



# Development of a cryogenic veto system for CEvNS detection

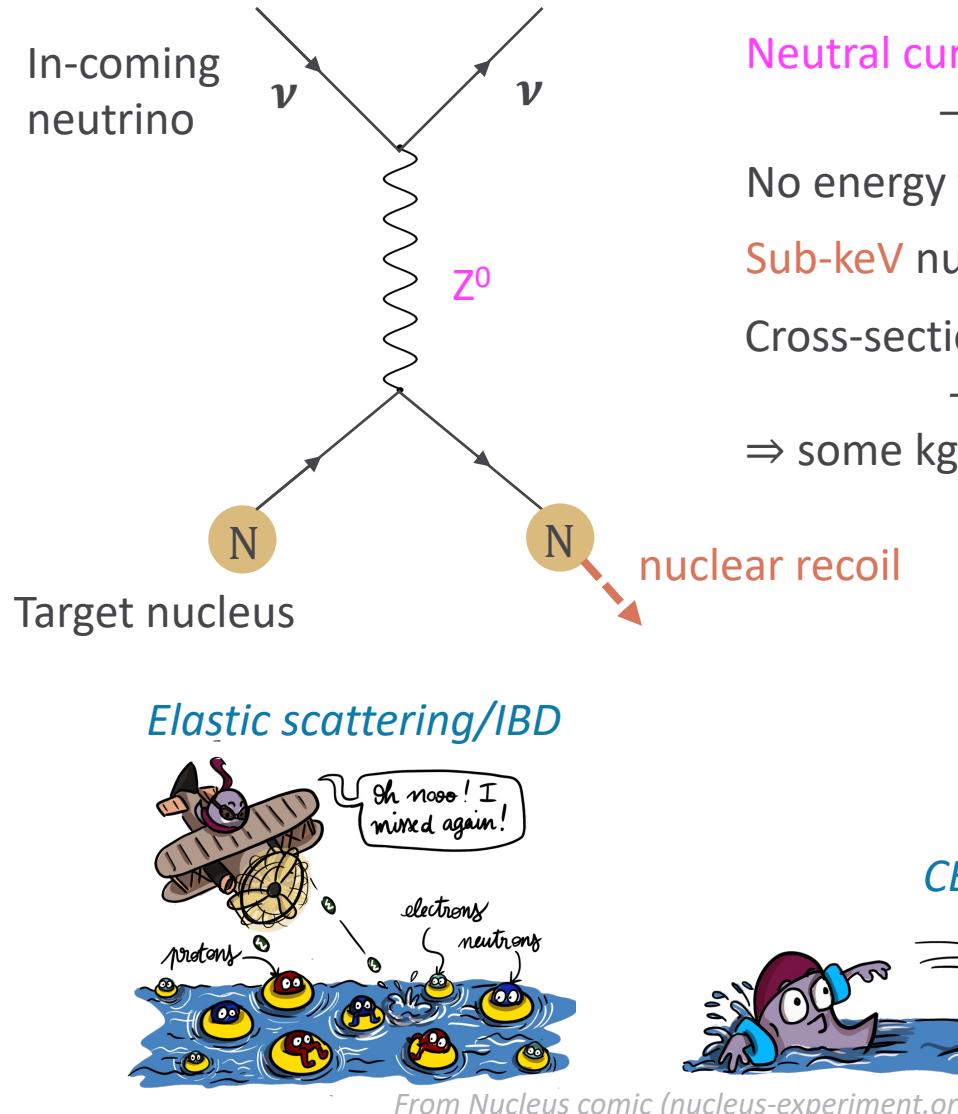
in the scope of the NUCLEUS experiment

P2IO BSM-Nu workshop, May 24th, 2023



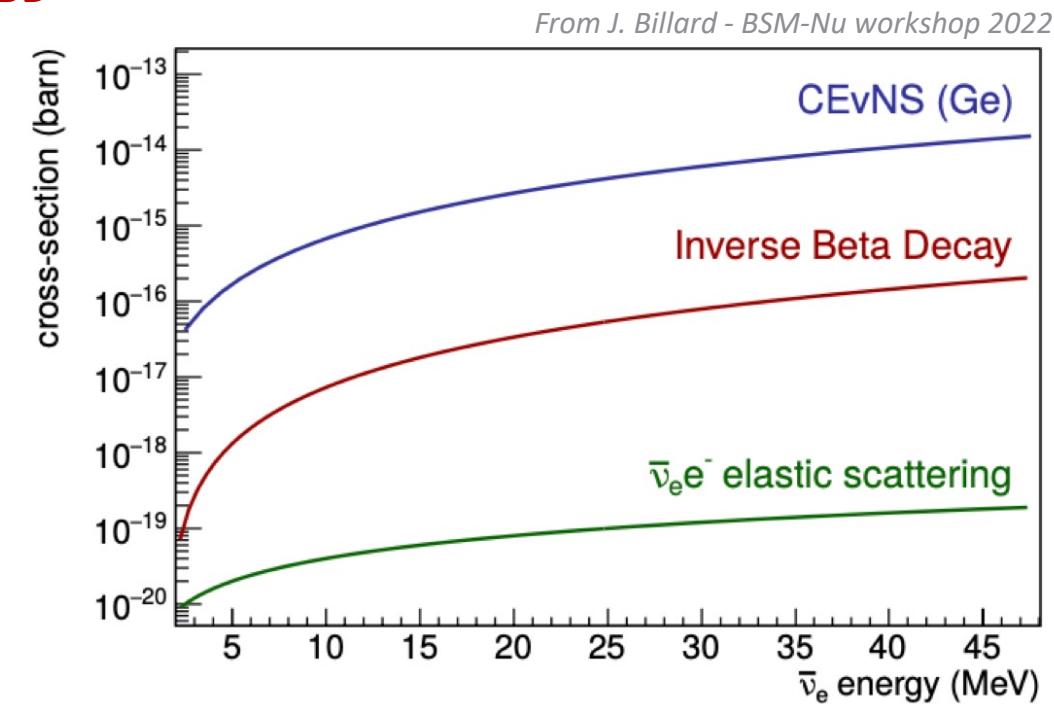
Chloé Goupy  
IRFU, CEA, Université Paris-Saclay

# Coherent Elastic Neutrino-Nucleus Scattering (CE $\nu$ NS)



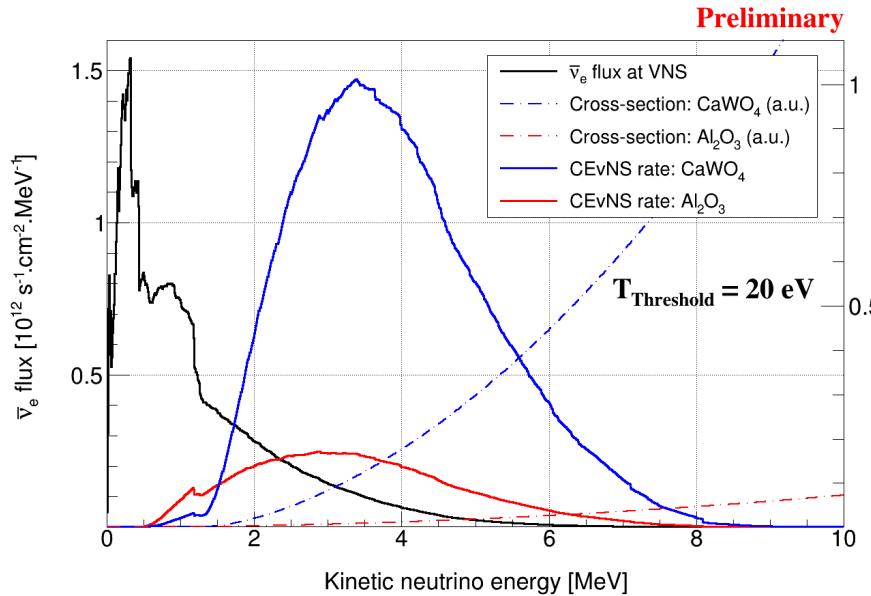
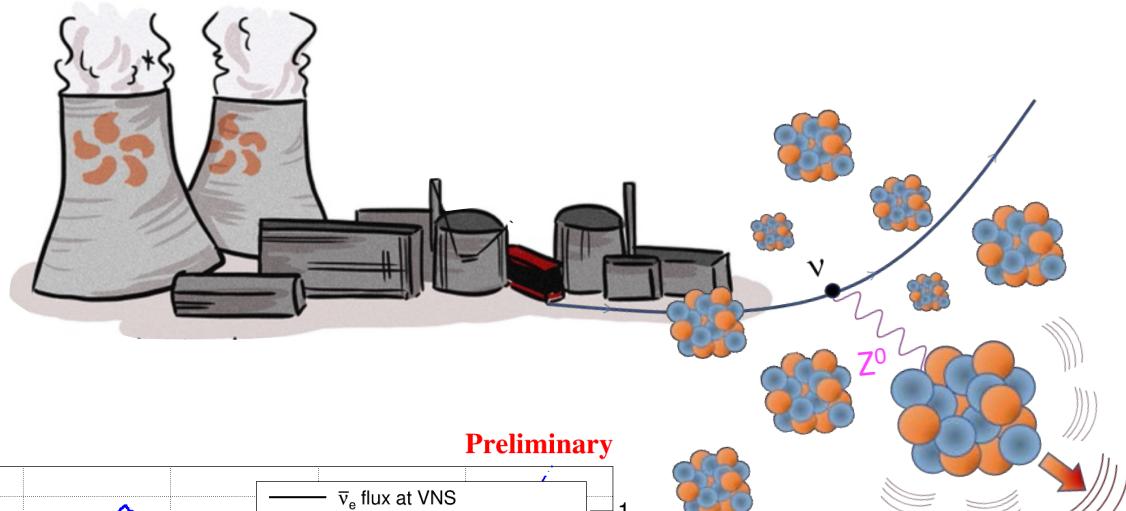
Neutral current interaction  
→ flavor independent  
No energy threshold  
Sub-keV nuclear recoils  
Cross-section proportional to  $N^2$   
→ 1000x larger than IBD  
⇒ some kg/g-scale detectors

$$\frac{d\sigma}{dE_r} = \frac{G_F^2}{4\pi} Q_w^2 F^2(q^2) m(Z, N) \left(1 - \frac{E_r}{E_{r,max}}\right)$$



# Study CE $\nu$ NS from reactor (anti-)neutrinos

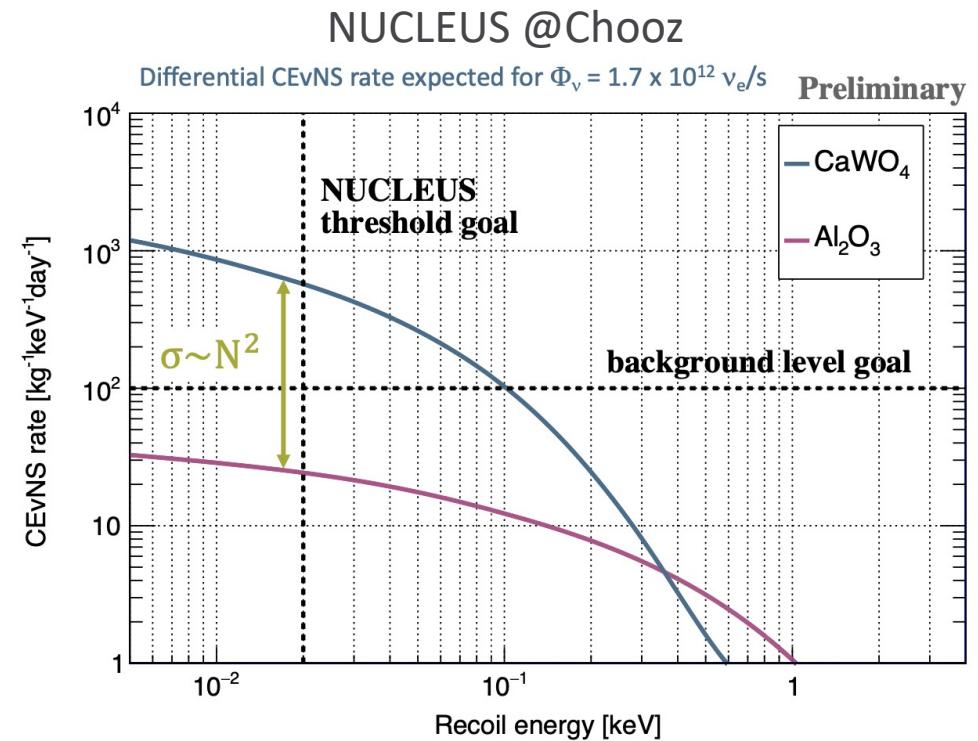
## Coherent Elastic Neutrino-Nucleus Scattering



Nuclear reactors: intense sources of  $\bar{\nu}_e$

$E_\nu < 10 \text{ MeV} \rightarrow$  fully coherent regime

→ Low thresholds detectors and low background counting rate required

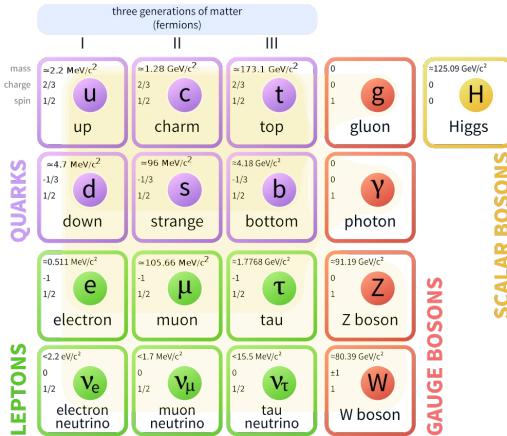


# CE $\nu$ NS, what for?

## 1- Probe for Standard Model

CE $\nu$ NS cross-section is a clean standard model prediction

### Standard Model of Elementary Particles

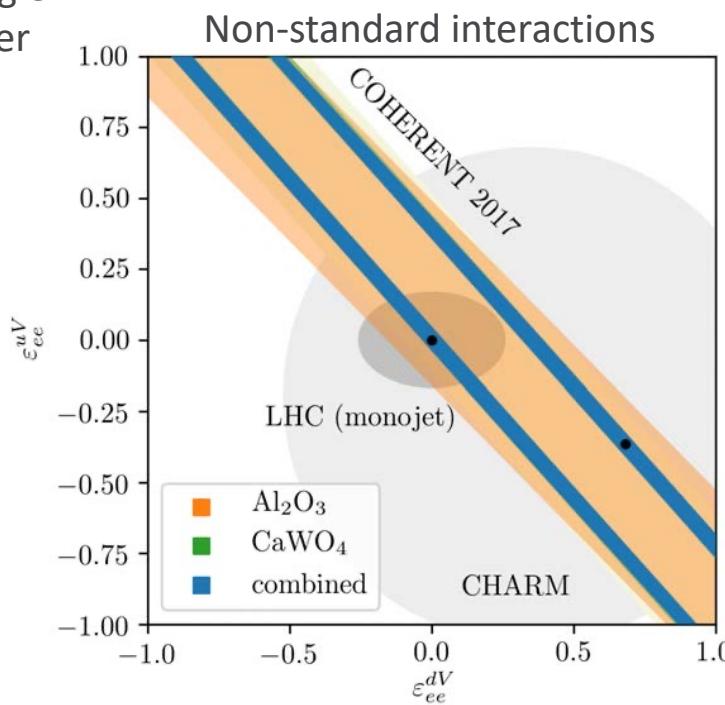
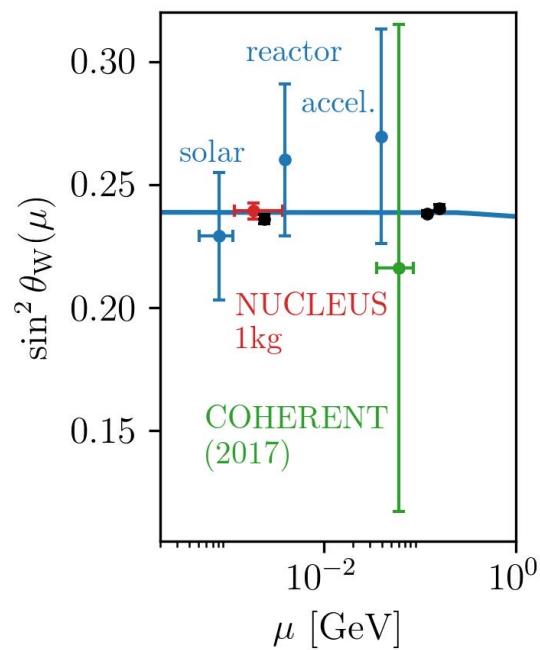


If deviations  
wrt SM

## 2- Beyond the SM

Sterile neutrino  
Neutrino electro-magnetic properties  
(e.g. magnetic dipole moment)

Measure the Weinberg angle  
at low momentum transfer



## 3- For Dark matter experiments:

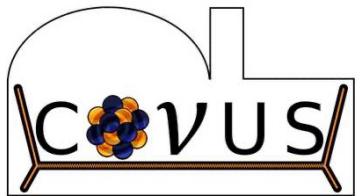
CE $\nu$ NS of solar or atmospheric neutrinos: irreducible background for experiment looking for WIMPs

$\Rightarrow$  "neutrino floor"

... and more!

