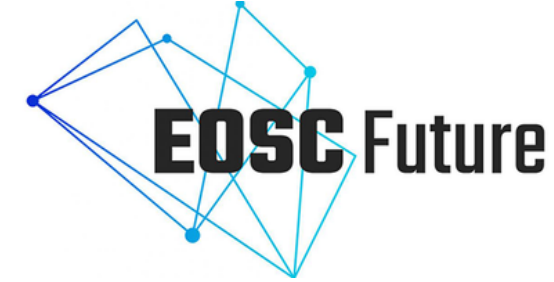


# The Dark Matter Science Project

JARED LITTLE, ON BEHALF OF THE DARK MATTER SCIENCE PROJECT RESEARCHERS



# ESCAPE, Open Science Projects

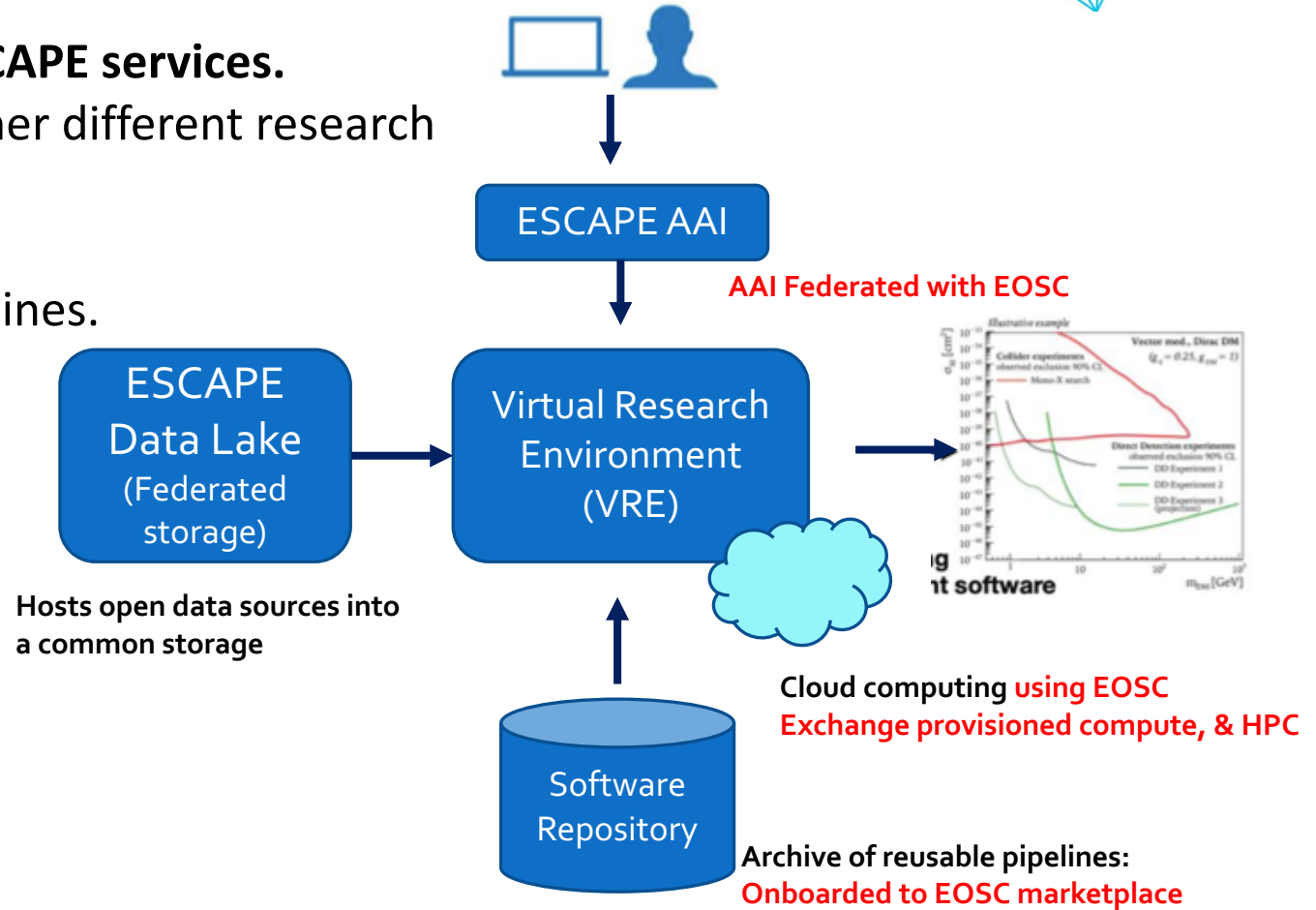


**Open Science Projects bring together different ESCAPE services.**

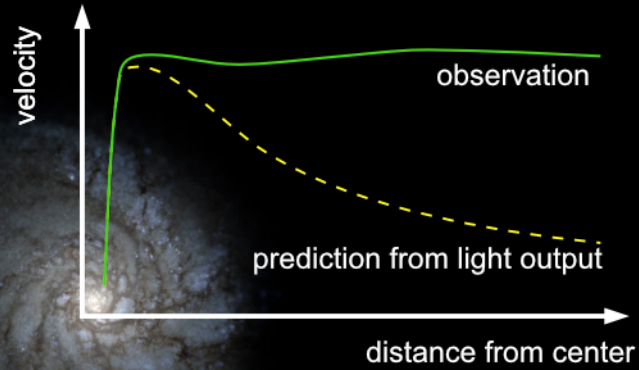
- ESCAPE is an EU-funded project to bring together different research infrastructures.
- Improve productivity of researchers.
- Gain new insights and innovation across disciplines.
- Two science projects being developed.

1. Dark Matter Science Project
2. Extreme Universe Science Project

See following talk by **G. Lamanna**

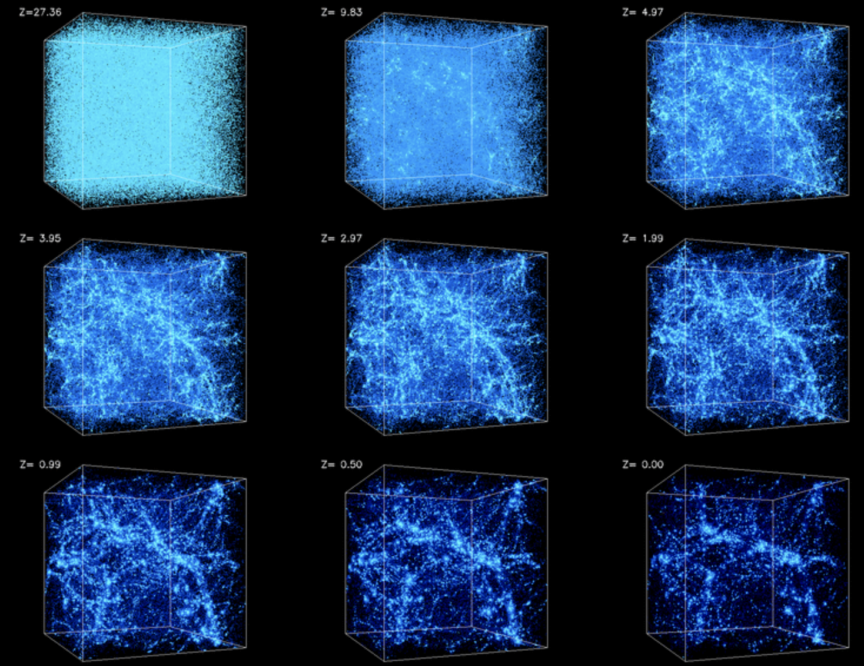


# Big science question: Dark Matter



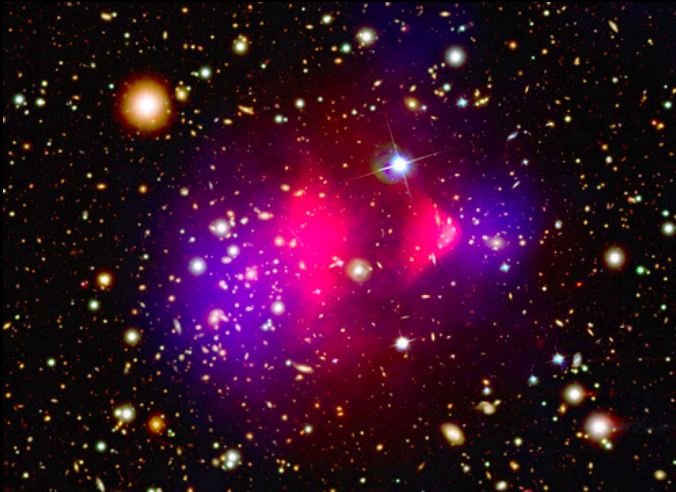
Vera Rubin,  
© Washington Times & Zuma

Wikipedia &  
Hopkins Research Group/Caltech



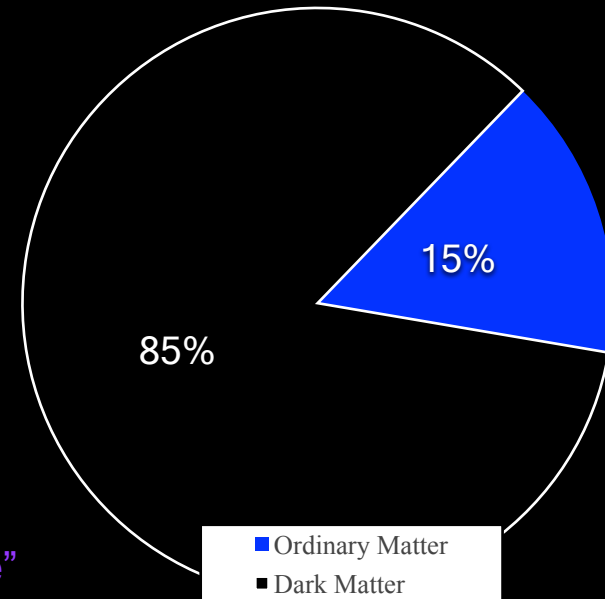
Simulations were performed at the National Center for Supercomputer Applications by A. Kravtsov and A. Klypin.

visible  
matter



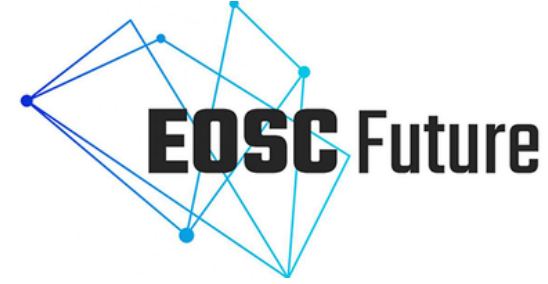
NASA/CXC/M. Weiss

“invisible”  
matter



<https://cordis.europa.eu/project/id/679305>

# Dark Matter Experiments



**Cutting edge dark matter experiments are increasingly unique**

- large, complex, costly experiments
- only one or a few experiments of each type worldwide



**Maximizing each experiment's science outputs is imperative:**

- **create** and store new analyses, datasets and results
- **combine** multiple results studying the same question
- **reinterpret** existing studies for new questions

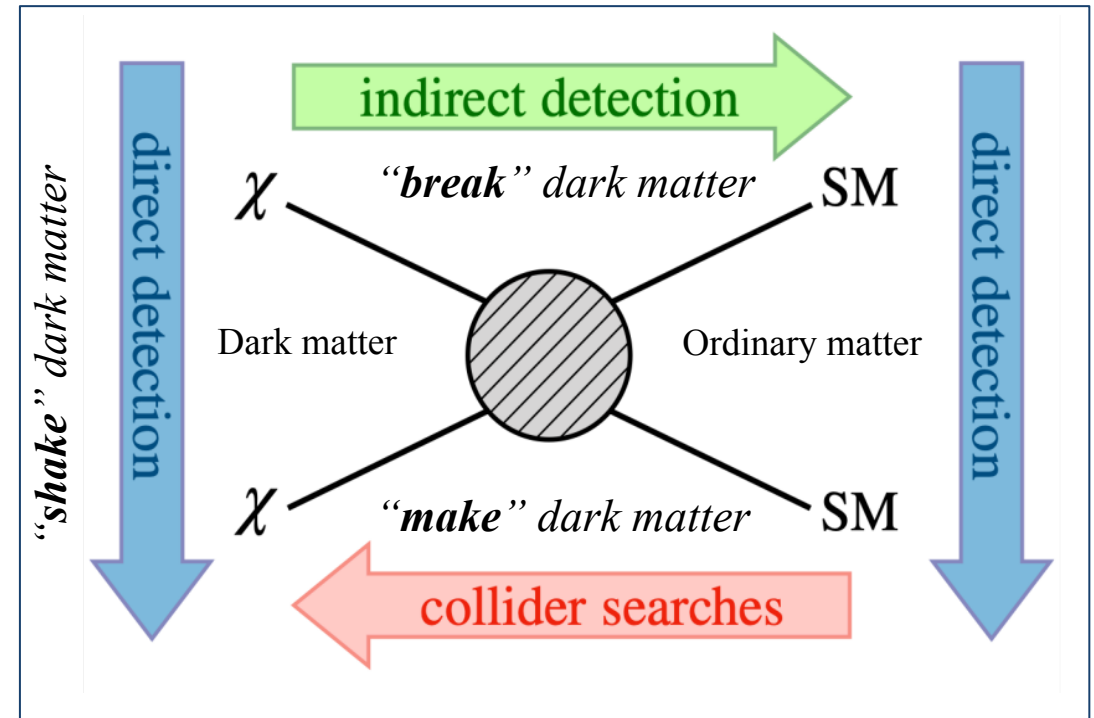


**The Dark Matter Science Project provides the community with [tools](#) to do all these tasks and allows [access to data and software](#) on the [EOSC](#) through [ESCAPE](#) infrastructure**

