



# Group meeting: Integration of WAGASCI in the OA

Andrés Muñoz  
21/03/2023

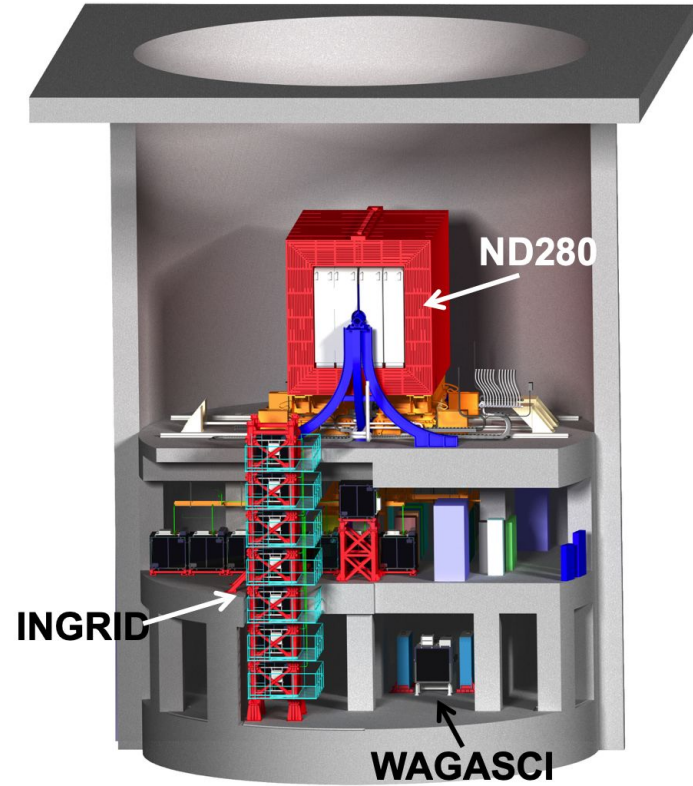




# Context & Introduction

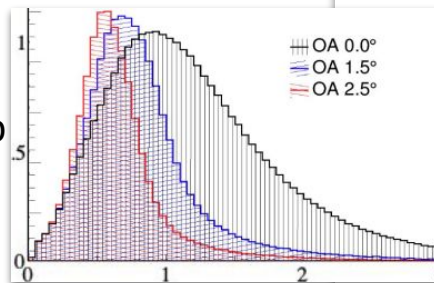
# Near detector complex

*REMINDER: since 2019, FGD2 is running with only 4 water bags (instead of 6)... but since 2019 we have WAGASCI/BM in place!*

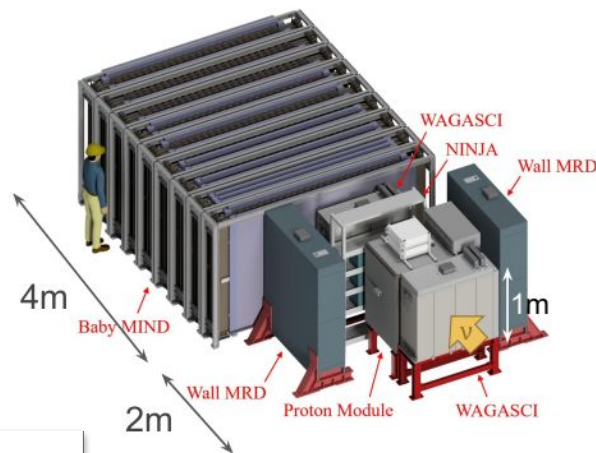


# WAGASCI

- Thus far only used for **xsec measurements** (link to detailed xsec analysis below)
- Could be interesting to exploit **water content** to constrain oxygen parameters
- A different **off-axis** could potentially help constrain flux systematics and energy dependant parameters



## WAGASCI, 1.5° off-axis



Recently added (2019)

Segmented cubic CH/H<sub>2</sub>O  
(WAGASCI) and  
SMRD+BabyMIND

**Magnetized detector**

Made of **80% of water (~0.5t)**<sub>4</sub>

# MOTIVATION

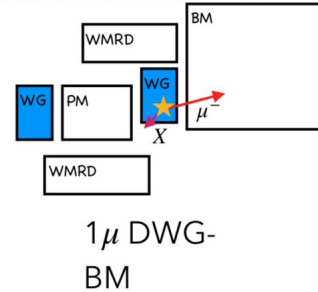
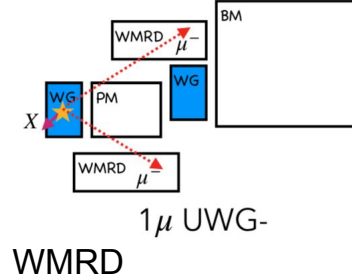
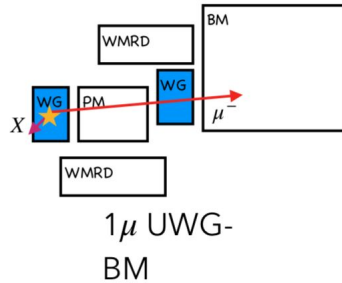
- The goal of this study is to perform sensitivity studies with the WAGASCI samples to see the additional constraints we might obtain in the OA (notably on spectral function oxygen parameters)  
<https://www.t2k.org/asg/meeting/2023/2023-02-02/waga>
- The first part of this study was **integrating** samples from WAGASCI (courtesy of Kenji Yasutome) to the analysis framework in the same format as the ND280 samples (GUNDAM for OA)
- Ongoing study consists of comparing post fit errors as a function of POT to see the constraints of adding WAGASCI samples



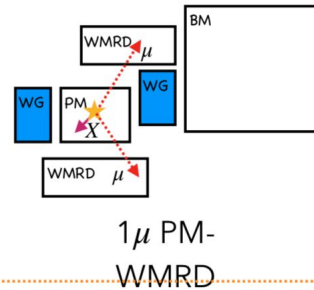
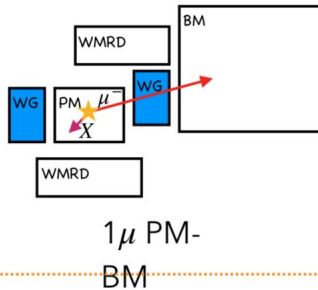
# Integrating WAGASCI samples for the OA

# WAGASCI Samples

## H<sub>2</sub>O Target



## CH Target



- Using samples defined by Kenji for his CC0pi xsec analysis (see his slides for more details)
- WAGASCI: 3 CC0pi + 1 CC1pi
- Proton module: 2 CC0pi + 1 CC1pi

