m_{\beta\beta} \rangle^2$$

Phase space: exactly calculable



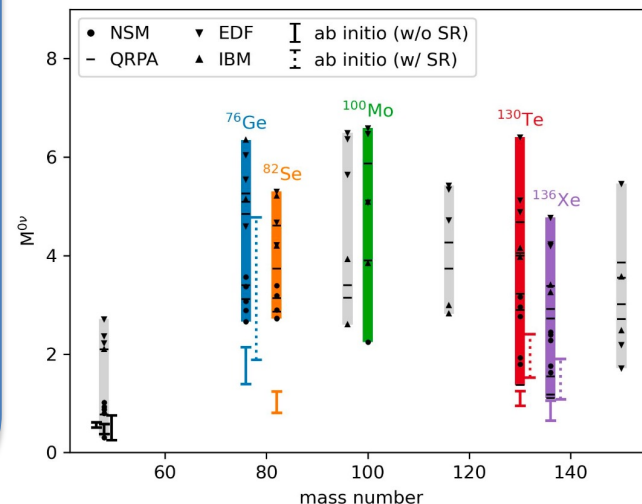
Phys. Rev. C 85, 034316 (2012)

$$g_A = \begin{cases} 1.269 & \text{Free nucleon} \\ 1 & \text{Quark} \end{cases}$$

$g_{A,\text{eff}} \sim 0.6 - 0.8$ to be taken (« quenching ») to describe β and $2\nu\beta\beta$ rates with current nuclear models

- Controversial
- Ab-initio calculation with unquenched g_A are required
- Progress ongoing

- NMEs encode the structure of the parent and daughter nuclides
- Several nuclear models



Credit Christoph Wiesinger – TAUP 2025

Phase space and nuclear matrix elements

$$1/\tau = G(Q,Z) g_A^4 |M_{\text{nuc}}|^2 \langle m_{\beta\beta} \rangle^2$$



WE WANT YOU

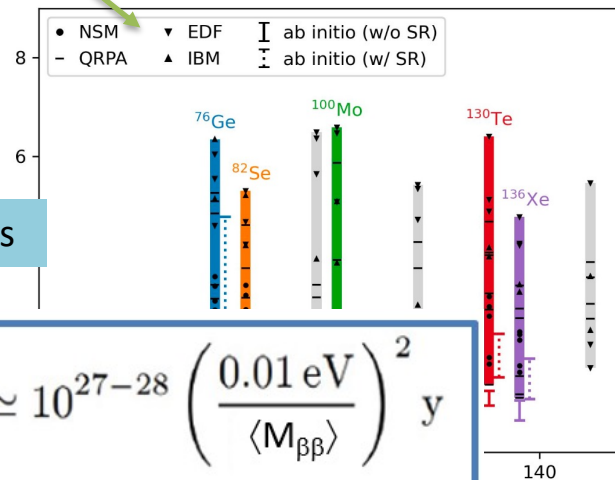
Phase space: exactly calculable

$$g_A = \begin{cases} 1.269 & \text{Free nucleon} \\ 1 & \text{Quark} \end{cases}$$

$g_{A,\text{eff}} \sim 0.6 - 0.8$ to be taken
(“quenching”) to describe

- Typically, a factor ~ 3 of spread for $|M_{\text{nuc}}|$ among the various models
- A spread of ~ 10 is expected in the half life predictions
- Again a factor ~ 3 of spread is expected for $\langle M_{\beta\beta} \rangle$

- Ab-initio calculation w
unquenched g_A are re
- Progress ongoing



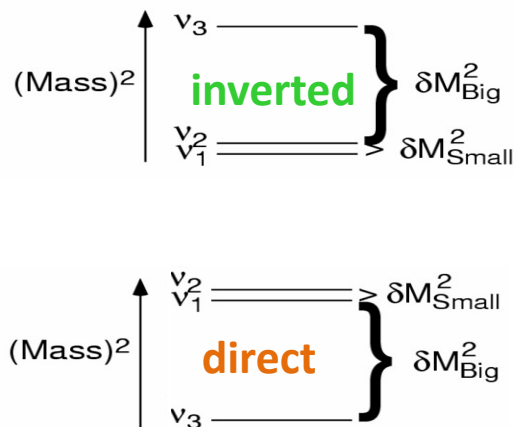
$$T_{1/2}^{0\nu} \simeq 10^{27-28} \left(\frac{0.01 \text{ eV}}{\langle M_{\beta\beta} \rangle} \right)^2 \text{ y}$$

Phys. Rev. C 85, 034316 (2012)

Credit Christoph Wiesinger – TAUP 2025

The effective Majorana mass

A reference plot



$\langle m_{\beta\beta} \rangle [\text{meV}]$

1000

100

10

1

$$\langle m_{\beta\beta} \rangle = \left| \left| U_{e1} \right|^2 m_1 + e^{i\alpha_1} \left| U_{e2} \right|^2 m_2 + e^{i\alpha_2} \left| U_{e3} \right|^2 m_3 \right|$$

IH ($\Delta m_{23}^2 < 0$)

NH ($\Delta m_{23}^2 > 0$)

Width of bands:

- Majorana phases
- Uncertainty on ν oscillations parameters

Phys. Rev. D90, 033005 (2014)

