

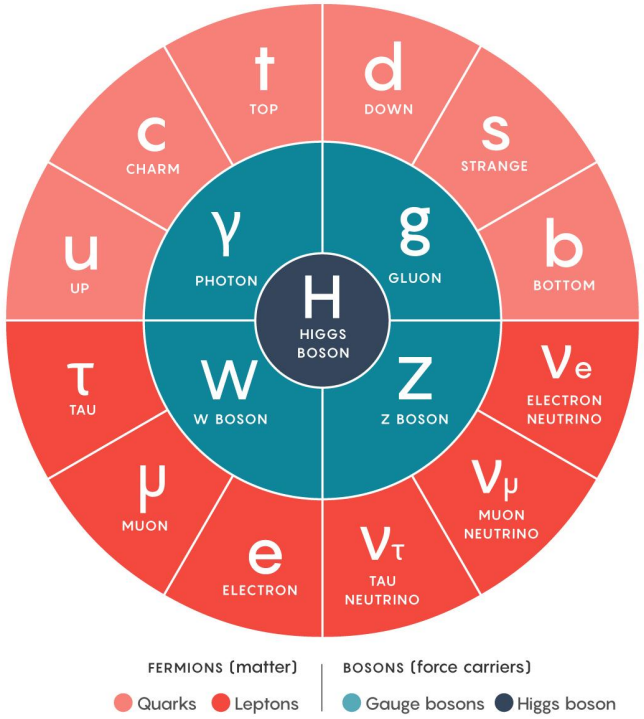
Event Reconstruction and Analysis in Water Cherenkov Detectors with Machine Learning

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What are Neutrinos ?

The Standard Model



- Neutrinos are mysterious

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \text{PMNS} \\ \text{matrix} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

“Flavour” (interaction) state Mass eigenstate

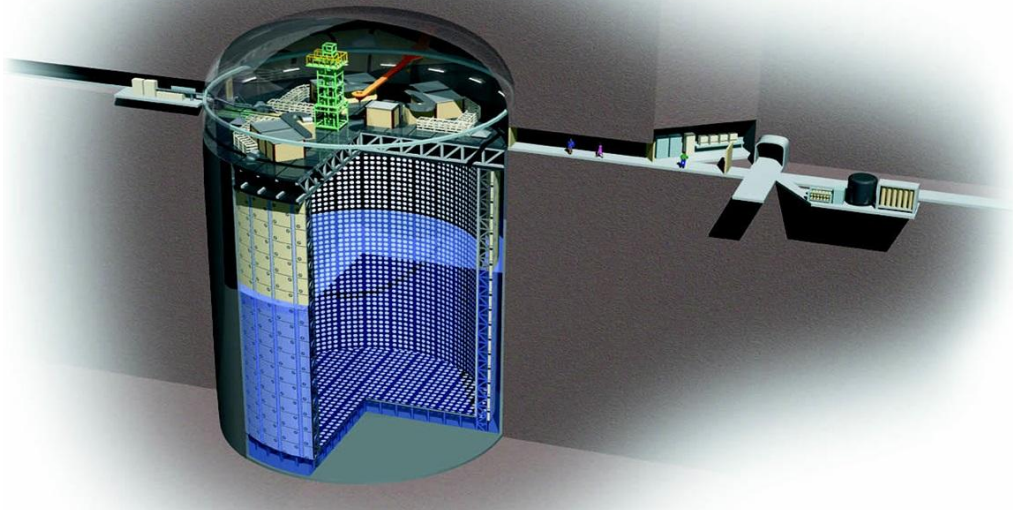
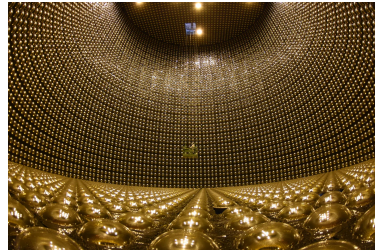
- Chargeless lepton : only interacts via the weak interaction : hard to observe
- Impressively low mass compared to other particles : less than 2eV for neutrinos vs 0.511MeV for the electron

Why is neutrino oscillation important to study ?