[http://www128.orz.nsl280/physics/xsec/subgroup/mse-fdmc/2022/isaameeting\\_oct31/wagasci-babyblind-isa](http://www128.orz.nsl280/physics/xsec/subgroup/mse-fdmc/2022/isaameeting_oct31/wagasci-babyblind-isa)

# Generic fitter for Upgraded Near Detector Analysis Methods (GUNDAM)

- Fitter framework for the next statistical analysis of T2K
  - Suite of applications for statistical analysis developed for ND280 upgrade
  - Is becoming the official fitter for ND fit (OA) and cross section analyses
- Framework designed to host multiple analysis using JSON/YAML configuration file for better traceability and validation of output → inputs easily shared
- Open source (LGPL) C++ code based on ROOT publicly available on [GitHub](#)
- First part was integrating WAGASCI samples into GUNDAM
- Development of readable inputs (from Kenji's xsec inputs)

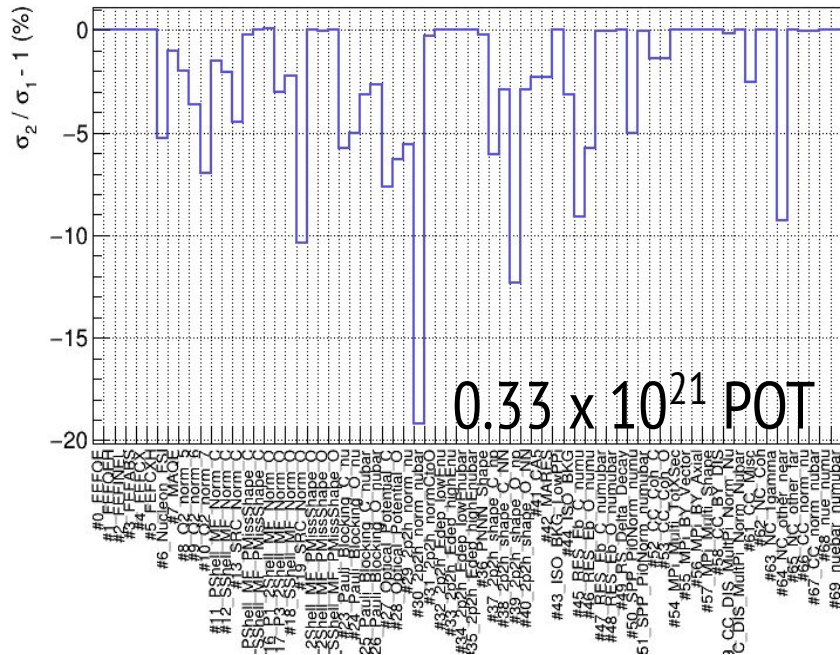


# New binning scheme

- Kenji's binning scheme adapted to his *Cross Section analysis*, only 1D in Pmu or Costhetamu
- Decided on a 2D binning that is more suitable to constrain systematic parameters
- **Idea:** make binnings where each bin has at least ~10 events per bin while taking into account detector resolution et reconstruction efficiency, see Kenji's TN [https://www.t2k.org/nd280/physics/xsec/xsecreviews/T2K-TN-455/review\\_for\\_fitter\\_method\\_fake\\_data\\_study/version\\_0\\_1\\_4/view](https://www.t2k.org/nd280/physics/xsec/xsecreviews/T2K-TN-455/review_for_fitter_method_fake_data_study/version_0_1_4/view)

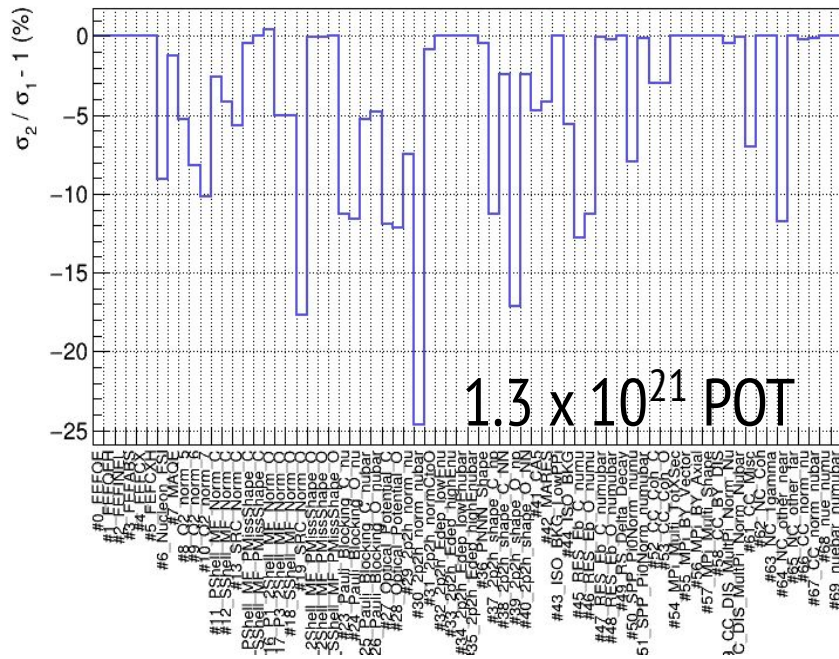
# Hesse post fit XSEC error ratio: 1D vs 2D binning

Comparing "Cross-Section Systematics" postFit parameters: "kenji"/Hesse [1] and "andresv2"/Hesse [2]



$\sigma_1$  = Kenji's 1D binning  
 $\sigma_2$  = My 2D binning

Comparing "Cross-Section Systematics" postFit parameters: "kenji"/Hesse [1] and "andresv2"/Hesse [2]