1

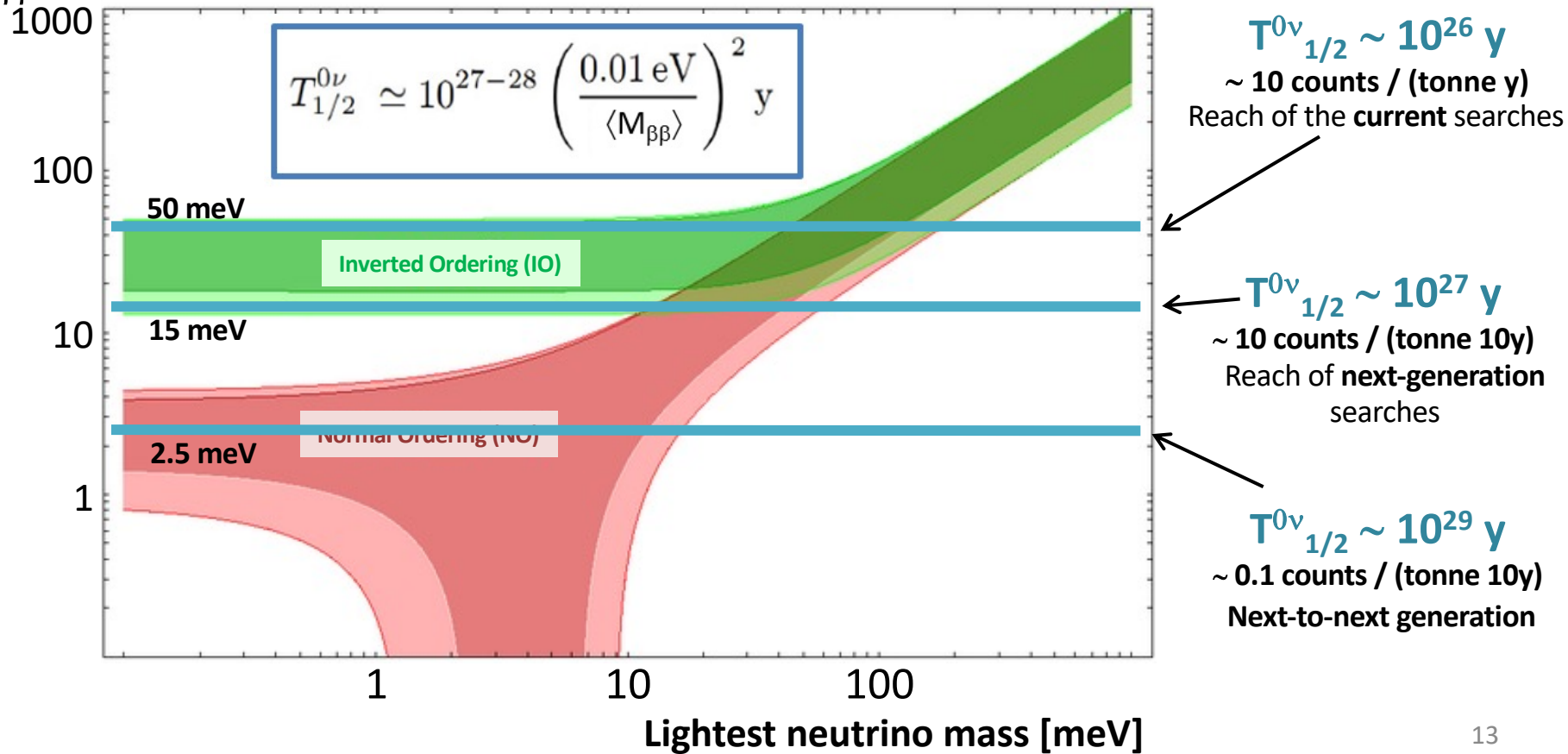
10

100

Lightest neutrino mass [meV]¹²

Experimental challenge

$\langle m_{\beta\beta} \rangle [\text{meV}]$



Next generation

$\langle m_{\beta\beta} \rangle [\text{meV}]$

1000

100

10

1

$$T_{1/2}^{0\nu} \simeq 10^{27-28} \left(\frac{0.01 \text{ eV}}{\langle M_{\beta\beta} \rangle} \right)^2 \text{ y}$$

Inverted Ordering (IO)

15 meV

Normal Ordering (NO)

$$T_{1/2}^{0\nu} \sim 10^{27} \text{ y}$$

$\sim 10 \text{ counts / (tonne } 10\text{y)}$
Reach of next-generation searches

Zero background/tonne-scale
are mandatory

1

10

100

Lightest neutrino mass [meV]

F: half-life sensitivity

Poisson limit

> 20 background counts

source mass live time energy resolution

$$F \propto (MT / b \Delta E)^{1/2}$$

background index

background counts @ $Q_{\beta\beta}$

$$M \times \Delta E \times T$$



Zero background

$$b \times M \times \Delta E \times T \ll 1$$

$$F \propto MT$$

Factors guiding isotope selection

➤ **High isotopic abundance (I.A.)**
and/or **easy enrichment**

➤ **High $Q_{\beta\beta}$**

➤ **Compatibility with
a beneficial detection
technique**

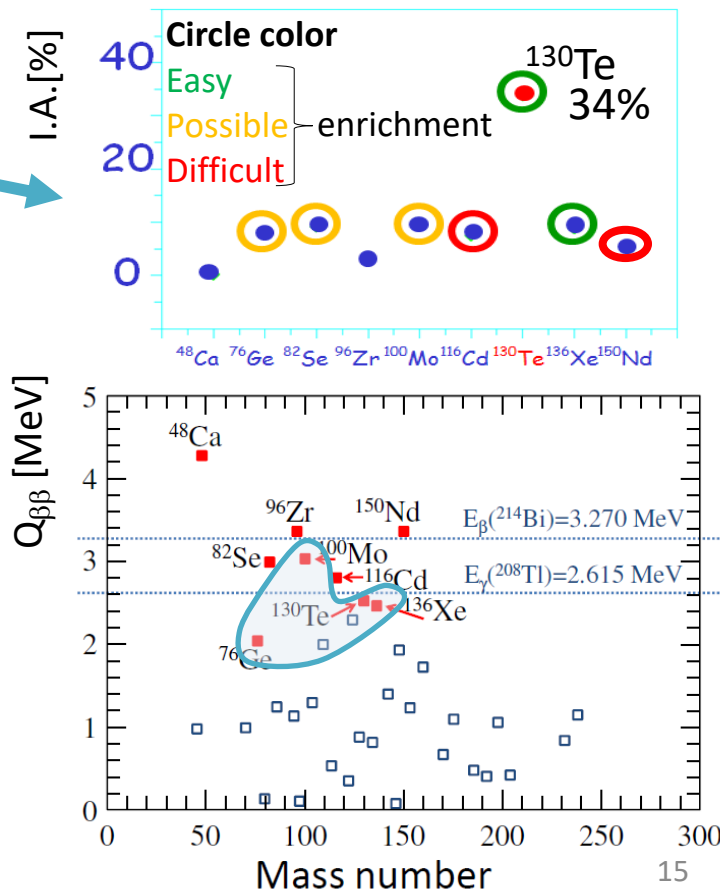
Larger phase space:
 $G(Q,Z) \propto Q^5$

Easier background
control

**High energy
resolution**

**Background
identification**

**Efficiency and
scalability**



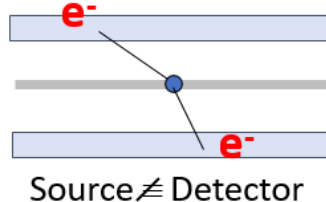
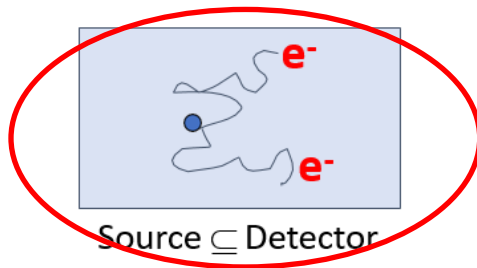
General features for $0\nu2\beta$ searches

Requests for the source

① **Large source** → tonne scale → $> 10^{27}$ nuclei

② **Maximize efficiency**

→ The option in which the source is separated from the detector is abandoned for next-generation experiments



However, this option may be interesting in case of discovery to investigate the mechanism of $0\nu2\beta$

Requests for the background

Generic measures as underground operation, shielding (passive and active), radiopurity of materials, vetos are common to $0\nu2\beta$ and other rare event search

Specific desirable features for $0\nu2\beta$

- High energy resolution
- Particle identification
- Tracking / Event topology
- Multi-site vs. single-site events
- Surface vs. bulk events
- Fiducial volume / Active shielding
- Final-state nucleus identification

Currently competing technologies (1)

①

Source dilution in a liquid scintillator

KamLAND-Zen (^{136}Xe) – SNO+ (^{130}Te)



- Re-use of existing infrastructures
- Large amount of isotopes (multi-ton)
- Isotope dilution (a few %)
- Energy resolution $\sim 10\%$ FWHM
- Rough space resolution

②

TPCs

EXO-200 – NEXT – nEXO (^{136}Xe)



- Large amount of isotopes (multi-ton)
- Full isotope concentration
- Energy resolution $\sim 1\% - 2\%$ FWHM
- Event topology

③

Semiconductor detectors

GERDA – LEGEND (^{76}Ge)



- Crystal array (~ 1 ton scale in total)
- (Almost) full isotope concentration
- Energy resolution $\sim 0.1\% - 0.2\%$ FWHM
- Particle identification
- Pulse shape discrimination

④

Bolometers

CUORE (^{130}Te) – AMoRE – CUPID (^{100}Mo)



Liquids and gases

Single crystals

Currently competing technologies (2)