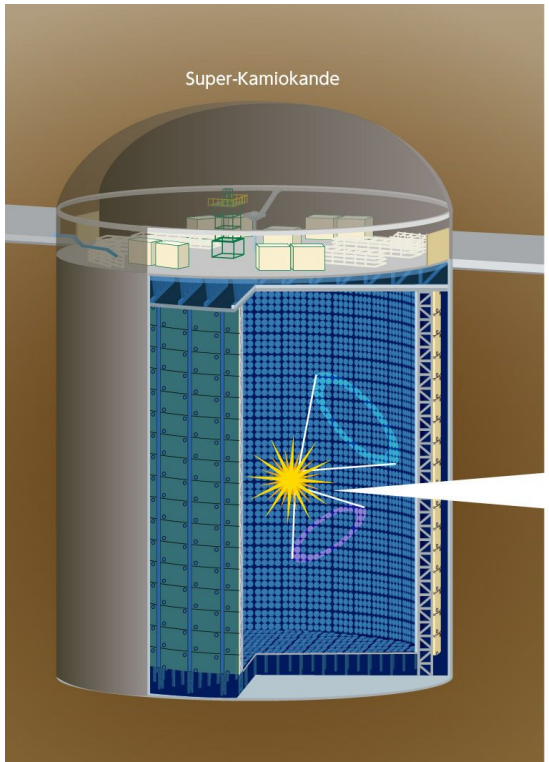
$$U_{\text{PMNS}} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix}}_{\text{atmospheric } \Delta m_{32}^2} \underbrace{\begin{pmatrix} \cos \theta_{13} & 0 & e^{-i\delta} \sin \theta_{13} \\ 0 & 1 & 0 \\ -e^{-i\delta} \sin \theta_{13} & 0 & \cos \theta_{13} \end{pmatrix}}_{\text{accelerators}} \underbrace{\begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{solar } \Delta m_{21}^2} \underbrace{\hspace{10em}}_{\text{reactors}}$$

- CP violation for leptons: Matter antimatter asymmetry
- Mass ordering is still unknown
- Neutrinos and neutrino experiments are important inside and outside particle physics too : proton decay, cosmology (CMB, Dark matter searches), astrophysics (SuperNovae)

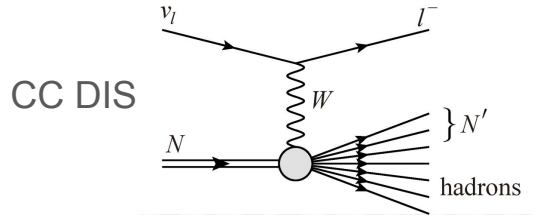
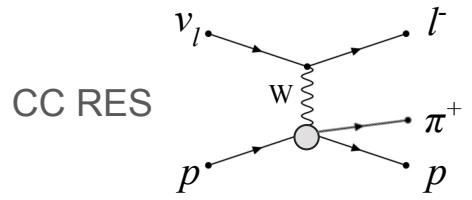
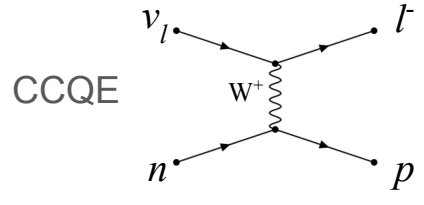
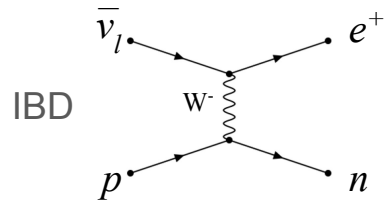
Neutrino experiments in Japan



A focus on Water Cherenkov detectors



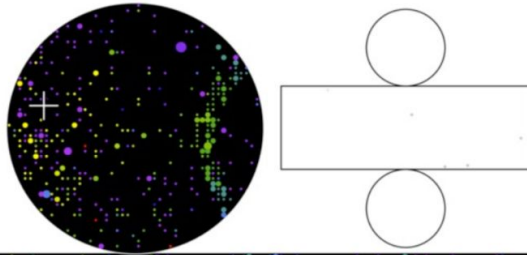
- Uses light to indirectly detect particles, via the Cherenkov effect
- The neutrinos will interact via different processes :



How do we understand what we see in the detector ?

