

Infering cosmological velocity field with graph neural network

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Internship with Hideki Tanimura



Presentation

- ▶ Studying at the ENS de Lyon
- ▶ 2nd year of master in Physics
- ▶ Actually : Intern at IPMU

Main interest in Physics : Theory, Cosmology and Condensed matter

Other interest : Doing art, Learning japanese, Music

Introduction

- ▶ Λ CDM model tells us that the 84% of matter is Dark Matter, not observable
- ▶ Galaxies are nested in dark matter halo, forming a cosmic web

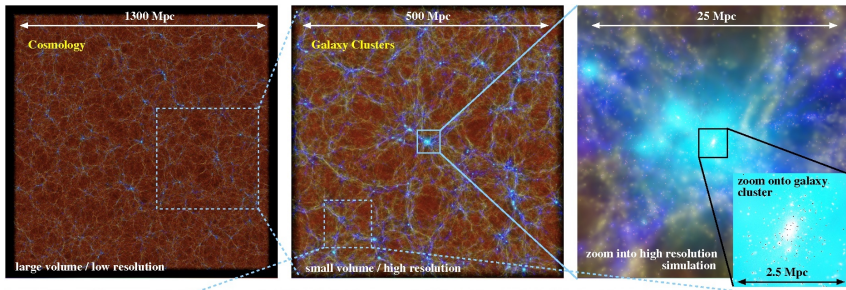


Figure: The cosmic web, in the Magneticum simulation

Introduction

- ▶ Goal of this internship : finding the velocity of these halo

Many applications, for example : correcting the Sunyaev-Zel'dovich effect on CMB observation

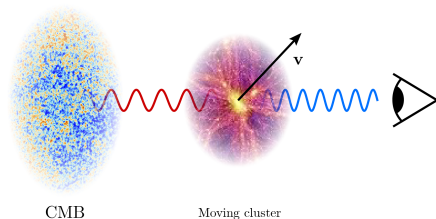


Figure: The Sunyaev-Zel'Dovich effect

But, problem :

- ▶ We cannot observe these halos
- ▶ Measuring peculiar velocity is difficult

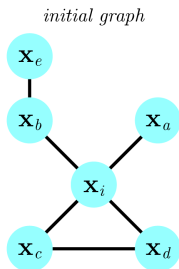
To remedy to that : let's try to use GNN

What is Graph Neural Network ?

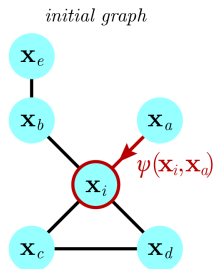
- ▶ Graph : a set of nodes linked by edges, how we want to model the cosmic web
- ▶ Neural Network : a function $\mathbb{R}^n \rightarrow \mathbb{R}^m$ with a set of parameters that can be optimised

Idea : training a GNN model on simulation (Magneticum) to find that velocity, in order to apply it to real data

