

Seismology of Core-Mantle Boundary Region of the Earth

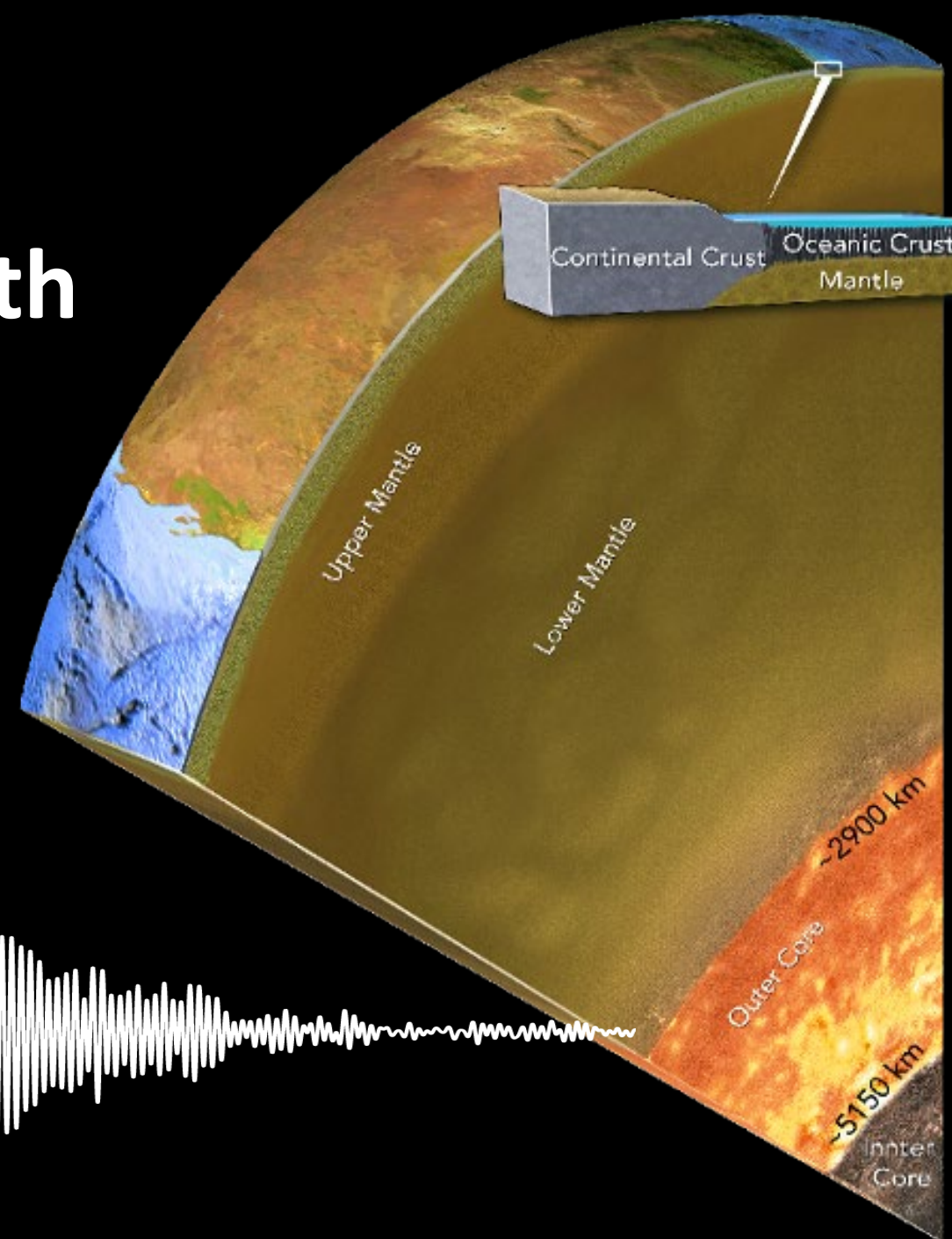
Doyeon Kim

Institute of Geophysics

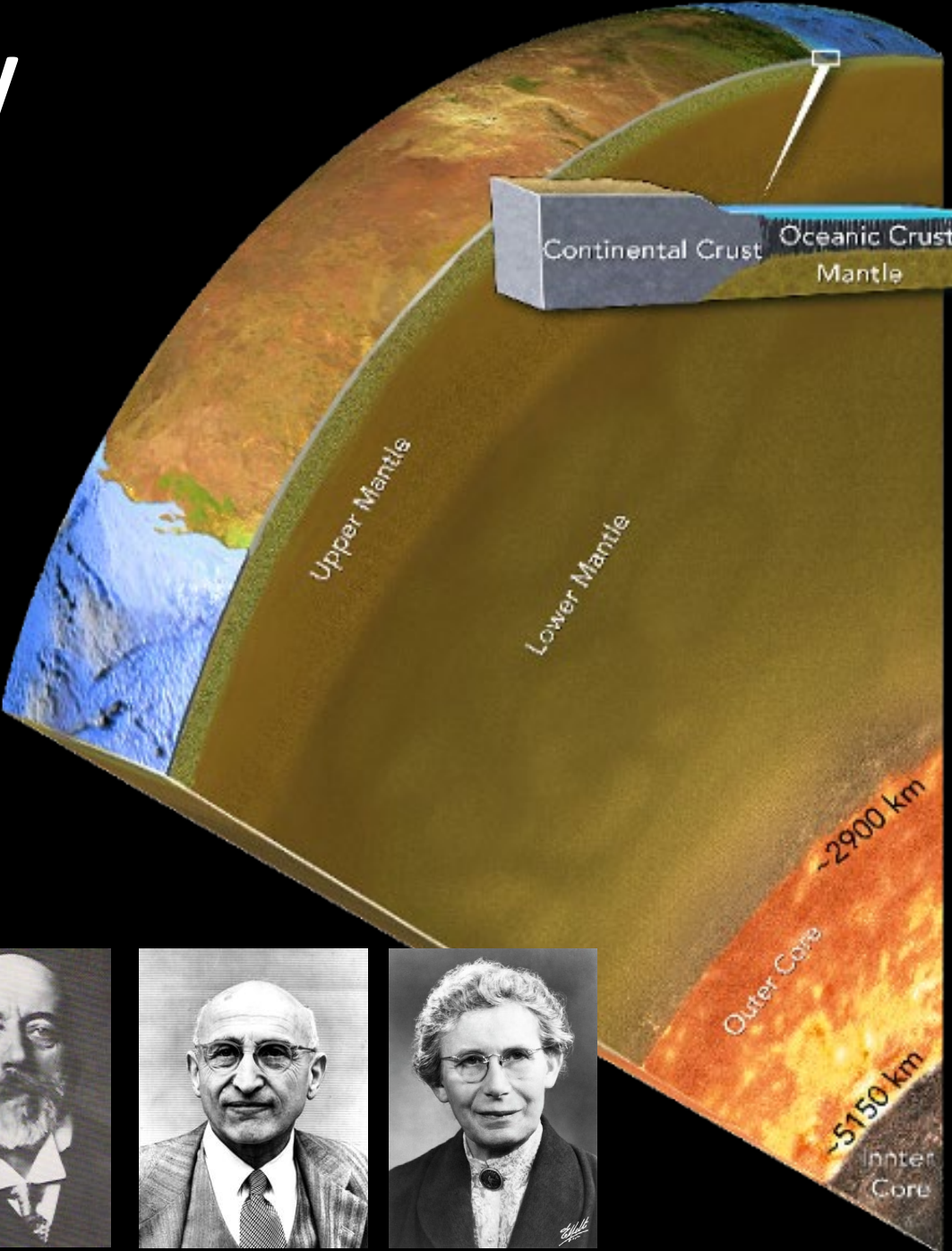
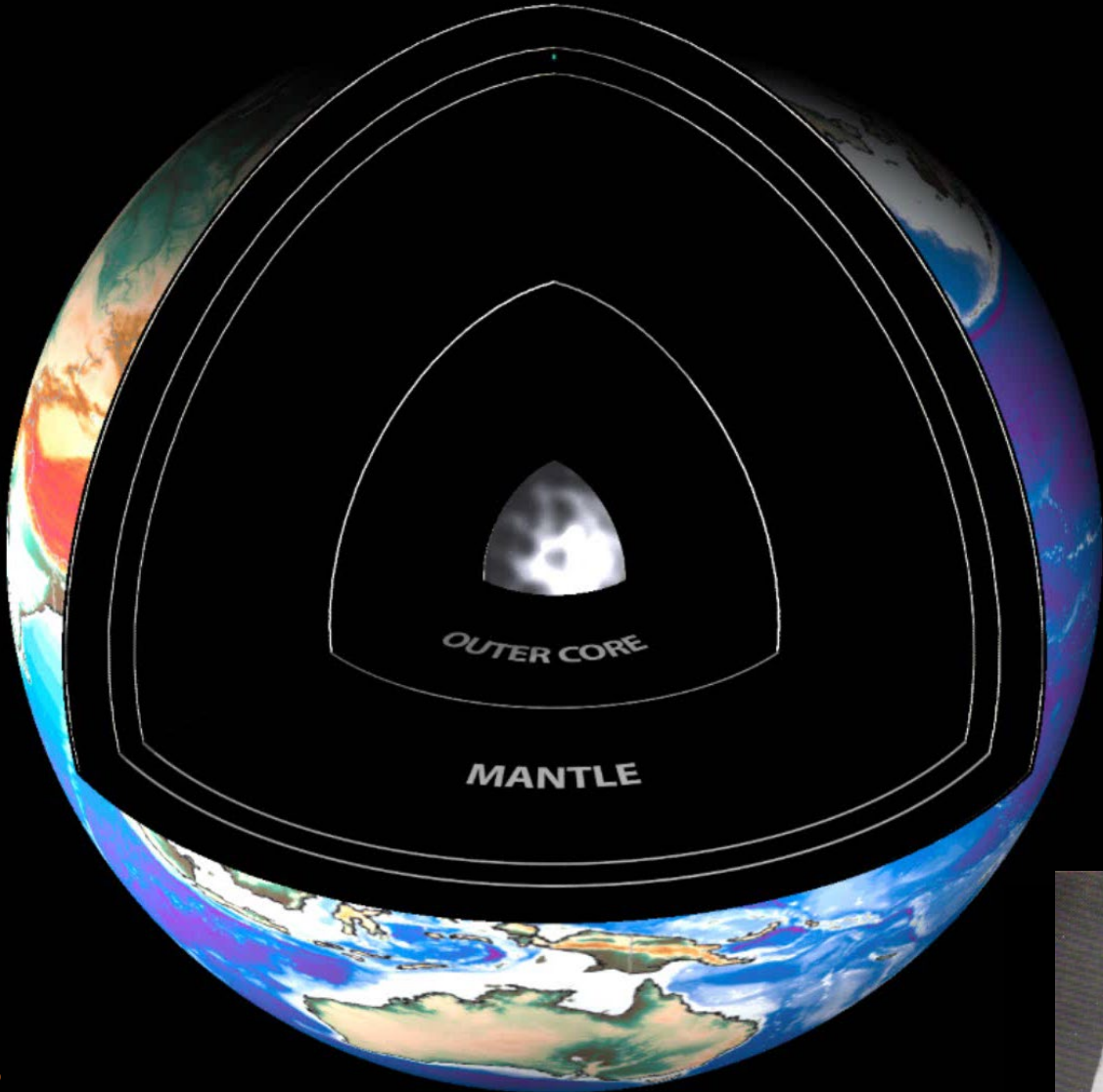
ETH Zurich

Imperial College London (Oct 1)

