

# Action-angle variables of post-Newtonian binary black holes

Sashwat Tanay

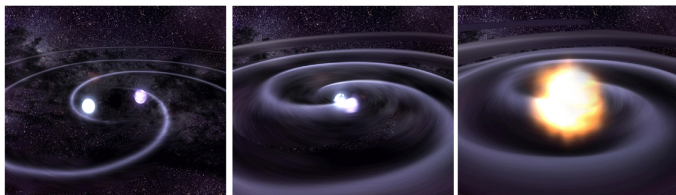
LUTH, CNRS/Paris Observatory/Universite Paris Cité

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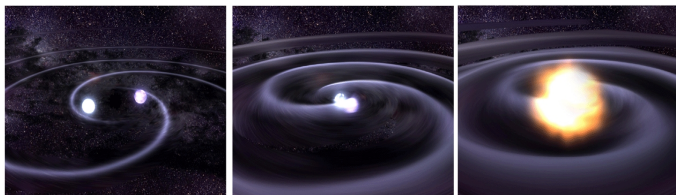
In collaboration with L. C. Stein, G. Cho, J. T. Gálvez Gherzi, and R. Samanta

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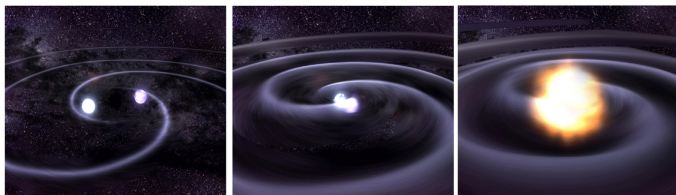
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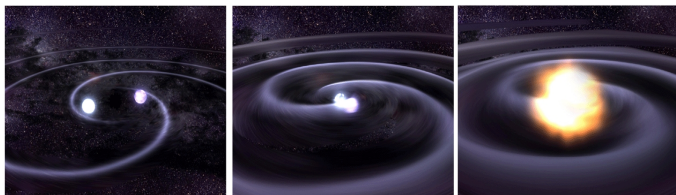
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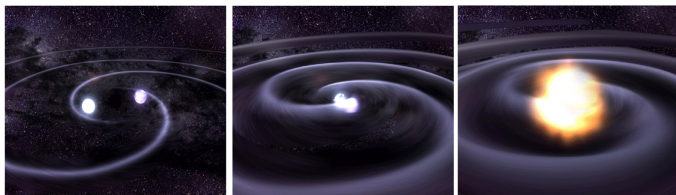
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- GWs are functions of **black hole trajectories** (*focus of the talk*).

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- **Example:** BBH Hamiltonian

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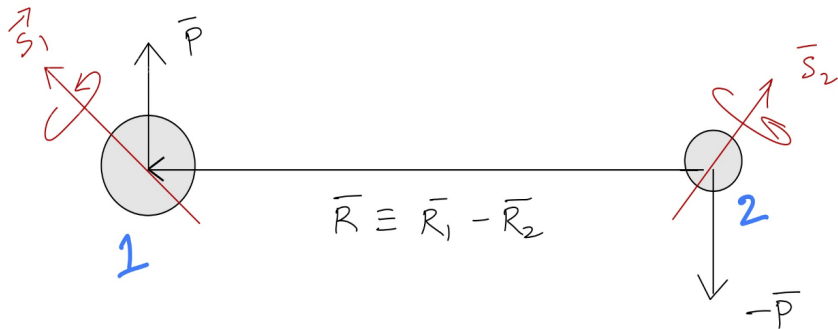
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- Each factor of  $1/c^2 \rightarrow$  one PN order.

# Phase space of spinning BBHs

COM FRAME



$\vec{R}, \vec{P}, \vec{S}_1, \vec{S}_2$

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- Solutions are crucial for fast GW template construction and data analysis.

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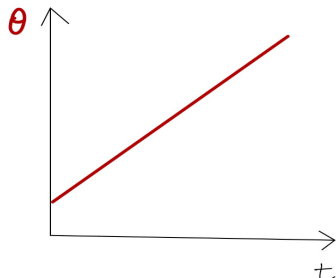
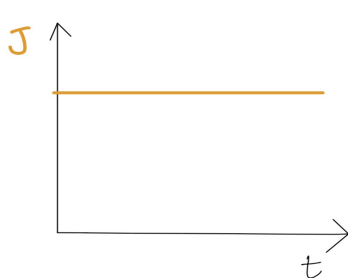
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**ACTION-ANGLES ARE COOL!**



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$$H = \underbrace{\left( \frac{P^2}{2\mu} - \frac{Gm_1 m_2}{R} \right)}_{\text{Newtonian}} + \frac{1}{c^2} F_1(\vec{R}, \vec{P}) + \frac{1}{c^3} F_2(\vec{R}, \vec{P}, \vec{S}_1, \vec{S}_2)$$

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- Extendable to higher PN via canonical pert. theory (Goldstein).

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- $m \equiv m_1 + m_2, \quad \mu \equiv m_1 m_2 / m, \quad \nu \equiv \mu / m, \quad \vec{L} \equiv \vec{R} \times \vec{P},$   
 $\sigma_1 \equiv (2 + 3m_2 / m_1), \quad \sigma_2 \equiv (2 + 3m_1 / m_2), \quad \vec{S}_{\text{eff}} \equiv \sigma_1 \vec{S}_1 + \sigma_2 \vec{S}_2,$   
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- $\mathcal{J}_5$  is very, very lengthy.

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## Refs:

- Papers: [2012.06586](#), [2110.15351](#), [2210.01605](#).
- Lecture notes: [2206.05799](#)
- Mathematica package:  
[github.com/sashwattanay/BBH-PN-Toolkit](https://github.com/sashwattanay/BBH-PN-Toolkit)
- [▶ YouTube video](#) on the package
- Contact: [sashwat.tanay@obspm.fr](mailto:sashwat.tanay@obspm.fr)



Thank you!  
Questions?