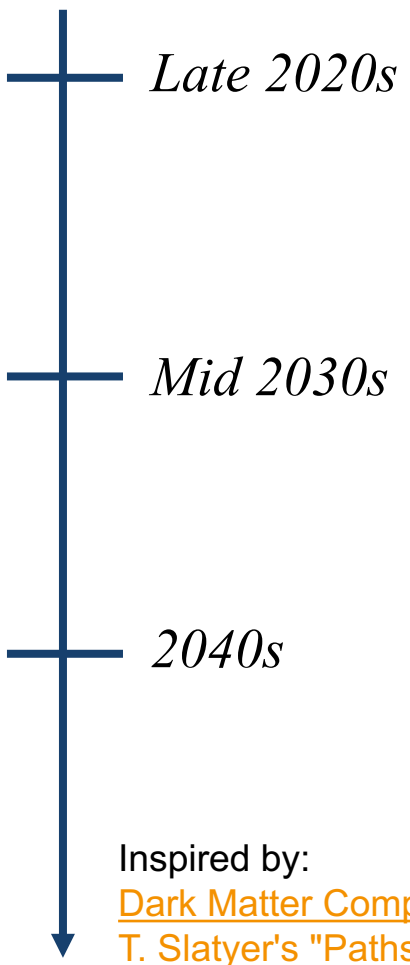
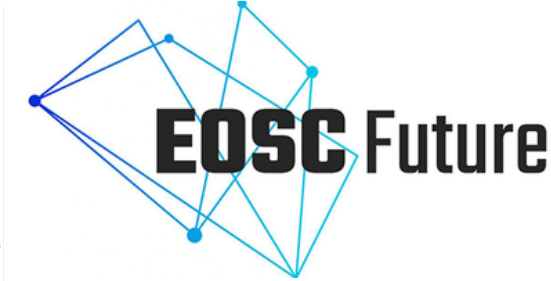
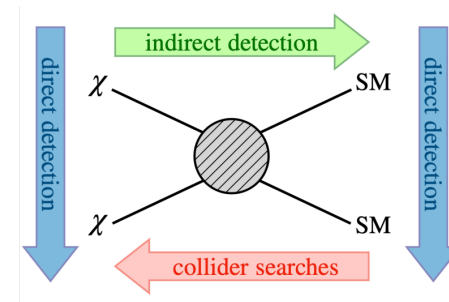
# Dark Matter: Complementary Approach

**A joint discovery of the nature of dark matter requires different experiments and inputs**

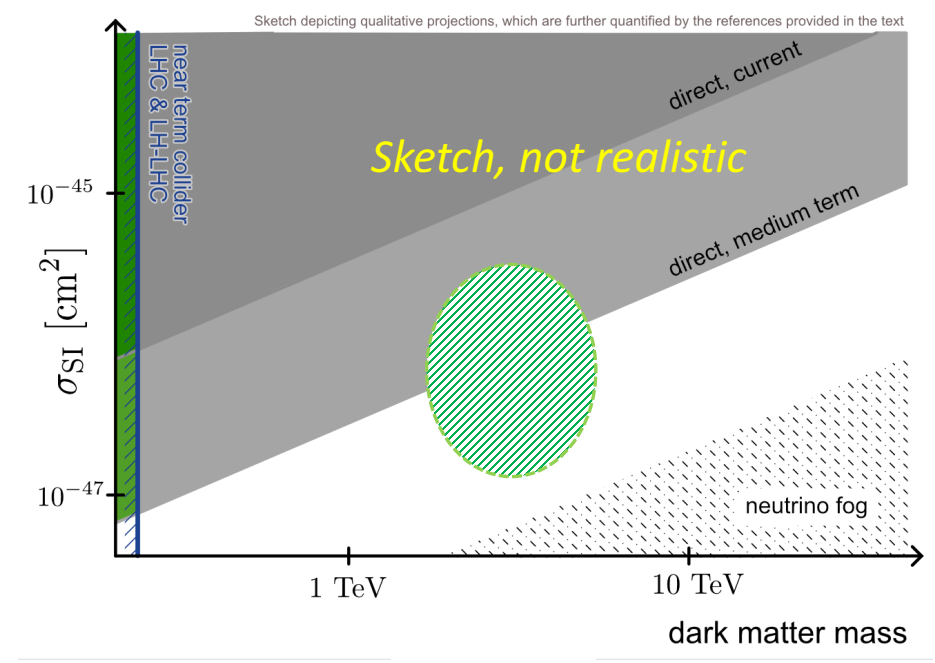
Experiments have **different** data sizes, workflows, data, and result sharing policies



# Example of a *discovery scenario*

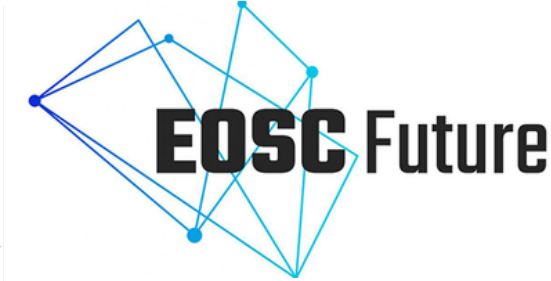
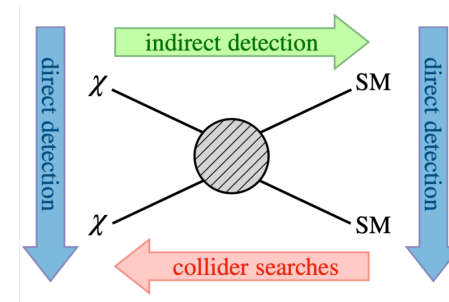


**Direct detection** experiment sees a hint of a signal, with characteristics compatible with WIMP DM



Inspired by:  
[Dark Matter Complementarity \(Snowmass report\), arXiv:2210.01770](#)  
[T. Slatyer's "Paths to discovery" talk at Snowmass 2022](#)

# Example of a *discovery scenario*



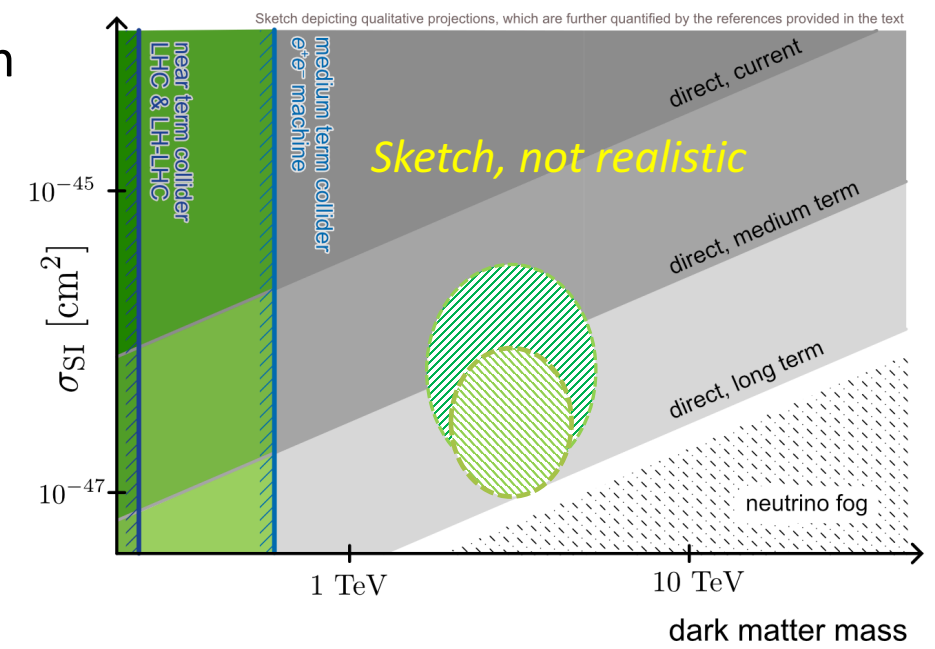
Late 2020s

Mid 2030s

2040s

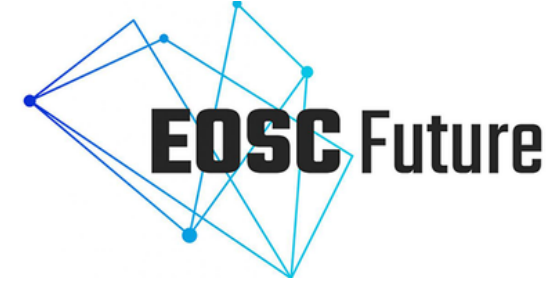
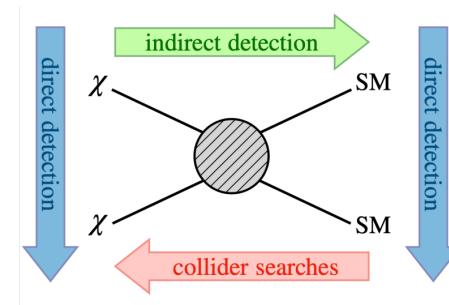
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**Direct detection** experiment (using another tech) confirms these hints



Inspired by:  
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# Example of a *discovery scenario*



Late 2020s

Mid 2030s

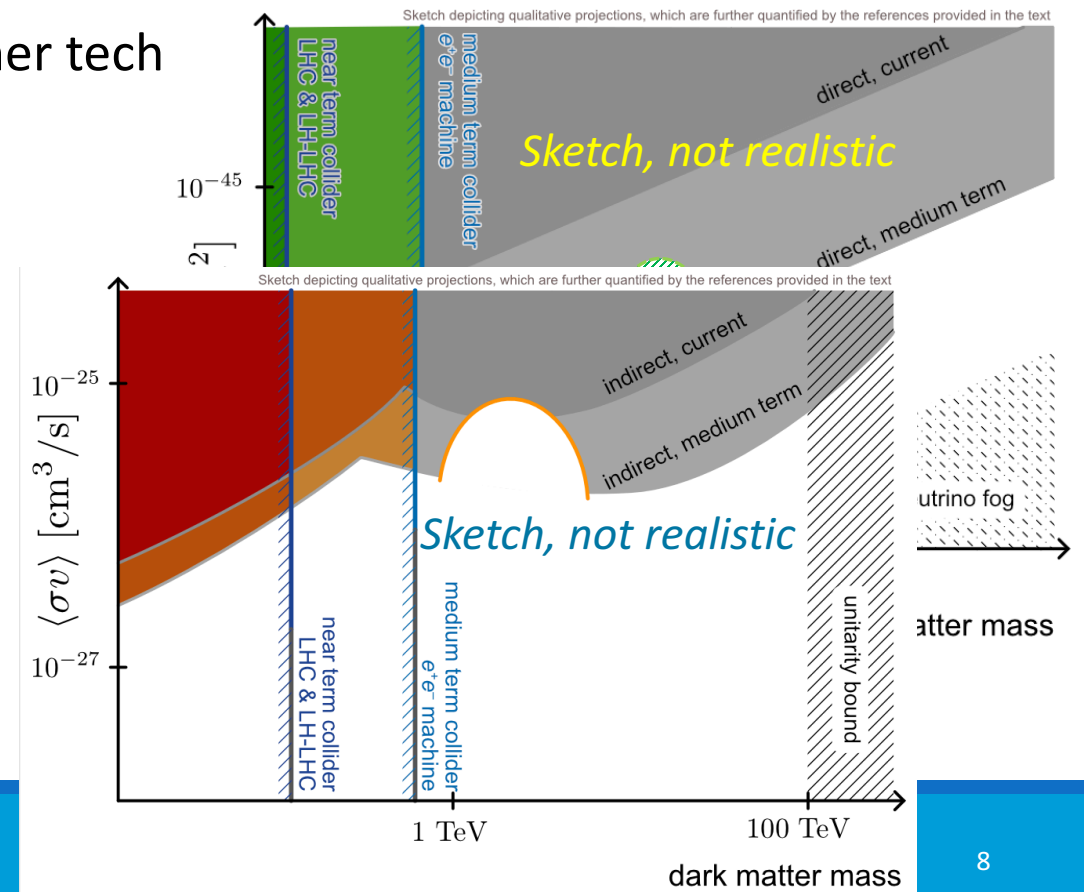
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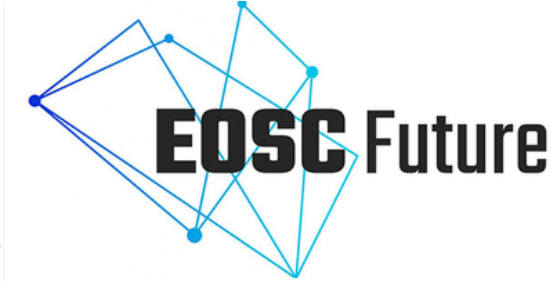
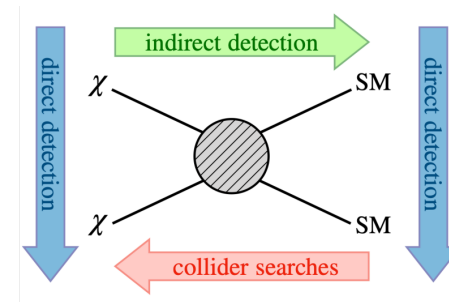
**Direct detection** experiment (using another tech confirms these hints