Multi-Messenger Tomography, Paris, 2023



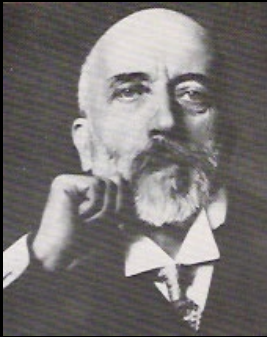
Layers of the Earth / Seismology



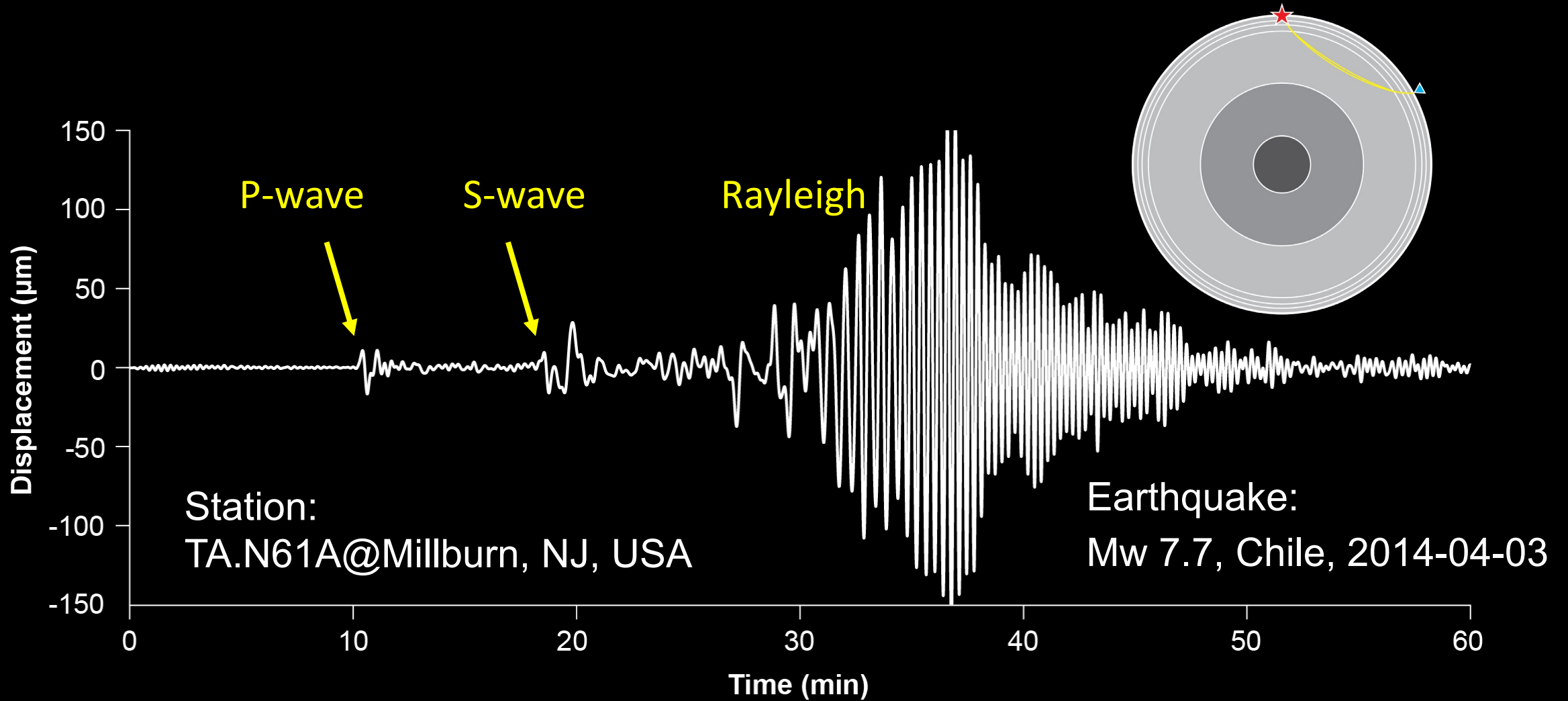
P-waves

S-waves

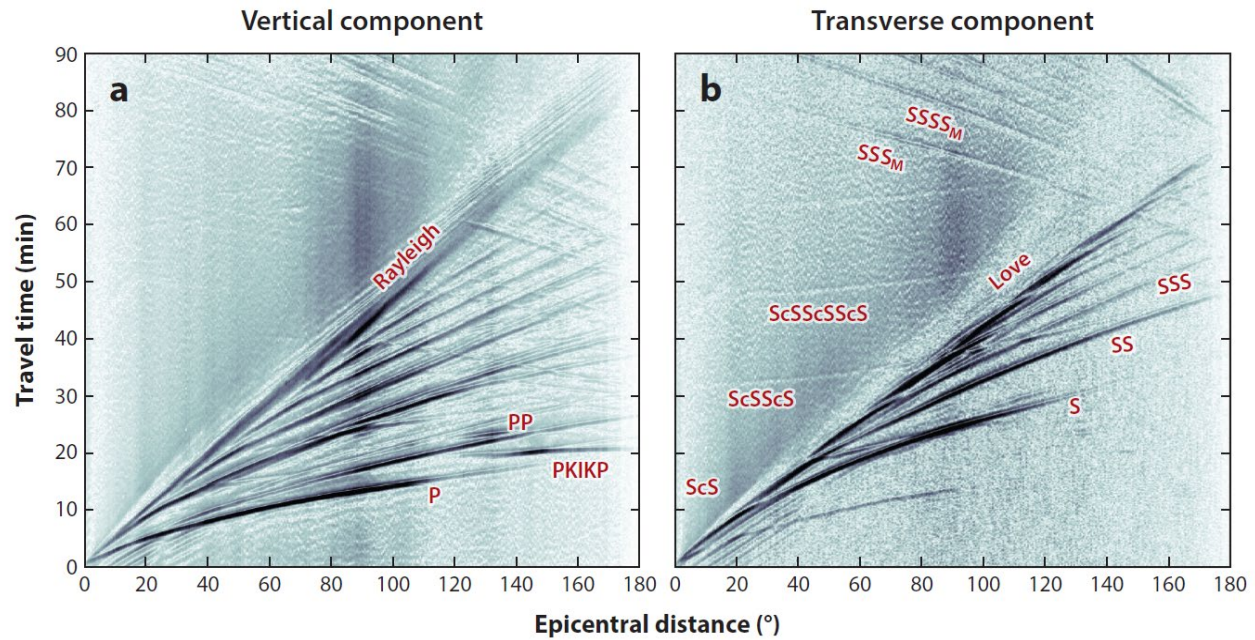
Movie by M. Thorne



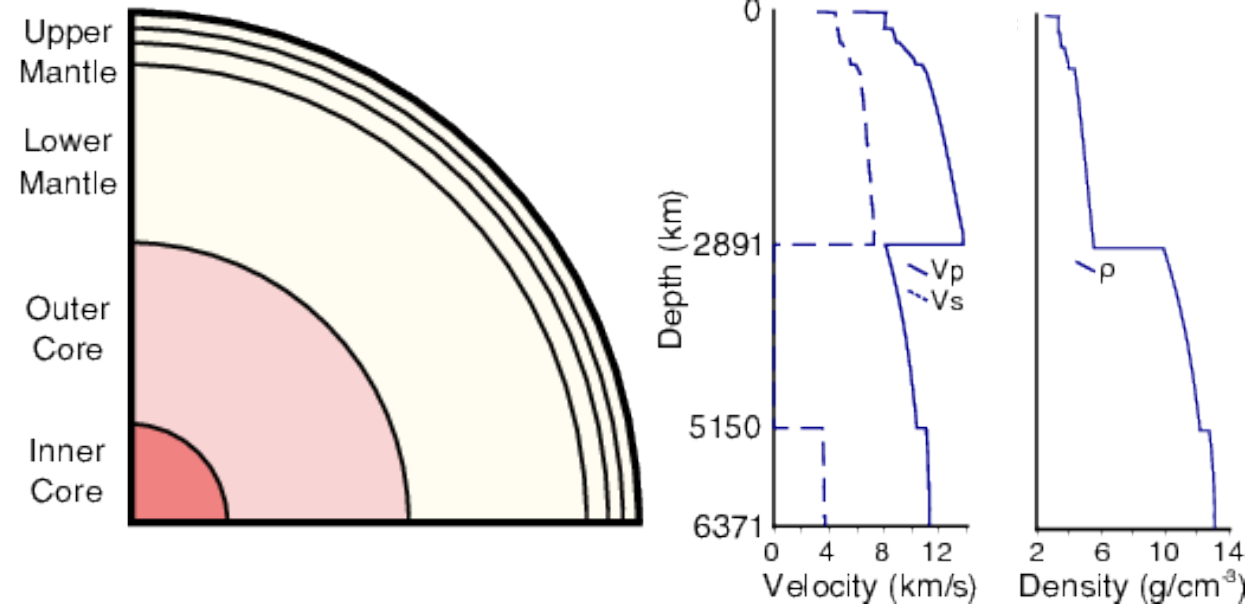
Seismogram from a single station



1-D Seismic Structure with Depth

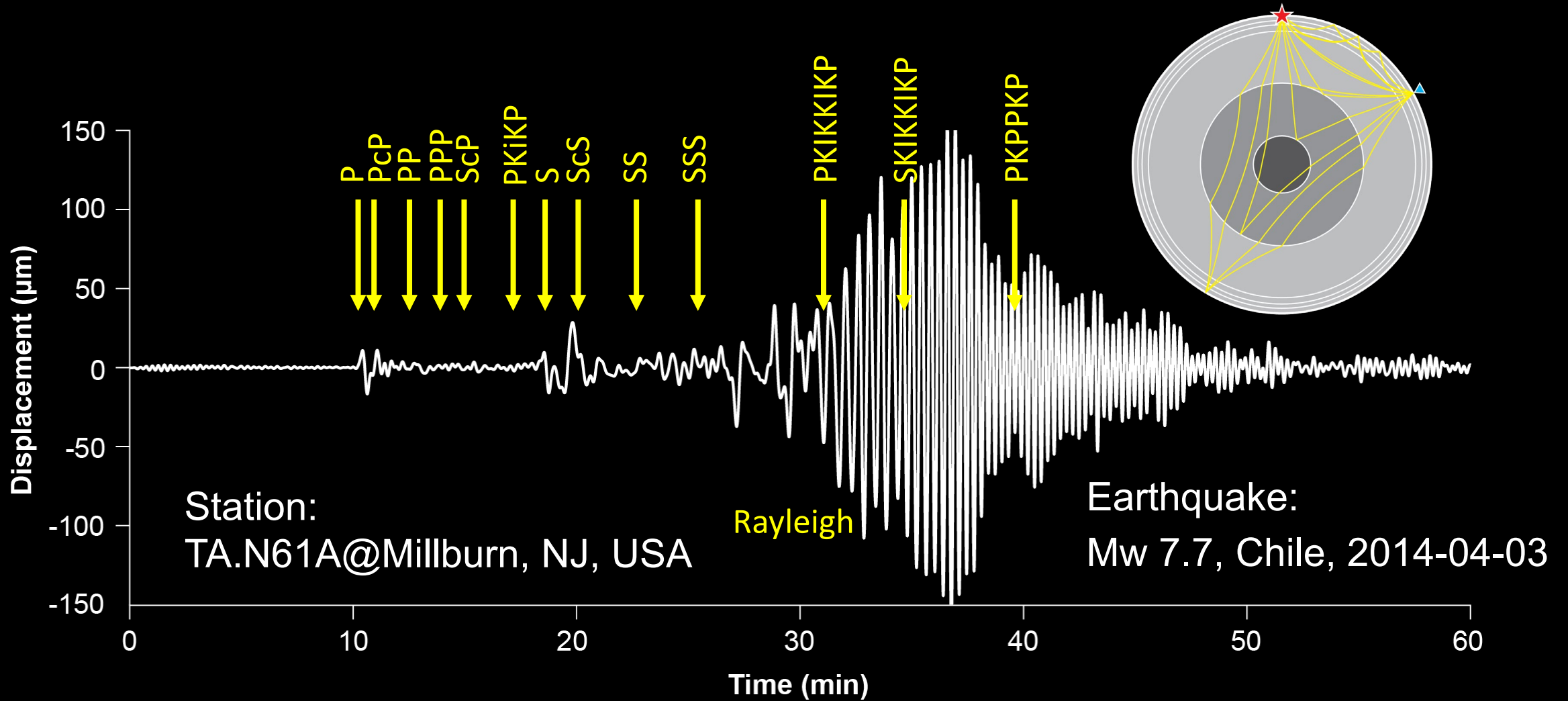


Traveltimes of seismic waves depend primarily on distance

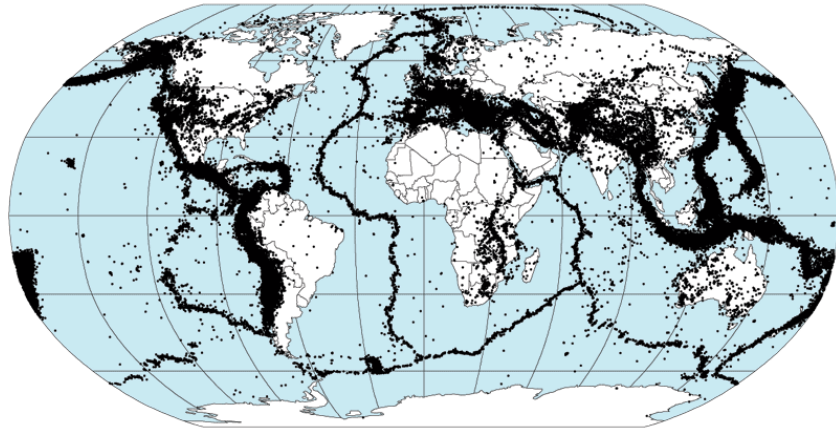


Velocity (and density) variations in the Earth depend primarily on depth

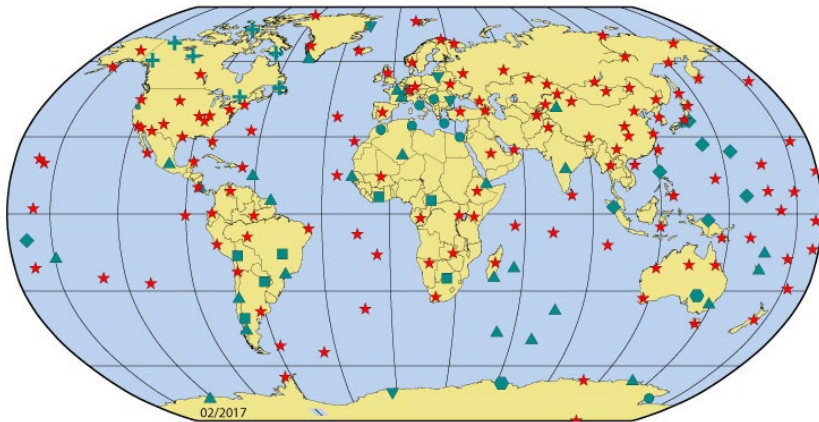
Seismogram from a single station



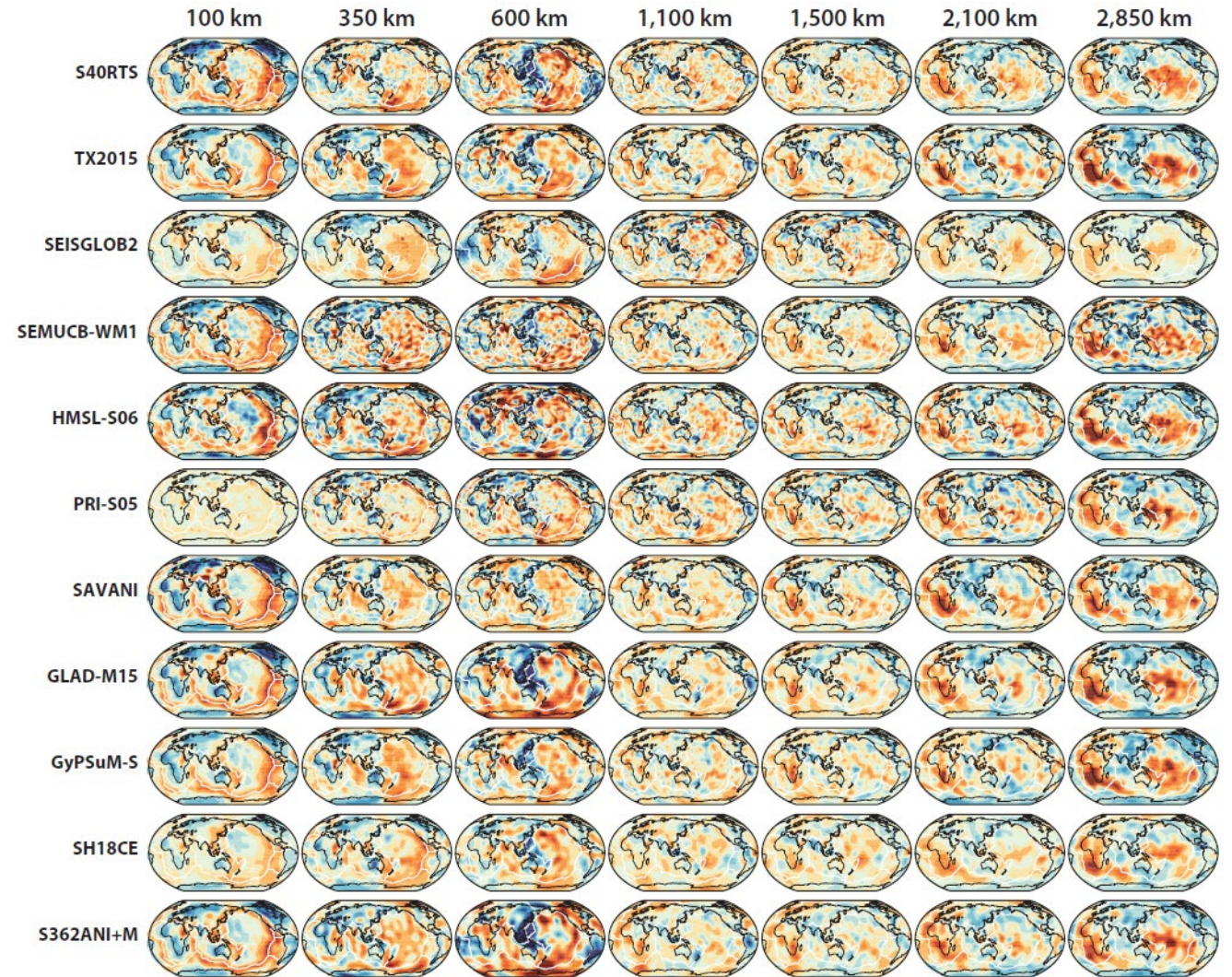
