

Neutrino-nucleus interaction event generators for neutrino oscillation experiments



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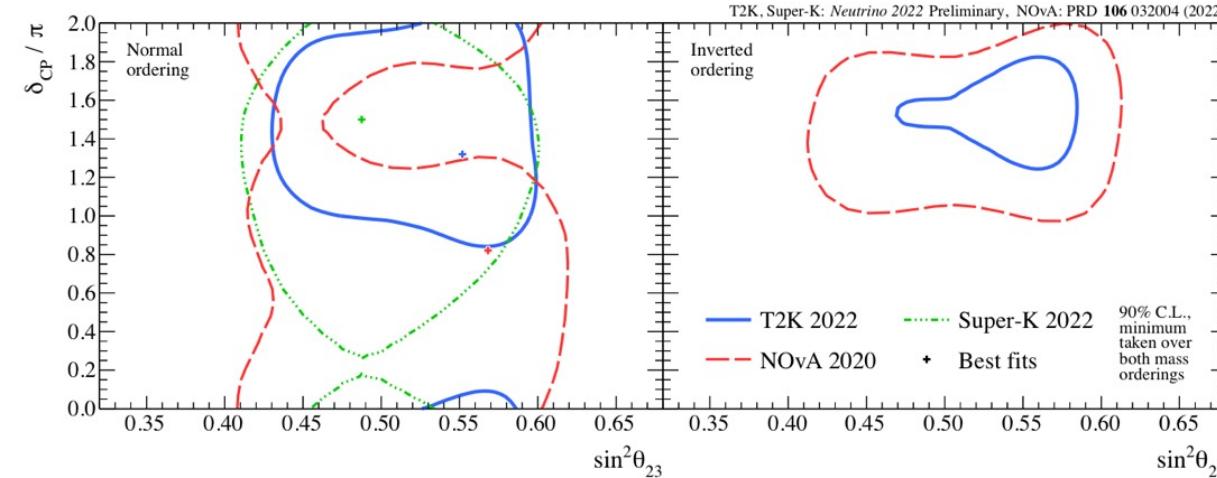


Outline

- Challenges for current and future long-baseline experiments
- Nuclear effects – what do we need to control?
- The role of neutrino event generators
- Success/limitations
- Solutions?

Current experiments – results and challenges

Long-baseline experiments are **uniquely suited to search for CP violation** in the lepton sector and study 3-flavor oscillations



- Measurements of CP violation are **severely limited by statistics** and knowledge of MO
- Now entering the **precision measurement era**

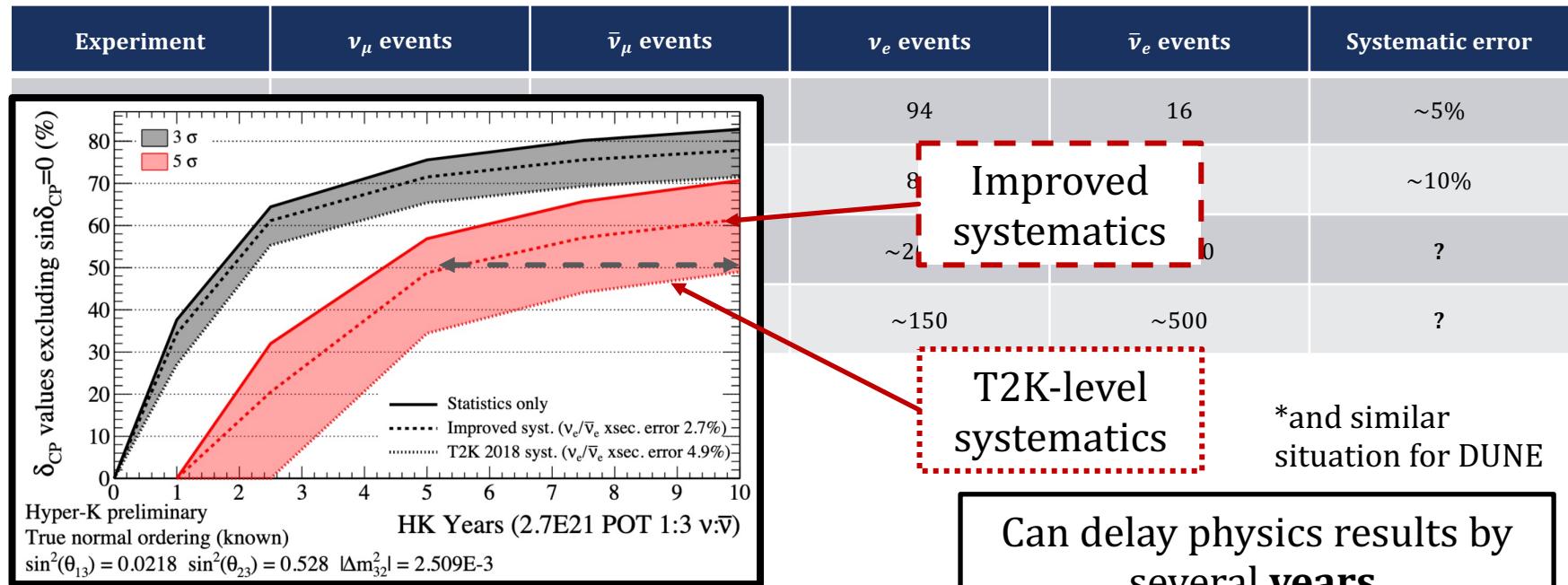
Are we prepared for future experiments?

Experiment	ν_μ events	$\bar{\nu}_\mu$ events	ν_e events	$\bar{\nu}_e$ events	Systematic error
T2K	318	137	94	16	$\sim 5\%$
NOvA	211	105	82	33	$\sim 10\%$

Are we prepared for future experiments?

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T2K	318	137	94	16	~5%
NOvA	211	105	82	33	~10%
Hyper-K	~10000	~14000	~2000	~2000	?
DUNE	~7000	~3500	~150	~500	?

Are we prepared for future experiments?



Sensitivity to exclude CP conserving values for δ_{CP}

Can delay physics results by several years.

Or prevent them altogether!

Are we prepared for future experiments?

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Hyper-K	~10000	~14000	~2000	~2000	Need ~1%-3%!
DUNE	~7000	~3500	~150	~500	Need ~1%-3%!

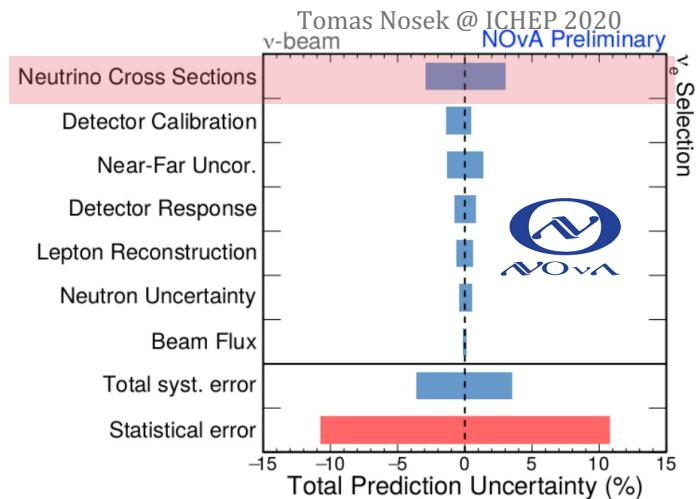
- **Woefully unprepared** at current level of systematics!
- Need dedicated, focused effort in order for future experiments not to be **pre-maturely limited by systematics**

Finding the culprit



Error source (units: %)	1R		MR		1Re			
	FHC	RHC	FHC	CC1 π^+	FHC	RHC	FHC	CC1 π^+
Flux	2.8	2.9	2.8		2.8	3.0	2.8	2.2
Xsec (ND constr)	3.7	3.5	3.0		3.8	3.5	4.1	2.4
Flux+Xsec (ND constr)	2.7	2.6	2.2		2.8	2.7	3.4	2.3
Xsec (ND unconstr)	0.7	2.4	1.4		2.9	3.3	2.8	3.7
SK+SI+PN	2.0	1.7	4.1		3.1	3.8	13.6	1.2
Total All	3.4	3.9	4.9		5.2	5.8	14.3	4.5

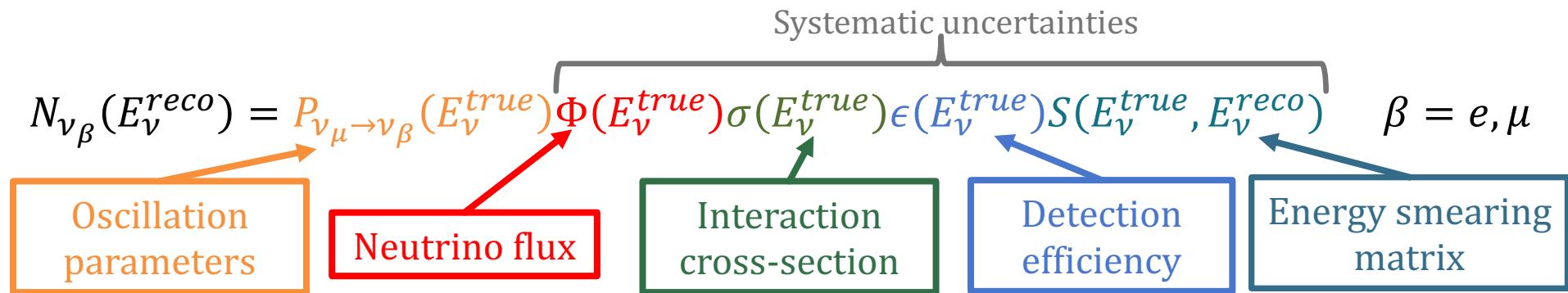
T2K Run 1-10, preliminary



The description of **neutrino-nucleus interactions** is the dominant source of systematic uncertainty for oscillation measurements

Neutrino cross-sections and oscillations

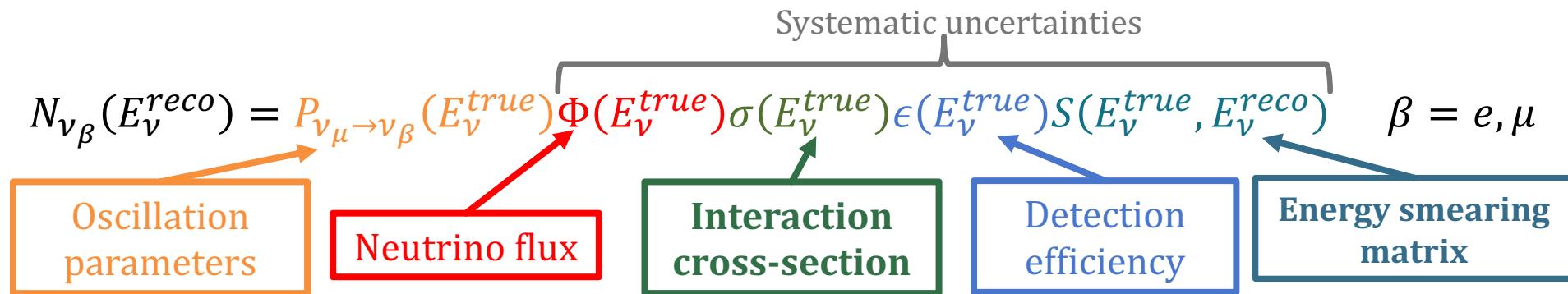
- Oscillation parameters are inferred from event spectra as a function of neutrino energy



- Constrain flux + cross-section systematics with near detector
- But **heavily rely on models** to predict near-to-far detector extrapolation + neutrino energy smearing matrix