**Better constraint with  
the new (2D) binning!**



First fits with WAGASCI  
samples with GUNDAM

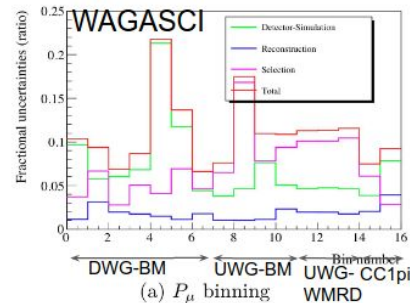
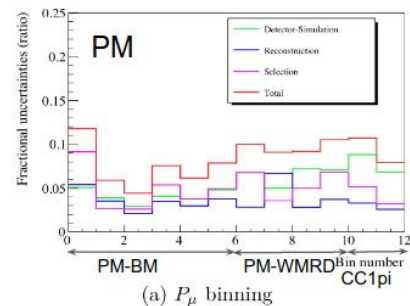
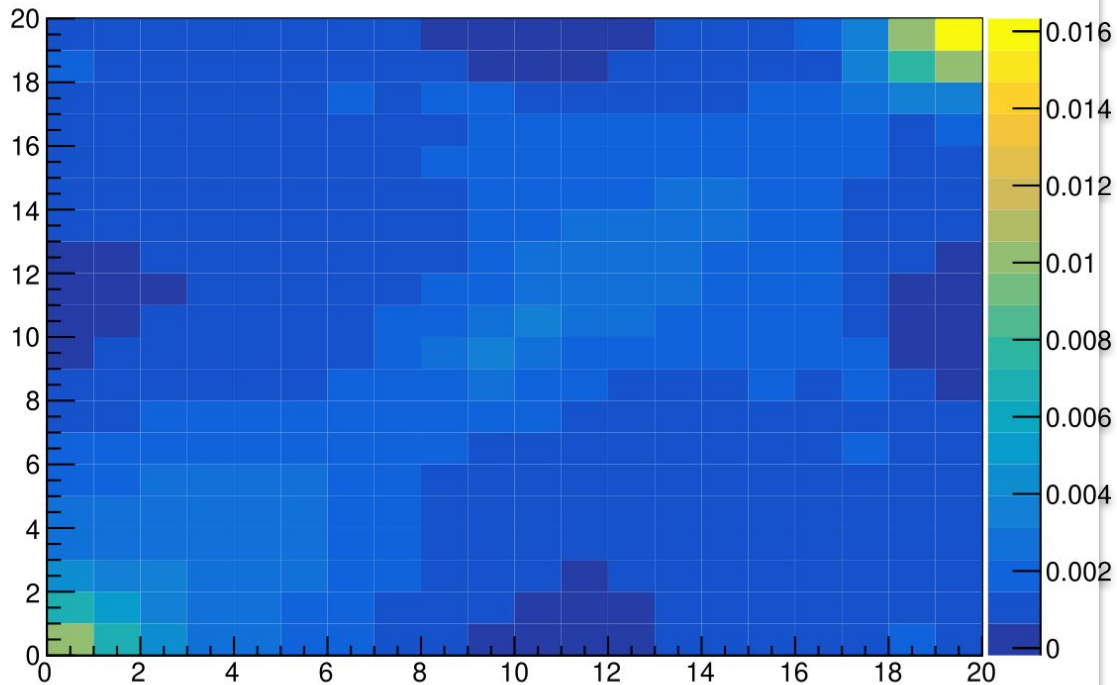


# WAGASCI fit

- Using Kenji's flux matrix (used in his xsec analysis)
- Cross-Section parameters same as in OA2022 (spline via WAGASCIReweight)
- Selected could be better adapted to current studies
- Detector smearing implemented directly in likelihood calculations (using detector performances in Kenji's TN [https://www.t2k.org/nd280/physics/xsec/xsecreviews/T2K-TN-455/review\\_for\\_fitter\\_method\\_fake\\_data\\_study/version\\_0\\_1\\_4/view](https://www.t2k.org/nd280/physics/xsec/xsecreviews/T2K-TN-455/review_for_fitter_method_fake_data_study/version_0_1_4/view) )
- Integration in GUNDAM complete and ready for future analysis

# WAGASCI fit

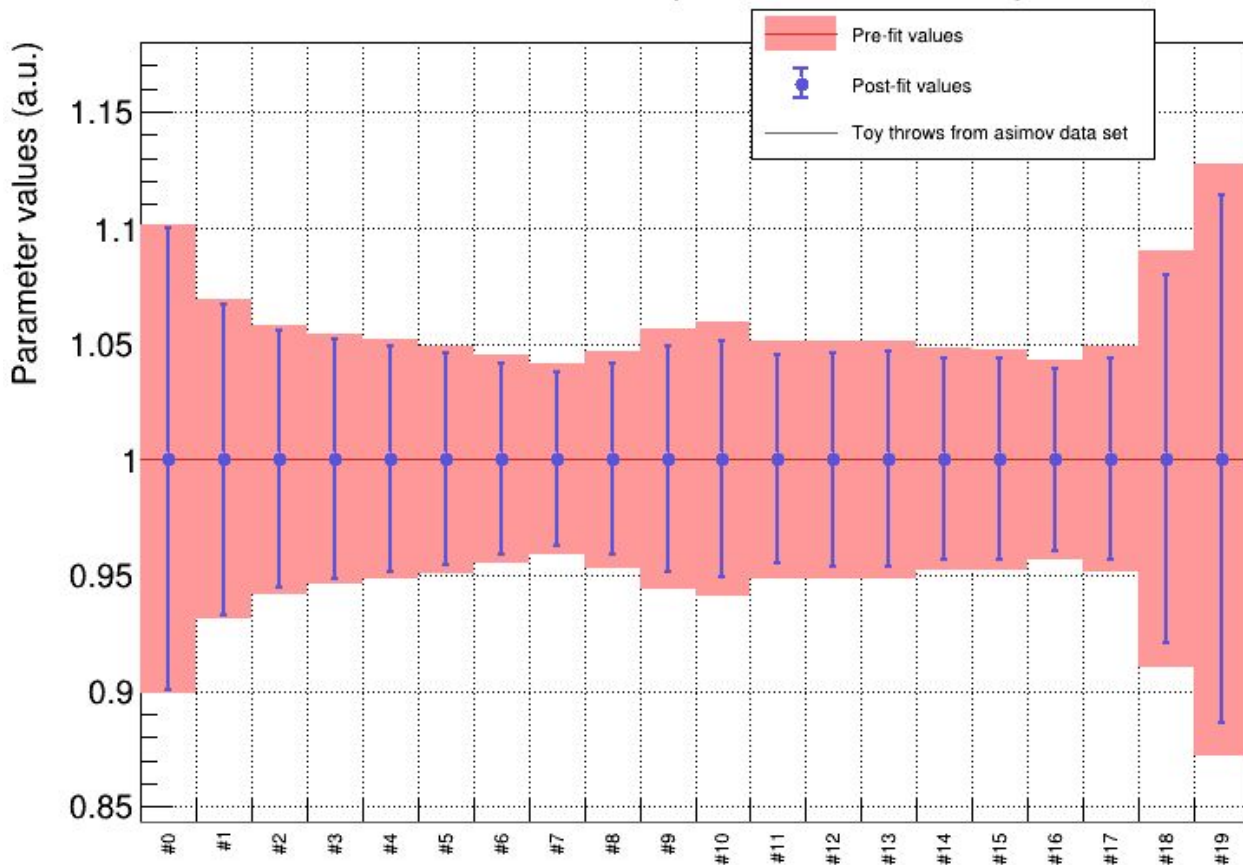
## Flux Cov Matrix



For sensitivity studies we can start by applying a  $\sim 10\%$  smearing factor on WAGASCI and PM samples