Earthquakes, Seismometers, 3D Earth Structures



We record and locate thousands of earthquakes each year

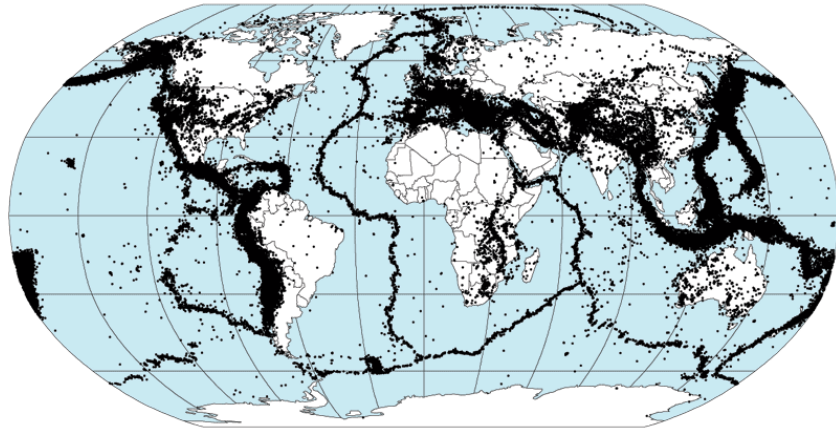


Global seismic networks provide open high-quality data

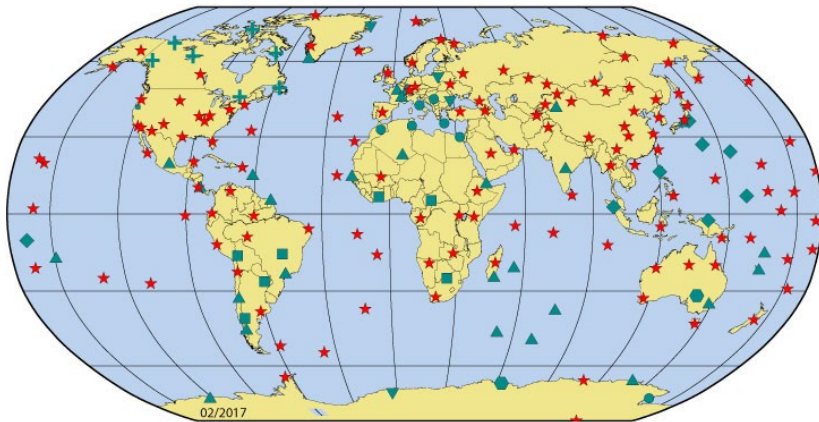


3D Earth structures from tomographic models

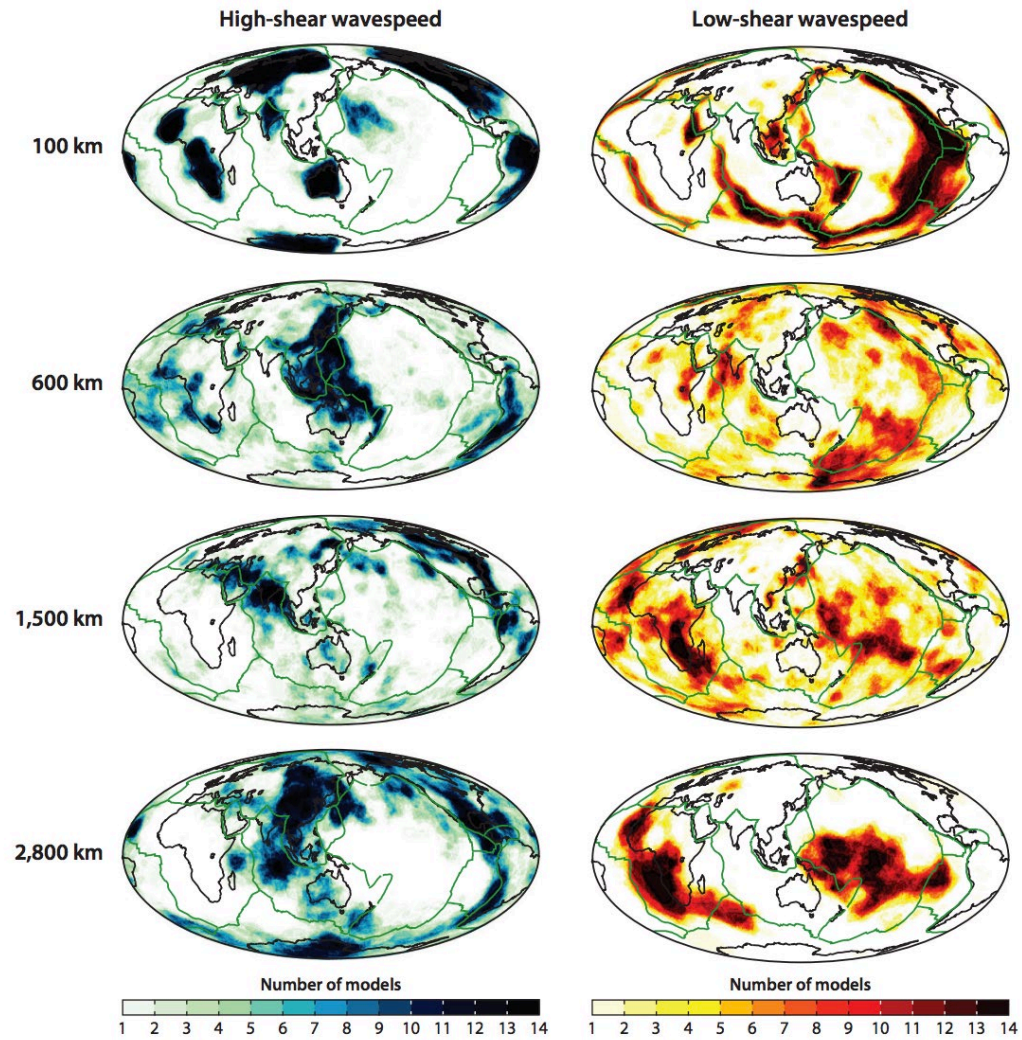
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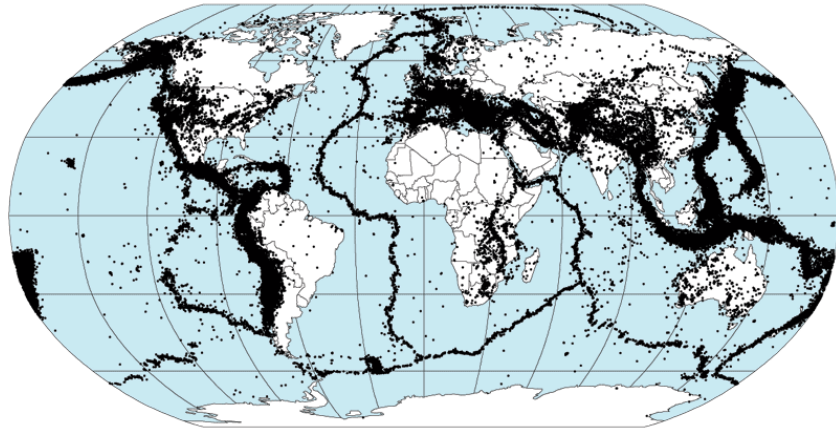
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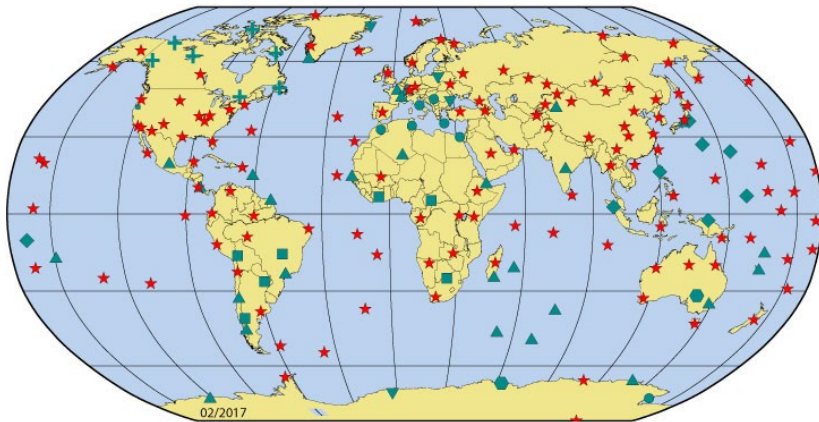
3D Earth structures from tomographic models

Ritsema & Lekic, AREPS (2020)

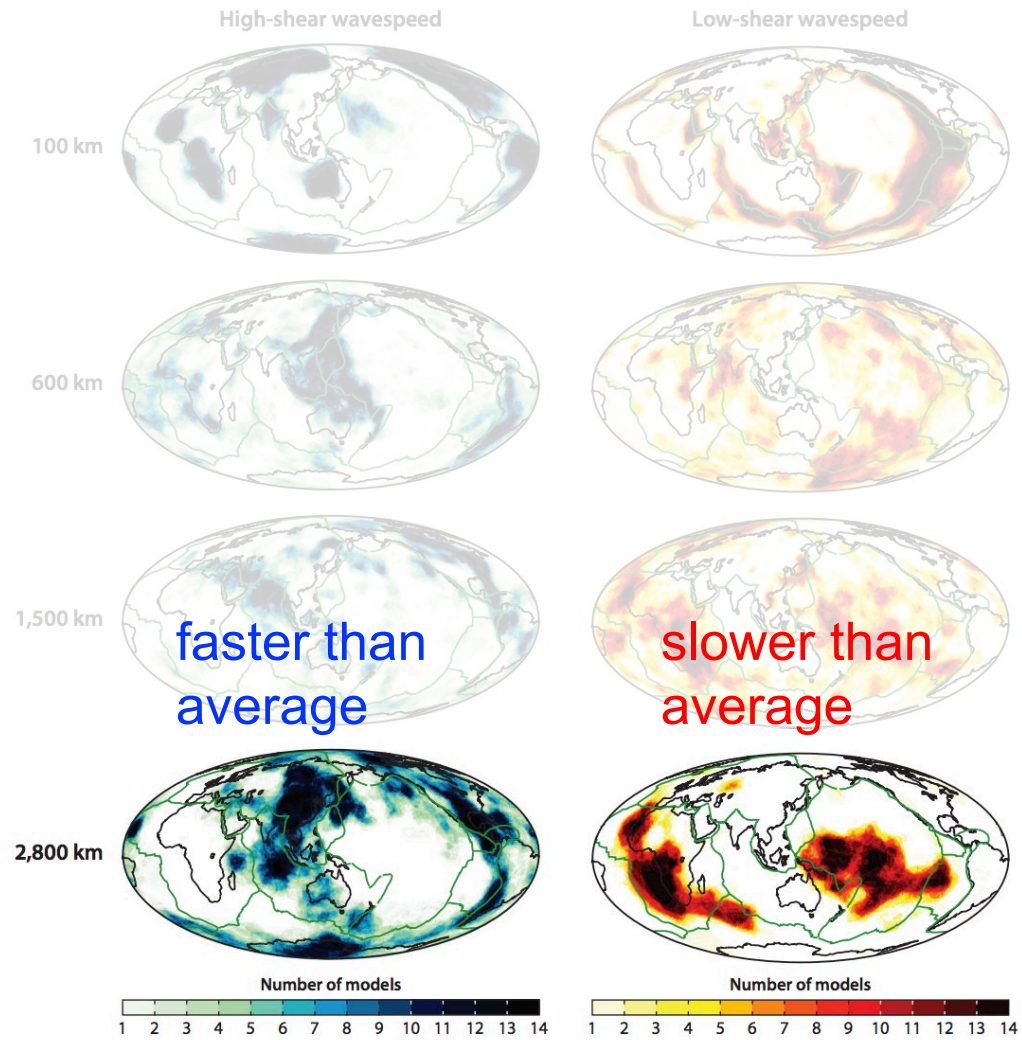
Earthquakes, Seismometers, 3D Earth Structures



