

KATRIN : from neutrino mass to dark matter

LAPP Annecy

Thibaut Houdy
12th of June, 2023

Qui suis-je?

Doctorat, APC/CEA-Irfu, U. Paris-Diderot (2014 – 2017)

- Etude des **neutrinos solaires (^8B)** et **stériles (eV)** dans **Borexino**

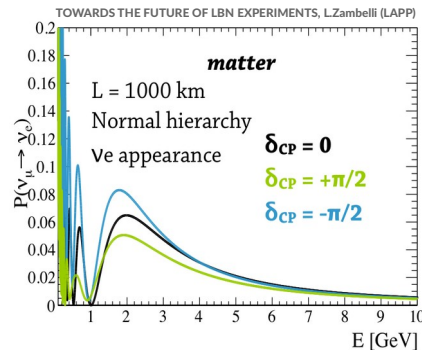
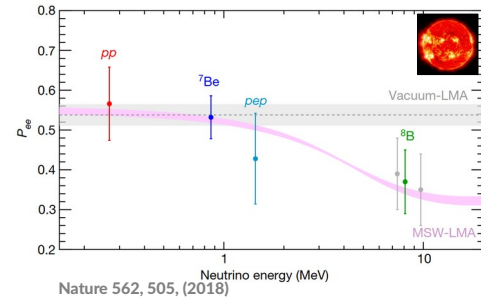
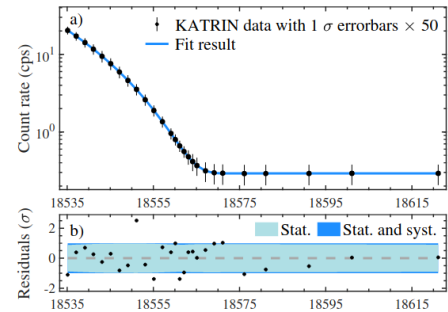
Postdoctorat, Max-Planck Physique, Munich, (2018 - 2021)

- Rejoins **KATRIN (masse du neutrino)**
- Recherche de **neutrinos stériles au keV** avec **KATRIN**

Maitre de conférence, l'U. Paris-Saclay, IJCLab (2021 -)

- Cours en Electromagnétisme, Instrumentations et Nucléaire & Particules. Responsable de la plateforme $E_2\text{PN}$
- Rejoins DUNE (**hiérarchie de masse**, **phase CP**, unitarité de la matrice PMNS)

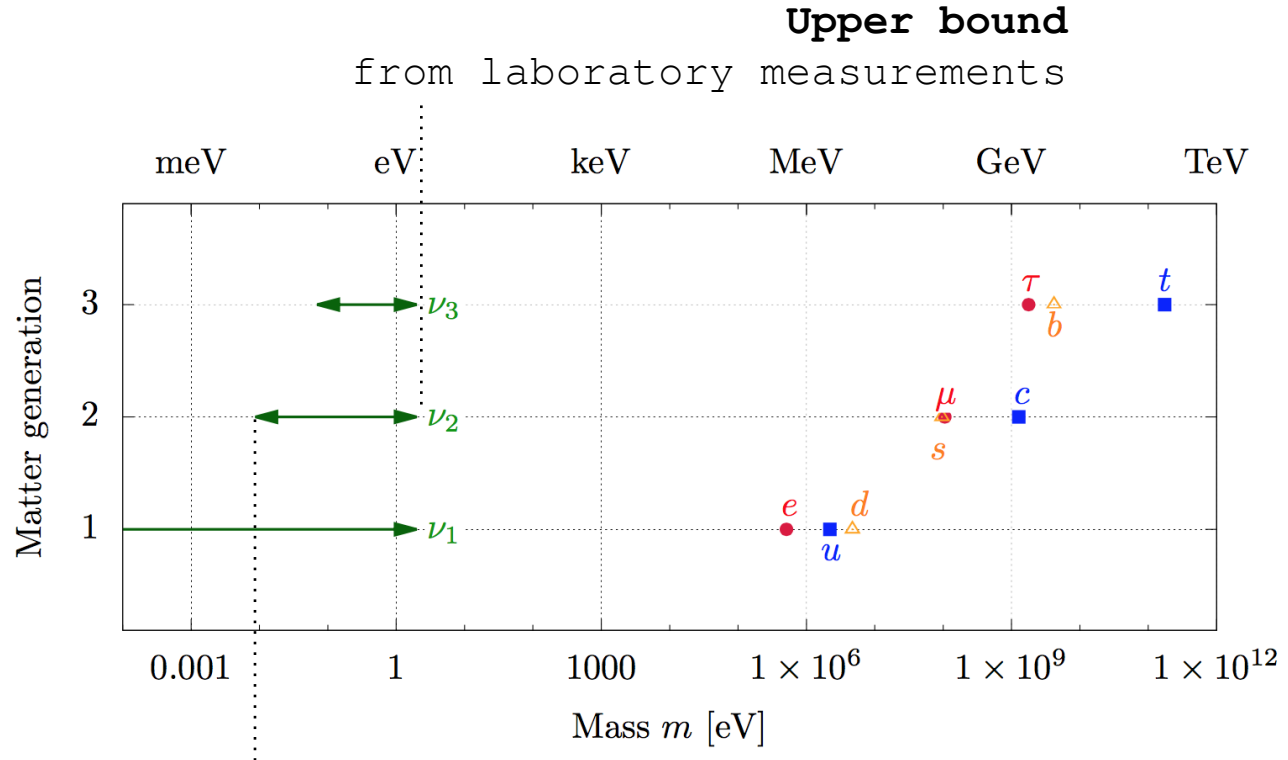
This talk is NOT made in the name of the KATRIN collaboration ;)



Neutrino mass status

$$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2} \left| \begin{array}{l} 7.42^{+0.21} \\ -0.20 \end{array} \right.$$

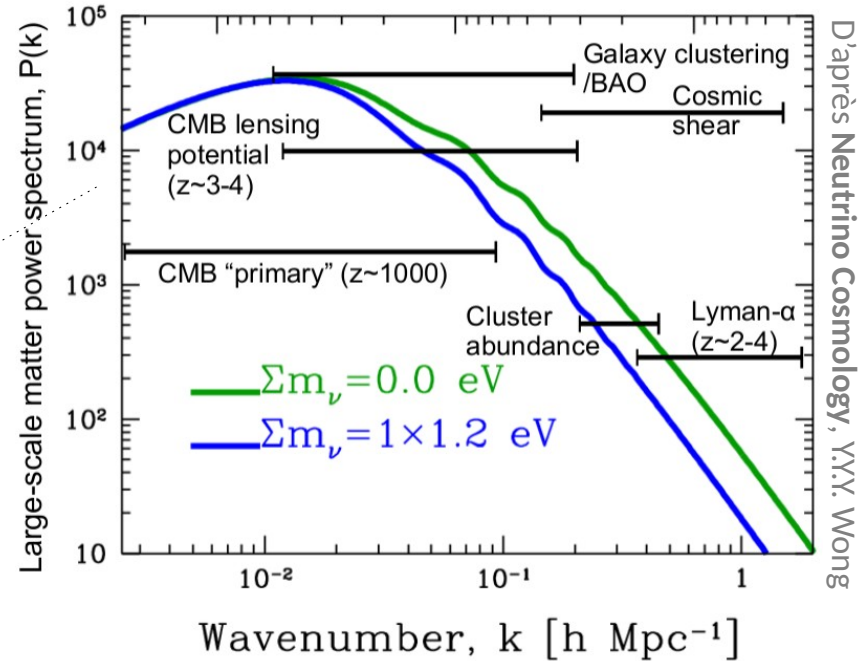
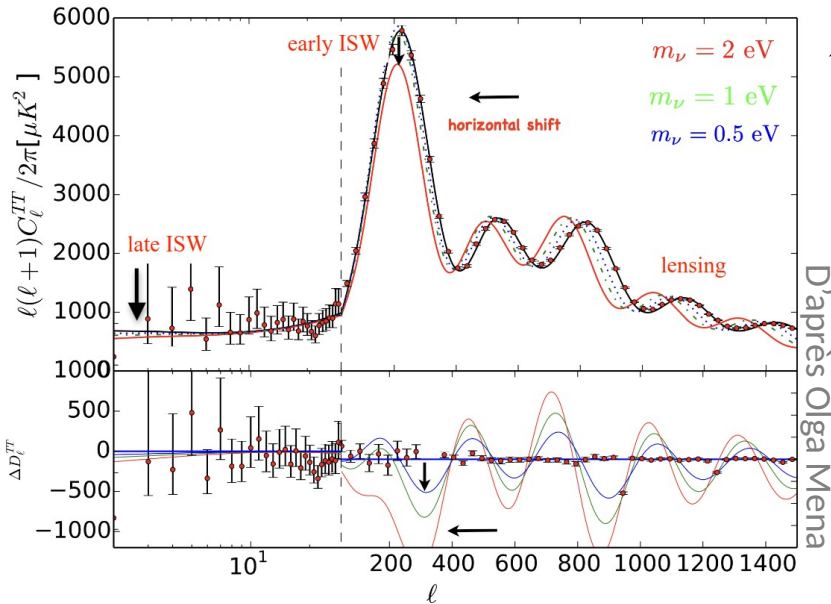
$$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2} \left| \begin{array}{l} +2.515^{+0.028} \\ -0.028 \end{array} \right.$$



Lower bound
from oscillation experiments

Neutrino mass measurements

- **Cosmology**
model dependent
 $\Sigma m_\nu < 120 \text{ meV}$
Ex : Planck, eBOSS



- Massive neutrinos wash out small scale structure formation after non-relativistic transition
- Measurement of the CMB lensing \rightarrow more massive, less structure, less sensing

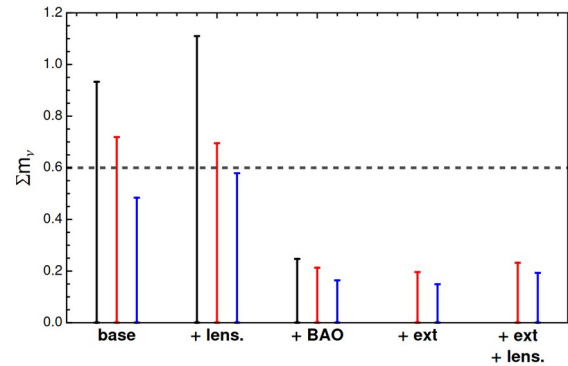
Neutrino mass measurements

- **Cosmology**

model dependent

$$\Sigma m_\nu < 120 \text{ meV}$$

Ex : Planck, eBOSS

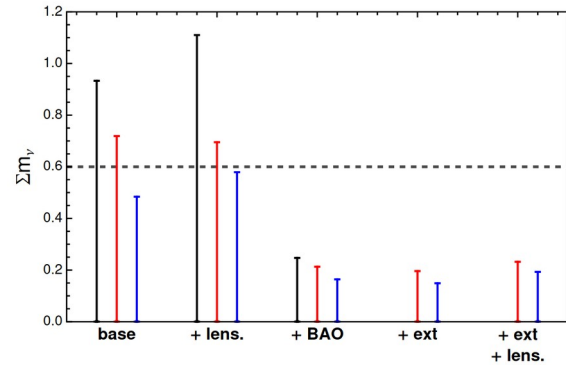


M.Lattanzi 2016 JPConf.:Ser:718 032008

Neutrino mass measurements

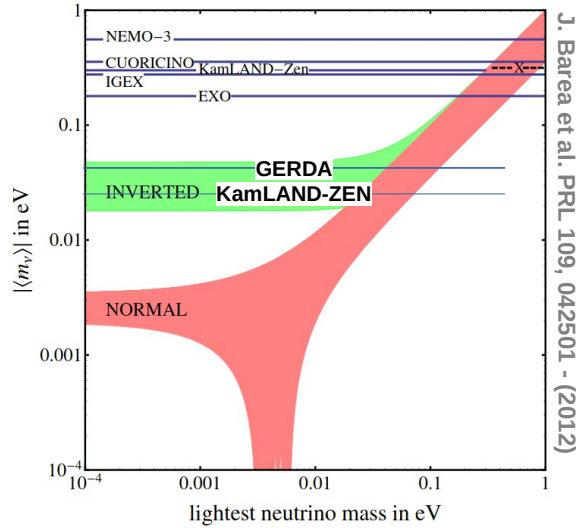
- **Cosmology**
model dependent
 $\Sigma m_\nu < 120 \text{ meV}$
Ex : Planck, eBOSS

- **$0\nu\beta\beta$**
laboratory based
 $m_{\beta\beta} < 100 \text{ meV}$
Ex : CUPID, LEGEND, etc



M.Lattanzi 2016 JPConf.Ser.718 032008

Neutrino mass measurements



- $0\nu\beta\beta$
laboratory based
 $m_{\beta\beta} < 100$ meV
Ex : CUPID, LEGEND, etc

Exclusion from measurement

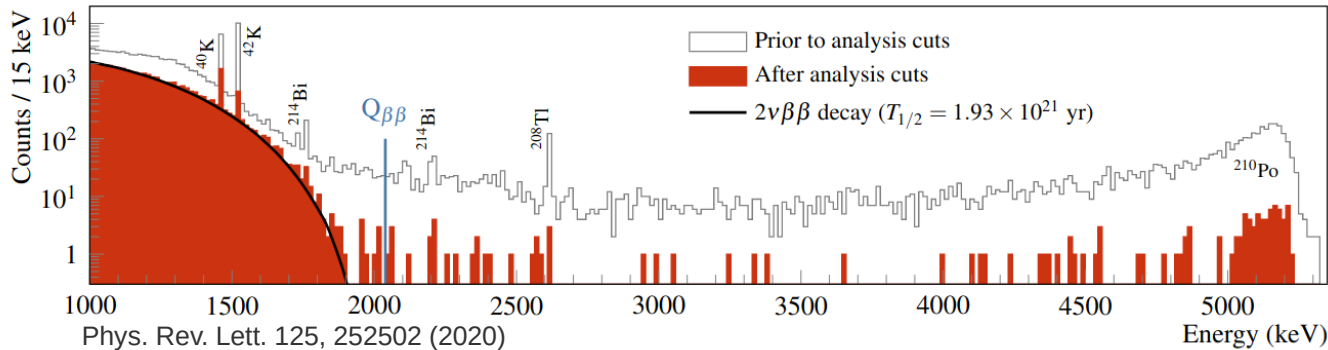
$$({}^{0\nu}T_{1/2})^{-1} \propto |M^{0\nu}|^2 m_{\beta\beta}^2$$

Matrix element not easily derived

$$m_{\beta\beta} = \left| \sum_{i=1}^3 U_{ei}^2 m_i \right|$$

BUT :

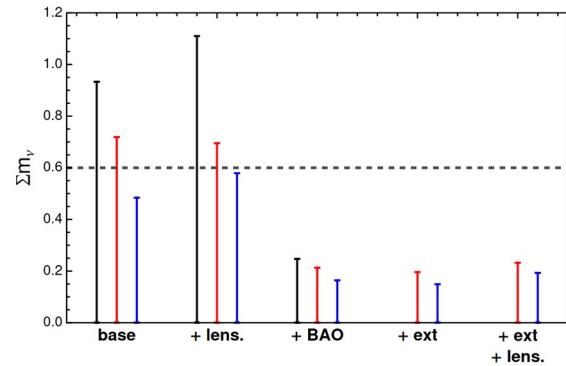
- only true if Majorana neutrino
- Depending on an unknown Majorana phase



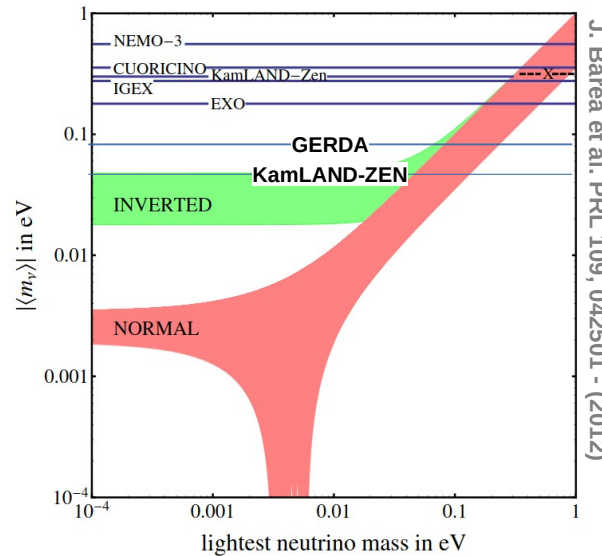
Neutrino mass measurements

- **Cosmology**
model dependent
 $\Sigma m_\nu < 120 \text{ meV}$
 Ex : Planck, eBOSS

- **$0\nu\beta\beta$**
laboratory based
 $m_{\beta\beta} < 100 \text{ meV}$
 Ex : CUPID, LEGEND, etc



M. Lattanzi 2016 JP Conf., Ser. 718 032008

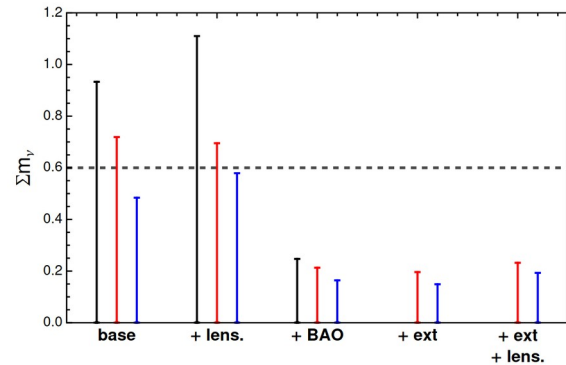


Neutrino mass measurements

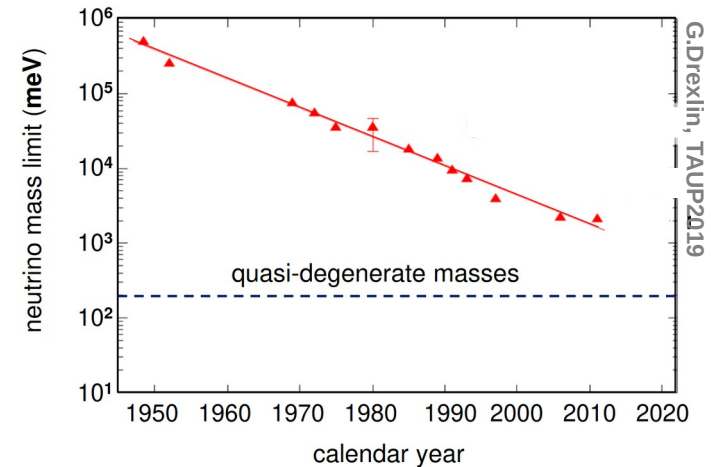
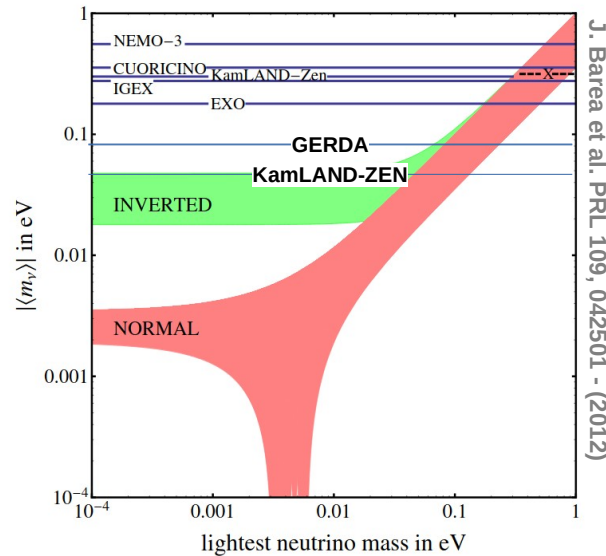
- **Cosmology**
model dependent
 $\Sigma m_\nu < 120 \text{ meV}$
 Ex : Planck, eBOSS

- **$0\nu\beta\beta$**
laboratory based
 $m_{\beta\beta} < 100 \text{ meV}$
 Ex : CUPID, LEGEND, etc

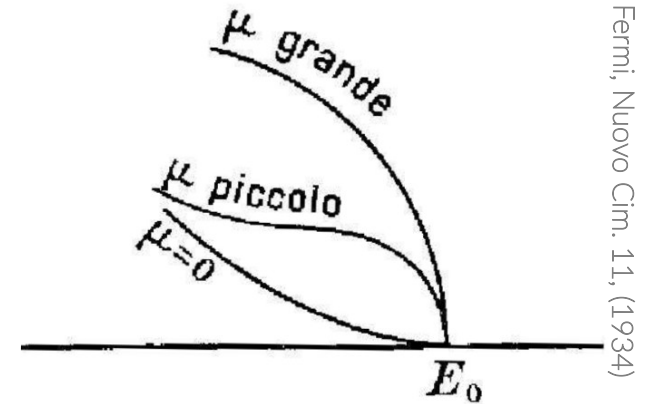
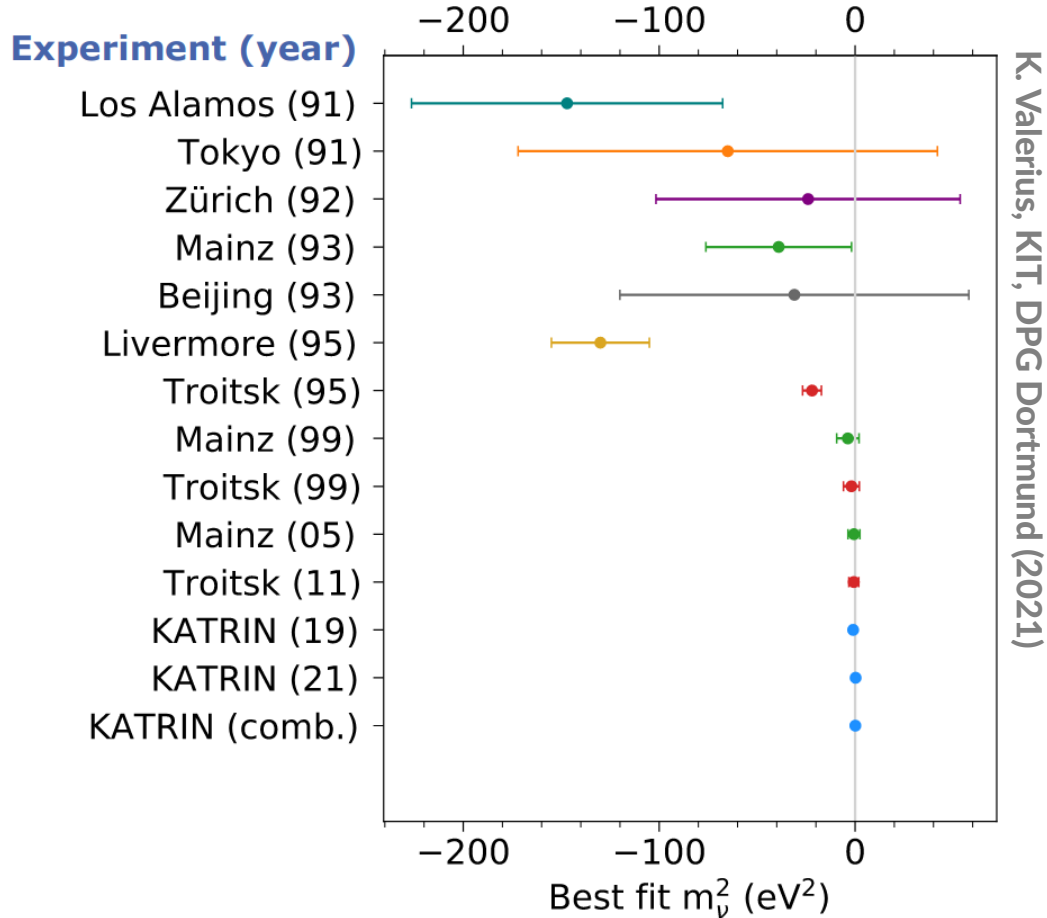
- **Kinematics of β -decay**
laboratory based
 $m_{\nu_e} = 50\text{-}200 \text{ meV}$
 Ex : ECHO, KATRIN, Project8