Neutrino cross-sections and oscillations

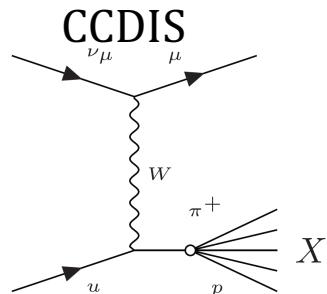
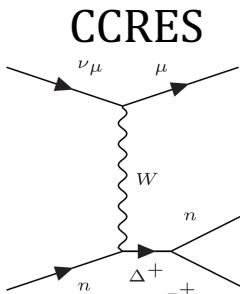
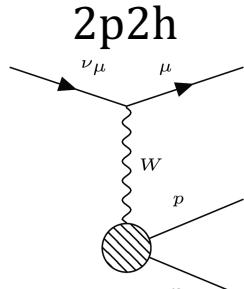
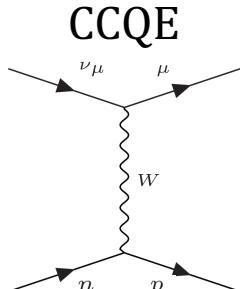
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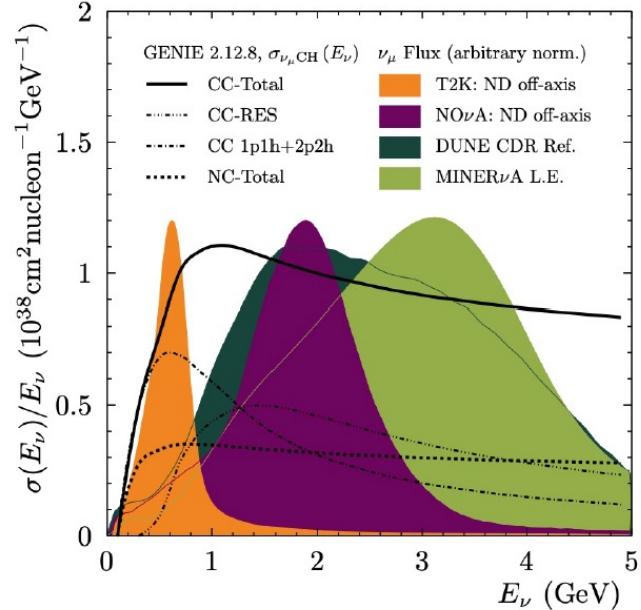
- Constrain flux + cross-section systematics with near detector
- But **heavily rely on models** to predict near-to-far detector extrapolation + neutrino energy smearing matrix

Same – but different!

Charged current channels essential for
reliable flavor identification
Same for most experiments!



But...

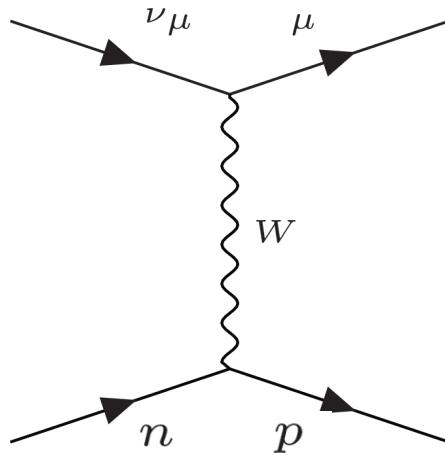


- Their relative contribution varies across experiments (different **flux**)
- **Uncertainties** related to **nucleon-level processes**

From the nucleON to the nucleUS

- Nucleons are bound inside complex nuclei

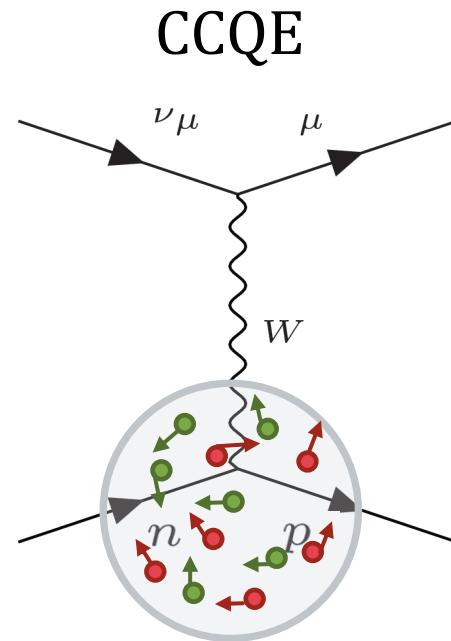
CCQE



"Simplest" CC interaction

From the nucleON to the nucleUS

- Nucleons are bound inside complex nuclei

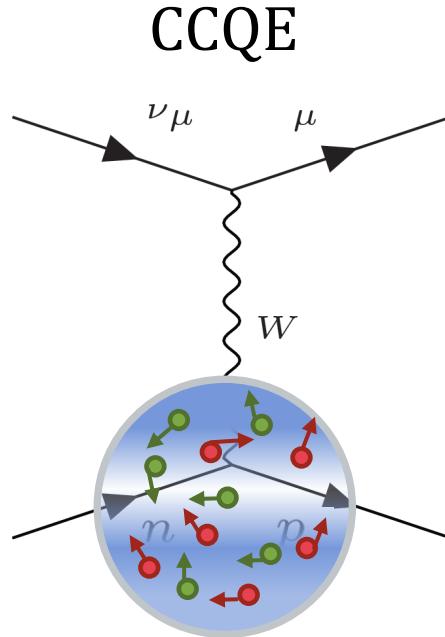


Nuclear effects

- Fermi motion

From the nucleON to the nucleUS

- Nucleons are bound inside complex nuclei

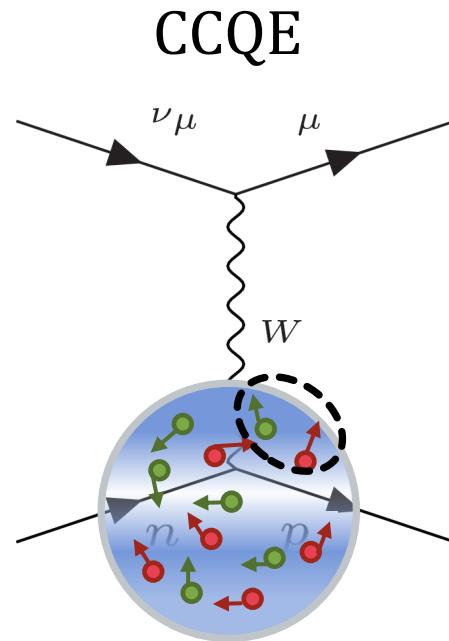


Nuclear effects

- Fermi motion
- Nuclear/optical potential

From the nucleON to the nucleUS

- Nucleons are bound inside complex nuclei

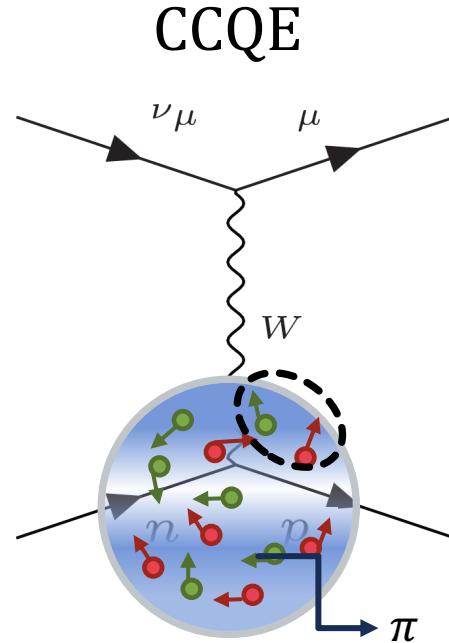


Nuclear effects

- Fermi motion
- Nuclear/optical potential
- Correlations between nucleons

From the nucleON to the nucleUS

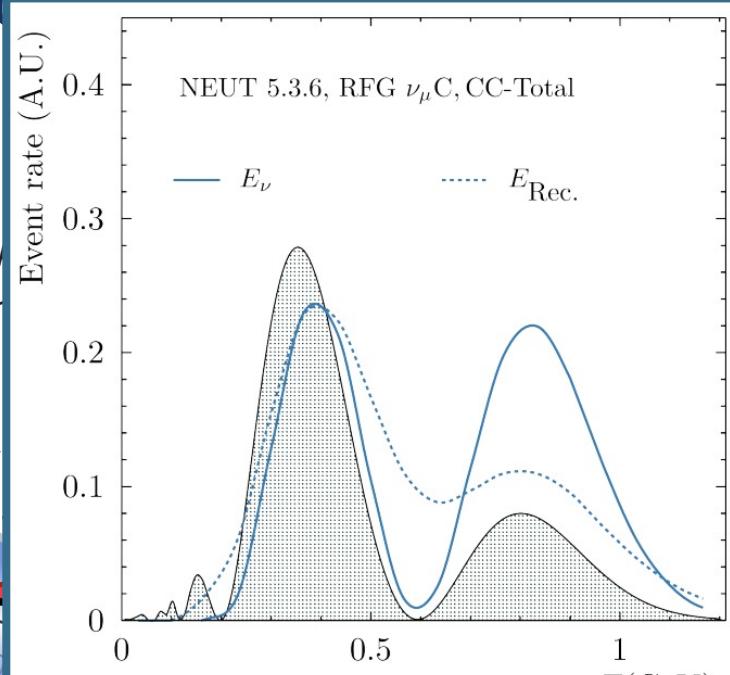
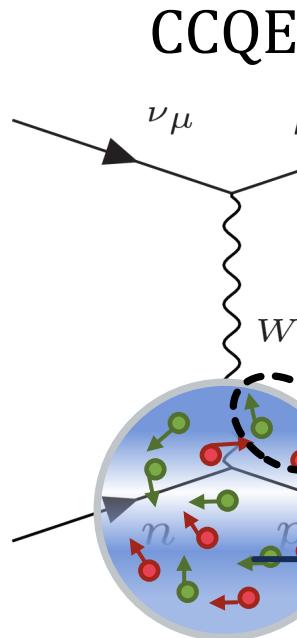
- Nucleons are bound inside complex nuclei



Nuclear effects

- Fermi motion
- Nuclear/optical potential
- Correlations between nucleons
- Interactions with the nuclear medium (Final State Interactions or FSI)

From the nucleON to the nucleUS



Impact of nuclear effects on
reconstructed neutrino energy
(oscillated spectrum)

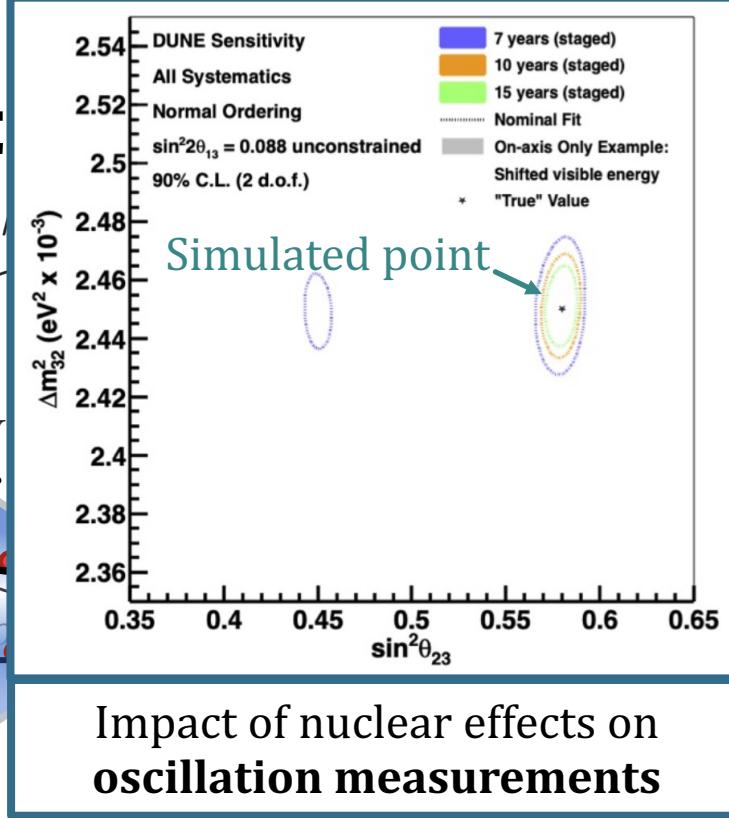
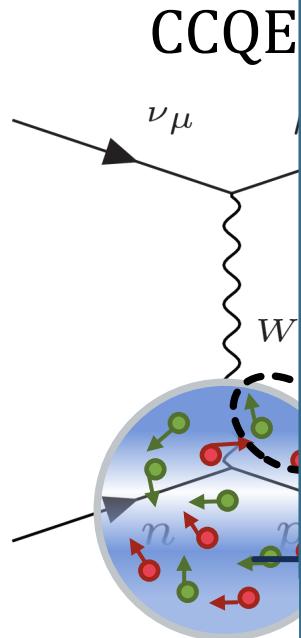
inside complex nuclei

potential

nucleons

nuclear medium (Final
FSI)