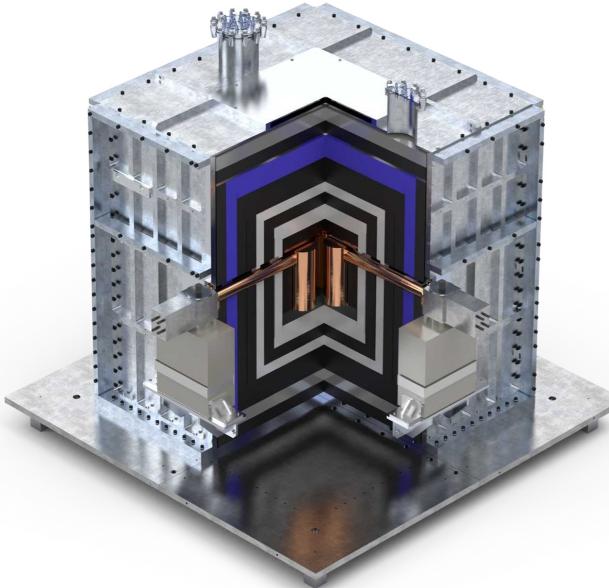
- long range detection:
- Supernovae neutrinos
  - Solar neutrinos
  - Reactor monitoring

# Reactor CE $\nu$ NS experiments

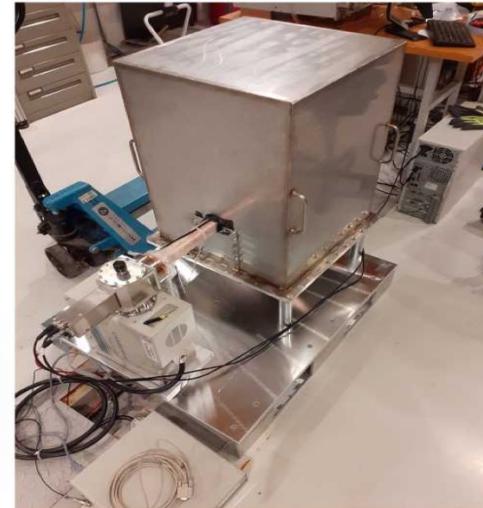


@Brokdorf reactor (KBR) (Germany)

Significant **overburden**  $\approx 24$  m.w.e.



**Target detectors:**  
Ge  $\rightarrow$  ionization  
(threshold:  $1\text{keV}_{\text{nr}}$ )

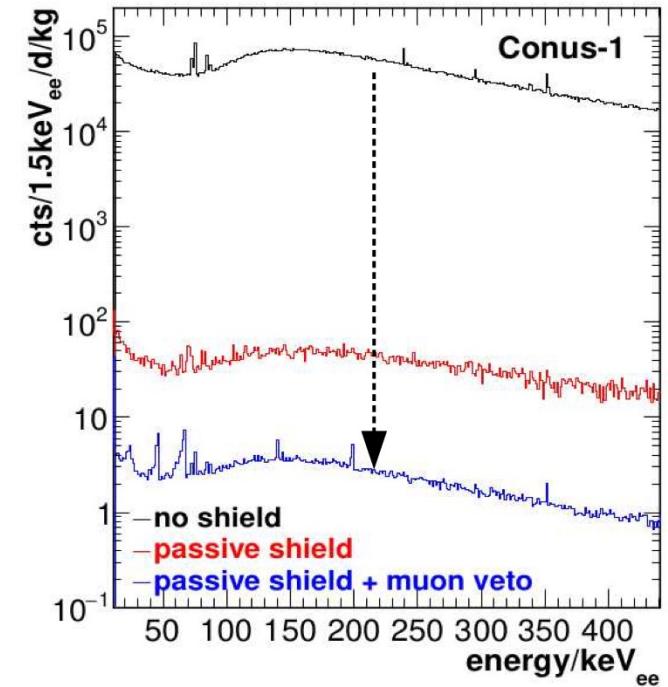


## Outer Shielding

- Borated Polyethylene
- 25 cm radiopure Pb
- Muon Veto
- Stainless Steel

## Background measurement

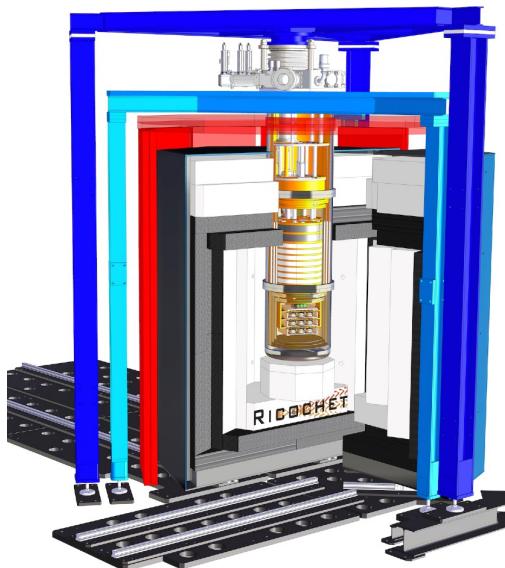
Overall background suppression via **passive** and **active** shield



total bkg suppression (w/o PSD):  $>10^4 \times$   
remaining bkg rate in ROI:  $O(10)$  cts/d/kg  
i.e.  $(0.3, 1.0)$  keV<sub>ee</sub>

From W. Maneschg (Magnificent CE $\nu$ NS, March 2023)

# Reactor CE $\nu$ NS experiments



**Target detectors:**  
Ge (& Si ?)  
→ ionization and heat  
(target RMS:  $20 \text{ eV}_{\text{ee}}$ ,  $10 \text{ eV}_{\text{nr}}$ )

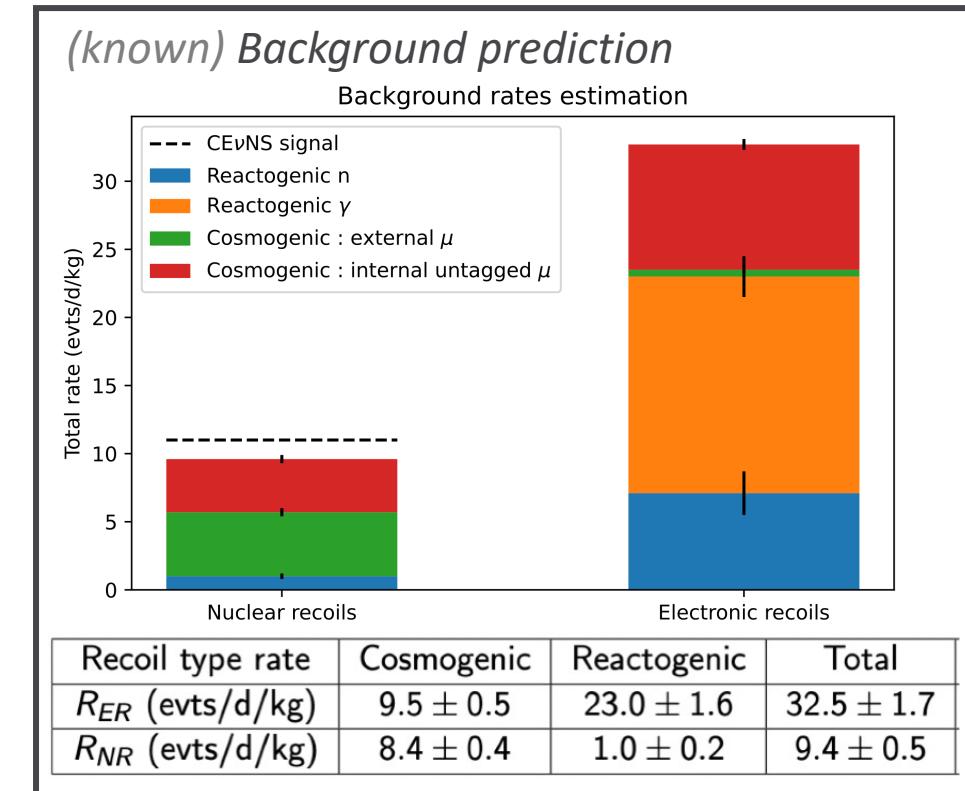
@ILL-H7 nuclear reactor site (Grenoble)

Significant overburden  $\approx 15 \text{ m.w.e.}$



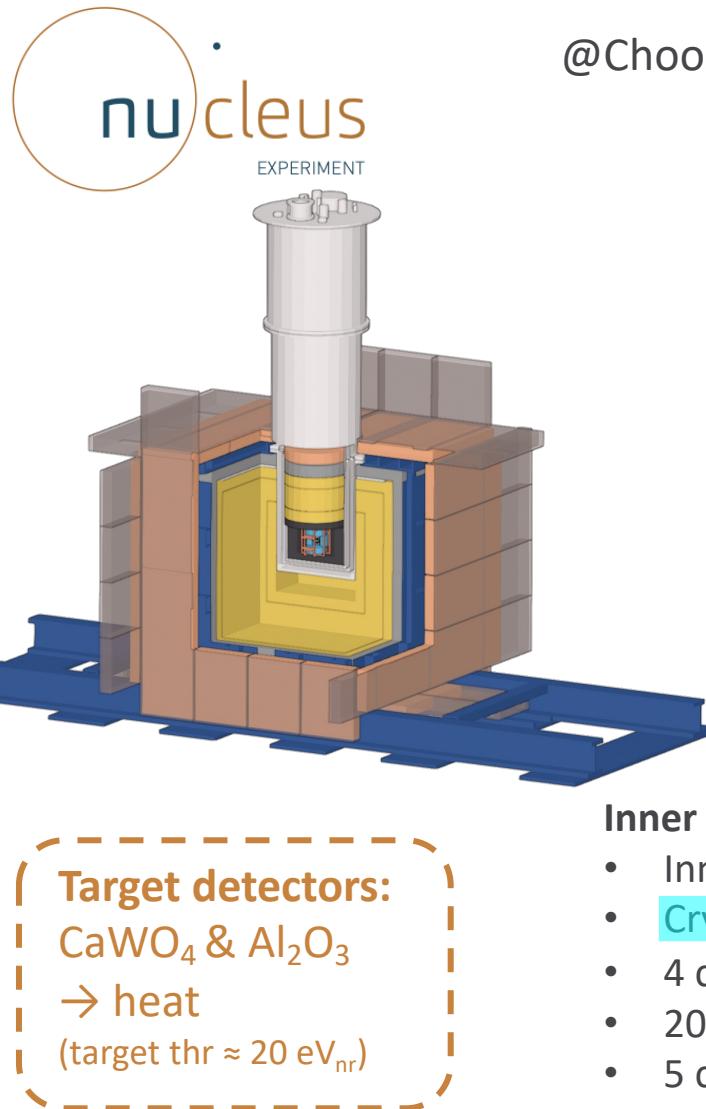
- Inner Shielding:**
- 30 cm PE/Cu
  - 15 cm Pb/Cu
  - Cryogenic Muon Veto
  - Mu-Metal

- Outer Shielding**
- 35 cm PE
  - 20 cm Pb
  - Muon Veto
  - Soft Iron



From G. Chemin and J. Billard (Magnificent CE $\nu$ NS, March 2023)

# Reactor CE $\nu$ NS experiments



@Chooz power plant (France, Givet)



## Inner Shielding:

- Inner Veto [This talk](#)
- [Cryogenic Outer Veto](#)
- 4 cm B<sub>4</sub>C
- 20 cm Borated PE/Cu
- 5 cm Pb/Cu
- Cryogenic Muon Veto

## Outer Shielding

- 20 cm Borated PE
- 5 cm Pb
- Muon Veto

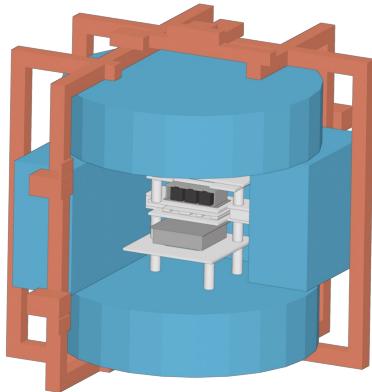
Shallow overburden  $\approx 3$  m.w.e.  
and small foot print required: a few m<sup>2</sup>

## (known) Background prediction

| Background contribution<br>Rates in kg <sup>-1</sup> d <sup>-1</sup> ( <i>Preliminary</i> ) | CaWO <sub>4</sub> array |                     |                |
|---|-------------------------|---------------------|----------------|
|   | 10 – 100 eV             | 100 eV – 1 keV      | 1 keV – 10 keV |
| Ambient gammas  | $0.5^{+0.9}_{-0.3}$     | $4.1^{+1.7}_{-1.4}$ | $92 \pm 7$     |
| Atmospheric muons   | $1.2^{+0.9}_{-0.8}$     | $2.7^{+1.3}_{-1.1}$ | $9.3 \pm 1.9$  |
| Atmospheric neutrons  | $\approx 9$             | $\approx 24$        | $\approx 90$   |
| <b>Total</b>  | $\approx 11$            | $\approx 30$        | $\approx 190$  |
| CE $\nu$ NS signal  | $\approx 30$            | $\approx 9$         | –              |

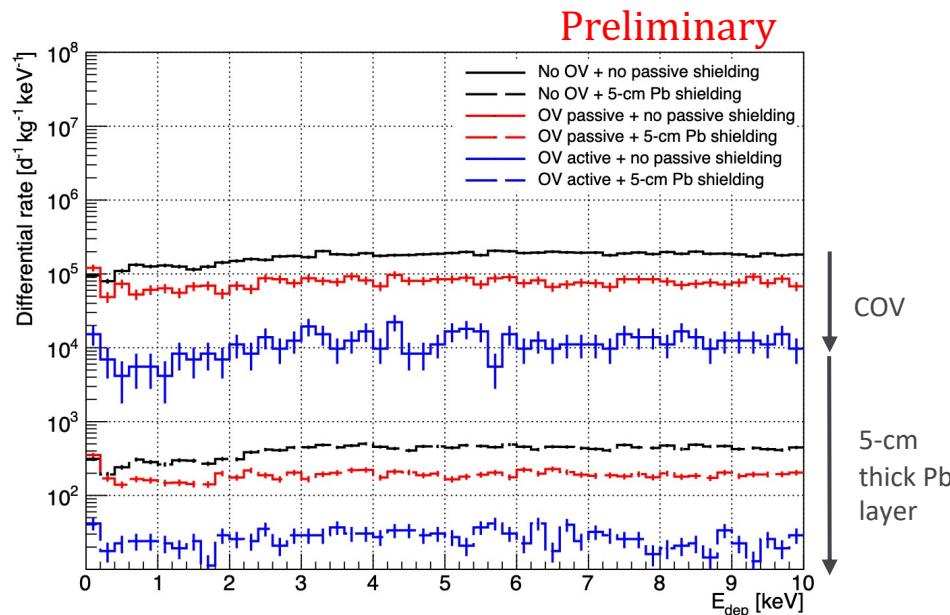
Estimated background rate in ROI (10 eV – 100 eV)  
 $\approx 120$  ev/kg/d/keV

# Development of a germanium cryogenic veto

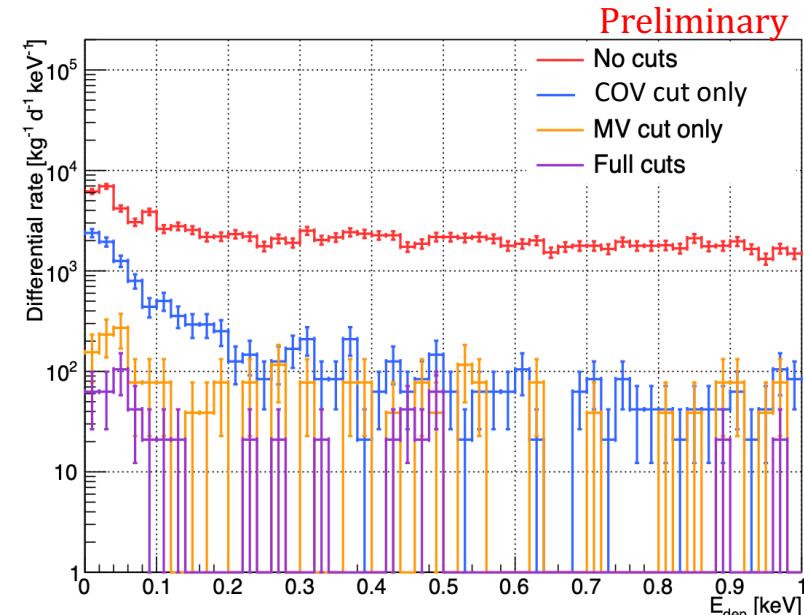


- 6 HPGe crystals
- $4\pi$ -coverage active veto
- Fast detector response
- Anti-coincidence with bolometric detectors
- $O(10\text{keV})$  threshold
- Compactness

⇒ Good gamma (and moderate neutron and muon) veto



Ambient gamma contribution to the residual background in  $\text{CaWO}_4$  crystals with different shielding configurations



Muon-induced contribution to the residual background in  $\text{CaWO}_4$  crystals with different active vetoes (simulation without the  $\text{B}_4\text{C}$  layer)

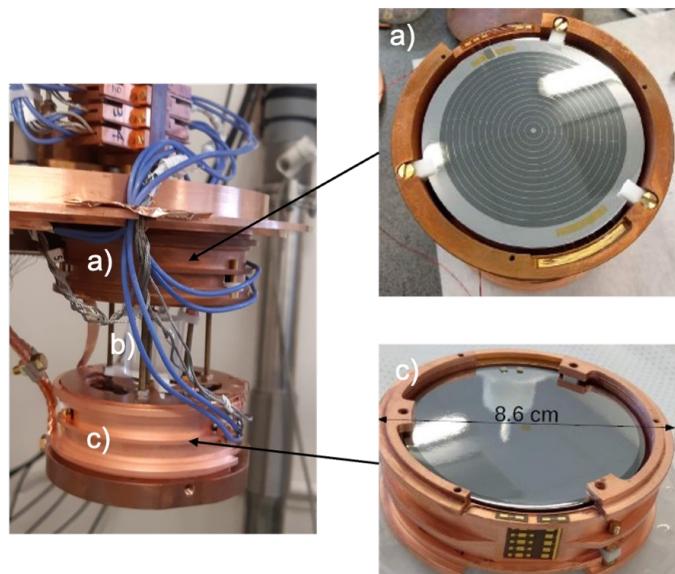
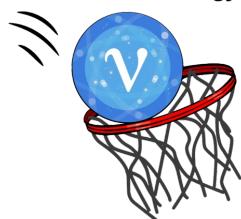
# Prototyping a germanium cryogenic veto for CE $\nu$ NS