**Indirect detection** experiment observes signals of DM annihilation





# Example of a *discovery scenario*



Late 2020s

Mid 2030s

2040s

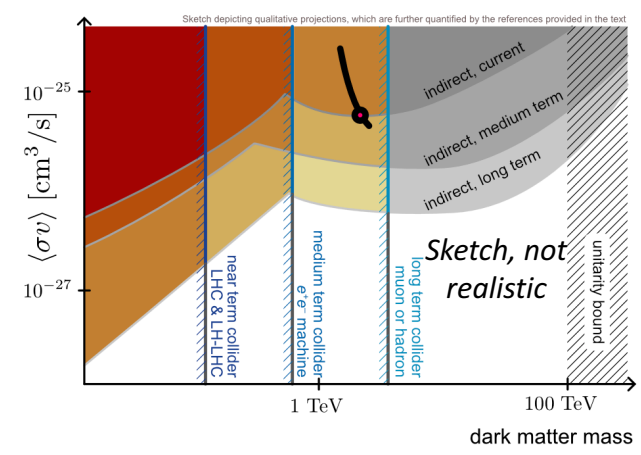
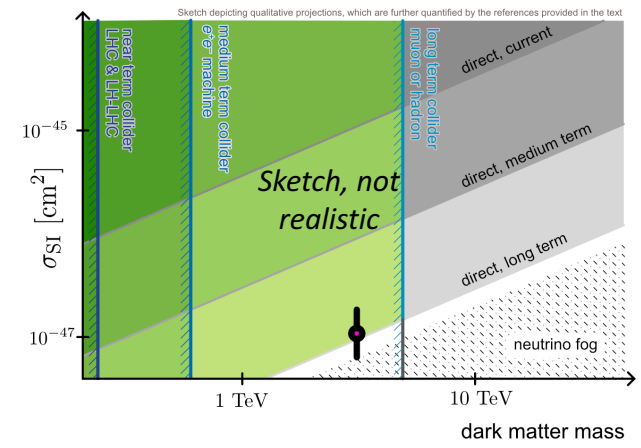
**Direct detection** experiment sees a hint of a signal, with characteristics compatible with WIMP DM

**Direct detection** experiment (using another technique) confirms these hints

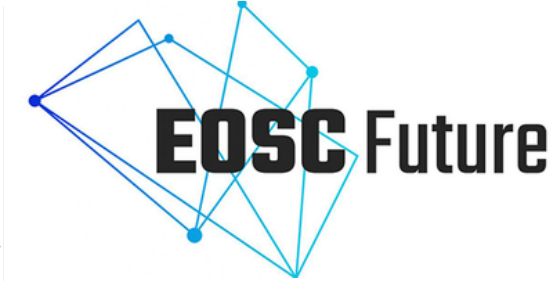
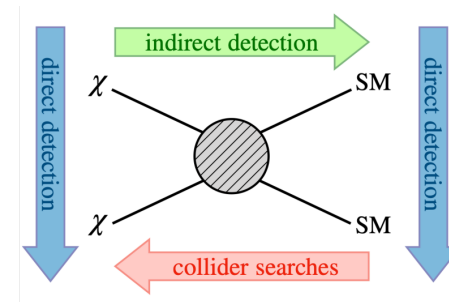
**Indirect detection** experiment observes signals of DM annihilation

**Future collider**, built to target particles with the mass of the putative DM candidate, sheds light on interactions between DM and ordinary matter

Inspired by:  
[Dark Matter Complementarity \(Snowmass report\), arXiv:2210.01770](#)  
[T. Slatyer's "Paths to discovery" talk at Snowmass 2022](#)



# Example of a *discovery scenario*



Late 2020s

Mid 2030s

2040s

**Direct detection** experiment sees a hint of a signal, with characteristics compatible with WIMP DM

**Direct detection** experiment (using another technique) confirms these hints

**Indirect detection** experiment observes signals of DM annihilation

**Future collider**, built to target particles with the mass of the putative DM candidate, sheds light on interactions between DM and ordinary matter

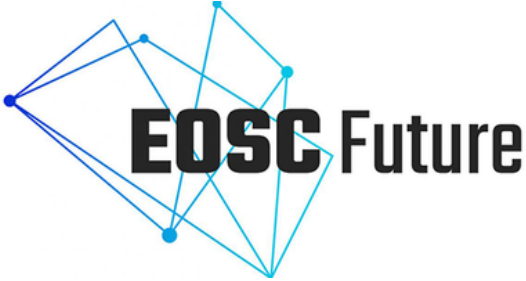
Such a scenario requires **interoperable and reproducible analyses**

- **comparison and combination** of results from different experiments
- **end-to-end workflows** available for cross-checks

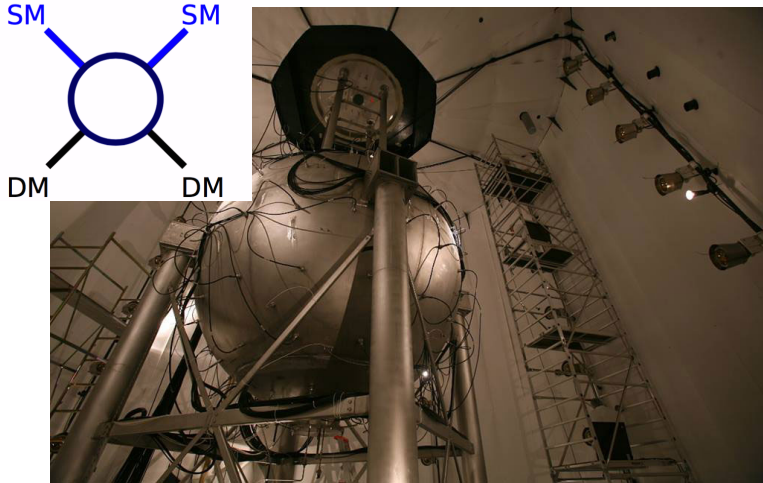
Inspired by:  
[Dark Matter Complementarity \(Snowmass report\), arXiv:2210.01770](#)  
[T. Slatyer's "Paths to discovery" talk at Snowmass 2022](#)

**With the Dark Matter Science Project, we build a *prototype that fulfills these requirements***

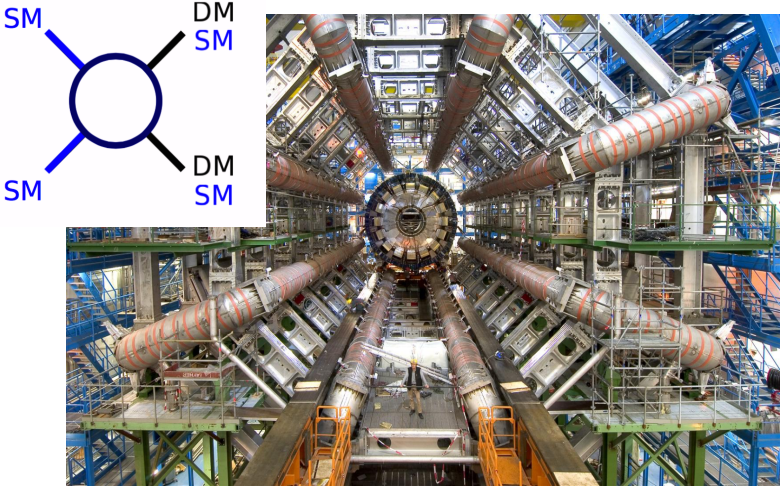
# Experiments involved in the Dark Matter Science Project



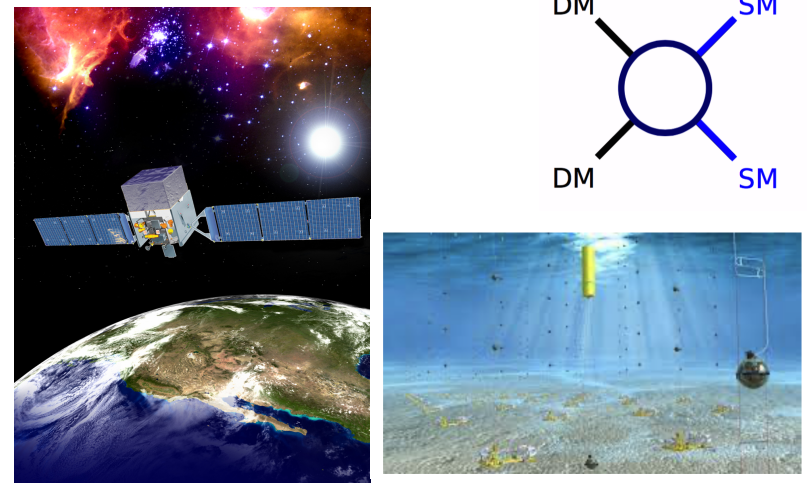
## Direct detection: DarkSide



## Colliders: ATLAS @ LHC



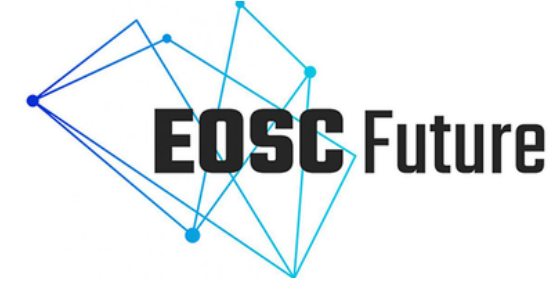
## Indirect detection: FermiLAT, KM3NeT



*...and their evolutions: DarkSide-20k / Argo, ATLAS @ HL-LHC, CTA*  
*Some of the **analysis & ML tools** necessary for these evolutions are also part of this Science Project*

With the Dark Matter Science Project,  
we understand the **computing and analysis challenges** of some of the future DM experiments

# Science outputs of the DM TSP

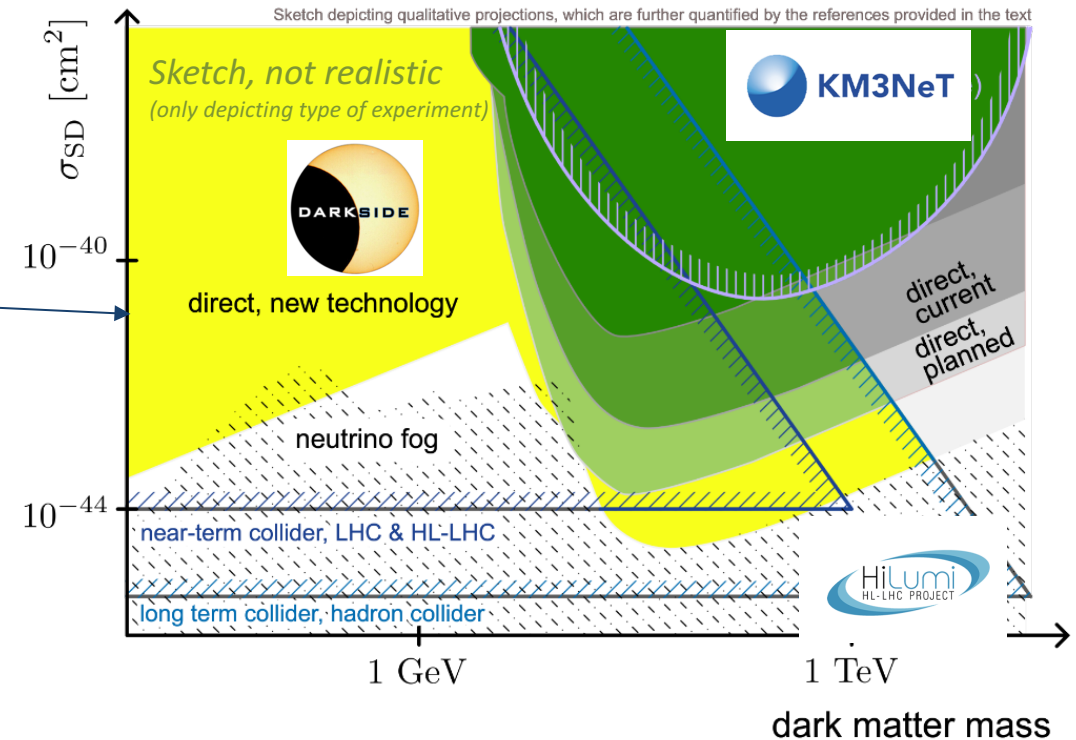


## *Planned domain science outputs*

- Individual results and publications
- Plots highlighting complementarity of different experimental efforts
- Combination of experimental results

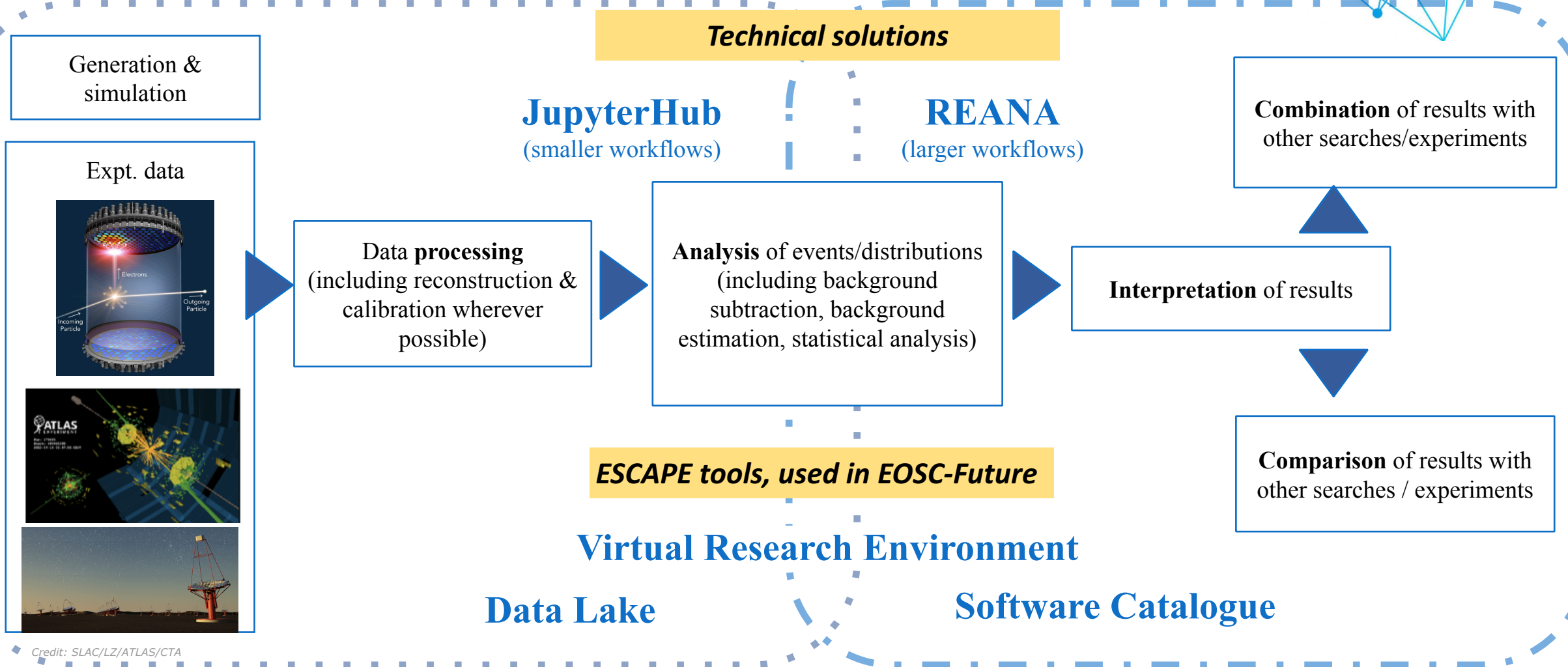
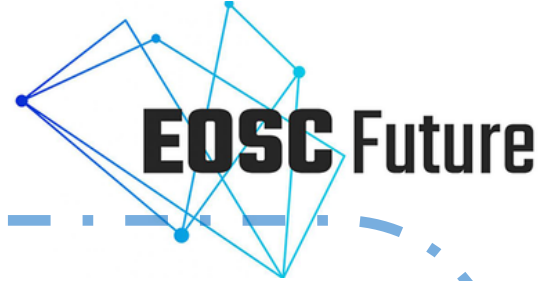
## *Data and software objects + pipelines*