From the nucleON to the nucleUS



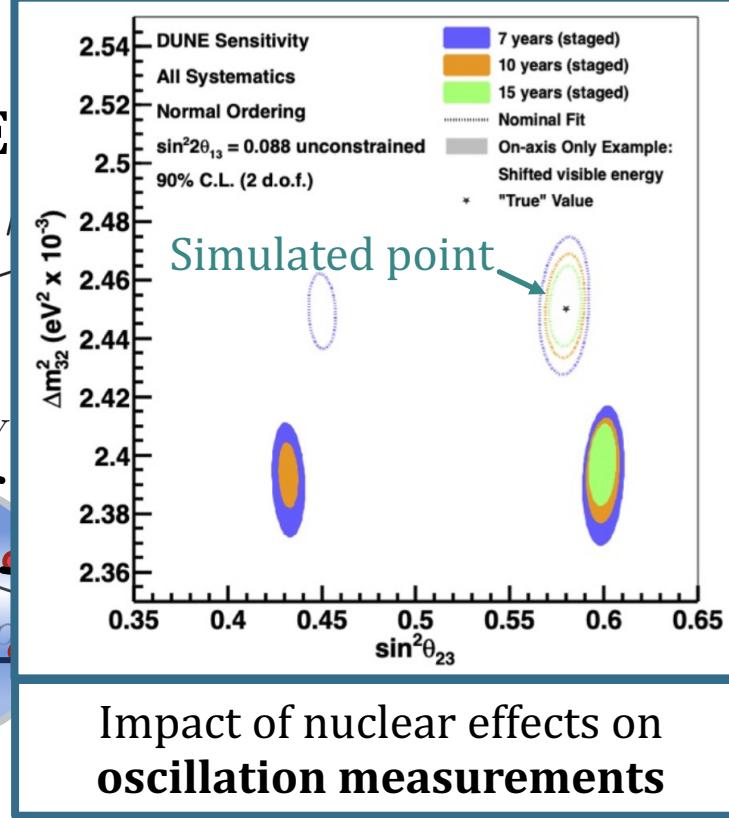
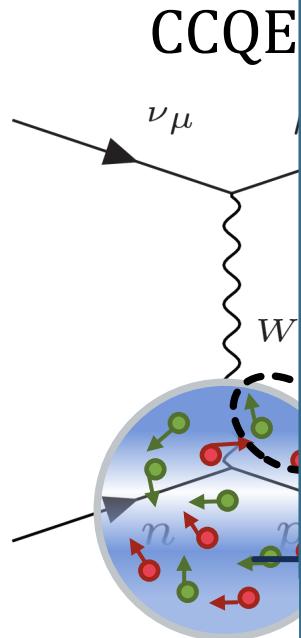
inside complex nuclei

ntial

n nucleons

nuclear medium (Final FSI)

From the nucleON to the nucleUS



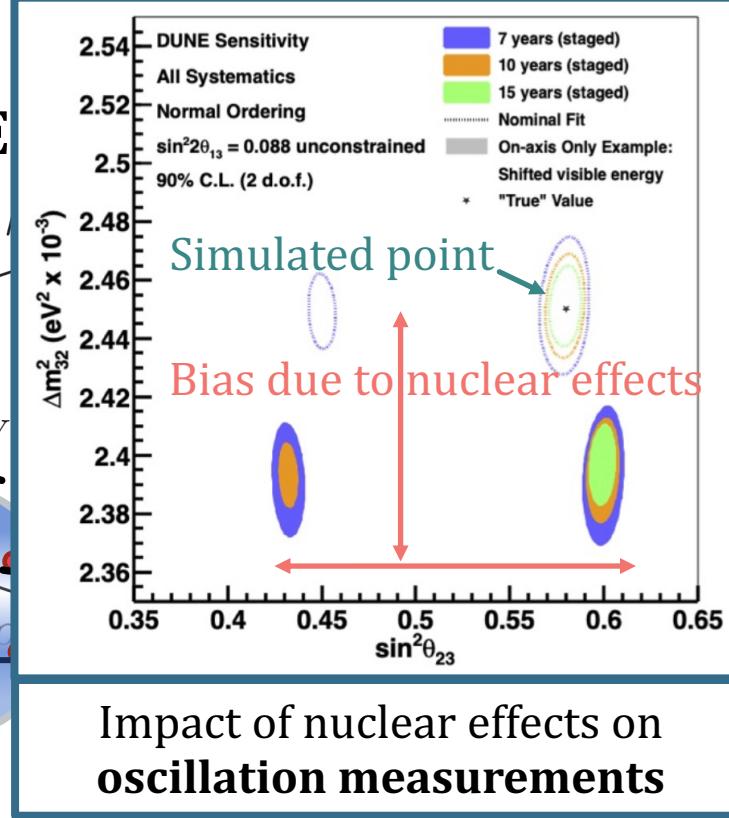
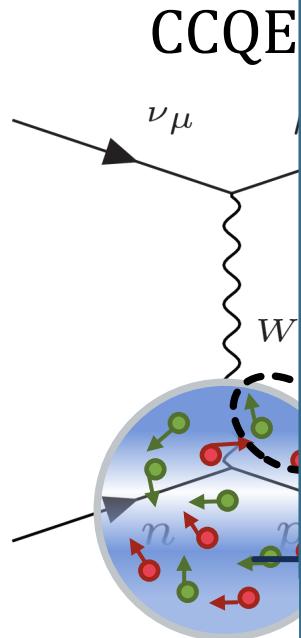
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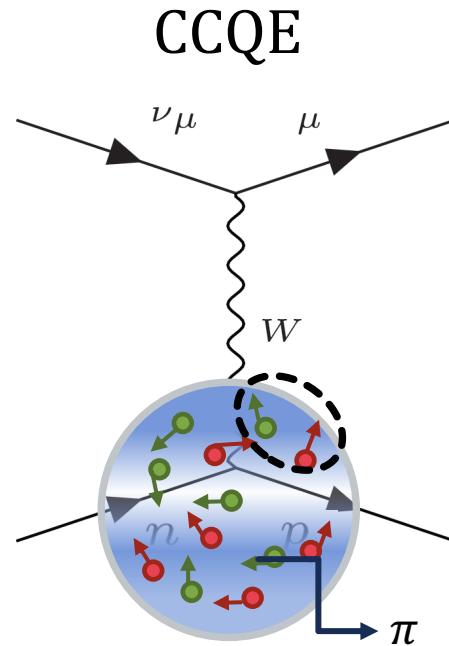
inside complex nuclei

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n nucleons

nuclear medium (Final FSI)

From the nucleON to the nucleUS



- Nucleons are bound inside complex nuclei

Nuclear effects

- Fermi motion
 - Nuclear/optical potential
 - Correlations between nucleons
 - Interactions with the nuclear medium (Final State Interactions or FSI)
- + so many targets (C, H, O, Ar), so little data...**
- + different models for each type of interaction/effect/flavor**

The role(s) of neutrino interaction generators

- Provide a (sufficiently) accurate **description** of neutrino-nucleus interactions
 - To be used as input model to oscillation analyses/model backgrounds
- **Compare** theoretical predictions to experimental measurements
 - To benchmark our models
- **Propagate uncertainties** related to the modelling of neutrino-nucleus cross-sections



Neutrino interaction generators*

*Most widely used by accelerator-based experiments

Nucl.Instrum.Meth.A 614 (2010) 87-104

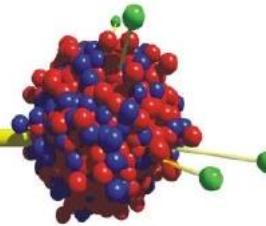


Nucl.Phys.B Proc.Suppl. 159 (2006) 211-216



Phys.Rept. 512 (2012) 1-124

GiBUU

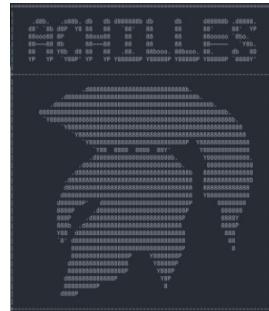


Eur.Phys.J ST 230 (2021) 24-44469-4481



NEUT

Phys.Rev.D 107 (2023) 3, 033007



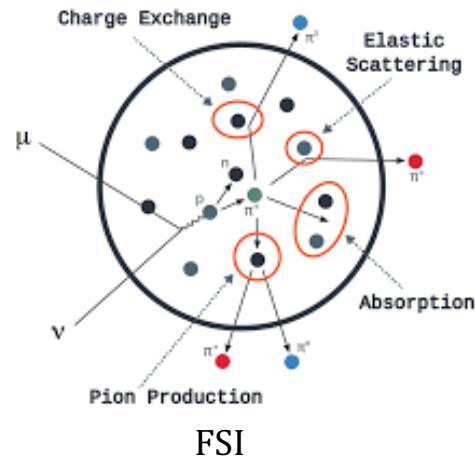
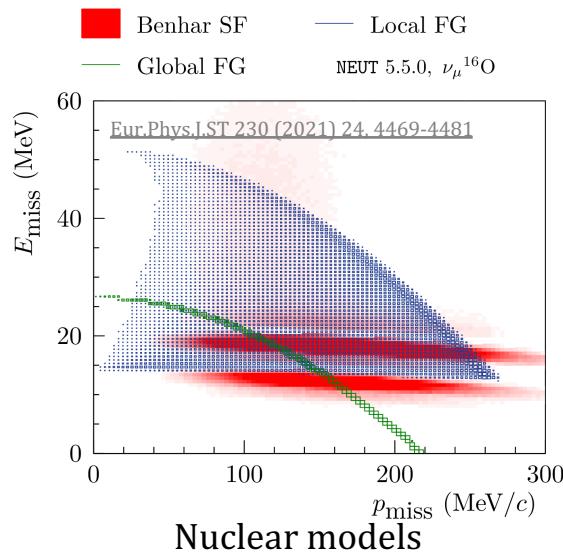
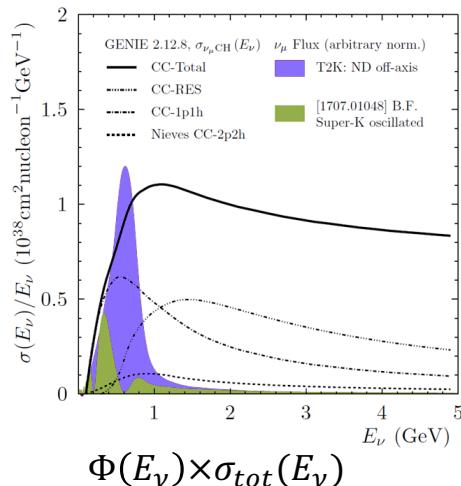
How do generators work?