## Experimental setup:

2 HPGe crystals (ionization channel) → a) and c)  
+ 1  $\text{Li}_2\text{WO}_4$  crystal in the center (heat channel (NTD),  
“target detector”) → b)



Bolometers At Sub KeV Energy Thresholds



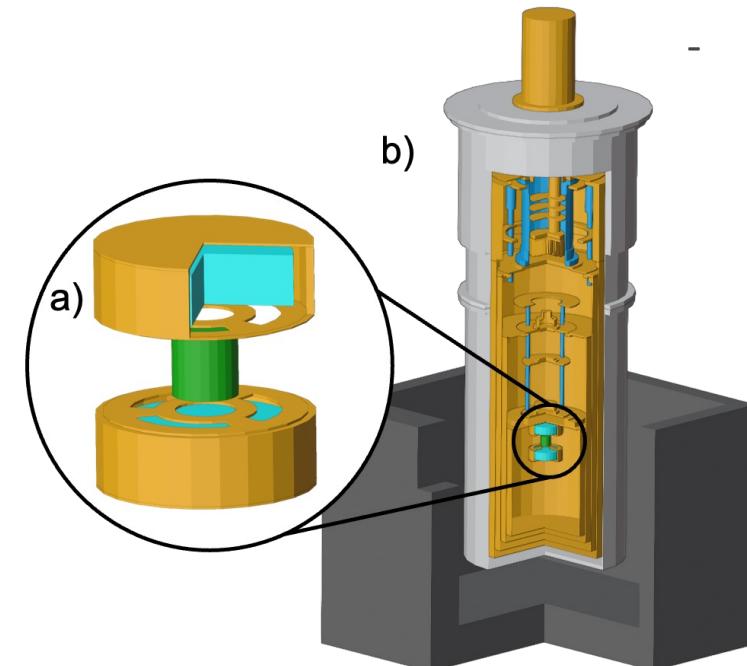
- Measurements with and without:
- 5-cm thick lead shielding,
  - Neutron source ( $^{252}\text{Cf}$ )
  - Gamma source ( $^{232}\text{Th}$ )



## Simulation of the setup

Data described by 3 components at surface:

- Environmental Gammas
- From cosmic rays:
  - Neutrons
  - Muons



# Prototyping a germanium cryogenic veto for CE $\nu$ NS

## Analysis method

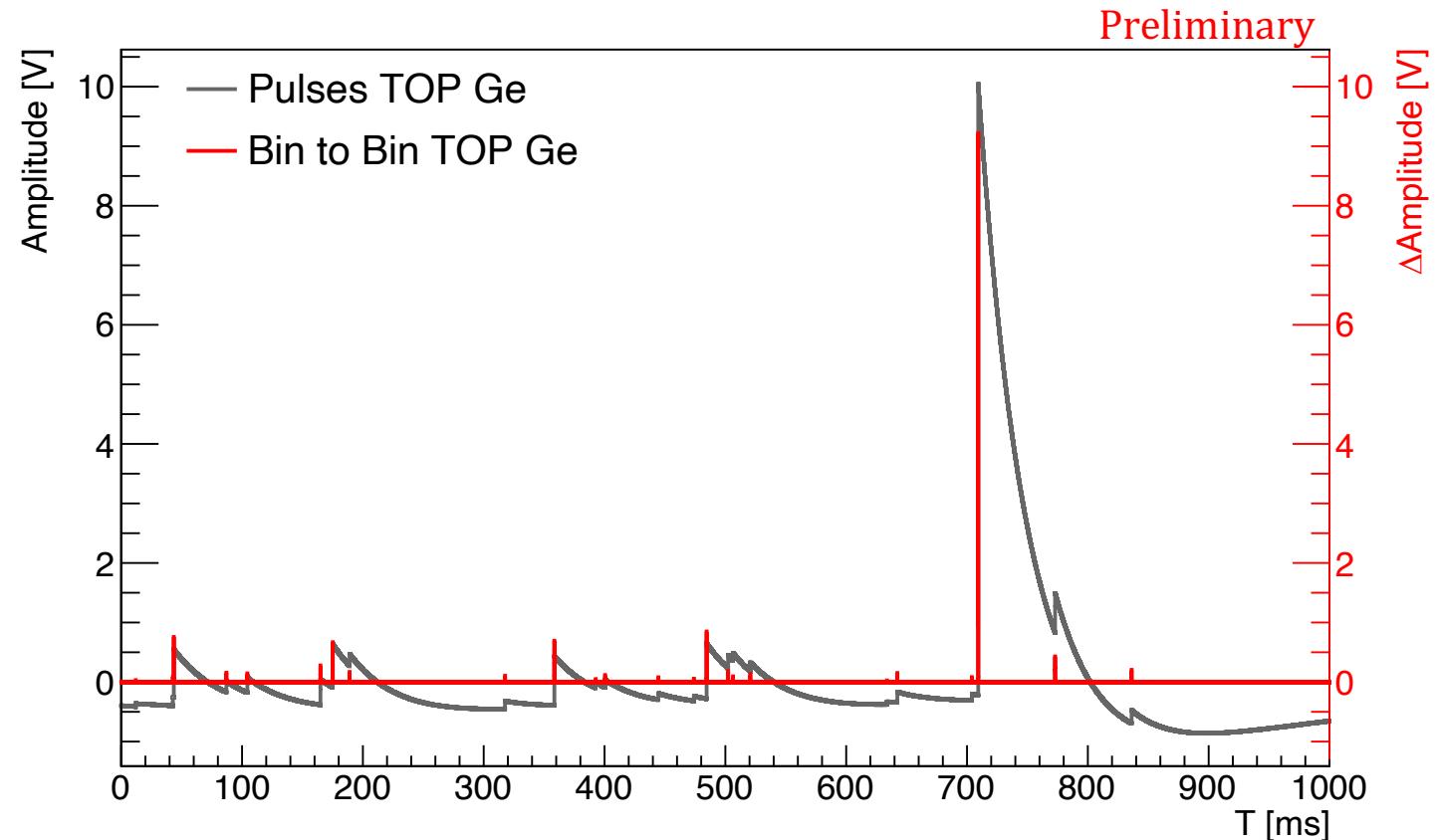
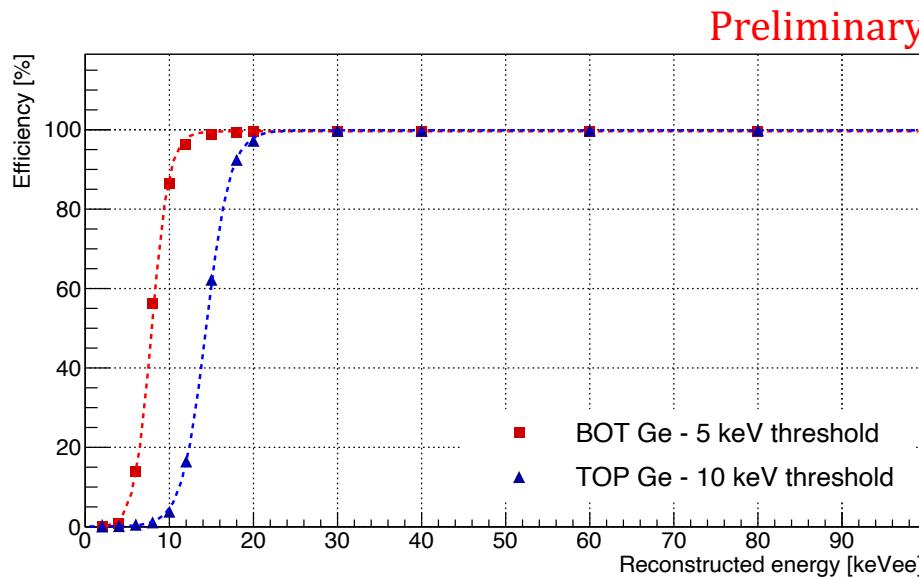
1- Top electrode – Bot electrode = Pulses

2- Bin-to-bin differentiation

- identify pulses rise time
- + flatten the baseline
- + deal with pile-up

3- Apply threshold

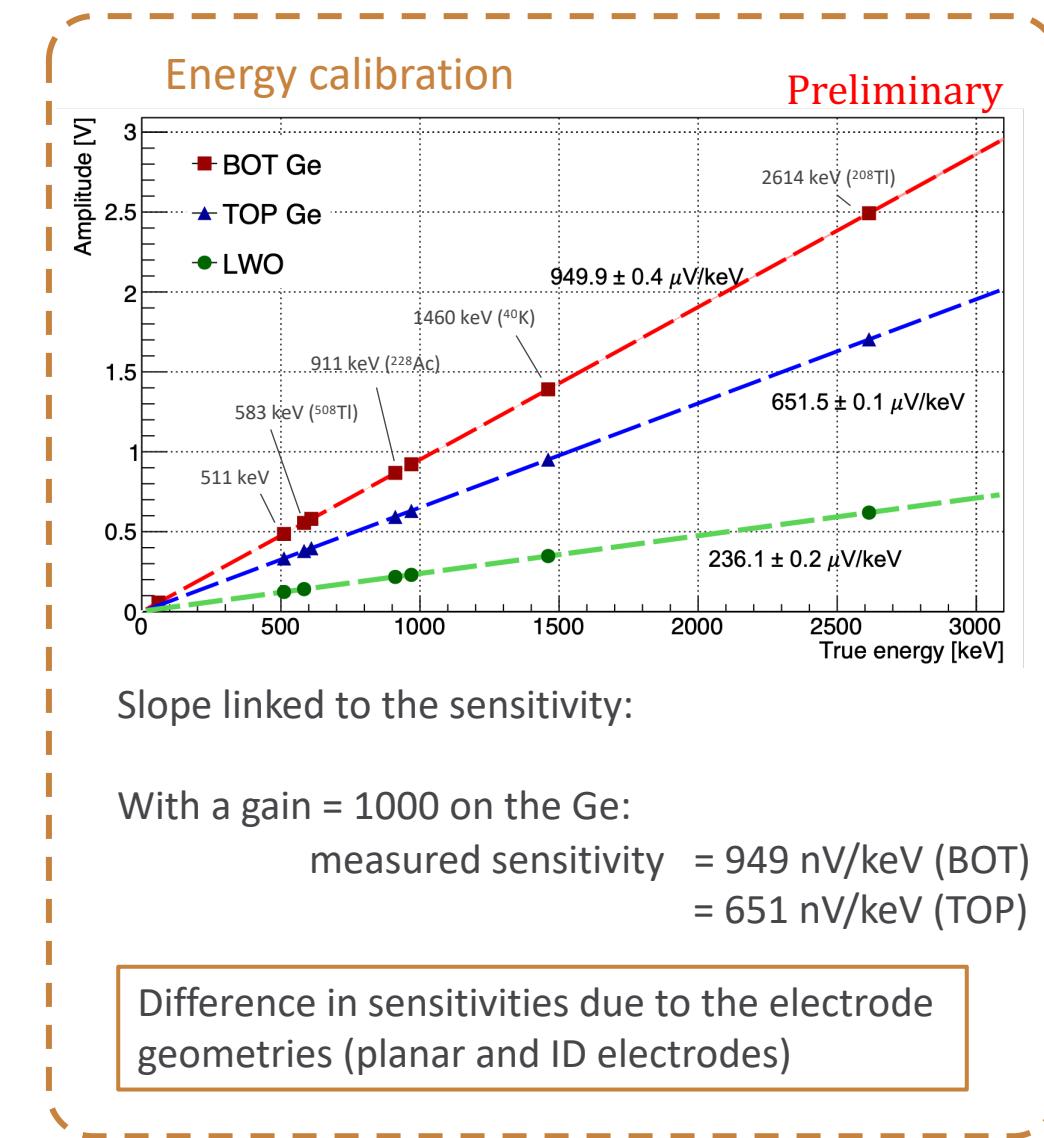
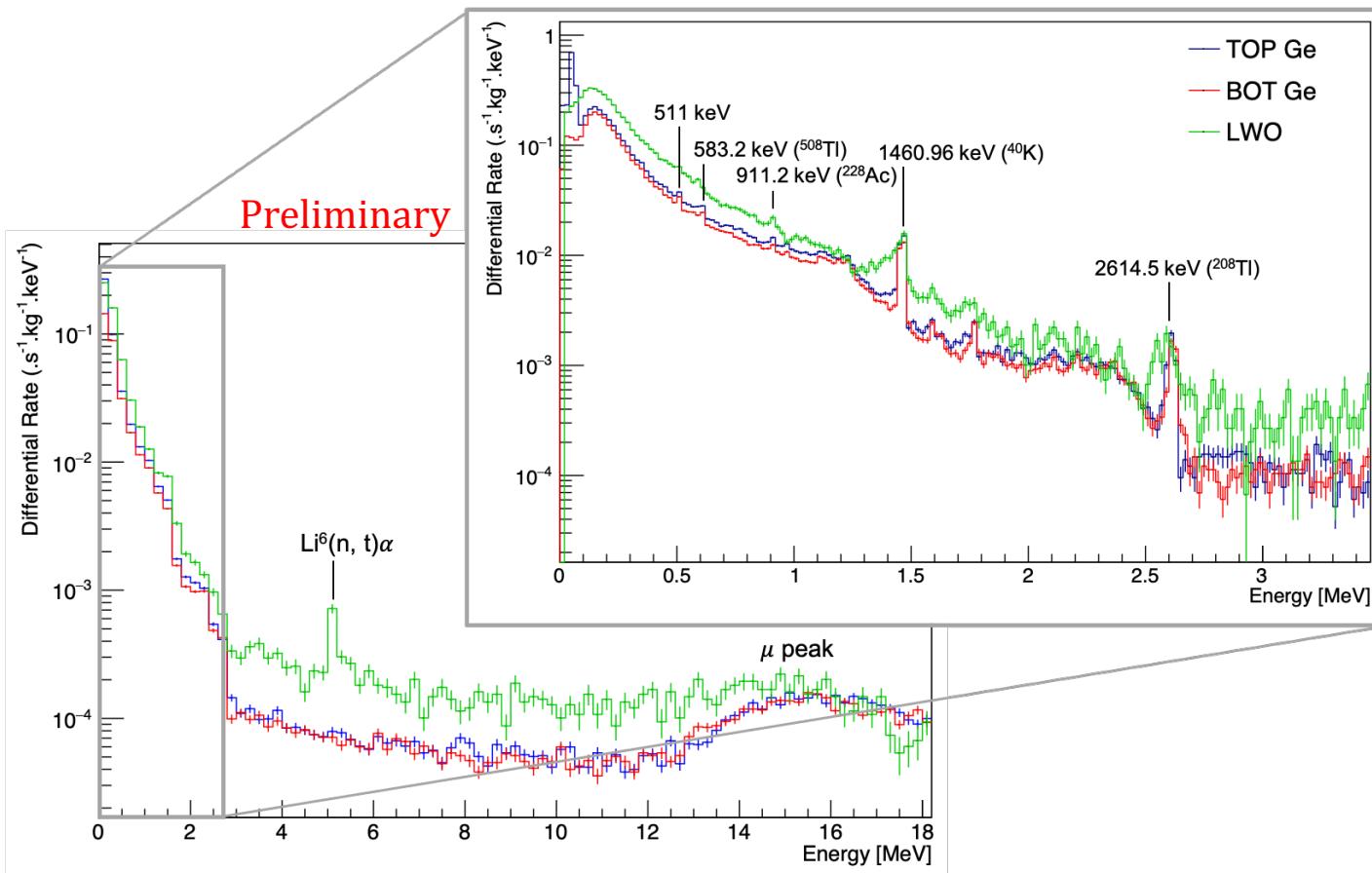
4- Calculate pulse amplitude



⇒ High reconstruction efficiency close to the threshold, handle the pile-up

# Prototyping a germanium cryogenic veto for CE $\nu$ NS

## Detector response



# Prototyping a germanium cryogenic veto for CE $\nu$ NS

Fit of the signal component fluxes on the data :

- Good understanding of our data and the signal component
- Simple analysis of the germanium detectors
- Correct simulation of the expected events

Preliminary

| Background contribution | Fluxes (/cm <sup>2</sup> /s)     |   |
|-------------------------|----------------------------------|---|
|                         | This work                        | Reference values                          |
| Atmospheric muons       | $(1.79 \pm 0.02) \times 10^{-2}$ | $(1.90 \pm 0.12) \times 10^{-2}$ from [1] |
| Environmental gammas    | $3.126 \pm 0.005$                | $3.2 \pm 0.3^*$                           |
| Atmospheric neutrons    | $(1.37 \pm 0.27) \times 10^{-2}$ | $1.34 \times 10^{-2}$ from [2]            |

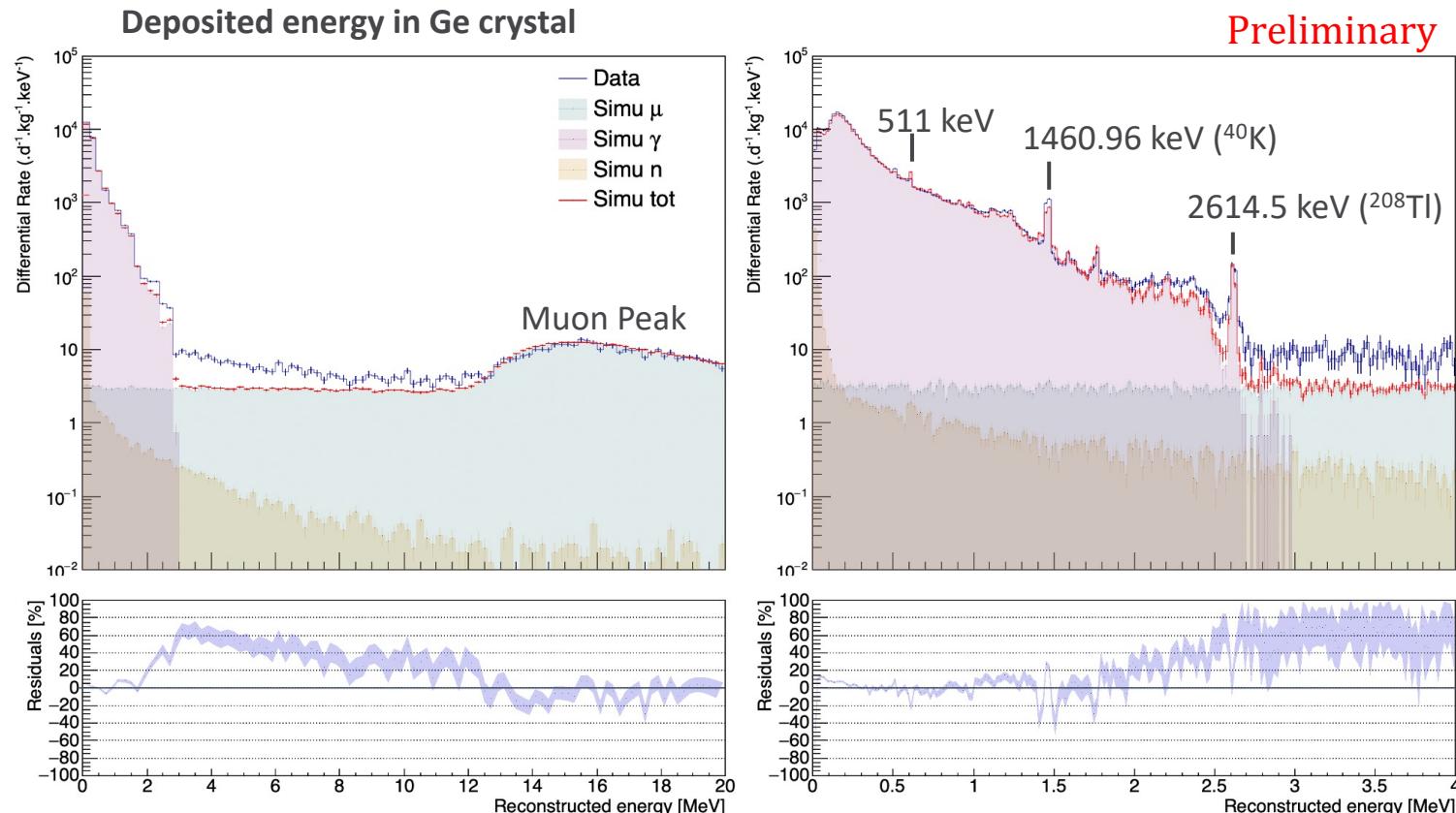
Table 1: Fitted values for the integrated fluxes of each background contribution.