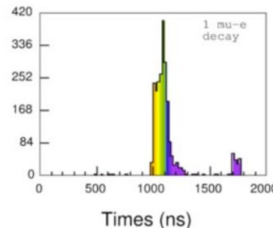
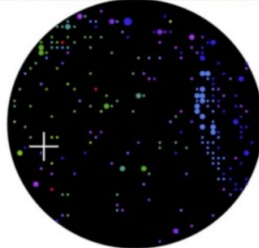
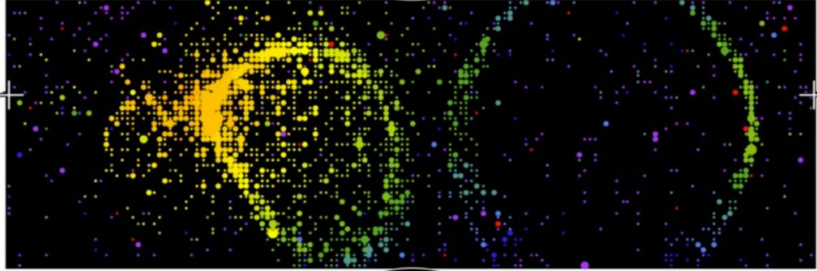
Super-Kamiokande IV

Run 999999 Sub 0 Event 897
14-02-16:01:13:20
Inner: 2370 hits, 5176 pe
Outer: 0 hits, 0 pe
Trigger: 0x07
D_wall: 453.0 cm
Evis: 513.4 MeV
2 e-like rings: mass = 341.8 MeV/c²



Time (ns)

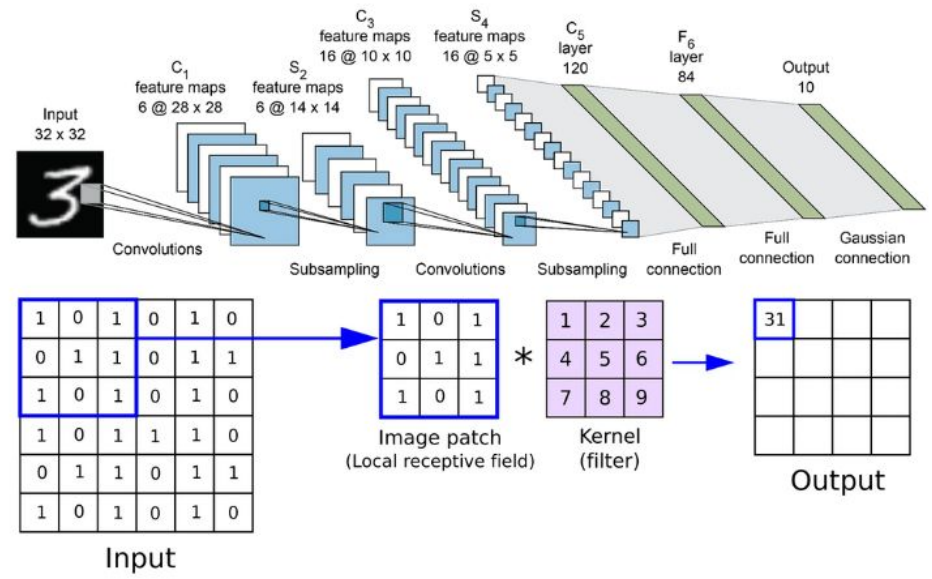
- < 974
- 974-987
- 987-1000
- 1000-1013
- 1013-1026
- 1026-1039
- 1039-1052
- 1052-1065
- 1065-1078
- 1078-1091
- 1091-1104
- 1104-1117
- 1117-1130
- 1130-1143
- 1143-1156
- >1156



- The detector acts like a giant camera only giving us charge and time information
- **Reconstruction methods** have been developed to recover the physics variables:
 - What particles ? Electron, photon, muon, pion...
 - What energy ? Position of origin in the detector ? Impact angle ?
- They are traditionally based on maximum likelihood estimation, based on what we know about the underlying processes. (FitQun)