- Data on the Data Lake, and software on the ESCAPE Software Catalogue
- Pipelines accessible via VRE



**Example sketch (*not* using ESCAPE experiments yet) highlighting direct detection, neutrino experiment indirect detection and collider complementarity**

# Analysis Workflows for the DM Science Project



Credit: SLAC/LZ/ATLAS/CTA

*Data sharing and data processing*

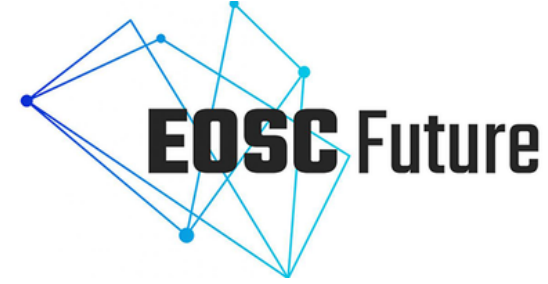
*Data analysis, preservation and interpretation*

# Dark matter at particle colliders: searches in the ATLAS experiment

Jared Little  
(LAPP)

Supervised by:  
Tanya Hrn'ova and Stephane Jezequel (LAPP),  
Caterina Doglioni  
(University of Manchester and Lund University)

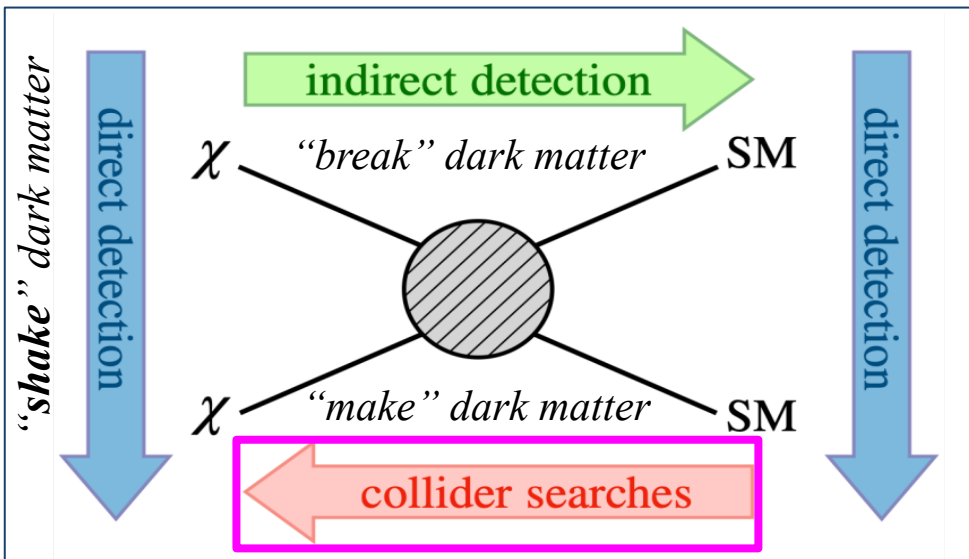
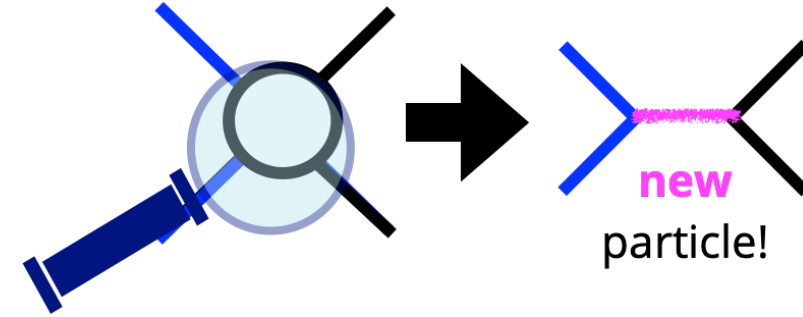
# DM Science Project - ATLAS



The **ATLAS Experiment**, along with CMS, are two general purpose detectors located on the Large Hadron Collider.

Wide range of physics investigated:

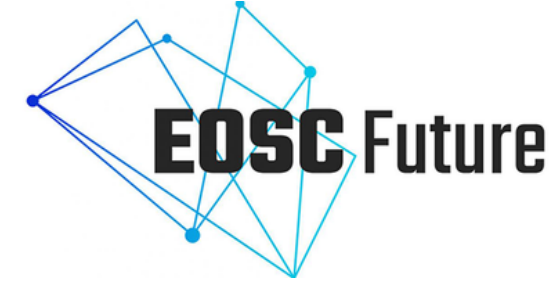
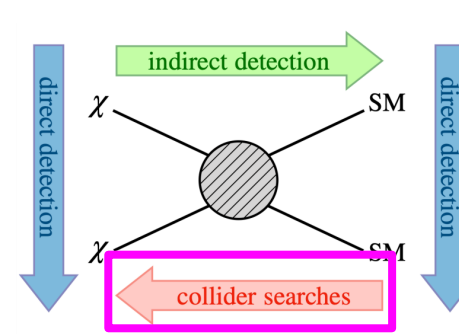
- Higgs discovered in 2012.
- Precision measurements on Standard Model properties.
- Searches for new physics, including particles that make up dark matter.



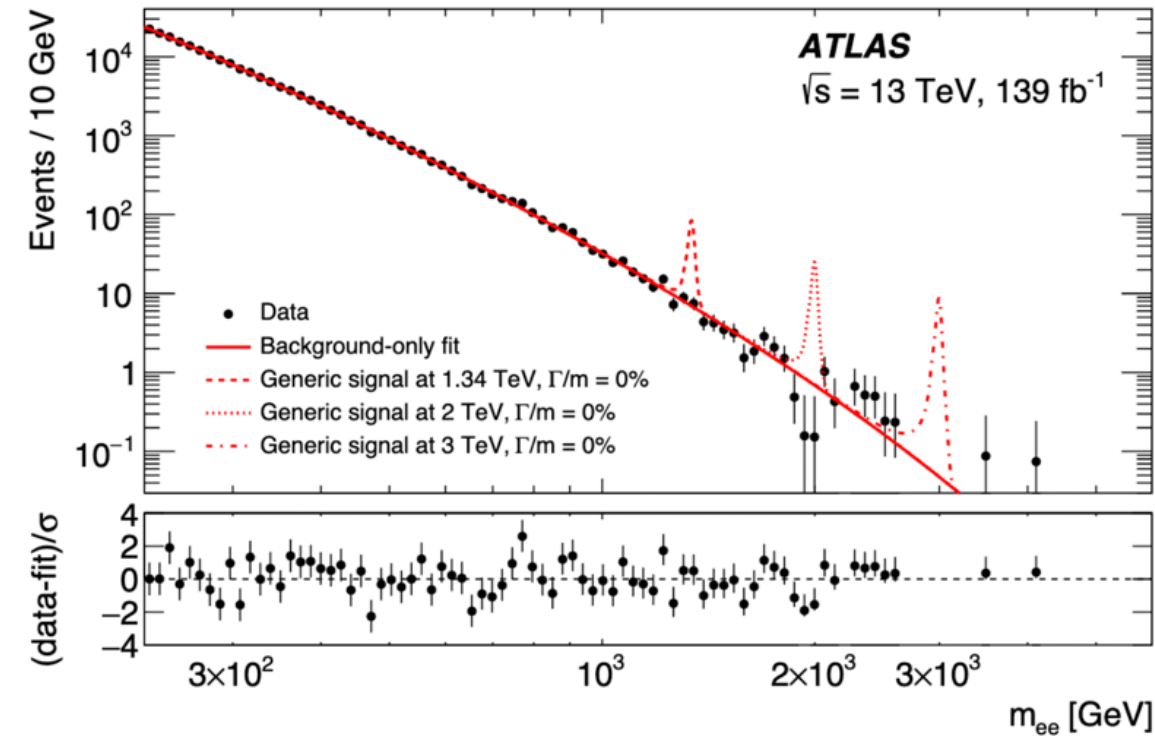
At the LHC, we are trying to “make” dark matter.

- By probing the interactions with ordinary matter, we can better understand the nature of DM.

# Inclusive Dilepton Resonance Search



Looking for a **bump** (= new particle) over the background of known particles



**DM mediator decays in two electrons**

→ search in di-electron final state

- No signal → constraints on the fiducial cross-section of a new  $Z'$  particle.

**Large backgrounds in the region below 1 TeV.**

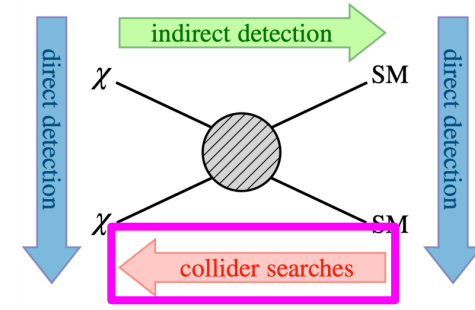
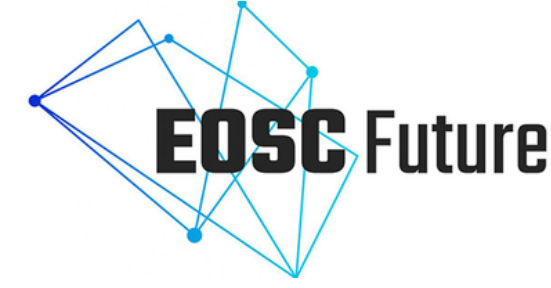
- Well-motivated dark matter models could have evaded detection!
- We can improve sensitivity to new physics by performing a more specific (*exclusive*) search

**Two projects within this TSP:**

1. Reinterpretation of inclusive resonance search in terms of dark matter mediators ✓
2. Exclusive  $Z'$ +MET analysis
  - New results expected soon!

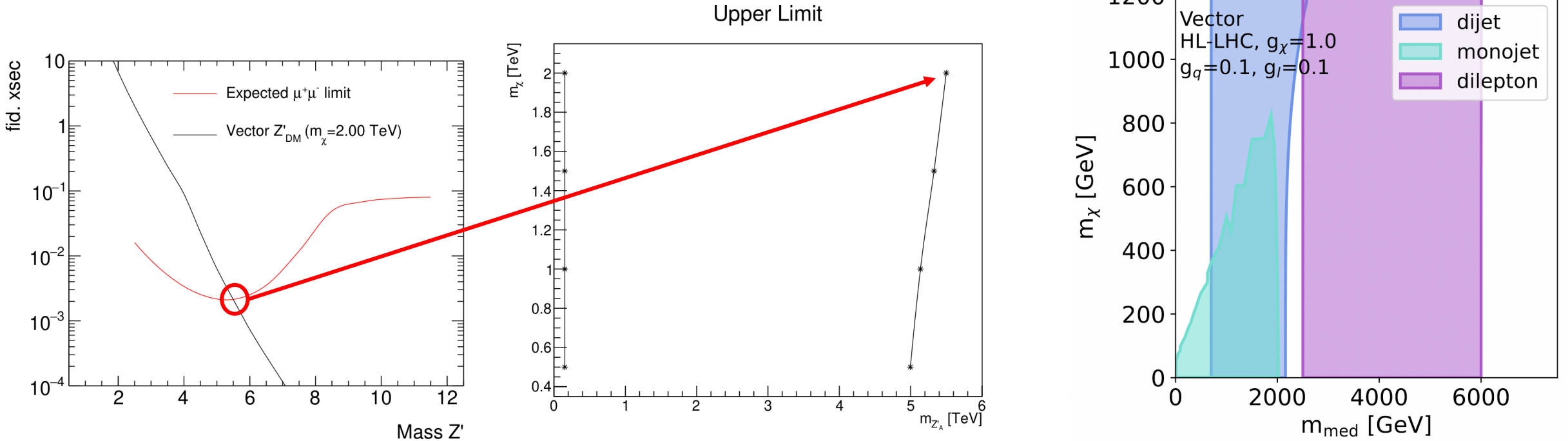


# Reinterpretation of the Resonance Search



Use the dilepton resonance search to constrain dark matter mediators.

