

# Probing the properties of Core-Mantle Boundary using Atmospheric Neutrino Oscillation

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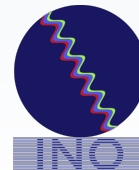
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**International Workshop on Multi-messenger Tomography of the Earth**

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**Collaborators:** Anil Kumar, Sanjib Kumar Agarwalla, Amol Dighe

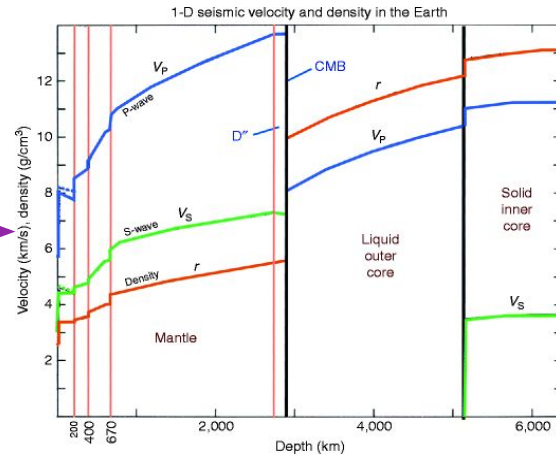
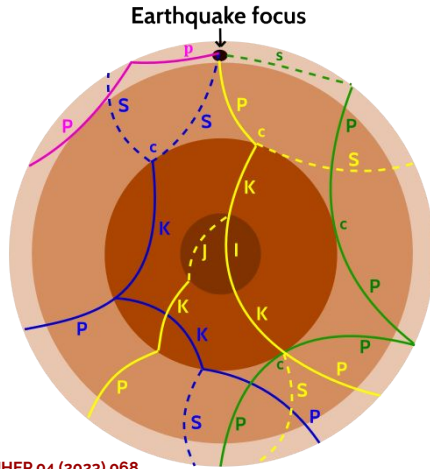


# Outline

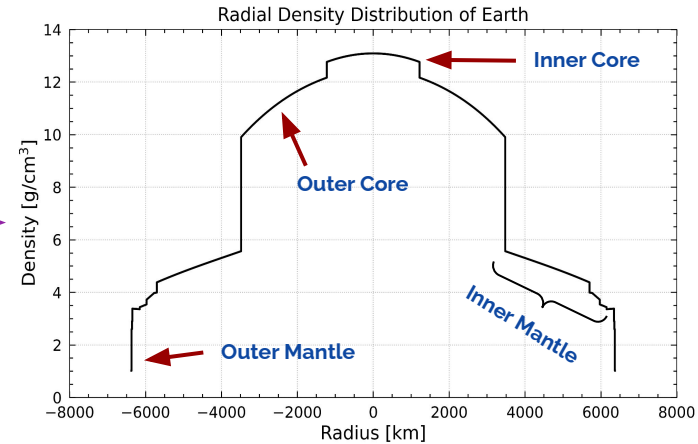
- Information About Earth's interior
- Atmospheric Neutrinos
- Earth's Matter Effects: Key to Probe Internal Structure of Earth
- ICAL at INO
- Neutrinos for Probing the Properties of Core-Mantle Boundary (CMB)
  - Locating CMB
  - Measuring Density Jump at CMB

# The Interior of Earth

- Information about the interior of Earth is obtained from indirect probes used in traditional **seismic** and **gravitational** studies → **Preliminary Reference Earth Model (PREM)**



"PREM" Dziewonski and Anderson (1981)



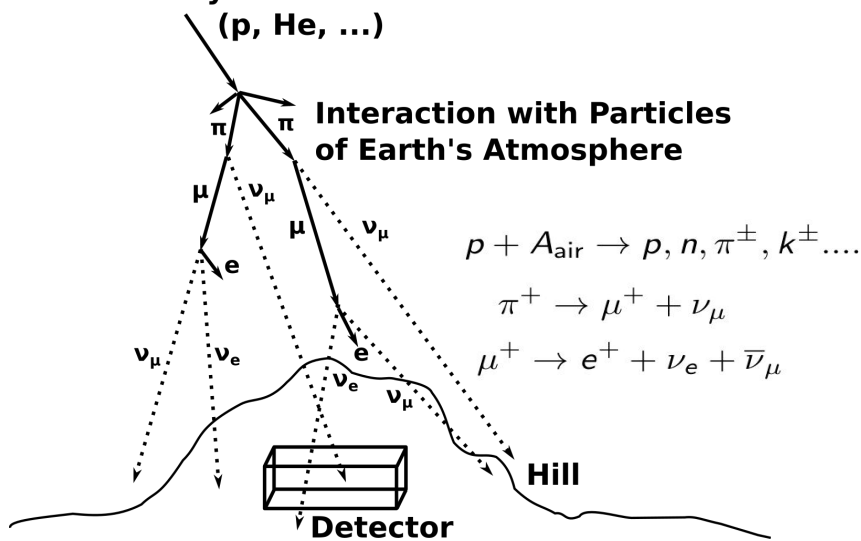
JHEP 04 (2023) 068

- Broadly classified: two concentric shell - the outer one is mantle, and the inner one with a much higher density is core
- Mantle consists of hot rocks of silicate and core is composed of metals like iron and nickel
- Outer core is expected to be liquid (absence of S-waves and decrease in the velocity of P-waves)
- Core-Mantle Boundary (CMB): the largest compositional discontinuity within the Earth at a depth of 2891 km
- The large density contrast across the CMB

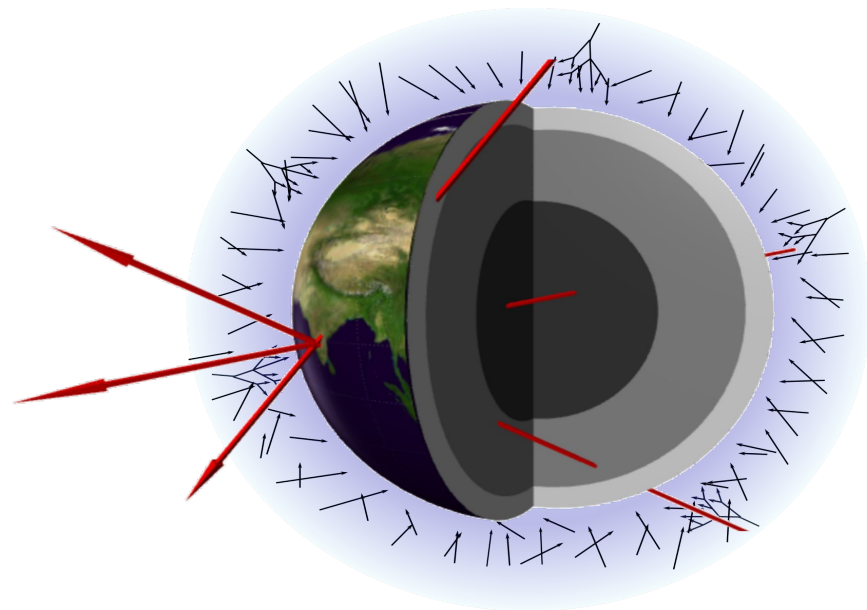
# Atmospheric Neutrinos

- Complementary information may be obtained using the observations of **neutrinos that pass through Earth**
- Neutrino exploit the **weak interactions** to probe the internal structure of Earth