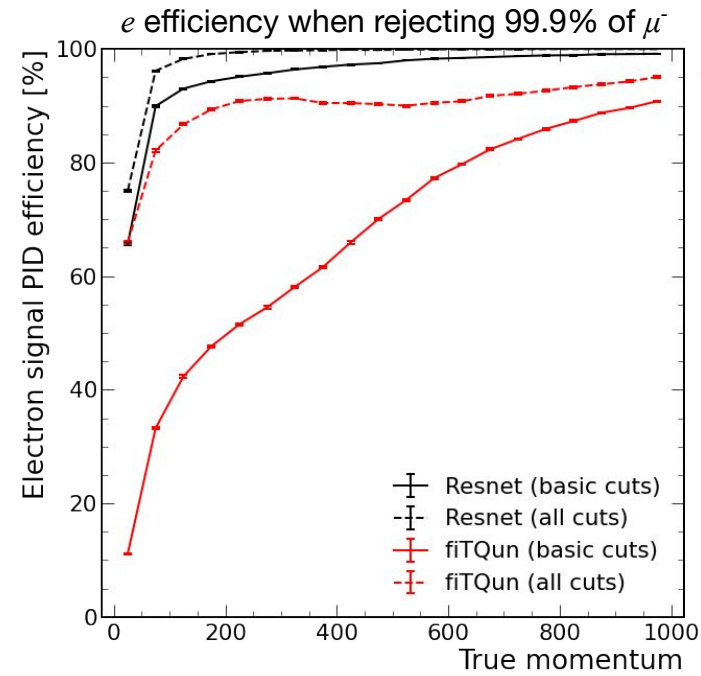
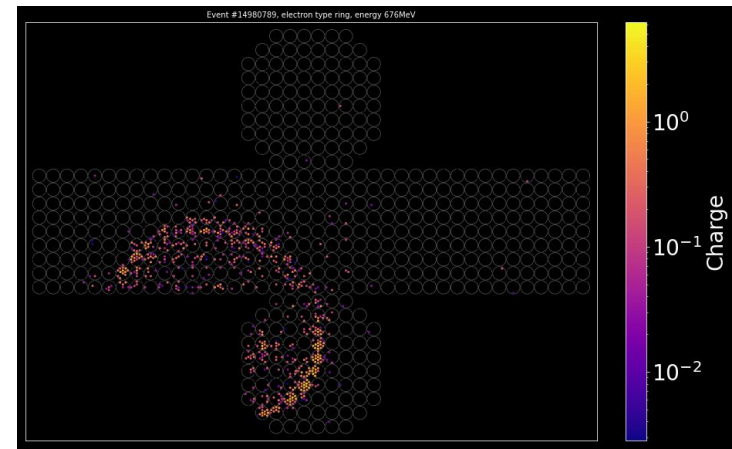
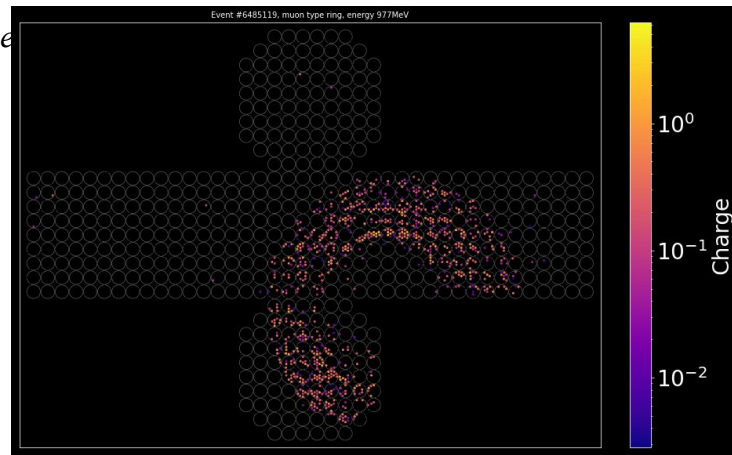
Machine Learning based reconstruction

- ML has seen an extreme rise of popularity and research effort in the past 5-6 years : some algorithms from Machine and Deep Learning can be adapted to neutrino detectors.



