

# Using primitive variables to evolve isolated neutron stars with pseudospectral methods

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Based on <https://arxiv.org/abs/2212.10853>

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October 17, 2023

- 1 Introduction
- 2 Conservative vs primitive variables
- 3 Application to numerical simulations: isolated neutron star oscillations
- 4 Conclusion and ongoing work

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- Development of high-resolution shock capturing (HSRC) methods

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- Valencia formulation (Banyuls et al. 1997): conservative formulation of GR-hydro equations
- Development of high-resolution shock capturing (HSRC) methods
- Conservative formulation comes with recovery procedures which use iterative algorithm to recover primitive variables from conservative ones:  $(D, S_i, \tau) = f(e, p, T, U_i)$ .
  - Be source of code failure (non-convergence of iterative algorithms, non-analytical EoS)
  - Computationally consuming (each grid point, each time step)

## Building a set of hydrodynamical equations

- Using non-conservative, physical variables ("primitive" variables): energy, pressure, velocity field, ...
- Testing them in a simple numerical setup: oscillations of barotropic spherically symmetric isolated NS

# Framework

## 3+1 formulation of GR

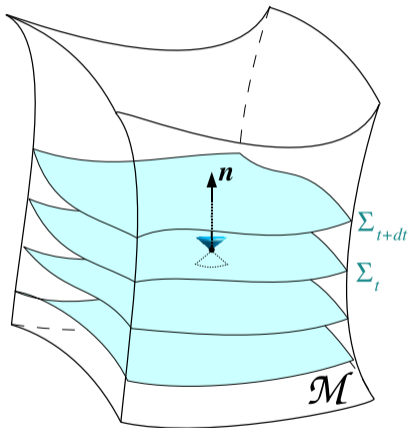


Figure: 3+1 foliation in spacelike hypersurfaces

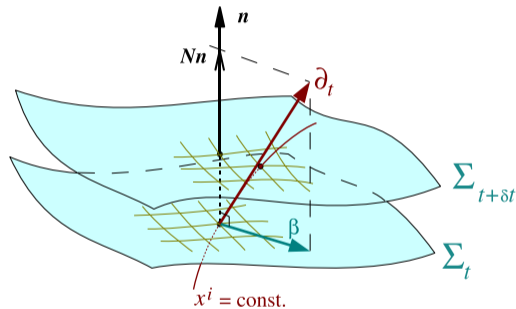


Figure: Illustration of the lapse and the shift

Figures credits: E.ourgoulhon



Physical variables are called "primitive" variables:

- $e$  the energy density,  $p$  the pressure
- $n_B$  the baryon number density
- $m_B$  a baryon mass
- $H = \ln\left(\frac{e+p}{m_B n_B}\right)$  the log-enthalpy
- $U_i$  the Eulerian velocity field
- $v_i$  the coordinate velocity field
- $c_s$  the sound speed

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# Conservative variables

The variables  $D = m_B n_B \Gamma^2$ ,  $S_j = (e + p) \Gamma^2 U_j$  and  $\tau = (e + p) \Gamma^2 - p$ , with  $\Gamma = (1 - U_i U^i)^{-1/2}$  the Lorentz factor, are conserved in the sense that  $\mathbf{u} = (D, S_j, \tau)$  obeys an equation that looks like

$$\partial_t \mathbf{u} + \text{div}(F(\mathbf{u})) = \text{source}$$

but the knowledge of  $e$ ,  $p$ ,  $U_i$  is compulsory to solve Einstein equations to compute the metric. The reference code CoCoNuT (Dimmelmeier et al. 2005) is based upon the conservative formulation of hydrodynamics.

**Recovery procedures are needed to recover the primitive variables in multiple steps of CoCoNuT and account for a significant part of the computation:**

- Solving for the metric
- Solving for the Riemann problems in each grid cell at each timestep

# Full GR equations, using only primitive variables

From the principles of stress-energy and baryon number conservation:

$$\nabla_{\mu}(n_B u^{\mu}) = 0, \quad \nabla_{\mu} T^{\mu\nu} = 0,$$

the following holds for a barotropic, non-reactive perfect fluid:

$$\begin{aligned} \partial_t U_i &= -v^j D_j U_i - D_i N - \frac{N}{\Gamma^2} \left( D_i H - \frac{\Gamma^2(1 - c_s^2)}{\Gamma^2 - c_s^2(\Gamma^2 - 1)} U_i U^j D_j H \right) \\ &\quad + U_j D_i \beta^j + U_i U^j D_j N \\ &\quad + \frac{N c_s^2}{\Gamma^2 - c_s^2(\Gamma^2 - 1)} U_i D_j U^j + \frac{N \Gamma^2 (c_s^2 - 1)}{\Gamma^2 - c_s^2(\Gamma^2 - 1)} U_i U^j U^l K_{jl} \\ \partial_t H &= -v^i D_i H - c_s^2 N \frac{\Gamma^2}{\Gamma^2 - c_s^2(\Gamma^2 - 1)} \left[ U^i U^j K_{ij} - \frac{U^i}{\Gamma^2} D_i H + D_i U^i \right] \end{aligned}$$

This new set of equations is **covariant** within the 3+1 formalism.

$e$ ,  $p$ ,  $c_s$  are recovered in a single call to the EoS (**no iteration**).