Sample neutrino
 $\Phi(E_\nu) \times \sigma_{tot}(E_\nu)$

Determine interaction channel

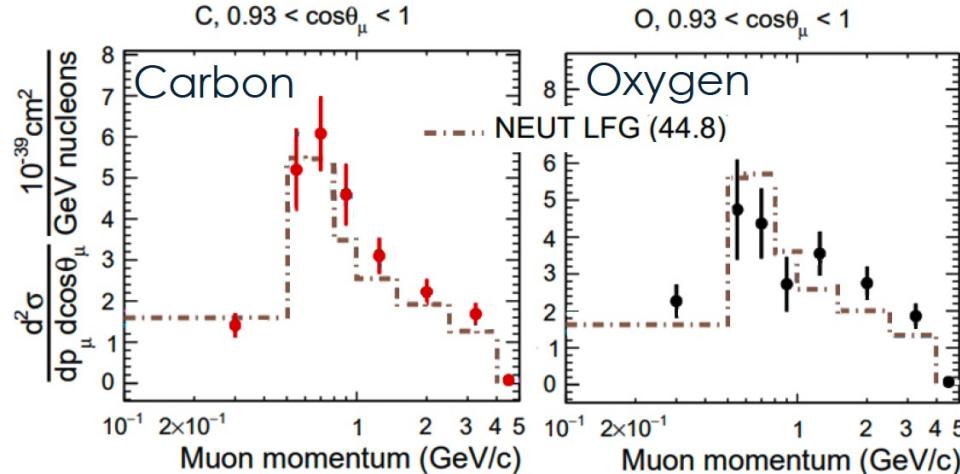
Sample kinematics
from model

Transport through
nucleus



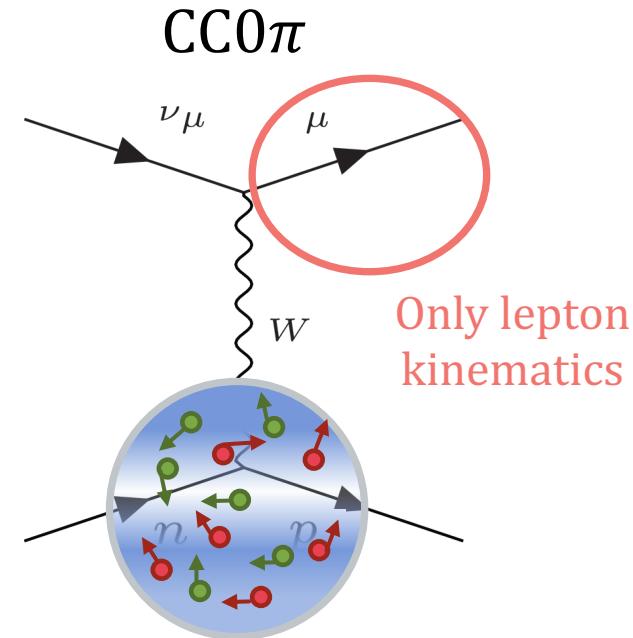
A success story

T2K CC0 π



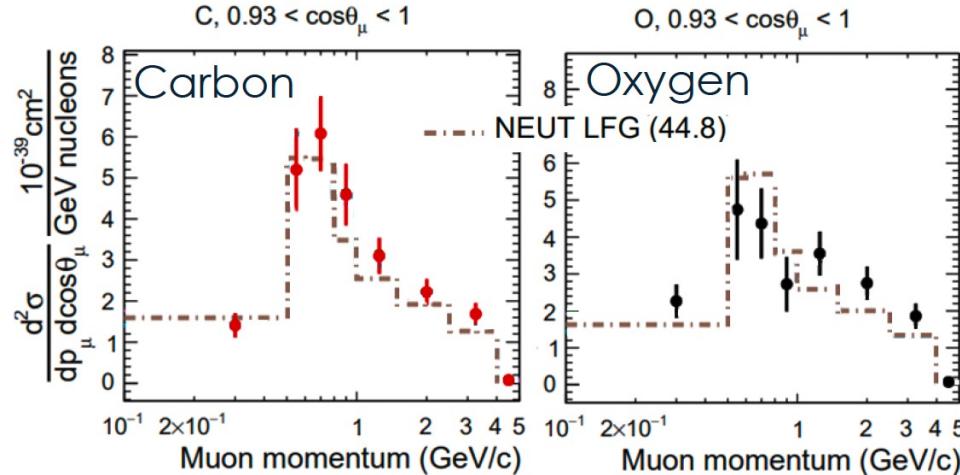
[Phys.Rev.D 101 \(2020\) 11, 112004](https://doi.org/10.1103/PhysRevD.101.112004)

Generators do a (reasonably) good job at predicting the cross-section for **inclusive** measurements



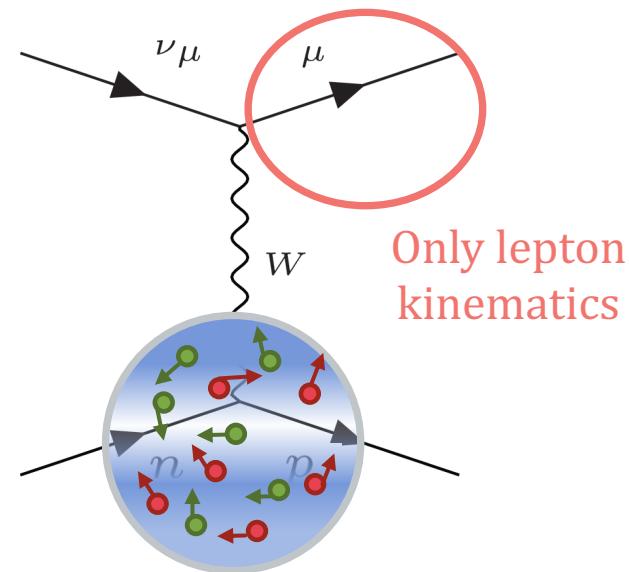
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[Phys.Rev.D 101 \(2020\) 11, 112004](#)

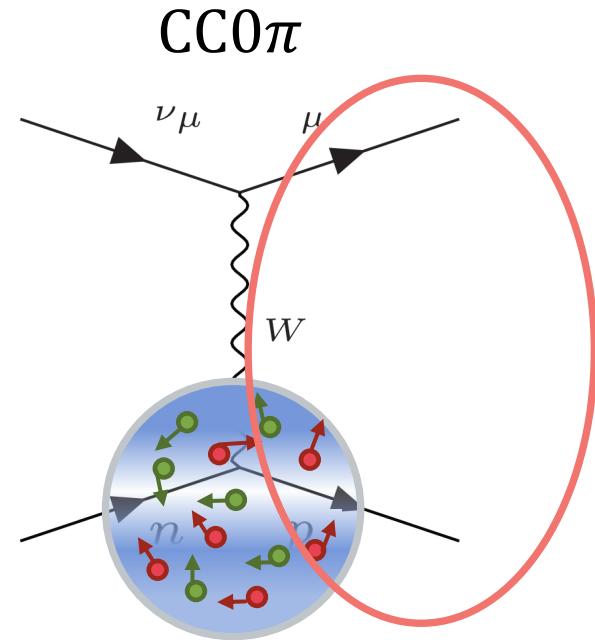
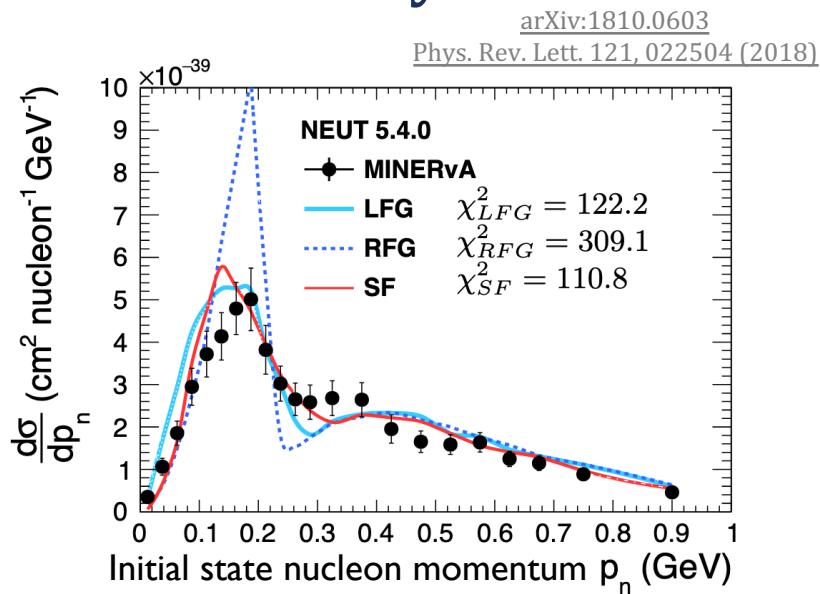
CC0 π



Generators do a (reasonably) good job at predicting the cross-section for **inclusive** measurements

...except at low q_0 ($\sim 15\%$ of events) - **more on that later!**

A success story?



But **none of them** offer a good quantitative prediction for (semi-)exclusive measurements

Lepton+hadron kinematics

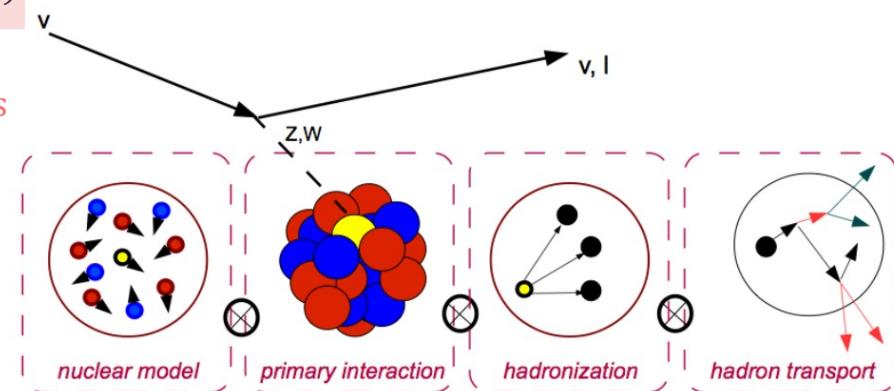
Reaching the limits of our approximations

- Generators (often) factorize the process based on the plane wave impulse approximation (PWIA)
 - IA – the neutrino interacts with an isolated nucleon.
 - PW – the final state nucleon exits the nucleus as a plane wave.

$$\frac{d^5\sigma_{v\ell}}{d\Omega(\hat{k}')d\Omega(p_N)dE_{\ell}'} \sim S(E_m, \mathbf{p}_m) L_{\mu\nu} W^{\mu\nu} \delta(\omega + M - E_m - E_{p'})$$

Nuclear spectral function

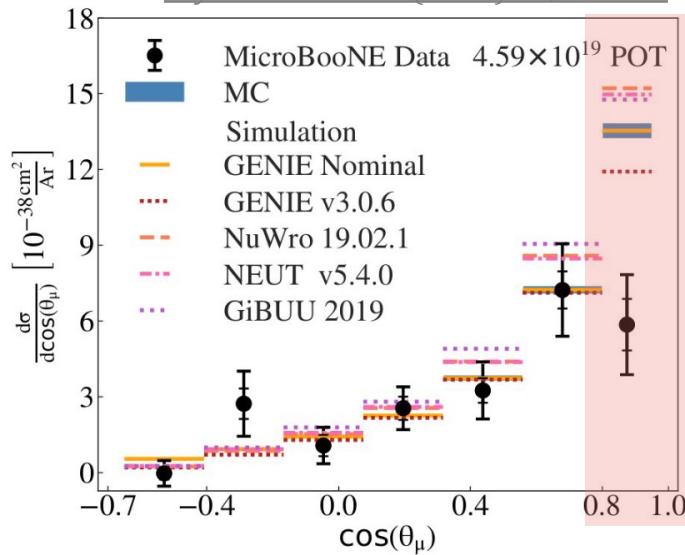
Single nucleon dynamics



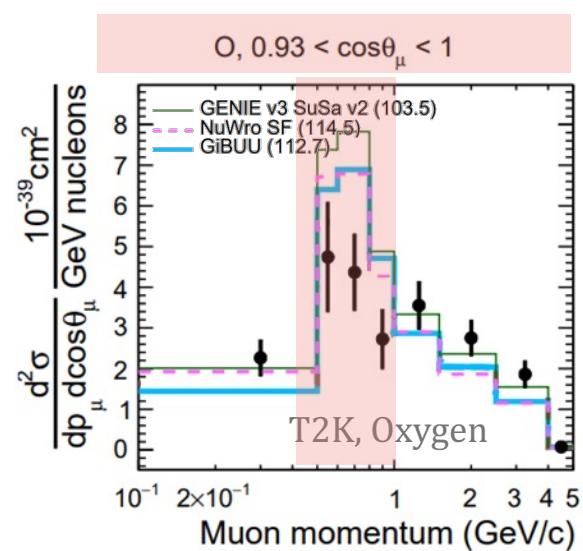
- X Nucleon correlations added in an ad-hoc way
- X FSI modelled through semi-classical cascades

How bad is it?