- + Much better performances than traditional methods
- + No physics knowledge required
- With current architectures, hard to estimate the uncertainty and the bias in the predictions
- Low interpretability

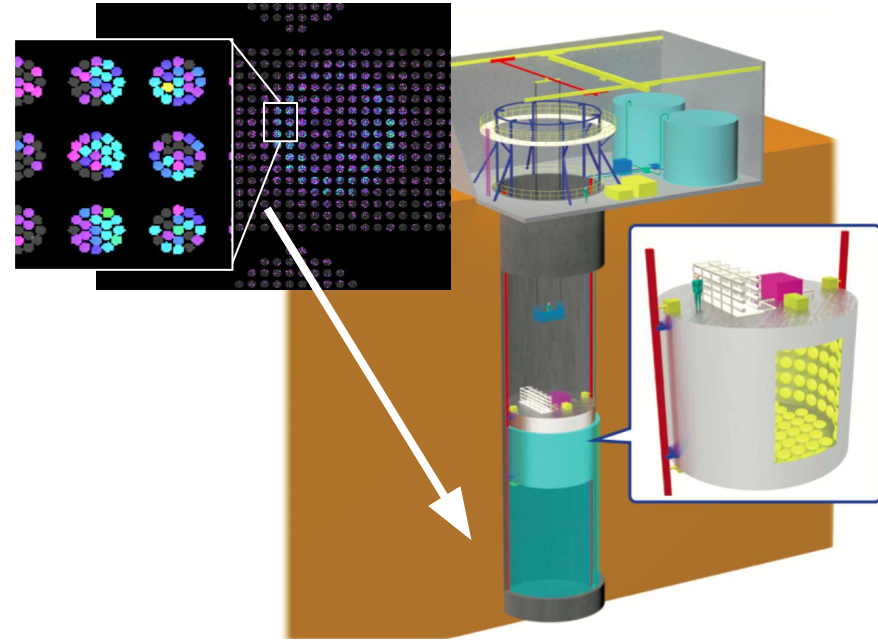
Current results



More than 99.9% efficiency needed for the physics analysis

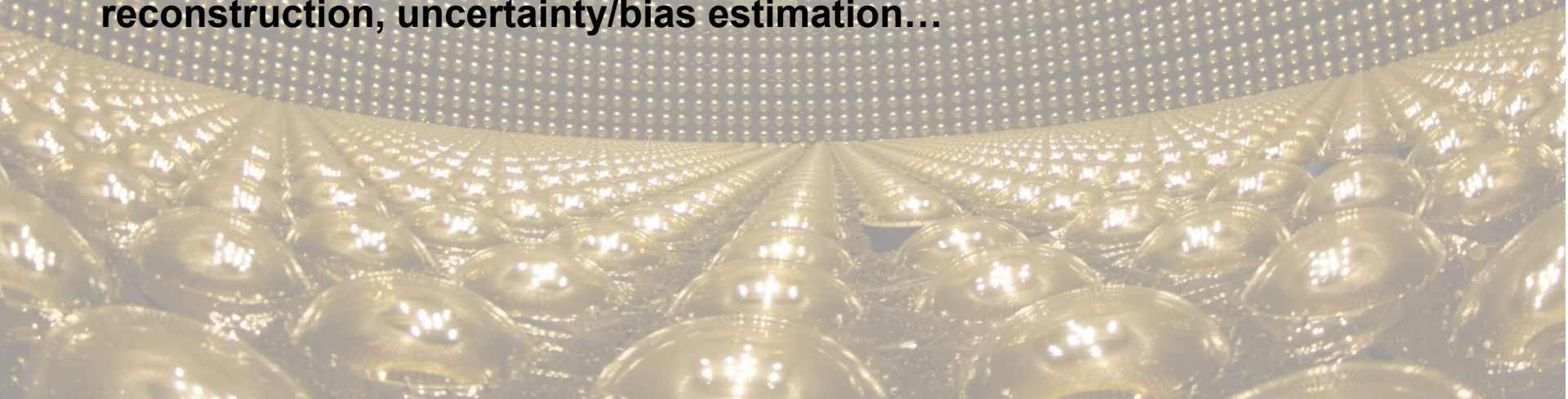
My work in the WatChMaL group

- Trying to improve and extend event classification
- Focused on the IWCD detector, a future detector of the HK experiment
- My goal is to implement ML event labelling within nuPRISM, a physics analysis framework for IWCD.