The errors given are fit errors calculated only from statistic errors.

\* Measured value in the lab with a high purity germanium spectrometer.

[1] Tang, et al. Physical Review, 2006

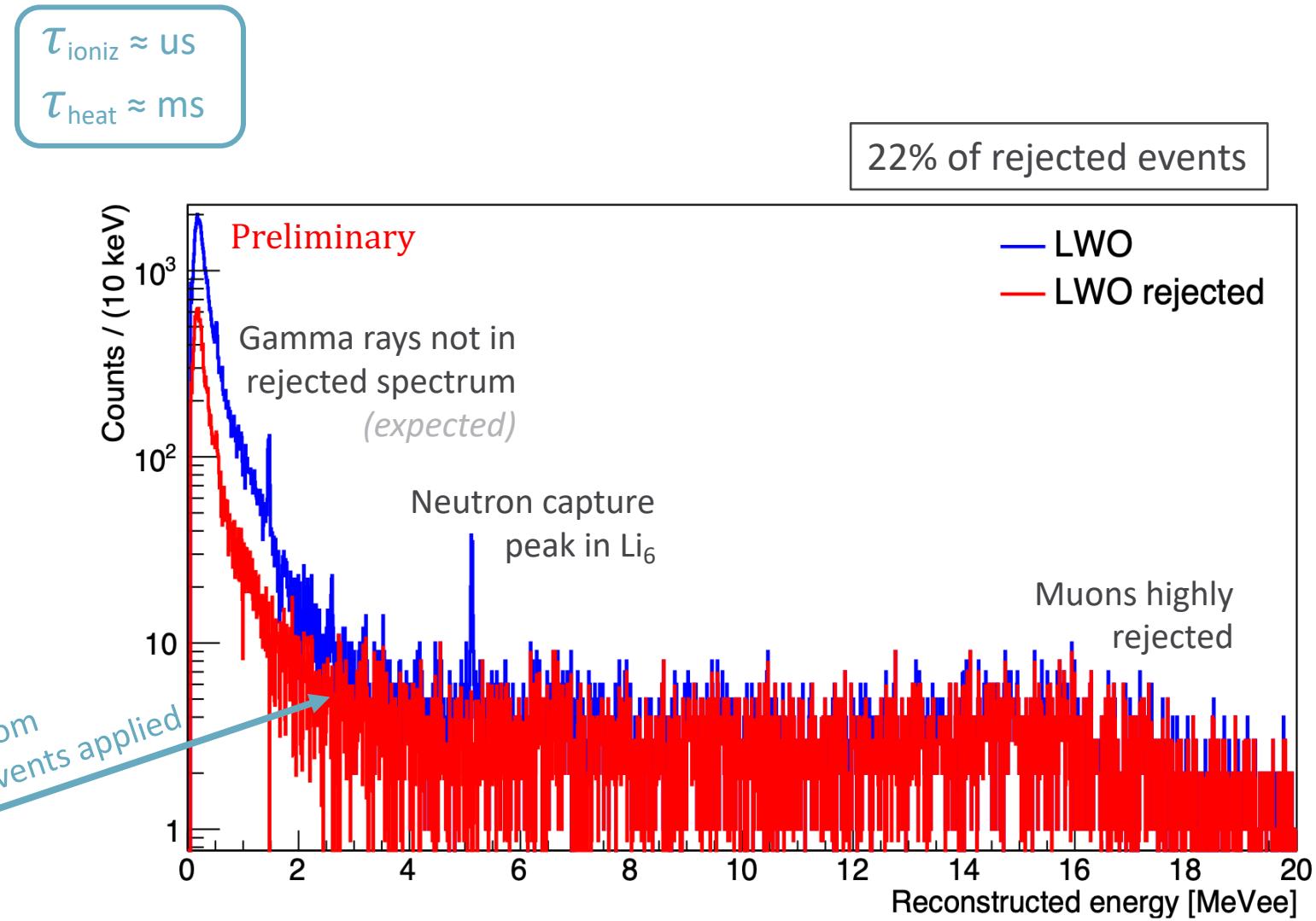
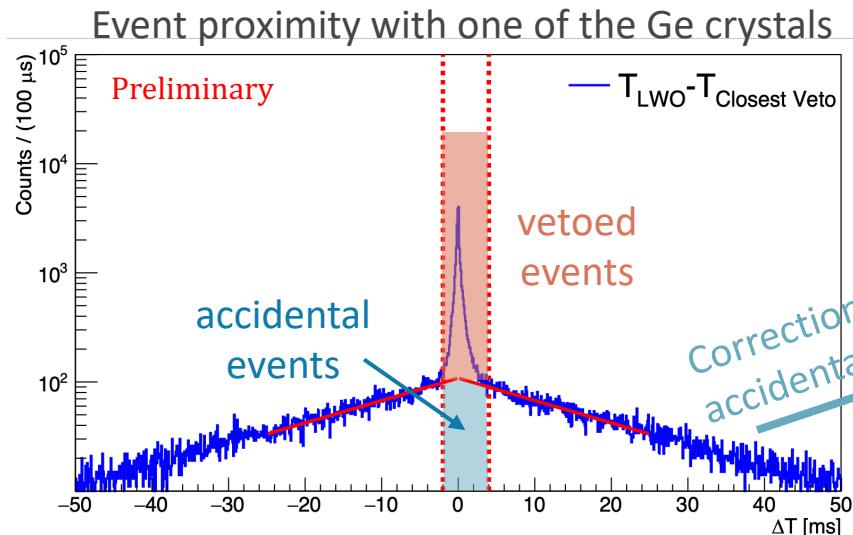
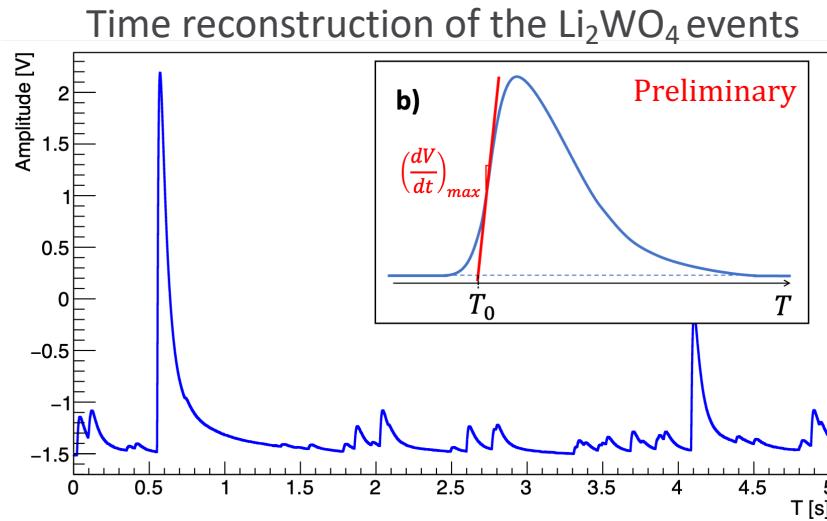
[2] Gordon, et al Nuclear Science, 2005



Article in preparation

# Prototyping a germanium cryogenic veto for CE $\nu$ NS

## Coincidences and veto in the data



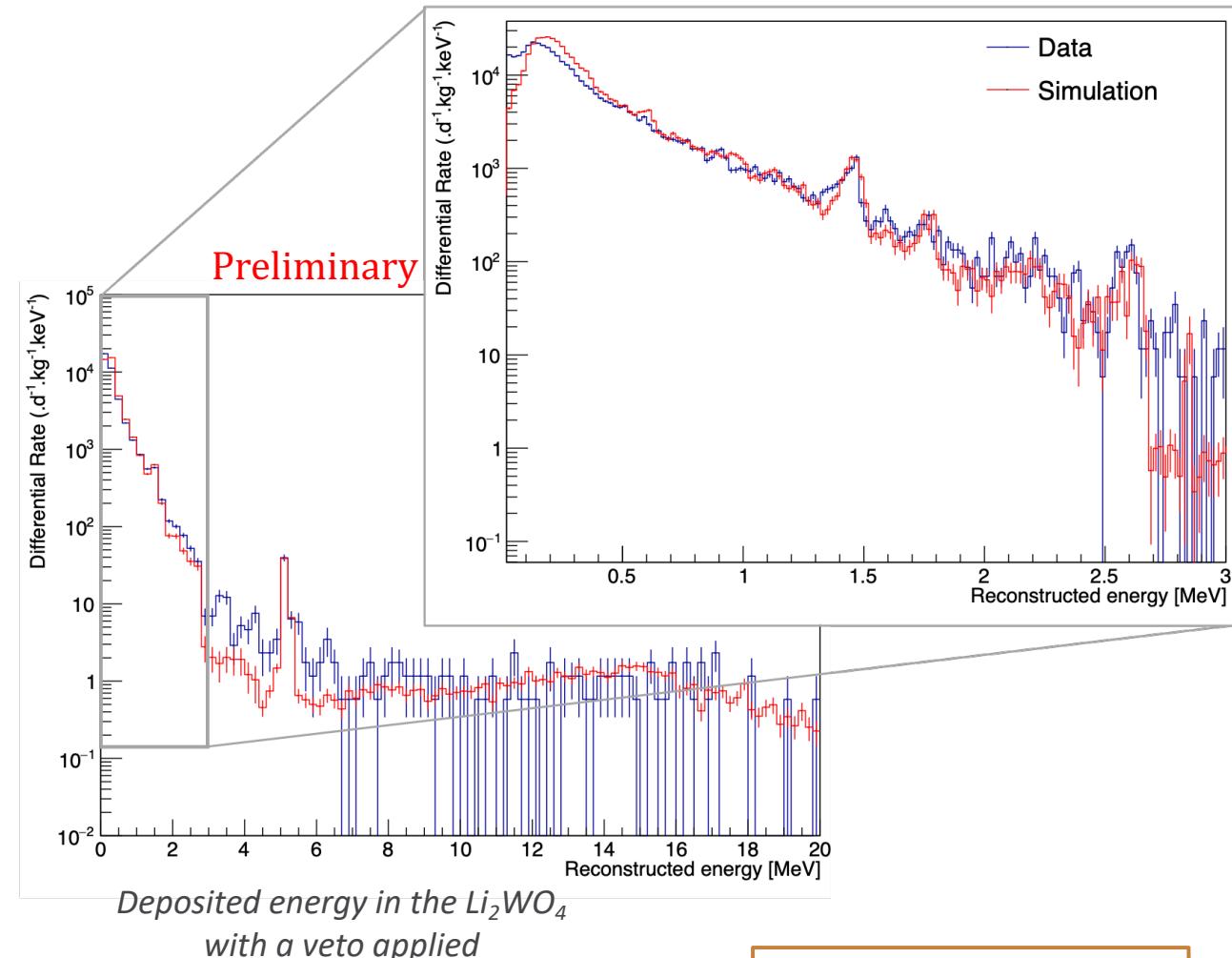
# Prototyping a germanium cryogenic veto for CE $\nu$ NS

## Comparison with the simulation

- Really good agreement between veto rejection obtained in the data and in the simulation

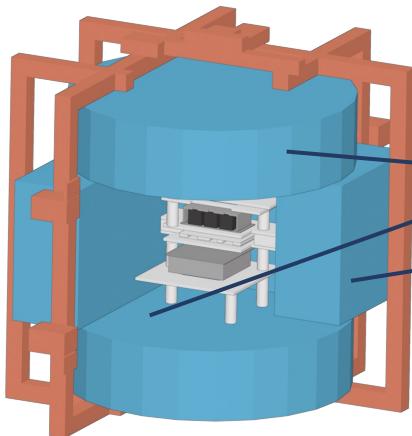
## Rejection power by the Ge veto in the Li<sub>2</sub>WO<sub>4</sub>

| Preliminary    | No shielding |              | With shielding |              |            |
|----------------|--------------|--------------|----------------|--------------|------------|
|                | Energy Range | Data         | Simulation     | Data         | Simulation |
| [0.05; 20] MeV | 22.0 ± 0.5 % | 22.4 ± 0.4 % | 30.8 ± 0.5 %   | 30.1 ± 0.7 % |            |
| [0.05; 3] MeV  | 20.4 ± 0.5 % | 22.0 ± 0.4 % | 26.3 ± 0.5 %   | 28.8 ± 0.8 % |            |
| [3; 10] MeV    | 93.1 ± 3.4 % | 84.2 ± 0.9 % | 93.2 ± 2.0 %   | 84.4 ± 2.0 % |            |
| [10; 20] MeV   | 78.4 ± 3.3 % | 81.7 ± 1.3 % | 80.2 ± 2.1 %   | 81.3 ± 2.2 % |            |



Article in preparation

# Towards a $4\pi$ veto for NUCLEUS

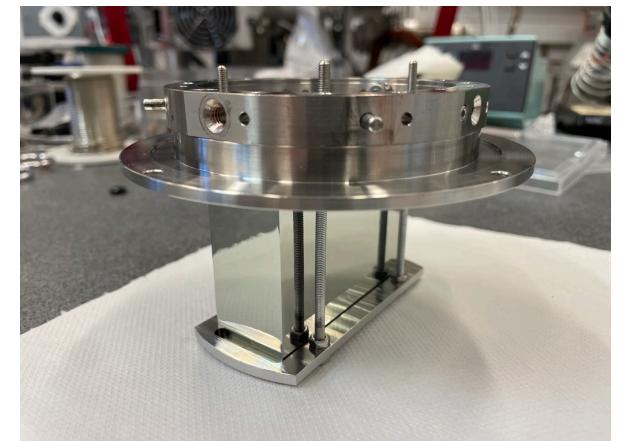


With a  $4\pi$ , 2cm-thick system: rejection efficiency estimated at 95%  
⇒ Development of this veto for NUCLEUS

Cylindric crystal (10-cm diameter, 2.5-cm thick, 1kg)  
Rectangular crystals (7 x 2.5 x 5 cm, 500g)