Phys.Rev.Lett. 125 (2020) 20, 201803

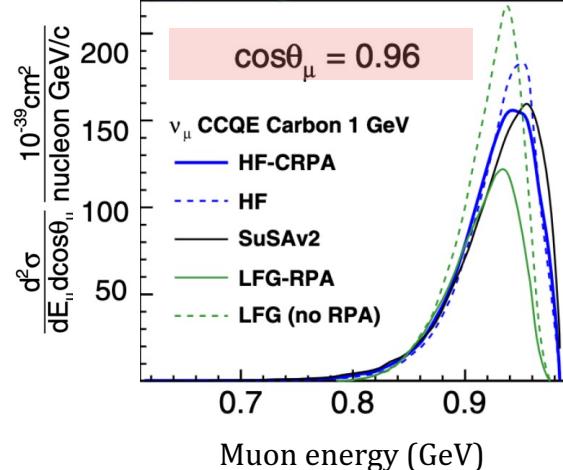


Phys.Rev.D 101 (2020) 11, 112004



Model disagreement

Phys.Rev.D 106 (2022) 7, 073001



These effects show up at forward scattering angles (low energy transfer)

PWIA breaks down

Increasingly important for oscillation measurements

The curse of dimensionality

- We ask generators to give predictions for X dimensions
- But theory inputs are usually inclusive (Y dimensions)
- If $X > Y$ – assign the kinematics for remaining X-Y dimensions using a best guess approach!

Channel	What we ask from the generators	What theorists give us
CCQE	$\frac{d^5\sigma}{dE_l d\cos\theta_l dE_N d\cos\theta_N d\cos\theta_{lN}}$	$\frac{d^5\sigma}{dE_l d\cos\theta_l dE_N d\cos\theta_N d\cos\theta_{lN}}$
2p2p	$\frac{d^8\sigma}{dq_0 dq_3 d\mathbf{p}_1 d\mathbf{p}_2}$	$\frac{d^2\sigma}{dq_0 dq_3}$
Resonant single meson production	$\frac{d^8\sigma}{dq_0 dq_3 d\mathbf{p}_N d\mathbf{p}_\pi}$	$\frac{d^2\sigma}{dQ^2 dW}$
Deep inelastic scattering	$\frac{d^N\sigma}{dx dy \prod_i d\mathbf{p}_i}$	$\frac{d^2\sigma}{dx dy}$

The curse of dimensionality

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Channel	What we ask from the generators	What theorists give us
CCQE	$\frac{d^5\sigma}{dE_ldcos\theta_ldE_Ndcos\theta_Ndcos\theta_{lN}}$	$\frac{d^5\sigma}{dE_ldcos\theta_ldE_Ndcos\theta_Ndcos\theta_{lN}}$
2p2p	$\frac{d^8\sigma}{dq_0dq_3dp_1dp_2}$	$\frac{d^2\sigma}{dq_0dq_3}$
Resonant single meson production	$\frac{d^8\sigma}{dq_0dq_3dp_Ndp_\pi}$	$\frac{d^2\sigma}{dQ^2dW}$
Deep inelastic scattering	$\frac{d^N\sigma}{dxdy \prod_i d\mathbf{p}_i}$	$\frac{d^2\sigma}{dxdy}$



What can we do about it?

- Consolidate strong ties with the theory community
 - Essential to iterate with theory developments as we make new, more exclusive measurements
- Take more data – make more **exclusive** measurements!
- Continue improving our generators
 - More sophisticated models exist but are not implemented yet – **active effort!**
- In the meantime, include sufficient freedoms in our analyses
 - Flexibility is key!

What can we do about it?

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- Take more data -
- Continue improving
 - More sophisticated analyses
- In the meantime,
 - Flexibility is key!

Systematic errors due to ν cross section and flux uncertainties are dominant ($\sim 3\%$) ...

It is faster and cheaper to pay a theoretician to reduce 2 % your systematics than building huge detectors!

!

yet - active effort!

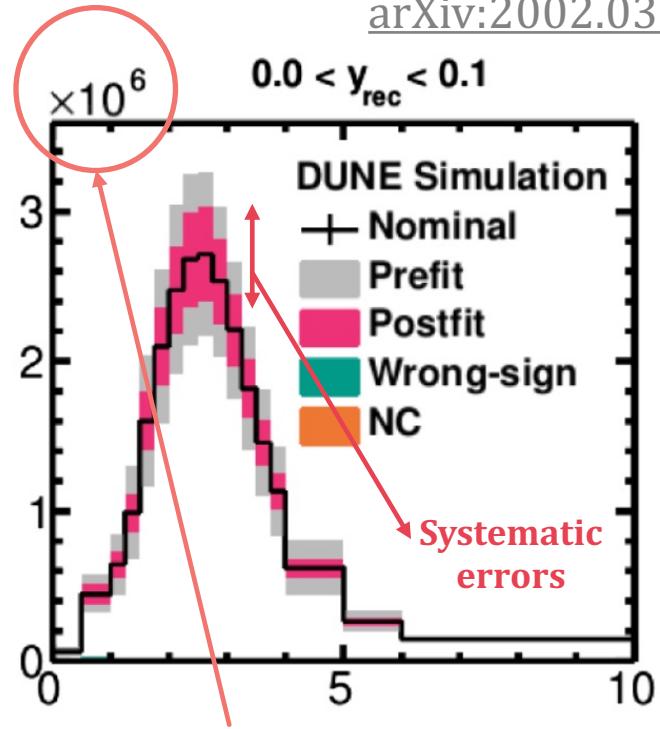
analyses

G. Megias @ NuInt 2018

Flexibility

arXiv:2002.03005

- Future experiments (DUNE/HK) will have unprecedented levels of statistics at the ND
 - Sensitive to the most minute details in our simulations
- To prepare for future analyses we need to include significantly more freedom than we do in current analyses



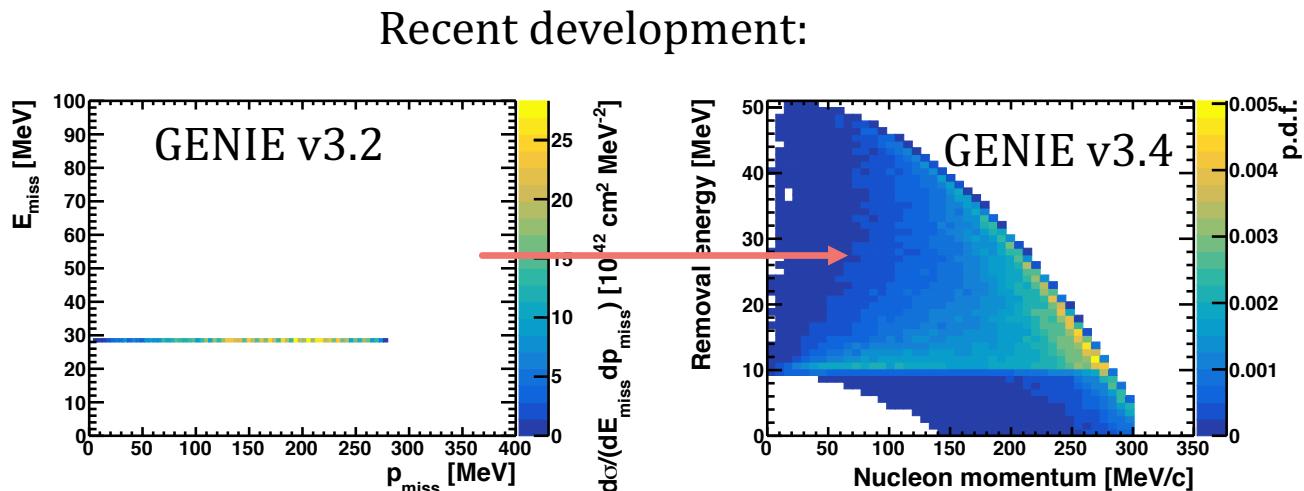
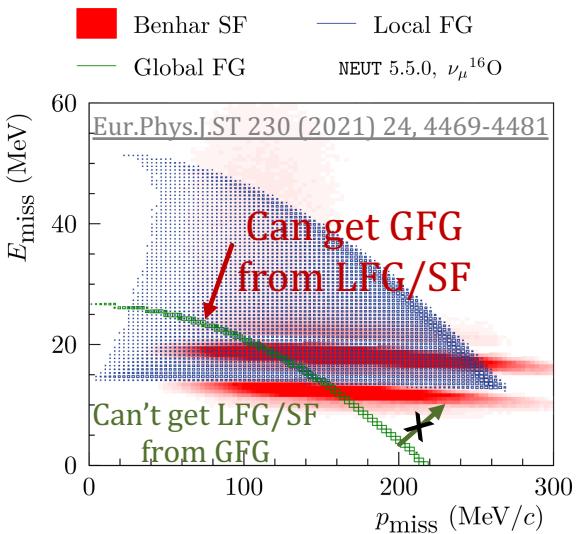
This plot **HAS** statistical error bars!

Flexibility

Example for nuclear model

Examples:

- Generating with an increased (but still motivated) phase space



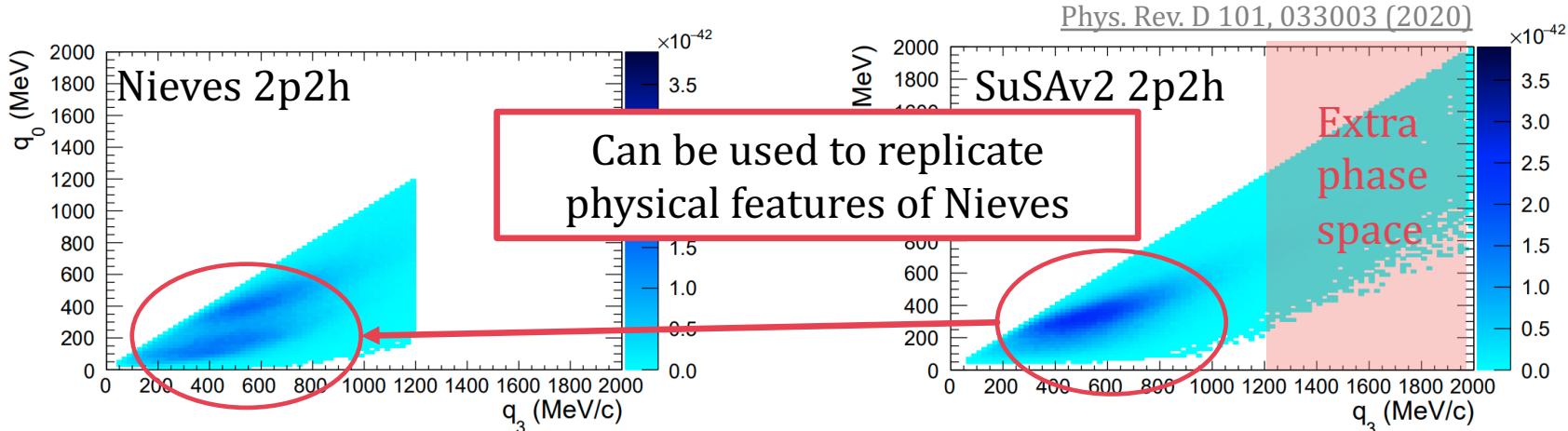
Will be used for DUNE
with new freedoms

Flexibility

Example for 2p2h

Examples:

- Generating with an increased (but still motivated) phase space



SuSA v2 has better agreement with electron scattering data

Sharing the tools

- Generators use different file formats and often different internal conventions
 - Experimental tools are usually built around one generator
 - Swapping generators is difficult
- Several efforts to mitigate this problem:
 - Downstream: **NUISANCE** – widely used tool to synthesize output from all generators and confront them to cross-section measurements
 - But loses interfacing with internal generator flux drivers/detector geometry packages
 - Upstream: Ongoing effort to develop a common format (**NuHEPMC**) which all generators would use



JINST 12 P01016
NUISANCE

Sharing the tools

- Increasingly important as oscillation experiments are starting to perform **joint fits**
 - T2K+SK atmospherics – easy to combine systematics since both use NEUT
 - T2K+NOvA – **extremely hard to combine systematics** because T2K uses NEUT and NOvA uses GENIE (but some of the physics is the same!)
- First step towards sharing tools: recent agreement reached between **DUNE+SBN experiment** (ICARUS and SBND) to use the **same** base model and systematic error tools
 - Easier since all are GENIE-based
 - Work is shared among all collaborations

Towards better communication

- NuInt workshops
 - Essential role in interfacing between theory and experimental/generator community
 - Next one in spring 2024 in Brazil
- NuSTEC
 - Collaboration between theorists and experimentalists – non exhaustive!
- Dedicated generator workshops:
 - TENSIONS, ECT*, Generators Workshop @ Fermilab...
- Communication with nuclear physics community
 - In particular, use of electron scattering data to benchmark models

Summary

- Neutrino-nucleus interaction uncertainties will be the dominant (limiting) source of systematic uncertainties for oscillation experiments
- The physics of these processes is complex and relies on neutrino interaction generator to provide a bridge between theory and experiments
- Current generators have come a long way, but are reaching the limits of their approximations
- Generators should be used to prepare for future experiments by including appropriate degrees of freedom in analyses
- Ongoing efforts to facilitate the communication between generators and their communities
- Imperative to ensure communication with theory community in this process

Thank you for your attention!



Supplementary slides

Neutrino interaction generators*



GENIE

- Arguably most widely-used
- Large active development team (collaboration)
- Support for experiment integration
- Primarily used by Fermilab-based experiments
- Has its own reweighting framework

NEUT

- Originally developed for cosmic neutrino backgrounds for proton decay
- Primarily used by T2K, SK, HK
- Closed-source (may change soon!)
- Small development team
- Has its own reweighting framework



*Most widely used by accelerator-based experiments

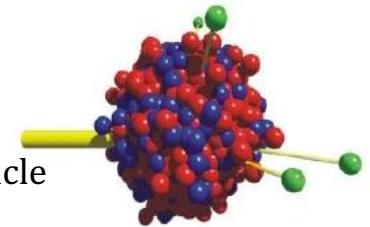
NuWro

- Wide range of available models
- Developed and maintained mostly by theorists
- Lack of reweighting framework
- Primarily used for comparisons to data



GiBUU

- Sophisticated quantum mechanical treatment of particle transport
- Very small development team
- Difficult to integrate with experimental suites



GiBUU

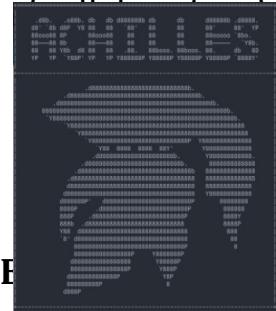
Neutrino interaction generators*

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GENIE

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NuWro

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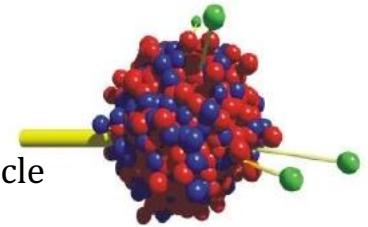
ACHILLES

- Developed mostly by theorists
- Takes inspiration from collider community experiments (FeynRules)
- Good agreement with electron scattering data

mechanical treatment of particle transport

- Very small development team
- Difficult to integrate with experimental suites

GiBUU



Long-baseline oscillation experiments

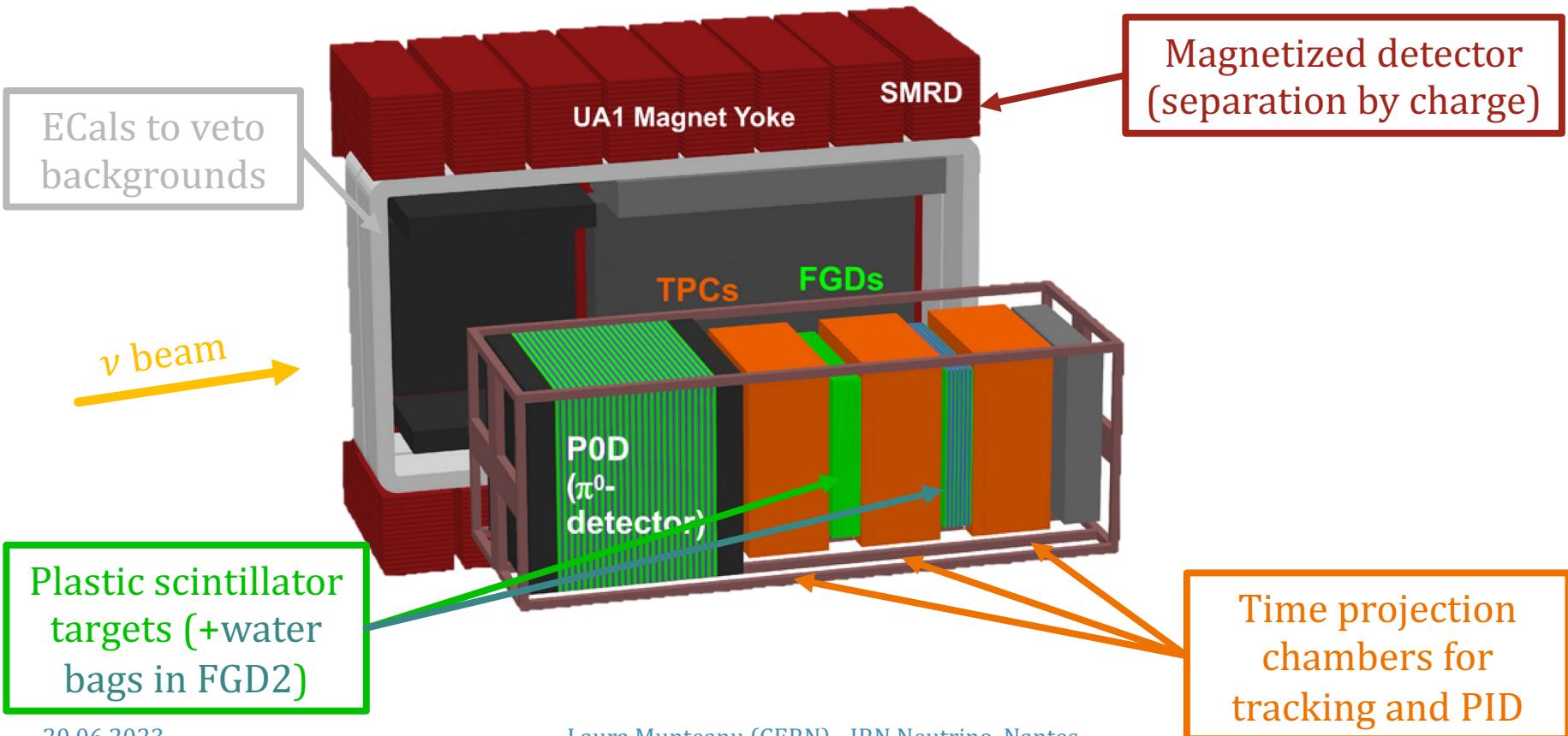
Experiment	Beam Energy	Baseline	Near detector	Far detector	
Current	T2K 	600 MeV (Narrow)	~300 km	Scintillator bars+water	Water Cherenkov
	NOvA 	1.2 GeV (Wide)	~800 km	Plastic scintillator	Plastic scintillator
Future	HyperK 	600 MeV (Narrow)	~300 km	Scintillator cubes + TBD	Water Cherenkov
	DUNE 	2.5 GeV (Wide)	~1200 km	Argon TPC + C/H STT	Liquid Argon TPC + TBD



Baseline

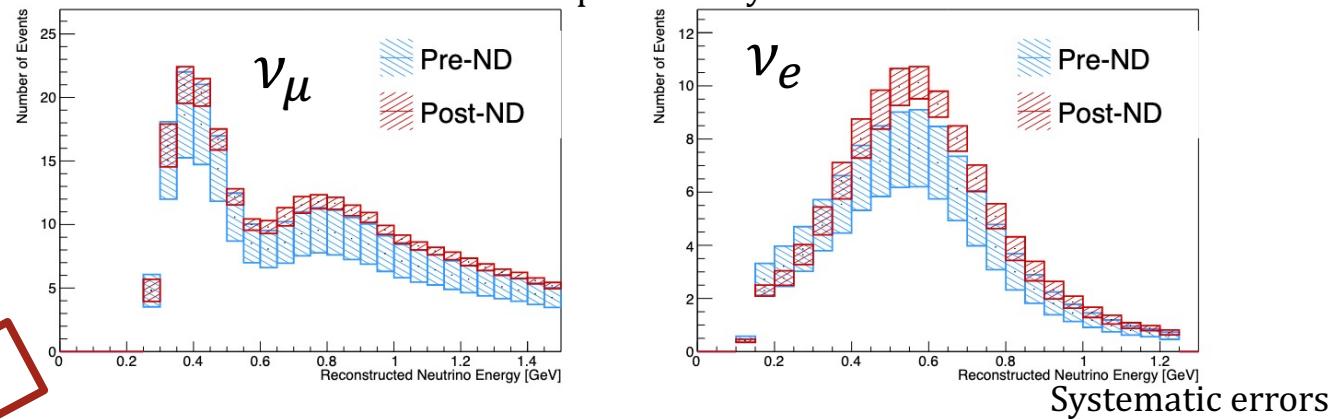
Neutrino Beam

The T2K near detector at 280 m: ND280



Impact of ND constraint on FD spectra and errors

T2K preliminary



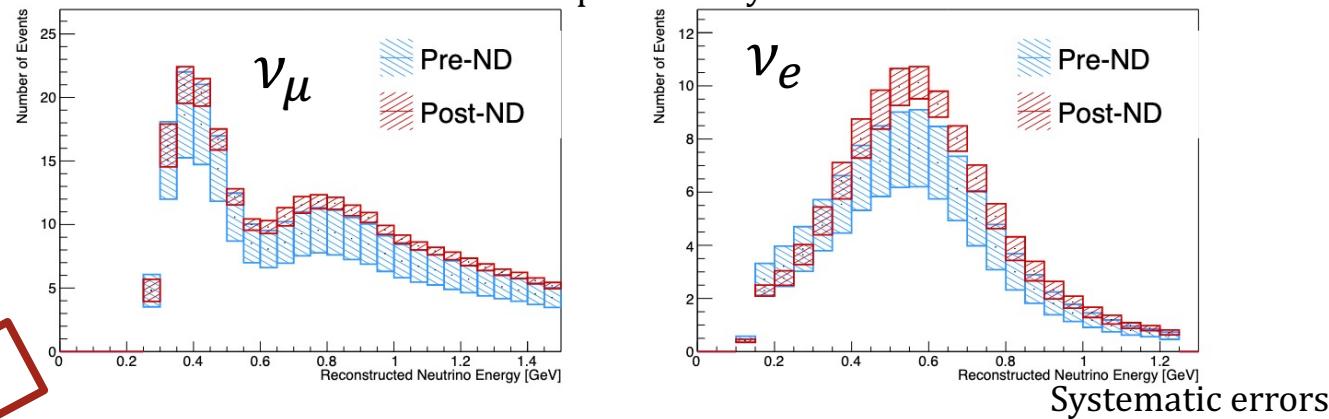
Before ND fit

Error source (units: %)	1R		MR		1Re		FHC/RHC	
	FHC	RHC	FHC	CC1 π^+	FHC	RHC		
Flux	5.0	4.6	5.2		4.9	4.6	5.1	4.5
Cross-section (all)	15.8	13.6	10.6		16.3	13.1	14.7	10.5
SK+SI+PN	2.6	2.2	4.0		3.1	3.9	13.6	1.3
Total All	16.7	14.6	12.5		17.3	14.4	20.9	11.6