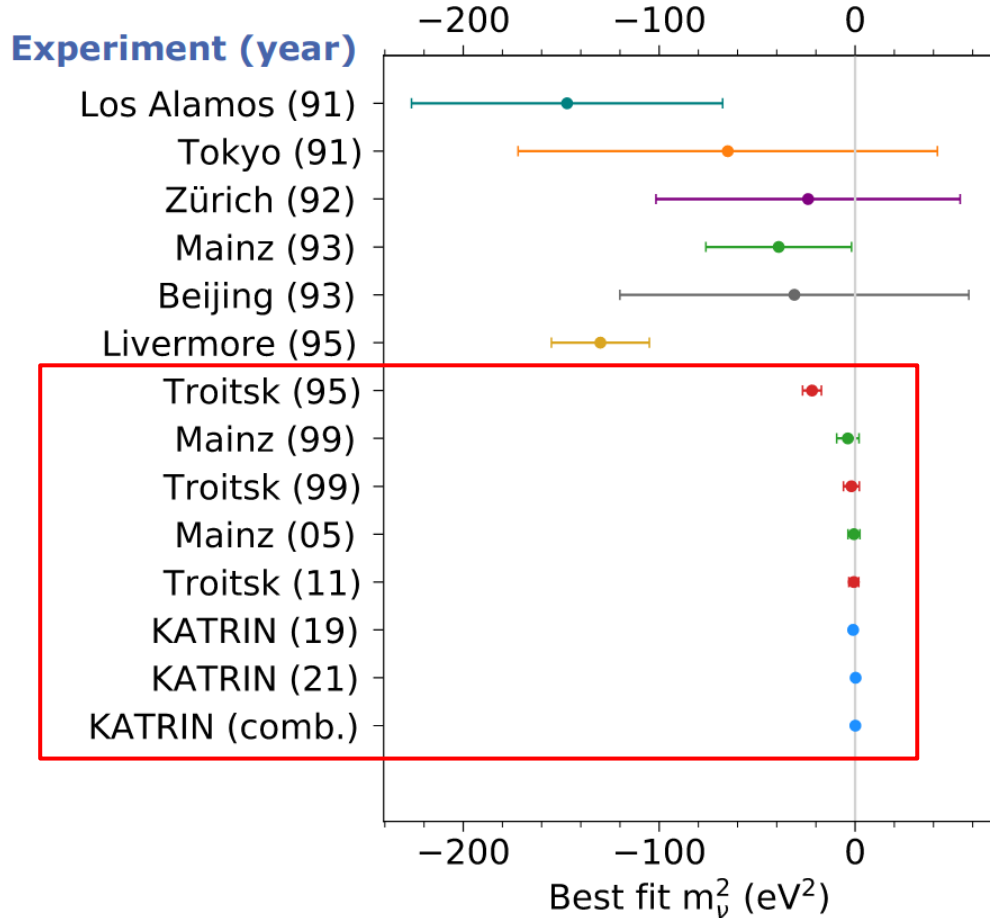
M.Lattanzi 2016 JPConf.:Ser.718 032008



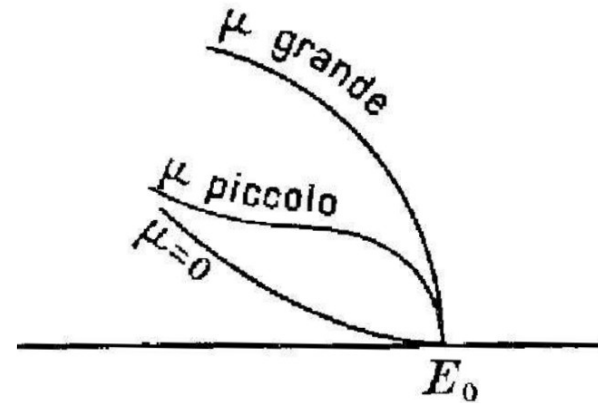
Determining m_ν from β -decay



Determining m_ν from β -decay



K. Valerius, KIT, DPG Dortmund (2021)



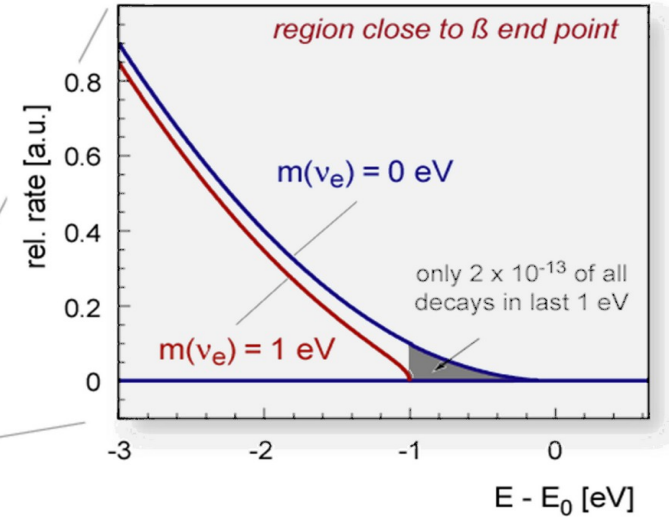
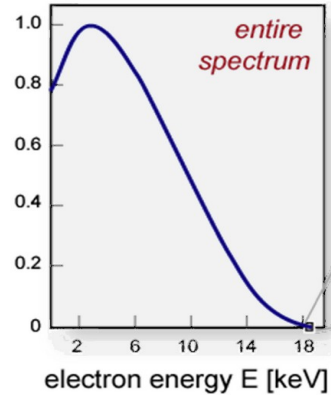
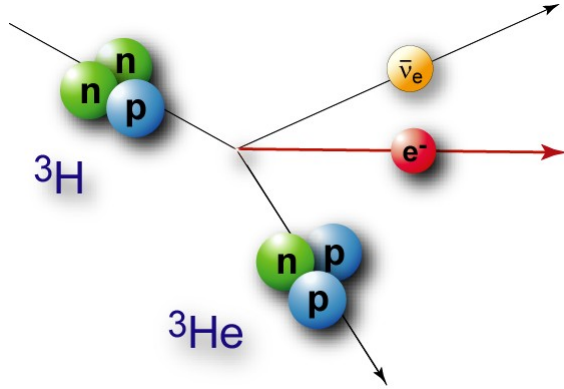
Fermi, Nuovo Cim. 11, (1934)

- Troitsk
- Mainz
- KATRIN

→ MAC-E filters with tritium source

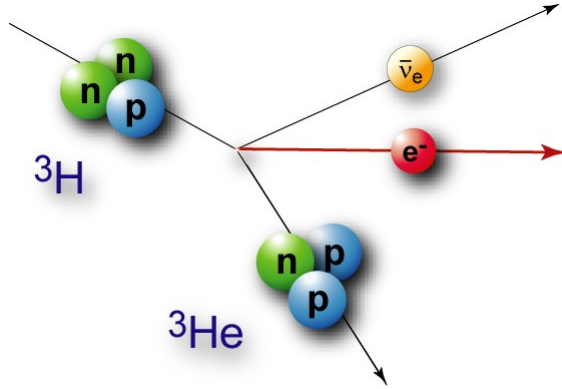
Determining m_ν from β -decay: KATRIN

General Idea



Determining m_ν from β -decay: KATRIN

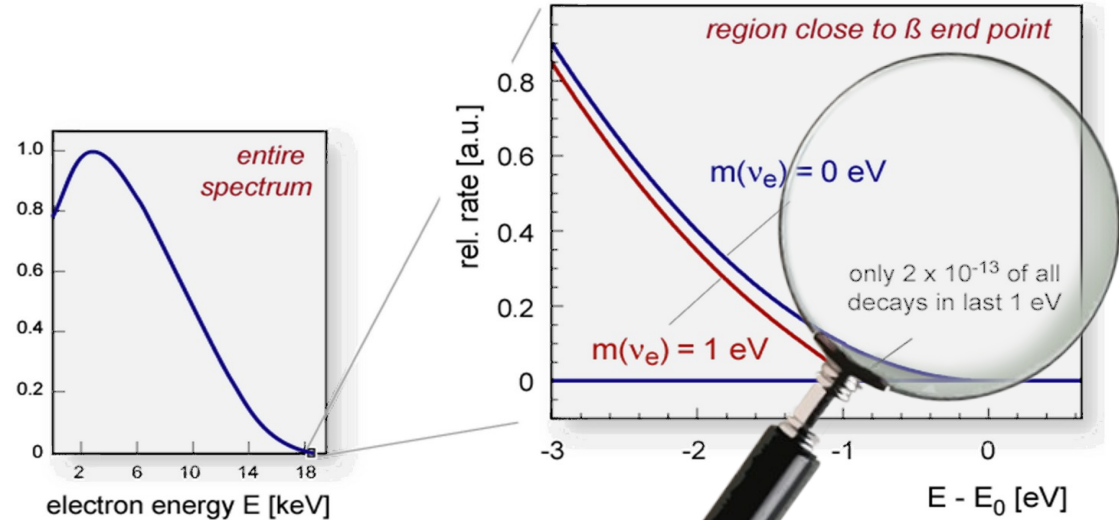
General Idea



- Ultra-strong β -source 10^{11} decays/s
- Low background level < 0.1 cps
- Excellent energy resolution ~ 1 eV
- Precise understanding of spectrum



16/06/2023



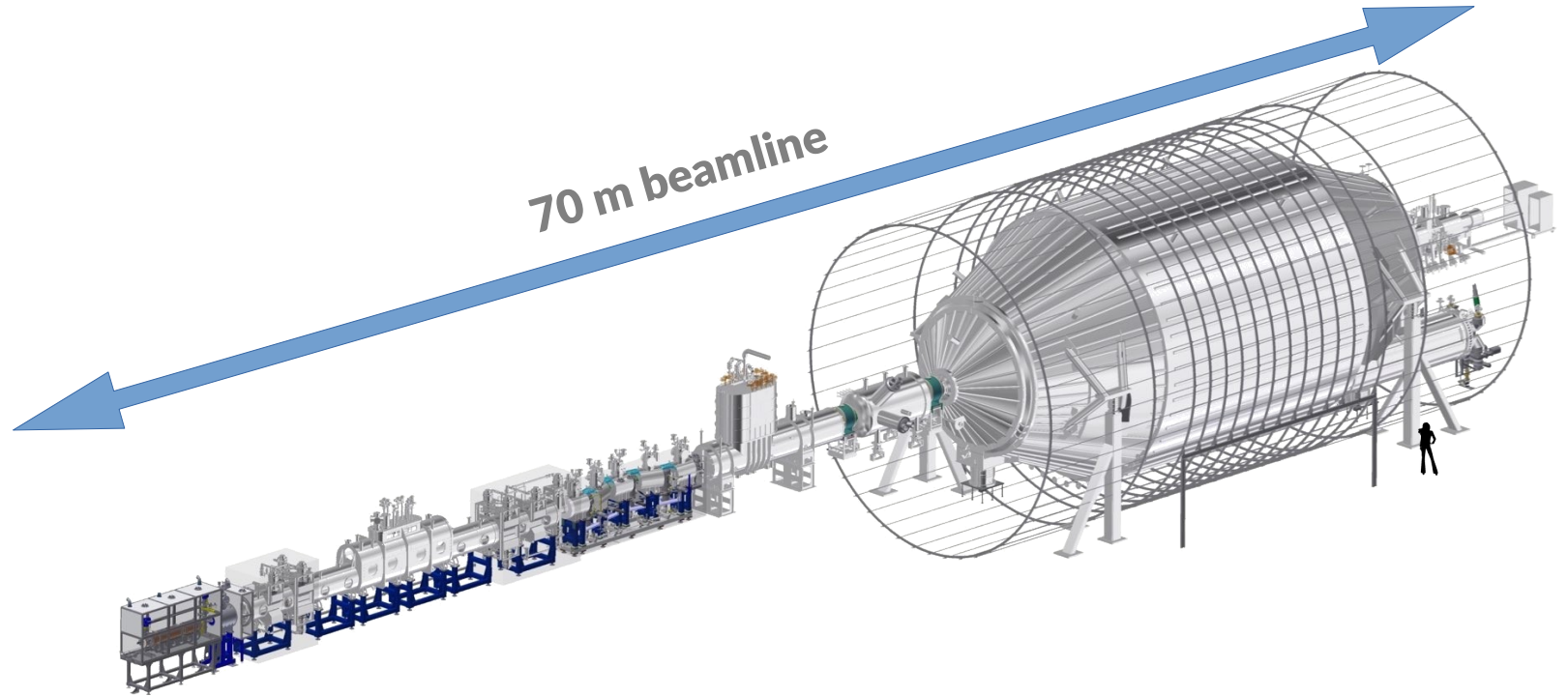
Karlsruhe TRitium Neutrino Experiment : KATRIN



- Experimental site : Karlsruhe Institute of Technology (KIT)
- International Collaboration (150 members)
- Design sensitivity: 0.2 eV (90% CL)



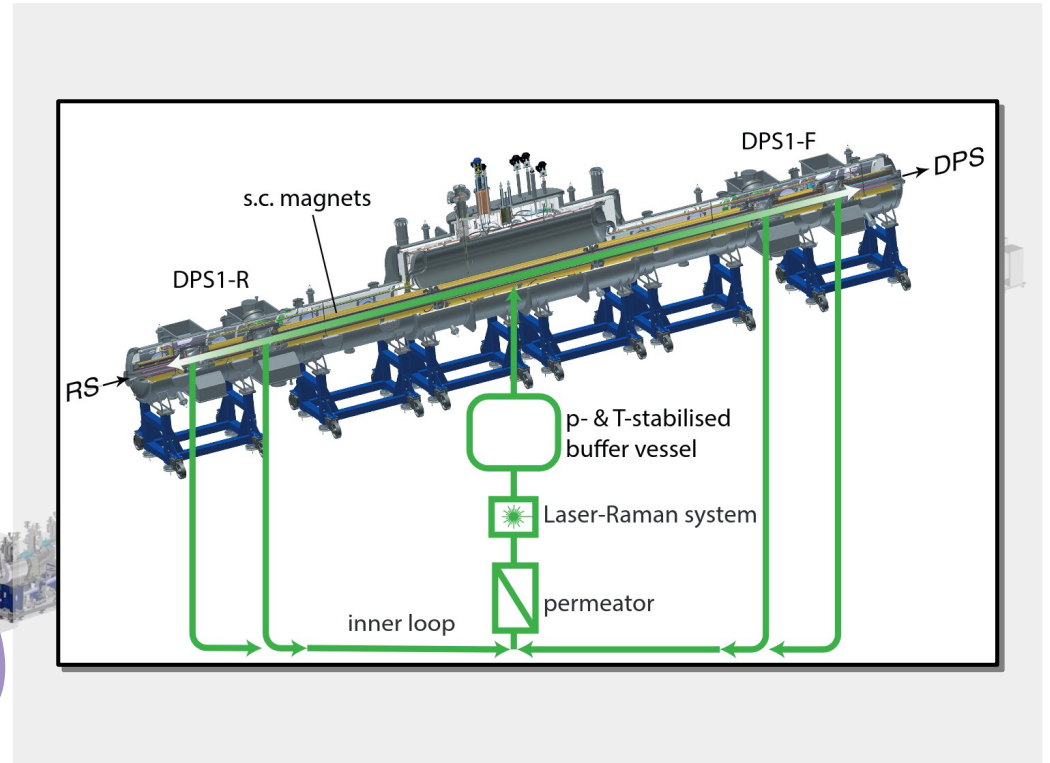
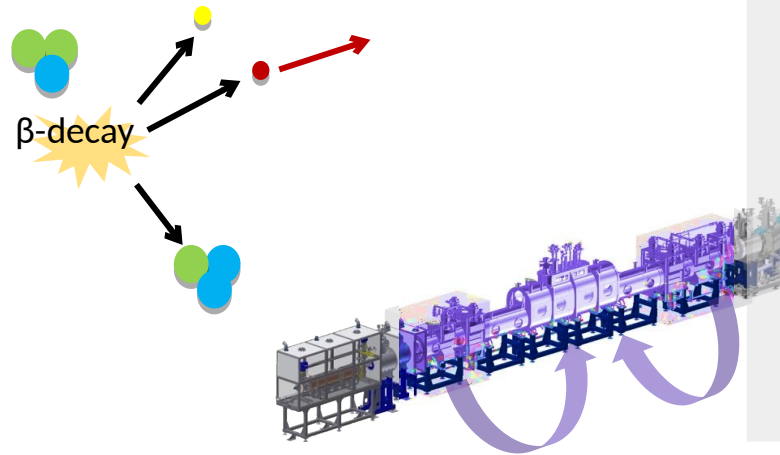
KATRIN Working Principle



KATRIN Working Principle

Windowless gaseous tritium source

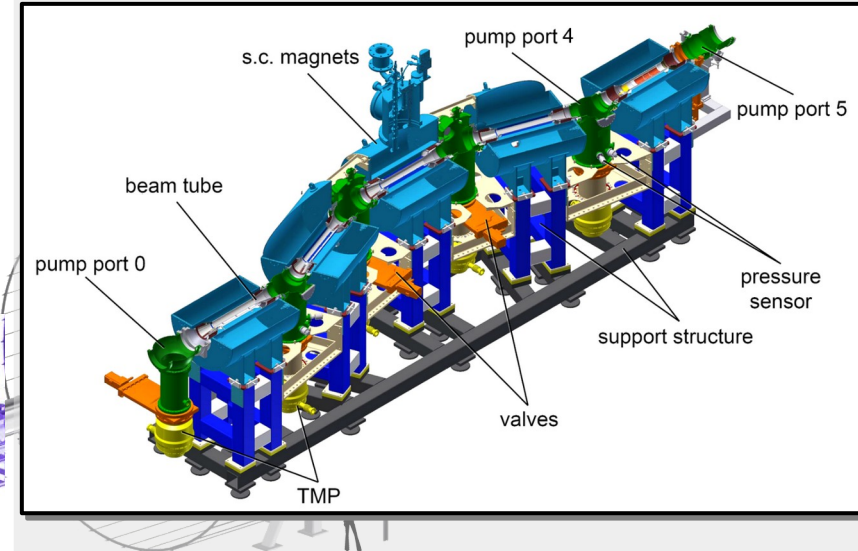
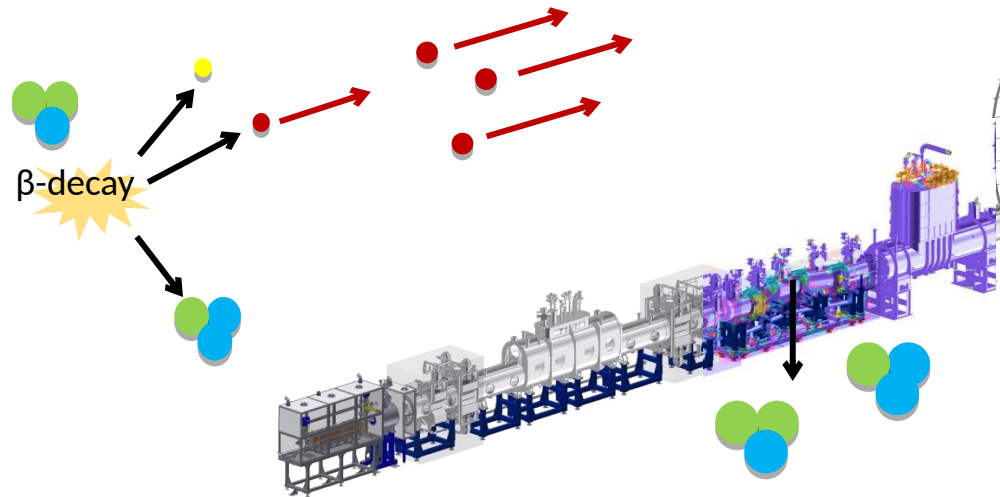
- molecular tritium in closed loop system
- 10^{11} decays/s



KATRIN Working Principle

Transport section

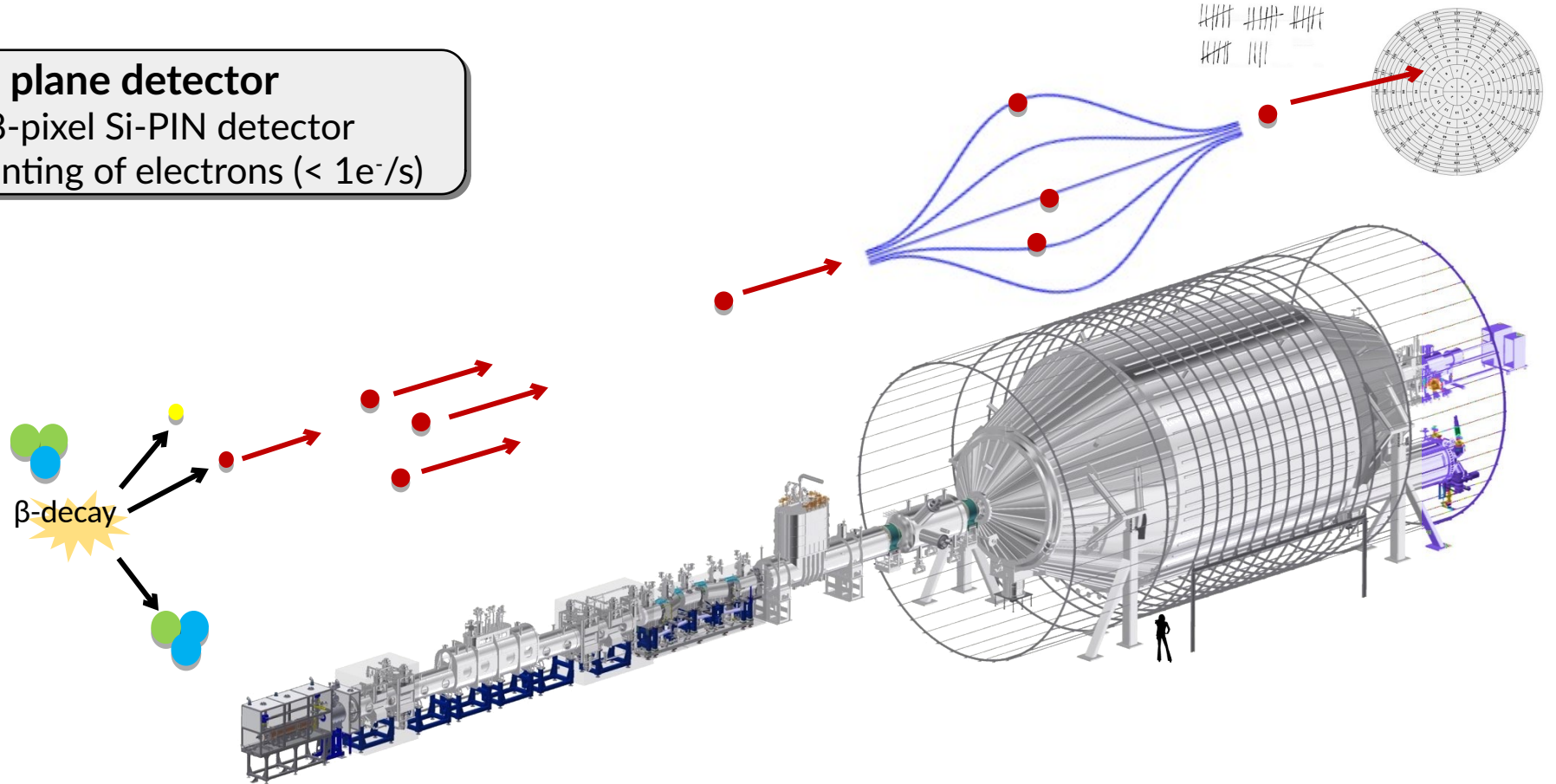
- magnetic guidance of electrons (@ 4 T)
- tritium flow reduction by $> 10^{14}$ + tritium ion removal



KATRIN Working Principle

Focal plane detector

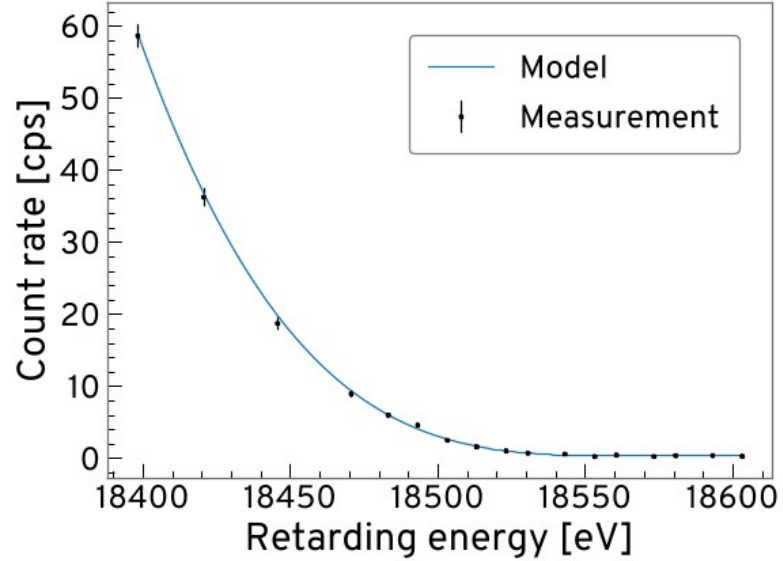
- 148-pixel Si-PIN detector
- counting of electrons ($< 1e^-/s$)