We record and locate thousands of earthquakes each year



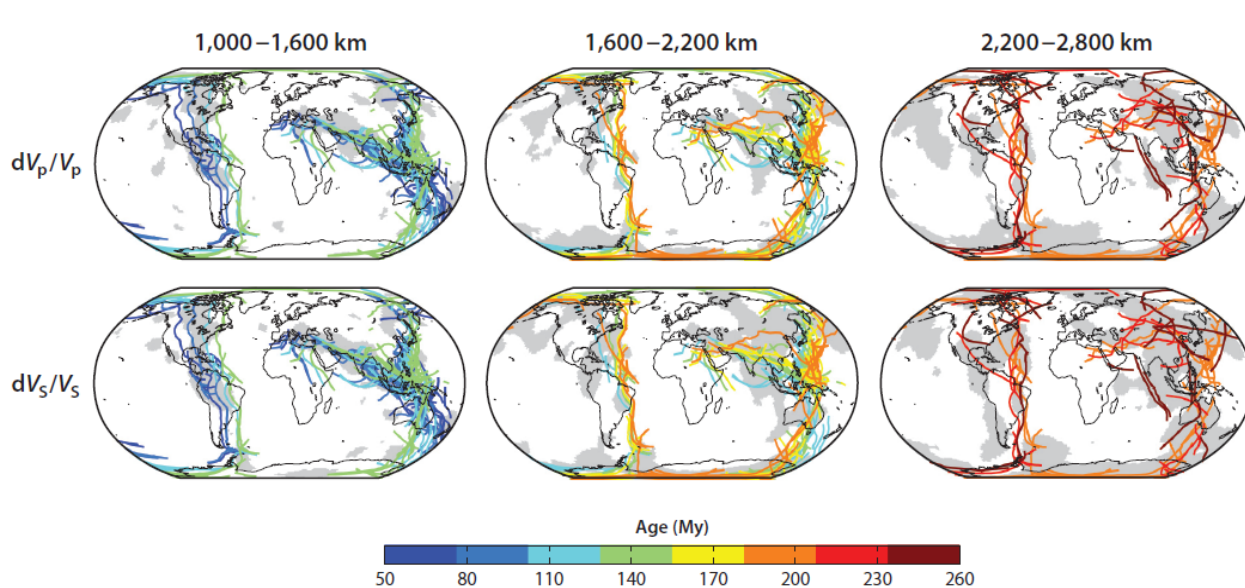
Global seismic networks provide open high-quality data



3D Earth structures from tomographic models

Ritsema & Lekic, AREPS (2020)

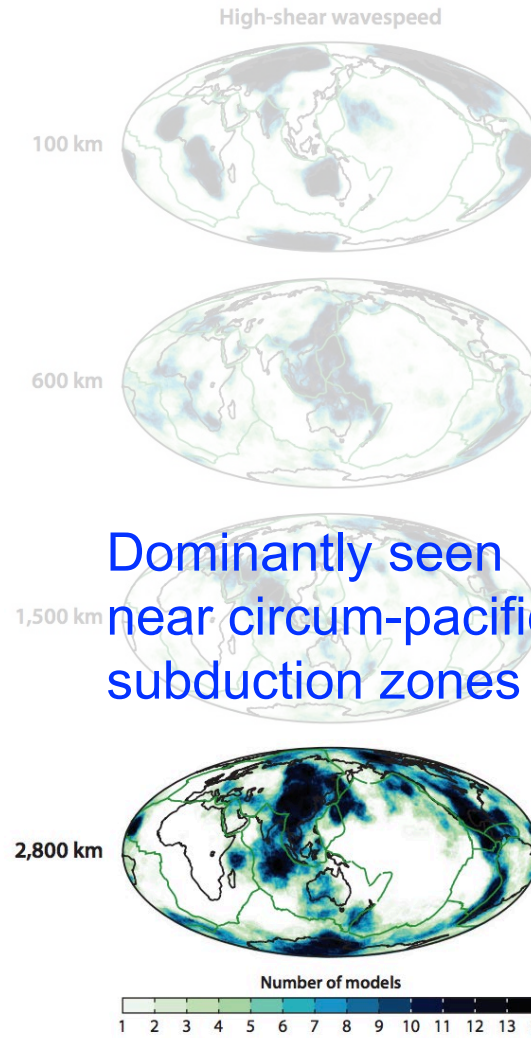
Large-scale Structures [~ 5000 km]: fast anomaly



Young et al, Nature (2019)

Subducted slabs in P- and S-wave models

Excellent correspondence between reconstructed locations of subduction zones and fast regions in lower mantle



Dominantly seen near circum-pacific subduction zones

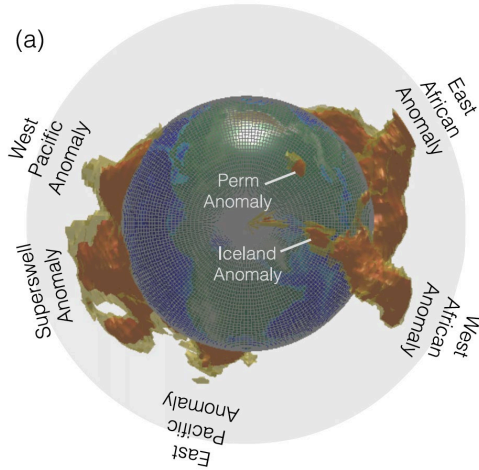
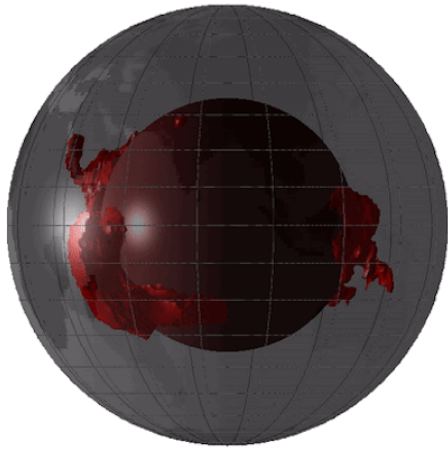
-Tomographic models reveal slabs of oceanic lithosphere subducting through the mantle

-Consistent between P- and S-wave models

-Well explained by past 200 Myr subduction history

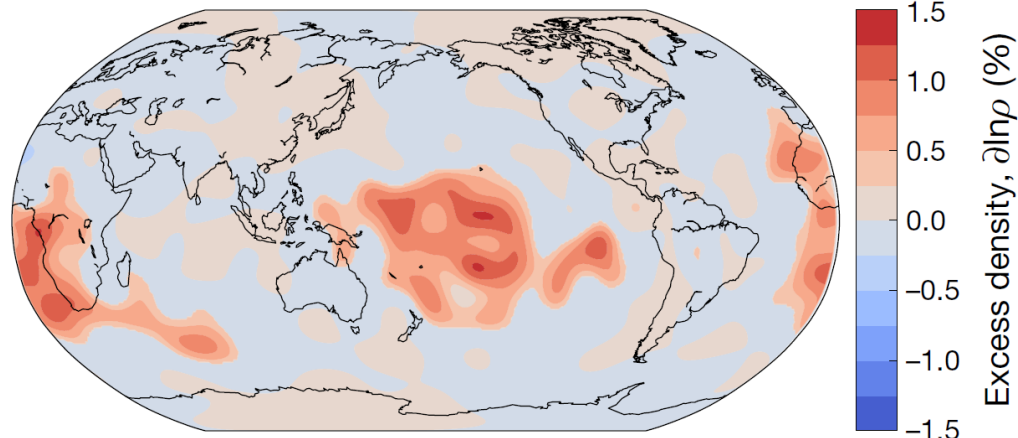
-When did the subduction begin? How far back to go? 130Ma? 240Ma?

Large-scale Structures [~ 5000 km]: **slow anomaly**



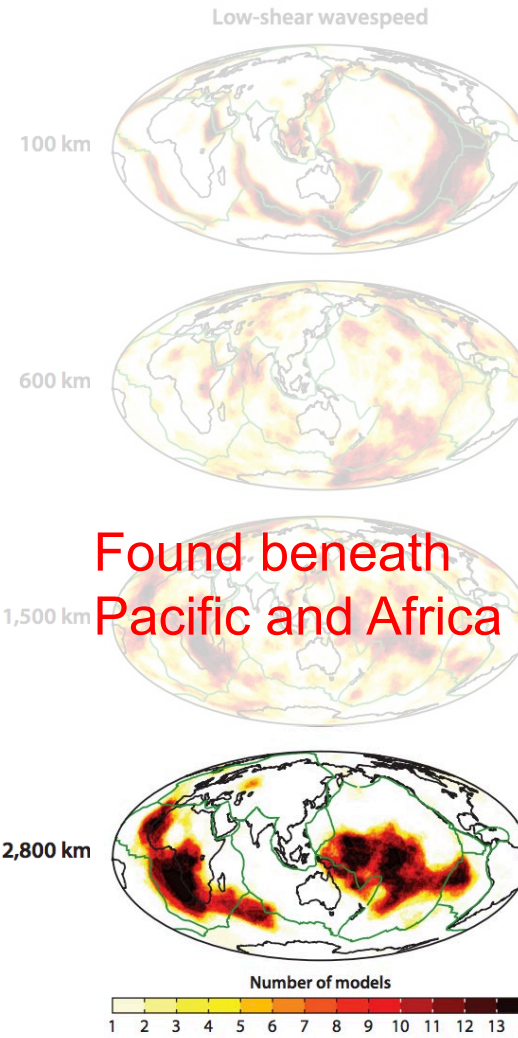
Cottaar & Lekic, *GJI* (2016)

Large low shear velocity provinces
($\sim 9\%$ of whole mantle mass)



Dense but seismically slow regions

Lau et al, *Nature* (2017)



Found beneath
Pacific and Africa

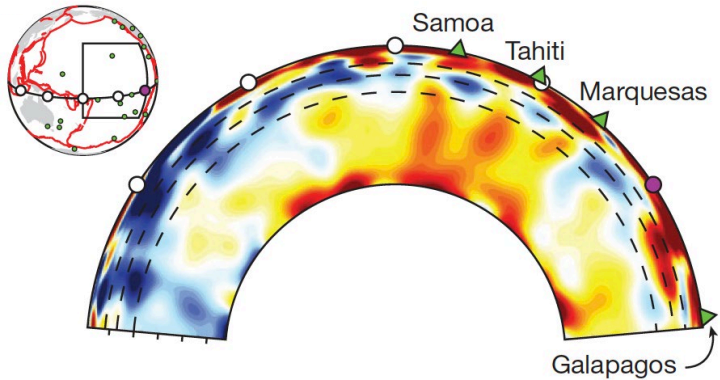
Extend no more
than ~ 900 km
above the CMB

-Traditionally viewed as mantle upwelling regions, “Super plumes”: isochemical?

-More recently observed in P-wave models / sharp edges support a compositional heterogeneity ($\sim \pm 3-5\%$ out of $\sim 2\%$)

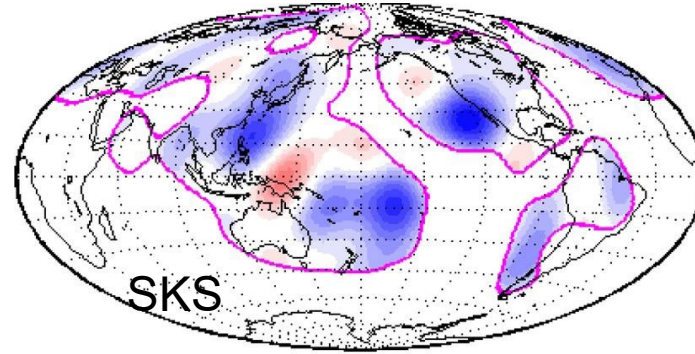
-Geophysical constraints (tidal response, normal modes, dynamic topography, CMB ellipticity, gravity) prefer denser material ($\sim 1\%$) at the base (~ 200 km) (e.g., Ishii & Tromp 1999; Trampert et al., 2004; Lau et al., 2017; Richards et al., 2023)

Large-scale Structures [~ 5000 km]: complexities

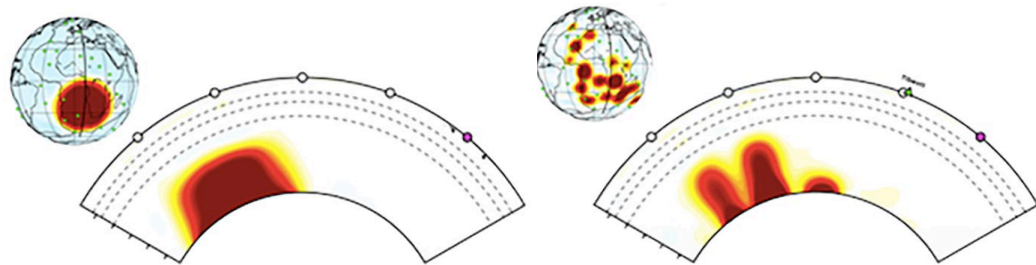


French & Romanowicz, Science (2015)

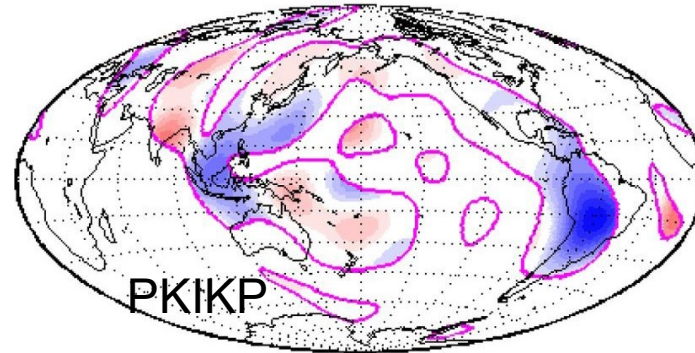
Seismic anomalies and resolution at the base of the mantle