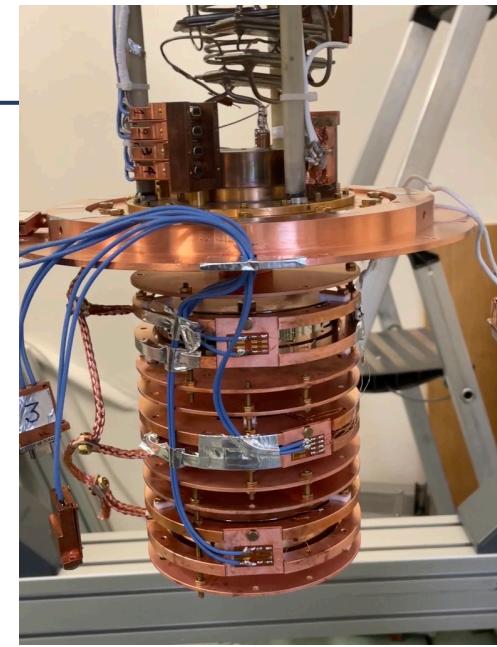
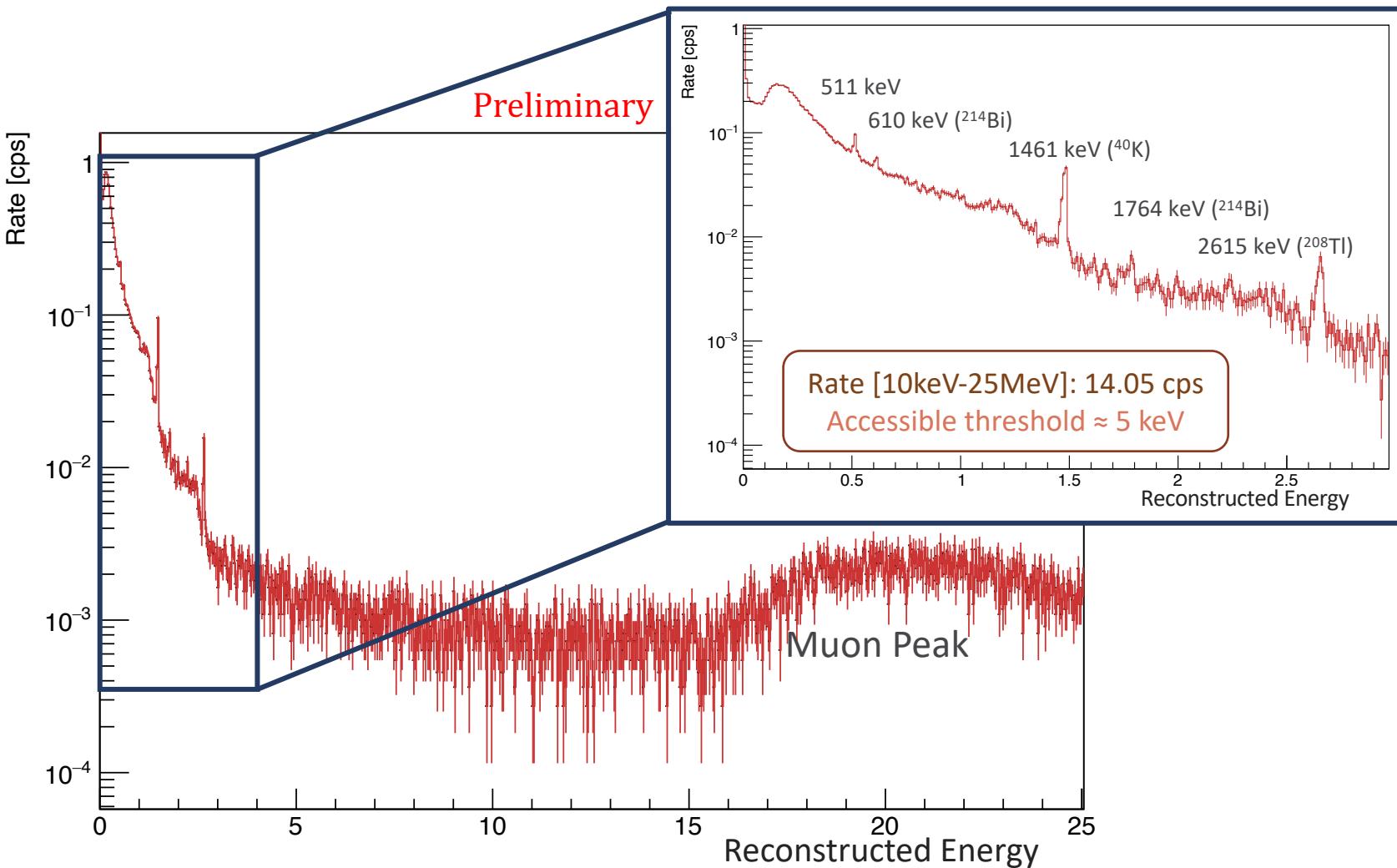


Electrode evaporation:  
→ 30 nm of a-Ge:H  
→ 200 nm Al electrode

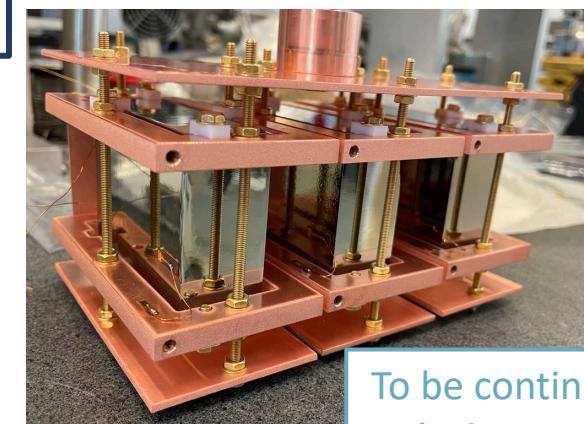


# Towards a $4\pi$ veto for NUCLEUS

## Test of the cylindric crystals @IJCLab



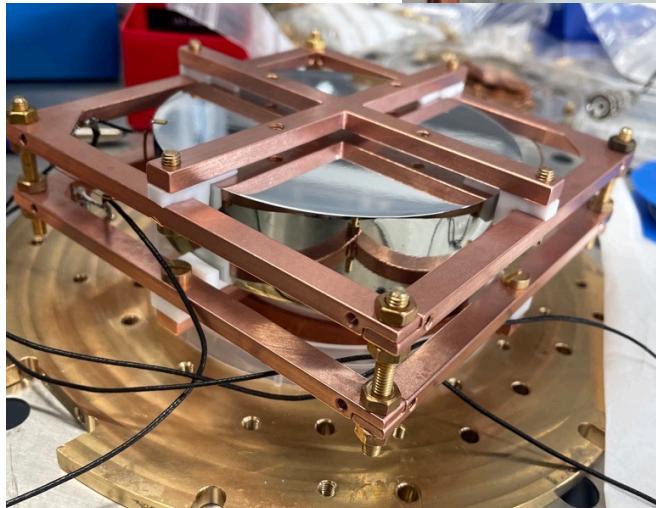
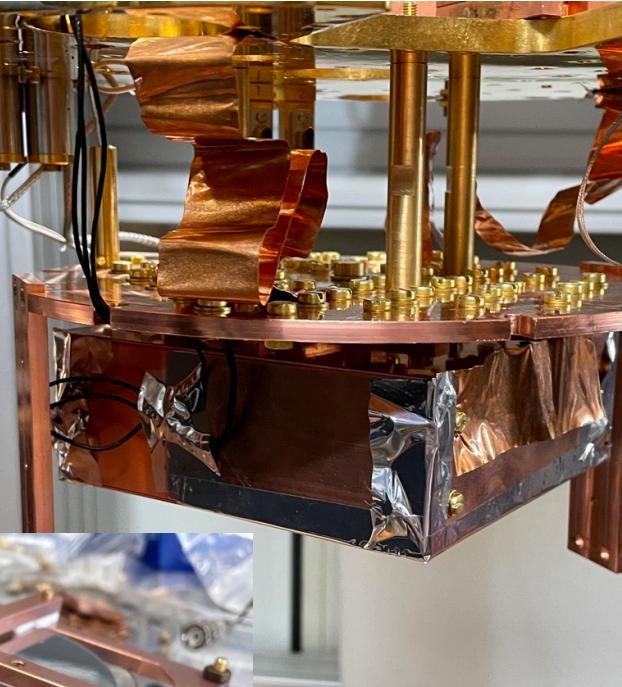
Test of the rectangular crystals...



To be continued at IJCLab  
in Actuator Cryostat

# Towards a $4\pi$ veto for NUCLEUS

In the NUCLEUS cryostat for the commissioning:



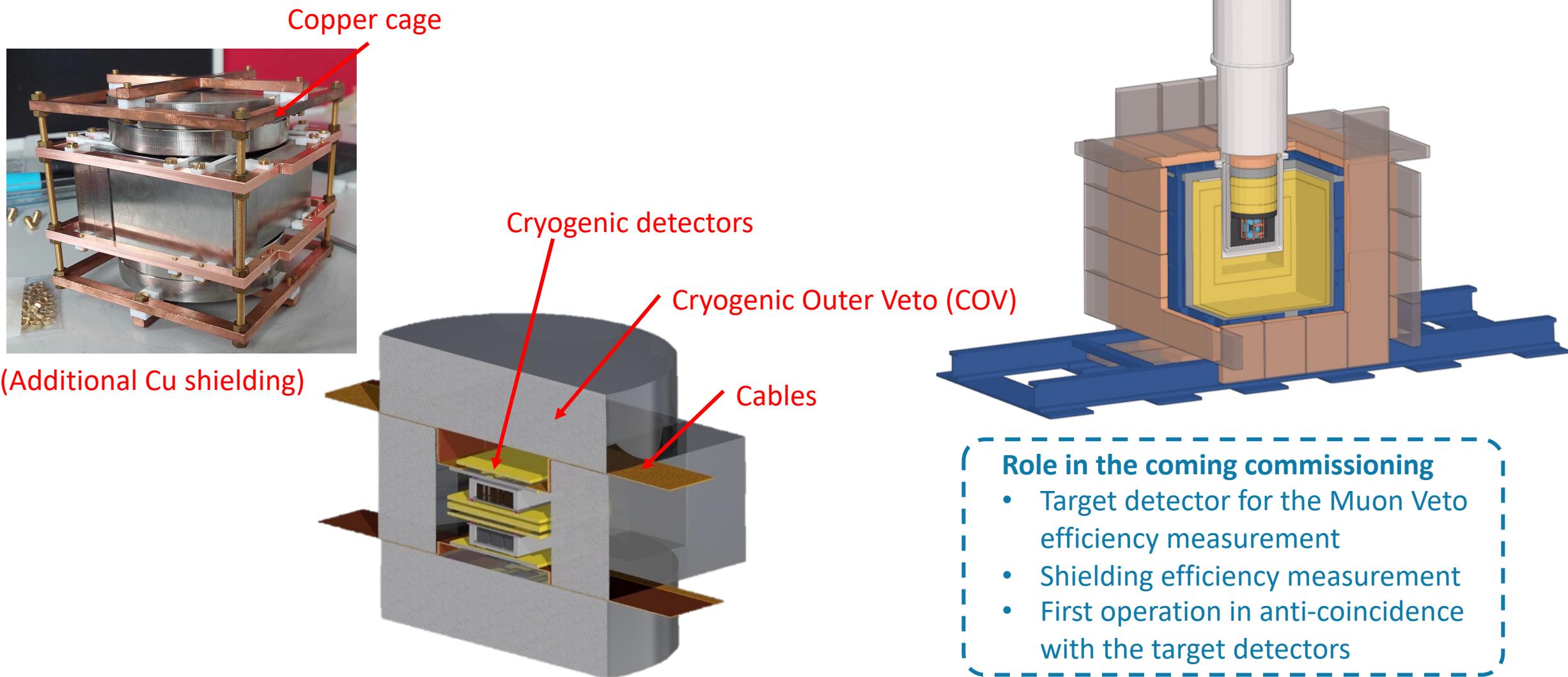
- First light in March 2023
- Base temperature (7mK) achieved

## Next steps:

- Long time measurement
- Background spectrum
- Final acquisition scheme
- Run in anti-coincidence with the NUCLEUS target detectors

# Towards a $4\pi$ veto for NUCLEUS

## Scaling up and role in the future commissioning plan



### Role in the coming commissioning

- Target detector for the Muon Veto efficiency measurement
- Shielding efficiency measurement
- First operation in anti-coincidence with the target detectors

# From blank assembly towards on-site installation

May 2023

Beginning 2024

## Blank Assembly & commissioning

Underground Lab at TUM



- Mechanical integration tests
- Calibrations at keV energies and below:
  - LED
  - XRF
  - Neutrons with CRAB  
(JINST 16 P07032 (2021))
- Detector performances
- Background studies at sub-keV (EXCESS)



Full COV installation

- Background measurement in the UGL
- Shielding efficiency characterization

2024

## NUCLEUS-10g physics run

Phase 1: observe CE $\nu$ NS

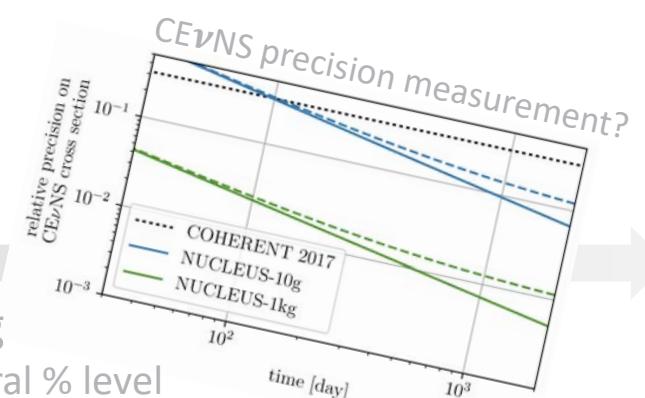
## On-site installation

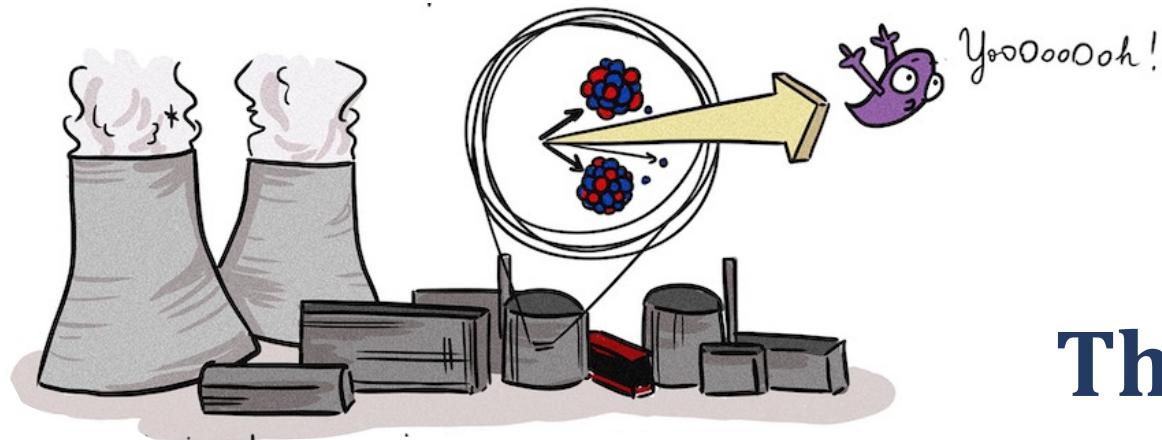
Chooz



## Towards NUCLEUS-1kg

Phase 2: measure CE $\nu$ NS at the several % level





# Thanks for your attention





# **Development of a cryogenic veto system for CEvNS detection**

in the scope of the NUCLEUS experiment

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Back-up slides

# CE $\nu$ NS, what for?

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## Differential cross section

$$\frac{d\sigma}{dE_r} = \frac{G_F^2}{4\pi} Q_w^2 F^2(q^2) m(Z, N) \left(1 - \frac{E_r}{E_{r,max}}\right)$$

$G_F$ : Fermi constant

$Q_w = N - Z(1 - 4\sin^2\theta_W) \sim N$ : Nuclear weak charge

$F$ : Nuclear form factor, depends on  $q^2$

$q$ : Momentum transfer

$m(Z, N)$ : Total mass of the nucleus

$E_r$ : nuclear recoil energy

$E_{r,max} = 2E_\nu^2/(m(Z, N) + 2E_\nu)$ : maximal recoil energy

## Non-standard interactions

$$\sigma \sim [Z(g_V^p + 2\varepsilon_{\alpha\alpha}^{uV} + \varepsilon_{\alpha\alpha}^{dV}) + N(g_V^n + \varepsilon_{\alpha\alpha}^{uV} + 2\varepsilon_{\alpha\alpha}^{dV})]^2$$