# Flux

## Pre-fit/Post-fit comparison for Flux Systematics

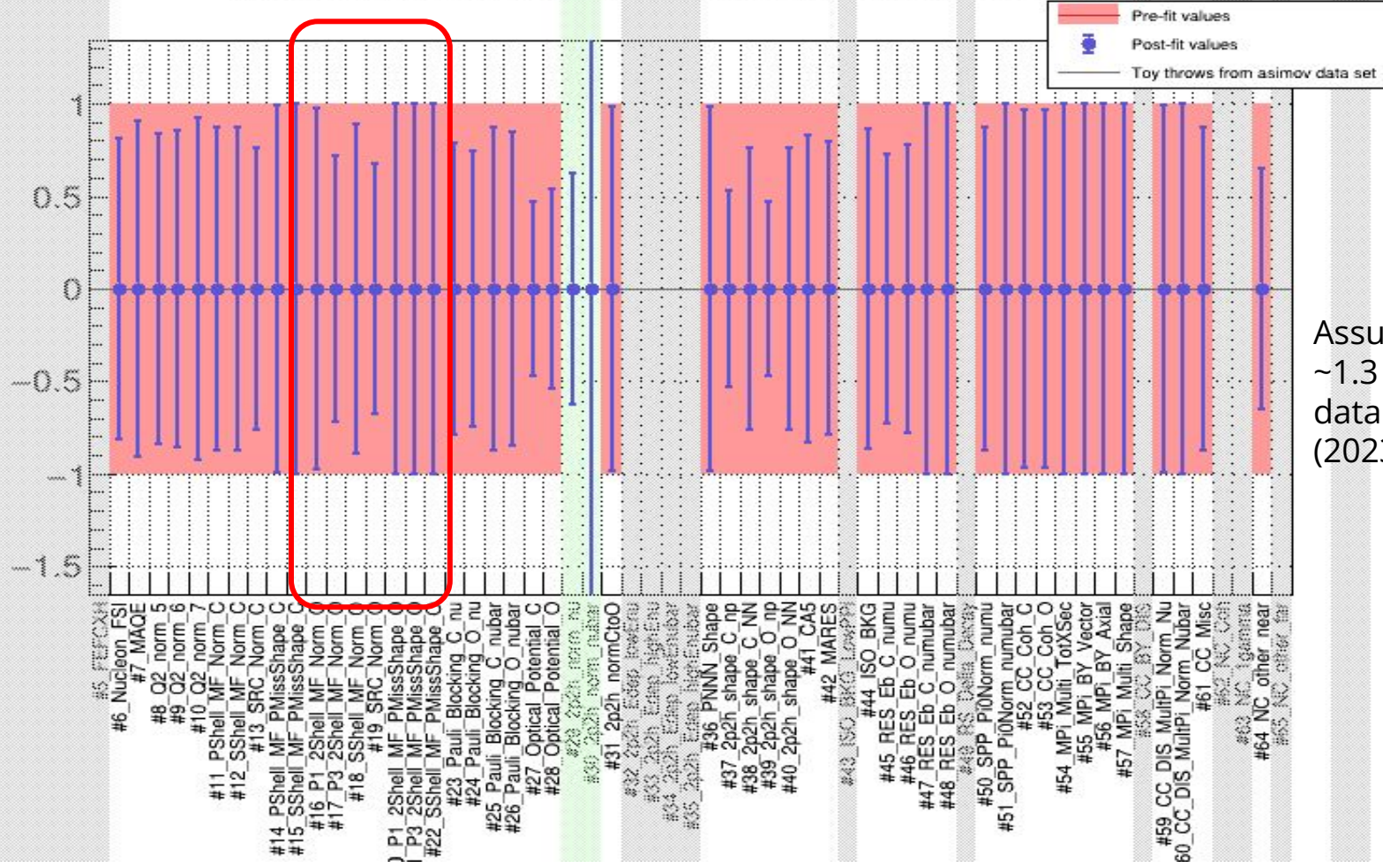


Assuming  $\sim 1.3$   
 $10^{21}$  data POT  
(2023 stat)



Parameter values (normalized to the prior)

### Pre-fit/Post-fit comparison for Cross-Section Systematics (normalized)



Assuming  
 $\sim 1.3 \cdot 10^{21}$   
data POT  
(2023 stat)