## Primary Cosmic Particles



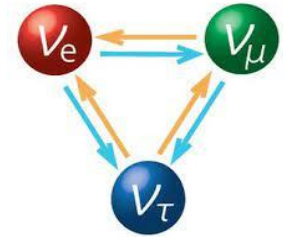
- Flux should be up-down symmetric
- Almost isotropic flux



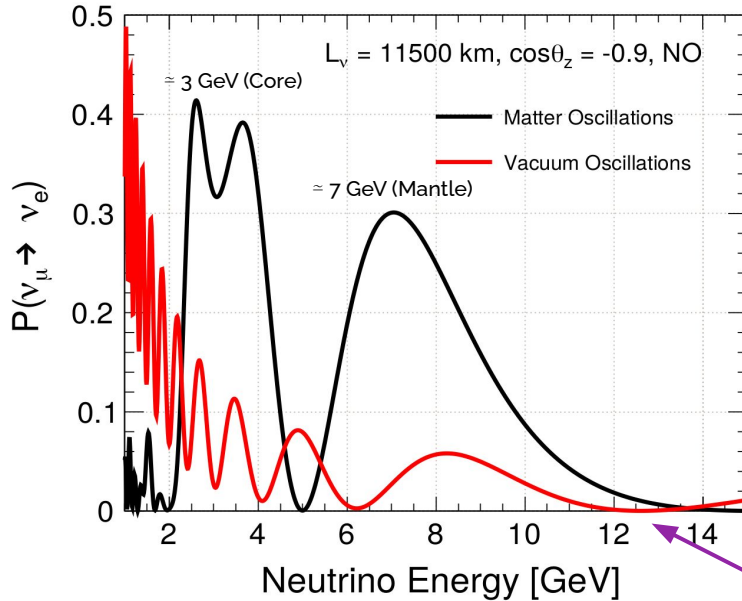
- Baseline: **15 km to 12757 km**
- Energy range: **0.1 GeV to ~10 TeV**

# Neutrino Oscillations

- Neutrino changes its flavor while propagating
- Neutrino oscillation probability get modified significantly in matter



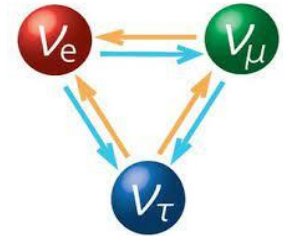
Earth's matter vs. Vacuum



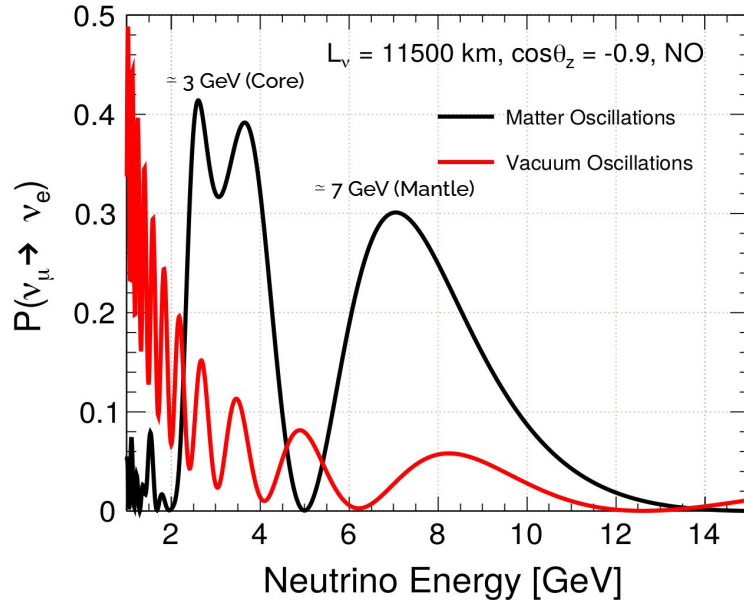
$$E_{\text{res}} = \frac{\Delta m_{31}^2 \cos 2\theta_{13}}{2\sqrt{2}G_F N_e} \simeq 7 \text{ GeV} \left( \frac{4.5 \text{ g/cm}^3}{\rho} \right) \left( \frac{\Delta m_{31}^2}{2.4 \times 10^{-3} \text{ eV}^2} \right) \cos 2\theta_{13}$$

# Neutrino Oscillations

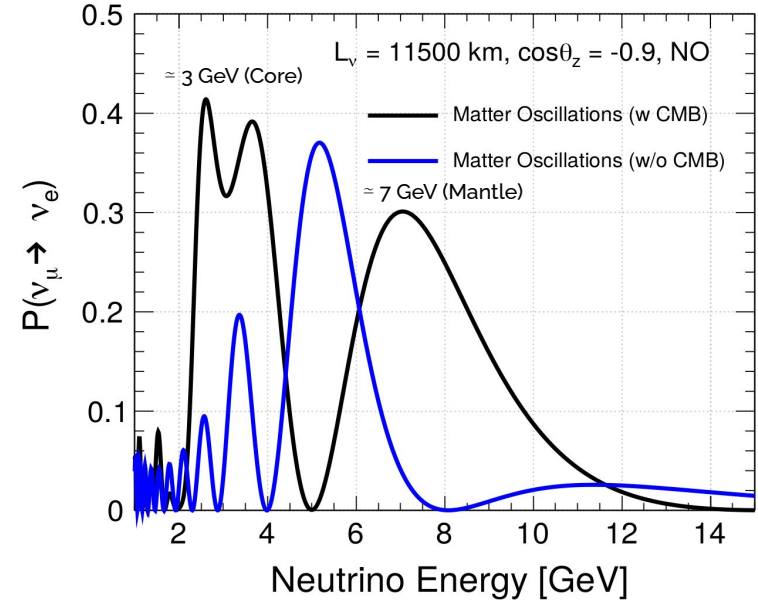
- Neutrino changes its flavor while propagating
- Neutrino oscillation probability get modified significantly in matter
- In absence of high density jump at CMB, the NOLR/PR resonance does not observed



Earth's matter vs. Vacuum



CMB vs. w/o CMB



# Matter Effects: Key to Probe Internal Structure of Earth

- The atmospheric neutrinos undergo coherent forward scattering with ambient electrons inside Earth which leads to the modification of neutrino oscillation probabilities