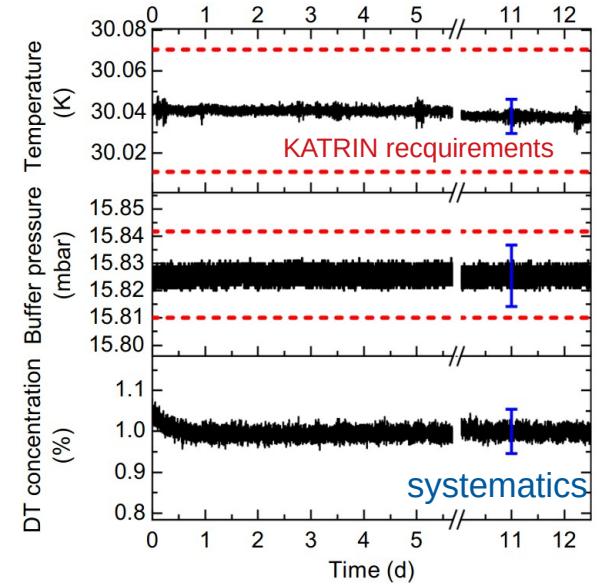
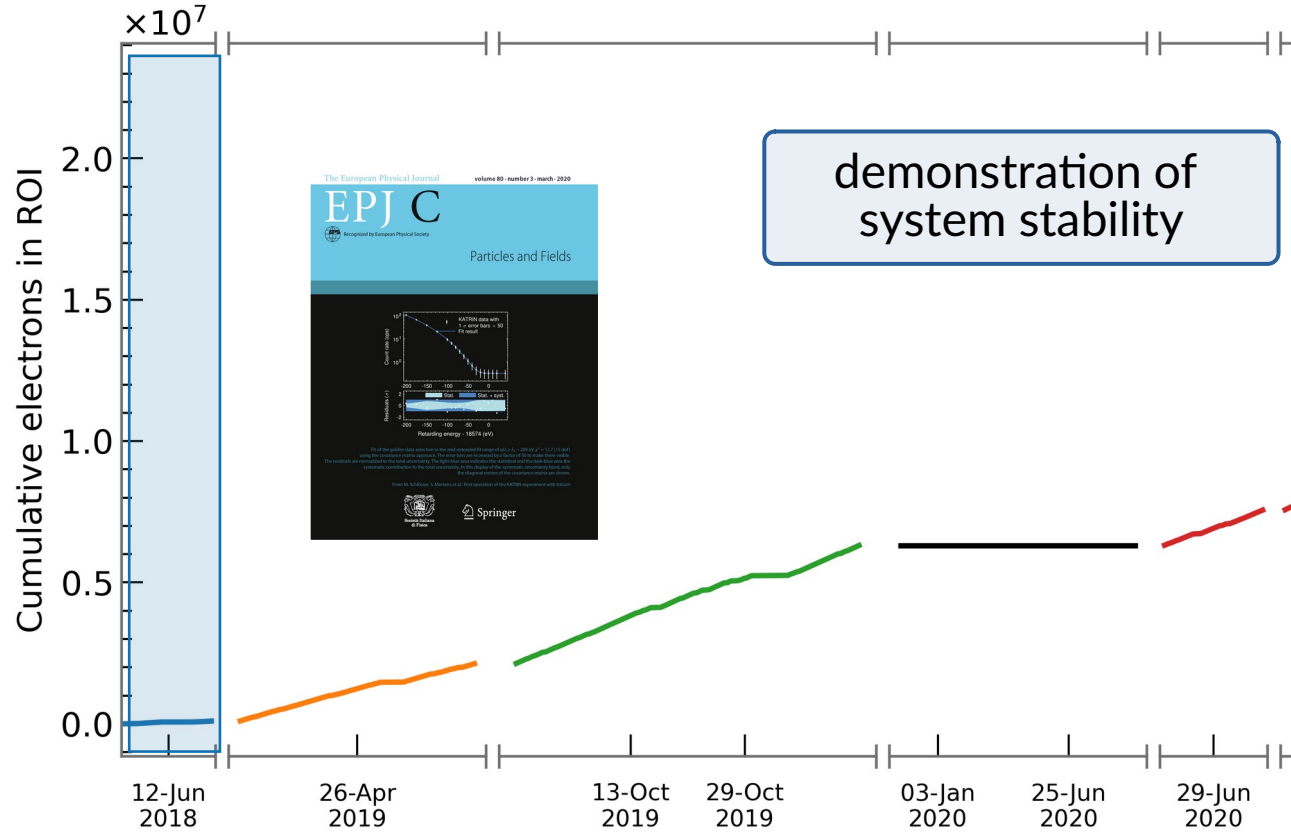
KATRIN Working Principle

Focal plane detector

- 148 pixel Si PIN detector

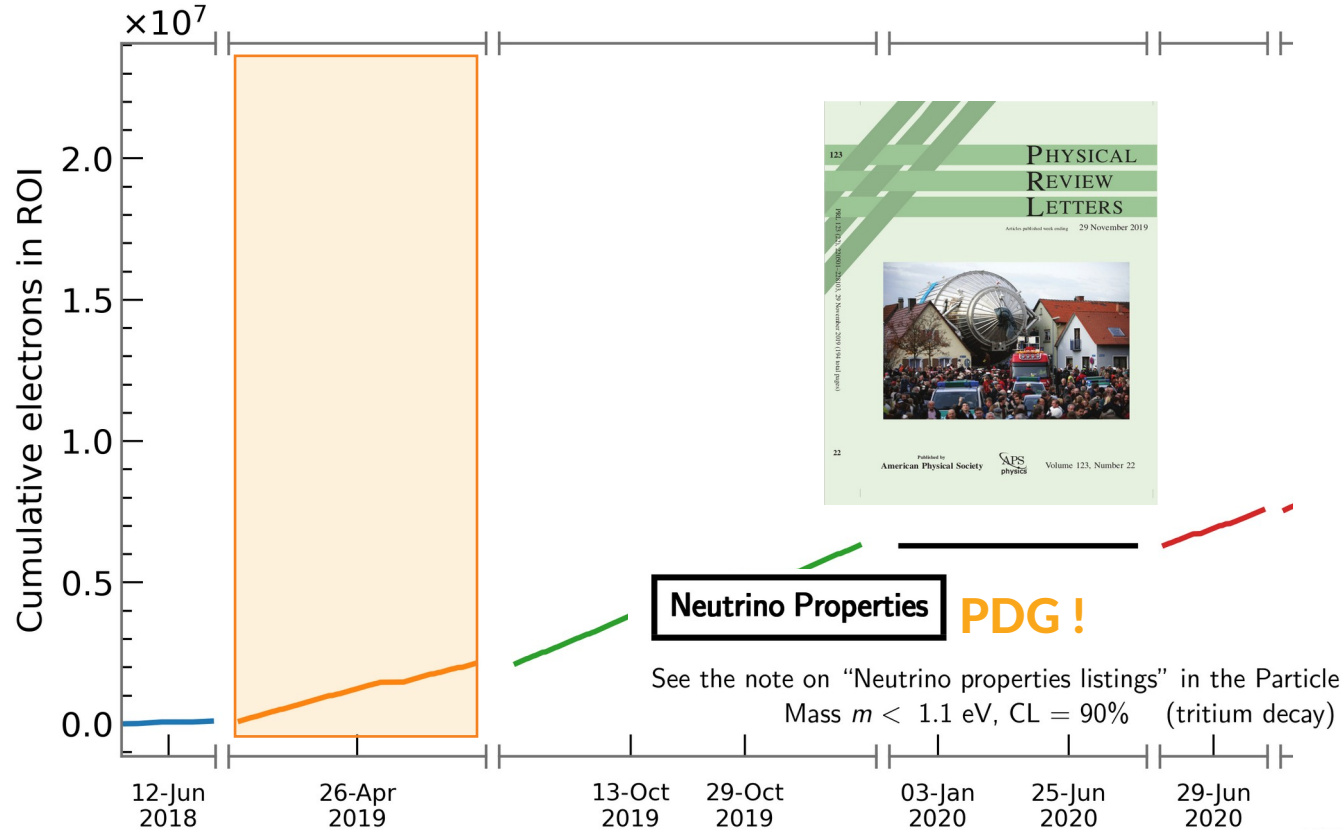


First tritium campaign

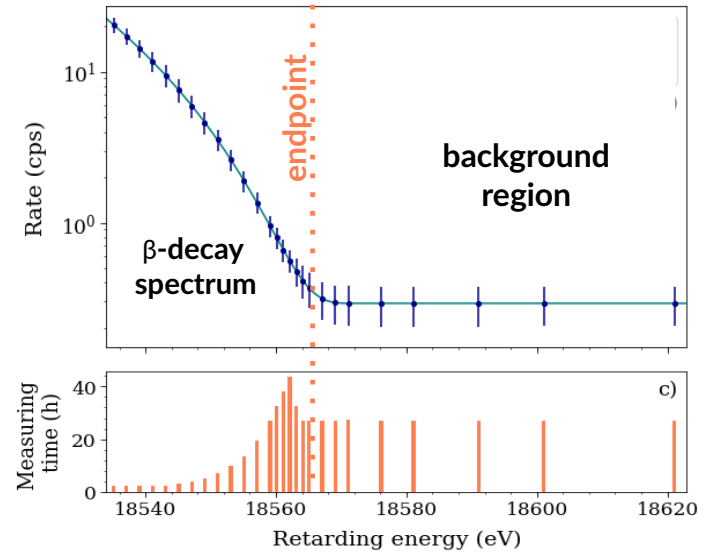


- First operation of the KATRIN experiment with tritium. *Eur. Phys. J. C* 80, 264 (2020)

1st neutrino mass campaign

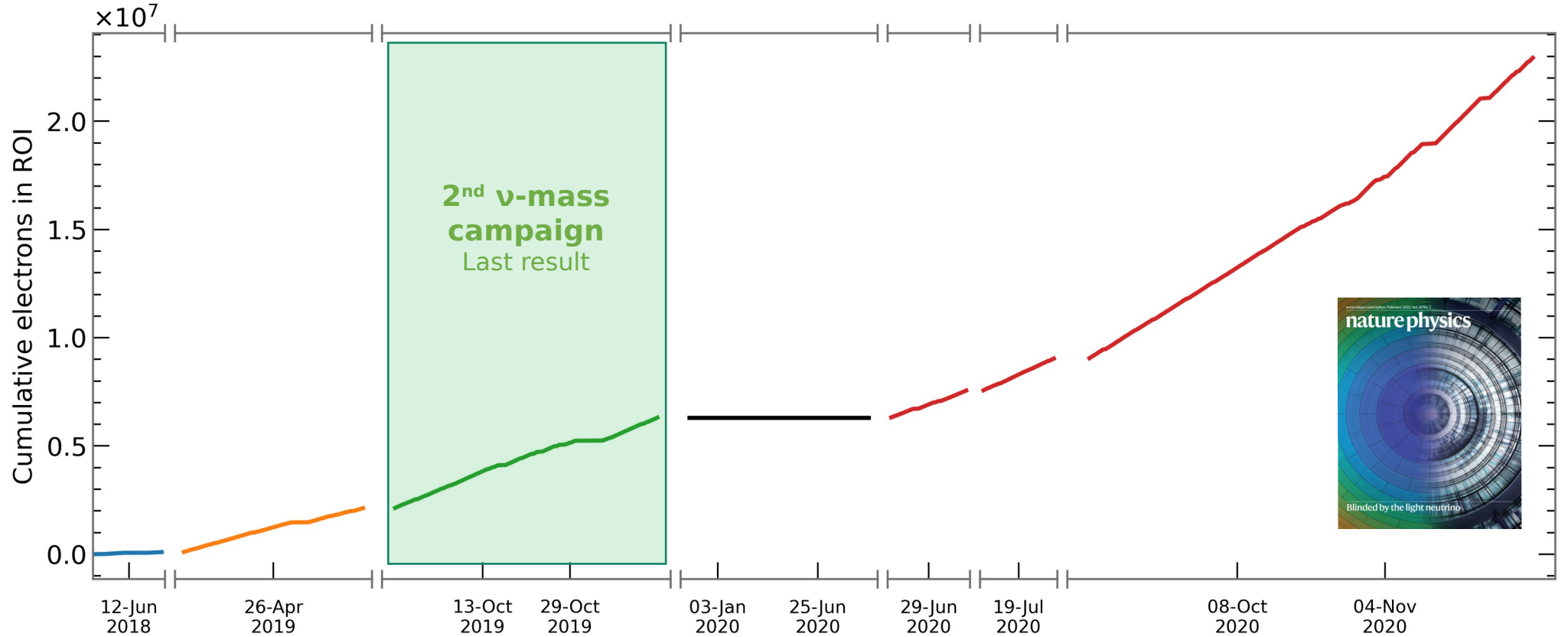


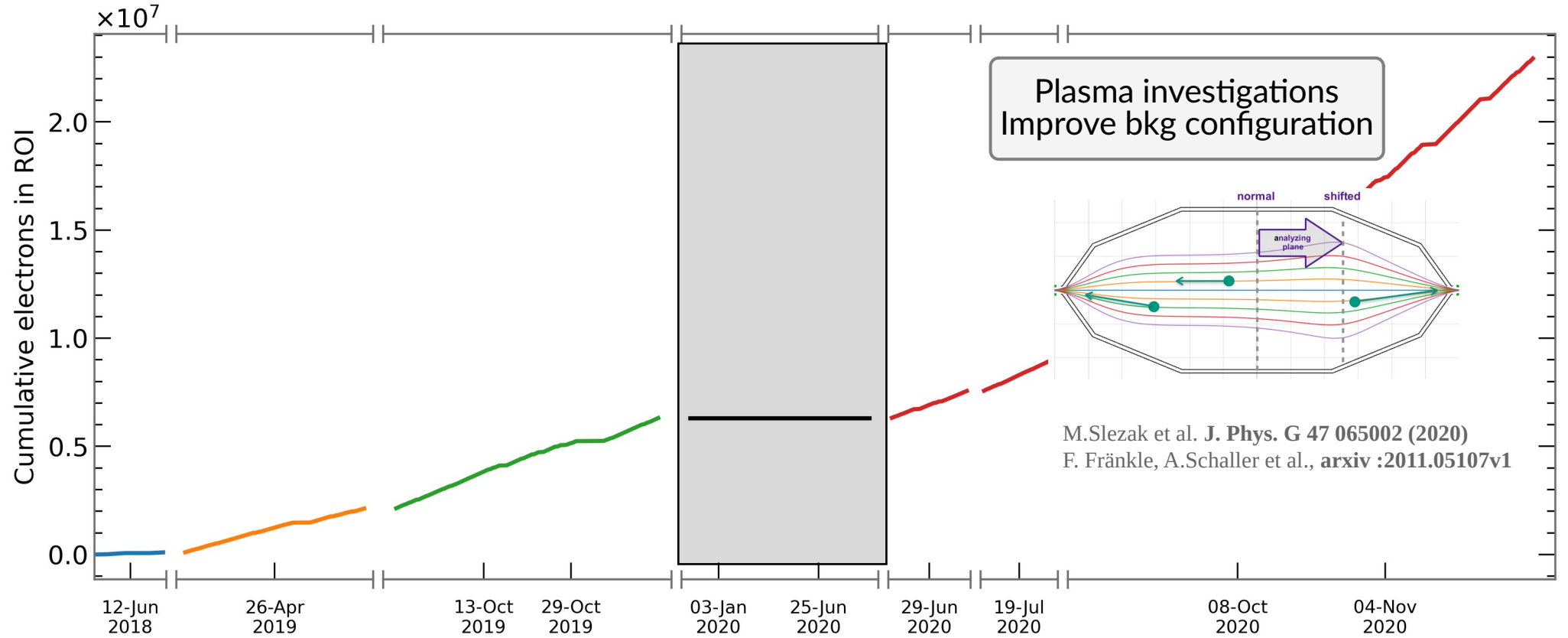
- Measurement time: **22 days**
- Gas density: **22%**
- Isotopic purity: **97.5% tritium**
- Source activity: **$2.45 \cdot 10^{10}$ Bq**
- Total statistics: **$2 \cdot 10^6$ e^s**



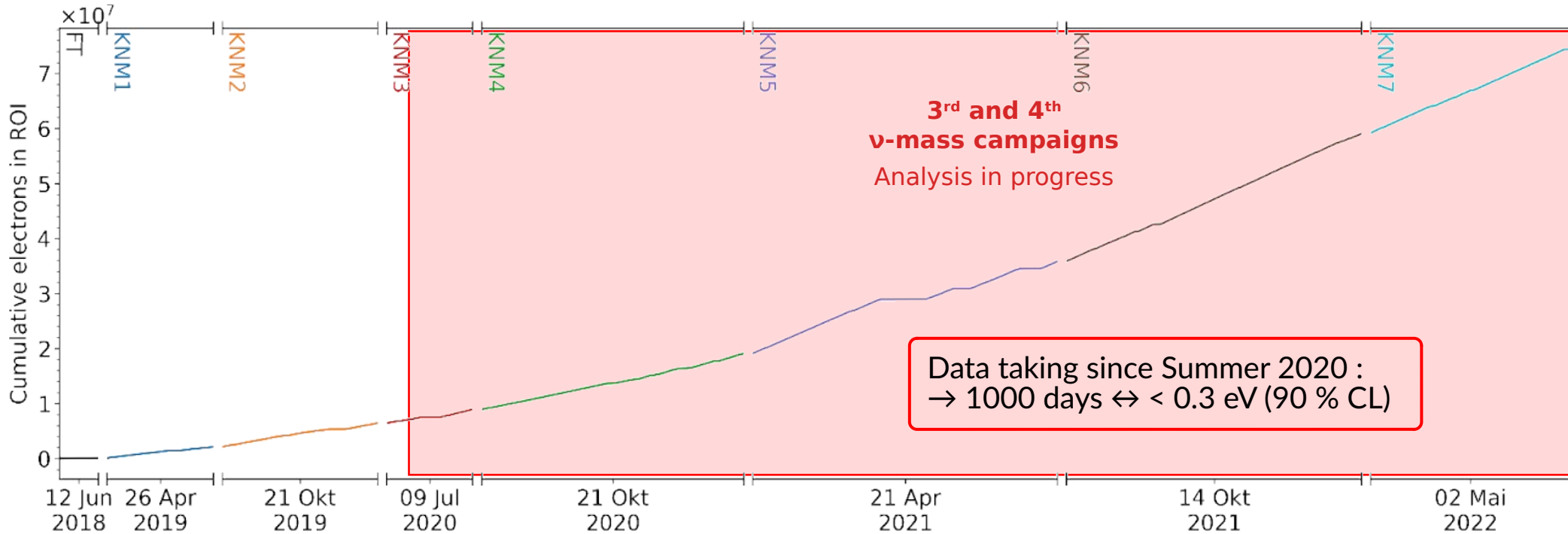
- *Improved Upper Limit on the Neutrino Mass from a Direct Kinematic Method by KATRIN, KATRIN Collaboration, Phys. Rev. Lett. 123, 221802 (2019)*

2nd neutrino mass campaign

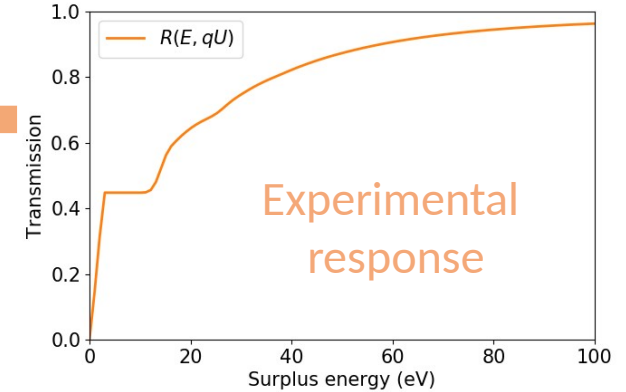
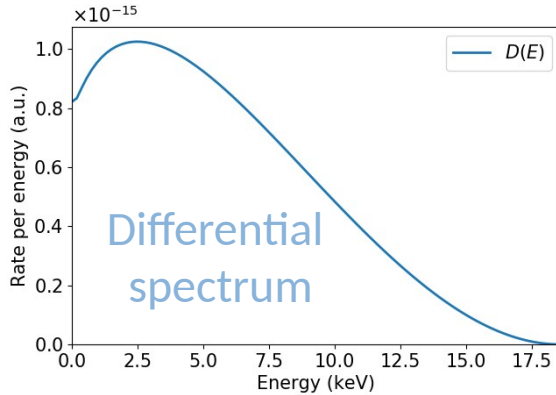




Cruising mode

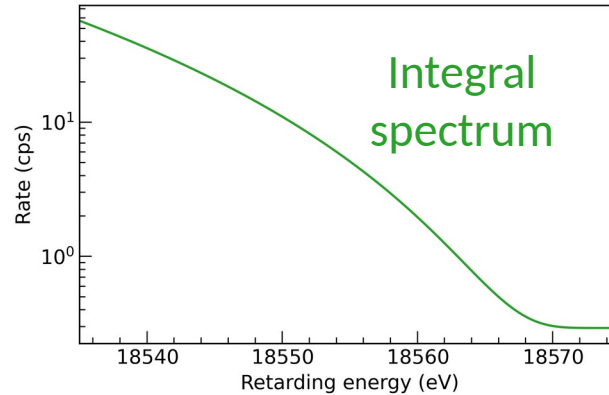


Model



$$\Gamma(qU) \propto \mathbf{A} \cdot \int_{qU}^{E_0} D(E; m_\nu^2, E_0) \cdot R(qU, E) dE + \mathbf{B}$$

- Fermi theory
- Molecular final states
- Doppler broadening
- Radiative corrections
- ...

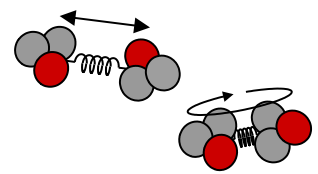


- Spectrometer resolution
- Scattering in the source
- Synchrotron radiation
- ...

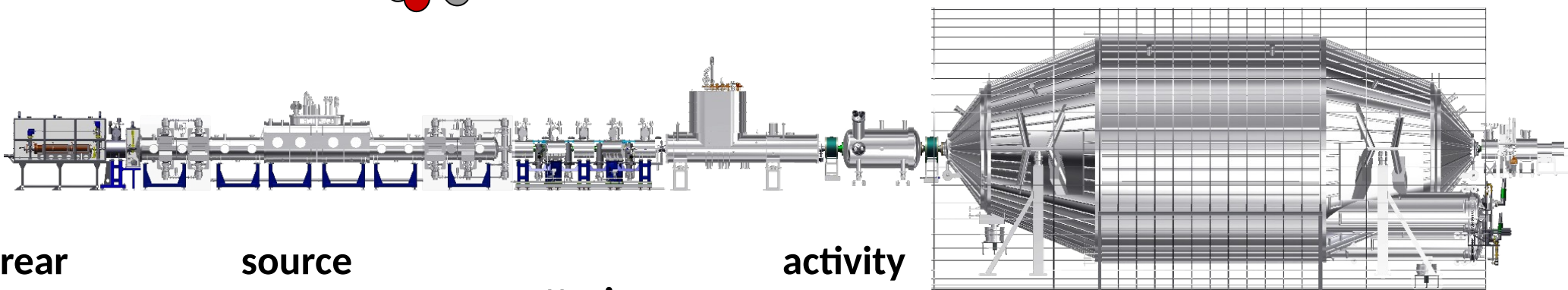
Analysis methods for the first KATRIN neutrino-mass measurement Phys. Rev. D 104, 012005 (2021)

Systematic uncertainties

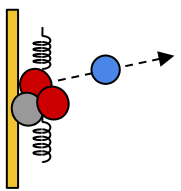
molecular final states



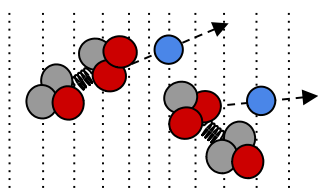
magnetic fields



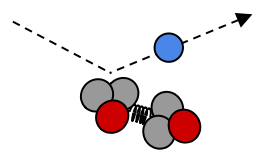
rear wall



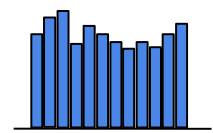
source potential



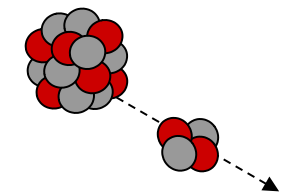
scattering



activity fluctuations



background



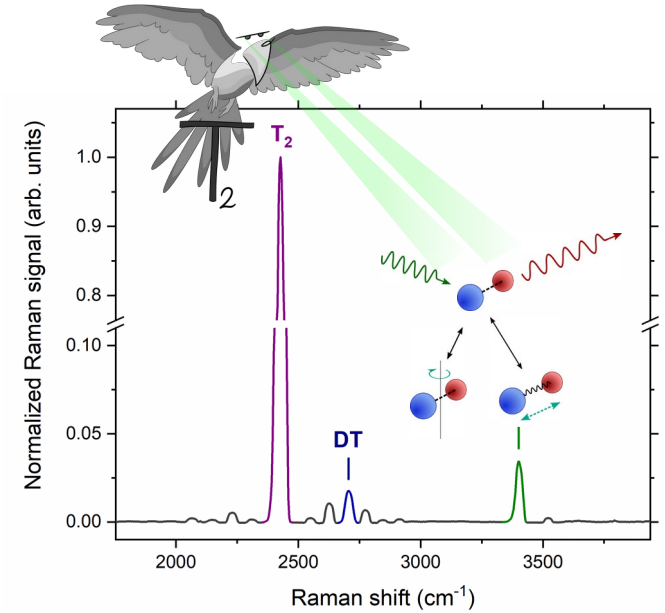
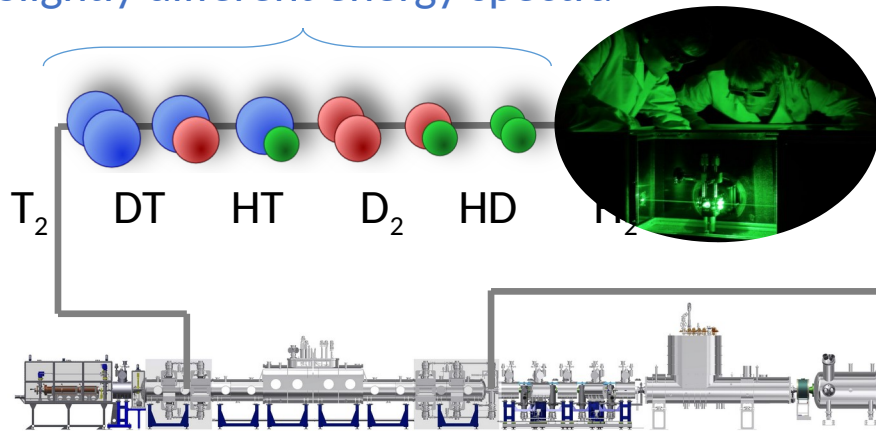
- overdispersion
- correlation
- slope