# First joint fit results

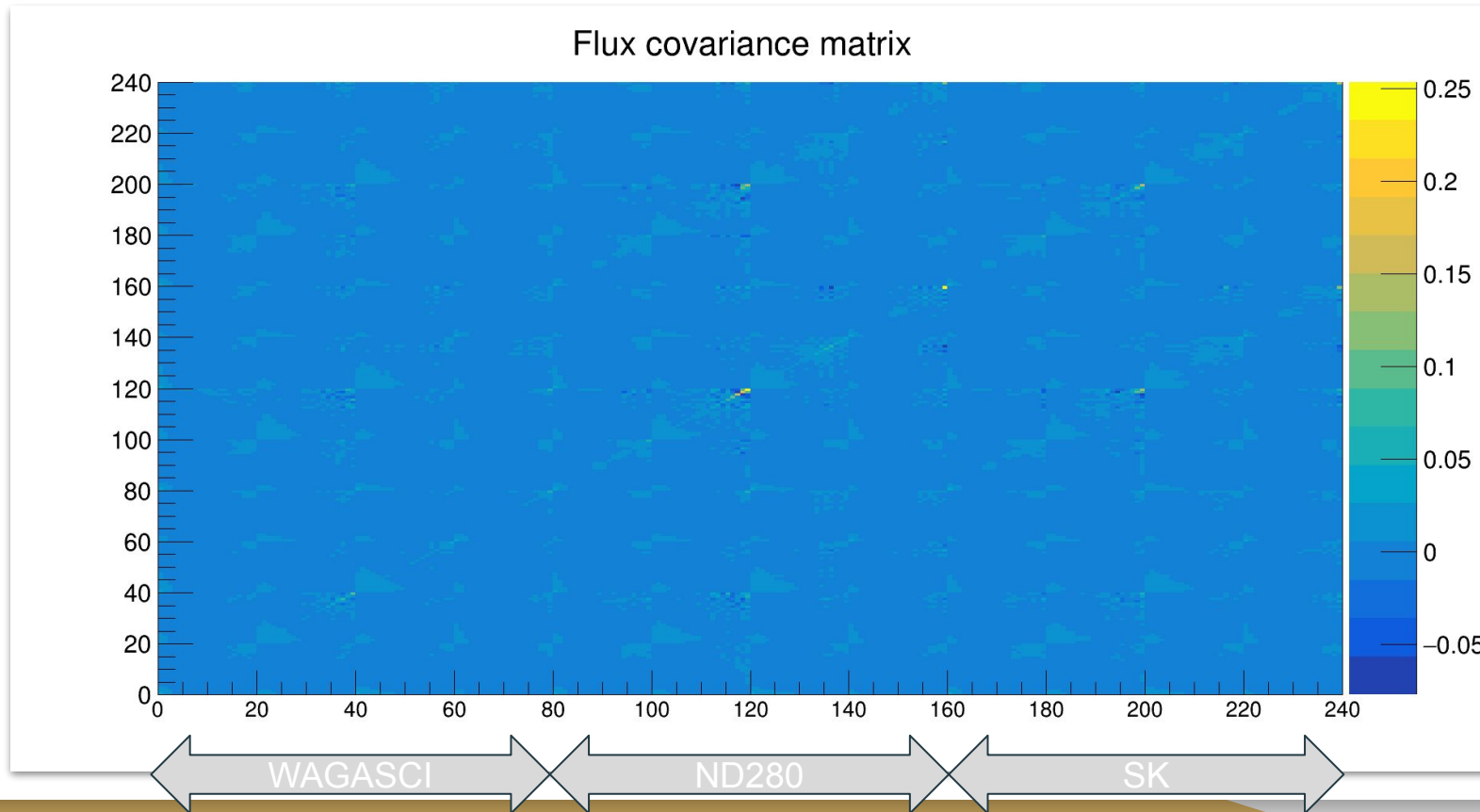


# Current fit configuration (still running)

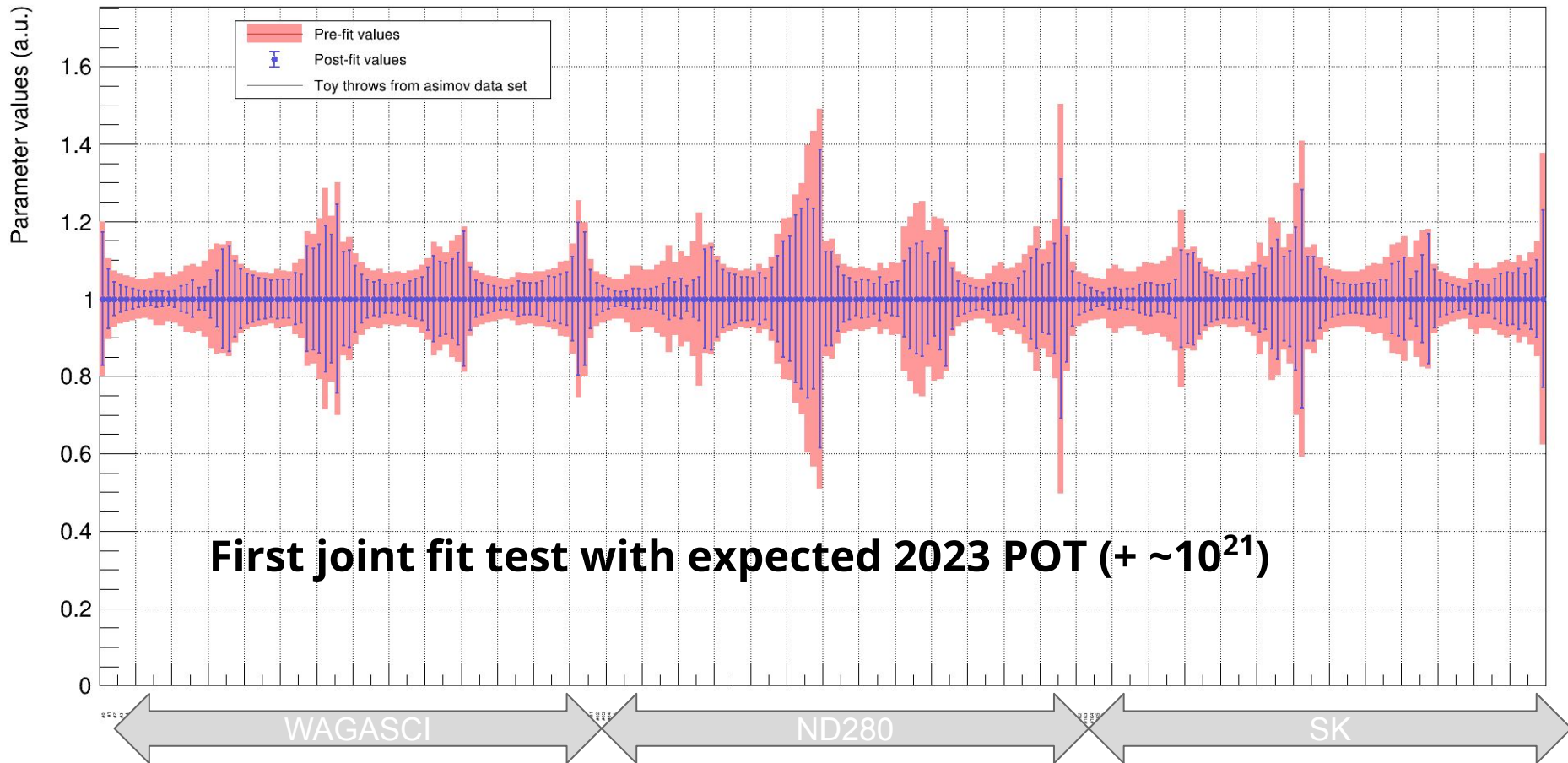
- Samples:
  - **FGD1/FGD2** (same rootfiles as OA2022)
  - **sFGD** (as in official sensitivity studies from Jaafar)
  - **WAGASCI** (as in Kenji's analysis, but 2D binning)
- So far, *CC 0pi* and *CC 1pi* only
- Detector smearing enabled for sFGD and WAGASCI
- Parameters: Flux and Cross-Sections