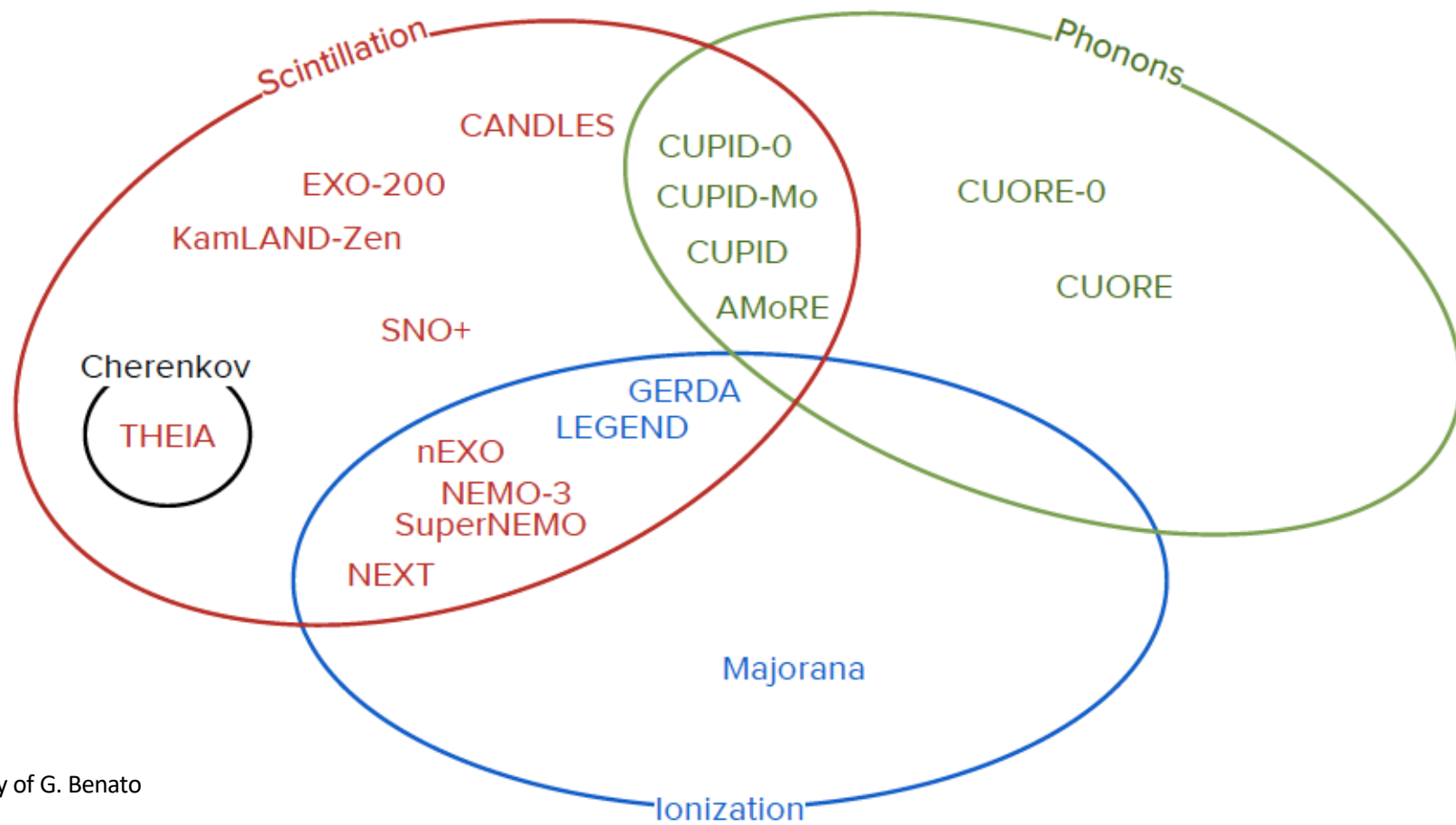


Readout channel	Energy resolution	Particle identification	Sensitivity to position	Applicable to multiple isotopes
Ionization	0.1-1%	Only in gas	Yes	Not really
Phonons	~0.2%	Nope	Nope	Yes
Scintillation	Few %	α vs β	In liquids and gases	Yes
Cherenkov	Forget it!	Visible only for β 's	Maybe	Yes

Take-away messages:

- The best detector technology does not exist
- The combination of readout channels is very welcome
 - Good resolution with one, background rejection with the other

Currently competing technologies (3)



Experimental status

KamLAND-Zen 800 - $T_{1/2} > 3.8 \times 10^{26}$ y

arXiv:2406.11438 (2024)

GERDA - $T_{1/2} > 1.8 \times 10^{26}$ y

Phys. Rev. Lett. 125, 252502 (2020)

EXO-200 - $T_{1/2} > 3.5 \times 10^{25}$ y

Phys. Rev. Lett. 123, 161802 (2019)

MAJORANA dem. - $T_{1/2} > 8.3 \times 10^{25}$ y

Phys. Rev. Lett. 130, 062501 (2023)

CUORE - $T_{1/2} > 3.5 \times 10^{25}$ y

TAUP 2025 (2025)

CUPID-0 - $T_{1/2} > 4.6 \times 10^{24}$ y

Phys. Rev. Lett. 129, 111801 (2022)

AMORE-I - $T_{1/2} > 2.9 \times 10^{24}$ y

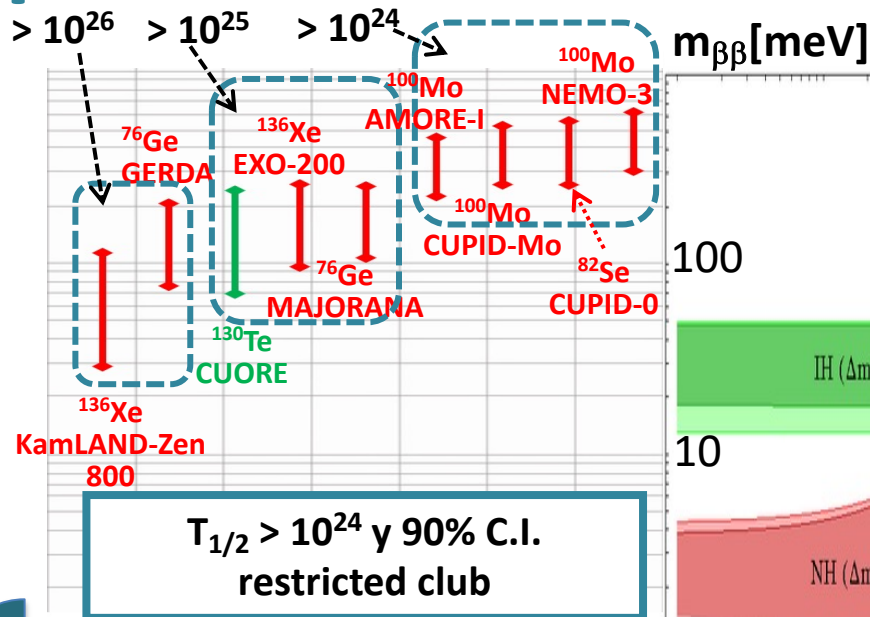
Phys. Rev. Lett. 134, 082501 (2025)

CUPID-Mo - $T_{1/2} > 1.8 \times 10^{24}$ y

Eur. Phys. J. C 82, 1033 (2022)

NEMO-3 - $T_{1/2} > 1.1 \times 10^{24}$ y

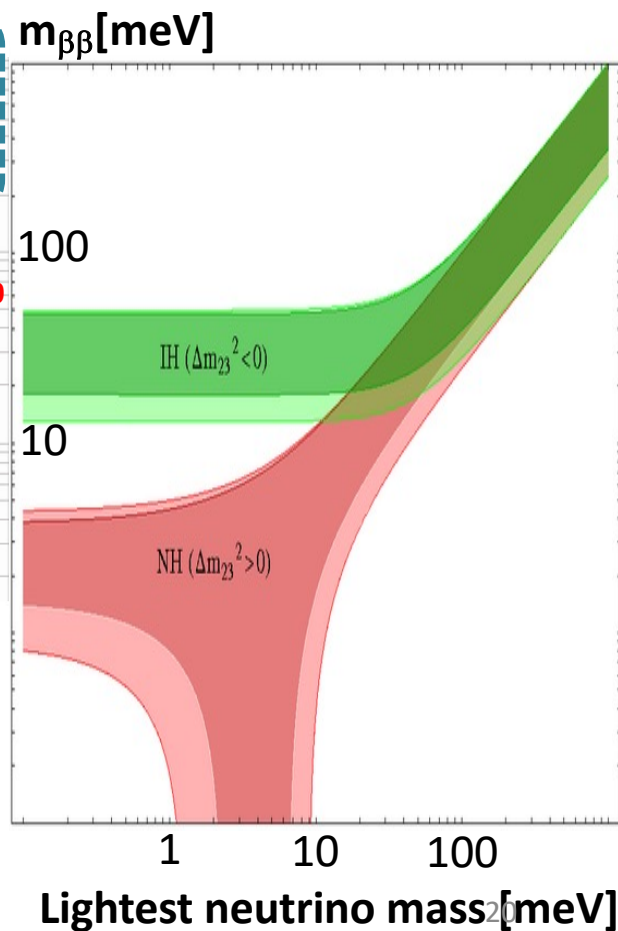
Phys. Rev. D 92, 072011 (2015)