- Assuming a non-zero coupling to leptons, a neutral mediator associated with a dark sector would produce an excess in the dilepton invariant mass distribution.

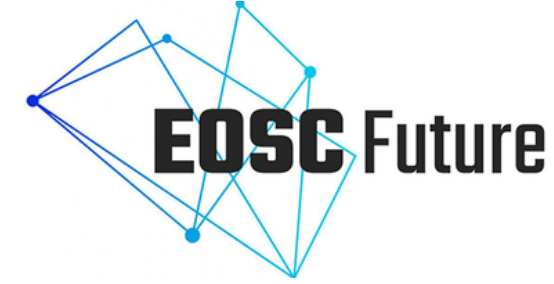
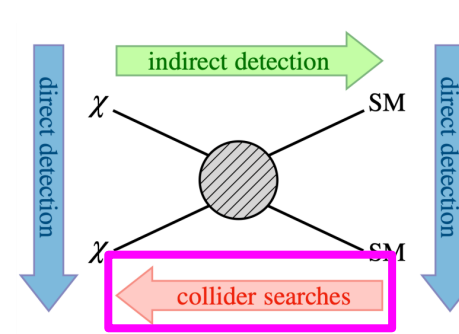


Results included in this paper: <https://arxiv.org/abs/2206.03456>  
 (prepared within the US prioritization effort "Snowmass")

# Reinterpretation of the Resonance Search

This reinterpretation was set up with REANA, sending the jobs to a remote computer.

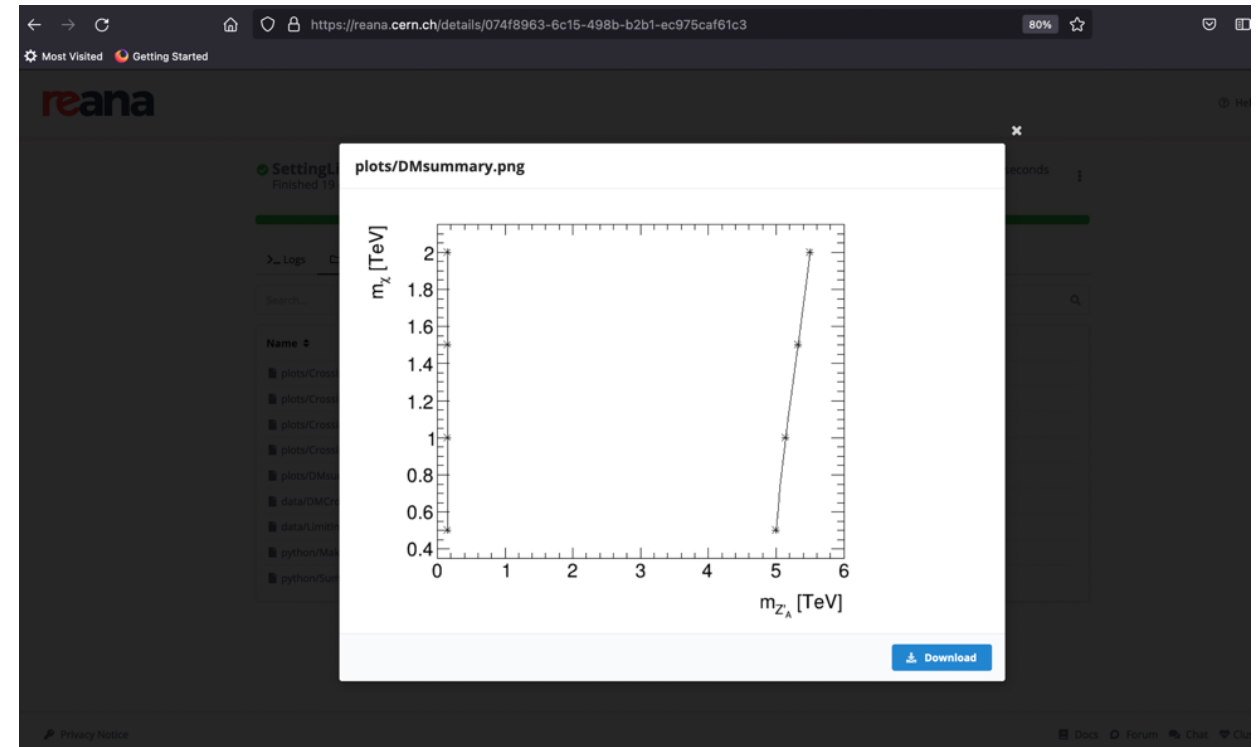
- Multiple stage workflows can be sent, passing the output to the following stage.



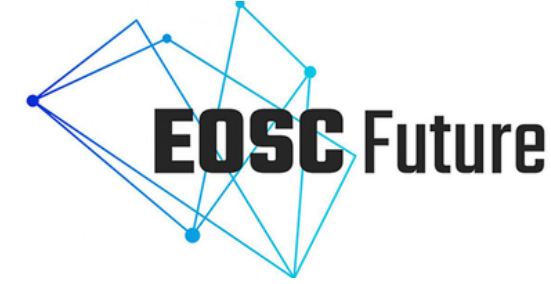
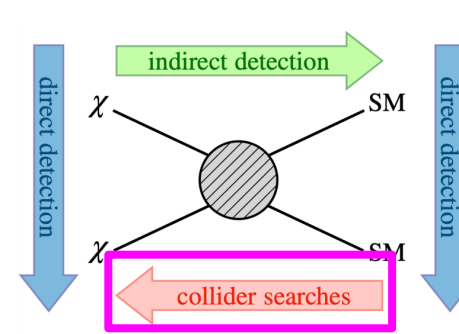
Name	Modified	Size
plots/Crossing_DM1p00_fsl.png	2022-06-13T18:54:15	11974
plots/Crossing_DM0p50_fsl.png	2022-06-13T18:54:15	12082
plots/Crossing_DM2p00_fsl.png	2022-06-13T18:54:15	12051
plots/Crossing_DM1p50_fsl.png	2022-06-13T18:54:15	12048
plots/DMsummary.png	2022-06-13T18:54:35	9128
data/DMCrossSectionGraphs_axial_massmass.root	2022-06-13T18:52:50	26404
data/LimitInterpolator_CL95_14TeV.root	2022-06-13T18:54:15	17439
python/MakeLimit.py	2022-06-13T18:52:50	8760
python/Summary.py	2022-06-13T18:52:50	2041

```

====> Verifying REANA specification parameters...
-> SUCCESS: REANA specification parameters appear valid.
====> Verifying workflow parameters and commands...
-> SUCCESS: Workflow parameters and commands appear valid.
====> Verifying dangerous workflow operations...
-> SUCCESS: Workflow operations appear valid.
====> Verifying compute backends in REANA specification file...
-> SUCCESS: Workflow compute backends appear to be valid.
SettingLimits: 27
====> SUCCESS: File /python/MakeLimit.py was successfully uploaded.
====> SUCCESS: File /python/Summary.py was successfully uploaded.
====> SUCCESS: File /data/DMCrossSectionGraphs_axial_massmass.root was successfully uploaded.
====> SUCCESS: File /python/MakeLimit.py was successfully uploaded.
====> SUCCESS: File /python/Summary.py was successfully uploaded.
====> SUCCESS: File /data/DMCrossSectionGraphs_axial_massmass.root was successfully uploaded.
====> SUCCESS: SettingLimits has been queued
jovyan@jupyter-little:~/atlas-dm-reinterpretations
    
```

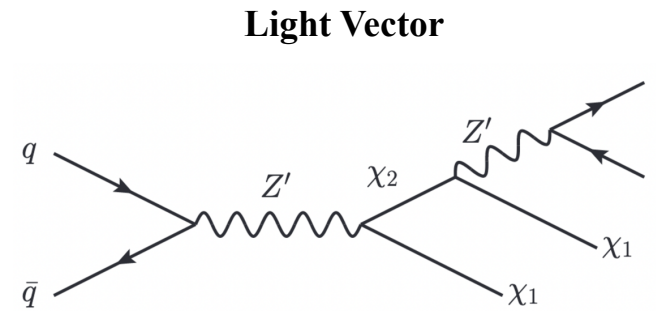
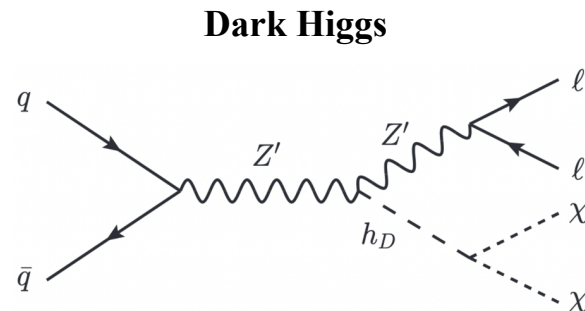


# Exclusive $Z'$ +MET Analysis



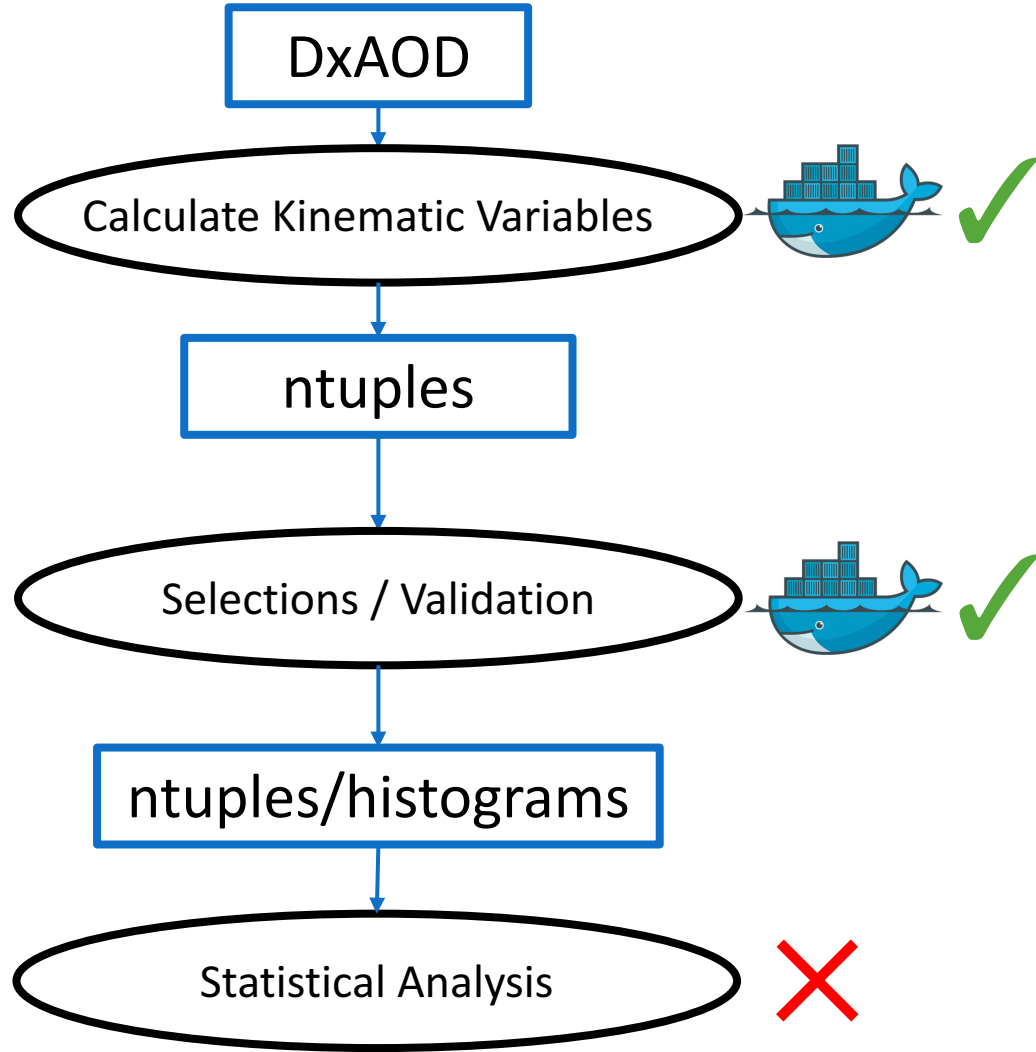
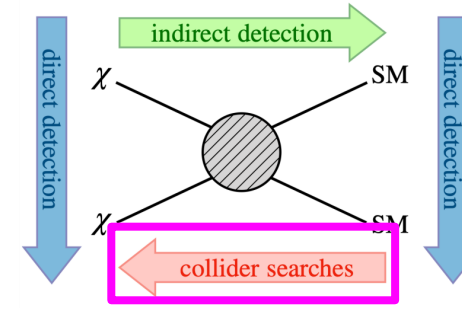
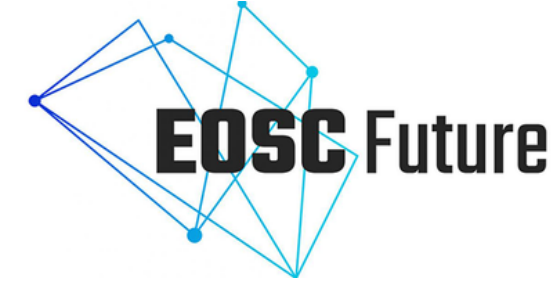
Search targeting dilepton resonances in the  $ll$ +MET final state.


- Searching for well-motivated models that could have escaped detection up to this point.
- Benchmark models help guide our analysis techniques, but we aim to stay as general as possible.
  - **Reproducible and reinterpretable results** are necessary for **collaboration**.
- By targeting dilepton events with MET in the final state, we will be more sensitive in the low-mass regions where the dilepton analysis was dominated by Standard Model events.