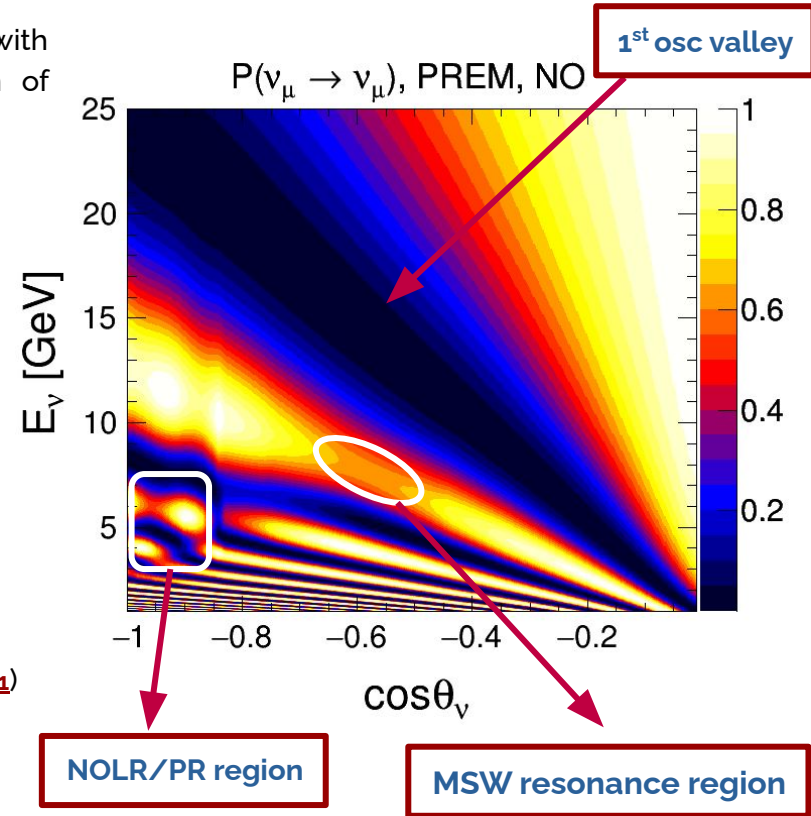
Electron number density

$$V_{CC} = \pm\sqrt{2}G_F N_e \approx \pm 7.6 \times Y_e \times 10^{-14} \left[ \frac{\rho}{g/cm^3} \right] \text{eV}$$

$$Y_e = N_e / (N_p + N_n)$$

$\rho$  denotes the matter density  
+1 (-1) for neutrino (antineutrino)

- MSW resonance ([L. Wolfenstein, PRD 17 \(1978\) 2369](#))
- Neutrino Oscillation Length Resonance (NOLR) ([Petcov, PLB 434 \(1998\) 321](#))  
or Parametric Resonance (PR) ([Akhmedov, NPB 538 \(1999\) 25](#))

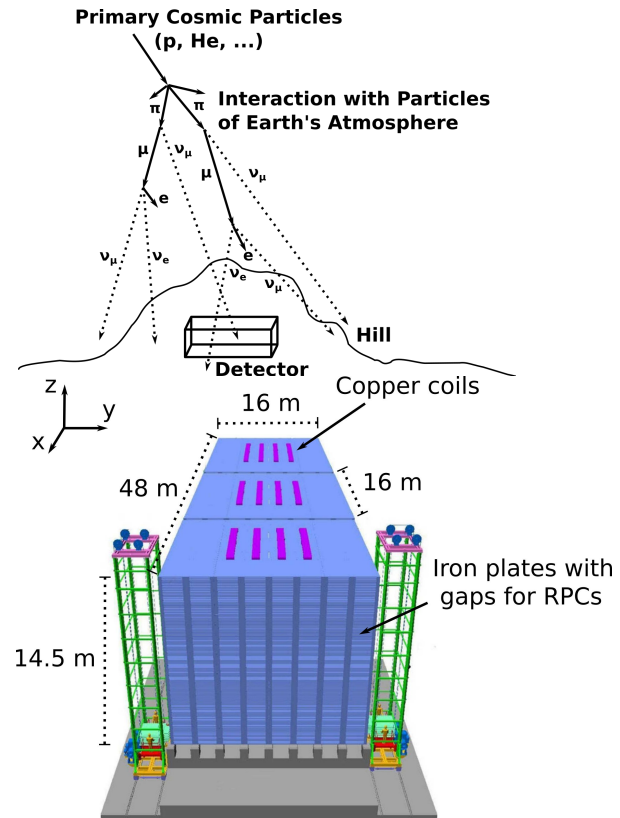


# Iron Calorimeter Detector (ICAL) at INO

- **ICAL@INO:** 50 kton magnetized iron calorimeter detector at the proposed India-based Neutrino Observatory (INO)
- **Location:** Bodi West Hills, Theni District, Tamil Nadu, India
- **Aim:** To determine neutrino mass ordering and precision measurement of atmospheric neutrino oscillation parameters
- **Source:** Atmospheric neutrinos and antineutrinos in the multi-GeV range of energies over a wide range of baselines
- **Uniqueness:** Charge identification capability helps to distinguish  $\mu^-$  and  $\mu^+$  and hence,  $\nu_\mu$  and  $\bar{\nu}_\mu$
- **Muon energy range:** 1 – 25 GeV
- **Muon energy resolution:** ~ 10%
- **Baselines:** 15 – 12000 km
- **Muon zenith angle resolution:** ~ 1°

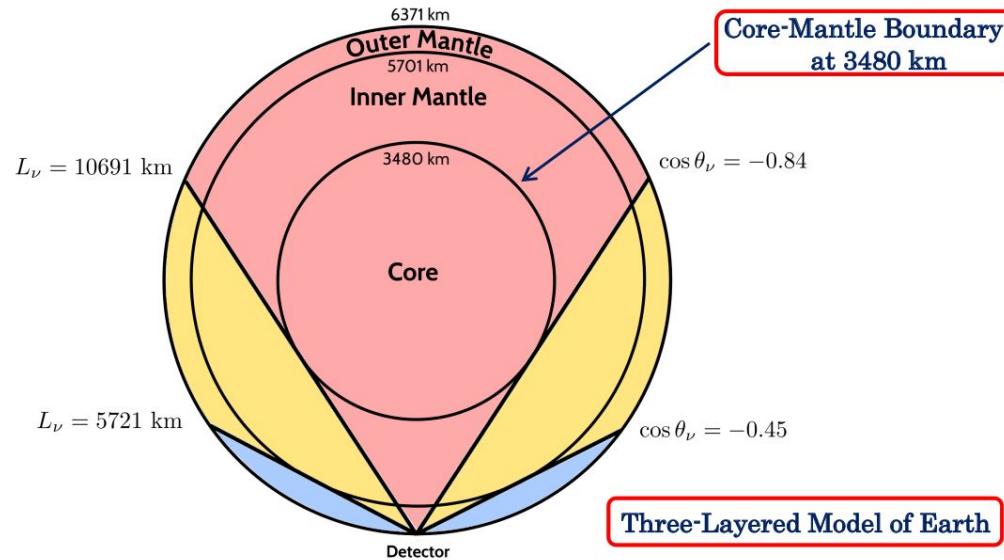
Next talk by  
Dr. Anil Kumar

[Pramana 88 \(2017\) 5, 79, arXiv:1505.07380](#)