All these experiments stopped
except **CUORE**
(expected end: mid-2026)

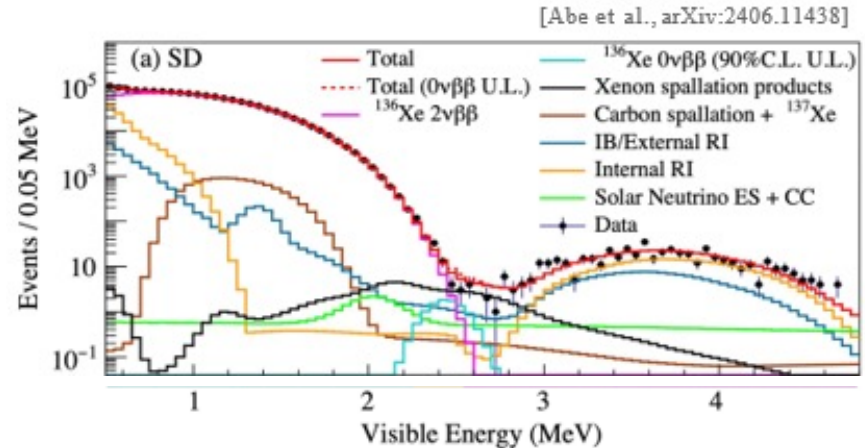
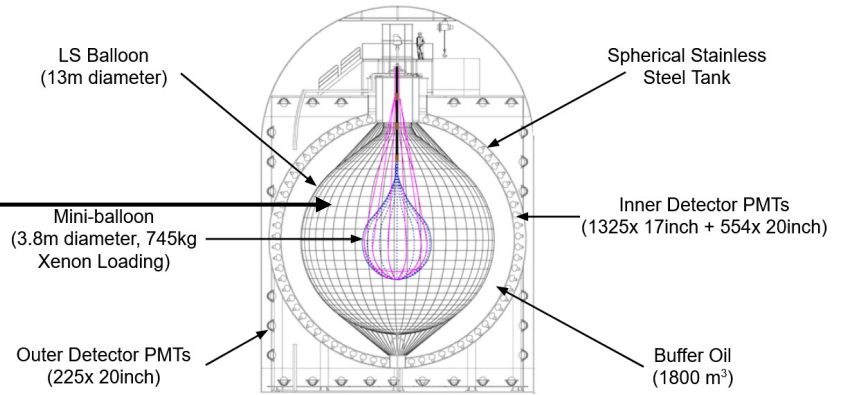
Waiting for next generation!

LEGEND200 already in data taking
[1.9×10^{26} y]



KamLAND-Zen

- **1000 ton liquid scintillator detector**
- Re-use facility used for ν oscillation experiment
- **Nylon balloon** with $^{\text{enr}}\text{Xe}$ -loaded scintillator
- **KamLAND-Zen-800**, 2nd phase
 - **Large isotope mass**, 750 kg of 91% ^{136}Xe
 - **Poor energy resolution**, 4% at 2.5 MeV
 - ➔ Final result, **world-best constraint**
 - $T_{1/2}(^{136}\text{Xe}) > 3.8 \cdot 10^{26} \text{ yr}$ (90% CL)
 - ➔ $m_{\beta\beta} < [28, 122] \text{ meV}$ (90% CL)
 - $> 2.6 \cdot 10^{26} \text{ yr}$ (90% CL, sensitivity)



- **KamLAND2-Zen**, planned detector upgrade ➔ After 5 years: $T_{1/2}^{0\nu} > 2 \times 10^{27} \text{ yr}$

SNO+

- **780 ton liquid scintillator detector**
Re-use facility used for solar ν experiment
- Staged $^{\text{nat}}\text{Te}$ loading

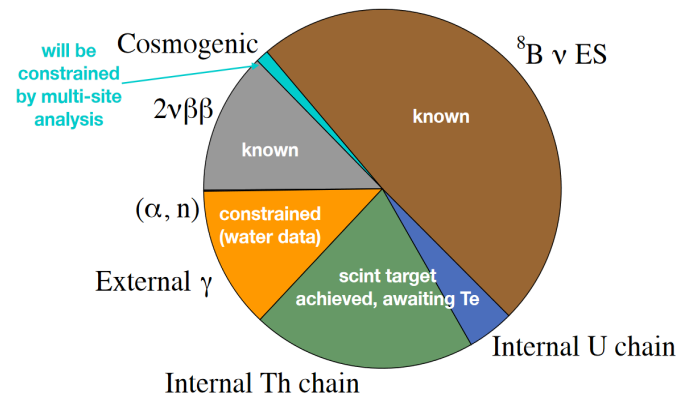
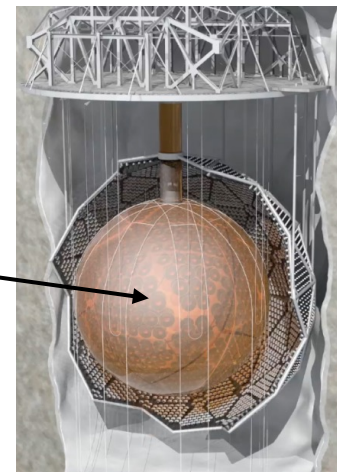
- **Large isotope mass,**
e.g. 1.3 ton of ^{130}Te for 0.5% loading
- **Poor energy resolution, 4% at 2.5 MeV**

- Water phase completed, scintillator phase ongoing, **0.5% loading in preparation**

➔ Projected sensitivity

$$T_{1/2}(^{130}\text{Te}) > 2.0 \cdot 10^{26} \text{ yr (90\% CL), 3 yr with 0.5\%}$$

$$T_{1/2}(^{130}\text{Te}) > 7.4 \cdot 10^{26} \text{ yr (90\% CL), 5 yr with 1.5\%}$$



Background budget
(low concentration → dominated by ^8B solar ν)

nEXO

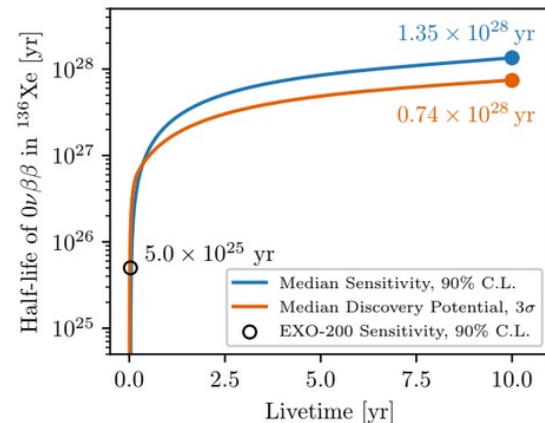
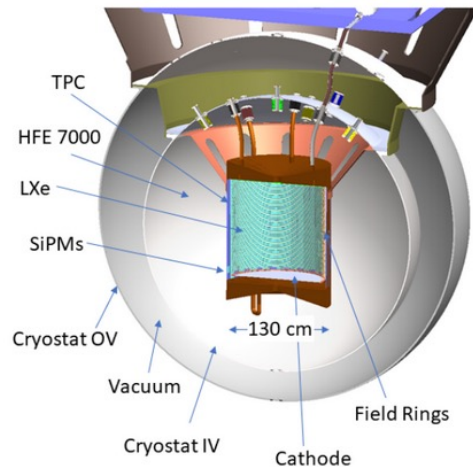
- **Liquid ^{enr}Xe time projection chamber,**
 - building on EXO-200
 - **Large isotope mass, 5 ton of 90% ¹³⁶Xe**
 - **Reasonable energy resolution, 1% at 2.5 MeV**