D'après Thierry Lasserre - EDSU 2022

Source composition

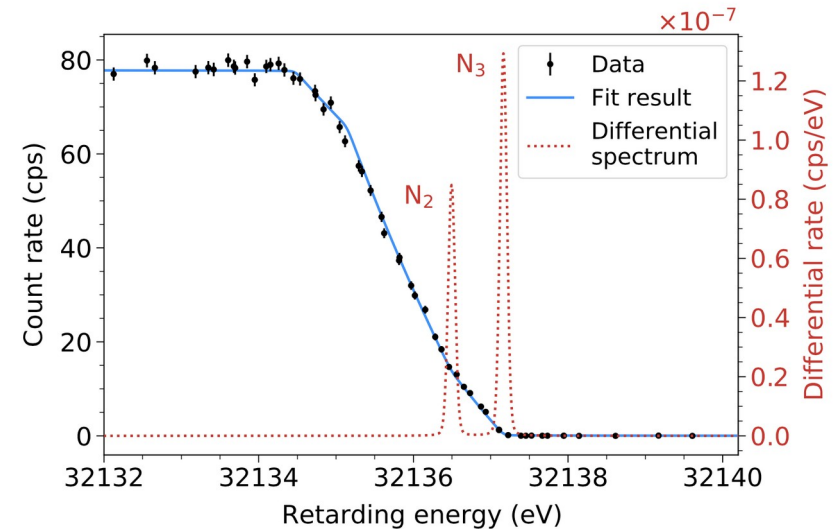
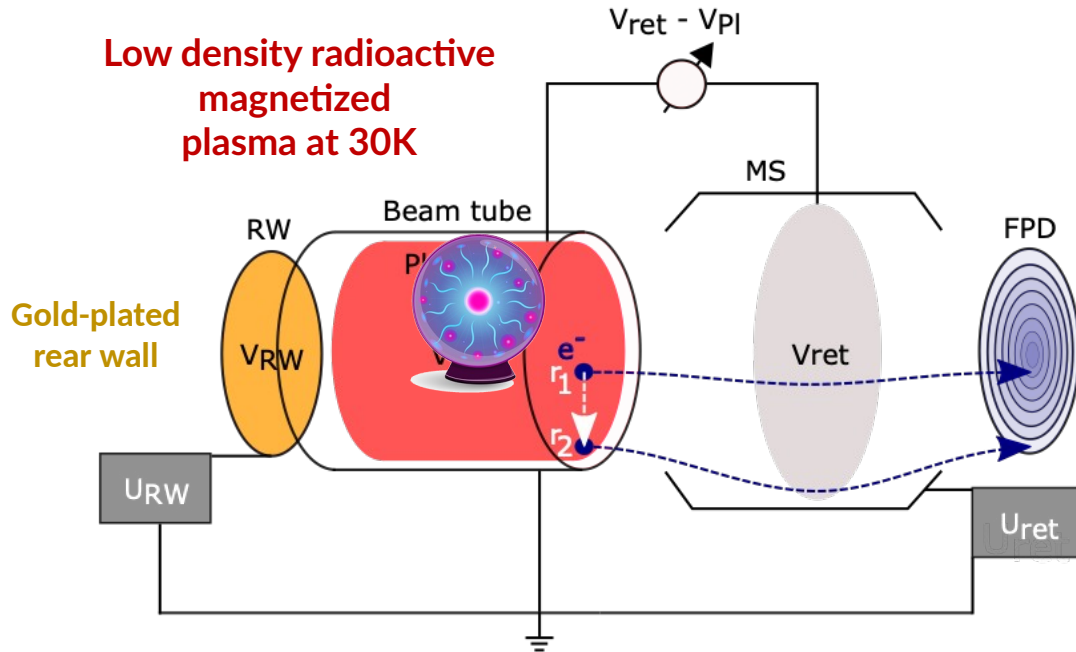
- Laser Raman IR Spectroscopy
- High purity and stability established (98.7 %)

Slightly different energy spectra



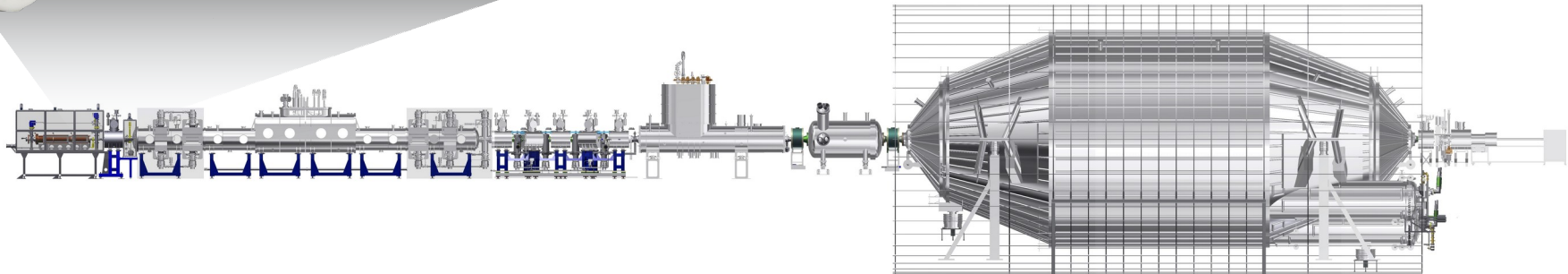
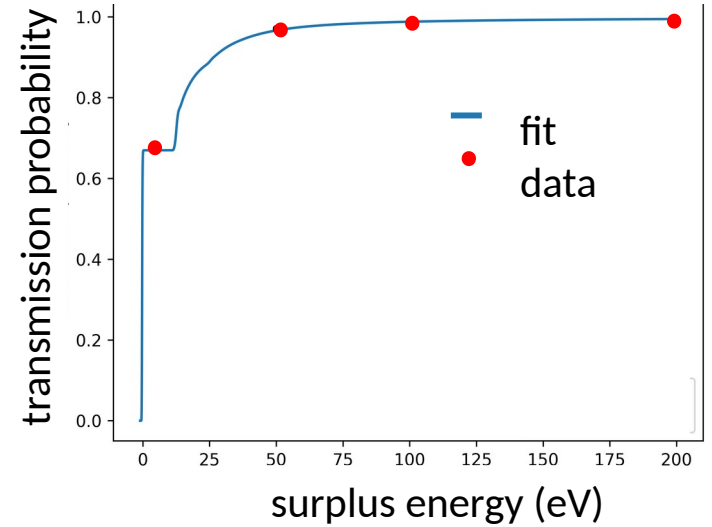
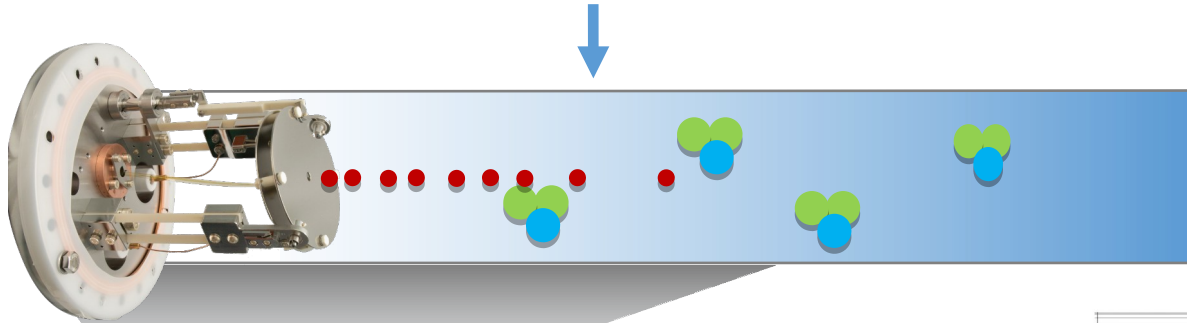
Source potential

- Spectroscopy through E threshold $\rightarrow qU_{\text{spectrometer}}(V_{\text{ret}}) - qU_{\text{source}}(V_{\text{pi}})$
- Absolute value does not affect the shape (endpoint free fit parameter)
- Calibration using gaseous $^{83\text{m}}\text{Kr}$ injected in the source, Ring-wise analysis for r-dependancy

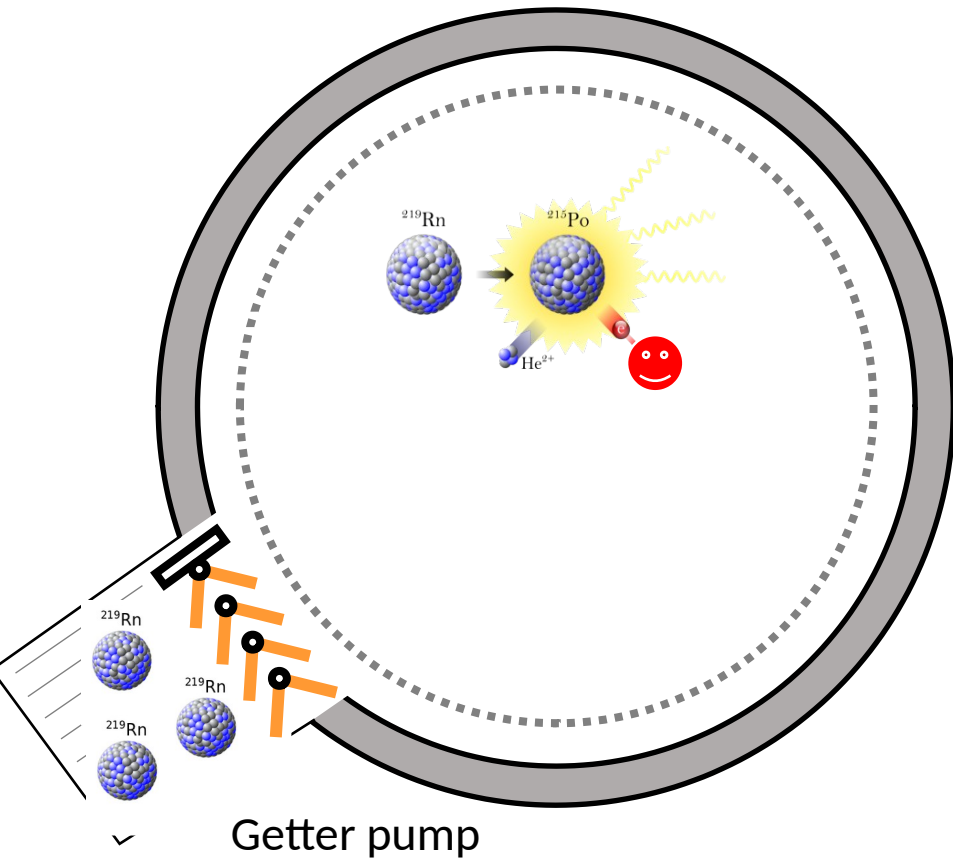



Source density

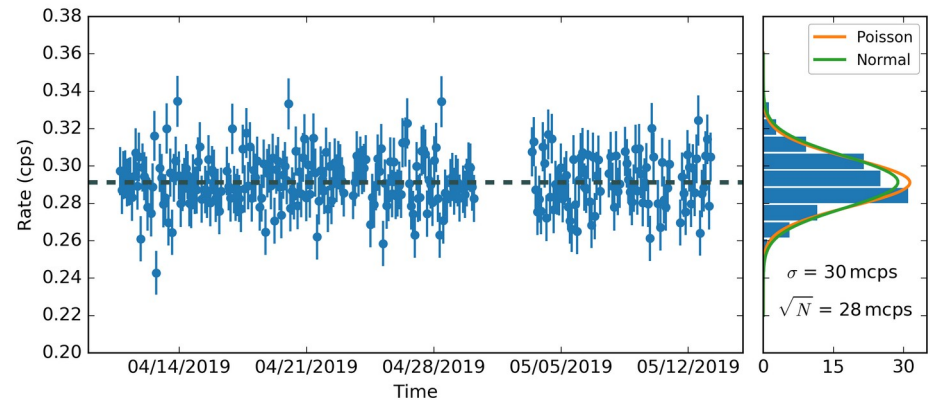
- High-intensity electron gun
- Column density $0.84 \times 1.1 \times 10^{21}$ molecules/m⁻²
- (0.1)%-ish stability of density observed
- Characterized with electron gun



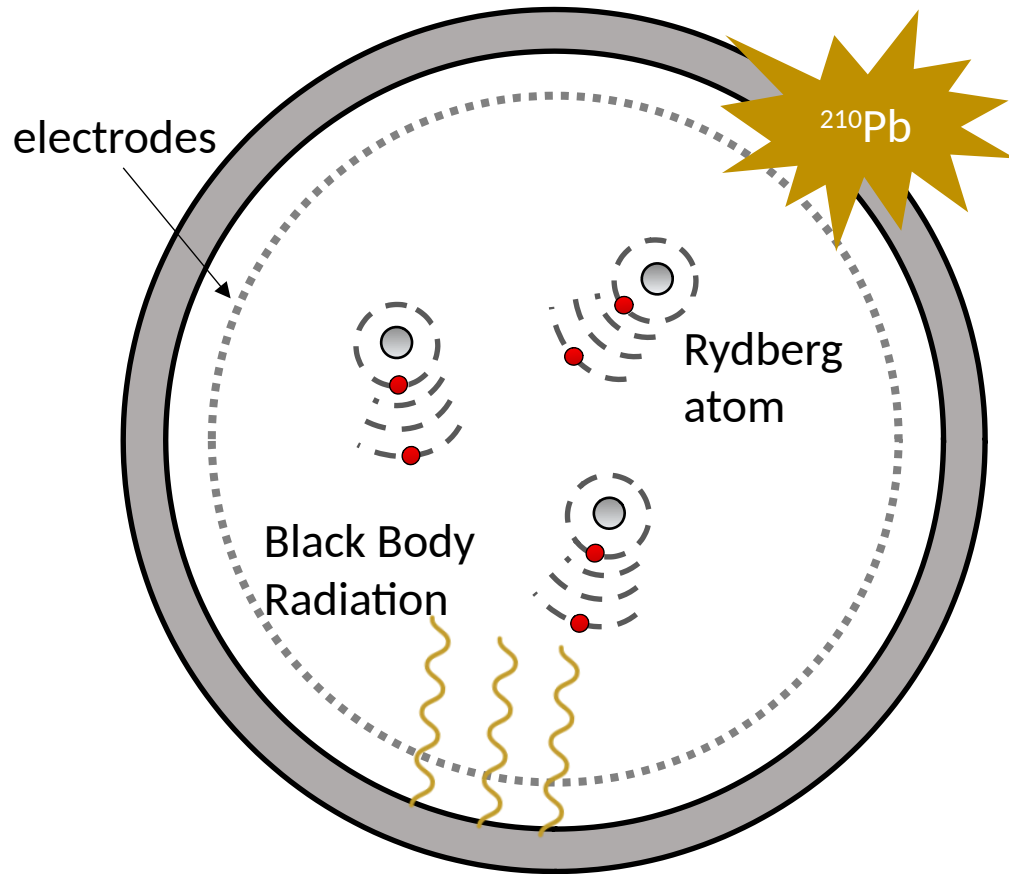
Radon-induced backgrounds



- NEG pumps radon emanation
- α -decays of single ^{219}Rn atoms (3.96 s)
- Low energy e^- emission inside spectrometer
- Effective reduction via nitrogen-cooled baffle system
- 10% Non-Poissonian rate over-dispersion 

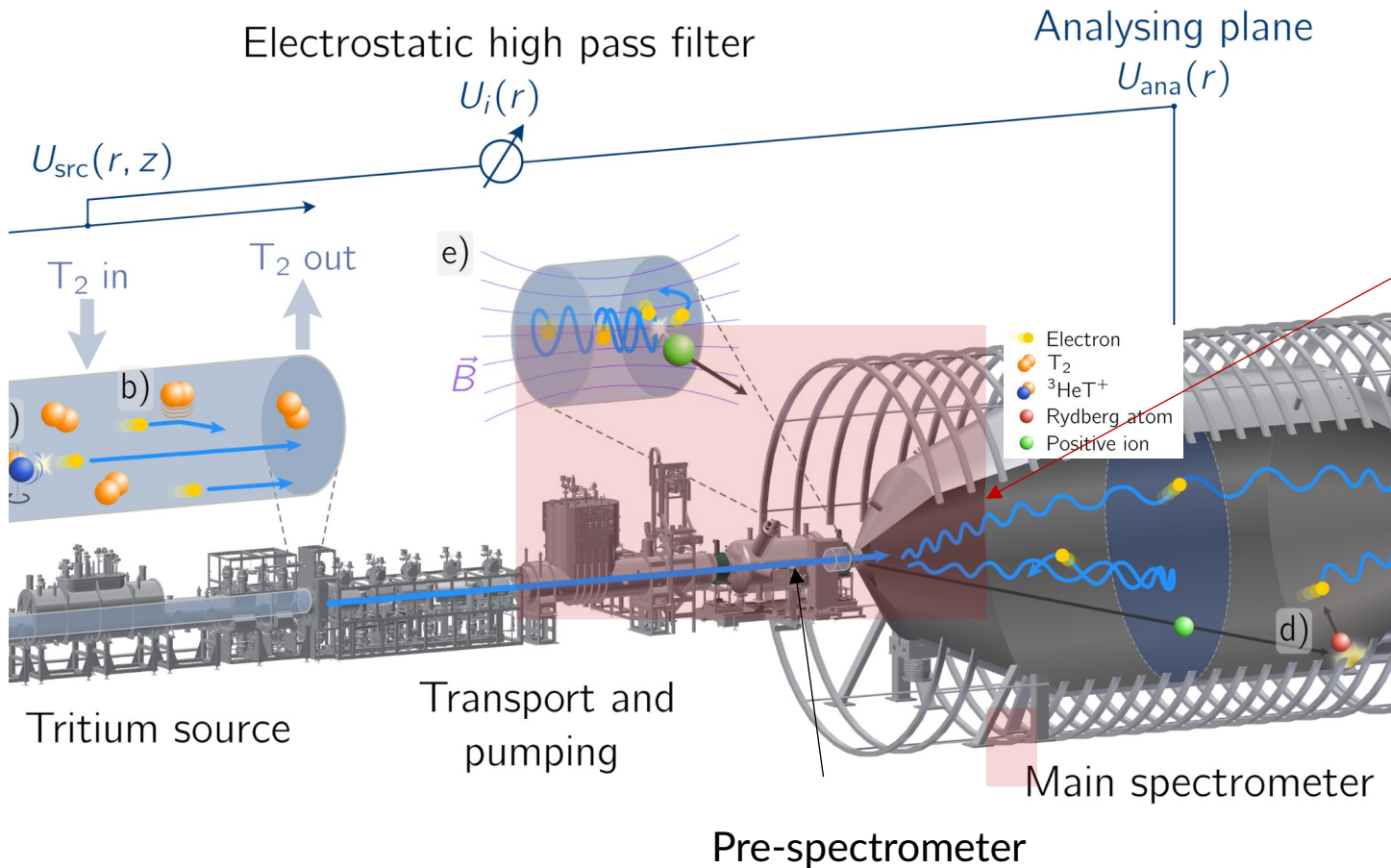


Neutral Excited Atoms



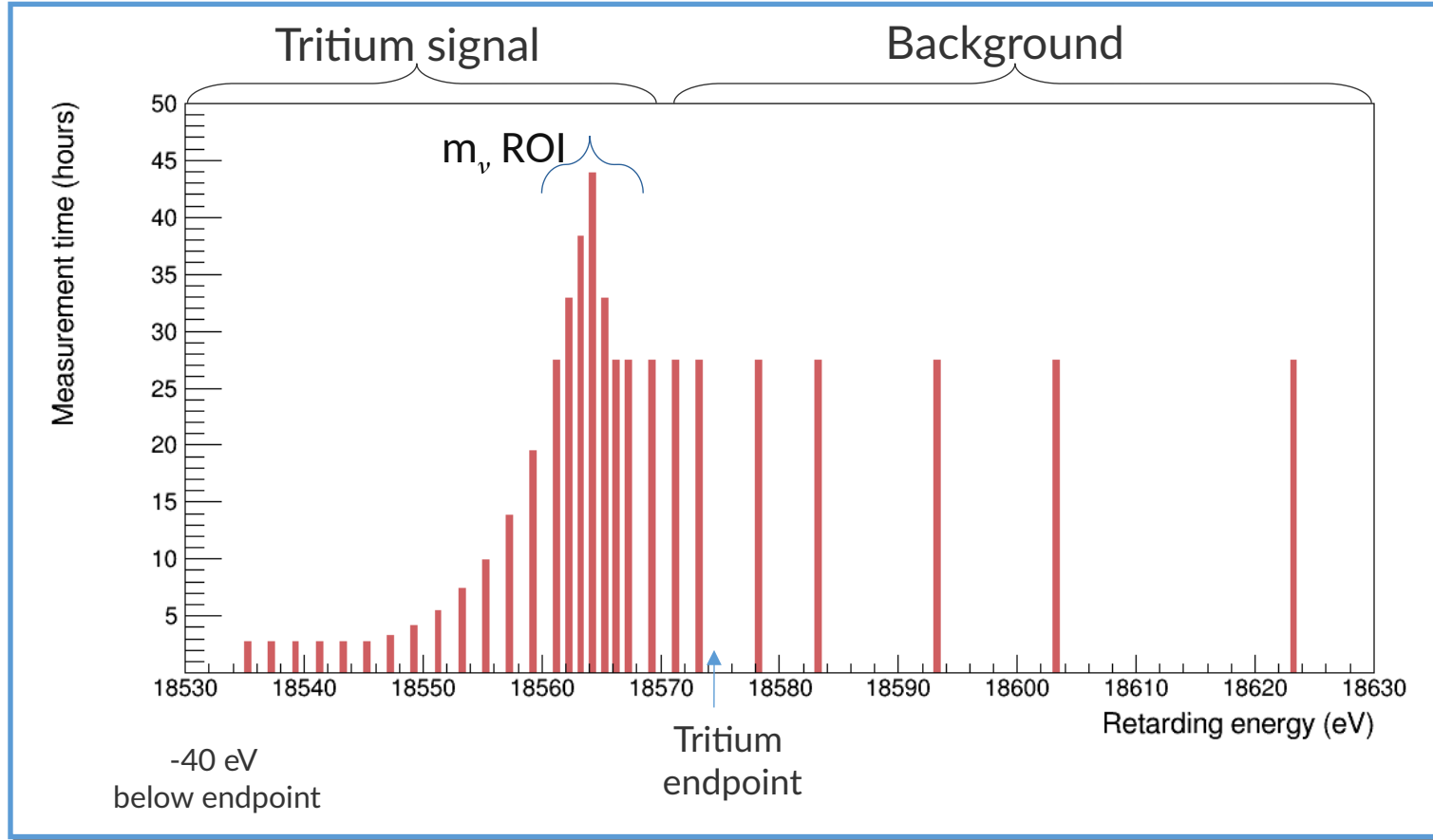
- Radon exposition during construction
→ ^{210}Pb surface contamination
- Rydberg atoms sputtered off from the spectrometer surfaces by ^{210}Pb α -decays
- Ionisation by thermal radiation
- Low energy e^- emission inside spectrometer
- Scale as the spectrometer volume...