$$g_V^p = +\frac{1}{2} - 2\sin^2\theta_W$$

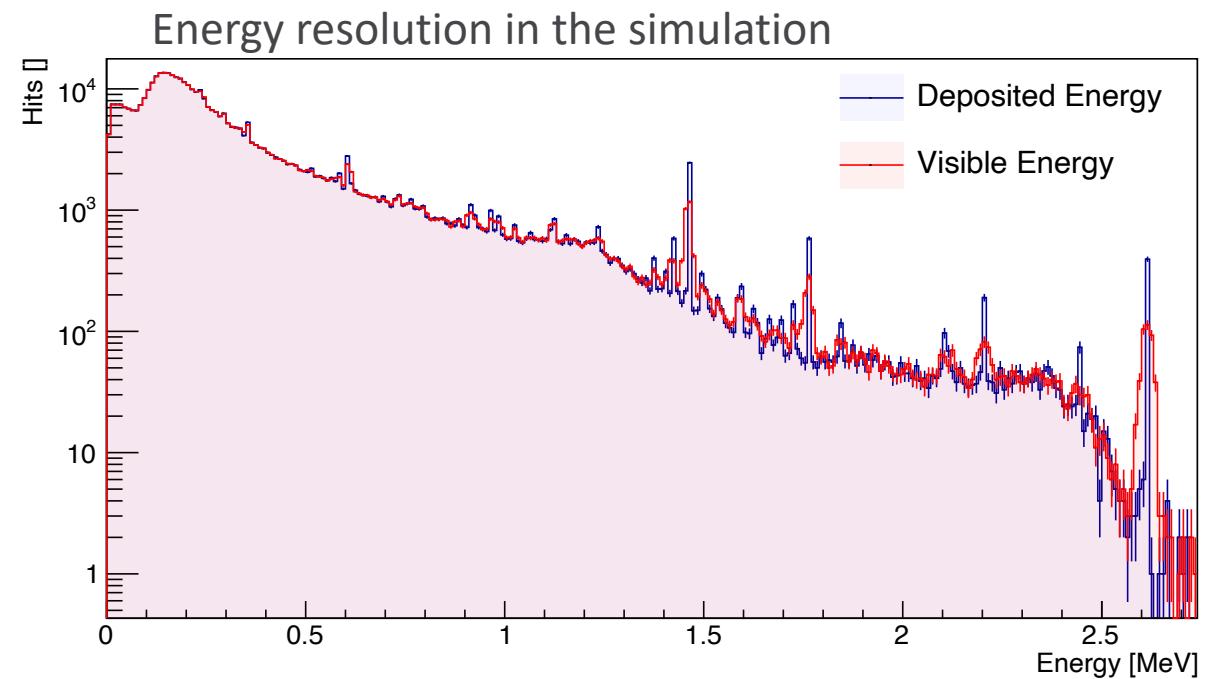
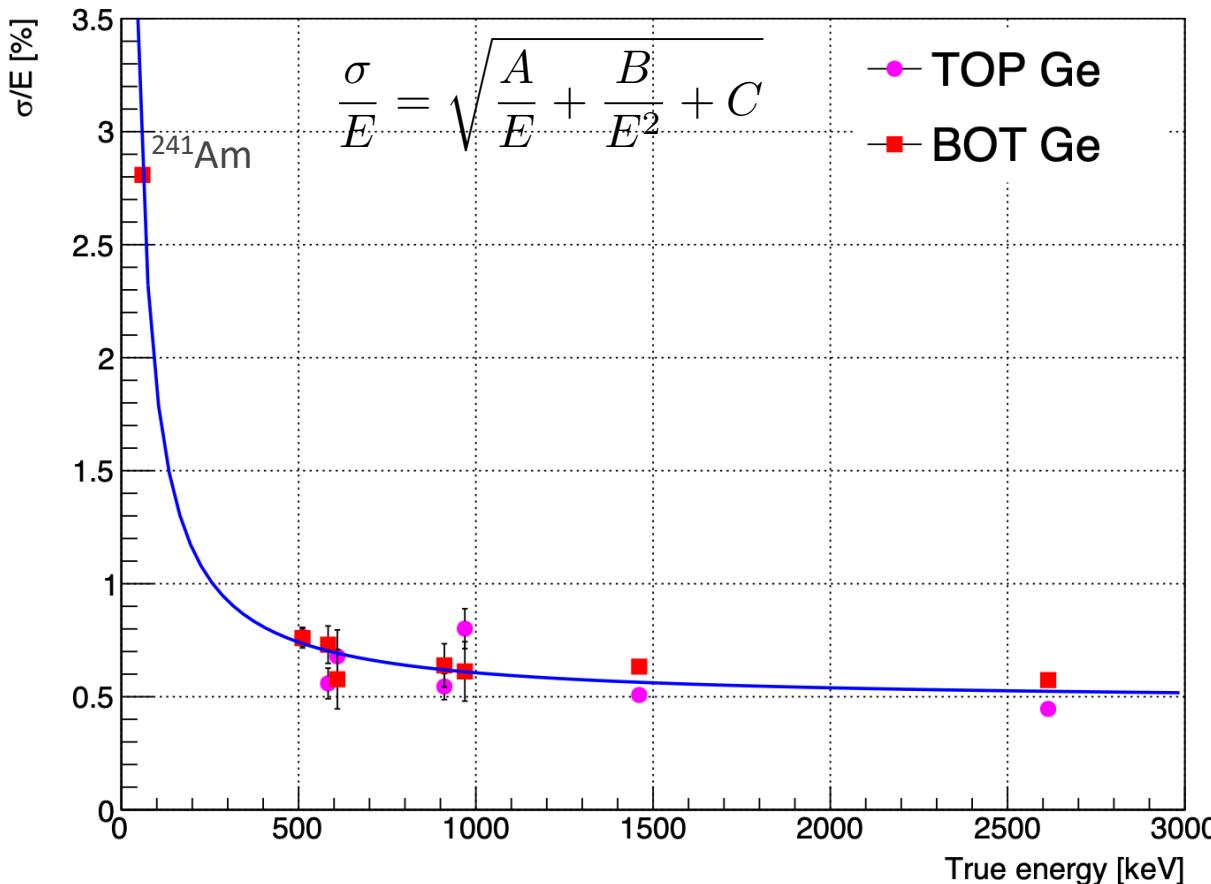
# Prototyping a germanium veto

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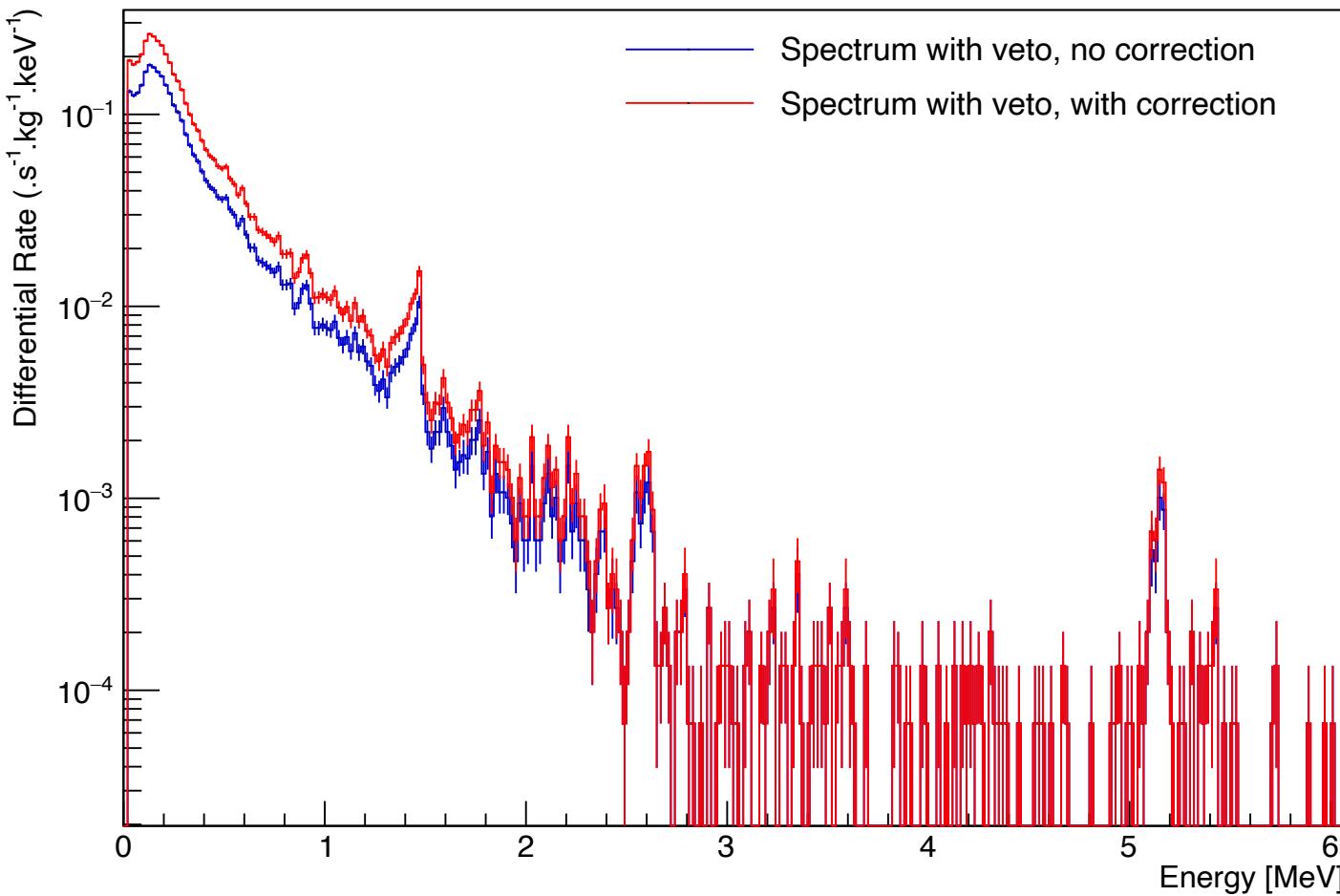
Analysis details

# Energy resolution

## Energy resolution



# Accidental events in veto



Higher rejection if no accidental correction is applied  $\Rightarrow$  Not the true coincidence rate.

- correction important in the gamma range
- less important at higher energy ( $> 5 \text{ MeV}$ )

# COV prototype simulation

---

Multi-steps fitting procedure

# Multi-steps fitting procedure

## Principle

$$\text{Minimize } \chi^2 = \sum_{det.} \frac{(D_{det} - M_{det})^2}{\sigma_{D_{det}}^2 + \sigma_{M_{det}}^2}$$

Sum over three detectors  
Stat. errors only

Data = reconstructed energy spectra

Model = scaled simulations

$$M_{det} = \phi_\mu \times S_{\mu,det} + \phi_n \times S_{n,det} + \phi_\gamma \times S_{\gamma,det}$$

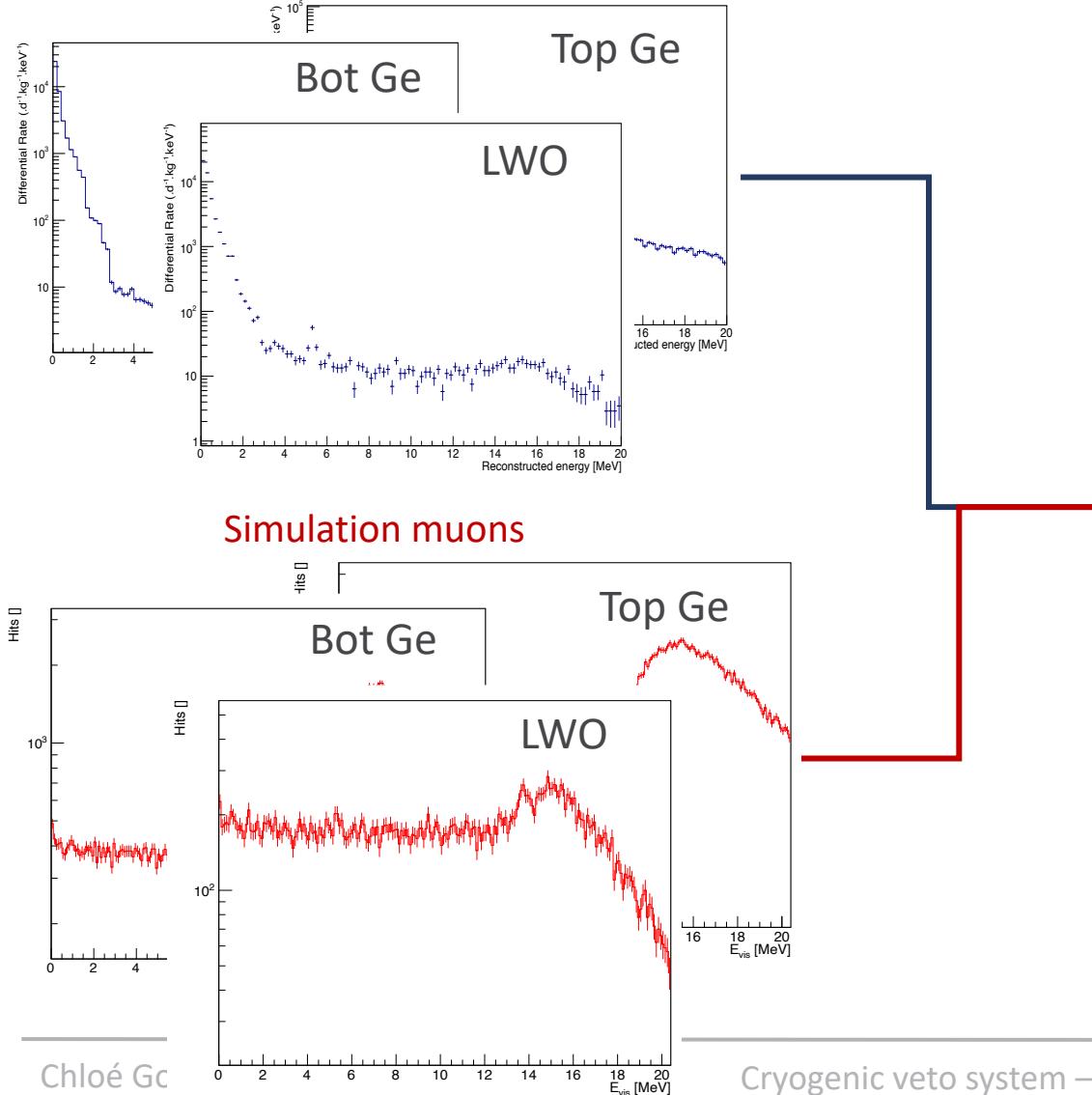
Fitted parameters

From MC simulations (fixed)

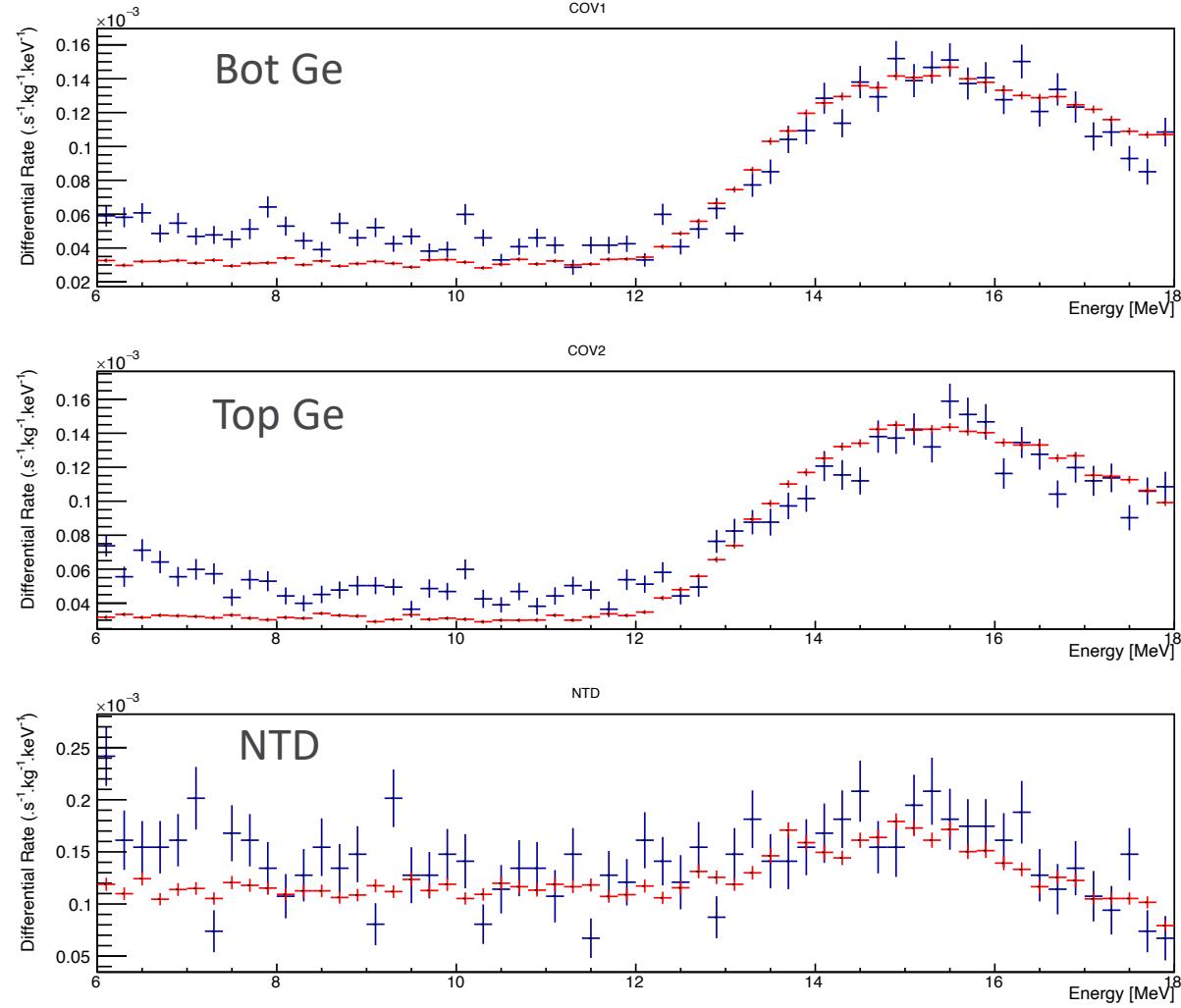
# Multi-steps fitting procedure

## Step 1 : Muons

Calibrated and time normalized data



## Fit Muon – 6-18MeV

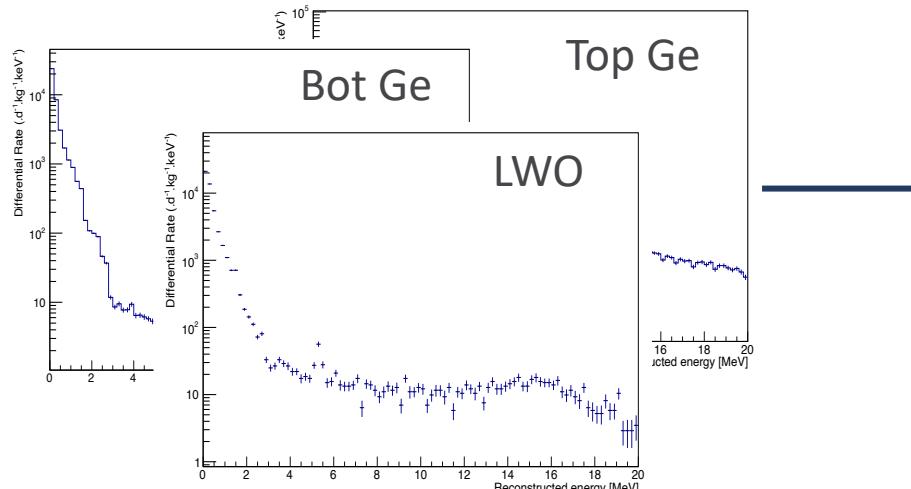


$$\phi_\mu$$

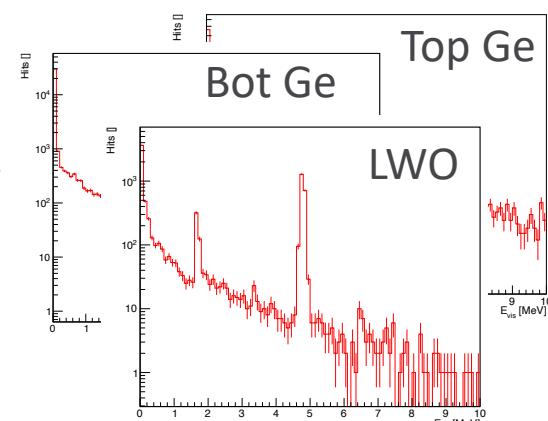
# Multi-steps fitting procedure

## Step 2 : Neutron from Li capture peak

Calibrated and time normalized data



Simulation neutron (shifted)