➔ Projected sensitivity, 10yr
 [Adhikari et al., J.Phys.G 49 (2022) 1, 015104]

$$T_{1/2}({}^{136}\text{Xe}) > 1.35 \cdot 10^{28} \text{ yr (90\% CL)}$$

$$T_{1/2}({}^{136}\text{Xe}) = 7.4 \cdot 10^{27} \text{ yr (3}\sigma) \quad \text{Discovery sensitivity}$$

- **Development of Ba tagging**
 [Chambers et al., Nature 569 (2019) 7755, 203-207;
 Ray et al., Atoms 12 (2024) 12, 71]



Background dominated by Rn outgassing and intrinsic radioactivity

NEXT

- **High pressure (10-15 bar) gas ^{enr}Xe TPC**
 - Primary scintillation ($t_0 \rightarrow z$ coordinate)
 - Electroluminescence for energy resolution and for tracking

Staged approach

- **NEXT-White (2015-2021) – LSC, Spain**
5 kg prototype

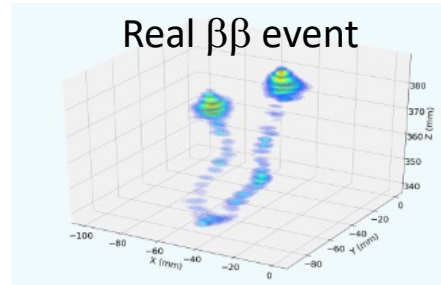
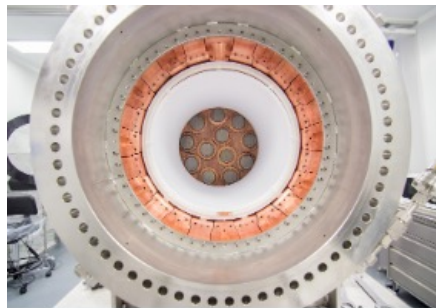
- **NEXT-100 (2023-2029) – LSC, Spain**

Upscaling of NEXT-White \rightarrow 97 kg

- Currently, operation at 4 bar
- Operation at 10 bar by 2026
- Good energy resolution $<1\%$

Projected sensitivity

$T_{1/2}(^{136}\text{Xe}) > 4 \cdot 10^{25} \text{ yr}$ (90% CL) – 3 year
[JHEP 2016, 159 (2016)]



- **NEXT-HD (High Definition) – start in 2030**

Up to 1 ton enriched Xe gas at 20 bar

Target sensitivity: 10^{27} y

5 ton \times y

- **NEXT-BOLD (Barium On Light Detection)**

- Ba tagging by **SMFI** (Single Molecule Fluorescence Imaging) was proved – demonstrator under R&D
- Background free [JINST 18 P08006 (2023)]

Target sensitivity: 10^{28} y

10 ton \times y

LEGEND

A. Leder at this conference

➤ High-purity ^{enr}Ge detectors in active liquid argon shield, building on:

- GERDA, **lowest background**, background-free [Agostini et al., PRL 125 (2020) 25, 252502]
- MAJORANA, **best energy resolution**, 2.5 keV (FWHM) at 2.0 MeV [Arnquist et al., PRL 130 (2023) 6, 062501]

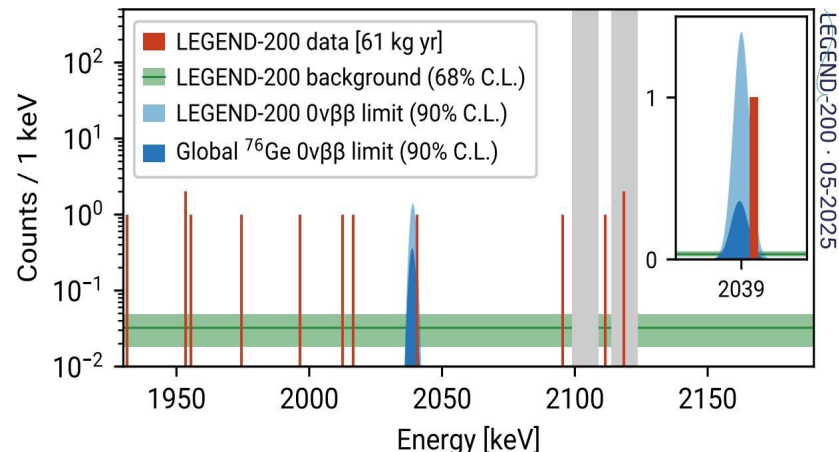
➤ **LEGEND-200**, up to 200 kg of 90% ⁷⁶Ge in

- Upgraded GERDA infrastructure, improved light read-out
- Data taking

➔ Projected discovery sensitivity, 5 yr

[Agostini et al.,
Rev.Mod.Phys. 95 (2023) 2, 025002]

$$T_{1/2}(^{76}\text{Ge}) = 1.5 \cdot 10^{27} \text{ yr } (3\sigma)$$



➤ **LEGEND-1000**, new infrastructure in preparation at LNGS

- **Sizeable isotope mass**, 1 ton of 90% ⁷⁶Ge

➔ Projected discovery sensitivity, 10 yr

➤ [Abgrall et al., arXiv:2107.11462]

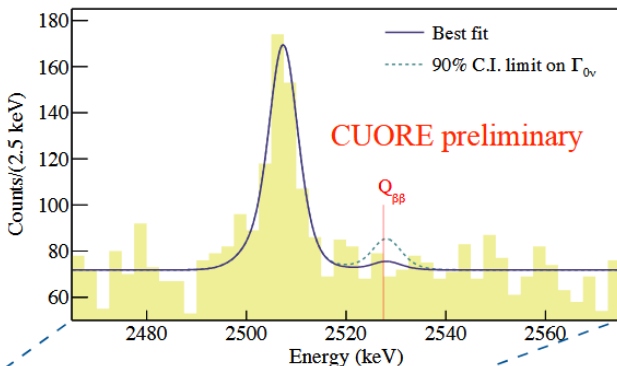
$$T_{1/2}(^{76}\text{Ge}) > 1.6 \cdot 10^{28} \text{ yr } (90\% \text{ CL})$$

$$T_{1/2}(^{76}\text{Ge}) = 1.3 \cdot 10^{28} \text{ yr } (3\sigma)$$

CUORE - LNGS, Italy

- Array of natural TeO_2 crystals operated as bolometers at 10 mK
 - Built on the precursor CUORICINO experiment
 - 988 TeO_2 crystals, arranged in 19 towers – 206 kg of ^{130}Te
 - Energy resolution $\Delta E \sim 7.8$ keV FWHM @ $Q_{\beta\beta}$

Results: $T_{1/2} > 3.5 \times 10^{25}$ y (90%) – 2039 kg y
[TAUP 2025]

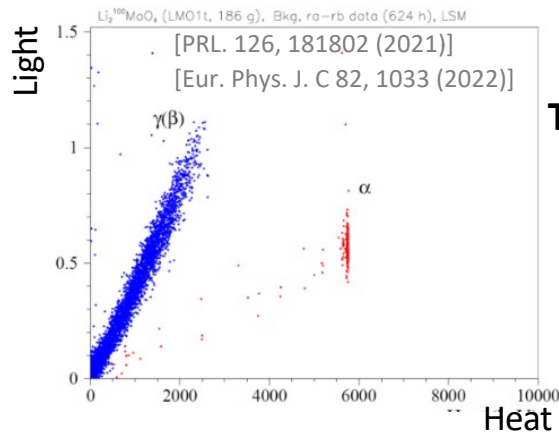


Background dominated by energy-degraded surface α 's

CUPID demonstrators

CUPID-Mo - LSM, France

- Pure bolometers → Scintillating bolometers
- $^{130}\text{Te} (\text{TeO}_2) \rightarrow ^{100}\text{Mo} (\text{enriched Li}_2\text{MoO}_4)$
 $Q_{\beta\beta} = 3034$ keV > 2.6 MeV (Reject external γ background)
- 20 Li_2Mo_4 crystals – 2.26 kg of ^{100}Mo
- Energy resolution $\Delta E \sim 7.7$ keV FWHM @ $Q_{\beta\beta}$