Hyper-Kamiokande

- **ML Reconstruction is a promising way for event reconstruction in neutrino experiments (other experiments like NoVa, US are already using ML based reconstruction in their analysis)**
- **Still a lot of work to do to include ML in T2K and SK analysis : multi-ring reconstruction, uncertainty/bias estimation...**

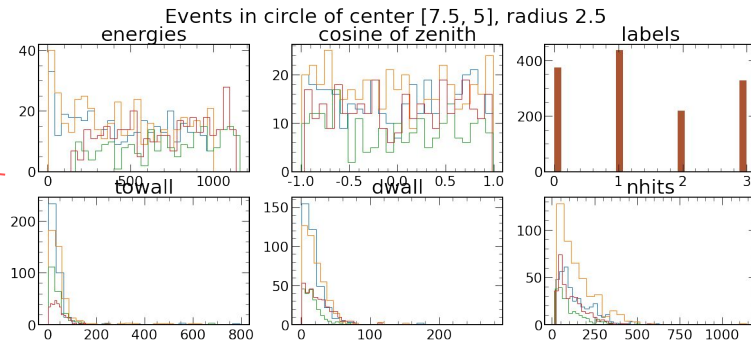
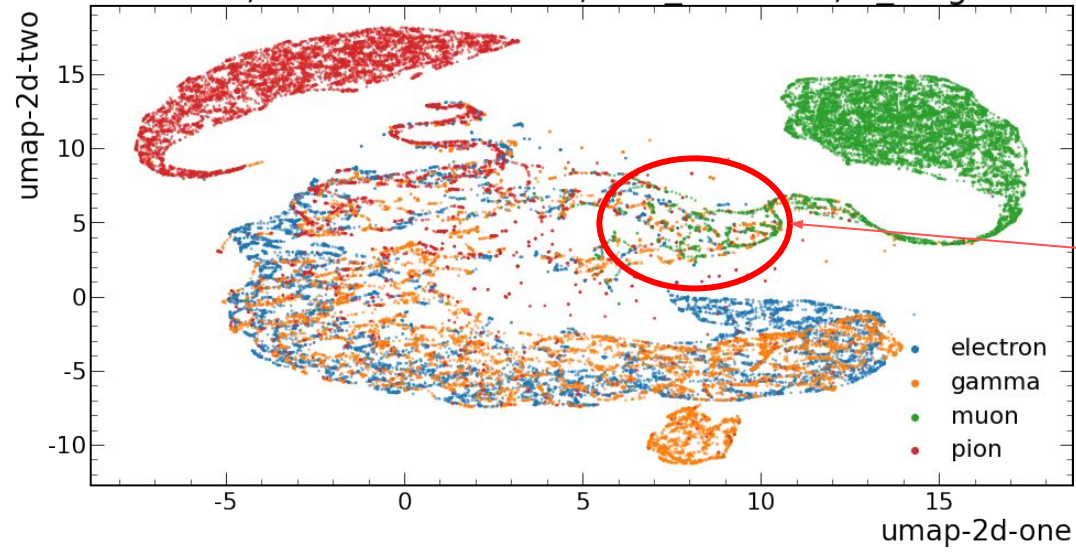


Backup slides

Class separation in latent space

- One of the goals of my internship was to design a ML model able to separate different classes (muon, electron, pion, gamma induced ring) in an abstract latent space

UMAP results, metric : euclidean, min_dist : 0.1, n_neighbors : 85



nuPRISM and IWCD in oscillation analysis

- We fit binned Poisson likelihood to the event rate to predict the cross section of muon and electron neutrinos, as well as their ratios
- Will help reduce uncertainties in the oscillation fit at the far detector