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# Frequency extraction

The equations are implemented in a **spherically symmetric, semi-Lagrangian C++** code based upon LORENE: **pseudospectral methods** (space) + explicit **finite-differences** (time).

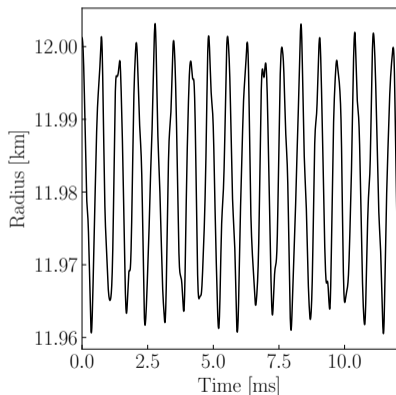


Figure: Radius vs time

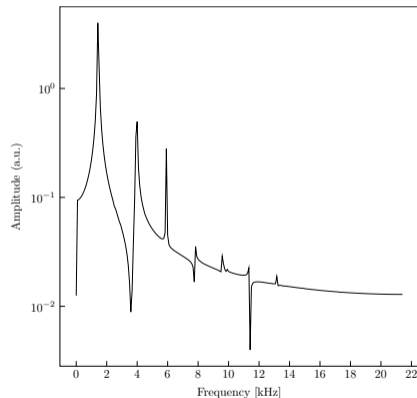


Figure:  $\hat{R}(f)$  spectrum

# Validation of the code: frequency extraction

- Reproduction of polytropic frequencies given in Font et al. 2002 and Hartle & Friedman 1975 with less than 1% precision in full GR:

	$\kappa$	$\gamma$	$H_c [c^2]$	$M [M_\odot]$	$R$ [km]	Fund. [kHz]	1st ov. [kHz]	2nd ov. [kHz]
Font et al. [25]	100	2	0.2279	1.4	14.15	1.450	3.958	5.935
This work	100	2	0.2279	1.401	14.16	1.442	3.954	5.915
Relative difference						0.6%	0.1%	0.3%
Hartle & Friedman [26]	7.308	5/3	$6.720 \times 10^{-2}$	×	×	0.824	1.94	2.86
This work	7.308	5/3	$6.720 \times 10^{-2}$	0.4866	16.49	0.823	1.95	2.86
Relative difference						0.1%	0.5%	0.0%

- In Cowling approximation:

	$\kappa$	$\gamma$	$H_c [c^2]$	$M [M_\odot]$	$R$ [km]	Fund. [kHz]	1st ov. [kHz]	2nd ov. [kHz]
Font et al. [25]	100	2	0.2279	1.4	14.15	2.696	4.534	6.346
This work	100	2	0.2279	1.401	14.16	2.685	4.548	6.339
Relative difference						0.4%	0.3%	0.1%

# Validation of the code: Migration test

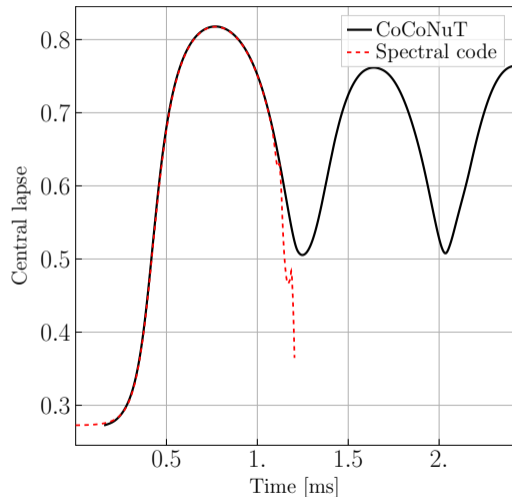


Figure: Central lapse

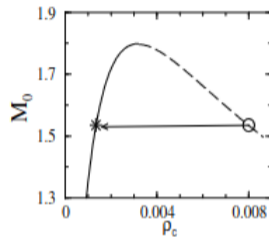


Figure: Migration test principle

- Shock formation during the oscillations
- Gibbs phenomenon: the spectral code crashes
- 10s (spectral code) vs 6min (CoCoNuT)



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- New set of **covariant** 3+1 equation for any configuration of the perfect fluid with primitive variables
- Primitive variables allow to simulate subsonic, smooth flows
- No shock treatment + spectral methods = shocks are prohibitive (cf migration test)
- 3D code is **under development**.

Thank you !  
arxiv:2212.10853

The set of Newtonian equations is:

$$\partial_t \rho + \nabla_i (\rho v^i) = 0$$

$$\partial_t v^i + v^j \nabla_j v^i = -\nabla^i (h + \Phi)$$

$$\Delta \Phi = 4\pi \rho$$

$$h = \frac{\kappa \gamma}{\gamma - 1} \rho^{\gamma-1}$$

# Frequency extraction : principle

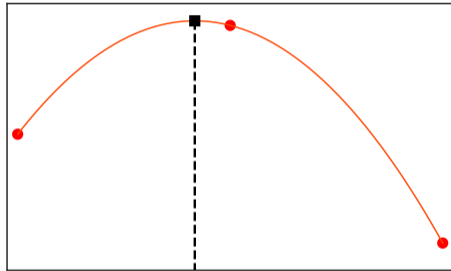


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