T2K Run 1-10, preliminary

Impact of ND constraint on FD spectra and errors

T2K preliminary



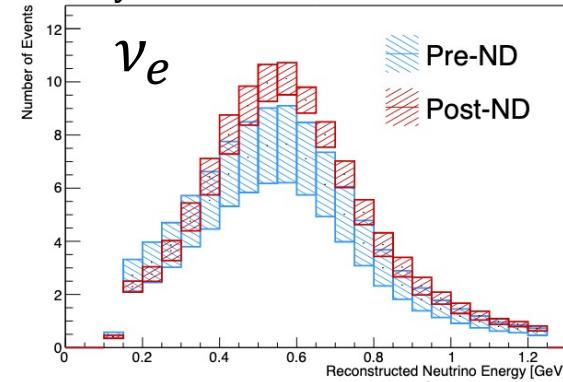
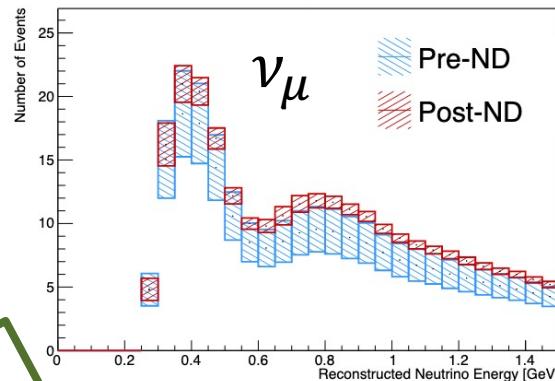
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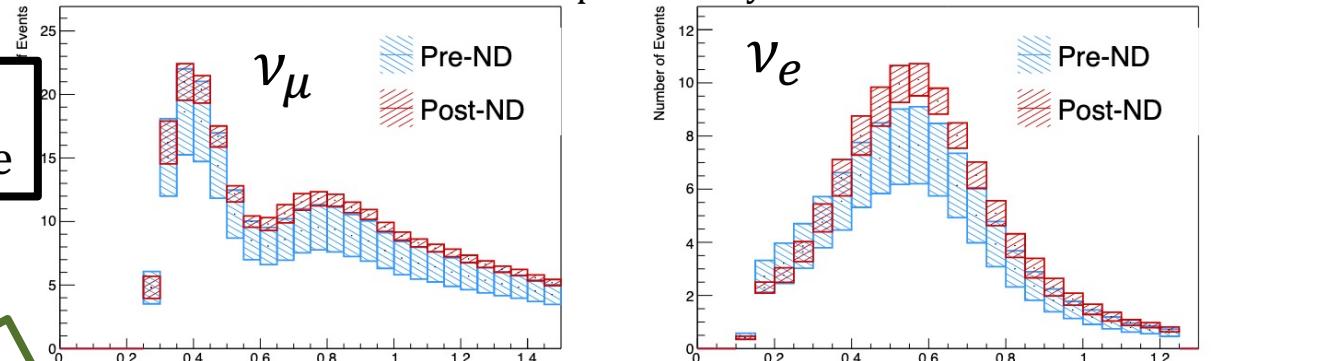
After ND fit

Systematic errors

Error source (units: %)	1R		MR		1Re		FHC/RHC	
	FHC	RHC	FHC	CC1 π^+	FHC	RHC		
Flux	2.8	2.9	2.8		2.8	3.0	2.8	2.2
Xsec (ND constr)	3.7	3.5	3.0		3.8	3.5	4.1	2.4
Flux+Xsec (ND constr)	2.7	2.6	2.2		2.8	2.7	3.4	2.3
Xsec (ND unconstr)	0.7	2.4	1.4		2.9	3.3	2.8	3.7
SK+SI+PN	2.0	1.7	4.1		3.1	3.8	13.6	1.2
Total All	3.4	3.9	4.9		5.2	5.8	14.3	4.5

Impact of ND constraint on FD spectra and errors

T2K preliminary



But also tuned
spectrum shape

After ND fit

Systematic errors

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	FHC	RHC	FHC	CC1 π^+	FHC	RHC	
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The nuclear ground state

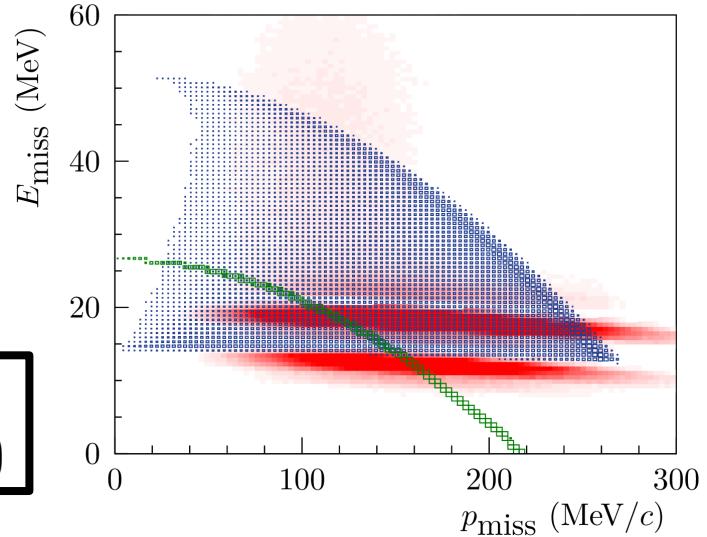
$$\frac{d^5\sigma_{\nu\ell}}{d\Omega(\hat{k}')d\Omega(p_N)dE_{\ell}'} \sim S(E_m, \mathbf{p}_m) \epsilon_{\mu\nu} W^{\mu\nu} \delta(\omega + M - E_m - E_{p'})$$

E_{miss}
(~nucleon removal energy)

p_{miss}
(~nucleon momentum)

[Eur.Phys.J.ST 230 \(2021\) 24, 4469-4481](#)

Benhar SF Local FG
Global FG NEUT 5.5.0, $\nu_\mu {}^{16}\text{O}$



$$\vec{p}_{\text{miss}} = \vec{p}_\nu - \vec{p}_\mu - \vec{p}_p$$

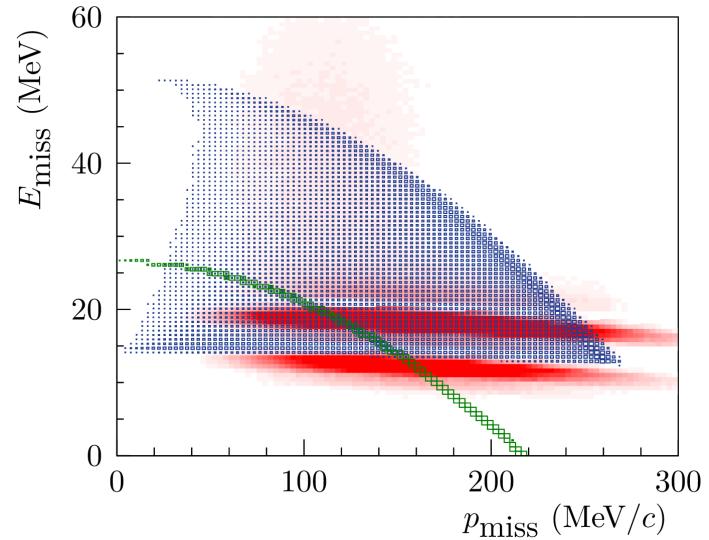
$$E_{\text{miss}} = \omega - T_p^{\text{pre-FSI}} - \Delta m_{n \rightarrow p} - T_{\text{nucr remnant}}$$

The nuclear ground state

- Multiple ground state models are available in neutrino interaction generators

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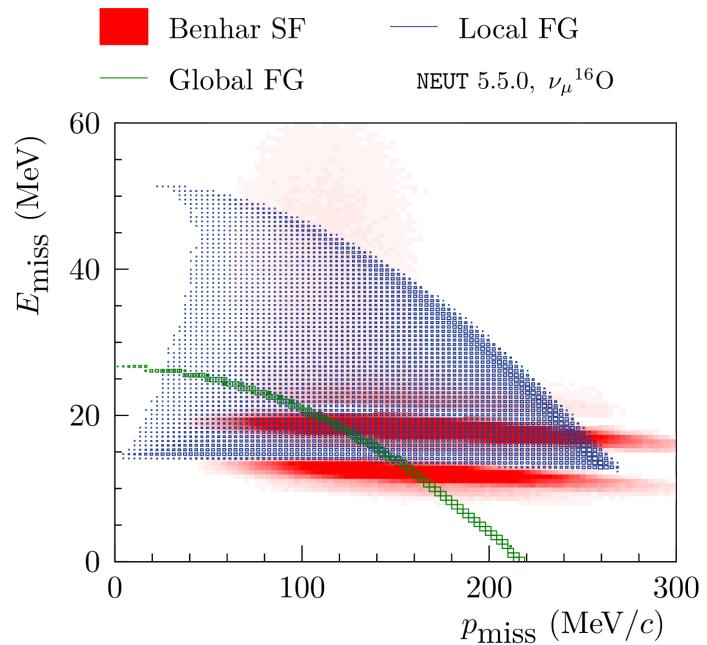
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The nuclear ground state

- Multiple ground state models are available in neutrino interaction generators
- RFG assumes the nucleus is a “box”-like potential and nucleons behave like a Fermi gas
 - Simplistic, widely used in the early days of neutrino oscillations, low predictive power

Eur.Phys.J.ST 230 (2021) 24, 4469-4481



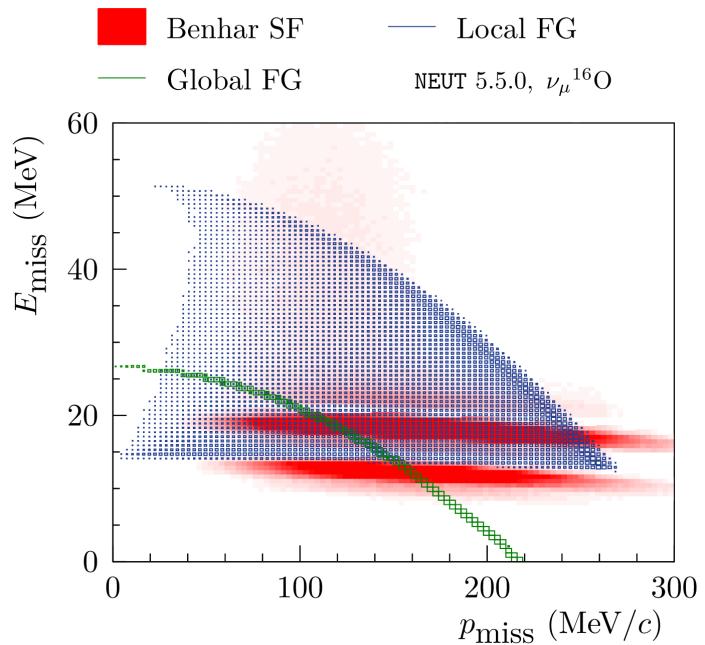
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 - Better predictive power than RFG, allows to explore other models

Eur.Phys.J.ST 230 (2021) 24, 4469-4481



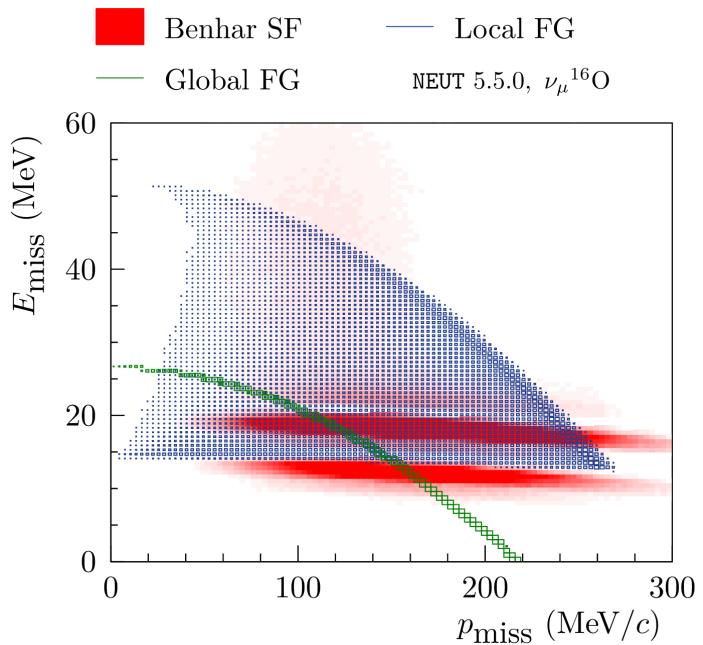
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 - Accounts for nuclear shell structure
 - Tuned to electron scattering data
 - Good predictive power for final state particles (but not perfect!)