Results:
 $T_{1/2} > 1.8 \times 10^{24}$ y (90%)
2.71 kg y

Reject
 α background

CUPID-0 - LNGS, Italy Zn^{82}Se
First scintillating
bolometer demonstrator

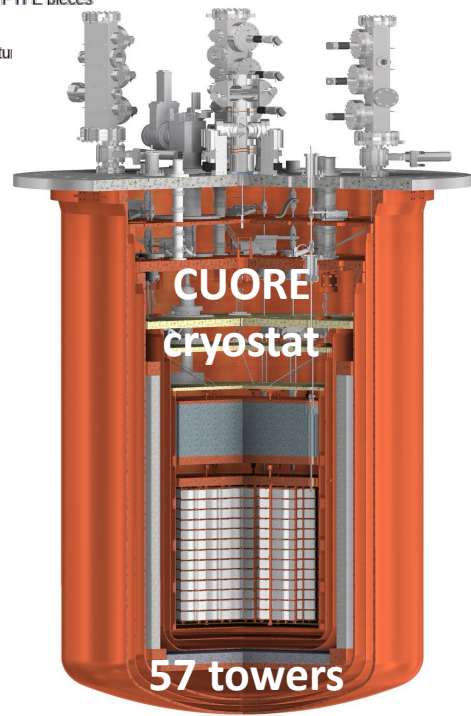
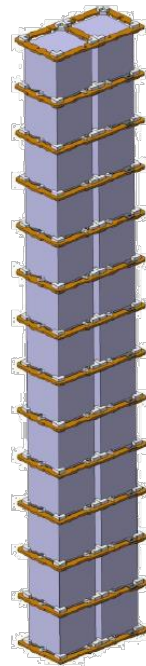
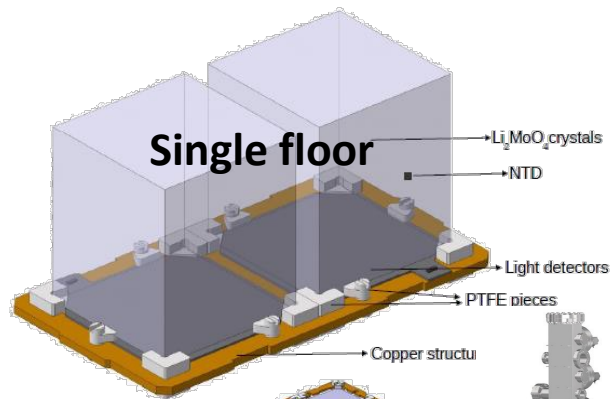
[PRL, 129, 111801 (2022)]

$T_{1/2} > 4.6 \times 10^{24}$ y

CUPID

^{enr}Mo-containing scintillating bolometers

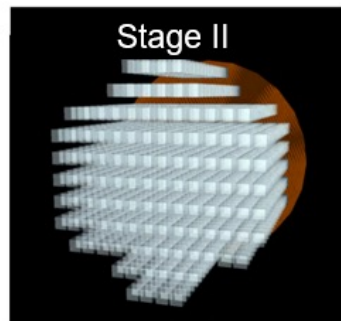
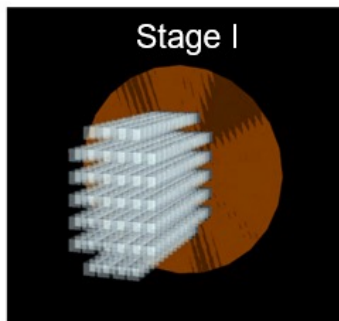
- Heat readout based on **NTD Ge sensors**
- Energy resolution $\Delta E \sim 5 \text{ keV FWHM @ } Q_{\beta\beta}$
- Exploit CUORE infrastructure changing isotope
- Adopt CUPID-Mo technology
- Single module: $\text{Li}_2^{100}\text{MoO}_4$ **45×45×45 mm – 280 g**
- **1596 crystals**
- **240 kg of ^{100}Mo** with >95% enrichment
- **Bolometric Ge light detectors** as in CUPID-Mo, CUPID-0 upgraded with **Neganov-Trofimov-Luke mode** to achieve the required rejection of $2\nu 2\beta$ random coincidences [**BINGO/CROSS development**]



CUPID – Timeline and sensitivity

➤ Staged development

- Stage-I → 1/3 of crystals (2030-2033)
- Stage-II → remaining 2/3 of crystals (2034)



CUPID Stage-I

➤ 3 σ - discovery sensitivity

$$T_{1/2}^{0\nu} = 0.2 \times 10^{27} \text{ yr}$$

$$m_{\beta\beta} < (26-44) \text{ meV}$$

FULL CUPID

➤ Exclusion sensitivity 90% C.I.

(10 y livetime - 240 kg ^{100}Mo + 5 keV FWHM)

$$T_{1/2}^{0\nu} > 1.4 \times 10^{27} \text{ yr}$$

$$m_{\beta\beta} < (10-17) \text{ meV}$$

➤ 3 σ - discovery sensitivity

$$T_{1/2}^{0\nu} = 1 \times 10^{27} \text{ yr}$$

$$m_{\beta\beta} < (12-21) \text{ meV}$$

R&D ongoing to reduce the BI by a further order of magnitude (BINGO and others), making possible

➤ CUPID-1T

10 y livetime - 1000 kg ^{100}Mo + 5 keV FWHM

$$T_{1/2}^{0\nu} > 9.2 \times 10^{27} \text{ yr}$$

$$m_{\beta\beta} < (4.0-6.9) \text{ meV}$$

Deep inside the normal
ordering region

Beyond CUORE and CUPID: BINGO



Techniques for background rejection in future Li_2MoO_4 / TeO_2 based experiments 3 main innovations

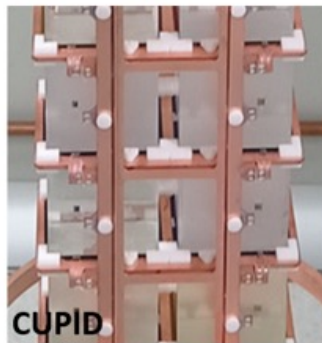
- **Internal active shield** (ultrapure BGO scintillators)
 - mitigate γ background in TeO_2
- **Revolutionary assembly** to reject surface background
- **Enhanced-sensitivity (Neganov-Trofimov-Luke)** light detectors
 - mitigate pile-up
 - read out the BGO active shields
 - detect Cherenkov light in TeO_2

MINI-BINGO in Modane (demonstrator)