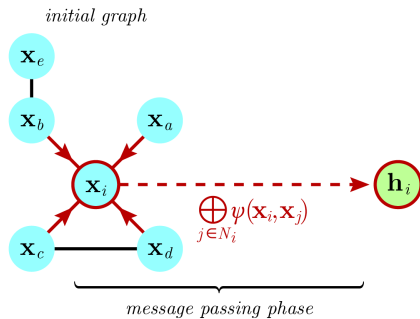
How does a Graph Neural Network works



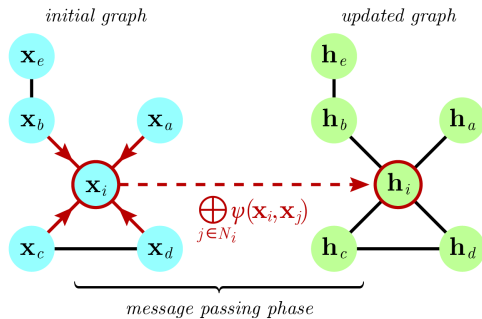
How does a Graph Neural Network works



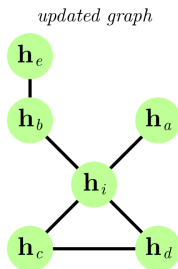
How does a Graph Neural Network works



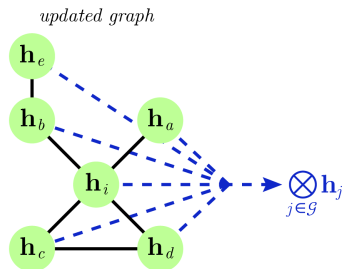
How does a Graph Neural Network works



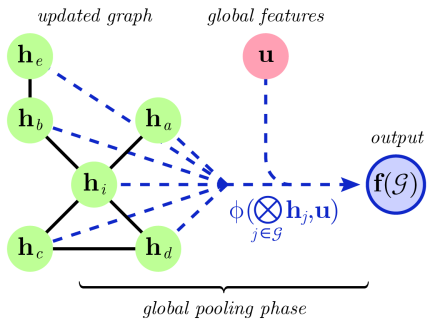
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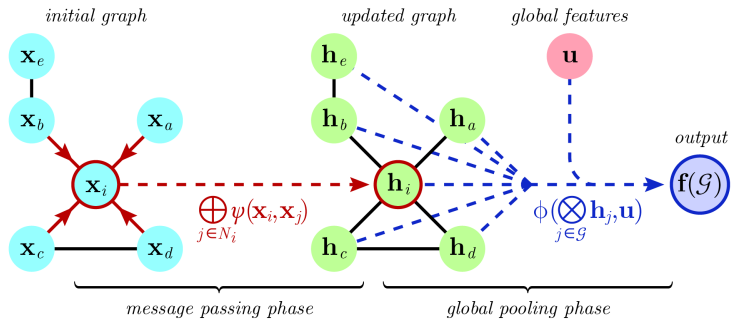
How does a Graph Neural Network works



How does a Graph Neural Network works



How does a Graph Neural Network works



Applying GNN to clusters of galaxies

- ▶ Nodes : all galaxies within R_{max} of center halo
- ▶ Node features : position, and mass
- ▶ Global features : total number of galaxies + total mass
- ▶ Edges : two galaxies are linked if they are close enough $d < d_n$

Problem of GNN → MANY hyperparameters

- ▶ learning rate, weight decay, n_{layers} , $n_{neurons}$, n_{mp} , size of the hidden vectors
- ▶ plus choice of aggregation functions \oplus and \otimes
- ▶ plus d_n , R_{max}

Applying GNN to clusters of galaxies

How to choose d_n :

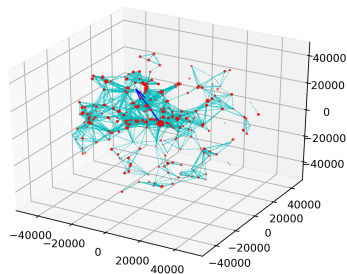
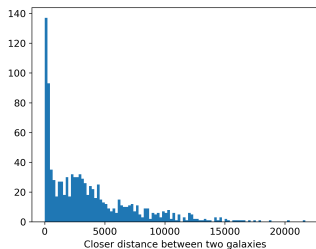


Figure: Plot of the distribution of the closest distance between two galaxies, and the graph obtained with $d_n = 15$ Mpc

Results

Training in /home/balbert/data/sims/Magneticum/box0/

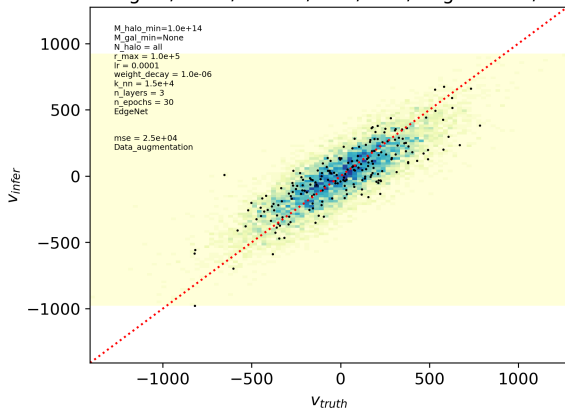


Figure: Accuracy of the tuned GNN

And now ?

- ▶ Trying to compare the accuracy with other features
- ▶ Trying to see if there is a way to interpret the GNN