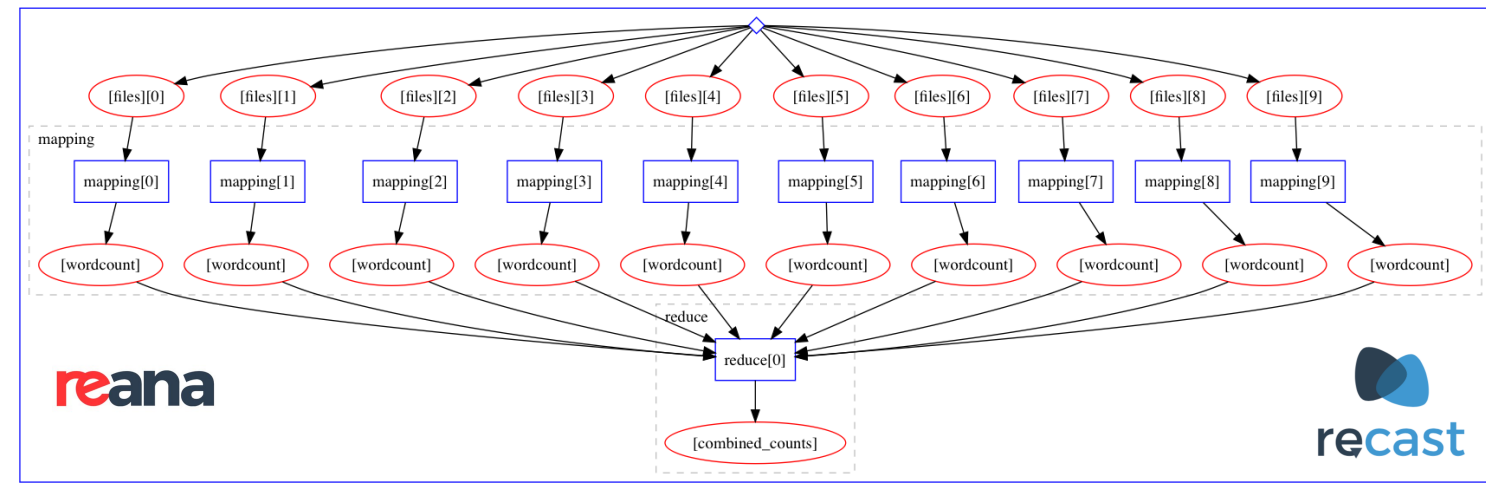


<https://arxiv.org/pdf/1504.01386.pdf>

# Exclusive $Z'$ +MET Analysis

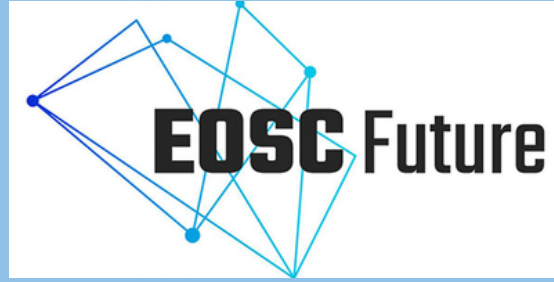


  $\geq 10 TB$





SAPIENZA  
UNIVERSITÀ DI ROMA



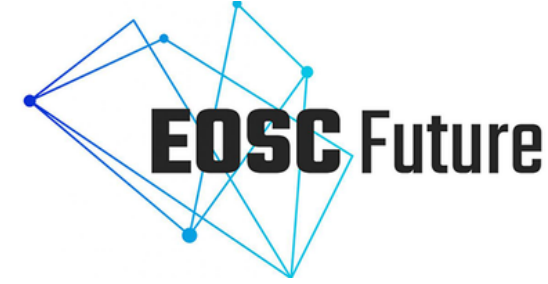
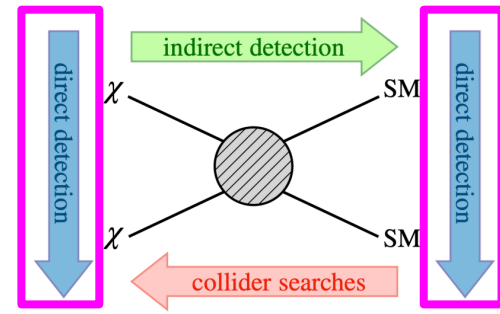
# Dark matter direct detection: DarkSide plans and results



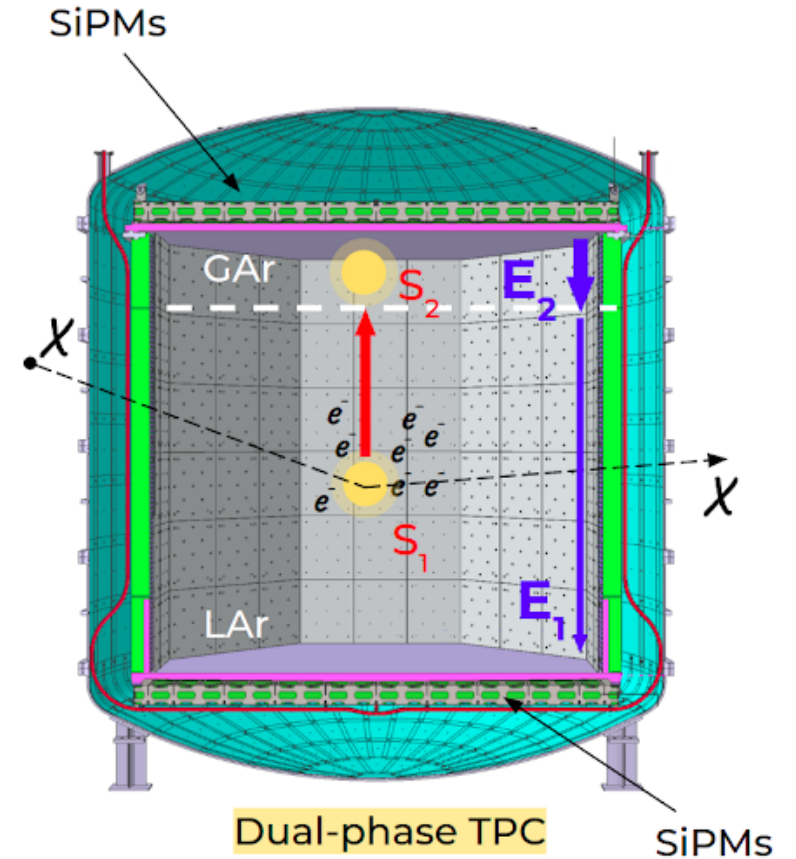
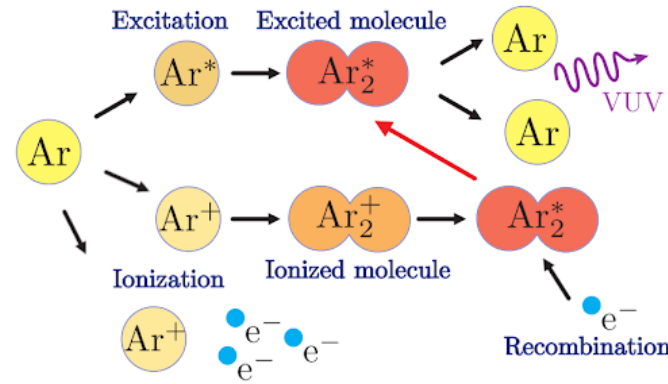
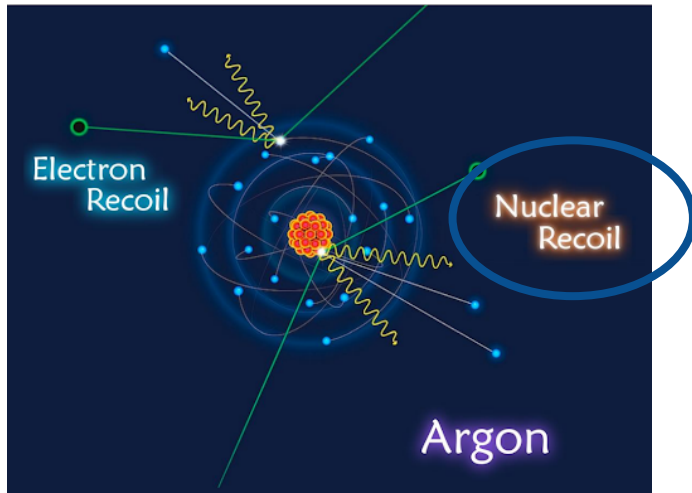
Maria Adriana Sabia (INFN/La Sapienza)  
Paolo Salomone (INFN/La Sapienza)  
Marco Rescigno (INFN)  
Valerio Ippolito (INFN)



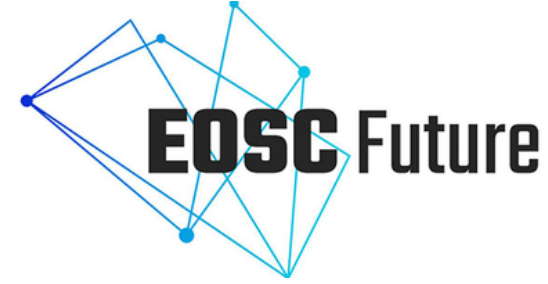
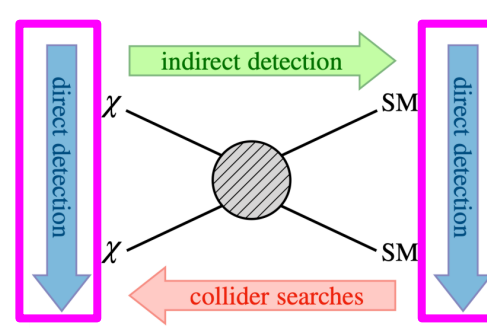
# Direct Detection with a LAr TPC



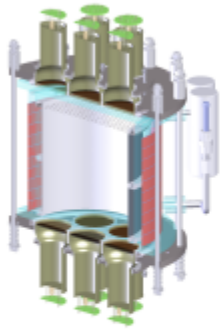
- DM as WIMP-like particle produces a **nuclear** or an **electron recoil**.
- Elastic scattering with Argon Nuclei results in **Scintillation & Ionization**.



# DarkSide Experimental Program

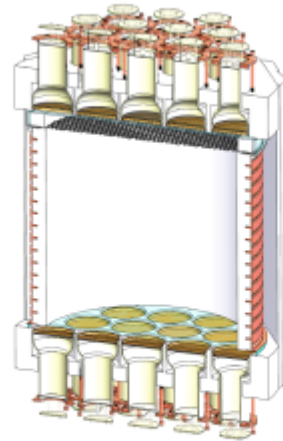


2012



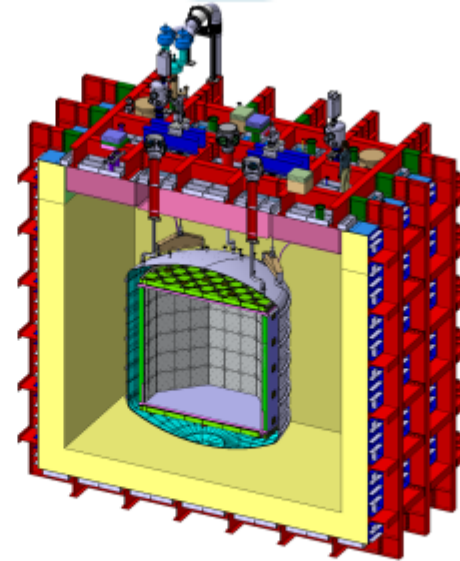
DarkSide-10

2013 - 2018



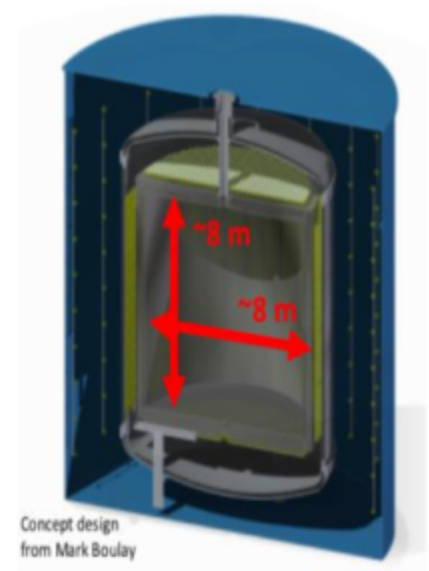
DarkSide-50  
16 t d

2026 - 2036



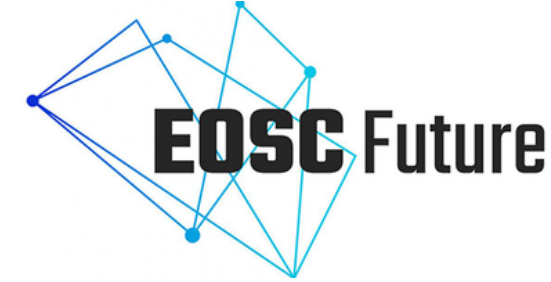
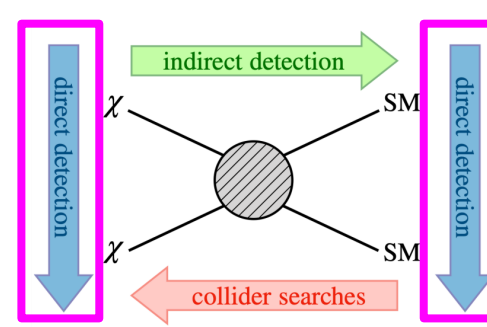
DarkSide-20k  
200 t yr

2030+



Argo  
3000 t yr

# DarkSide Plans



- Implemented a reanalysis tool for a high-mass search on the **VRE platform**.
  - Output: **DarkSide50** exclusion curve for WIMP-nucleon cross section.
- Low mass analysis to be implemented.
- Different **theoretical models** (WIMP halo, argon response...) can be inserted by the user to produce different limit results.
- Working towards first open implementation.