Penning-trap induced Background



- Both pre- and main spectrometers, operated at high voltage
- create a Penning trap
- Stored electrons create ions⁺, which can escape the trap into the main spectrometer → background
- Trap emptied with an e⁻-catcher system after each sub-scan
- Can induce background dependency with sub-scan length

Scanning strategy

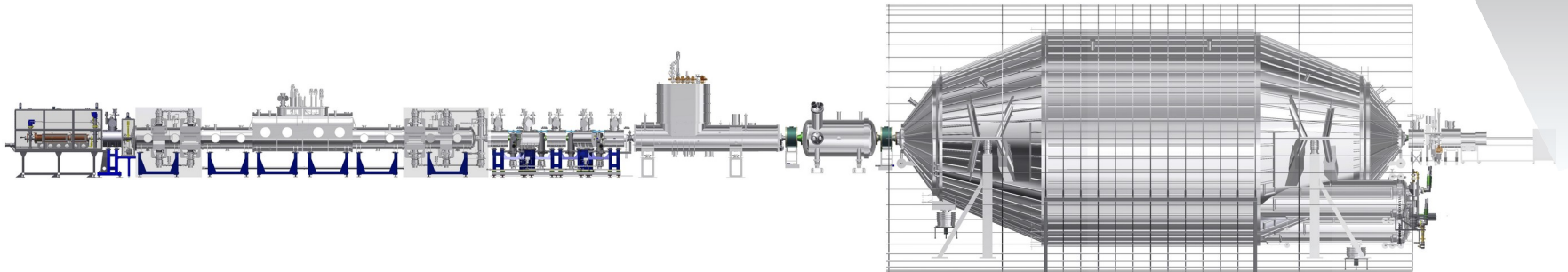
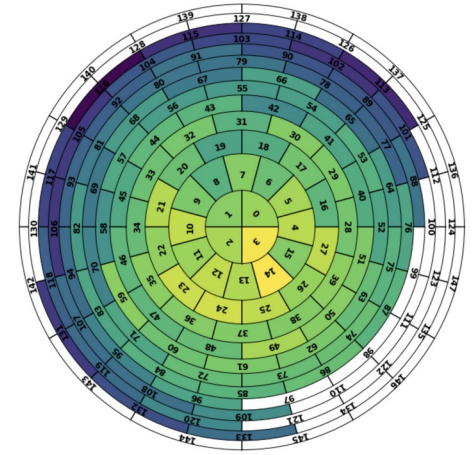


X 361 β -scans

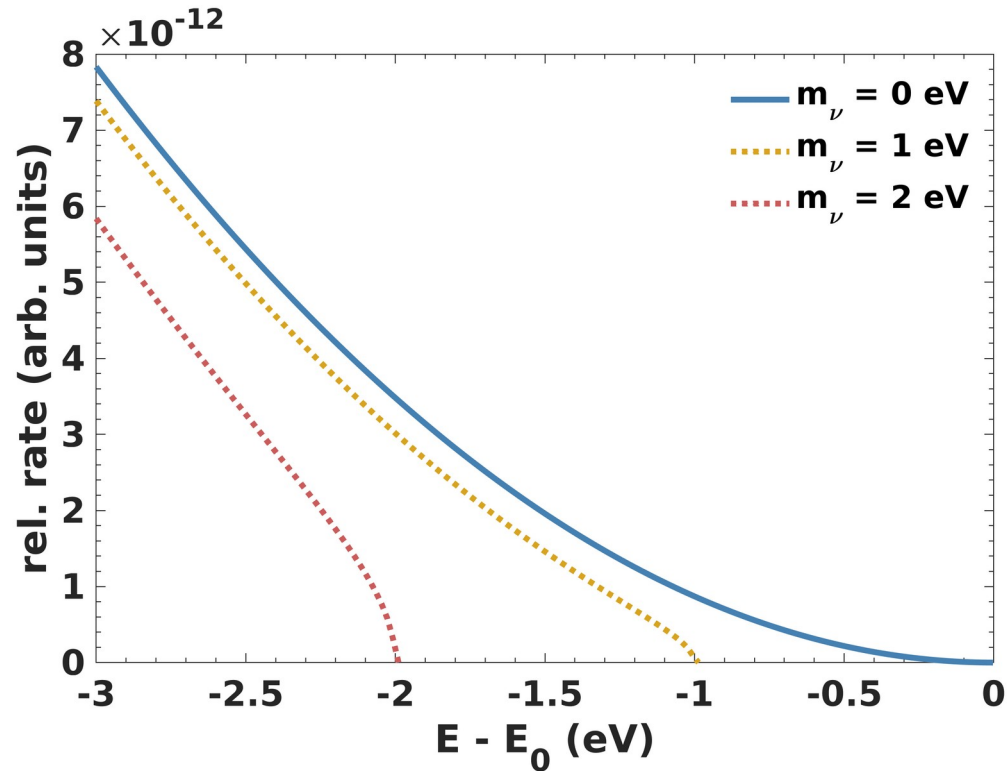
Focal plane detector

- multi-pixel silicon array
- 117/148 (79%) of all pixels used
- detection efficiency of 90%
- negligible retarding-potential dependence of efficiency
- negligible intrinsic background (~ 1 mcps)

➤ One β -decay spectrum for each pixel

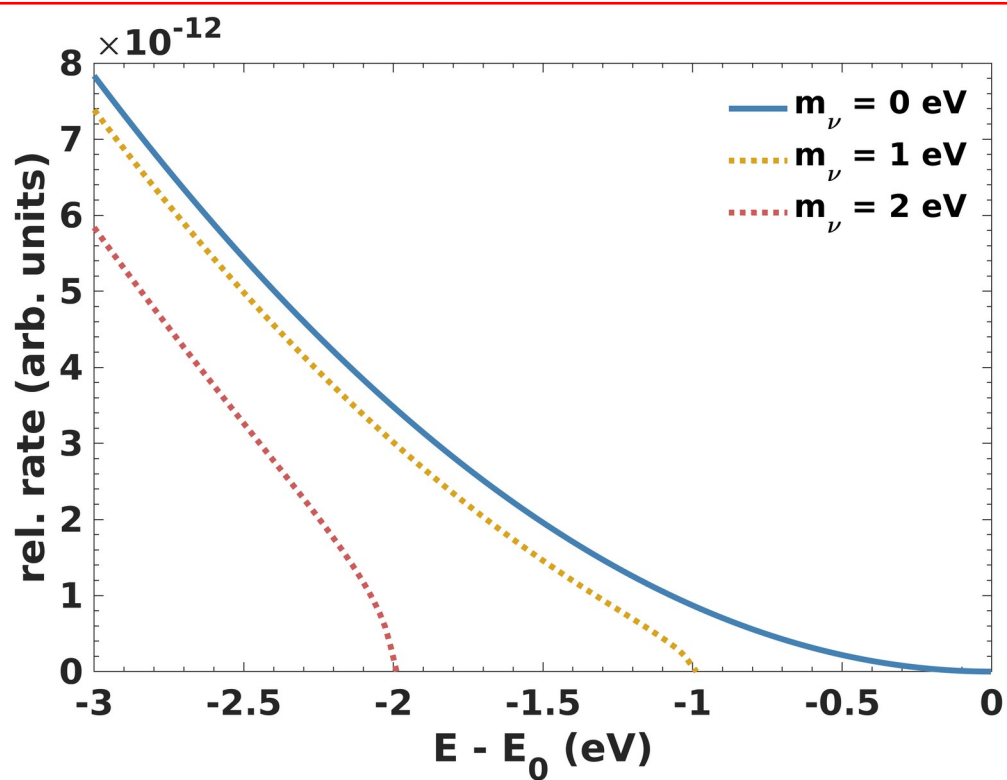


Fitting the mass imprint



Courtesy of Thierry Lasserre

Fitting the mass imprint



Courtesy of Thierry Lasserre

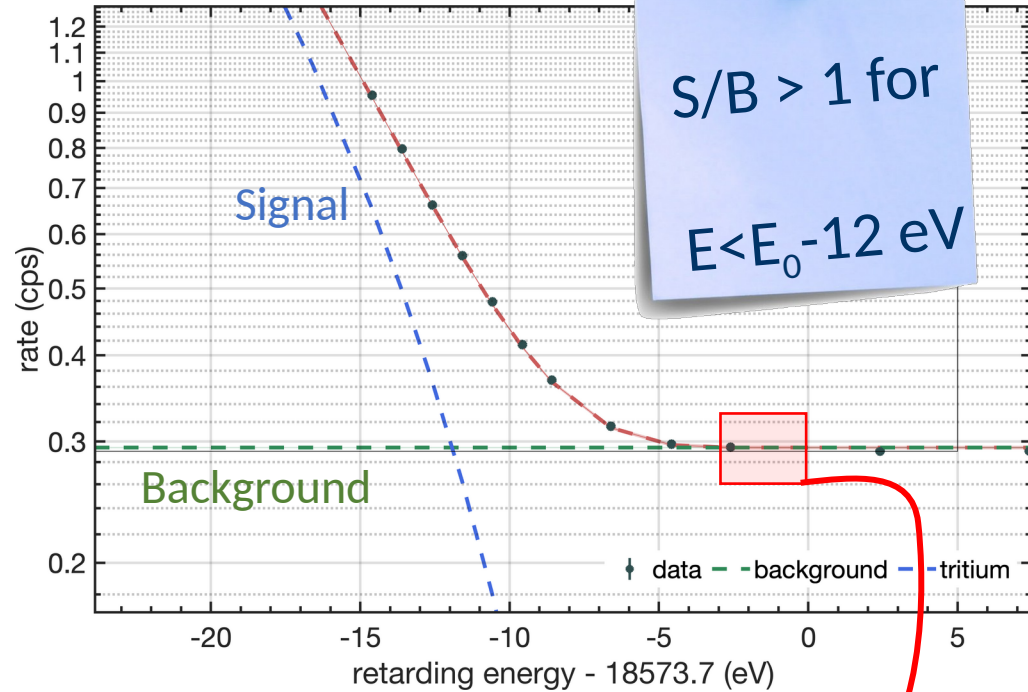
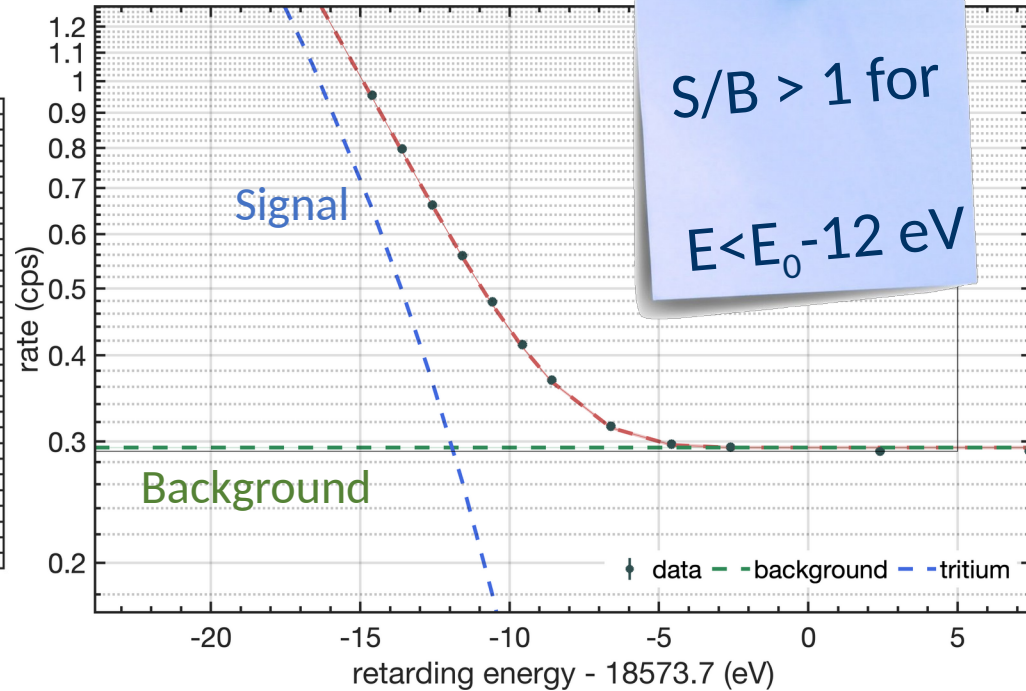
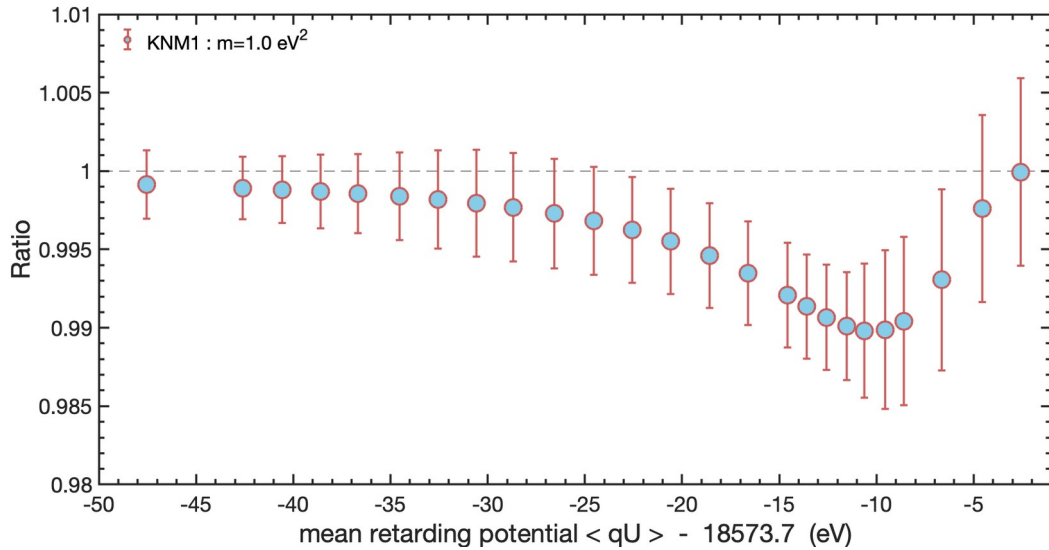


Illustration using 1st neutrino mass campaign

Fitting the mass imprint



Courtesy of Thierry Lasserre

*Illustration using 1st
neutrino mass
campaign*

Fitting the mass imprint

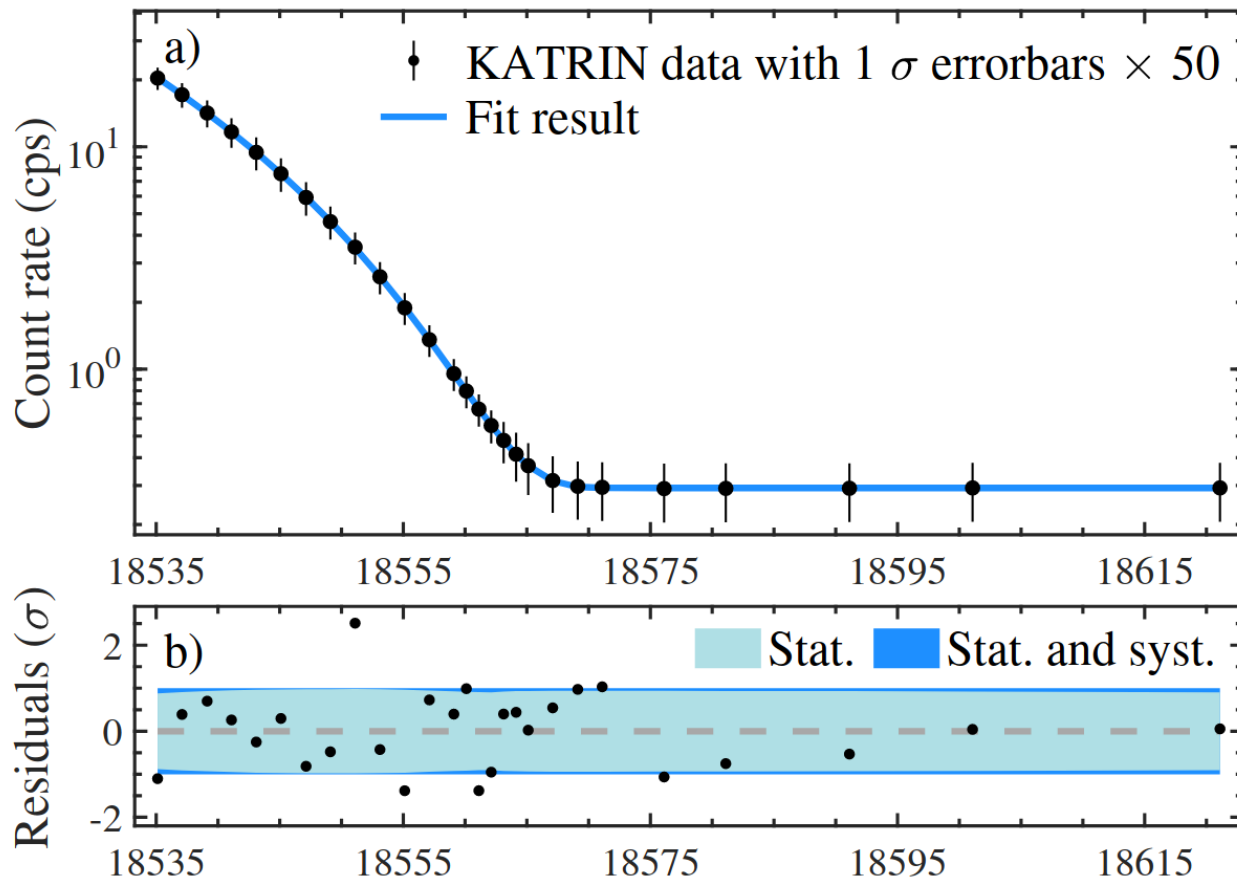
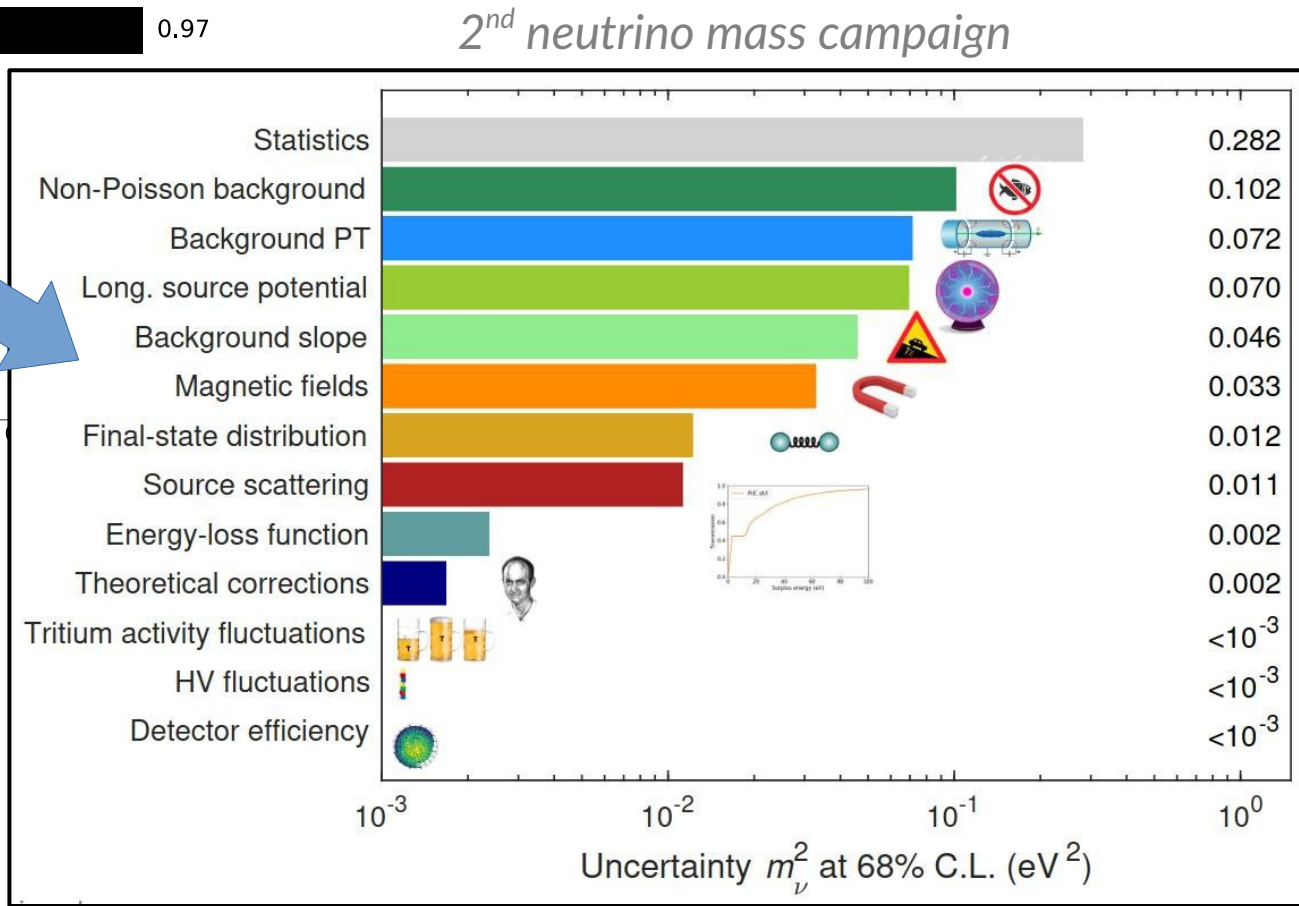
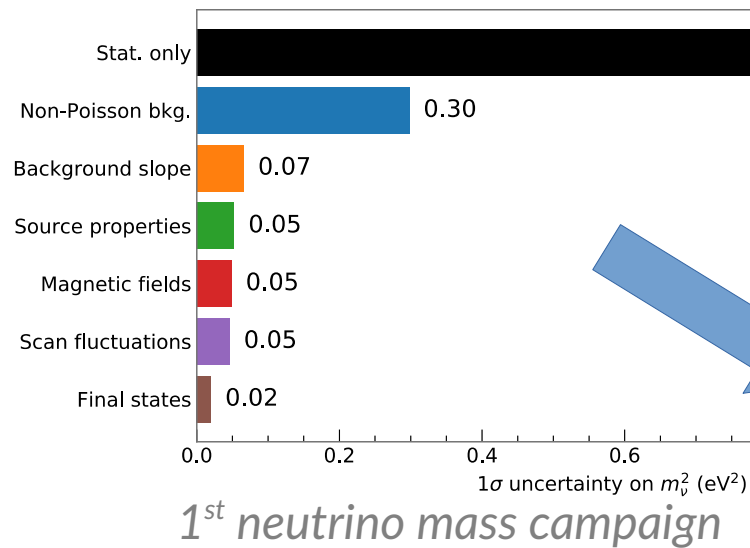
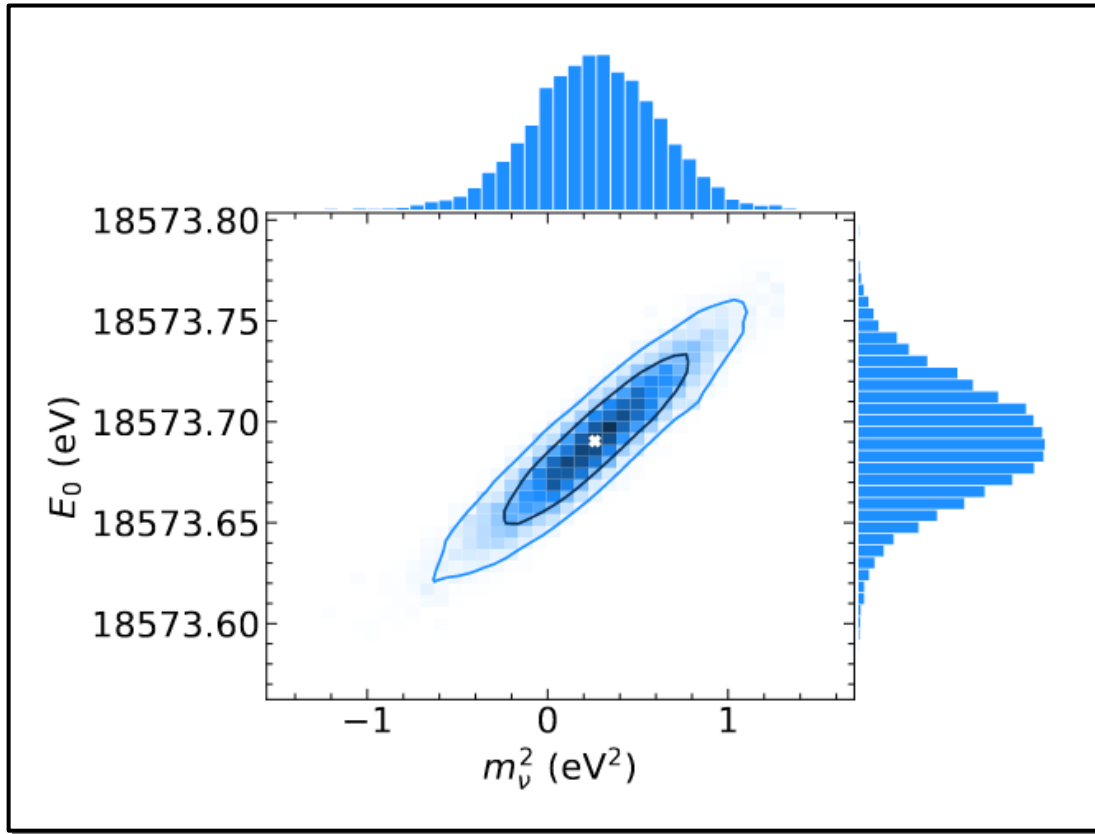