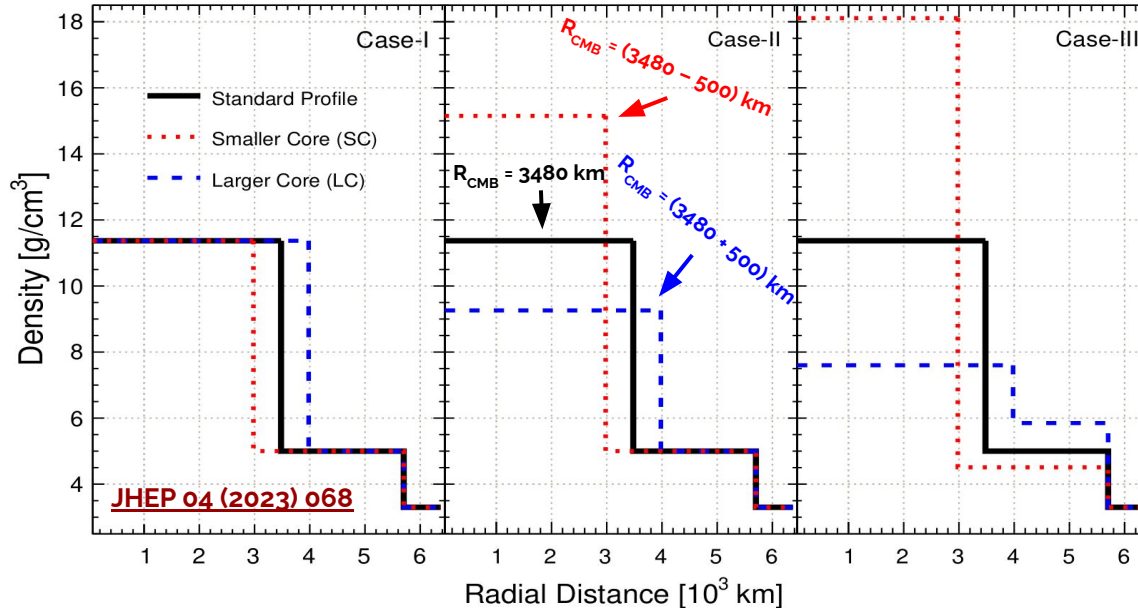


# Locating Core-Mantle Boundary (CMB) Using Atmospheric Neutrinos



Based on: A. K. Upadhyay, A. Kumar, S. K. Agarwalla, and A. Dighe [JHEP 04 \(2023\) 068](#)

# A Few Toy Models of Earth with Varying CMB



**Case-I:** Densities of all layers fixed and  $M_{\oplus}$  is not invariant

**Case-II:** Densities of inner & outer mantle fixed. Core density varies to keep  $M_{\oplus}$  invariant

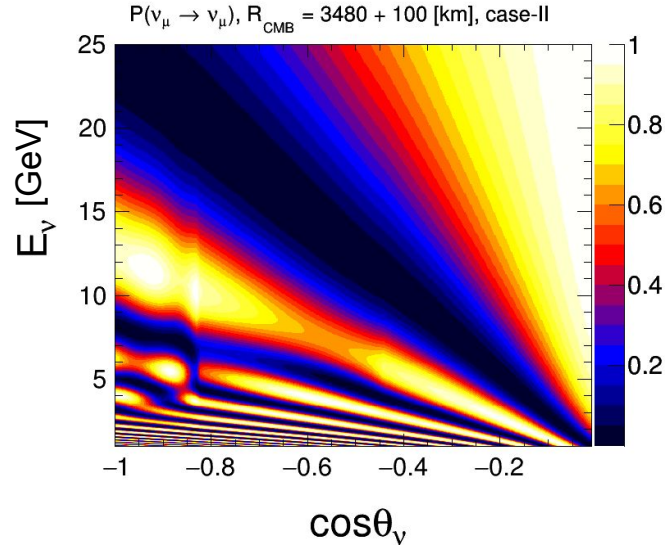
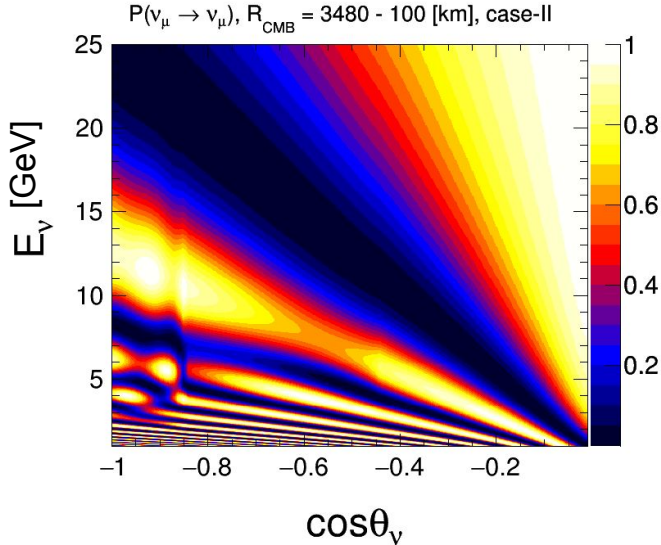
**Case-III:** Core & inner mantle densities vary to keep their masses fixed. Outer mantle density fixed &  $M_{\oplus}$  invariant

# Effect of CMB Variation on Oscillograms

Case-II

Smaller Core (SC)

Larger Core (LC)



- For smaller core, the NOLR/PR shifts to left and patterns shrink
- For larger core, the NOLR/PR shifts to right and patterns broaden
- No modification in MSW region

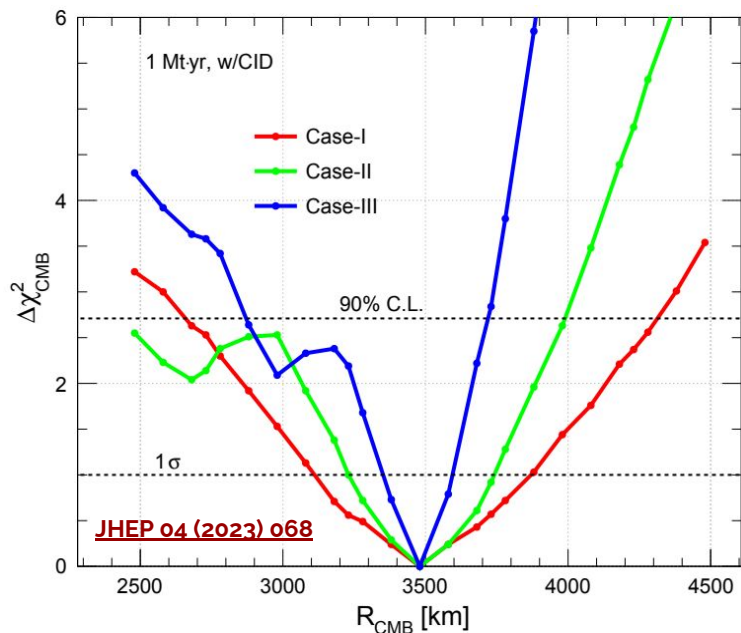
[JHEP 04 \(2023\) 068](#)