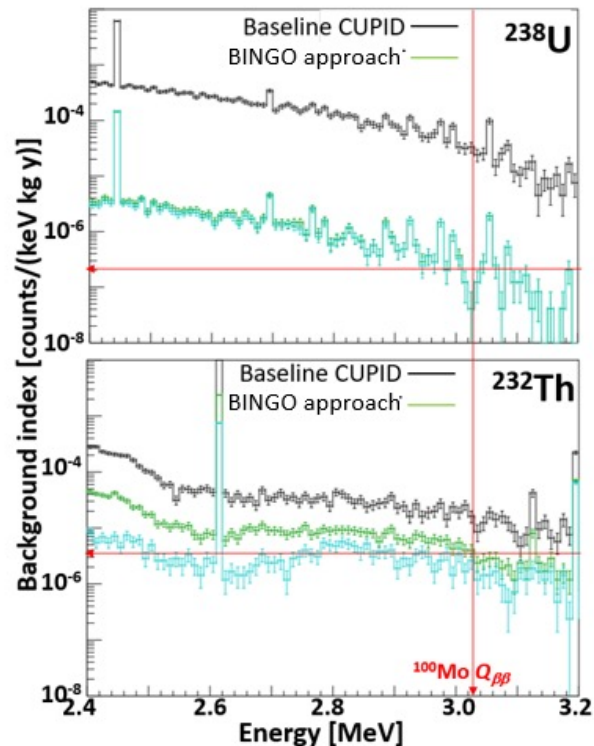
12 Li_2MoO_4 crystals

12 TeO_2 crystals

from 2026



Simulation CUPID vs. BINGO



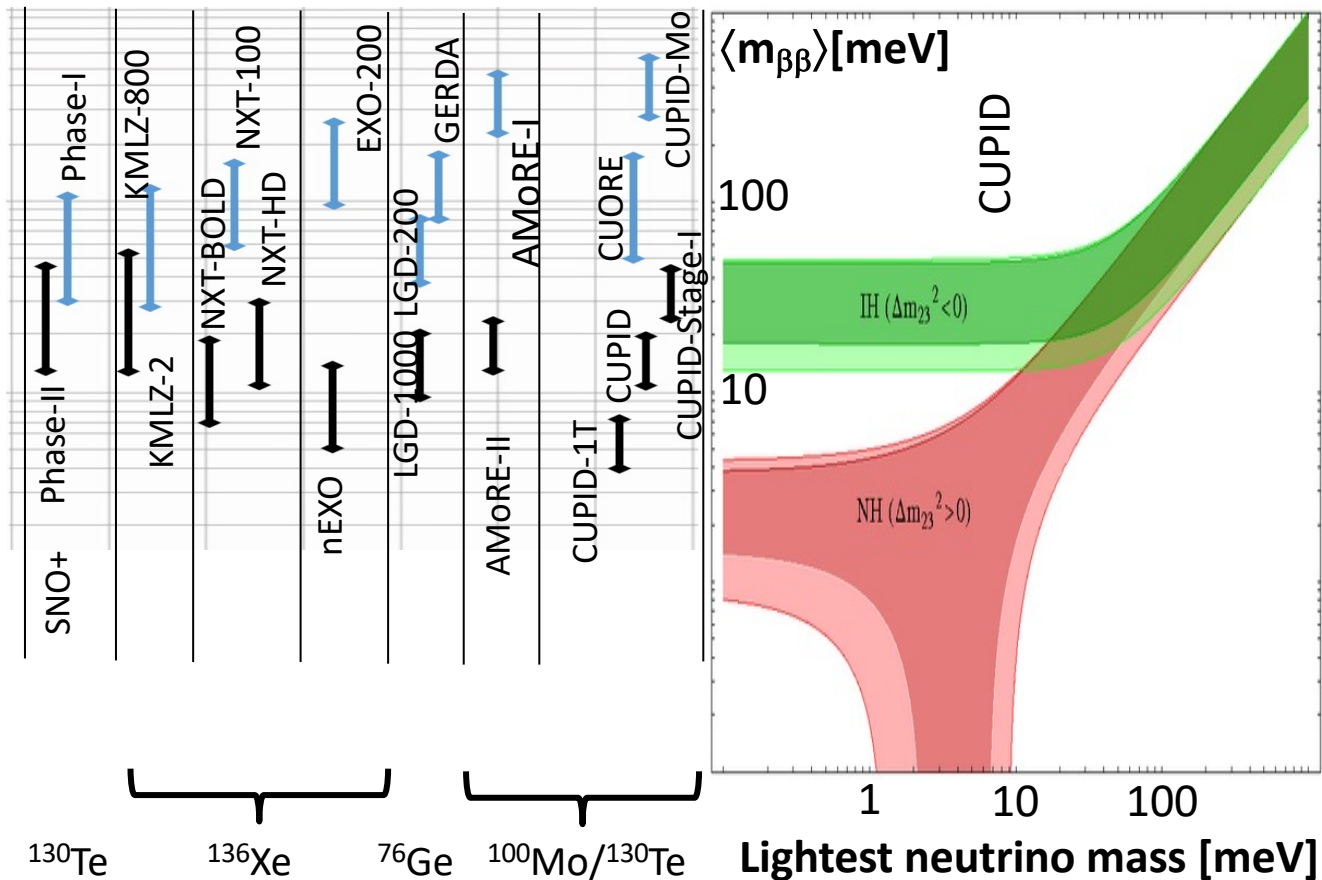
How we are moving forward

Current generation

(final sensitivity for concluded/running projects)

Next generation

(projects to be started on 5-10 years' time scale)



Conclusions

- $0\nu\beta\beta$ is a crucial process for particle physics and cosmology
- Several approaches and technologies make this field very active
- Many projects will extend their sensitivity in the next years
- Next-generation experiments have a good discovery potential

Neutrinoless double beta decay is a high-risk, high-reward pursuit

A hard search, but a discovery would reshape physics:
the Standard Model, mass, and the matter–antimatter asymmetry.



What lies ahead

From the current to the next generation

Emphasis on **7 promising research lines**

- | | | | |
|---|--|---|--|
| ① | KamLAND-Zen 400 → KamLAND-Zen 800 → KamLAND2-Zen | ^{136}Xe | Source dilution in a liquid scintillator |
| ② | SNO+ → SNO+-phase II | ^{130}Te | |
| ③ | EXO-200 → nEXO | ^{136}Xe | Xe TPCs |
| ④ | NEXT-White → NEXT-100 → NEXT-HD / NEXT-BOLD | | |
| ⑤ | $\left\{ \begin{array}{l} \text{GERDA} \\ \text{MAJORANA dem.} \end{array} \right\} \rightarrow \text{LEGEND-200} \rightarrow \text{LEGEND-1000}$ | ^{76}Ge | Semiconductor detectors – Ge diodes |
| ⑥ | AMORE-I → AMORE-II | ^{100}Mo | Bolometers |
| ⑦ | $\left\{ \begin{array}{l} \text{CUORE} \\ \text{CUPID-Mo} \\ \text{CUPID-0} \end{array} \right\} \rightarrow \text{CUPID} \rightarrow \text{CUPID 1t}$ | ^{130}Te
^{100}Mo
^{82}Se
^{100}Mo | |

Neutrino mass scale and effective neutrino mass

Cosmology, **single** and **double β decay** measure different combinations of the neutrino mass eigenvalues, constraining the neutrino mass scale

$$\Sigma \equiv \sum_{i=1}^3 m_i$$

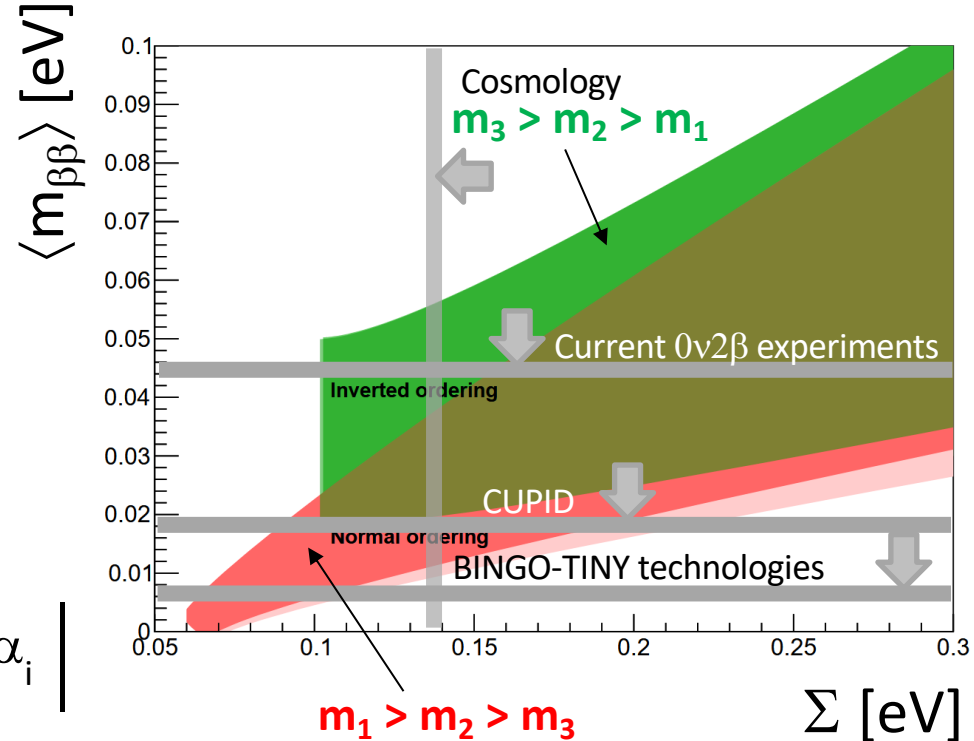
Cosmology- simple sum - pure kinematical effect

$$\langle m_{\beta} \rangle \equiv \left(\sum_{i=1}^3 m_i^2 |U_{ei}|^2 \right)^{1/2}$$