Illustration using 1st neutrino mass campaign

Systematics



Neutrino mass results



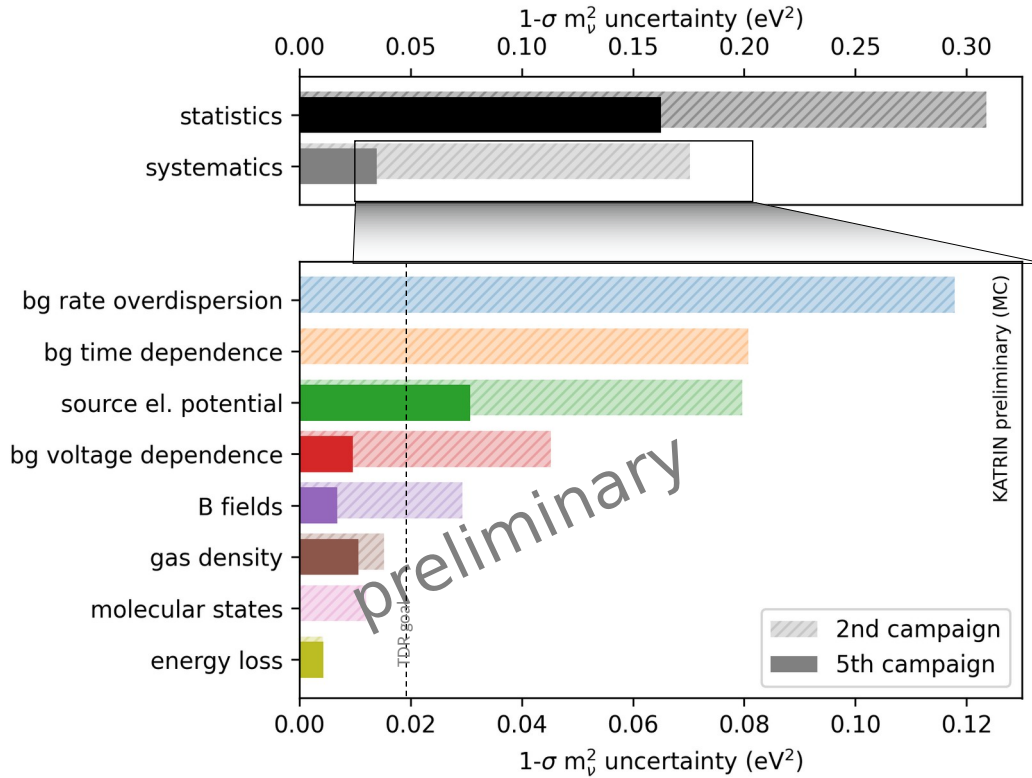
Gas density : 84 %
Sensitivity : $m_\nu < 0.7 \text{ eV (90\% CL)}$
Electrons in ROI : $4.2 \cdot 10^6$
Scan time : 31 days

$$\begin{aligned}\sigma_{\text{stat}} &\sim 0.28 \text{ eV}^2 \\ \sigma_{\text{syst}} &\sim 0.15 \text{ eV}^2 \\ \sigma_{\text{tot}} &\sim 0.35 \text{ eV}^2\end{aligned}$$

Results:

$$\begin{aligned}m_\nu^2 &= 0.26 \pm 0.35 \text{ eV}^2 \\ m_\nu &< 0.9 \text{ eV (90\% CL)}\end{aligned}$$

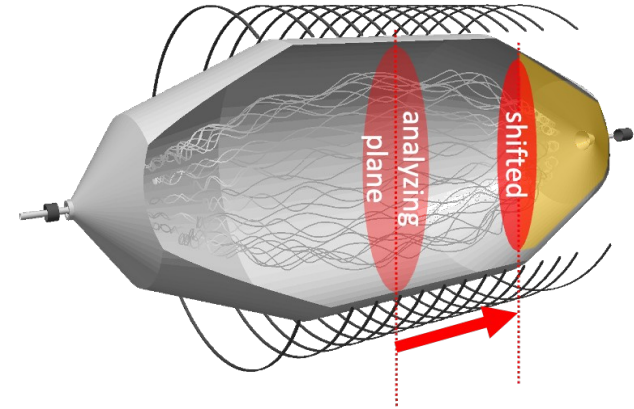
Recent achievements



D'après Thierry Lasserre - Moriond EW 2023

Major improvements:

- ✓ background reduction (2) via new EM field layout A. Lokhov et al, EPJC 82, 258 (2022)



- ✓ KNM 1 2 + 3 4 5 (2019 - 2022 data)

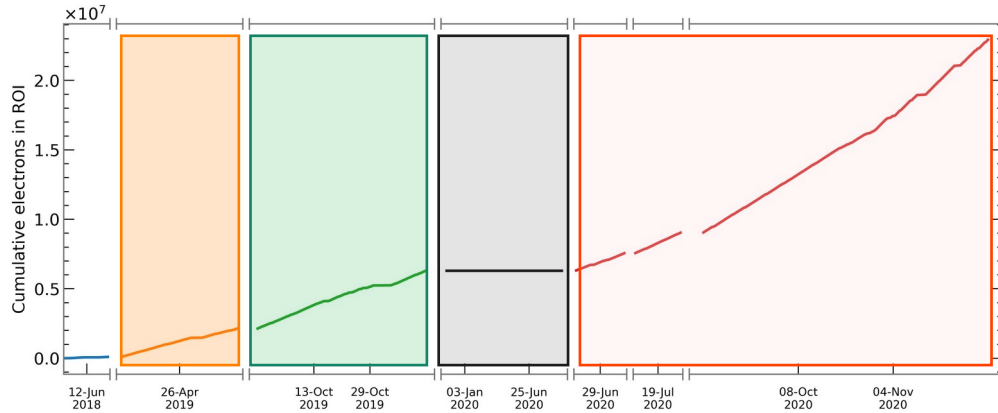
- ✓ 30 millions of electrons in ROI
- ✓ <0.5 eV sensitivity

- ✓ **Next results in summer 2023**



Conclusion

Neutrino mass with KATRIN



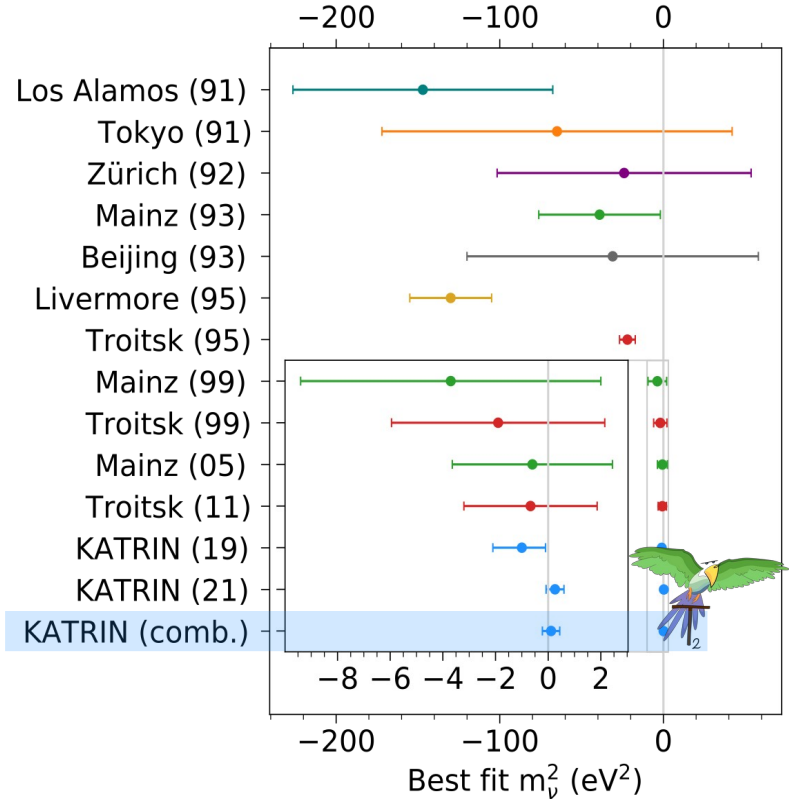
1st and 2nd campaign combined limit:

$$m_\nu < 0.8 \text{ eV (90\% CL)}$$

In Nature Physics 18, 160-166 (2022)

Future:

- Reduced background and systematics
- **New results coming soon (stats x6)!**
Stay tune for a follow up!



How to generate mass from the SM ?

- In the SM, fermions masses = Yukawa coupling between RH-LH and Higgs field
- No RH neutrinos exist in SM → neutrino massless
- To generate mass : need of RH partners
- A lot of different models – all must extend SM

$$- \mathcal{L}_{M_\nu} = M_{Dij} \bar{\nu}_{si} \nu_{Lj} + \frac{1}{2} M_{Nij} \bar{\nu}_{si} \nu_{sj}^c + \text{h.c.}$$

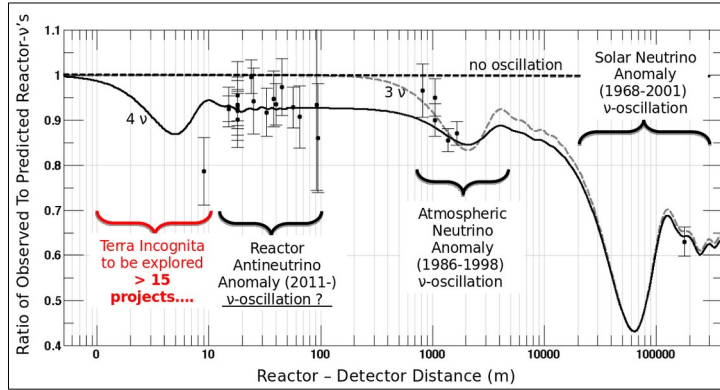
Dirac
mass term

Majorana
mass term

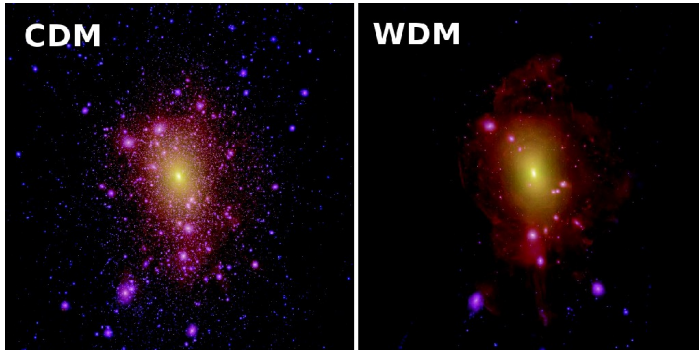
RH neutrinos → Sterile neutrinos

| | | |
|---|---|--|
| $\frac{2}{3}$ Left u up Right | $\frac{2}{3}$ Left c charm Right | $\frac{2}{3}$ Left t top Right |
| $-\frac{1}{3}$ Left d down Right | $-\frac{1}{3}$ Left s strange Right | $-\frac{1}{3}$ Left b bottom Right |
| 0 Left ν_e sterile neutrino Right | 0 Left ν_μ sterile neutrino Right | 0 Left ν_τ sterile neutrino Right |
| -1 Left e electron Right | -1 Left μ muon Right | -1 Left τ tau Right |

Is there a sterile neutrino ?



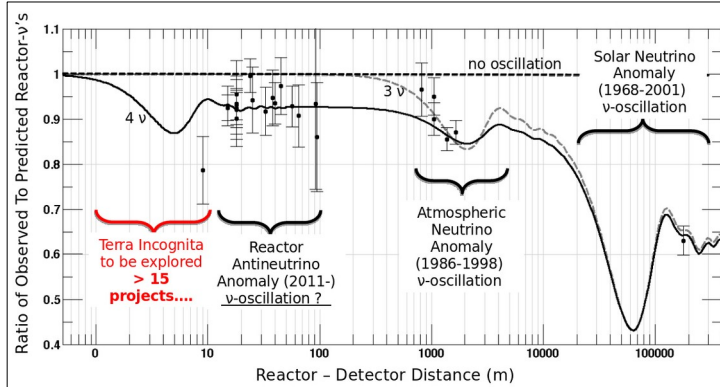
eV-scale:
Resolve anomalies in oscillation experiments



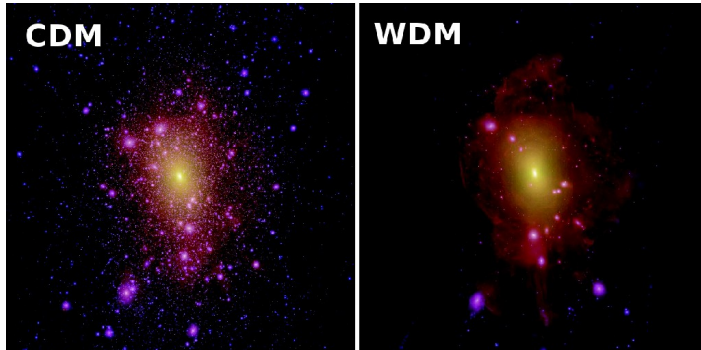
keV-scale:
Dark Matter candidate

| | | | | | |
|---------------------------------|---|----------------------------------|--|---------------------------------|--|
| 2/3 Left up Right | 2.4 MeV u | 2/3 Left charm Right | 1.27 GeV c | 2/3 Left top Right | 171.2 GeV t |
| -1/3 Left down Right | 4.8 MeV d | -1/3 Left strange Right | 104 MeV s | -1/3 Left bottom Right | 4.2 GeV b |
| < 1 eV Left ν_e | ~eV ? N₁ sterile neutrino | < 1 eV Left ν_μ | ~keV ? N₂ sterile neutrino | < 1 eV Left ν_τ | ~GeV ? N₃ sterile neutrino |
| -1 Left electron Right | 0.511 MeV e | -1 Left muon Right | 105.7 MeV μ | -1 Left tau Right | 1.777 GeV τ |

Is there a sterile neutrino ?



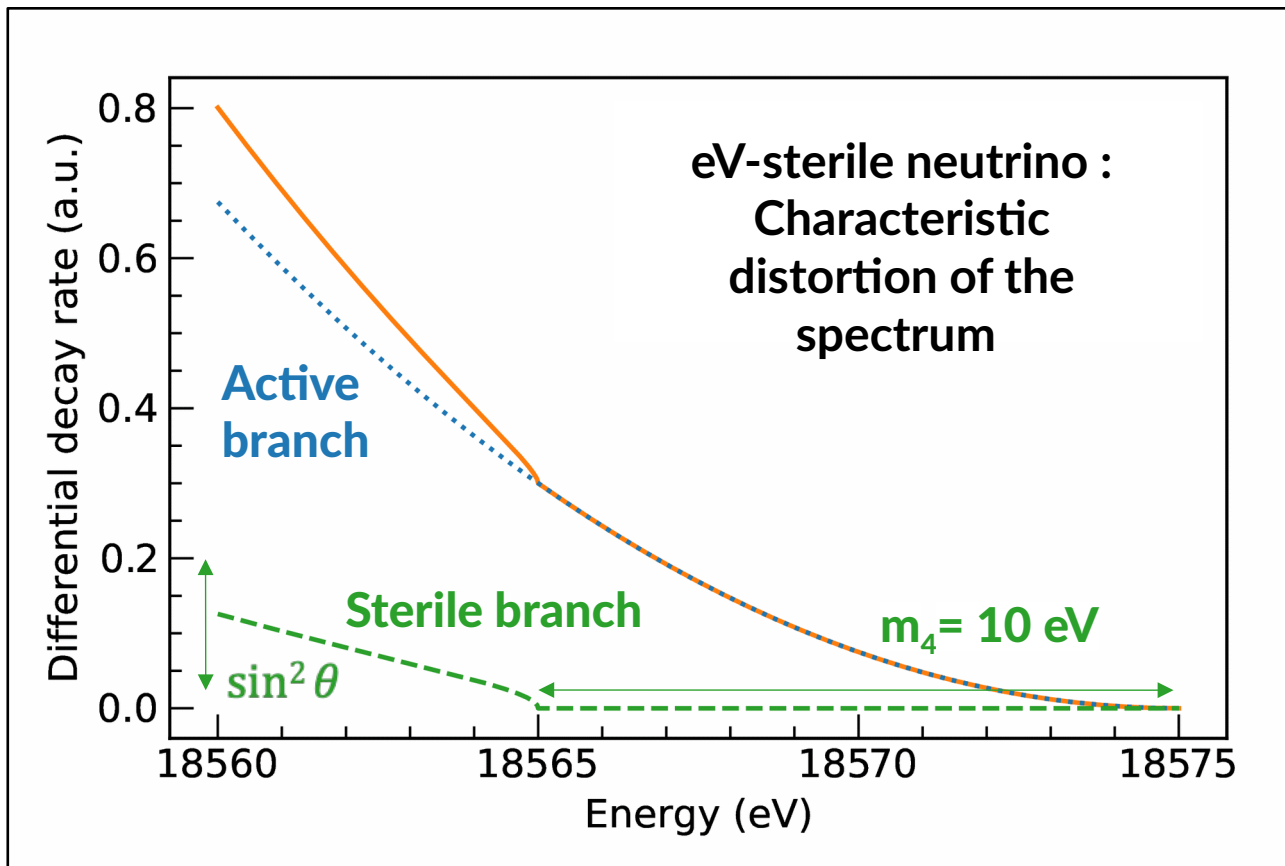
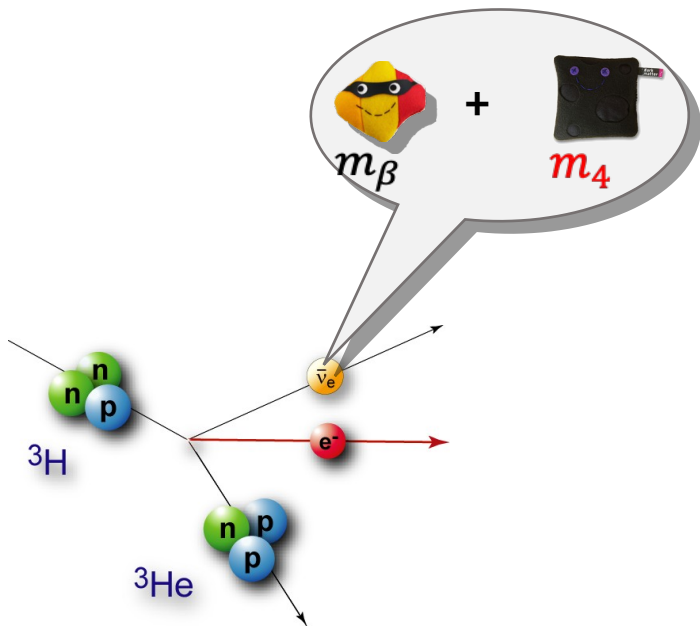
eV-scale:
Resolve anomalies in oscillation experiments



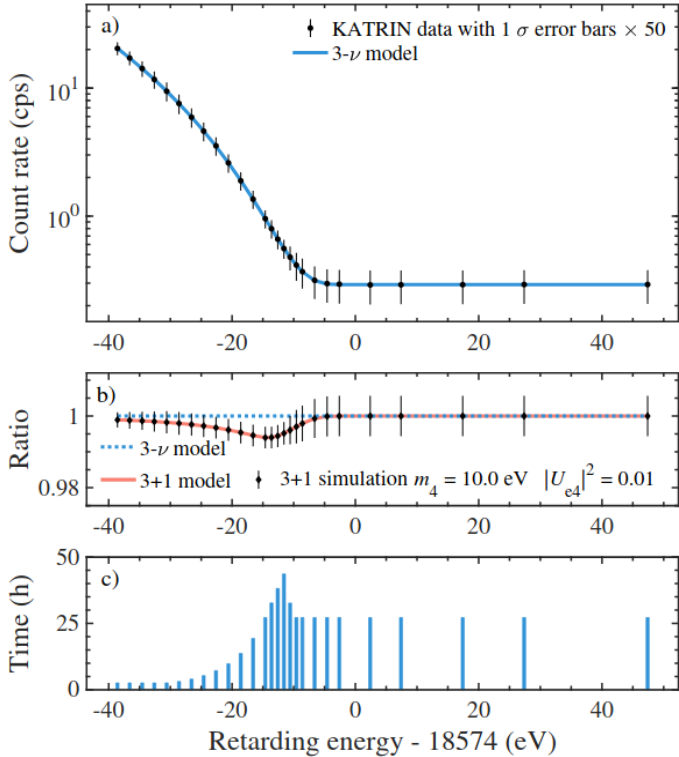
keV-scale:
Dark Matter candidate

| | | | | | |
|---------------------------------|--|----------------------------------|---|---------------------------------|--|
| 2/3 Left up Right | 2.4 MeV u | 2/3 Left charm Right | 1.27 GeV c | 2/3 Left top Right | 171.2 GeV t |
| -1/3 Left down Right | 4.8 MeV d | -1/3 Left strange Right | 104 MeV s | -1/3 Left bottom Right | 4.2 GeV b |
| 0 Left ν_e | < 1 eV N₁ sterile neutrino | 0 Left ν_μ | ~eV ? N₂ sterile neutrino | 0 Left ν_τ | ~keV ? N₃ sterile neutrino |
| -1 Left electron Right | 0.511 MeV e | -1 Left muon Right | 105.7 MeV μ | -1 Left tau Right | 1.777 GeV τ |

eV-sterile signature in β -decay

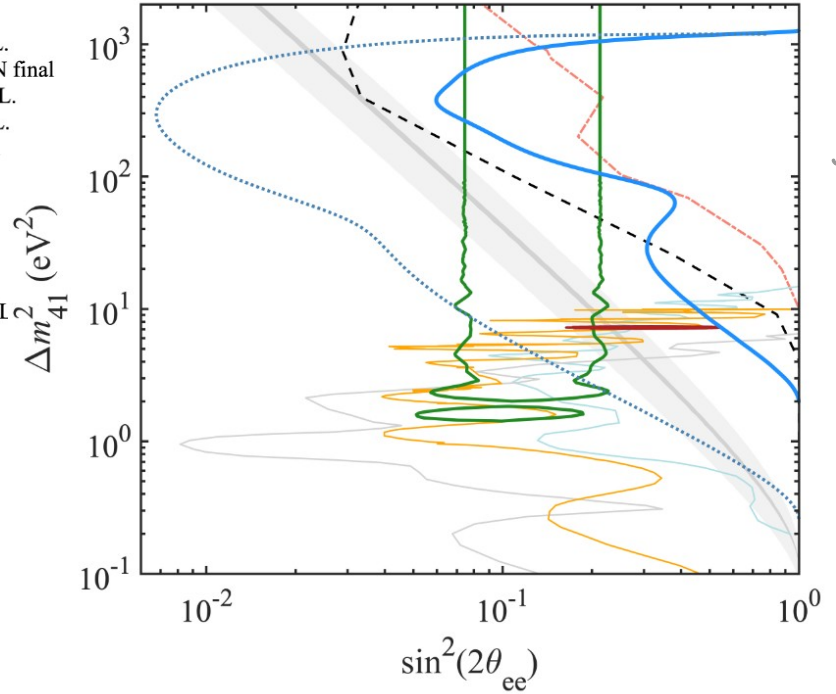


Sterile hunt with KATRIN



PRL 126, 091803

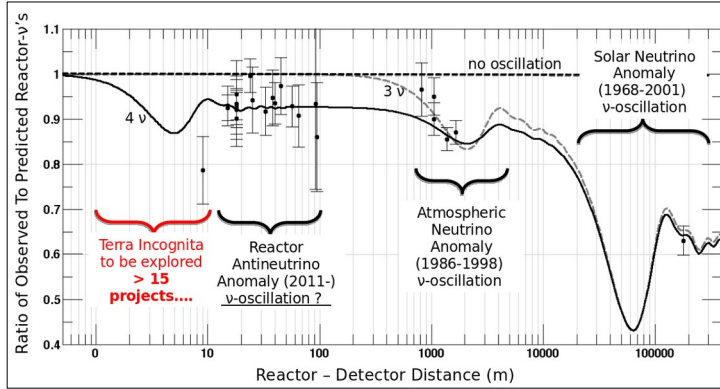
- Neutrino-4 2σ
- KATRIN 95% C.L.
- ⋯ Projected KATRIN final sensitivity 95% C.L.
- $0\nu\beta\beta$ NH 90% C.L.
- $0\nu\beta\beta$ IH 90% C.L.
- - Mainz 95% C.L.
- - Troitsk 95% C.L.
- - Prospect 95% C.L.
- - DANSS 95% C.L.
- - Stéréo 95% C.L.
- - RAA + GA 95% CI



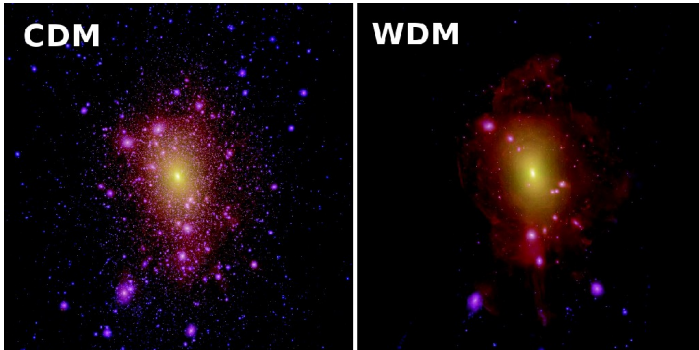
PRL 126, 091803

- Unique large window at high mass
- Complementary with Reactor experiments
- Exclude most of the favored phase-space in the next years

Is there a sterile neutrino ?