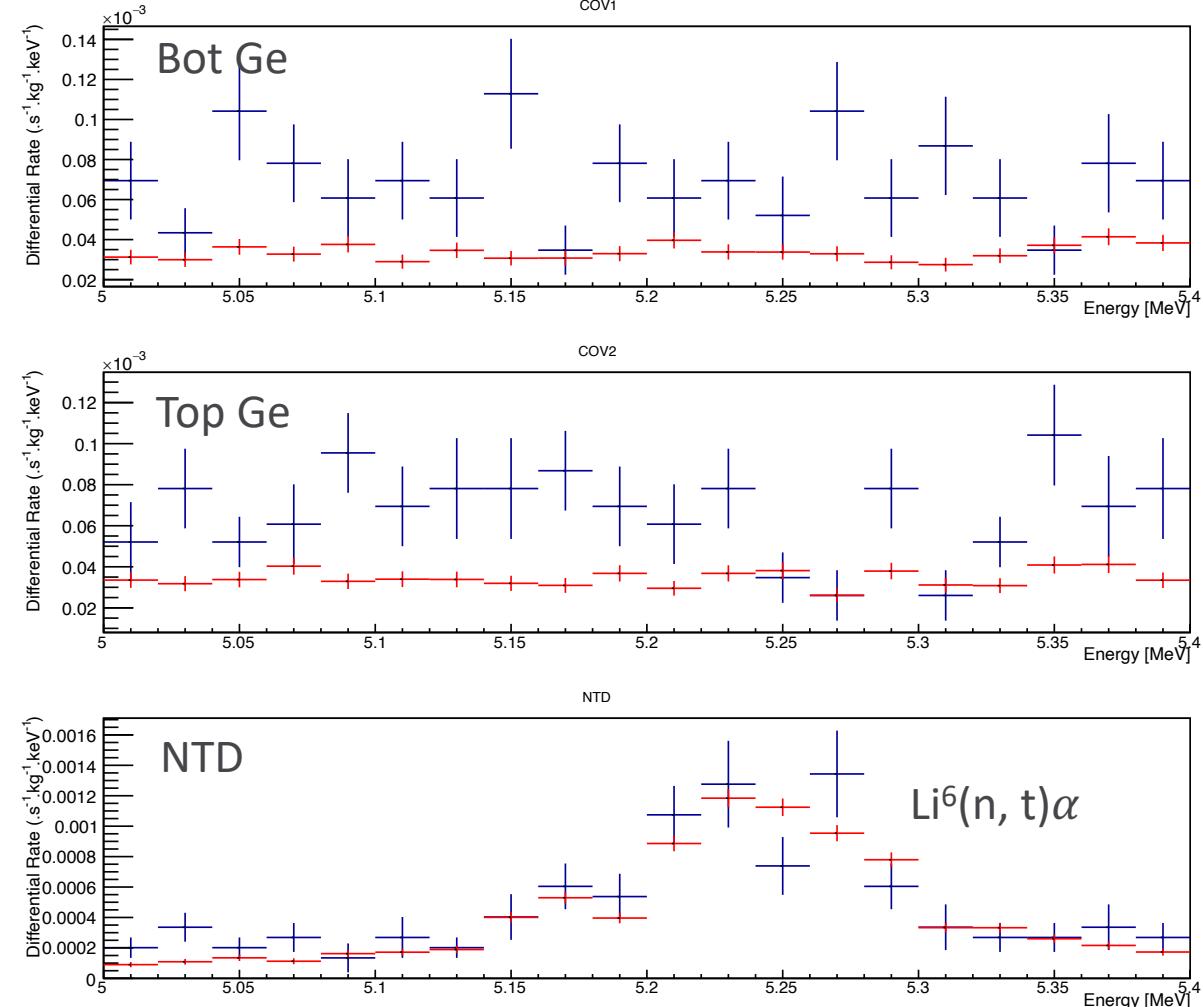


$\phi_\mu$

Simulation muons scaled with  $\phi_\mu$

$\phi_n$

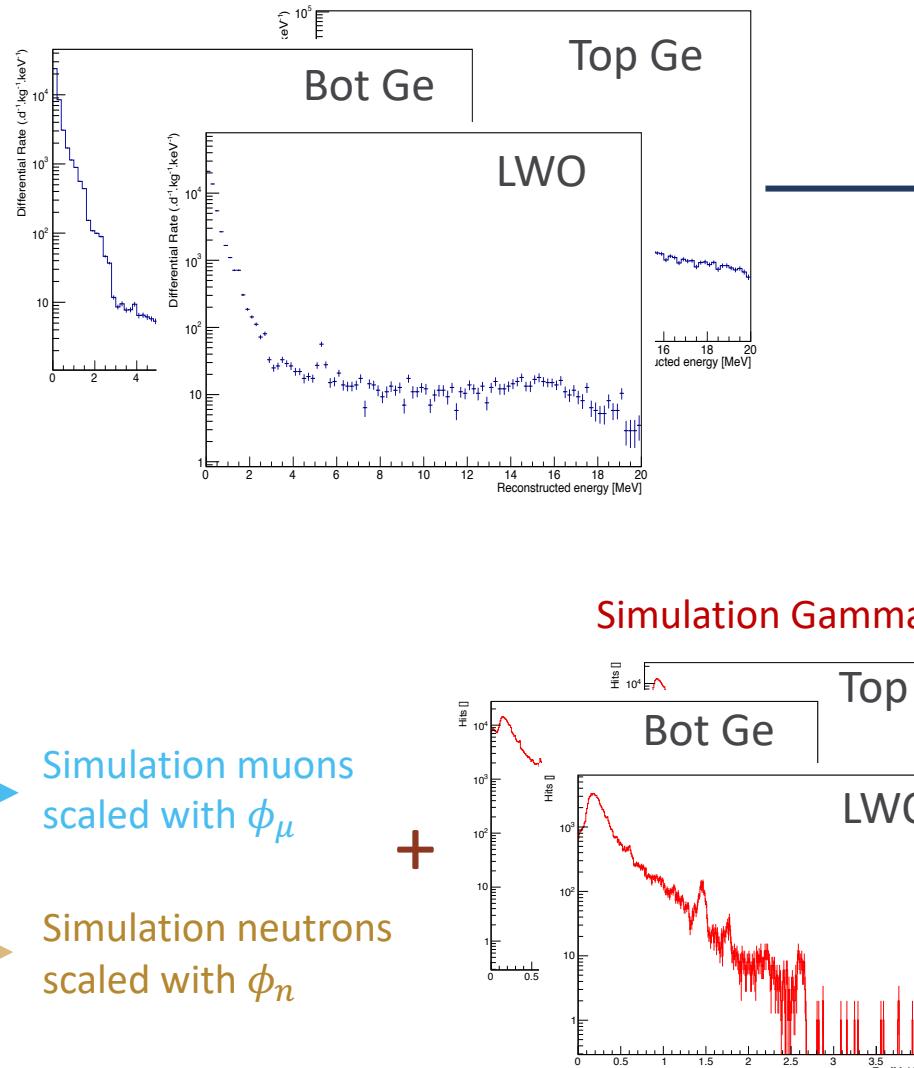
Fit Neutrons – 5-5.4MeV



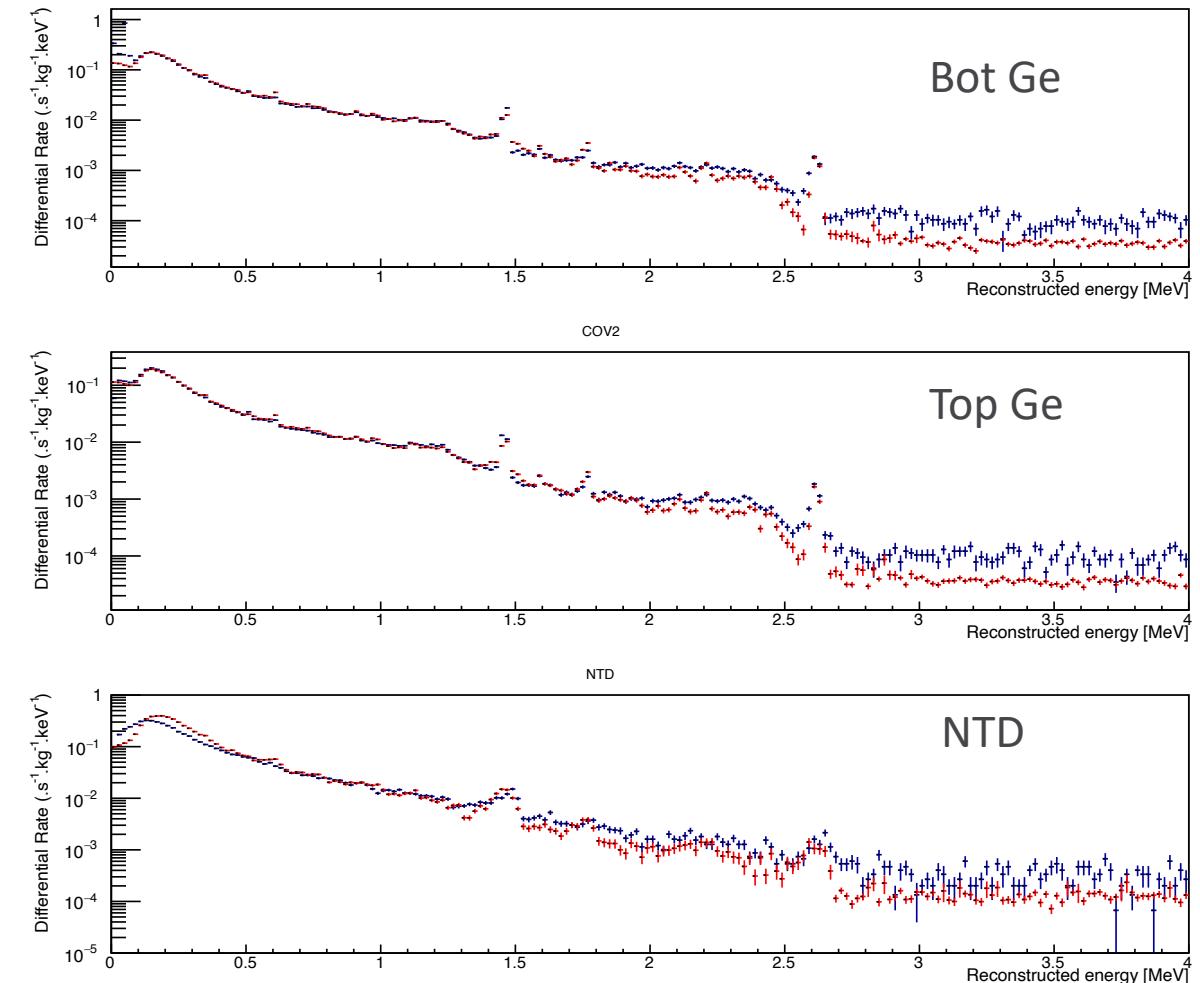
# Multi-steps fitting procedure

## Step 3 : Gammas

Calibrated and time normalized data



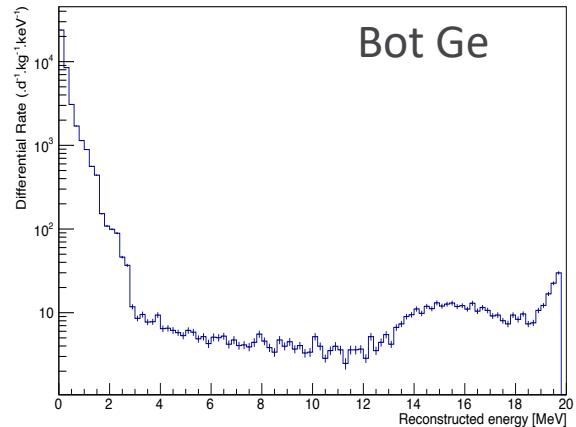
## Fit Gammas – 0.1-4MeV



# Multi-steps fitting procedure

## Step 4 : Am241

Calibrated and time normalized data

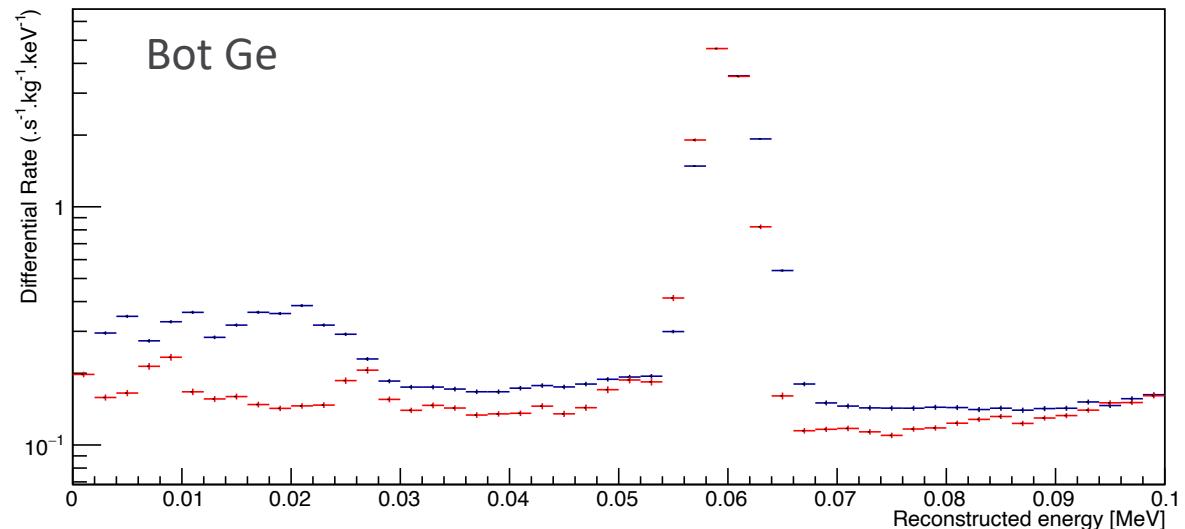


Bot Ge

Fit Am241 – 0-0.1MeV

COV1

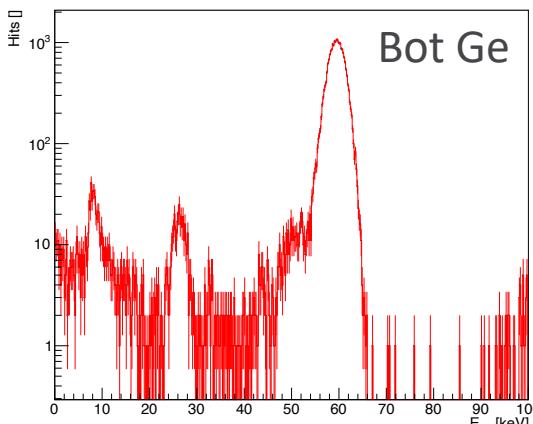
$\phi_{Am}$   
↑



Bot Ge

Simulation Am241

Bot Ge



$\phi_\mu$

Simulation muons  
scaled with  $\phi_\mu$

$\phi_n$

Simulation neutrons  