Opposite modification in NOLR/PR region for SC and LC

# Sensitivity for Locating Core-Mantle Boundary

$$\Delta\chi_{\text{CMB}}^2 = \chi^2(\text{modified } R_{\text{CMB}}) - \chi^2(\text{standard } R_{\text{CMB}})$$

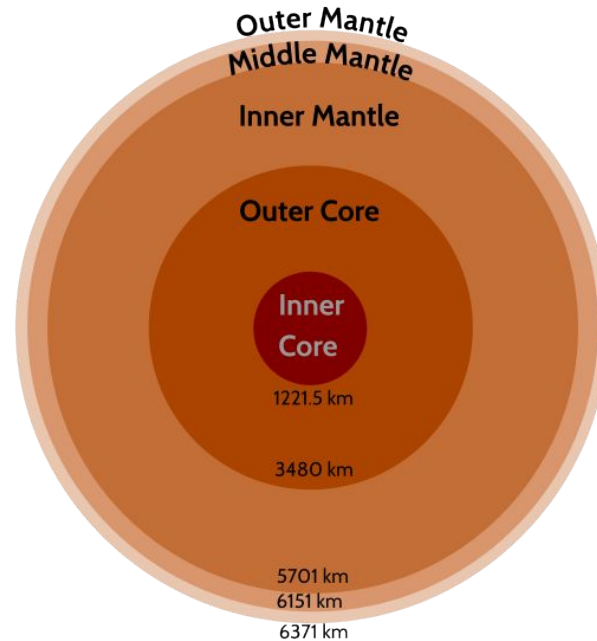
- 1 Mt·yr exposure at ICAL
- Marginalization over: (1) systematic uncertainties, (2) Oscillation parameter: (2a)  $\sin^2\theta_{23}$ : (0.36, 0.66), (2b)  $\Delta m_{\text{eff}}^2$ : (2.1, 2.6)  $\times 10^{-3} \text{ eV}^2$  (2c), mass ordering: (NO, IO)



		$\Delta\chi_{\text{CMB}}^2$	
		w CID	w/o CID
Case-I	$R_{\text{CMB}} = 2980 \text{ km}$	1.53	1.01
	$R_{\text{CMB}} = 3980 \text{ km}$	1.44	0.95
Case-II	$R_{\text{CMB}} = 2980 \text{ km}$	2.53	1.66
	$R_{\text{CMB}} = 3980 \text{ km}$	2.63	1.67
Case-III	$R_{\text{CMB}} = 2980 \text{ km}$	2.09	1.33
	$R_{\text{CMB}} = 3980 \text{ km}$	8.07	5.23

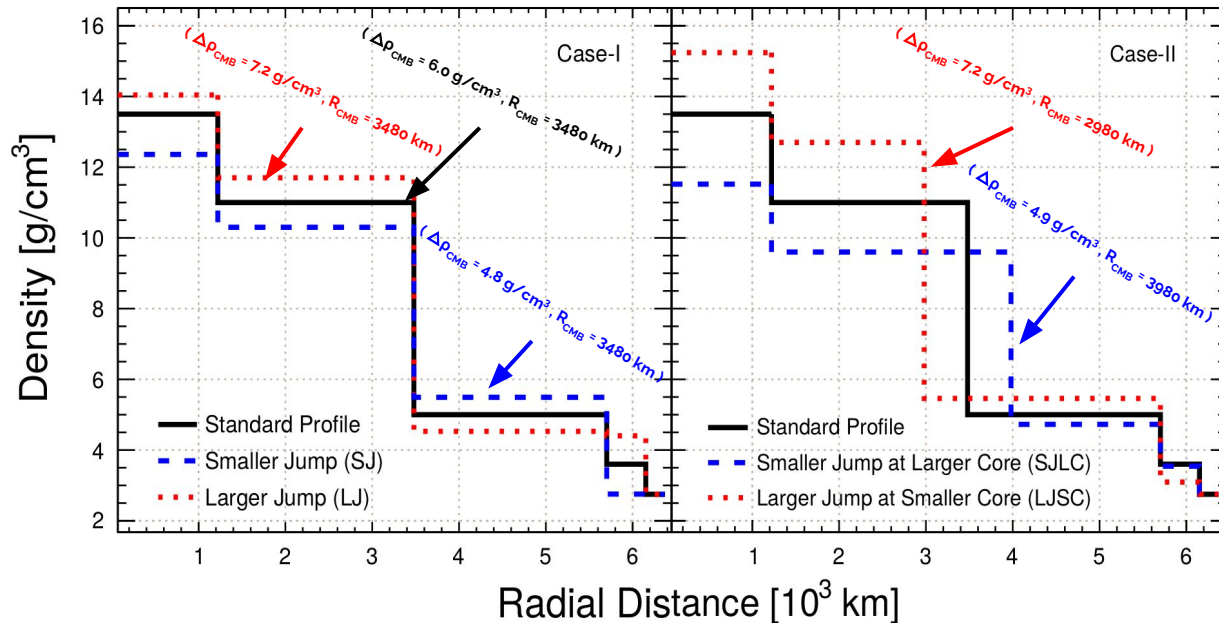
- $1\sigma$  precision of about  $\pm 380 \text{ km}$ ,  $\pm 250 \text{ km}$ , and  $\pm 120 \text{ km}$  for Case-I, Case-II, and Case-III, respectively.

# Measuring Correlation Between Density Jump at CMB and Location of CMB Using Atmospheric Neutrinos



- **Five layers:** Inner Core (IC), Outer Core (OC), Inner Mantle (IM), Middle Mantle (MM), and Outer Mantle (OM)
- **Density Jumps:** IC-OC, CMB, IM-MM, and MM-OM boundary;  $\Delta\rho_{\text{IC-OC}}$ ,  $\Delta\rho_{\text{CMB}}$ ,  $\Delta\rho_{\text{IM-MM}}$ , and  $\Delta\rho_{\text{MM-OM}}$

# Toy Models of Earth with Modifying $\Delta\rho_{\text{CMB}}$ & $R_{\text{CMB}}$



## Constraints:

- Mass of Earth is fixed
- Moment of inertia of Earth is fixed
- Hydrodynamic condition  
( $\rho_{\text{inner layer}} > \rho_{\text{outer layer}}$ ) is satisfied
- Density of outer mantle is fixed
- Electron number density ratio:  
 $Y_e = N_e / (N_p + N_n)$  or  $Z/A$  :  
 $Y_e$  (Inner Core) = 0.466  
 $Y_e$  (Outer Core) = 0.466  
 $Y_e$  (Mantle) = 0.494

**Case-I:** Density jump ( $\Delta\rho_{\text{CMB}}$ ) modify at the standard  $R_{\text{CMB}} = 3480$  km

**Case-II:** Density jump ( $\Delta\rho_{\text{CMB}}$ ) and location of CMB ( $R_{\text{CMB}}$ ) both modify simultaneously