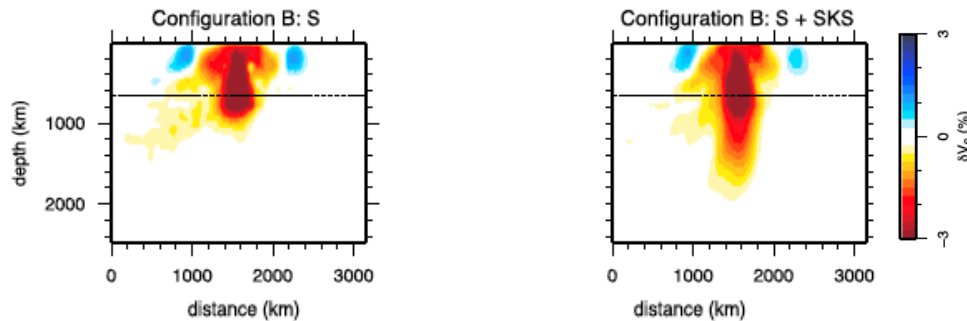
-Limited resolution: regularizations in tomographic inversions can generate blurred images and it is a challenge to robustly image structures smaller than ~ 1000 km scale



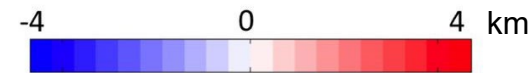
Davaille & Romanowicz, Tectonics (2020)



-Complexities with other existing structures: difficult to decouple mantle heterogeneity vs. CMB topography



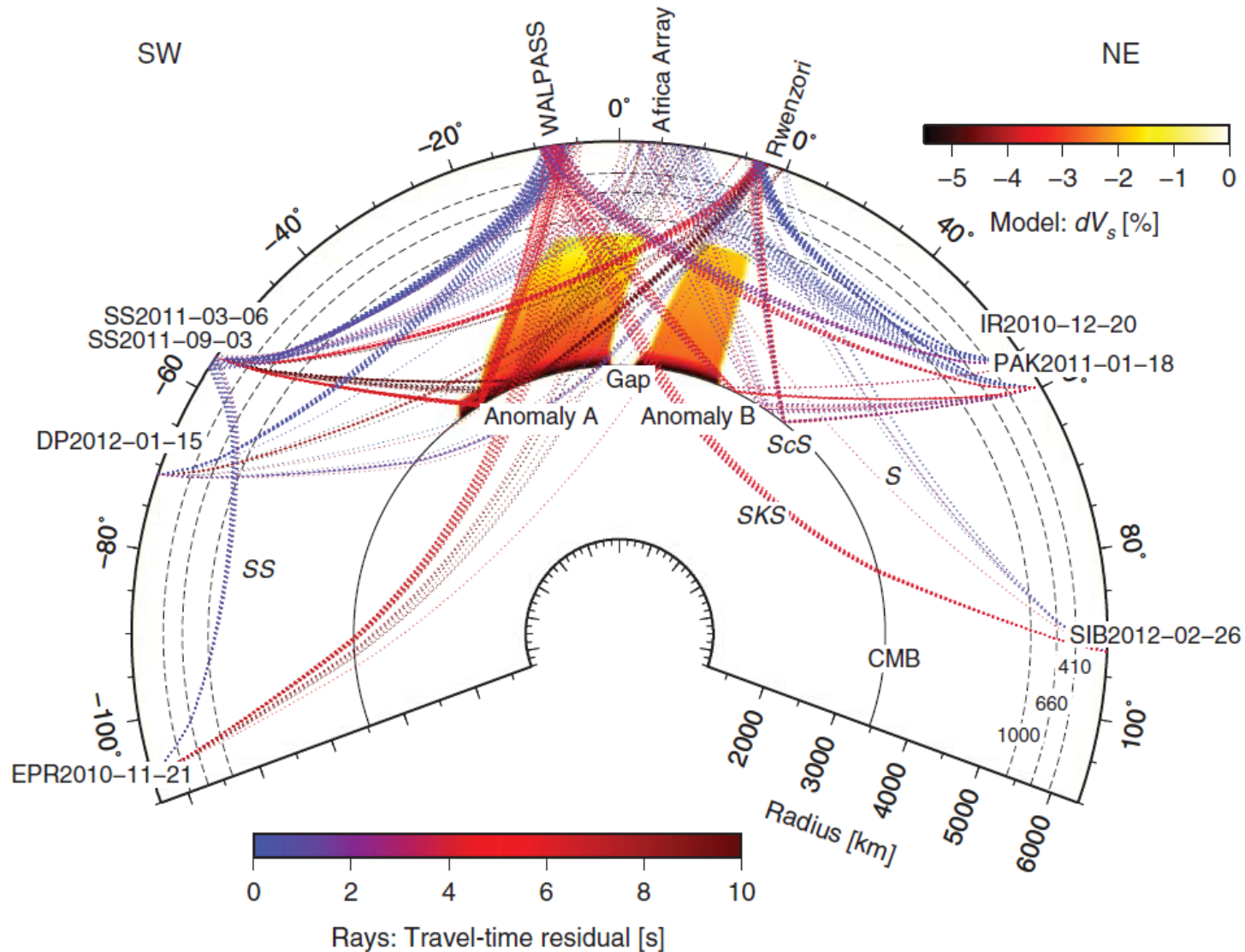
Maguire et al., GRL (2017)



CMB topographic patterns with varying lateral wavelengths

Colombi et al., GJI (2014)

Large-scale Structures [~ 5000 km]: complexities

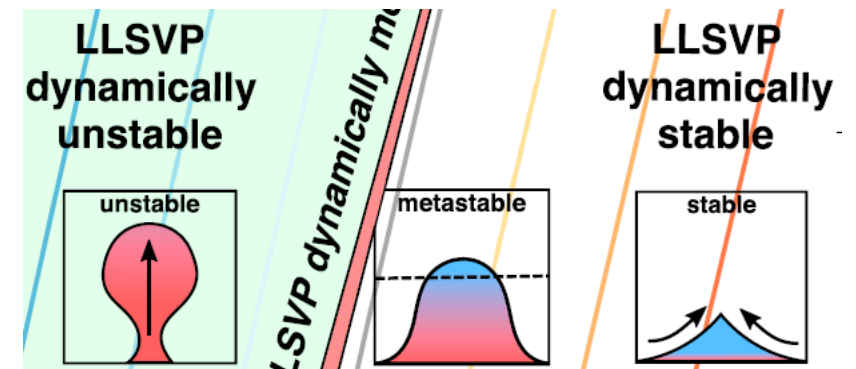


Kästle et al. (2017)

-Probing detailed morphology and small-scale intra-LLSVP structures are needed

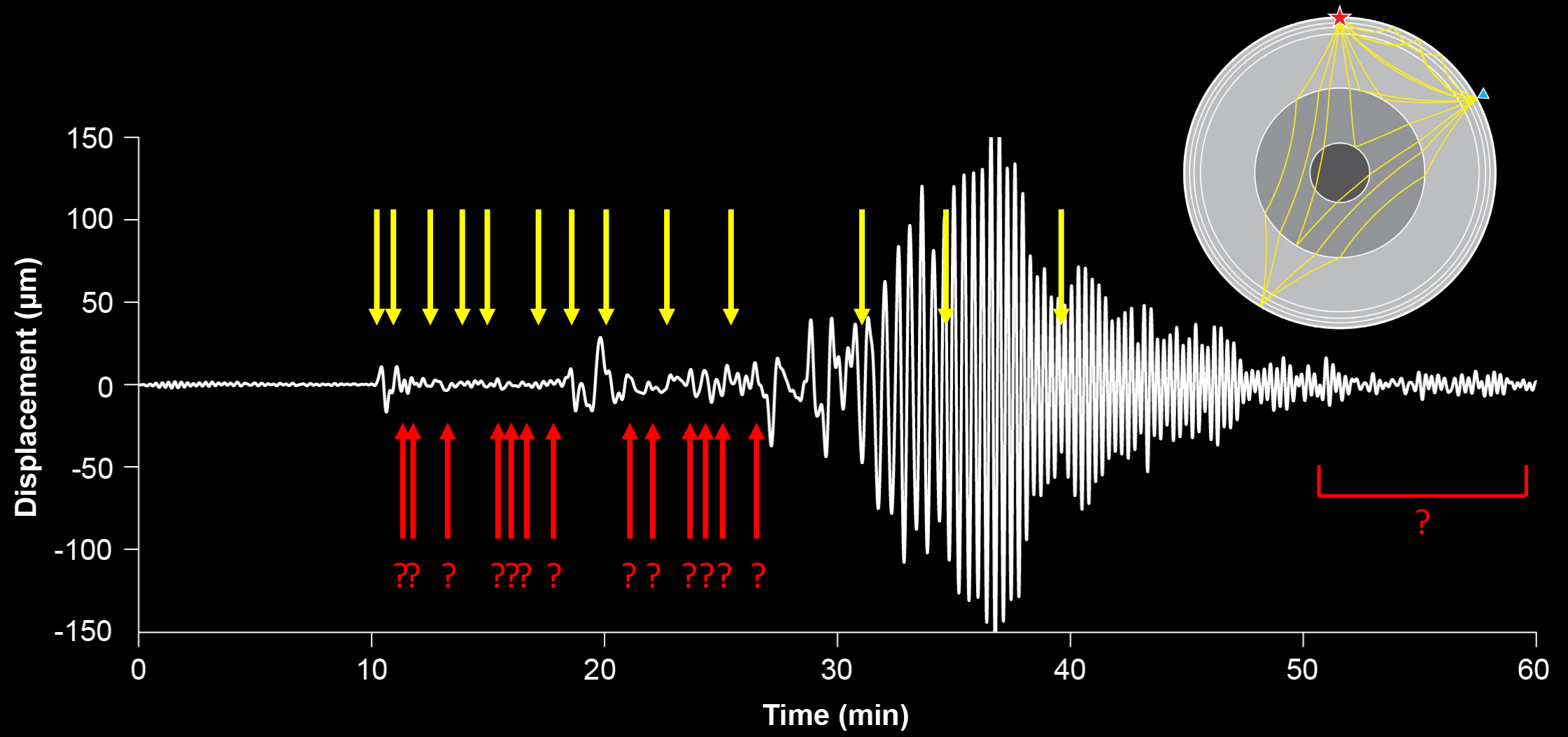
-Gap within African LLSVP: signature of individual plumes or piles boundary undulation?

-Often, such studies find stronger overall δV_s drops compared to global tomography

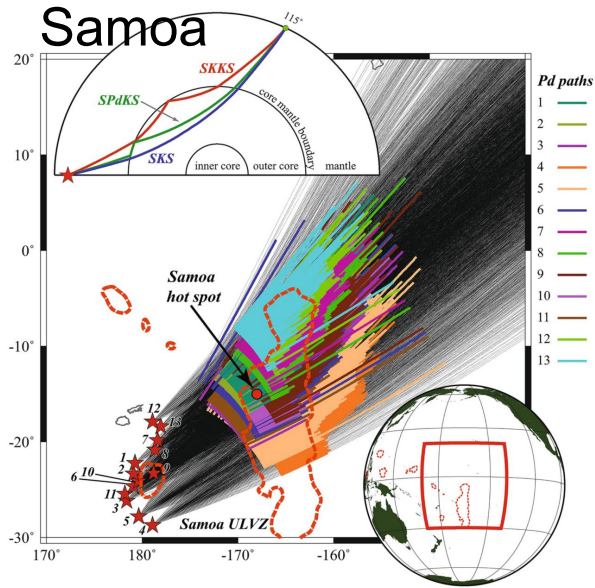


Wolf et al. (2015)

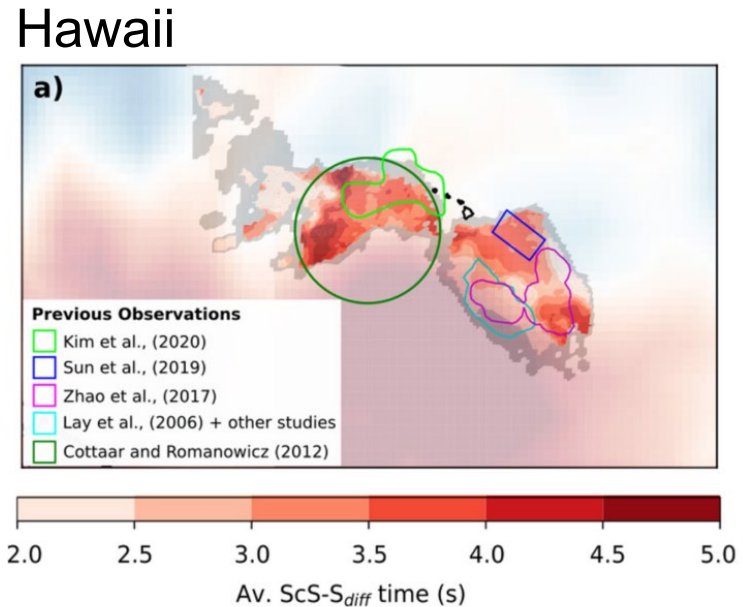
Holy Grail in Seismology: understading coda wiggles



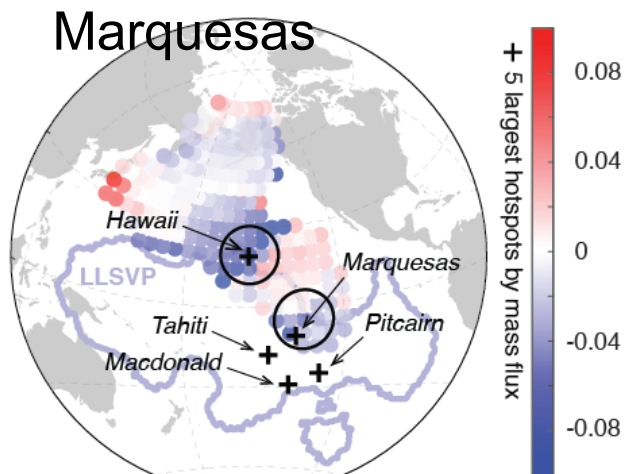
Meso-scale structures [~ 1000 km]: mega-ULVZs