β Decay – KATRIN - incoherent sum - **real** neutrino

$$\langle m_{\beta\beta} \rangle \equiv \left| \sum_{i=1}^3 m_i |U_{ei}|^2 e^{i\alpha_i} \right|$$

double β decay - coherent sum - **virtual** neutrino - Majorana phases



Many theoretical models for the nuclear matrix elements



WE WANT YOU

$$1/\tau = G(Q,Z) g_A^4 |M_{\text{nuc}}|^2 \langle m_{\beta\beta} \rangle^2$$

The $0\nu\beta\beta$ community in general assumes $g_A \approx 1.27$ (no quenching)

- Different **phenomenological many-body methods** using different approximations (e.g. limited number of nuclear shells), **significant spread**

[Agostini et al., Rev.Mod.Phys. 95 (2023) 2, 025002; ..]

- Experiments provide **range of $m_{\beta\beta}$ constraints**

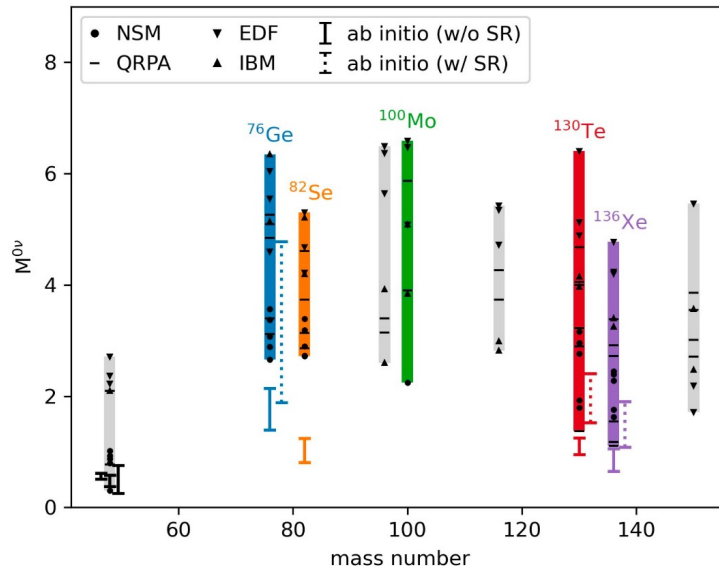
$$T_{1/2} = \dots \rightarrow m_{\beta\beta} = [\dots, \dots]$$

- First **ab initio calculations available**, may resolve quenching issue

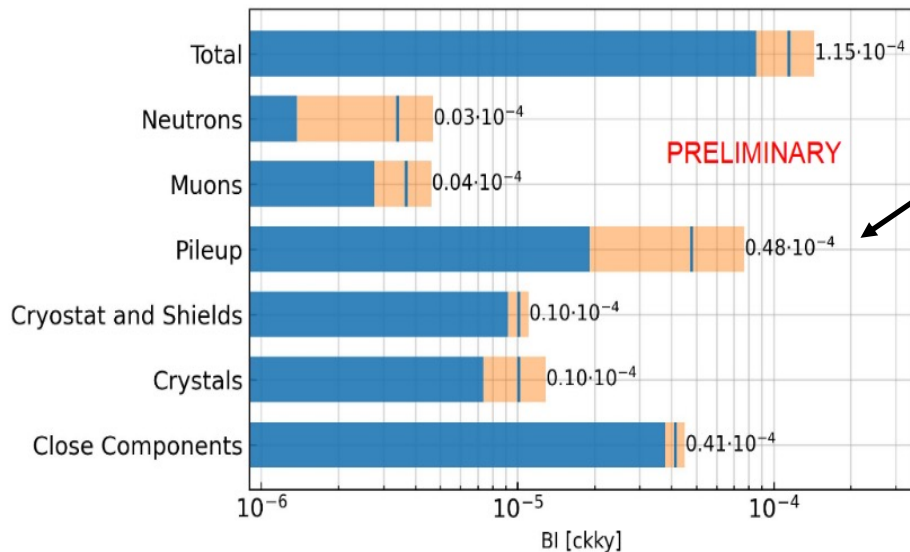
[Yao et al., PRL 124 (2020); Belley et al., PRL 126 (2021); Novario et al., PRL 126 (2021); Cirigliano et al., PRL 120 (2018); Belley et al., arXiv:2307.15156; Belley et al., PRL 132 (2024); ..]

- Effective field theory (EFT) analysis identified additional **short-range contribution**

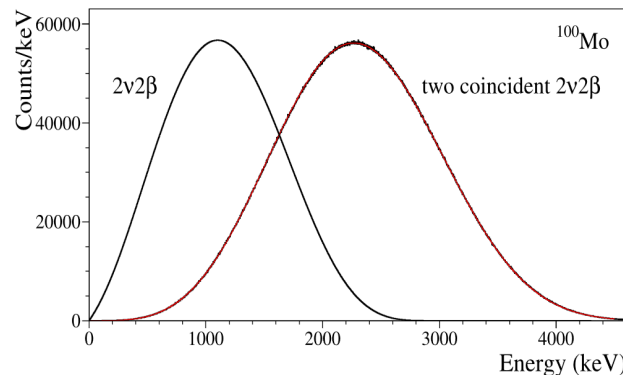
[Cirigliano et al., PRL 120 (2018) 20, 202001; ..]



CUPID - background



Dominant contributions: **pile-up events**
(random coincidences of ordinary $2\nu\beta\beta$ events)



➤ **Data driven:**
based on CUORE and CUPID-Mo background models
[Phys. Rev. D 110 (2024) 052003] [Eur. Phys. J. C 83 (2023) 675]

Conservative projections based on
current available data: **BI = 1.1×10^{-4} ccky**

- **Fastest $2\nu 2\beta$ decay: $T_{1/2} \sim 7.1 \times 10^{18} \text{ y}$**
- **Slow response of bolometric signals**
Mitigated by Neganov-Trofimov-Luke light detectors

**Two orders of magnitude
better than in CUORE**

AMoRE

enrMo-containing scintillating bolometers

- Heat readout based on **MMC sensors**
- Energy resolution $\Delta E \sim 10\text{-}15 \text{ keV FWHM @ } Q_{\beta\beta}$

➤ AMoRE-I (2020-2022)

- $13 \times {}^{48\text{depl}}\text{Ca}^{100}\text{MoO}_4$ (4.6 kg)
- $5 \times \text{Li}_2^{100}\text{MoO}_4$ (1.6 kg)

Results: $T_{1/2} > 2.9 \times 10^{24} \text{ y}$

[Phys. Rev. Lett. 134, 082501 (2025)]

➤ AMORE-II – from 2026 – phased approach

- Secured **110 kg of ${}^{100}\text{Mo}$** – $596 \times \text{Li}_2^{100}\text{MoO}_4$ crystals
- New cryostat and underground lab – work in progress

Projected sensitivity: $8 \times 10^{26} \text{ y}$ in 5 years