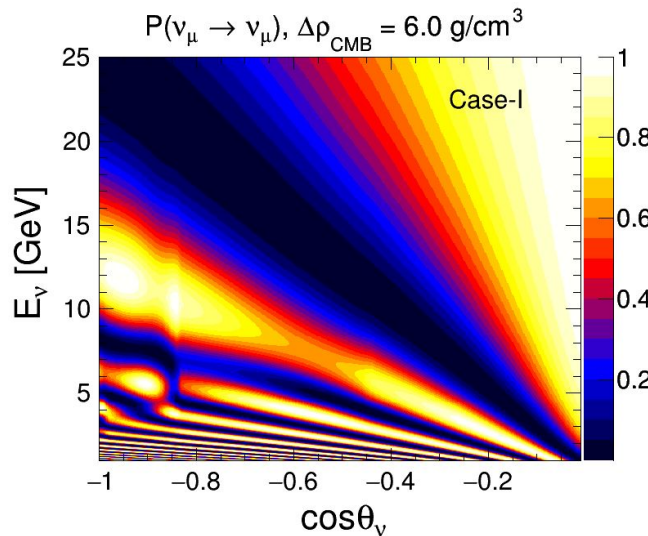
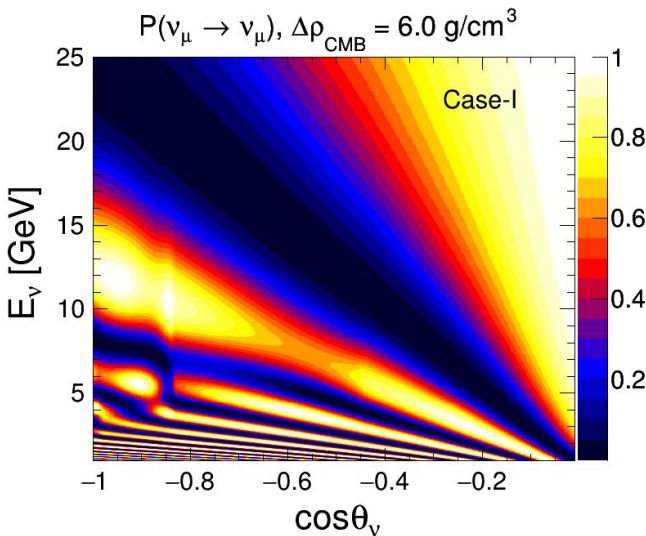
# Effect of $\Delta\rho_{\text{CMB}}$ Variation on Oscillograms

Smaller Jump (SJ)

Case-I

Larger Jump (LJ)

Observations:



• For SJ:

- NOLR/PR shift to left
- MSW patterns shrink to right

• For LJ:

- NOLR/PR shift to right
- MSW patterns stretched to left

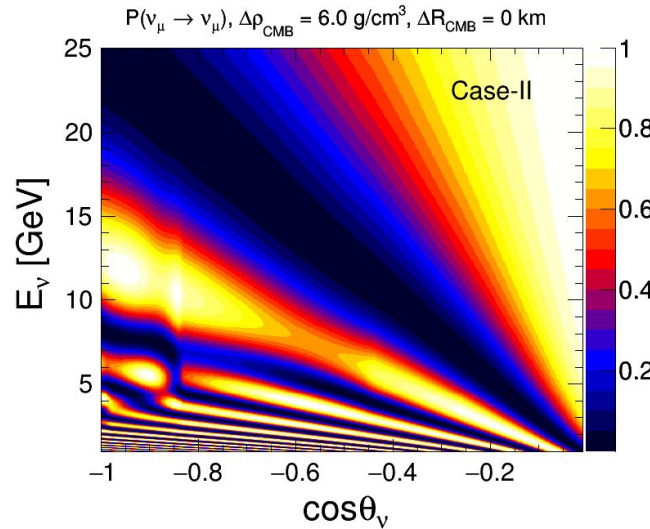
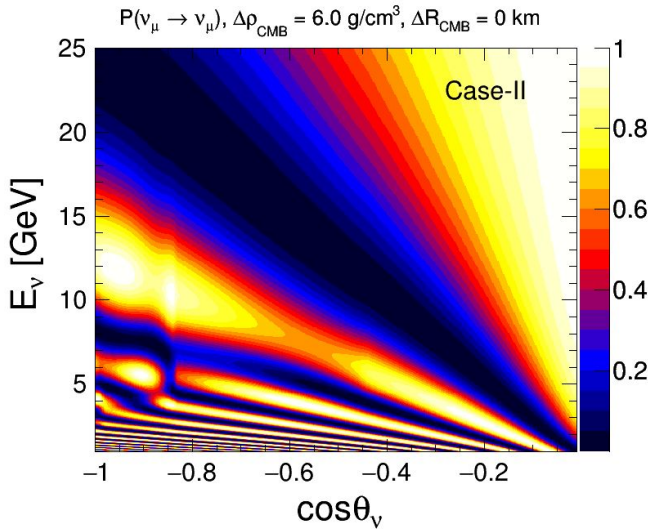
Opposite modification in MSW and NOLR/PR region for SJ and LJ scenarios

# Effect of $\Delta\rho_{\text{CMB}}$ & $R_{\text{CMB}}$ Variation on Oscillograms

Case-II

LJSC

SJLC



Observations:

- **For LJSC:**
  - NOLR/PR shifts to right and patterns shrink
  - MSW patterns shrink to right
- **For SJLC:**
  - NOLR/PR shifts to left and patterns broaden
  - MSW patterns stretched to left

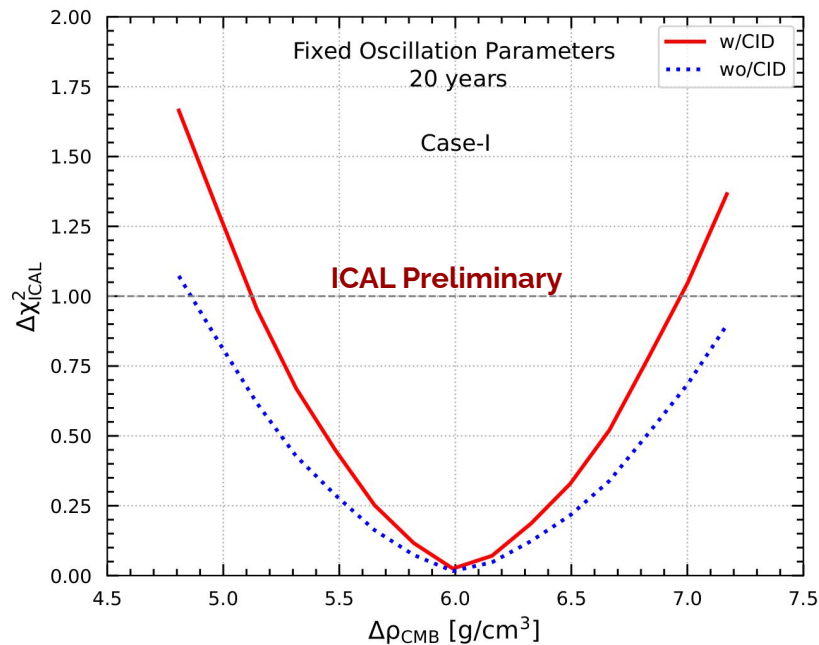
Opposite modification in MSW and NOLR/PR region for LJSC and SJLC scenarios