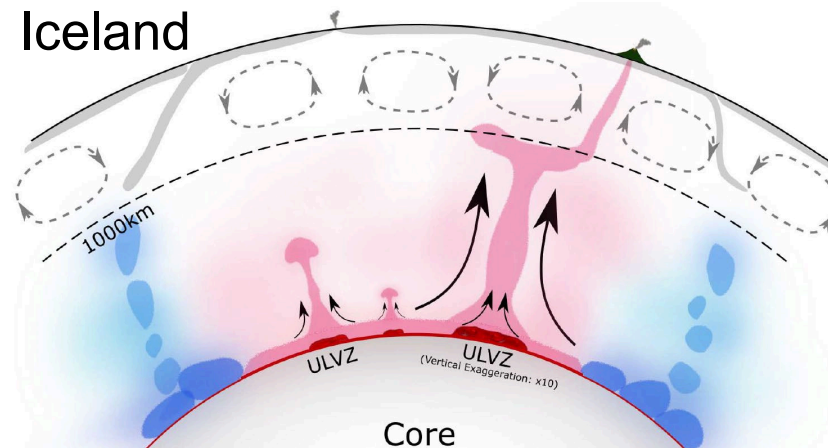
Krier, et al., JGR (2021)



Jenkins et al., EPSL (2021)



Kim et al., Science (2020)



Yuan & Romanowicz, Science (2017)

-Unresolved by tomographic inversions so largely inferred from distortions in the seismic wavefield (except for Perm / Kamchatka)

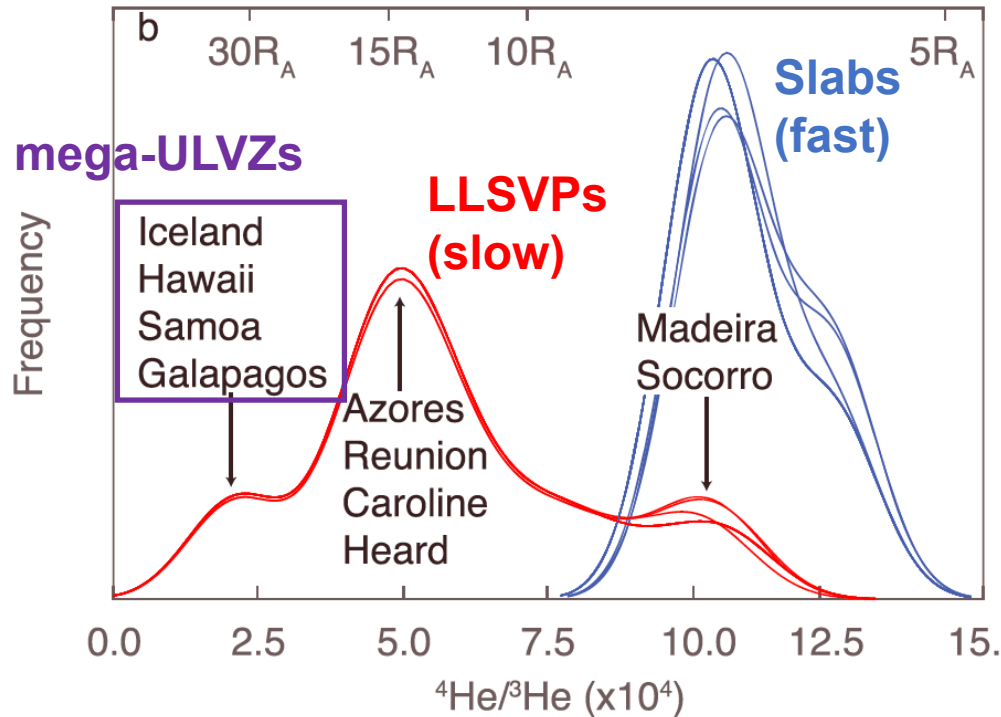
-Reduced S-wave speed up to 40%; P-wave constraints are relatively less explored

-Mega-ULVZs are only discovered below a few hotspot volcanoes (e.g., Cottarr & Romanowicz 2012; Thorne et al., 2013; Yuan & Romanowicz, 2017; Kim et al., 2020; Cottarr et al., 2022)

Hawaii/Samoa/Iceland/Marquesas/Galapagos

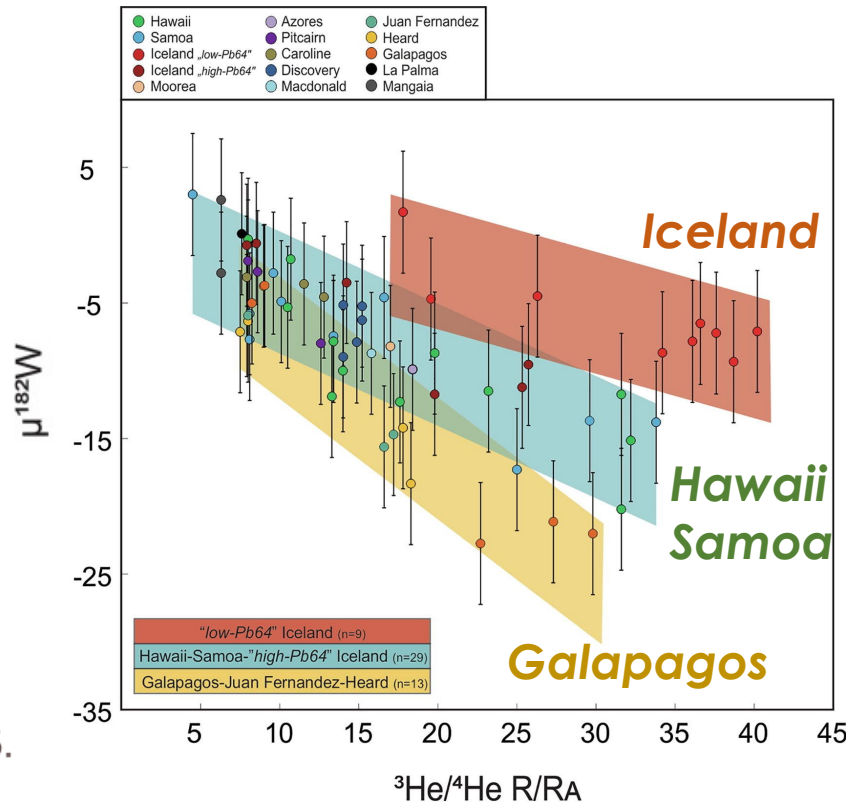
-Origin and composition of these structures are largely unknown

Geochemical Characteristics of Mega-ULVZs



Normal *Tungsten* & high $^3\text{He}/^4\text{He}$ suggests relatively un-degassed material that is associated with LLVPs

Williams et al., Science (2019)



Negative correlation of *Tungsten* vs. $^3\text{He}/^4\text{He}$ is seen at nearly all mega-ULVZs

Mundl-Petermeier et al., Geochimica et Cosmochimica Acta (2020)

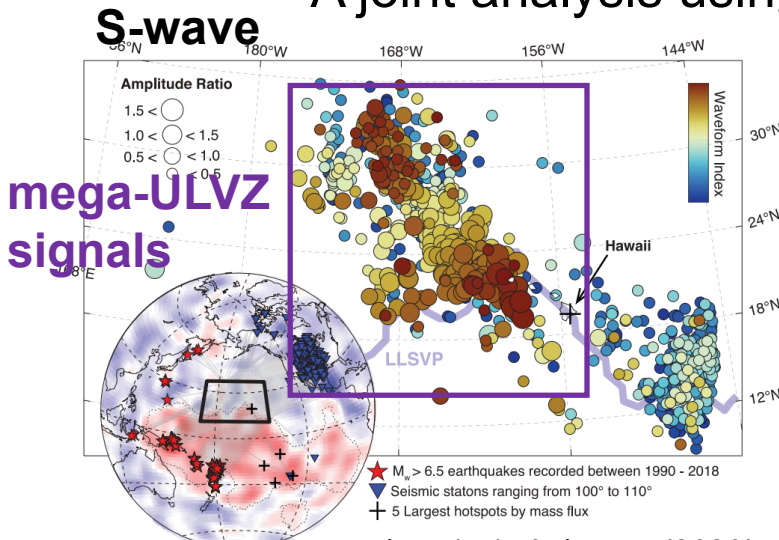
Some isotope system provides an important temporal constraint for primordial geochemical reservoirs

Mega-ULVZs may host primitive geochemical signatures?

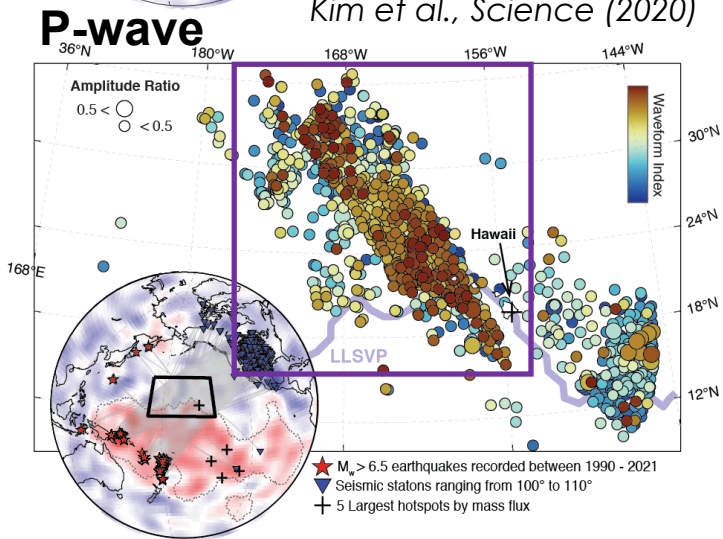
If mega-ULVZs are associated with partial melt, this will facilitate core-mantle isotope equilibration

One of a kind: Hawaiian mega-ULVZ

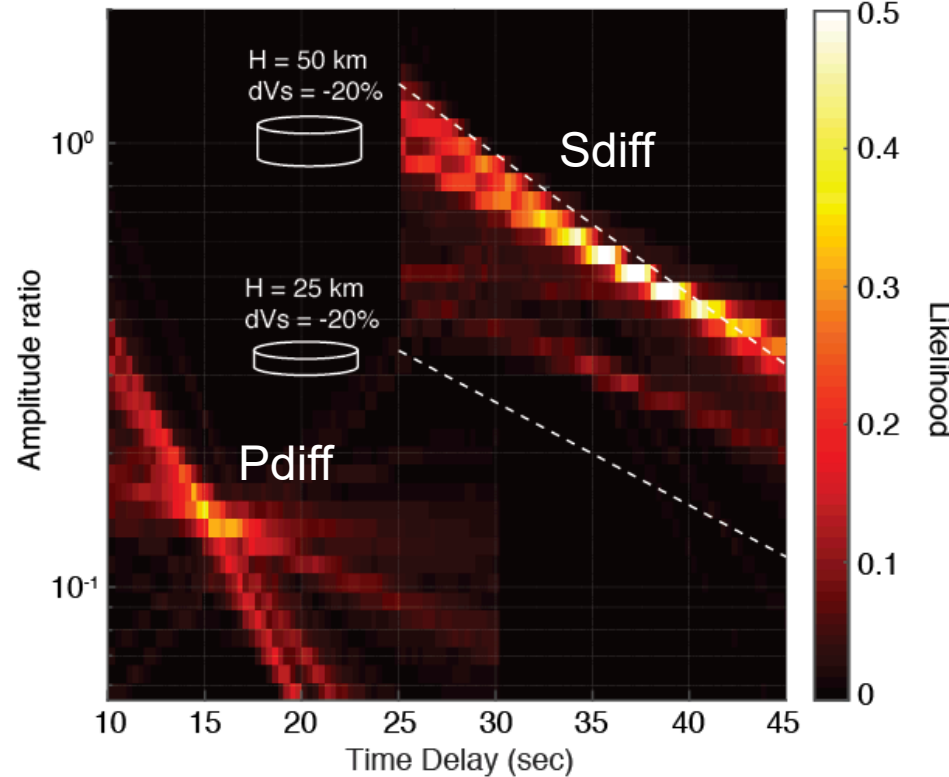
A joint analysis using Pdiff and Sdiff datasets



Kim et al., Science (2020)

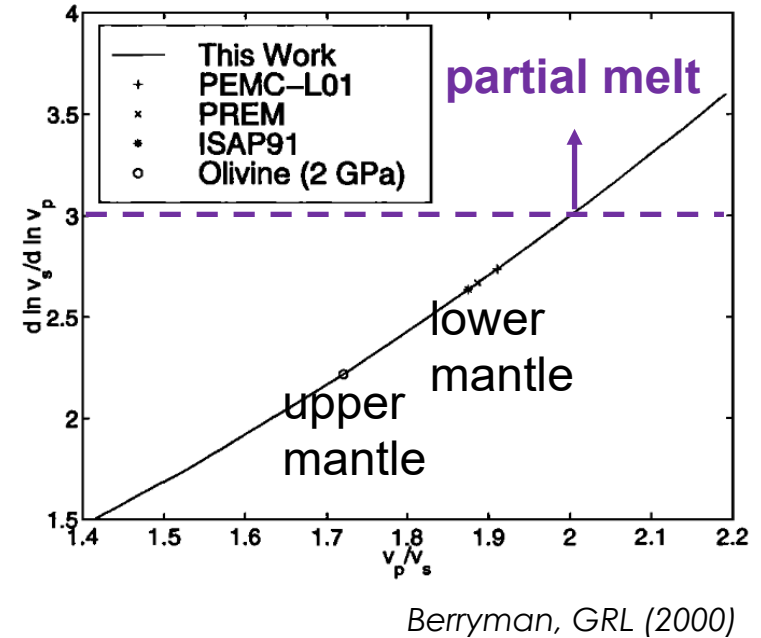


Kim et al., AGU meeting (2021)



P-wave amplitudes being much weaker ($d \ln V_s / d \ln V_p = 2-4$) but showing similar spatial pattern compared to the S-wave