# Caspar's flux matrix

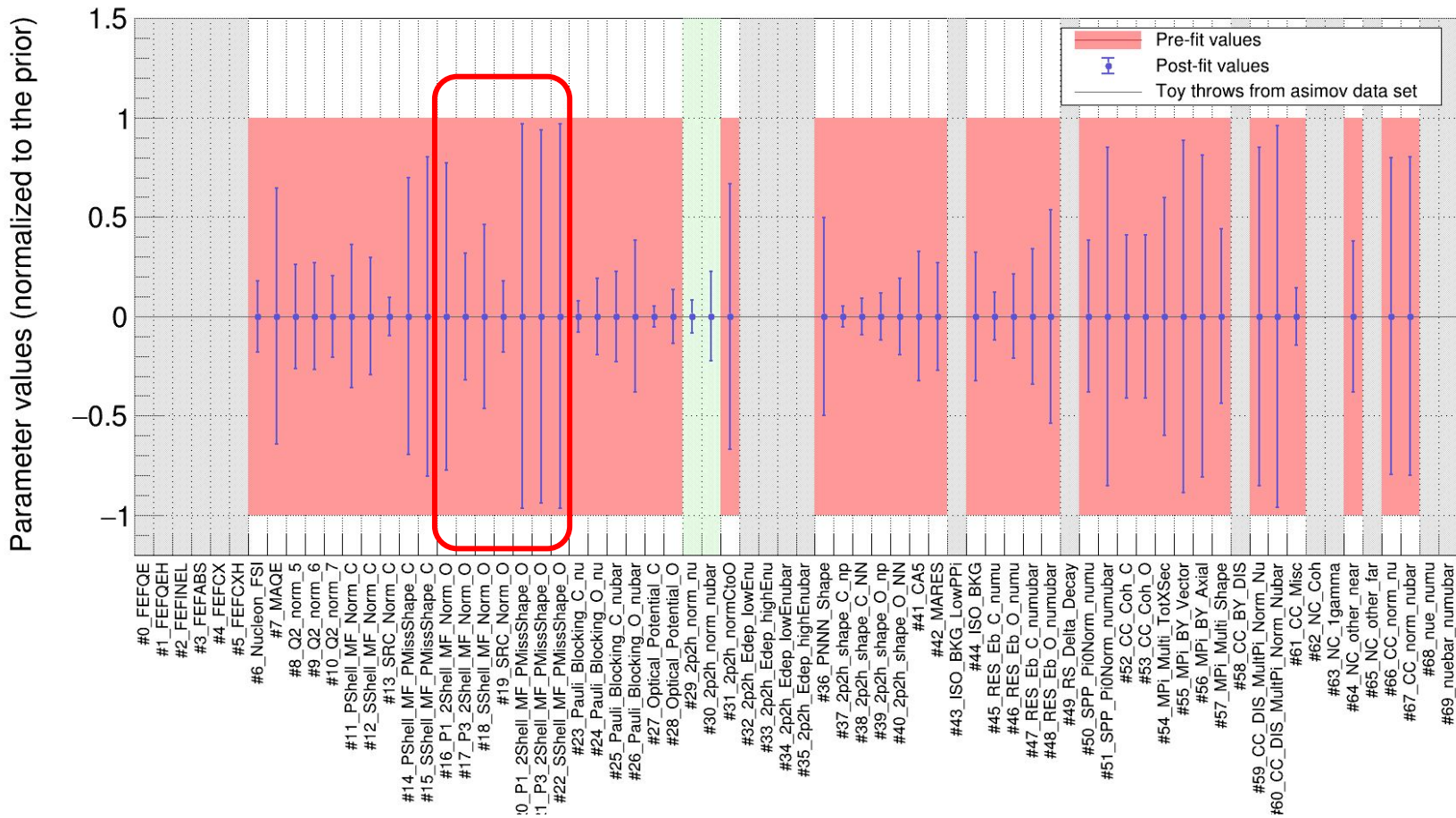
Caspar (xsec analyser/convener) developed a tool in collaboration with the beam group, to provide joint covariance matrices





# Pre-fit/Post-fit comparison for Flux Systematics



# Pre-fit/Post-fit comparison for Cross-Section Systematics (normalized)

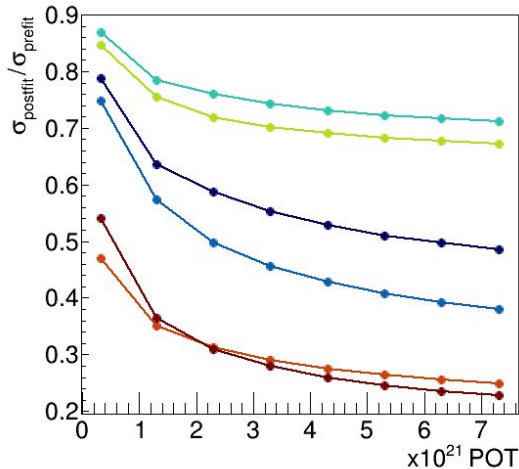




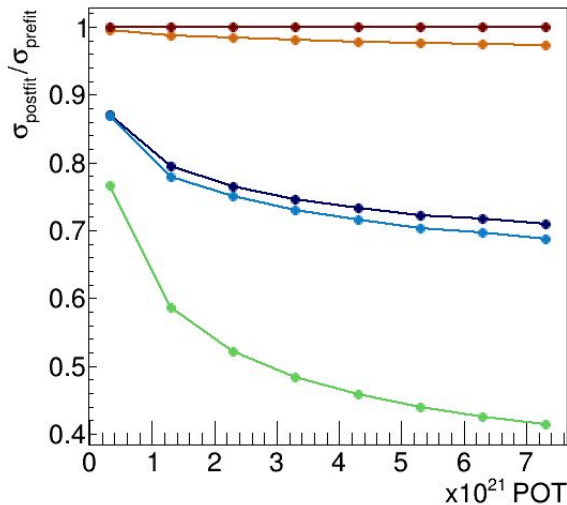
# Sensitivity studies with increasing POT

# POT studies WAGASCI

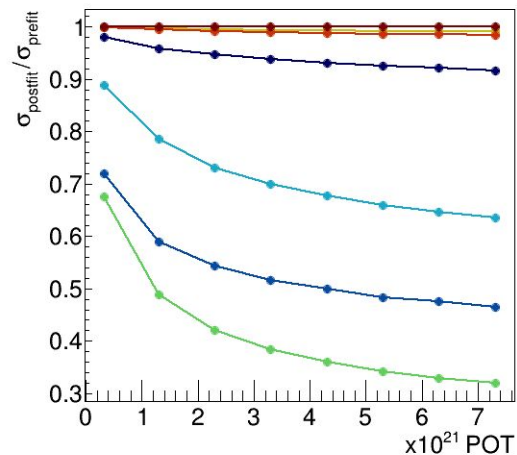
- Pauli Blocking C  $\nu$
- Pauli Blocking O  $\nu$
- Pauli Blocking C  $\bar{\nu}$
- Pauli Blocking O  $\bar{\nu}$
- Optical Potential C
- Optical Potential O