scaled with  $\phi_n$

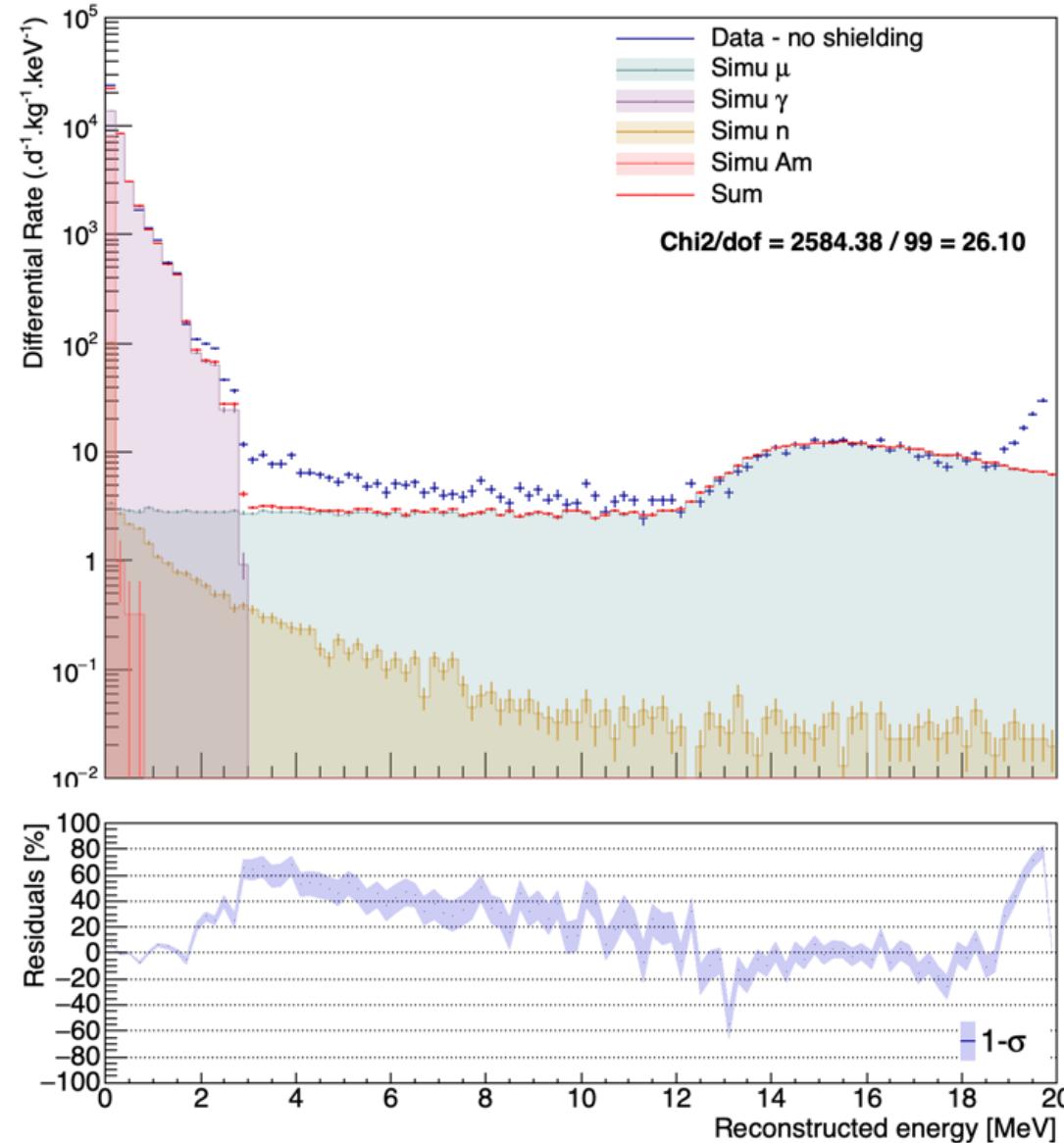
$\phi_\gamma$

Simulation gammas  
scaled with  $\phi_\gamma$

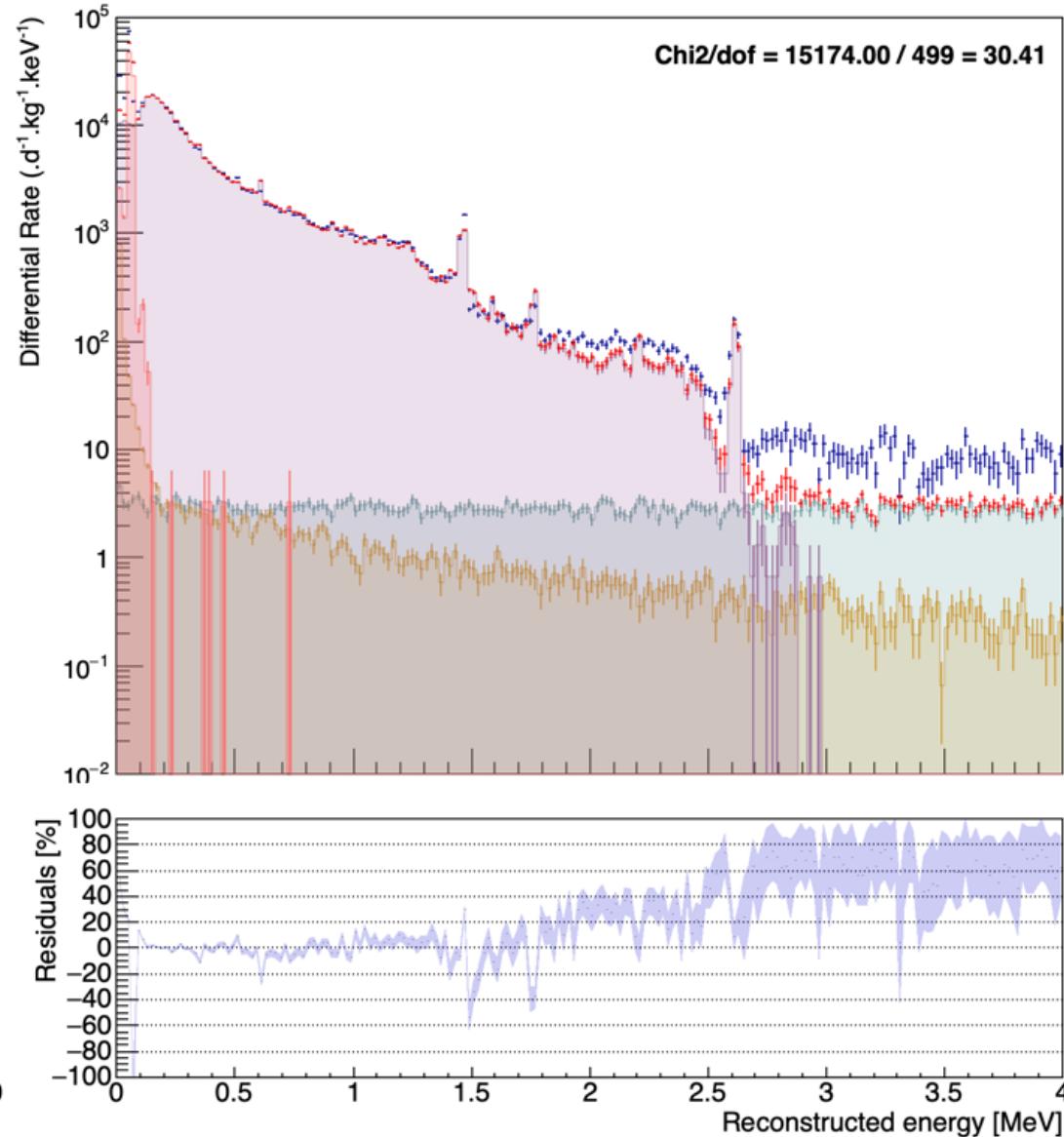
+

## Fit results – Bot Ge

Bot Ge (COV1) - 0-20MeV

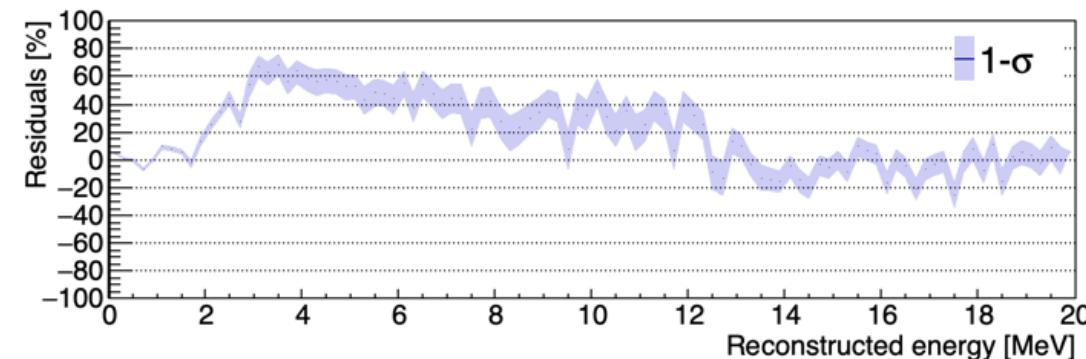
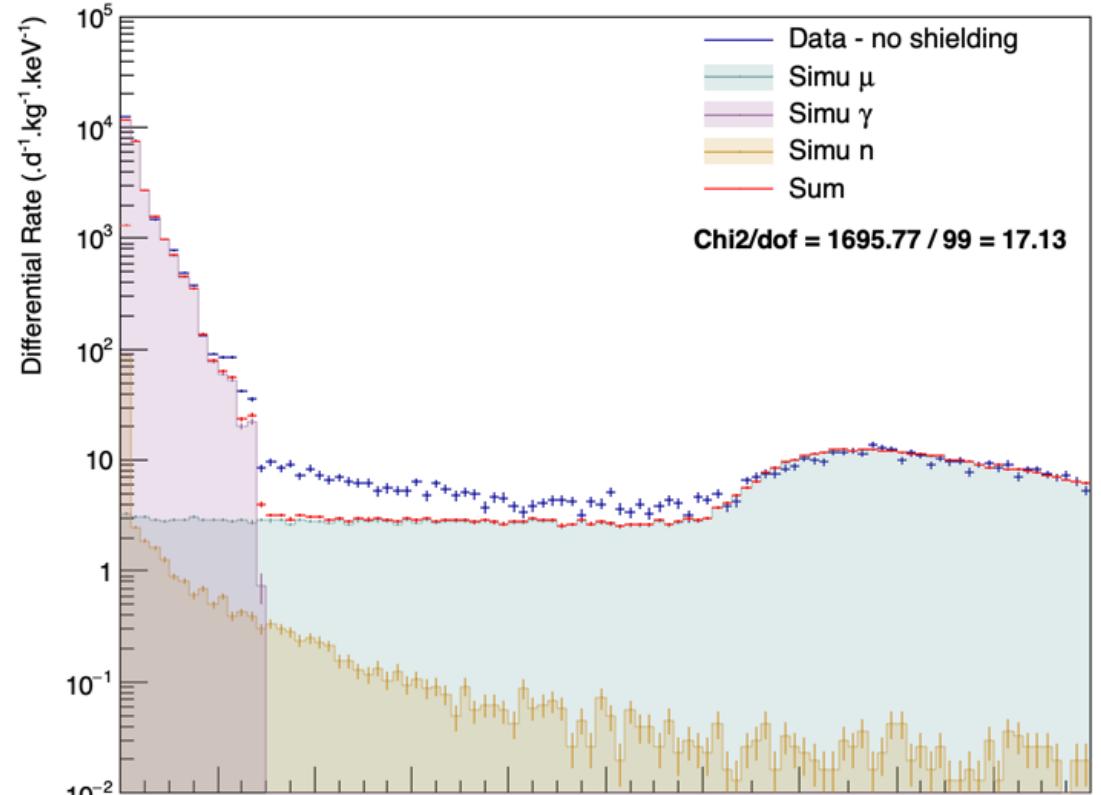


Bot Ge (COV1) - 0-4MeV

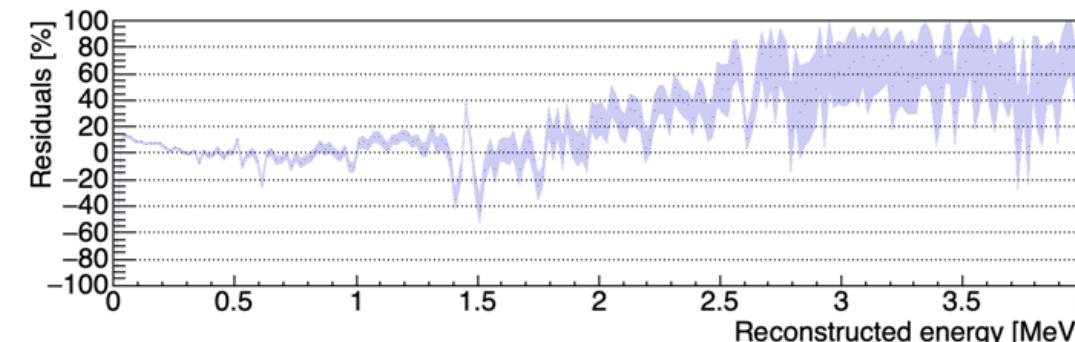
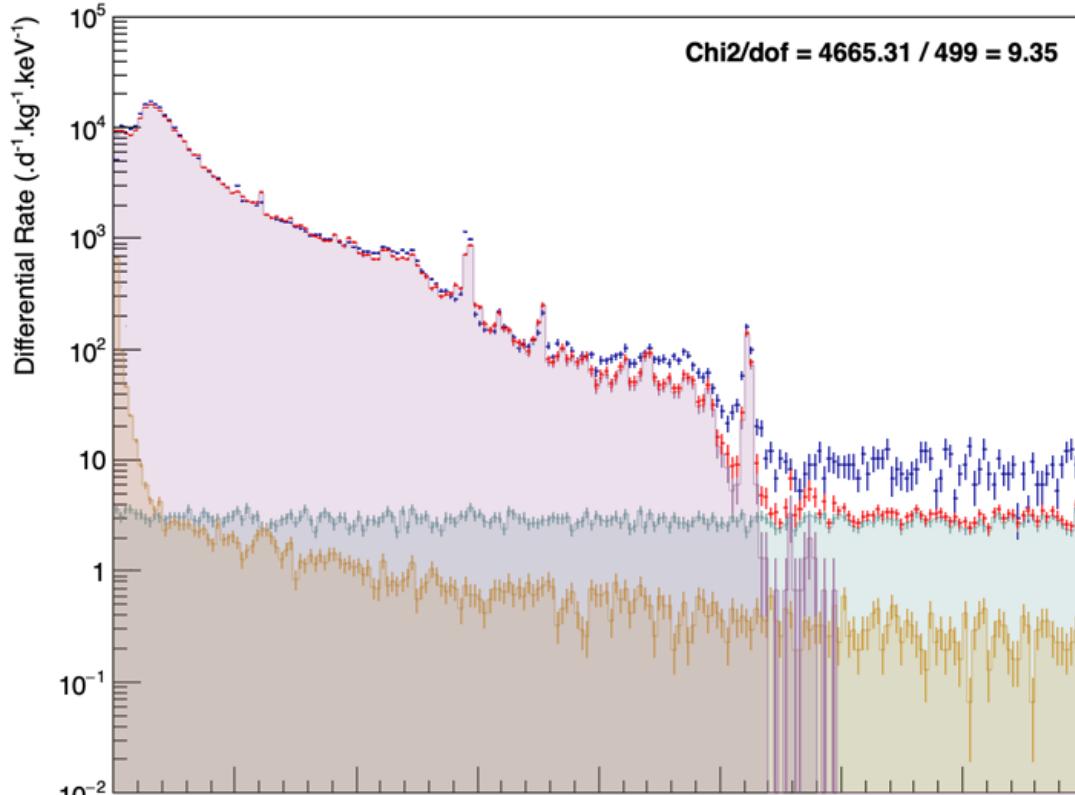


## Fit results – Top Ge

Top Ge (COV2) - 0-20MeV



Top Ge (COV2) - 0-4MeV

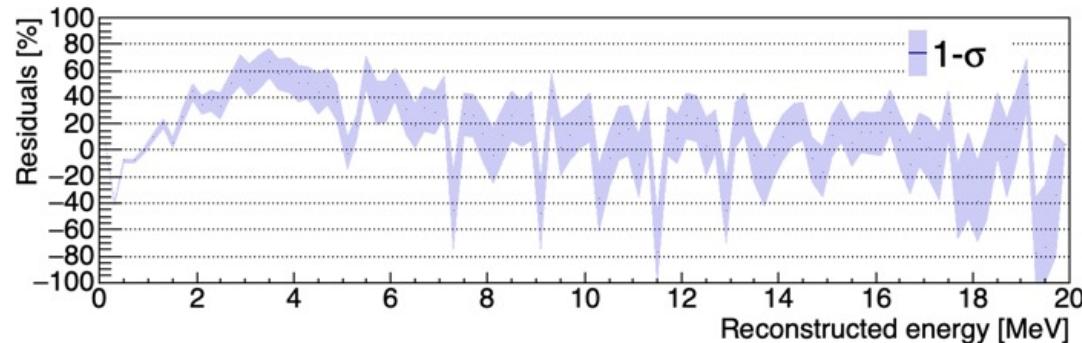
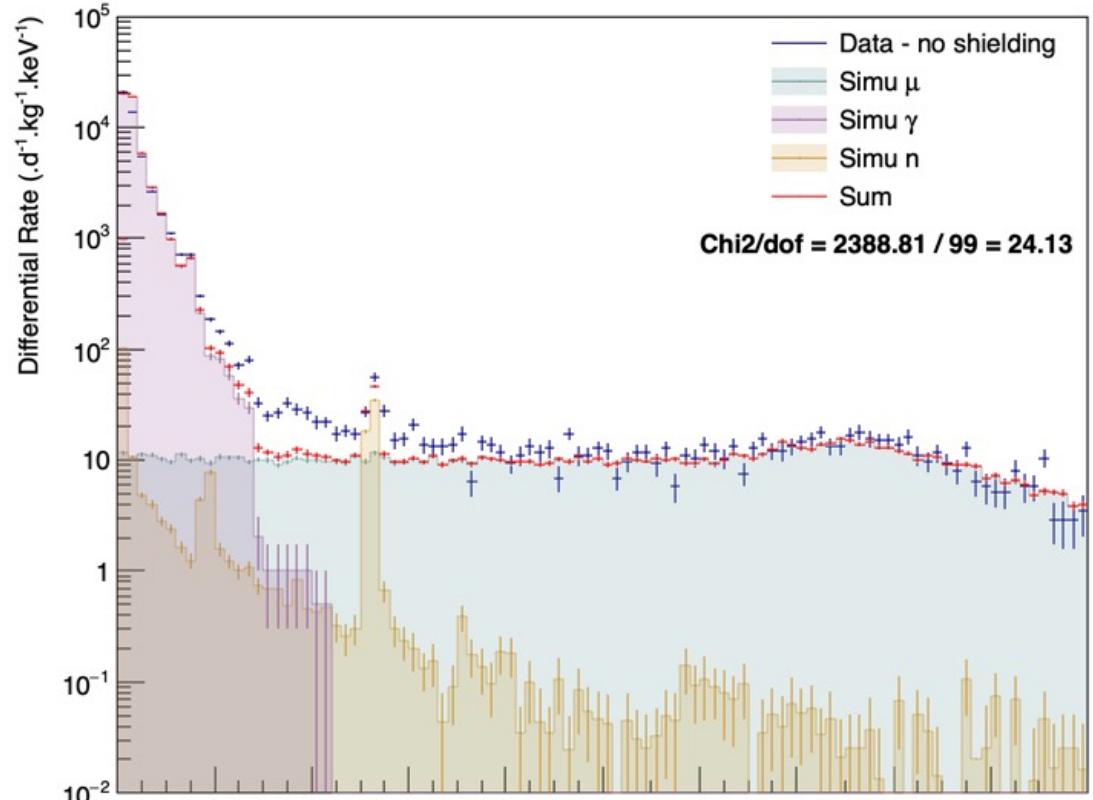


Remark:

Rate lower in Top Ge than in Bot Ge the cryostat has a shielding effect\*.

## Fit results - LWO

Mid LWO (NTD) - 0-20MeV



Mid LWO (NTD) - 0-6MeV