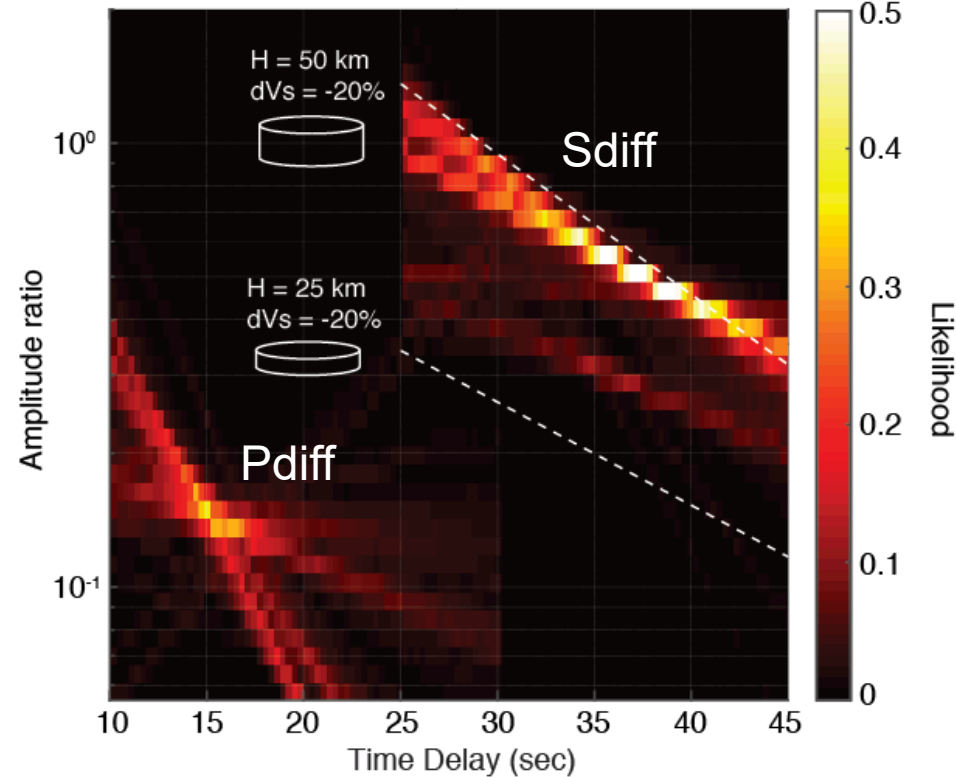
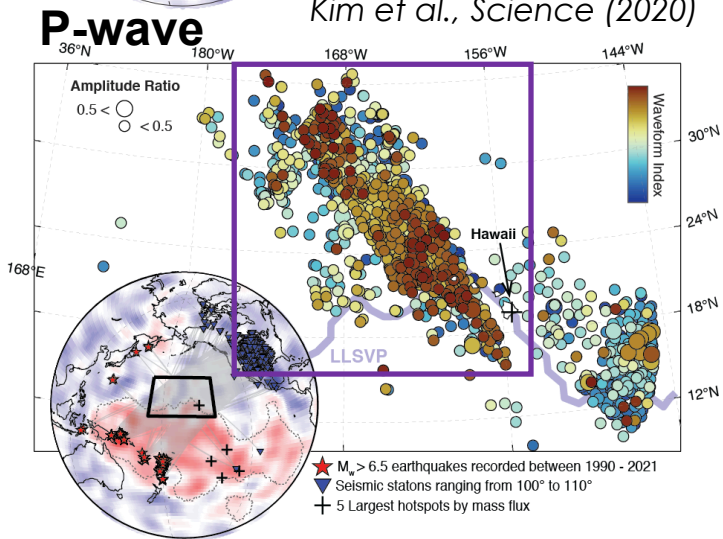
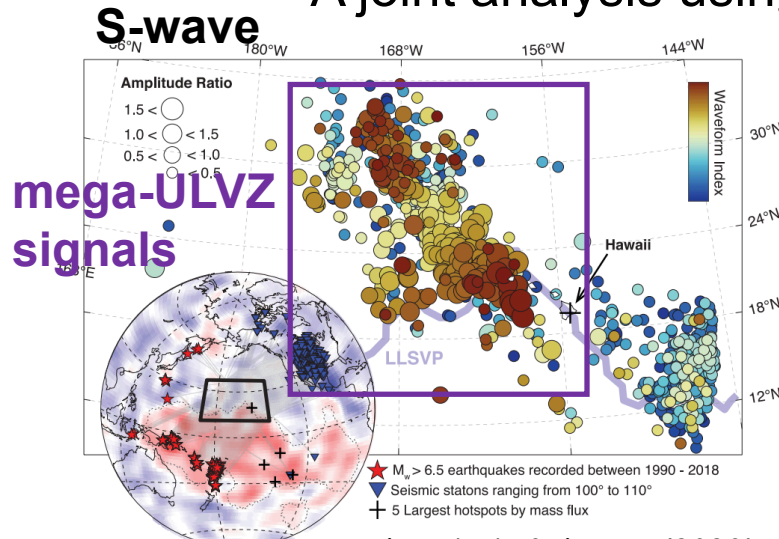
$d \ln V_s / d \ln V_p$ vs. V_p / V_s



Observed $d \ln V_s / d \ln V_p$ is close to what we would expect for partially molten materials at the CMB

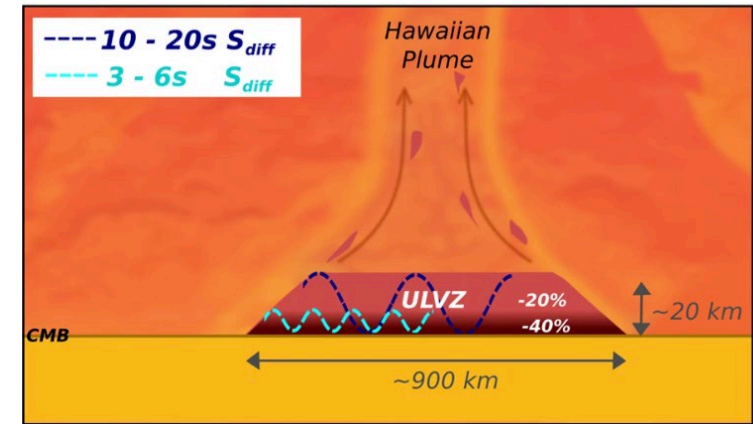
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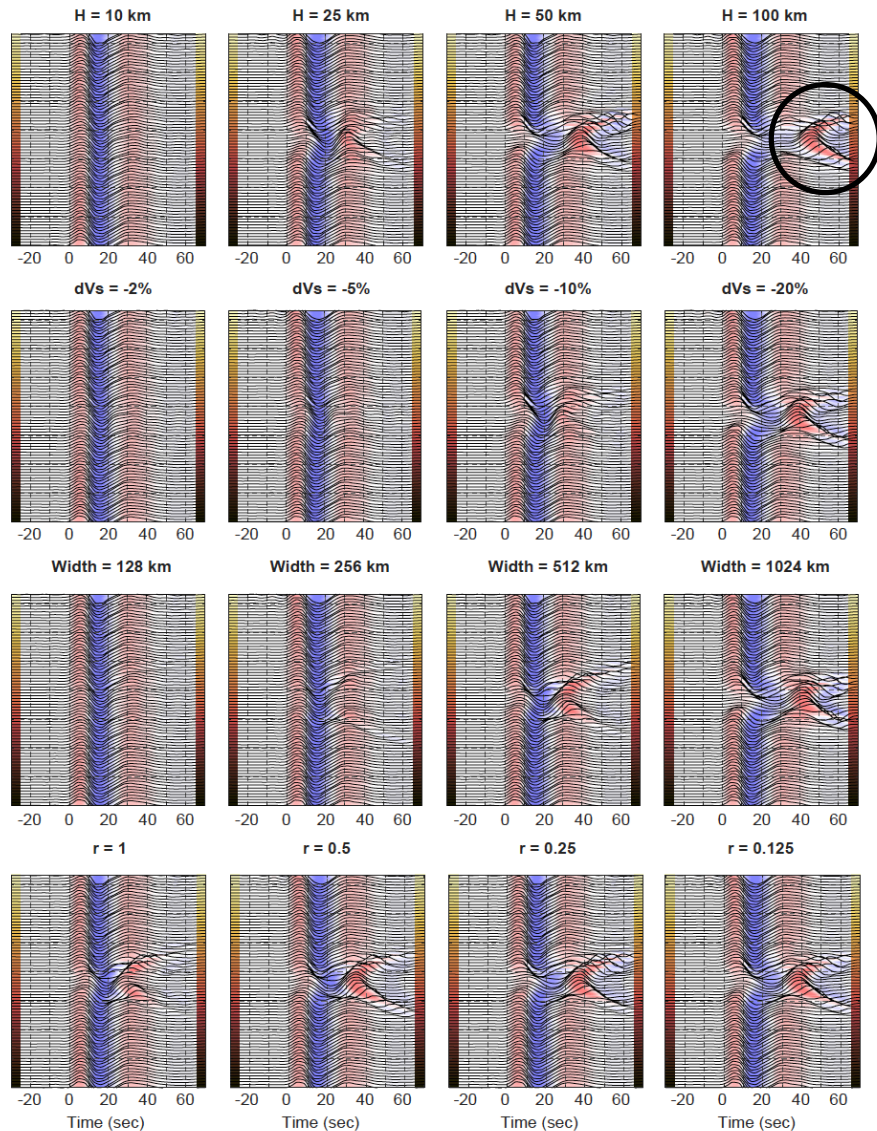
km-scale structures within the mega-ULVZ



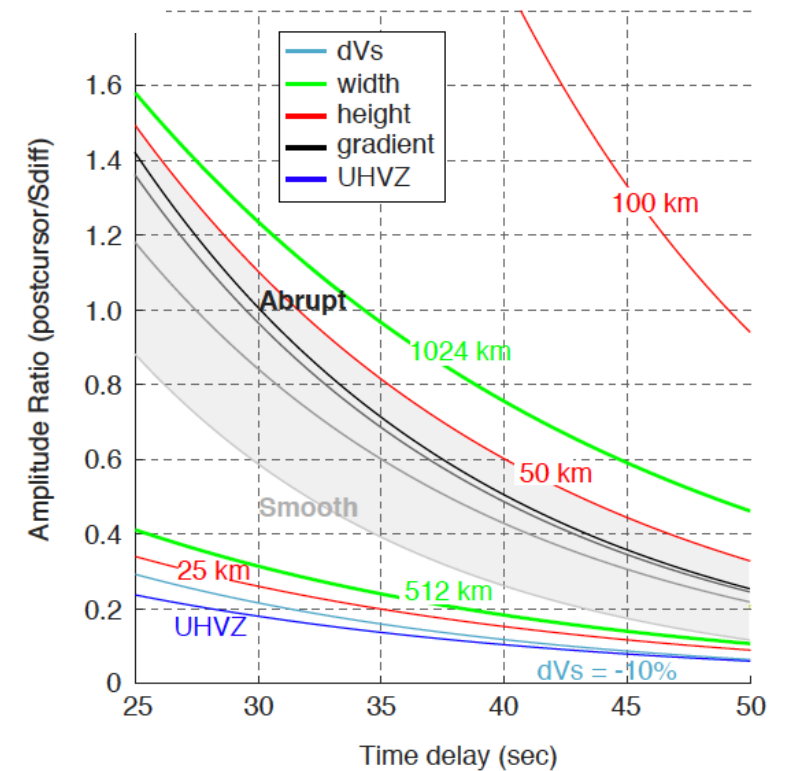
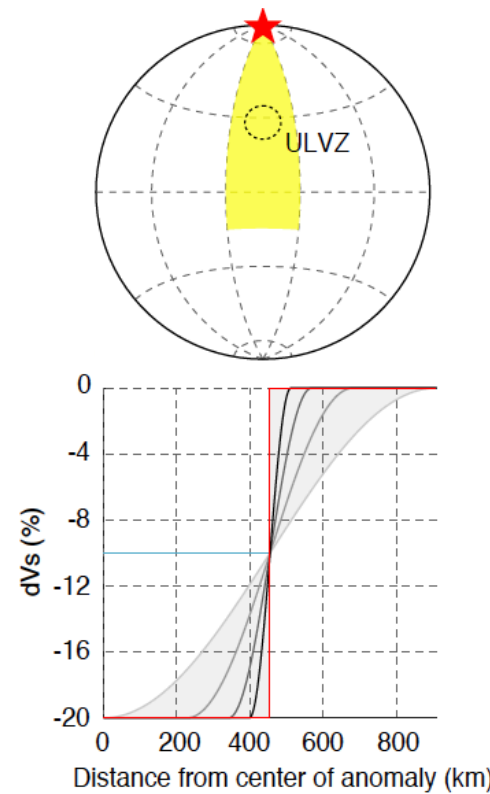
Li et al., Nat. Comm. (2022)

New seismic evidence supports a partially molten or chemical distinct material with increasing iron content towards the base of the CMB

Meso-scale structures [~ 1000 km]: challenges



Waveform simulations of S-wave diffracting the CMB

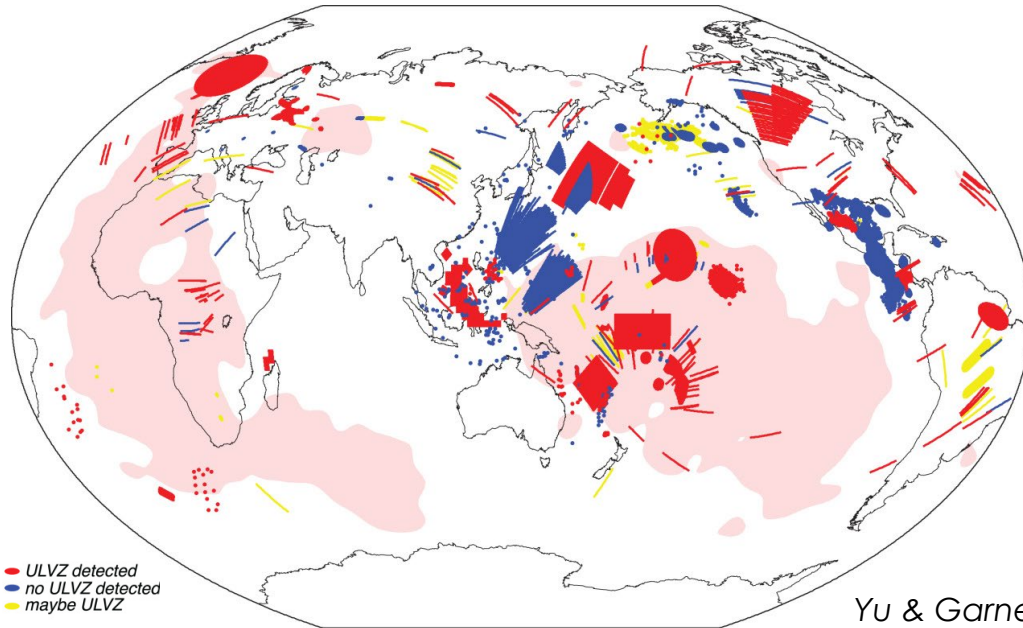


- Tradeoffs exist among physical parameters (velocity, density, size, shape, etc.) in waveform modeling
- Modeling with other seismic phases also suffer from non-uniqueness