# Indirect dark matter search with gamma rays and its association with VRE platform via open-science tools

Pooja Bhattacharjee

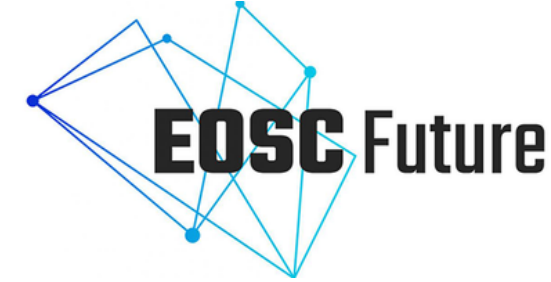
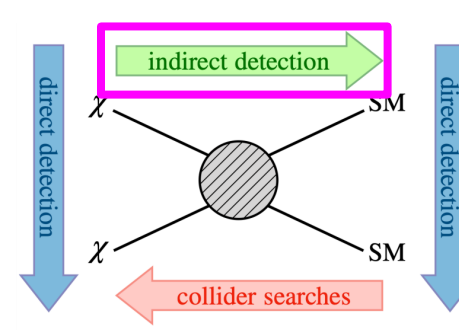
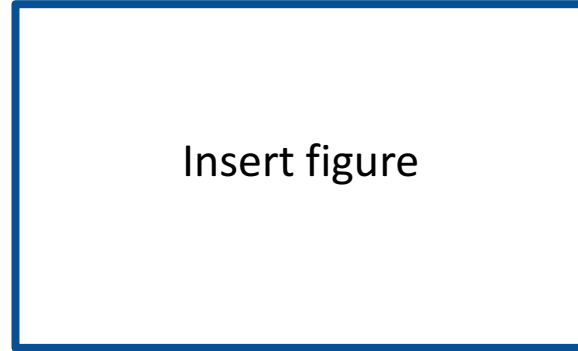
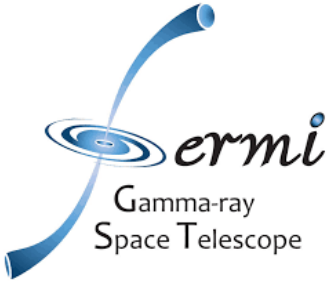
Laboratoire d'Annecy De Physique Des Particules (L.A.P.P)

Supervised by:

Francesca Calore

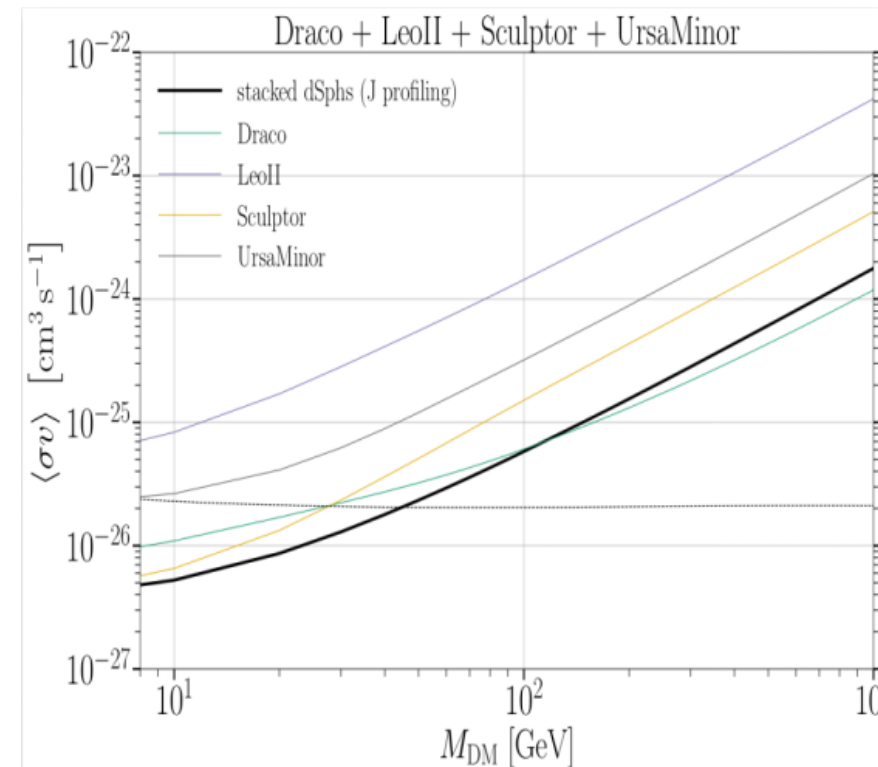
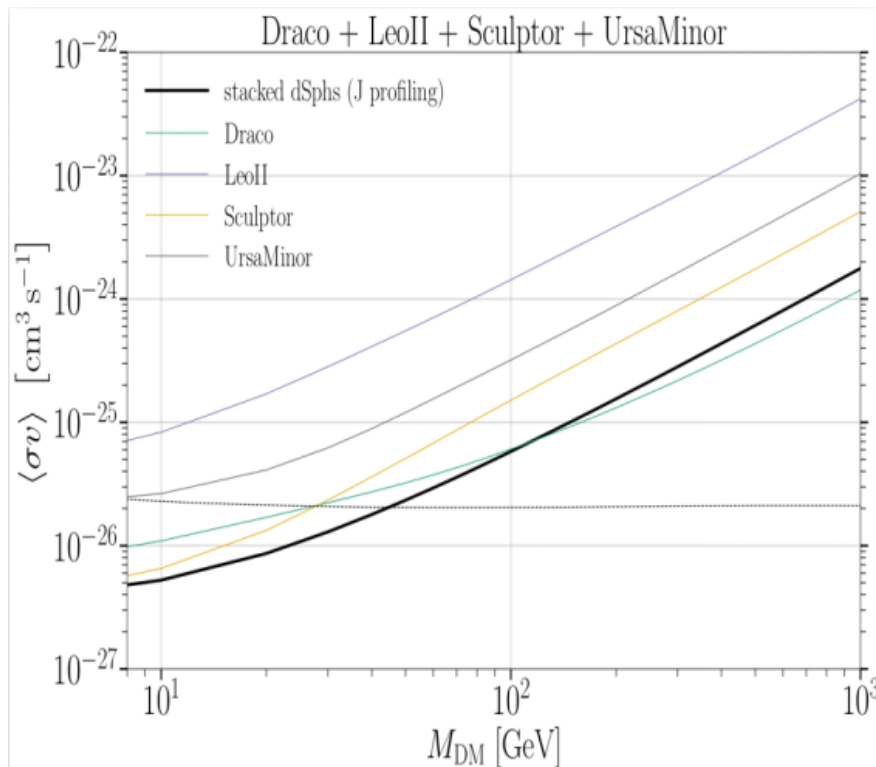
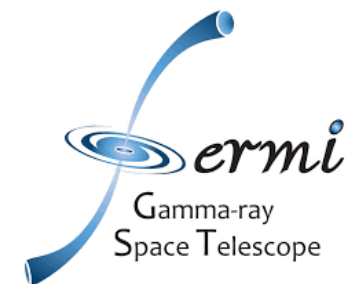
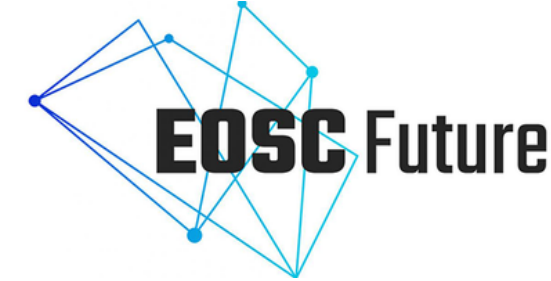
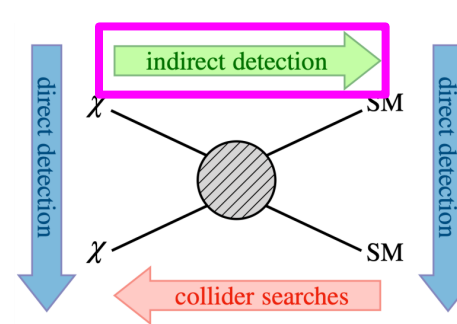
Laboratoire d'Annecy-le-Vieux de Physique Théorique (L.A.P.Th)

# Code Structure: MLFermiDwarfs



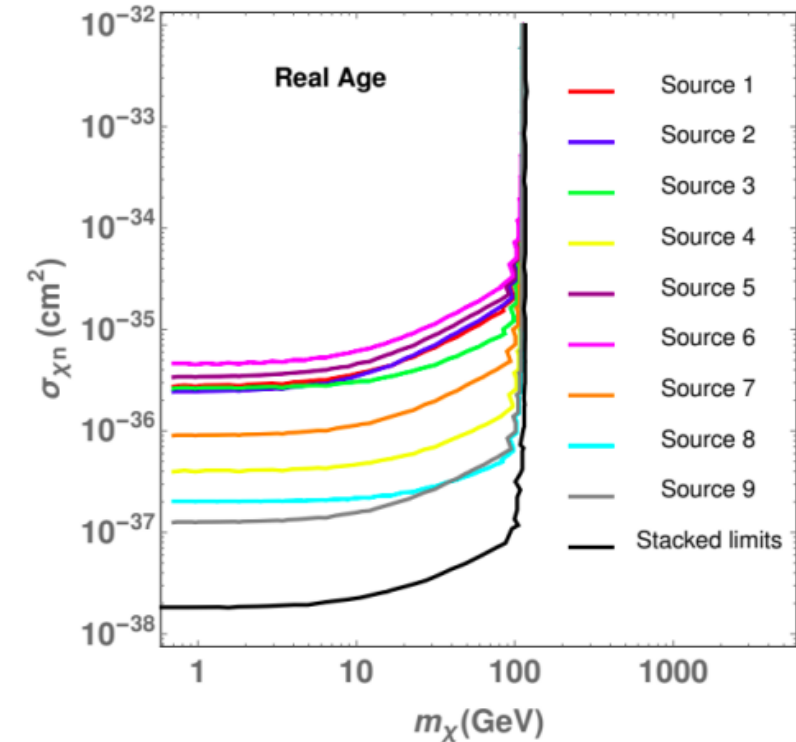
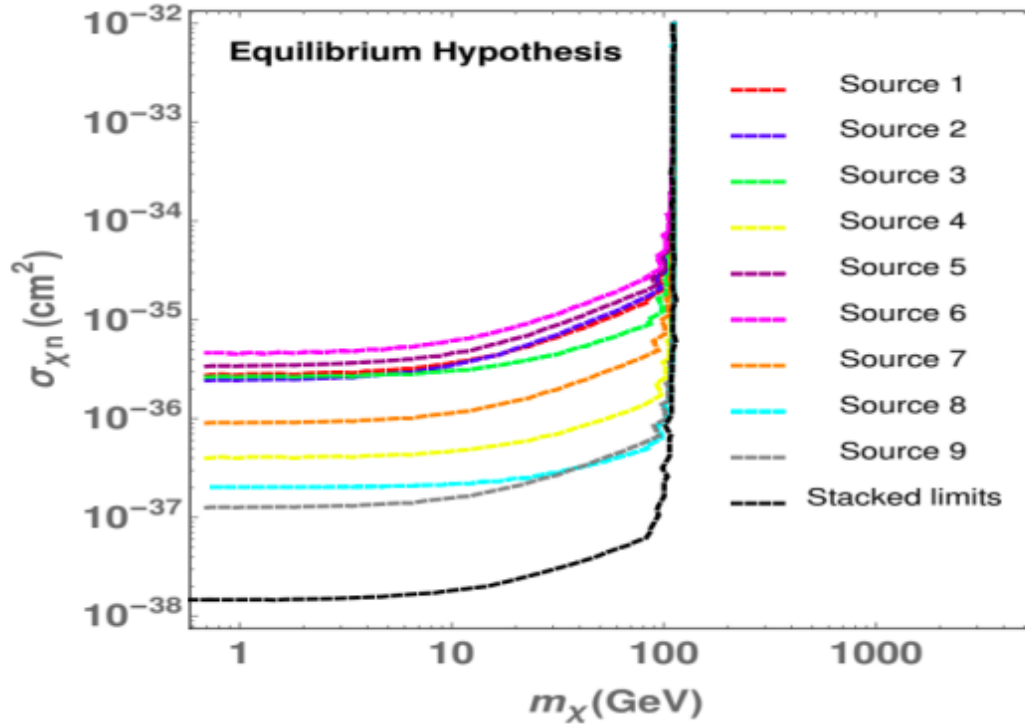
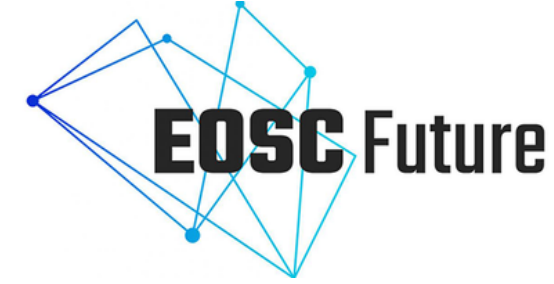
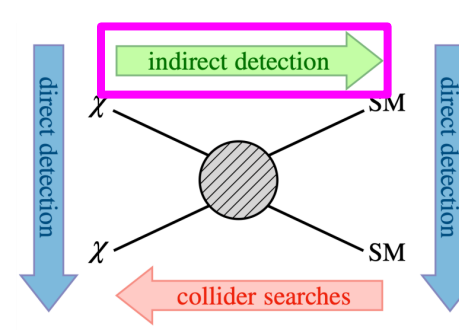
- Gamma-ray data from the Fermi Large Area Telescope (Fermi-LAT).
- The data and main processing software (Fermi Science Tools) are publicly accessible, and now fully available in the VRE.
- Code is entirely written in python 3 using well-known packages like scikit-learn.
- Package can be optimized from the command line enabling a quick check of the viability of a user-defined Dark Matter model.