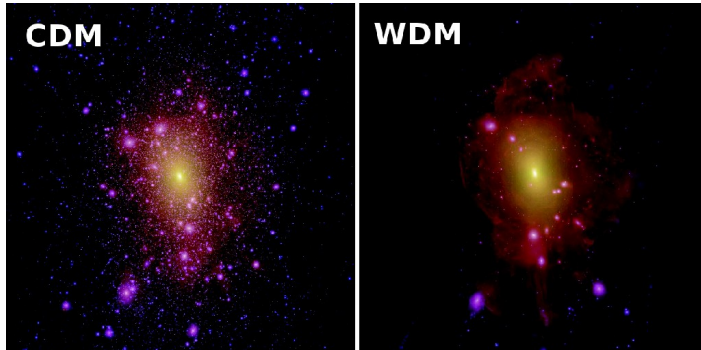
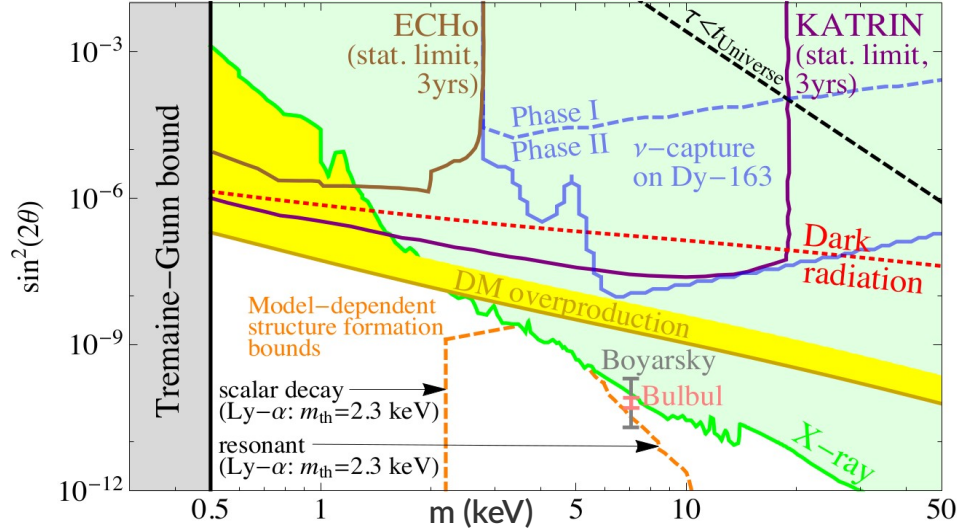
eV-scale:
Resolve anomalies in
oscillation experiments



keV-scale:
Dark Matter candidate

| | | | | | |
|--------------------------------------|--|---------------------------------------|---|--------------------------------------|---|
| 2/3 Left u up Right | 2.4 MeV | 2/3 Left c charm Right | 1.27 GeV | 2/3 Left t top Right | 171.2 GeV |
| -1/3 Left d down Right | 4.8 MeV | -1/3 Left s strange Right | 104 MeV | -1/3 Left b bottom Right | 4.2 GeV |
| < 1 eV Left ν _e | ~eV ? N ₁ sterile neutrino | < 1 eV Left ν _μ | ~keV ? N ₂ sterile neutrino | < 1 eV Left ν _τ | ~GeV ? N ₃ sterile neutrino |
| -1 Left e electron Right | 0.511 MeV | -1 Left μ muon Right | 105.7 MeV | -1 Left τ tau Right | 1.777 GeV |

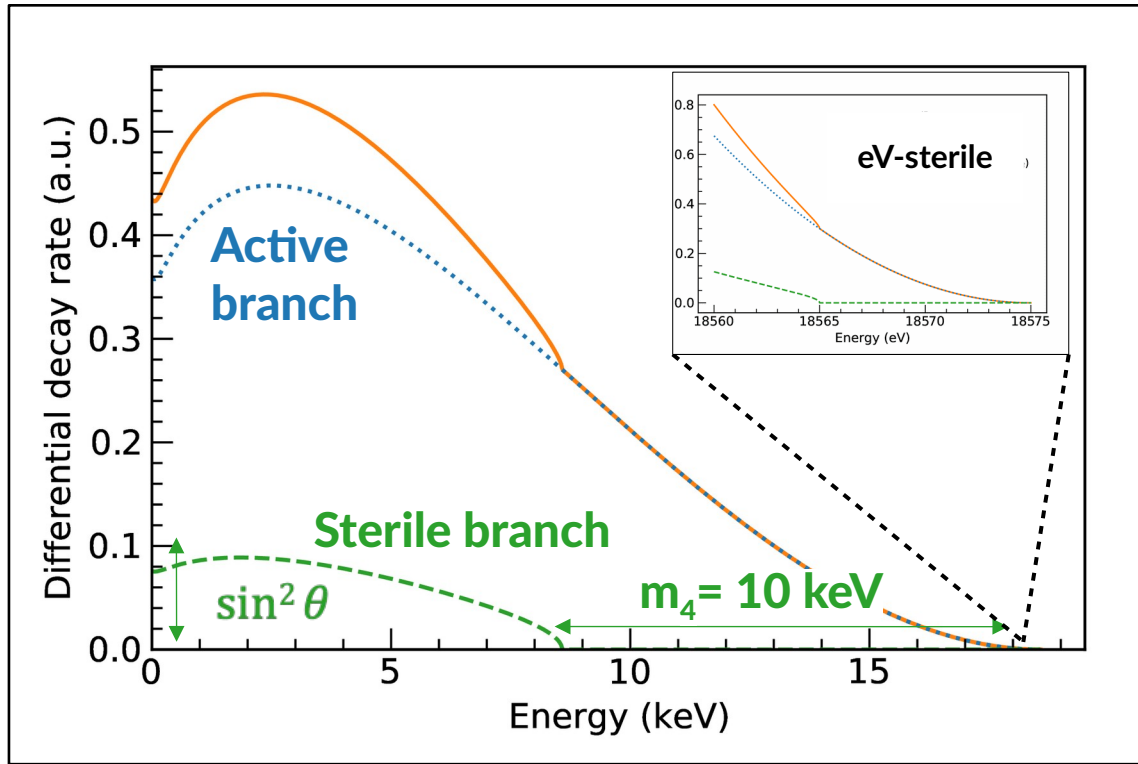
Is there a sterile neutrino ?



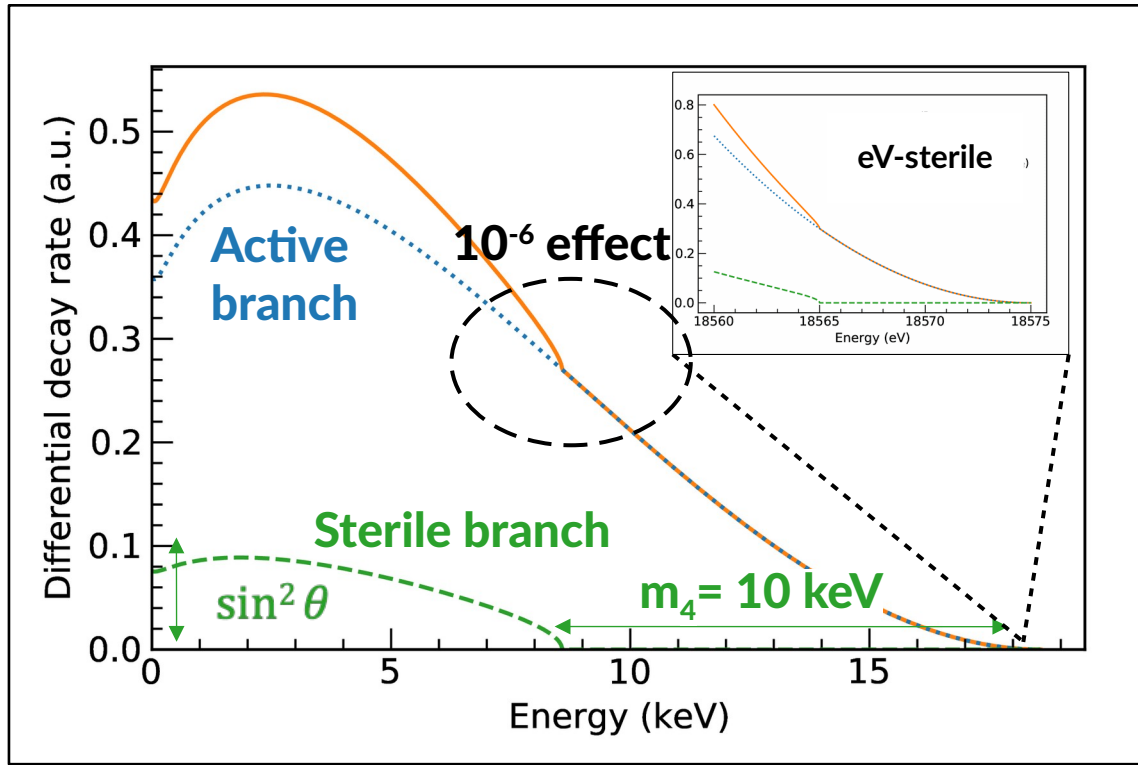
keV-scale:
Dark Matter candidate

| | | | | | |
|--------------------------------------|---|---------------------------------------|--|--------------------------------------|--|
| 2/3 Left u up Right | 2.4 MeV | 2/3 Left c charm Right | 1.27 GeV | 2/3 Left t top Right | 171.2 GeV |
| -1/3 Left d down Right | 4.8 MeV | -1/3 Left s strange Right | 104 MeV | -1/3 Left b bottom Right | 4.2 GeV |
| < 1 eV Left ν _e | ~eV ? N ₁ sterile neutrino | < 1 eV Left ν _μ | ~keV ? N ₂ sterile neutrino | < 1 eV Left ν _τ | ~GeV ? N ₃ sterile neutrino |
| -1 Left e electron Right | 0.511 MeV | -1 Left μ muon Right | 105.7 MeV | -1 Left τ tau Right | 1.777 GeV |

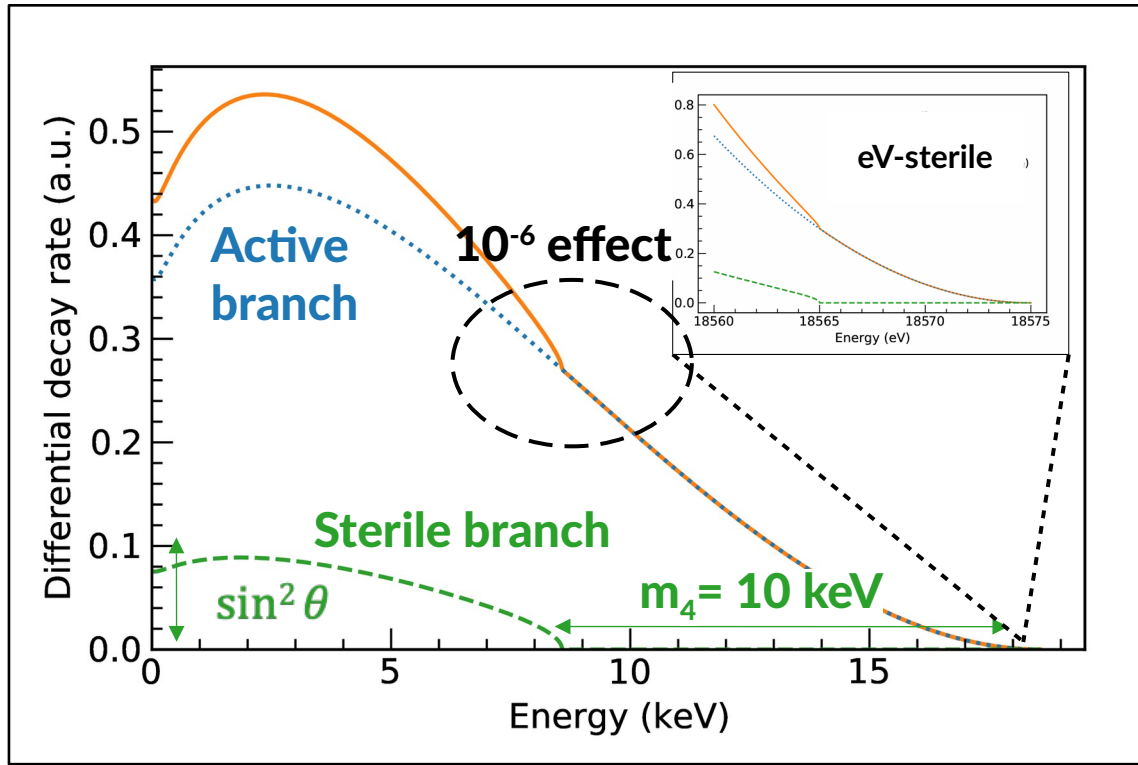
keV-sterile signature in β -decay



keV-sterile signature in β -decay



keV-sterile signature in β -decay



Stringent limit from **astrophysical** and **cosmological** observations ($\sin^2(\theta) < 10^{-7}$):

→ Dramatic **increase of the count rate** (up to $3 \times 10^8 \text{ Hz}$)

→ Integral and differential phases (detector with **good resolution**)

→ Highly **pixelised**

→ **new detector is needed: the TRISTAN project**

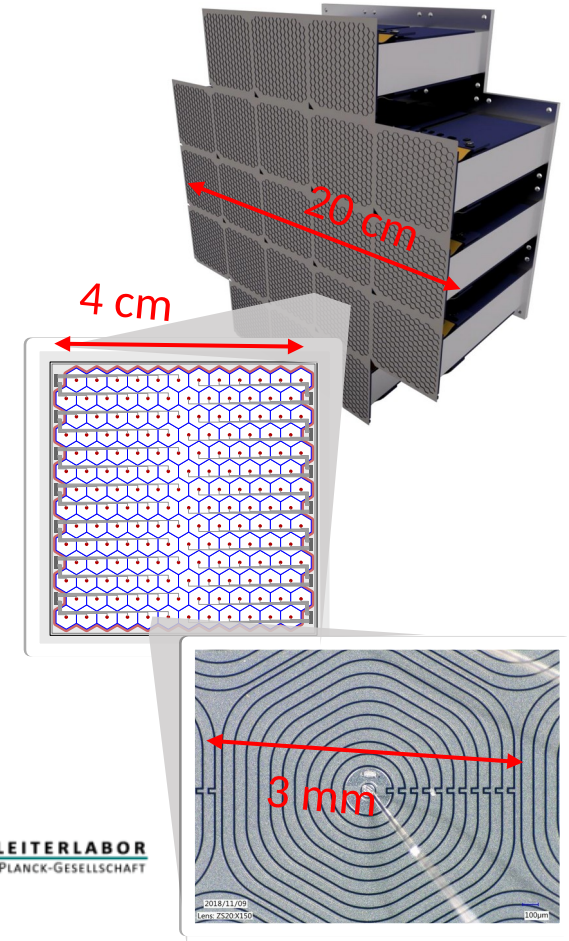
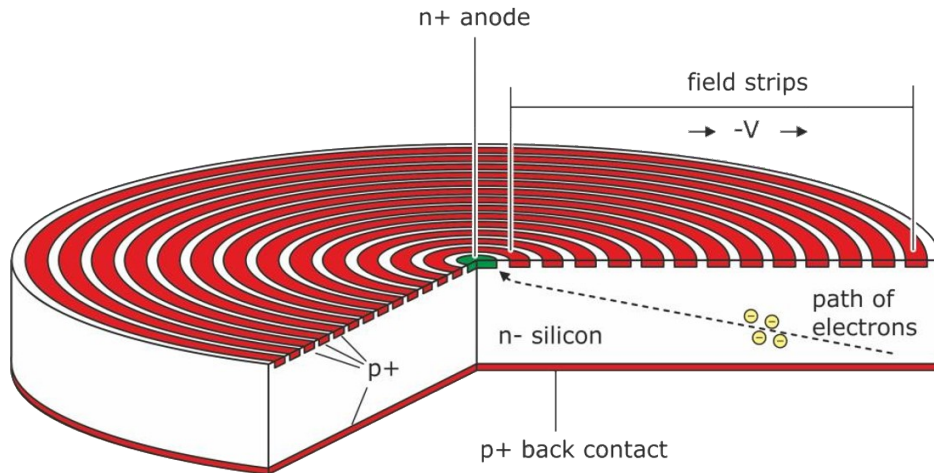


PI : S. Mertens

TRISTAN Project

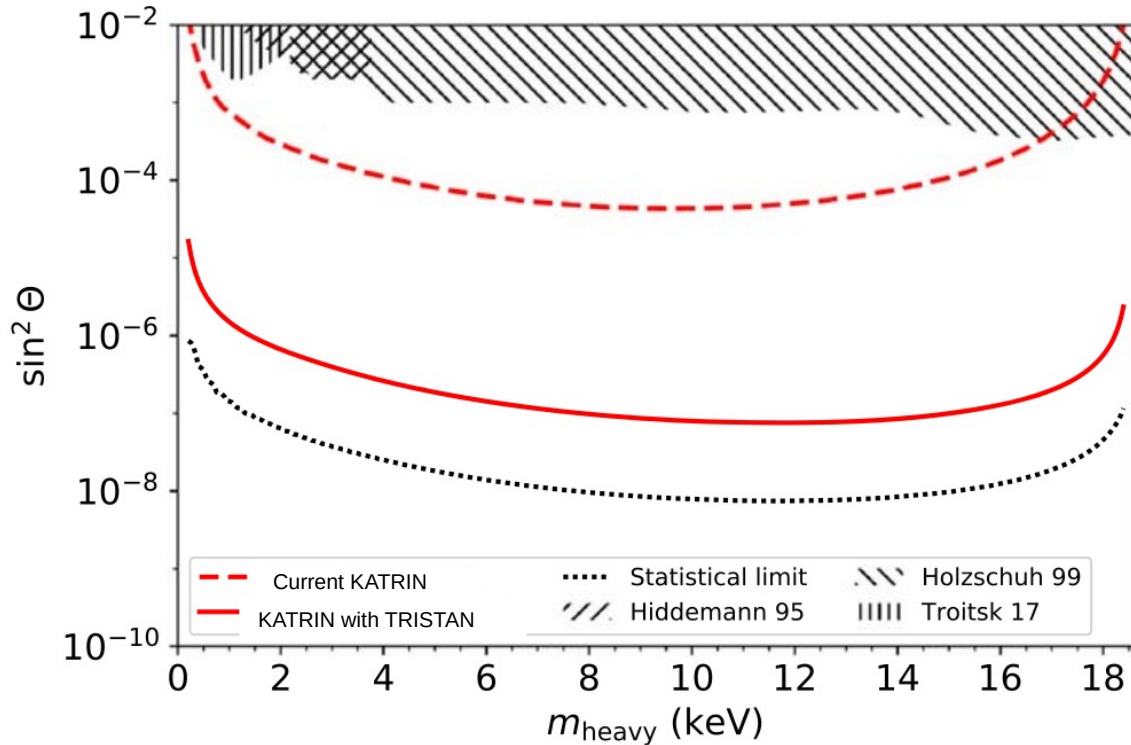
Capability of handling high rates ($> 3 \times 10^8$ cps)
+ Excellent energy resolution (300 eV @ 20 keV)

- Silicon Drift Detector (SDD) Technology
- Novelty: large number of pixels (about 3500)
- Novelty: application to high-precision β -spectroscopy

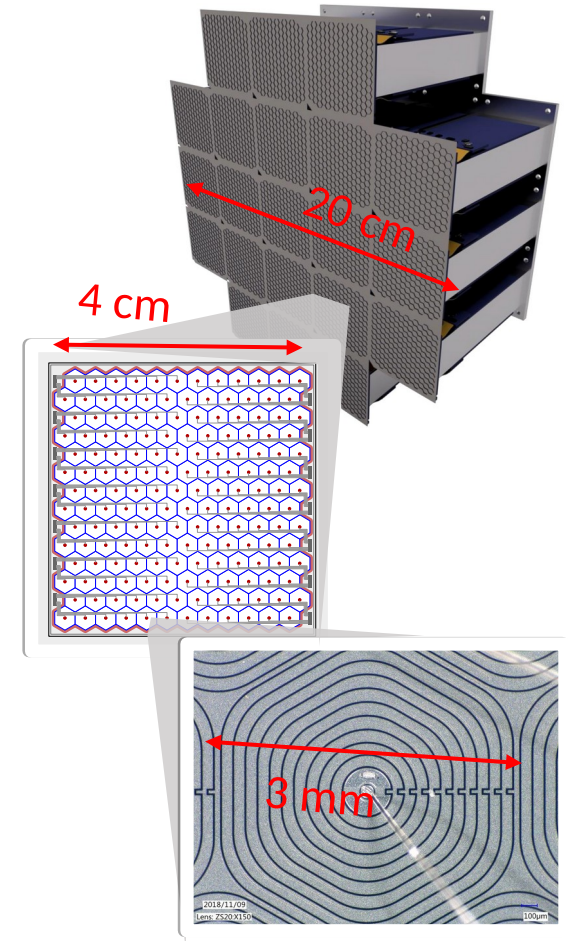


TRISTAN Project

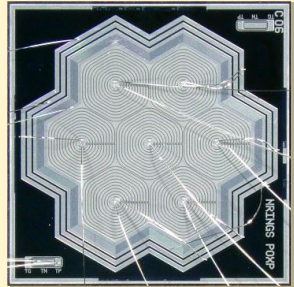
TRISTAN : Development of a large area SDD array and read-out system to look for keV-sterile neutrino with the KATRIN experiment



J.Phys. G46 (2019) no.6, 065203

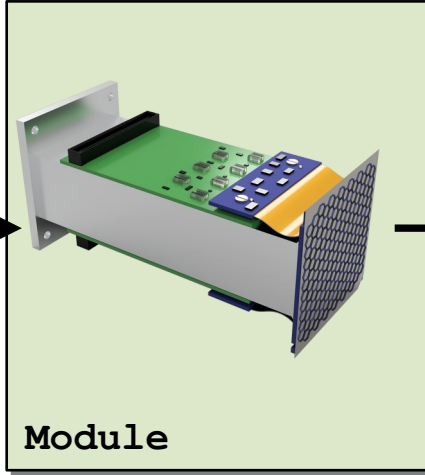


Staged approach



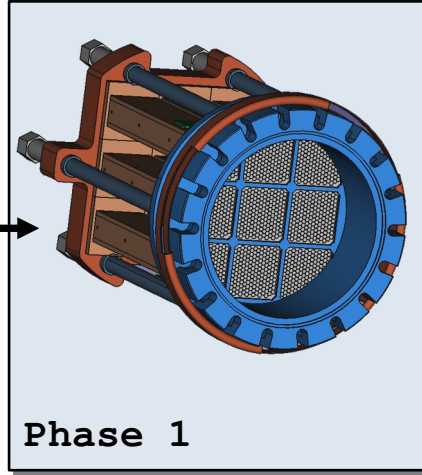
Prototypes

x24



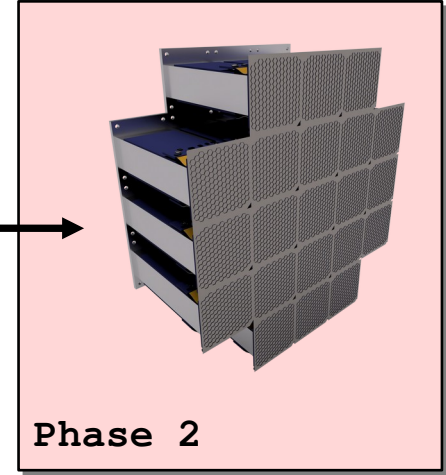
Module

x9



Phase 1

x2.5



Phase 2

Present

TRISTAN prototype

- 7-pixels with external CMOS

TRISTAN module

- 166-pixels with integrated JFET

Mini-TRISTAN

- 9 x modules
→ 1500 pixels

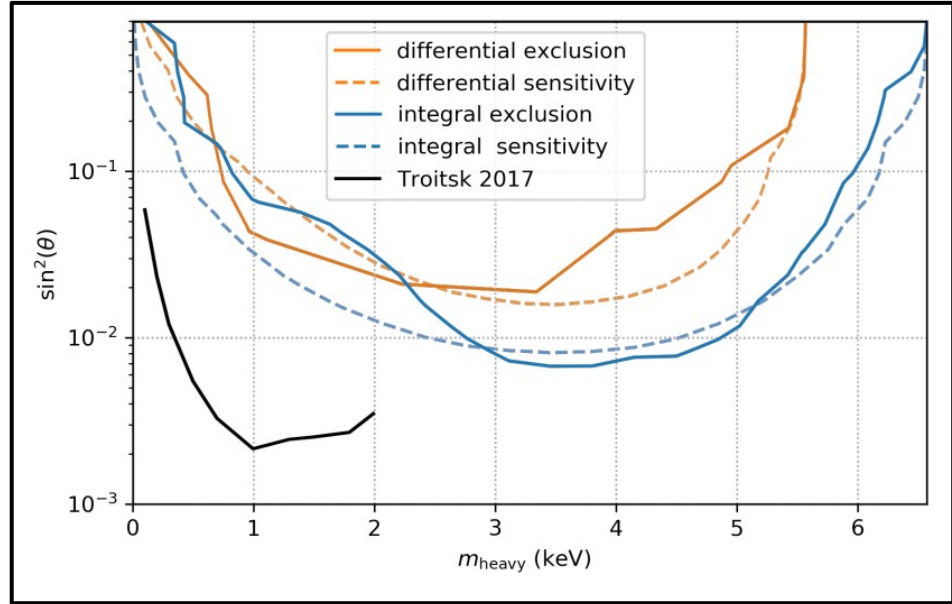
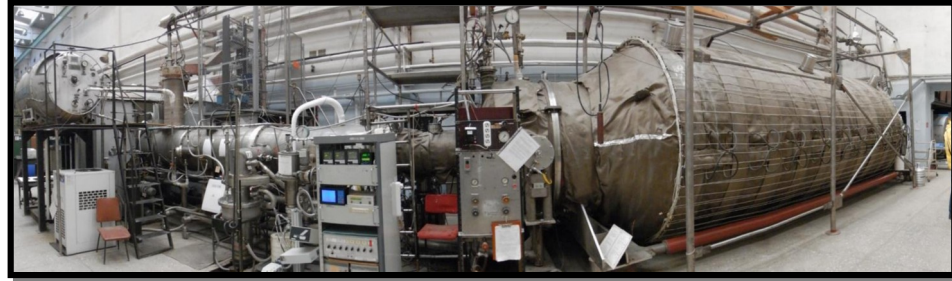
Full TRISTAN