# Sensitivity for Measuring Density Jump $\Delta\rho_{\text{CMB}}$ at Standard

$R_{\text{CMB}}$

Case-I

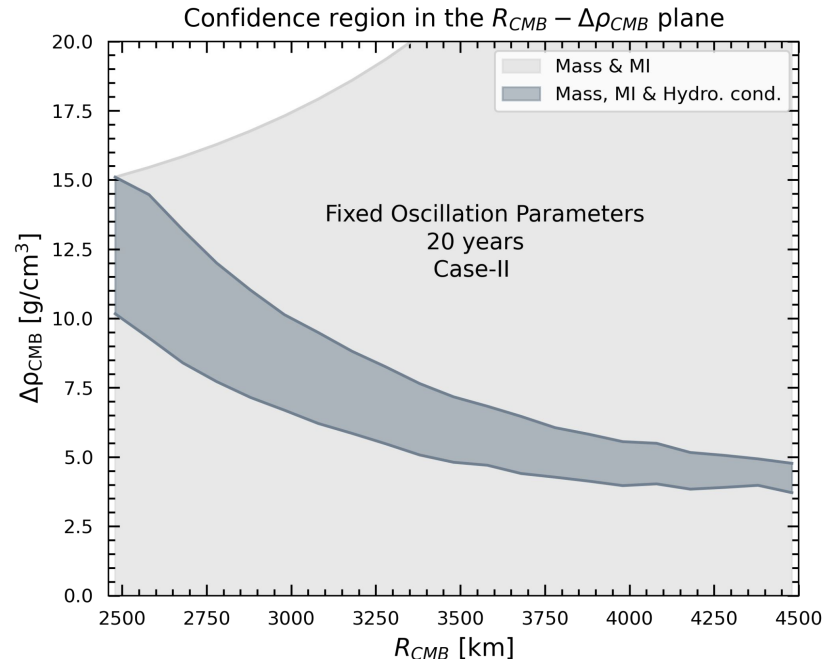


$6.0^{+0.97}_{-0.88} \text{ g/cm}^3$

Relative precision of about 15% at the  $1\sigma$  (w/CID)

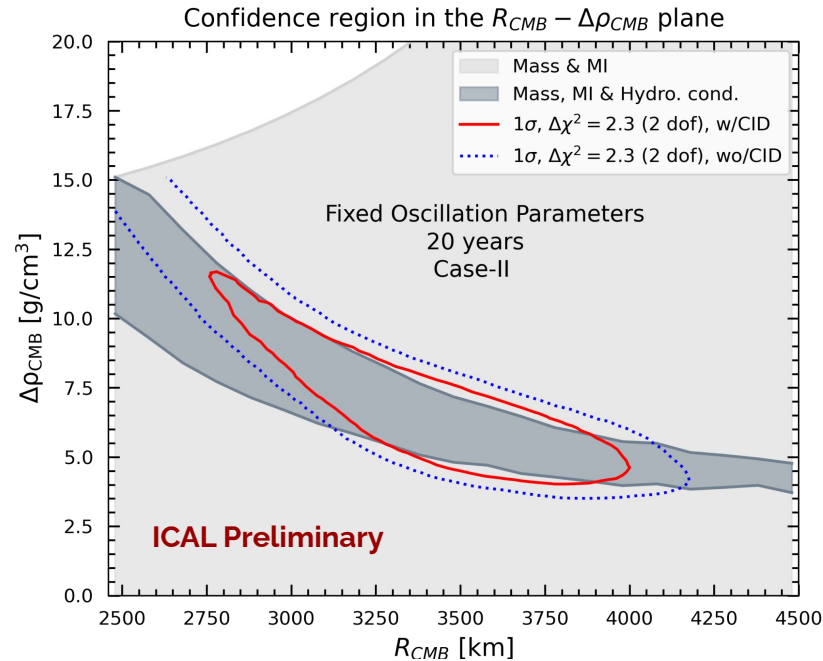
# Sensitivity for Measuring Correlation Between $\Delta\rho_{\text{CMB}}$ & $R_{\text{CMB}}$