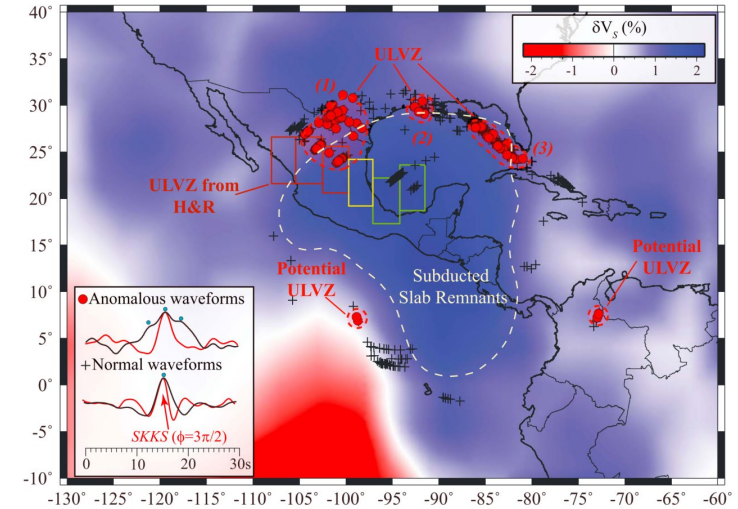
Small-scale Structures [$\sim 10\text{-}100\text{ km}$]: ULVZs

A global assessment of ULVZ studies until 2018

Some ULVZs may be associated with slabs



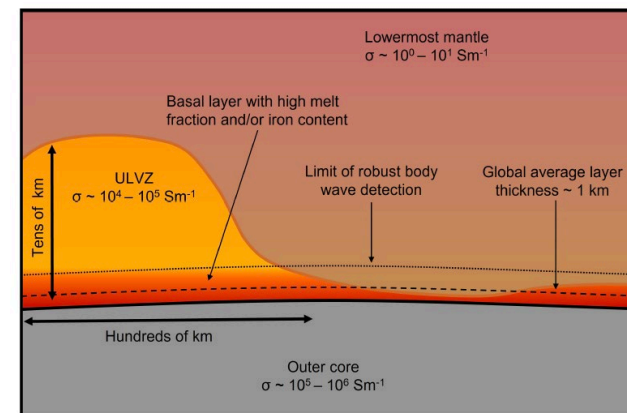
Yu & Garnero, G3 (2018)



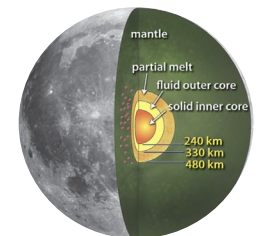
Thorne et al., GRL (2019)

- The total percentage of CMB area sampled by past ULVZ studies since 2018 is 17.1%
- Shows no spatial correlation between observed ULVZ locations and large-scale structures / hotspot volcanoes
- Studies using reflected / converted seismic data advocate a density increase

Globally molten layer at the CMB?

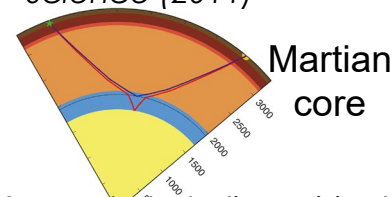


Russell et al., GRL (in revision)



Lunar core

Webber et al., Science (2011)

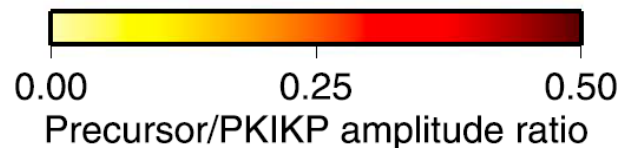
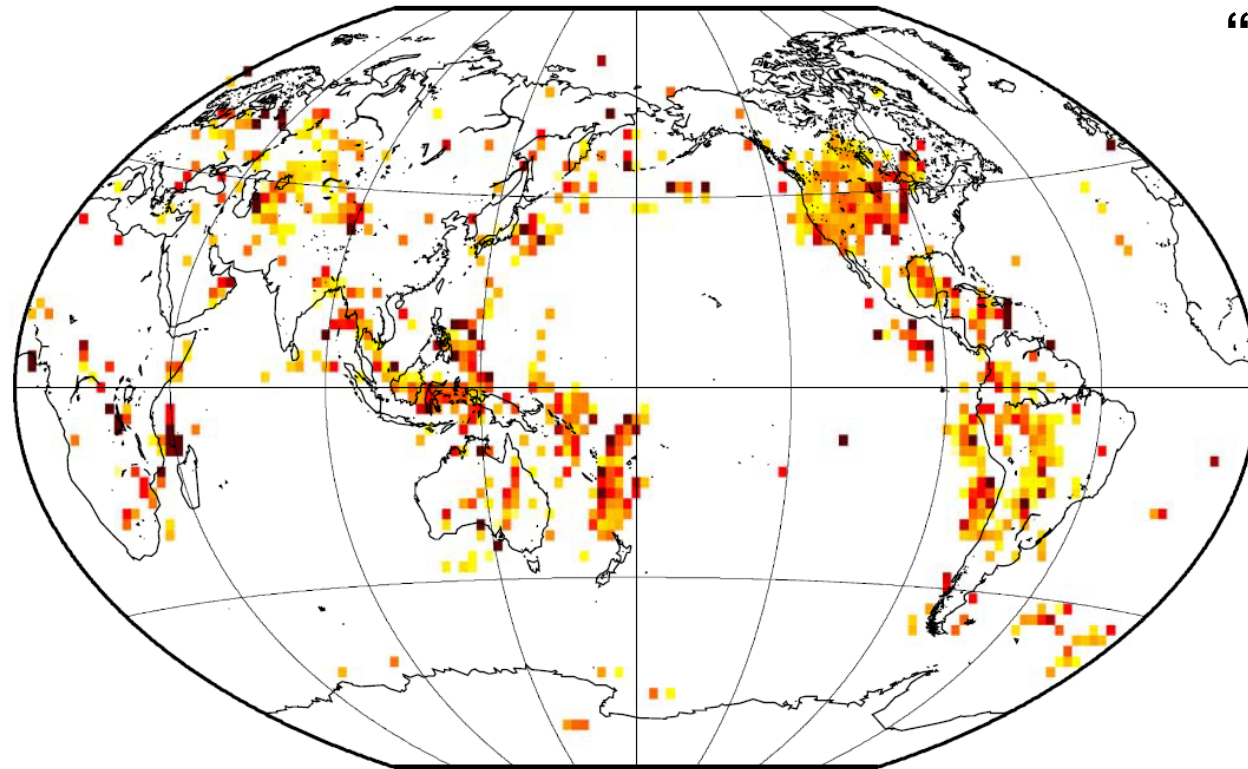


Martian core

Samuel et al., (in revision)

Small-scale Structures [~ 10 -100 km]: scatterers

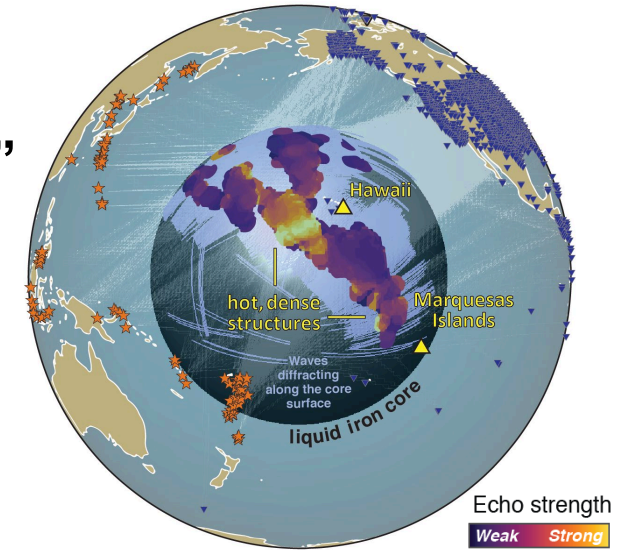
Global CMB scatterers from core-traversing phases



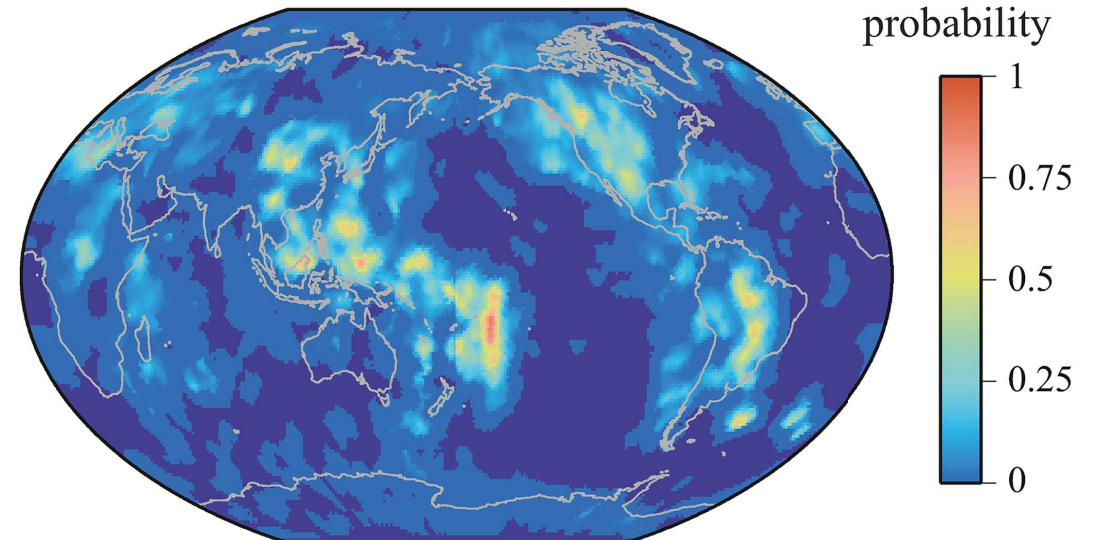
Waszek et al., GRL (2015)

“pervasive postcursors”

No strong correlation is found with known large lower mantle structures but may be related to widespread scattering ?



Kim et al., Science (2020)



Thorne et al., G3 (2021)

Key Takeaways

-Lower mantle hosts anomalous structures in various scales:

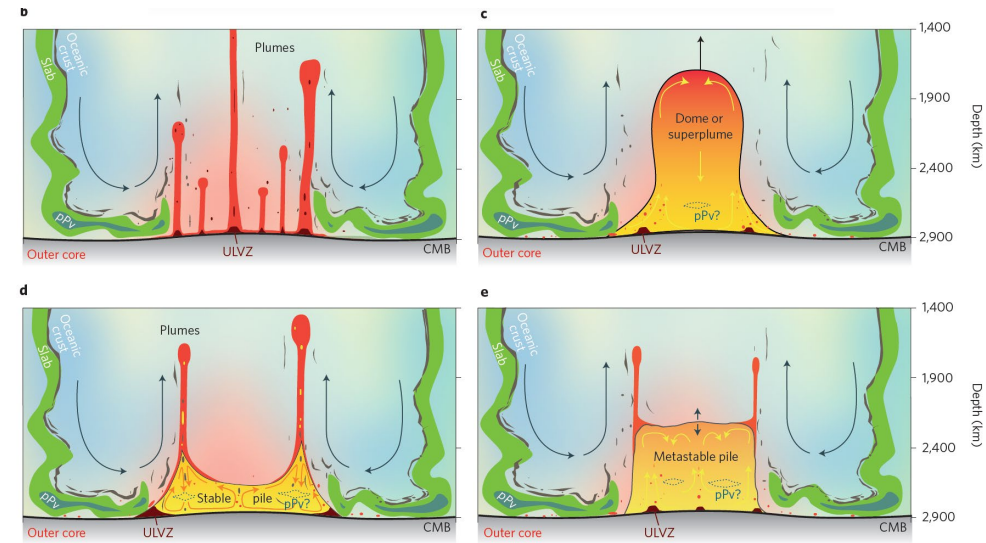
- LLSVPs [5000 km]
- Mega ULVZs [1000 km]
- ULVZs [100 km]
- CMB topography, unknown scatterers

-These structures are intimately tied to the fate of subducted slabs, origin of plumes, and nature of primitive geochemical reservoirs with many open questions to be answered

-Robustly imaging the lower mantle structure requires different types of seismic waves and multiple approaches

-Much of the lowermost mantle / CMB region remains unexplored (e.g., physical coverage illuminated by seismic waves, anisotropy, attenuation, etc.)

-Seismologists continue to work together with the larger terrestrial / planetary science community to uncover various mysteries of the deep planetary interior structures and dynamics



Garnero et al., Nat. Geo. (2017)

With many future planetary missions, we will learn so much more !

Apollo (1961-1972)

Farside Seismic Suite (2025)

Dragonfly (2026)

VERITAS (2028?)

EnVision (2028-2031?)

Phantom (TBD)

InSight (2018-2023)