- 21 x modules
→ 3500 pixels

Staged approach



Prototypes

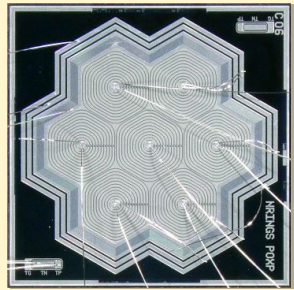


JINST14 P11013 (2019)

TRISTAN prototype

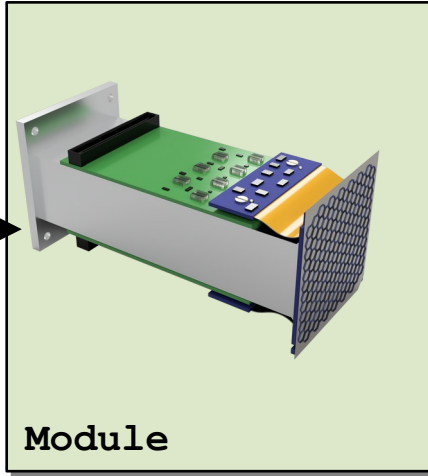
- 7-pixels with external CMOS

Staged approach



Prototypes

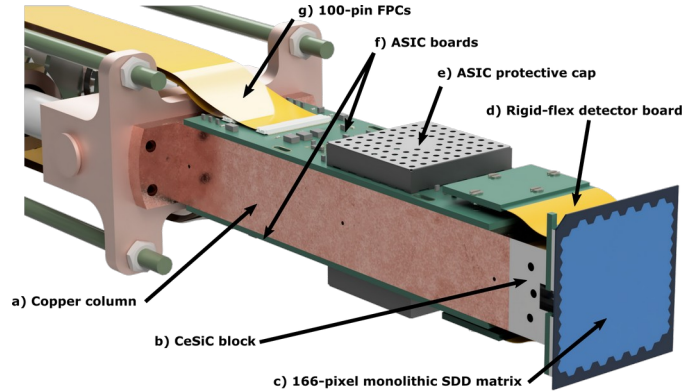
x24



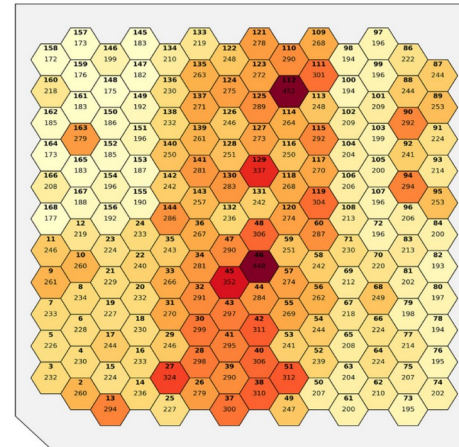
Module

TRISTAN module

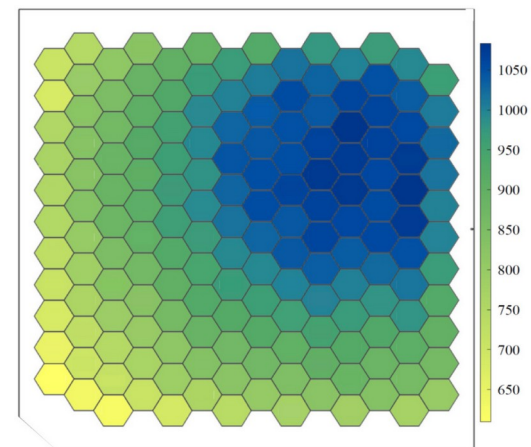
- 166-pixels with integrated JFET



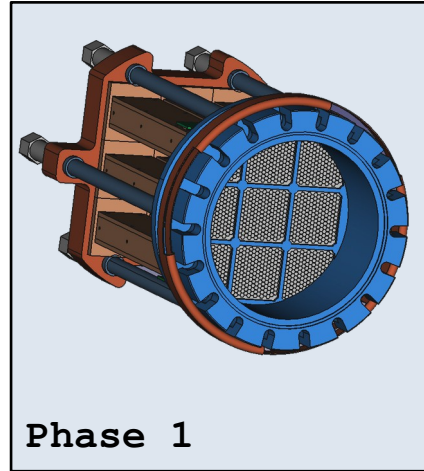
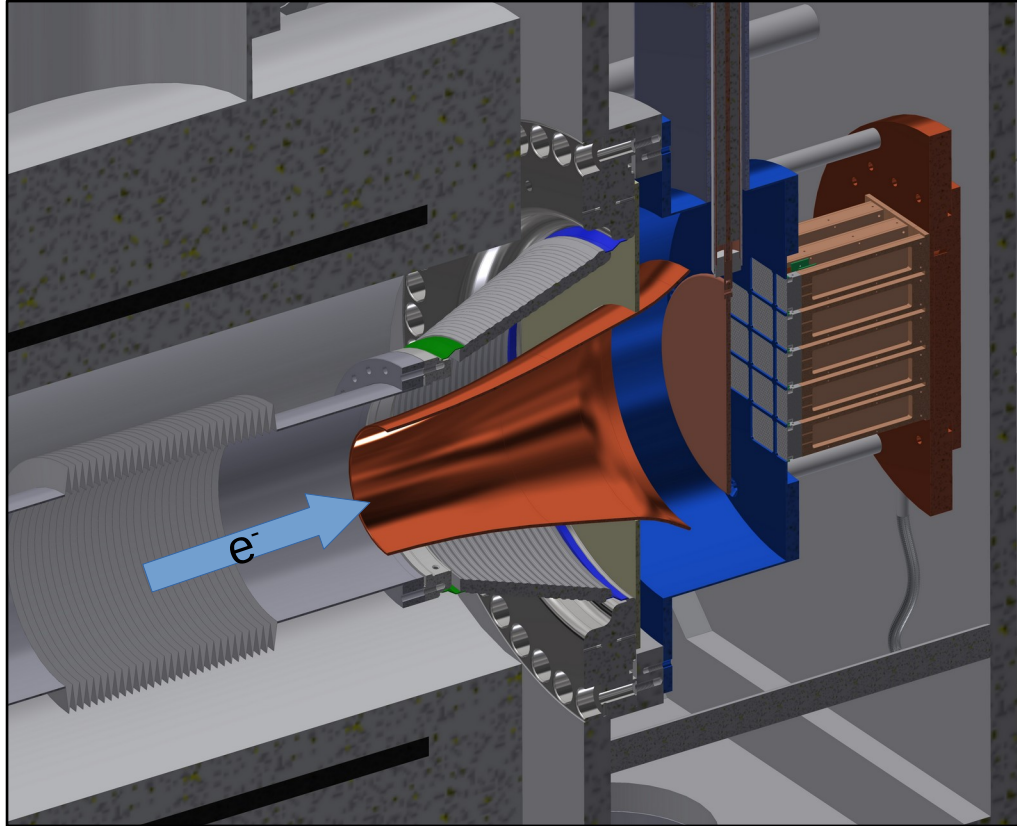
(a) Energy Resolution [eV FWHM]



(b) Count Rate [cps]



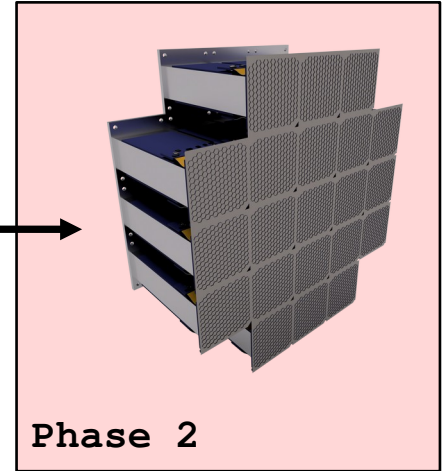
Staged approach



Mini-TRISTAN

- 9 x modules
→ 1500 pixels

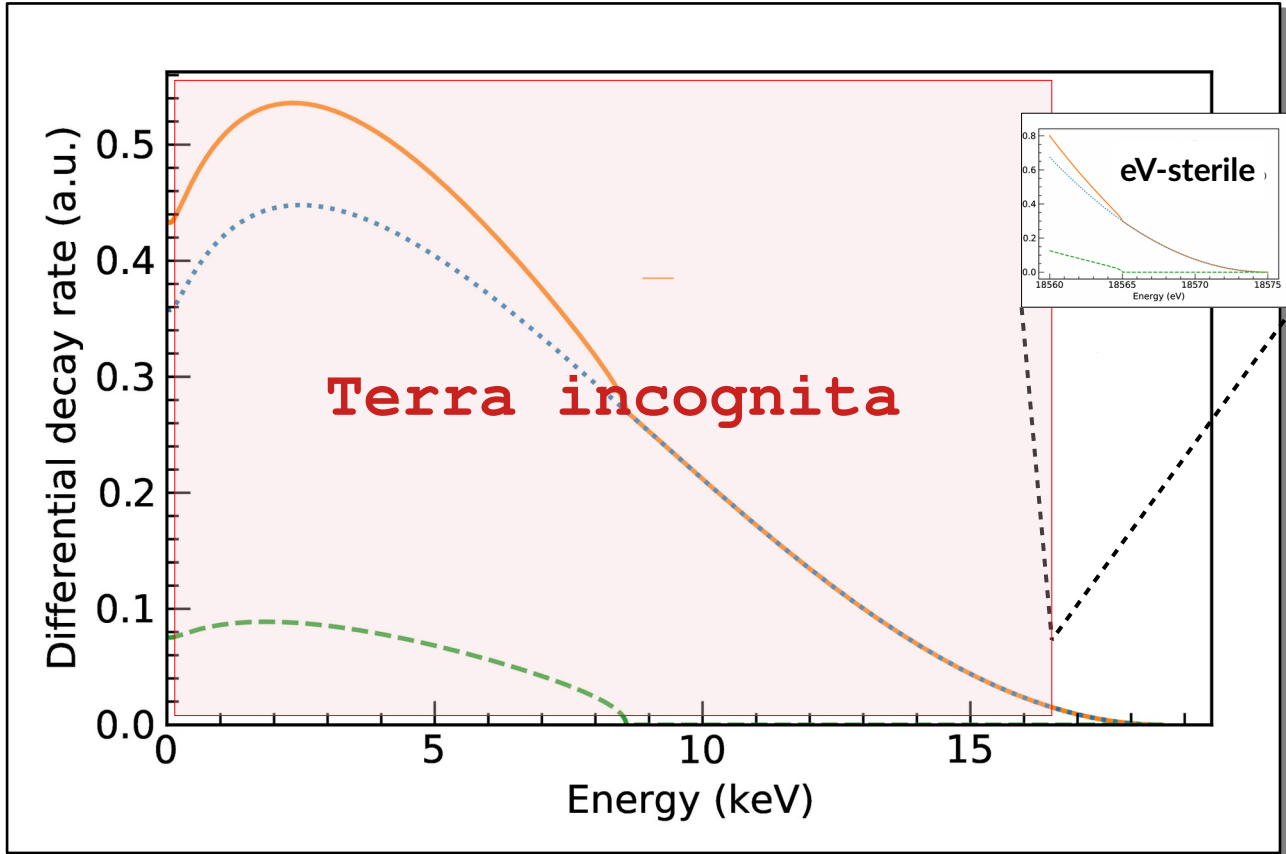
x2.5



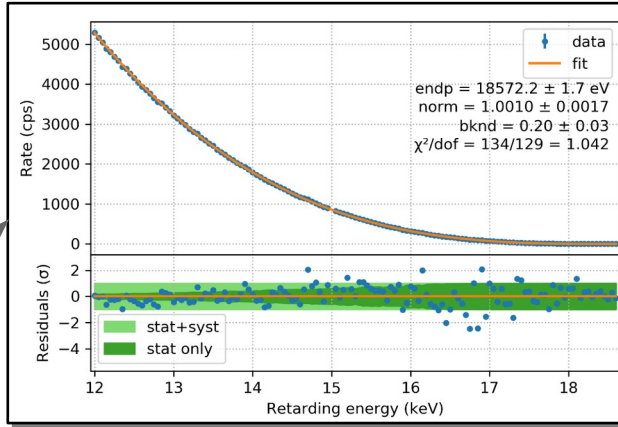
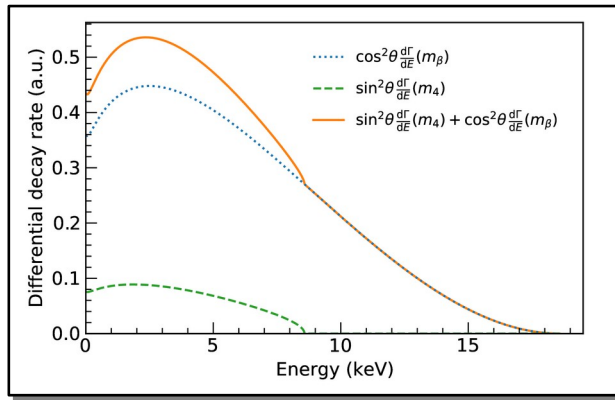
Full TRISTAN

- 21 x modules
→ 3500 pixels

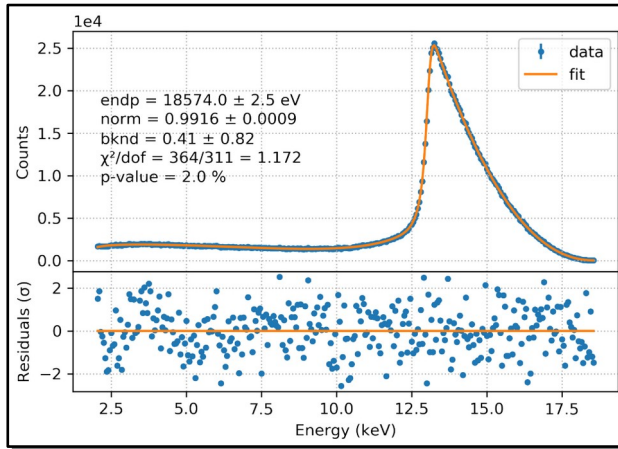
Deep Tritium Model



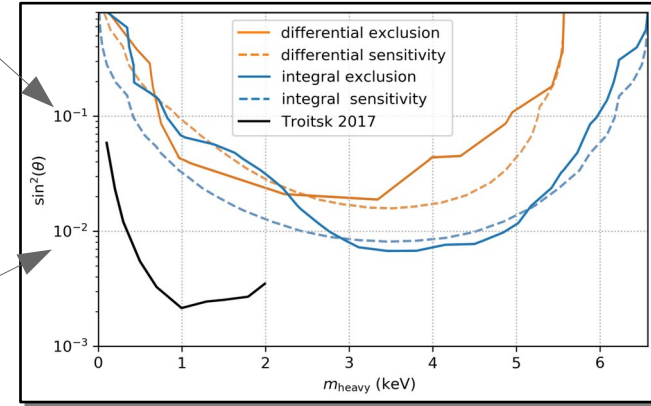
Deep Tritium Model



JINST 14 P11013 (2019)



JINST 14 P11013 (2019)



Deep Tritium Model : effects to consider

Effect can be different for differential and integral mode

Rear-Wall

- Back-scattering
- Auger electron emission
- Residual beta-activity

Source

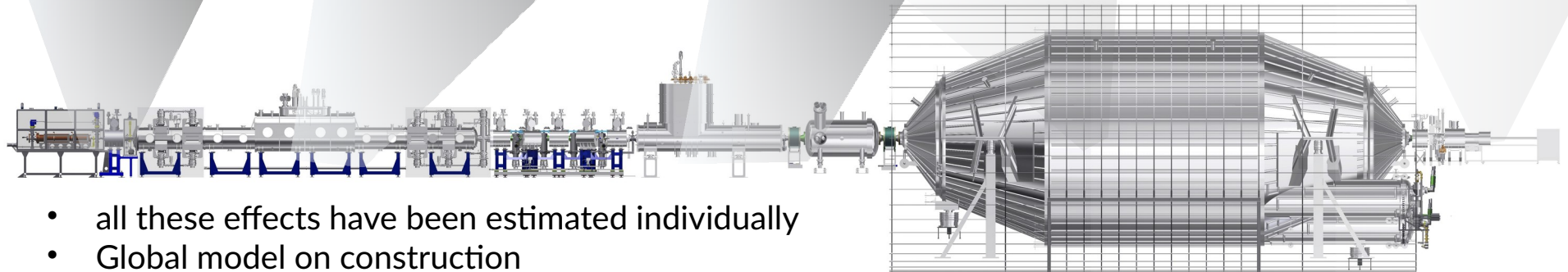
- Scattering
- Magnetic traps
- Plasma effects
- Stability
- Gas composition and impurities

Transport - Spectrometer

- Non-adiabaticity transport
- Synchrotron radiation
- HV stability
- Background
- B-field stab.

Detector

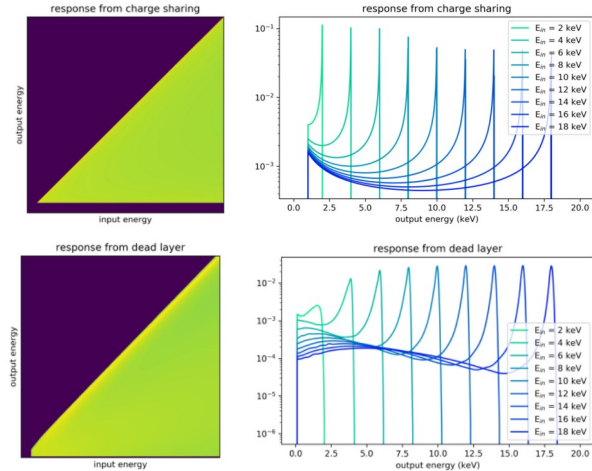
- SDD response
- Read-out resp
- ADC NL
- Post acceleration electrode
- Pile-up, backscattering
- Penning traps
- Stability



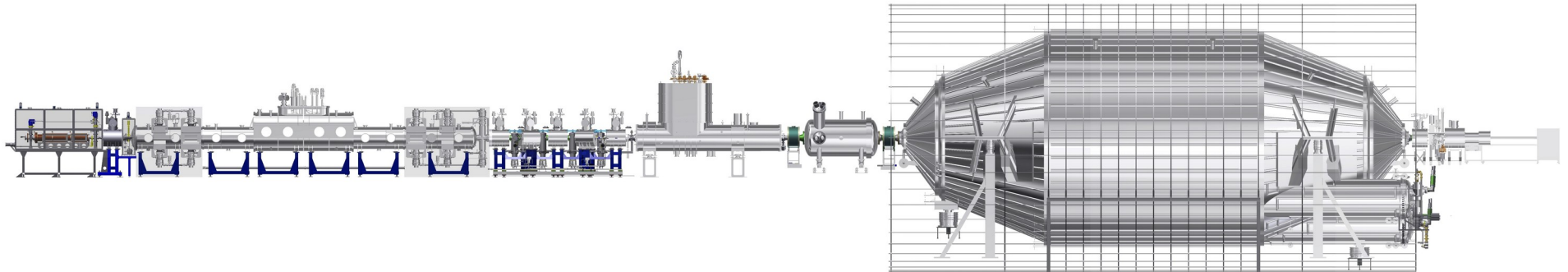
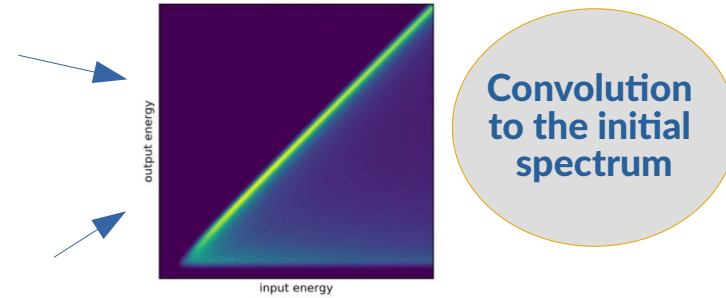
- all these effects have been estimated individually
- Global model on construction

Deep Tritium Model

Each physical effect described with a response matrix



Combined responses



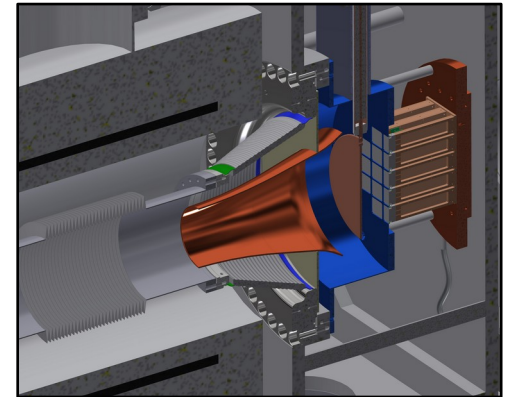
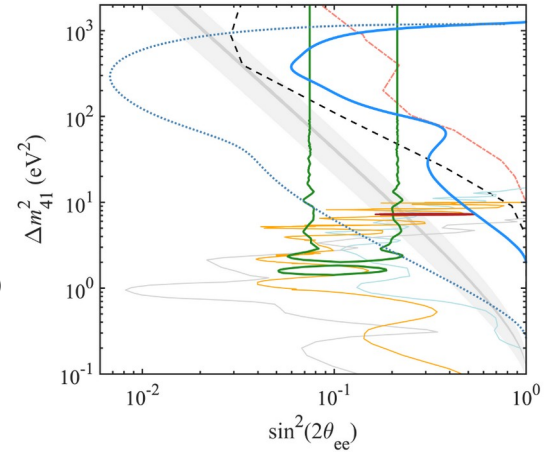
Conclusion

Sterile neutrinos with KATRIN

KATRIN has now presented a first study on eV sterile neutrinos

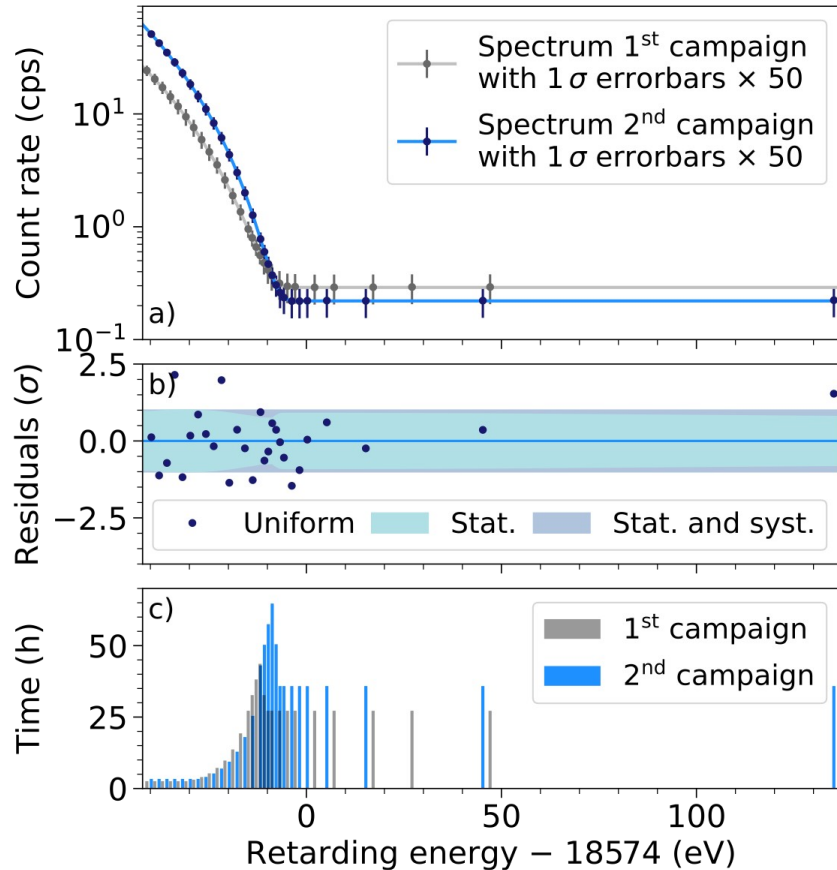
KATRIN with TRISTAN :

- feasibility of the SDD technology for the keV-sterile neutrino search has been demonstrated
 - with photons and electrons
 - with tritium in realistic conditions (Troitsk)
- A **first TRISTAN module** is being commissioned in KATRIN
- A complete **deep tritium model** is being built. New sensitivity studies will be done to reduce systematics
- The TRISTAN technology is also being studied to join solar axion search with the IAXO project



Thank you!

Combination of KNM1 & KNM2



KNM1: 1st campaign:

- total statistics: 2 million electrons
- background 290 mcps
- best fit: $1.0^{+0.9}_{-1.1} \text{ eV}^2$ (stat. Dom.)
 $m_\nu < 1,1 \text{ eV}$ (90% CL)

KNM2: 2nd campaign:

- total statistics: 4.3 million electrons
- background 220 mcps
- best fit: $0.26 \pm 0.34 \text{ eV}^2$ (stat. dom.)
 $m_\nu < 0.9 \text{ eV}$ (90% CL)

- Both KNM1 and KNM2 are statistically domin.
→ Treat them as independent data sets

$m_\nu < 0.8 \text{ eV}$ (90% CL)