- P Shell MF Norm C
- S Shell MF Norm C
- SRC Norm C
- P Shell MF  $p_{\text{miss}}$  Shape C
- S Shell MF  $p_{\text{miss}}$  Shape C

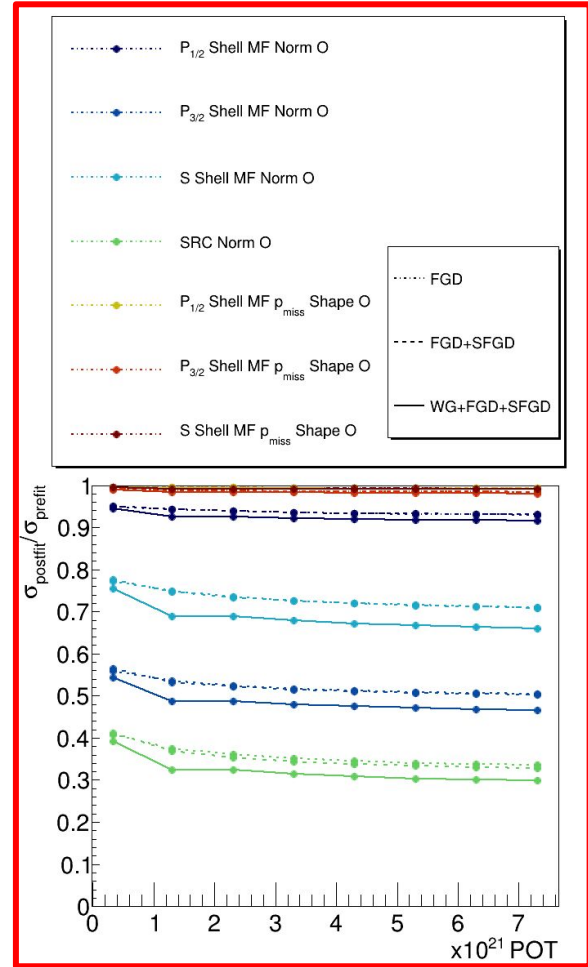
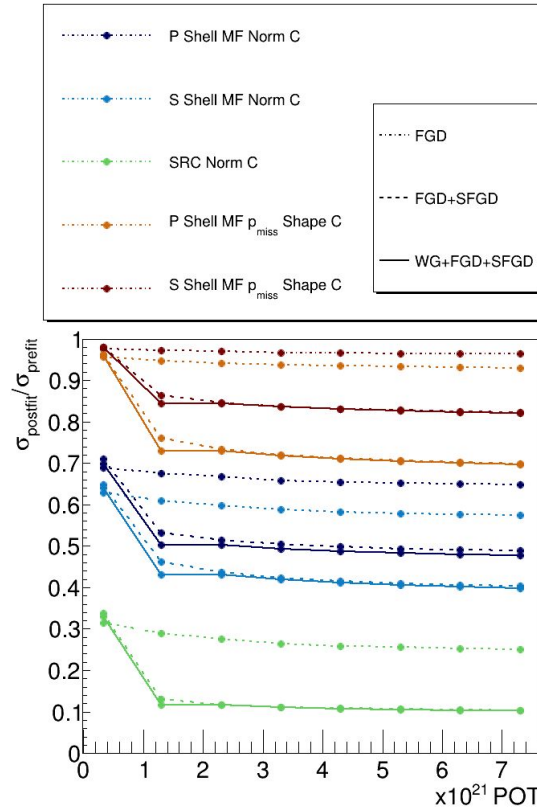


- $P_{1/2}$  Shell MF Norm O
- $P_{3/2}$  Shell MF Norm O
- S Shell MF Norm O
- SRC Norm O
- $P_{1/2}$  Shell MF  $p_{\text{miss}}$  Shape O
- $P_{3/2}$  Shell MF  $p_{\text{miss}}$  Shape O
- S Shell MF  $p_{\text{miss}}$  Shape O





# Adding WAGASCI to ND280 fit (Very preliminary)



# Conclusion

- The addition of WAGASCI in OA is definitely achievable: Inputs for sensitivity study ready and working
- Further work on the selection could give more promising results
- **Goal:** show ~complete results (with current selection) by next CM to motivate additional work around WAGASCI implementation in Highland and reconstruction/selection improvements



# What's next?

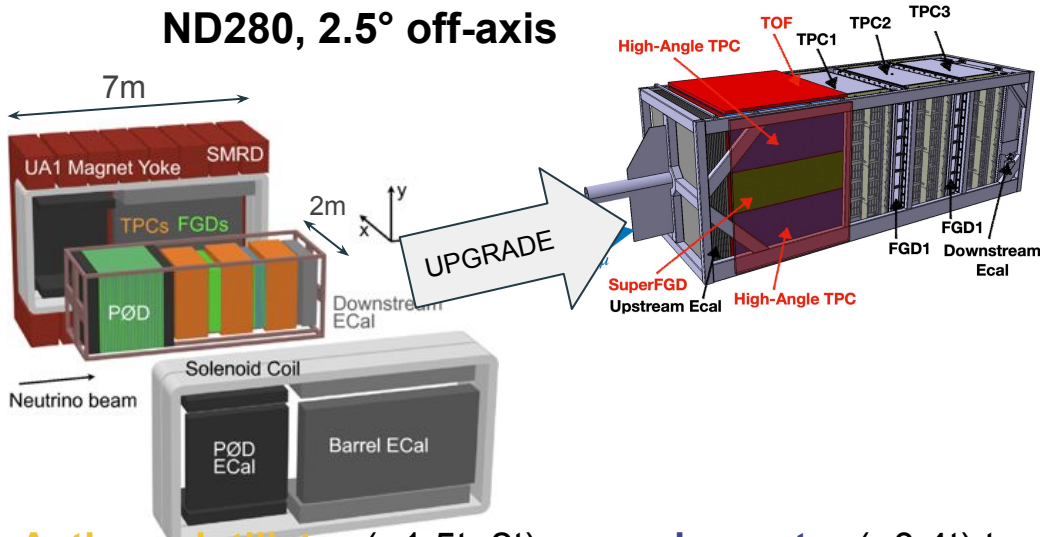
- We will need a new flux covariance matrix with the same format as OA
- We will try also adding prior O/C correlations (30%?)
- Ideally we would also like to propagate this to **SK** (ptheta)
- I'm planning to join the effort (Kenji, Cesar, John, Honjo-san,...) to improve current *WAGASCI reconstruction* (proton? Mom. by curvature?) and *selection*