## Case-II



# Sensitivity for Measuring Correlation Between $\Delta\rho_{\text{CMB}}$ & $R_{\text{CMB}}$

## Case-II

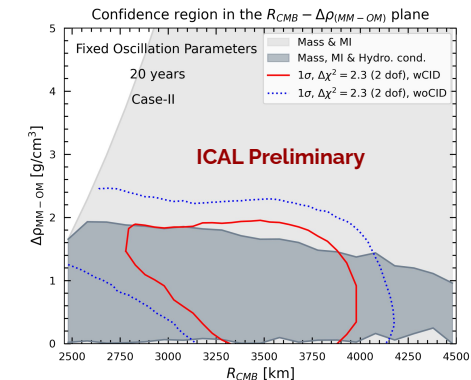
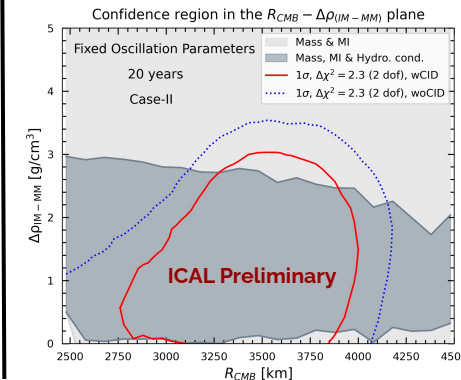
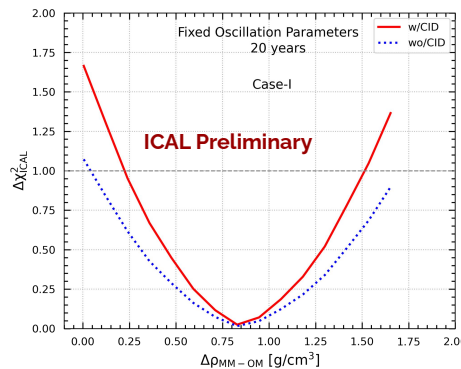
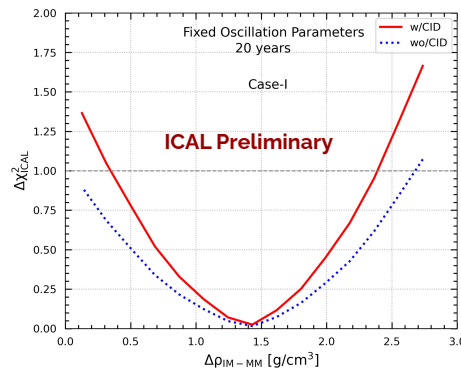
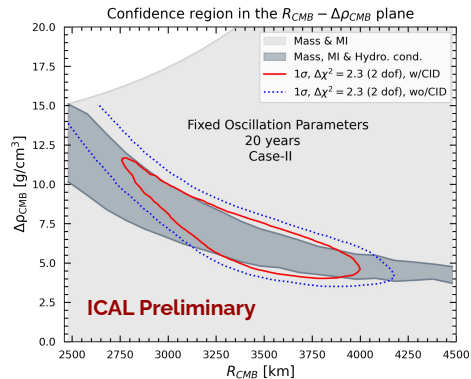
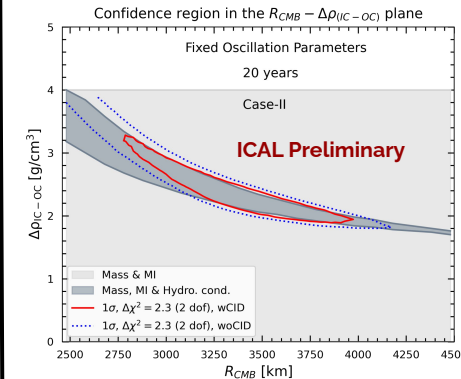
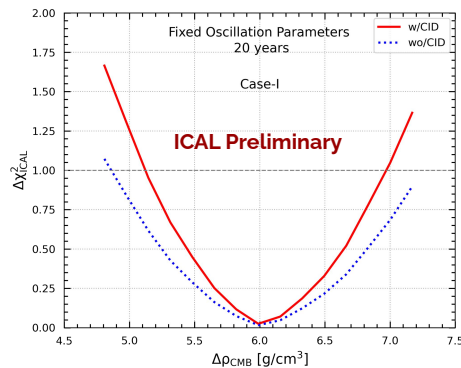
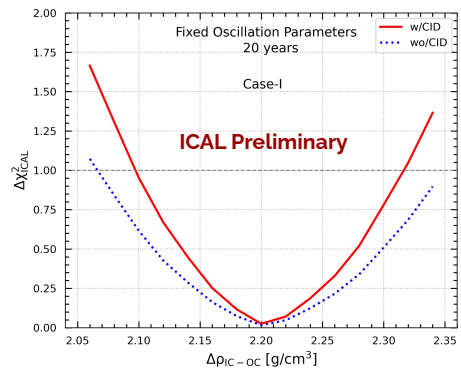


- Relative precision for **locating  $R_{\text{CMB}}$**  is around **9%** at the 1 $\sigma$  (w/CID) w.r.t standard  $R_{\text{CMB}}$
- Relative precision for constraining **density jump at CMB ( $\Delta\rho_{\text{CMB}}$ )** is around **25%** at the 1 $\sigma$  (w/CID) w.r.t standard  $\Delta\rho_{\text{CMB}}$

# Sensitivity for Measuring Correlation Between $\Delta\rho_{IC-OC}$ , $\Delta\rho_{CMB}$ , $\Delta\rho_{IM-MM}$ , $\Delta\rho_{MM-OM}$ , & $R_{CMB}$

## Case-I

## Case-II



# Summary

- In combination with gravitational and seismic studies, neutrino oscillations and absorption based measurements would pave the way for **“Multi-Messenger Tomography of Earth”**
- Atmospheric neutrinos have energies in the multi-GeV range where the Earth matter effects are significant, hence they would serve as probes of the internal structure of Earth
- ICAL can detect 331  $\mu^-$  and 146  $\mu^+$  core passing events in 10 years
- ICAL detector with 1000 kt·yr exposure would be able to **locate the CMB** with a precision of about  **$\pm 250$  km** at  $1\sigma$  confidence level
- ICAL detector with 1000 kt·yr exposure would be able to measure the density jump at inner-outer core boundary ( $\Delta\rho_{\text{IC-OC}}$ ) and CMB ( $\Delta\rho_{\text{CMB}}$ ) at the standard  $R_{\text{CMB}} = 3480$  km with a relative precision of about **5%** and **15%** at the  $1\sigma$  C.L, respectively
- ICAL would be able to constraint the allowed region between  $\Delta\rho_{\text{IC-OC}} - R_{\text{CMB}}$  around **6%** for  $R_{\text{CMB}}$  and **8%** for  $\Delta\rho_{\text{IC-OC}}$  w.r.t. standard  $\Delta\rho_{\text{IC-OC}}$  and  $R_{\text{CMB}}$  radius at  $1\sigma$  level with 1 Mt·yr exposure
- ICAL would be able to constraint the allowed region between  $\Delta\rho_{\text{CMB}} - R_{\text{CMB}}$  around **9%** for  $R_{\text{CMB}}$  and **25%** for  $\Delta\rho_{\text{CMB}}$  w.r.t. standard  $\Delta\rho_{\text{CMB}}$  and  $R_{\text{CMB}}$  radius at  $1\sigma$  level with 1 Mt·yr exposure

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We acknowledge the support from the Department of Atomic Energy (DAE), Department of Science and Technology (DST), and Science and Engineering Research Board (SERB), Govt. of India

**Thank you!**

# Backup

# Statistical Analysis

In this analysis, the  $\chi^2$  statistics is expected to give median sensitivity of the experiment in the frequentist approach.

$$\chi_-^2 = \min_{\xi_l} \sum_{i=1}^{N_{E' \text{ had}}^{\text{rec}}} \sum_{j=1}^{N_{E\mu}^{\text{rec}}} \sum_{k=1}^{N_{\cos \theta_\mu}^{\text{rec}}} \left[ 2(N_{ijk}^{\text{theory}} - N_{ijk}^{\text{data}}) - 2N_{ijk}^{\text{data}} \ln \left( \frac{N_{ijk}^{\text{theory}}}{N_{ijk}^{\text{data}}} \right) \right] + \sum_{l=1}^5 \xi_l^2$$

where,

$$N_{ijk}^{\text{theory}} = N_{ijk}^0 \left( 1 + \sum_{l=1}^5 \pi_{ijk}^l \xi_l \right)$$

Similarly,  $\chi_+^2$  is defined for  $\mu^+$

$$\chi_{\text{ICAL}}^2 = \chi_-^2 + \chi_+^2$$

# The Interior of Earth

Properties	Value ( $\pm$ )	Units
Earth's Mass	$(5.9722 \pm 0.006) \times 10^{24}$	kg
Mean moment of inertia	$(8.01736 \pm 0.00097) \times 10^{37}$	kg m <sup>2</sup>
Earth's mean radius	6371.23 (0.01)	km
Core-mantle boundary	3483 $\pm$ 5	km
Inner-outer core boundary	1220 $\pm$ 10	km

- [W.F. McDonough, Elsevier \(2003\), p. 547](#)
- [W. Chen et. al. Journal of Geodesy 89,179–188 \(2015\)](#)
- [B. Luzum et.al. Celestial Mechanics and Dynamical Astronomy 110, 293–304 \(2011\)](#)