# MLFermiDwarfs



- Based on the Fermi-LAT data set used for paper **A. Alvarez et al. JCAP09 (2020) 004**.
- MLFermiDwarfs code is accessible from <https://gitlab.in2p3.fr/escape2020/virtualenvironment/mlfermilatdwarfs>

# Brown Dwarf Analysis



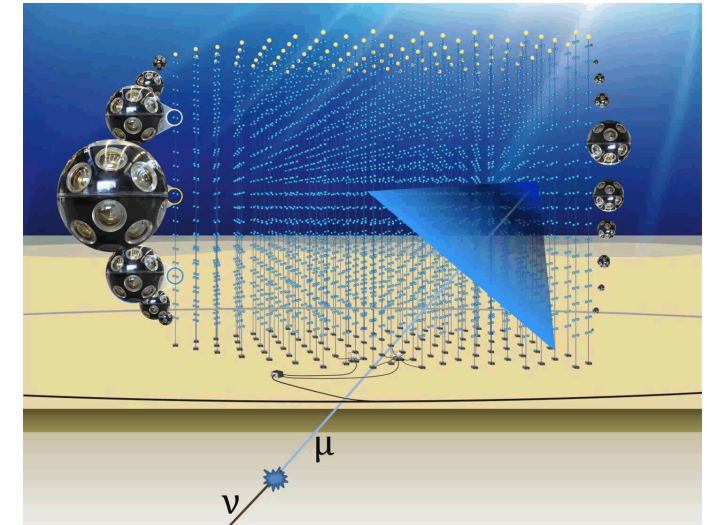
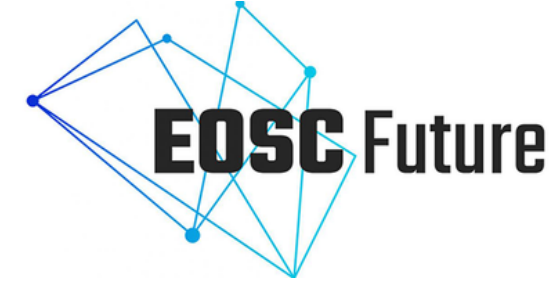
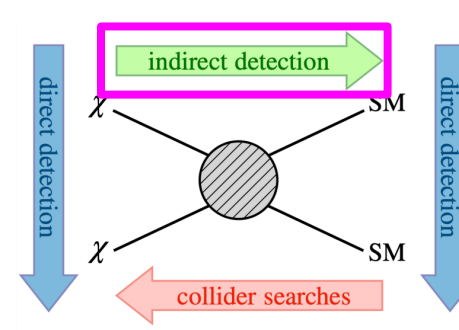
- Based on the recent Published paper on **Bhattacharjee et.al, PRD,107, 043012, 2023.**
- Code is accessible from <https://gitlab.in2p3.fr/escape2020/virtual-environment/brown-dwarfs-gamma>

# Instrument Response Function of KM3Net for point-source analysis

Mikhail Smirnov  
(Friedrich-Alexander University FAU-ECAP)

Supervised by:  
Kay Graf  
Friedrich-Alexander University FAU-ECAP

# IRF Concept



• Instrument **R**esponse **F**unction of neutrino telescope provides a quantitative estimation of the event rate and the background rate

• It allows to avoid extensive MC simulations each time for a new configuration of neutrino source

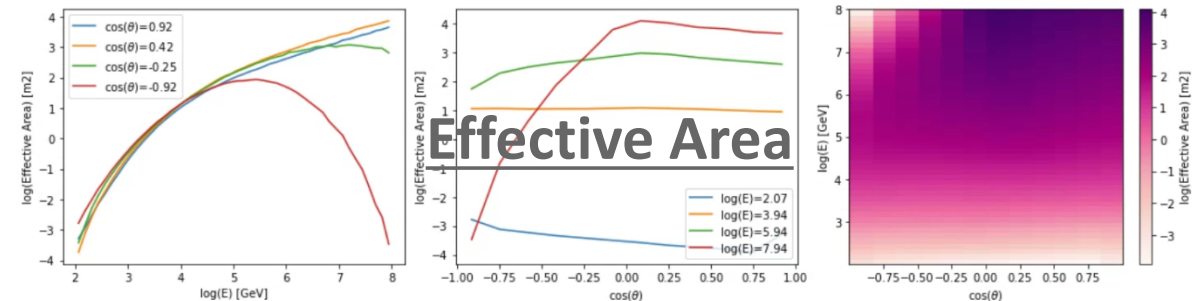
• It supports different configurations of neutrino sources:

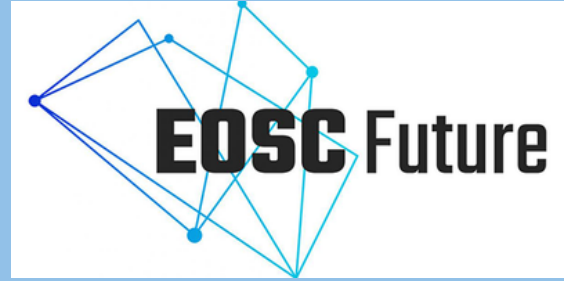
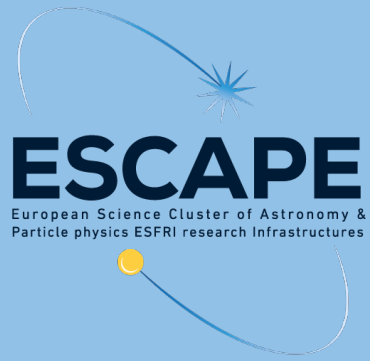
- Point source with power law  $E^{-a}$
- Diffuse source
- Extended source

• Compatibility with [gammapy](#) will give an easy combination with other gamma experiments like CTA

• Active development of the [km3irf](#) python package

```
pip install km3irf
from km3irf import utils
new_plot = utils.DrawAeff()
new_plot.peak()
```





# Summary



ESCAPE - The European Science Cluster of Astronomy & Particle Physics ESFRI Research Infrastructures has received funding from the European Union's Horizon 2020 research and innovation programme under Grant Agreement no. 824064.

## Upcoming challenges for DM TSP:

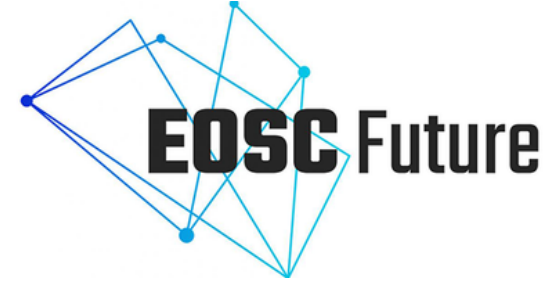
- Onboard new analyses requiring:
  - very **large amounts of data**, o(TB) → turn to pledged resources
  - more **complex workflows** → stress-test VRE & EOSC cell, using REANA
- Guarantee **restricted data access** until embargos lifted → use EOSC core **authentication**
- Publish software and pipelines on **OSSR**
- Use Gambit software for **combination** of results (subject to time/resources)
- Expand use cases to **real-time analysis on constrained infrastructure** (subject to time/resources)

## Consolidation (common across DM and Extreme Universe Science Project)

- Widening participation of **scientists to Open Science tools**
  - more support needed for combination of results, training, documentation
- Consolidation work on EOSC for **lasting infrastructure**
  - Integration of services with EOSC core
- **Strengthening cooperation and sharing experience** across Science Projects



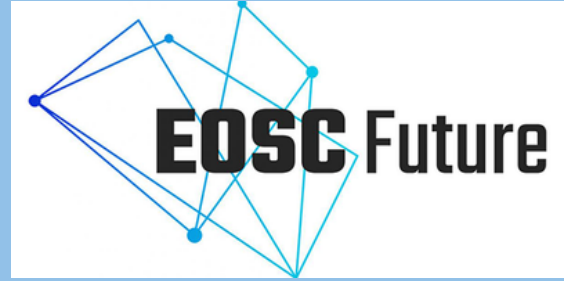
# Conclusions



With the **DM Science Project's analyses and tools** on the **VRE**, we are making progress towards:

- production of **new scientific results** discovering or constraining dark matter hypotheses
  - providing other communities with the necessary understanding to reproduce the analysis
- **comparing and combining** results from different experiments
- demonstrating **FAIR data and interoperable workflows** as an example for the community
- building a **working prototype cell for the European Open Science Cloud**
  - providing a testing ground for software & computing that can be explored by future experiments

**Thanks to the Dark Matter SP Team!**



# Backup



ESCAPE - The European Science Cluster of Astronomy & Particle Physics ESFRI Research Infrastructures has received funding from the European Union's Horizon 2020 research and innovation programme under Grant Agreement no. 824064.

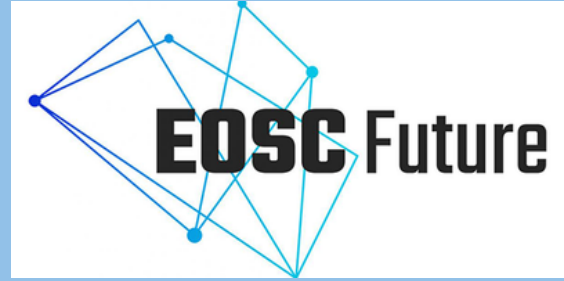
# Machine learning tools for big data compression

Axel Gallén, Alexander Ekman  
(Lund University)

Supervised by:  
Caterina Doglioni  
University of Manchester and Lund University

# Insert if any

Link to talk at CHEP



# Extreme Universe

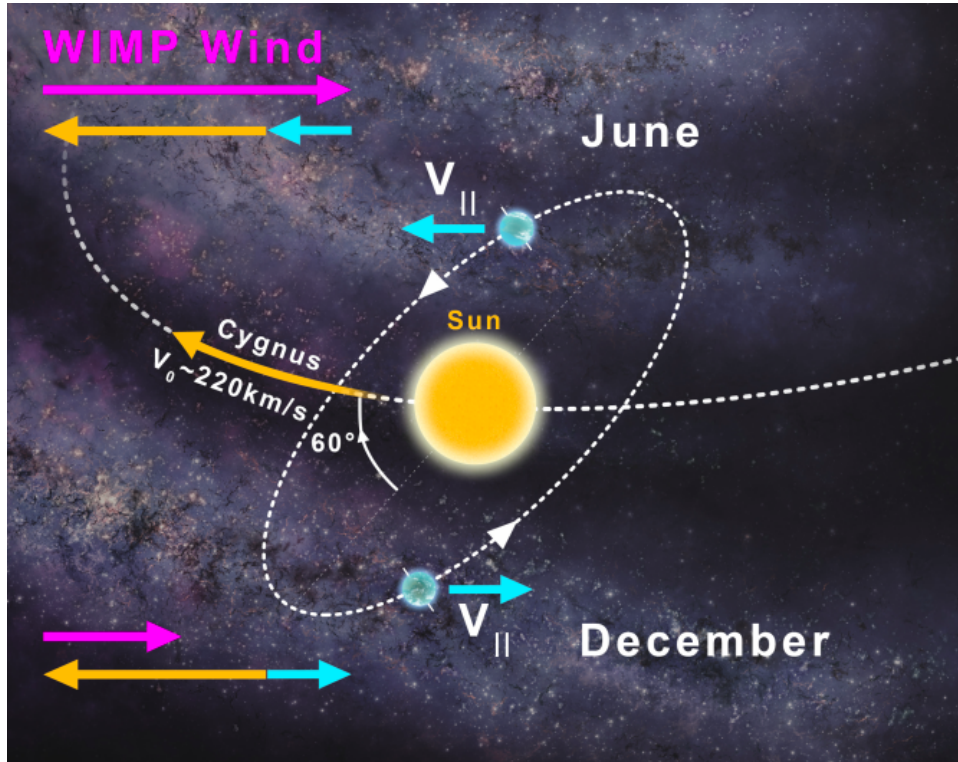
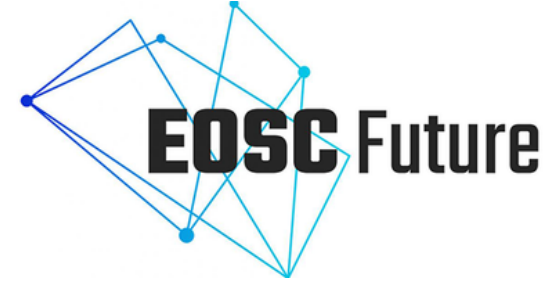
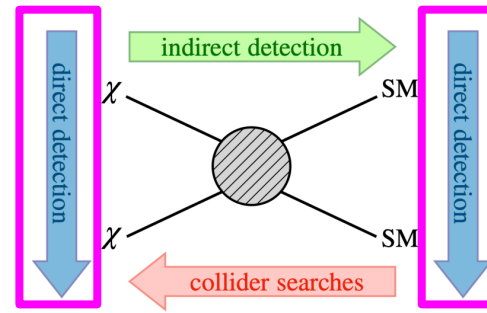


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1-2 slides for EU SP?

# Dark Matter Direct Detection

## WIMP Rate and Cross Section limits



$$\frac{dR}{dE_r} = \frac{MT}{2m_W \mu_N^2} \times \underbrace{\sigma_{Wn}}_{\text{Particle physics}} \times \frac{\mu_N^2}{\mu_p} A^2 \times \underbrace{F^2(E_r)}_{\text{Nuclear physics}} \times \underbrace{\rho_0 \times \int_{v_{\min}}^{v_{\max}} \frac{f(\vec{v})}{v} d^3v}_{\text{Astrophysics}}$$