BACK UP

# REMINDER: T2K off-axis near detectors

## ND280, 2.5° off-axis



Active scintillator (~1.5t+2t) + passive water (~0.4t) targets

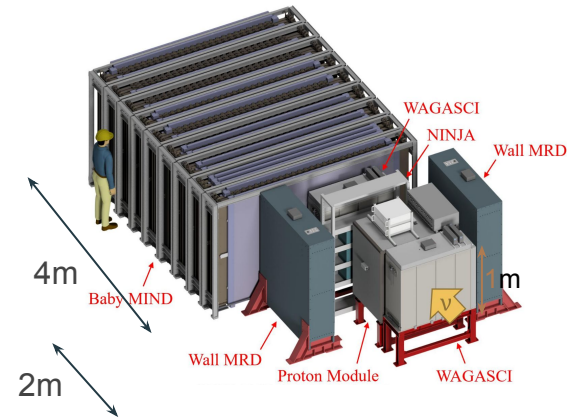
Tracking with 3 TPC

Magnetized for charge and momentum measurements

**Ecal** to distinguish tracks from showers

Used for OA and xsec measurements

## WAGASCI, 1.5° off-axis



Recently added (2019)

Segmented cubic CH/H<sub>2</sub>O  
(WAGASCI) and SMRD+BabyMIND

Magnetized detector

Made of ~0.2t-1t water and ~0.5t of  
CH target

# Sensitivity at the ND

- FGD1+2 : Current ND fit, no additional samples
- ..... SFGD+FGD1+2  $\mu$  only : Add to current ND fit SFGD samples binned in lepton kinematics
- SFGD+FGD1+2  $\mu$ +N : Add to current ND fit SFGD samples binned in (Evis,  $\delta p_T$ )



J. Chakrani

CCQE parameters,  
no O/C correlation

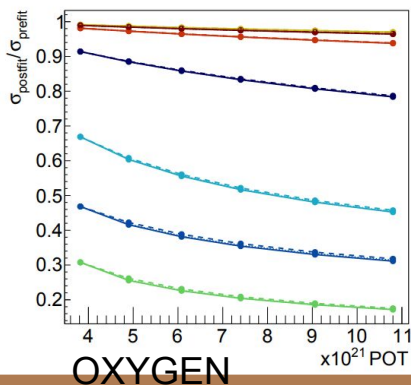
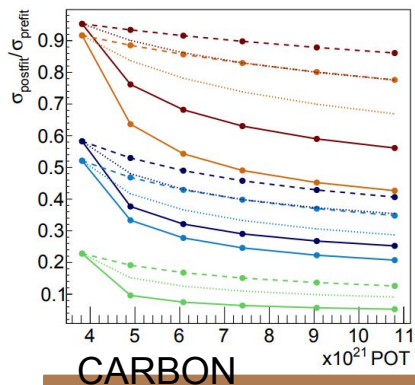
- P Shell MF Norm C
- S Shell MF Norm C
- SRC Norm C
- P Shell MF  $p_{miss}$  Shape C
- S Shell MF  $p_{miss}$  Shape C

- P<sub>1/2</sub> Shell MF Norm O
- P<sub>3/2</sub> Shell MF Norm O
- S Shell MF Norm O
- SRC Norm O
- P<sub>1/2</sub> Shell MF  $p_{miss}$  Shape O
- P<sub>3/2</sub> Shell MF  $p_{miss}$  Shape O
- S Shell MF  $p_{miss}$  Shape O

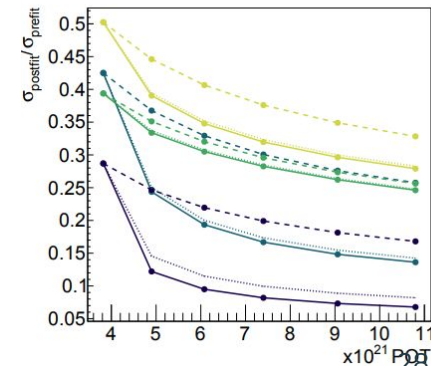
The sensitivity to oxygen xsec systematics depends on the assumed O/C correlation

CCQE parameters,  
O/C correlation (~70%)

- E<sub>b</sub> C  $\nu$
- E<sub>b</sub> C  $\bar{\nu}$
- E<sub>b</sub> O  $\nu$
- E<sub>b</sub> O  $\bar{\nu}$

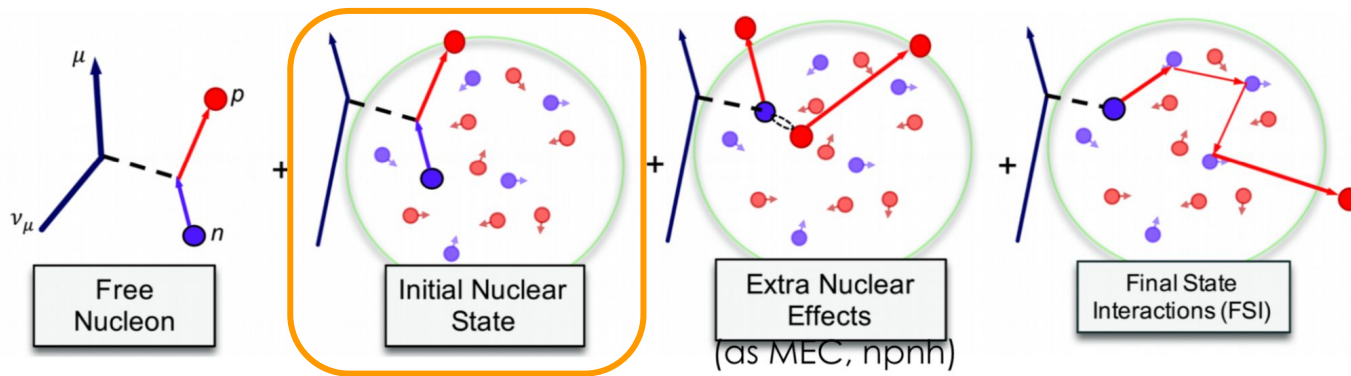


Adding SuperFGD (CH) events helps constraining oxygen systematics ONLY IF a prior correlation between O/C is known



# Parametrized estimation of uncertainties for exclusive predictions of neutrino-nucleus scattering: Initial Nuclear State

Jaafar



Parameterisation validated against existing xsec data

Used already in [T2K OA 2022](#) and in [Phys. Rev. D 105, 032010 \(2022\)](#)

Paper in preparation, expected in 2023

based on electron scattering data