Eur.Phys.J.ST 230 (2021) 24, 4469-4481



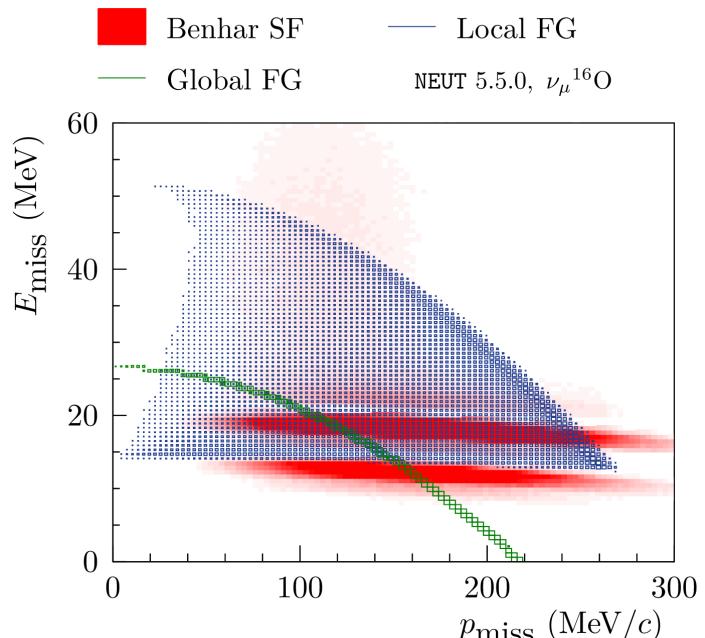
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- More sophisticated models exist but are not yet implemented in generators (**ongoing work**)

Eur.Phys.J.ST 230 (2021) 24, 4469-4481

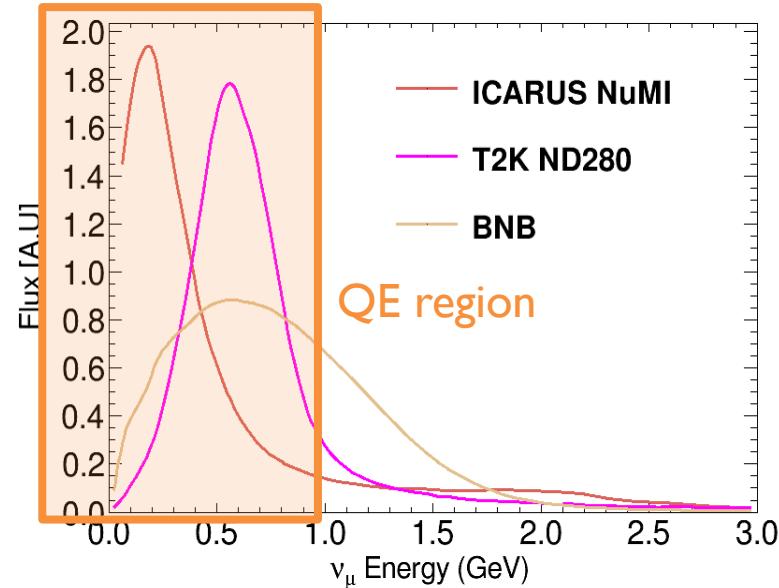


$$\vec{p}_{\text{miss}} = \vec{p}_v - \vec{p}_\mu - \vec{p}_p$$

$$E_{\text{miss}} = \omega - T_p^{\text{pre-FSI}} - \Delta m_{n \rightarrow p} - T_{\text{nuc. remnant}}$$

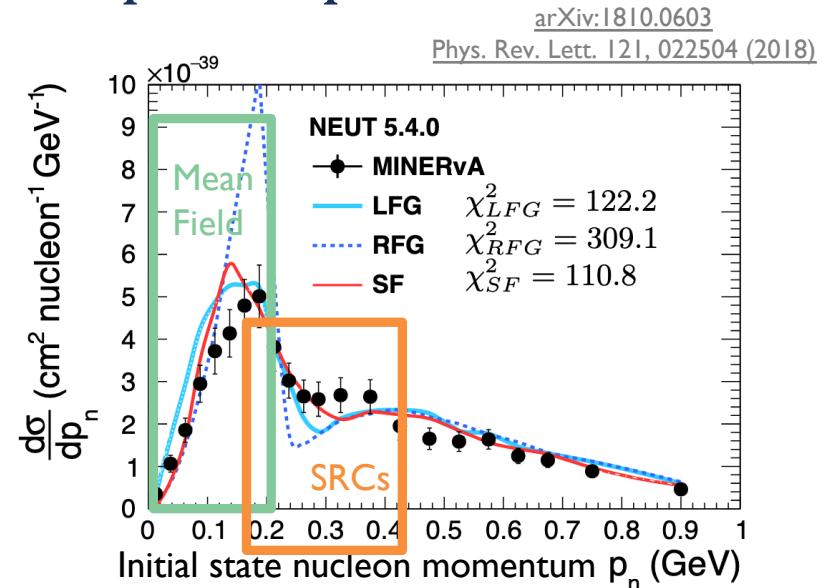
How does the nuclear ground state impact experiments?

- Most nuclear models predict removal energies which vary in a range of **10-50 MeV**
- Nucleon momenta span ranges of up to **hundreds of MeV**
- Mismodelling the removal energy causes a **direct bias** in reconstructed neutrino energy
 - Particularly important for oscillation measurements which evolve as a function of neutrino energy
 - But also for lower-energy physics (e.g. BSM physics)



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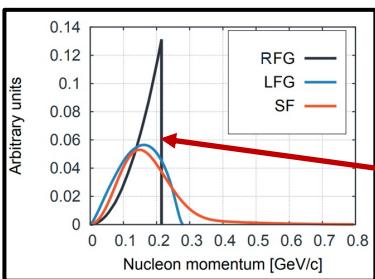
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 - Particularly important for oscillation measurements which evolve as a function of neutrino energy
 - But also for lower-energy physics (e.g. BSM physics)
- Mismodelling the nucleon momentum alters **predictions on hadron kinematics**
 - Will be crucial for experiments sensitive to the hadronic side of the final state



SRCs or “correlated tail” effects become important at nucleon momenta of $O(200\text{-}700 \text{ MeV})$

Example from T2K

2018



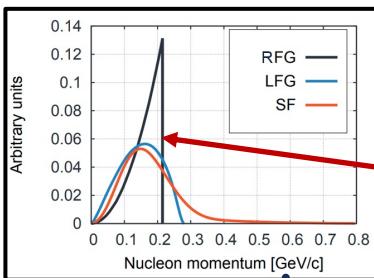
Type of Uncertainty	$\nu_e/\bar{\nu}_e$ Candidate Relative Uncertainty (%)
Super-K Detector Model	1.5
Pion Final State Interaction and Rescattering Model	1.6
Neutrino Production and Interaction Model Constrained by ND280 Data	2.7
Electron Neutrino and Antineutrino Interaction Model	3.0
Nucleon Removal Energy in Interaction Model	3.7
Modeling of Neutral Current Interactions with Single γ Production	1.5
Modeling of Other Neutral Current Interactions	0.2
Total Systematic Uncertainty	6.0

The systematic uncertainty on the predicted relative number of electron neutrino and electron antineutrino candidates in the SK samples with no decay electrons.

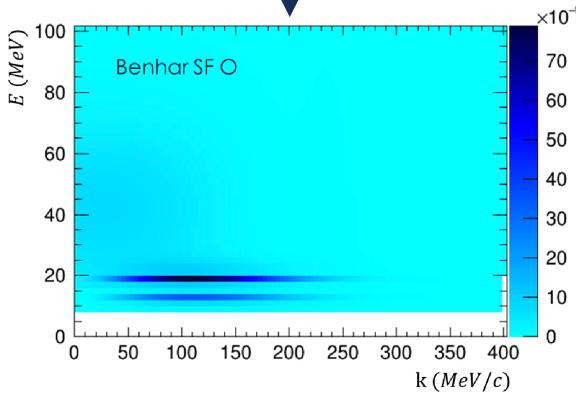
[Nature](#) volume 580, pages 339–344 (2020)

Example from T2K

2018



Change in
nuclear model



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Nature volume 580, pages 339–344 (2020) The 50th Anniversary



Now incorporated here

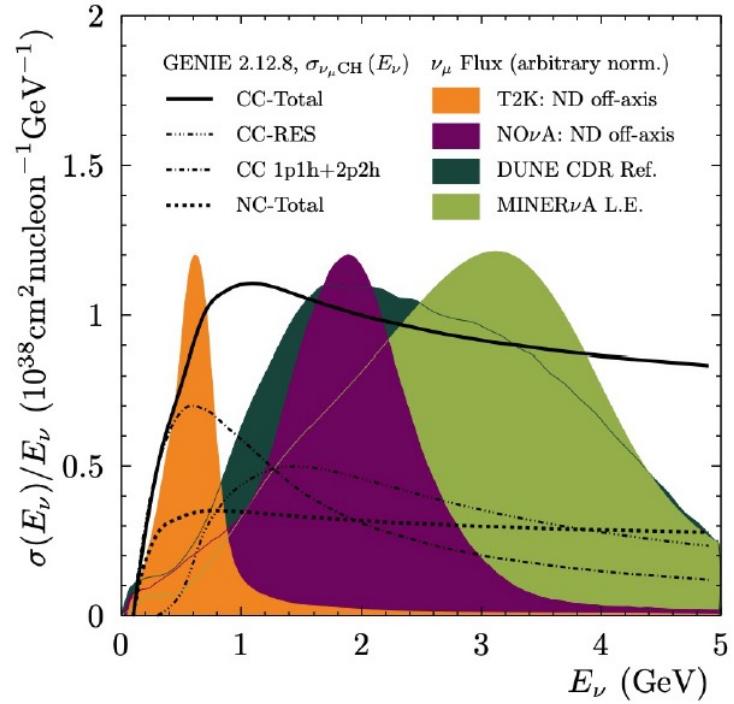
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Total All	3.4	3.9	4.9		5.2	5.8	14.3

T2K Run 1-10, preliminary

59

Ongoing work on DUNE

- DUNE is preparing oscillation sensitivity studies focused on its **near detectors**
- At DUNE statistics, we will be sensitive to some of **the most minute differences** in nuclear models
- DUNE doesn't have data yet
 - Need a **flexible model** which allows us to cover the variations suggested by other models and measurements
 - Need a model which is adapted for both oscillation studies and BSM physics



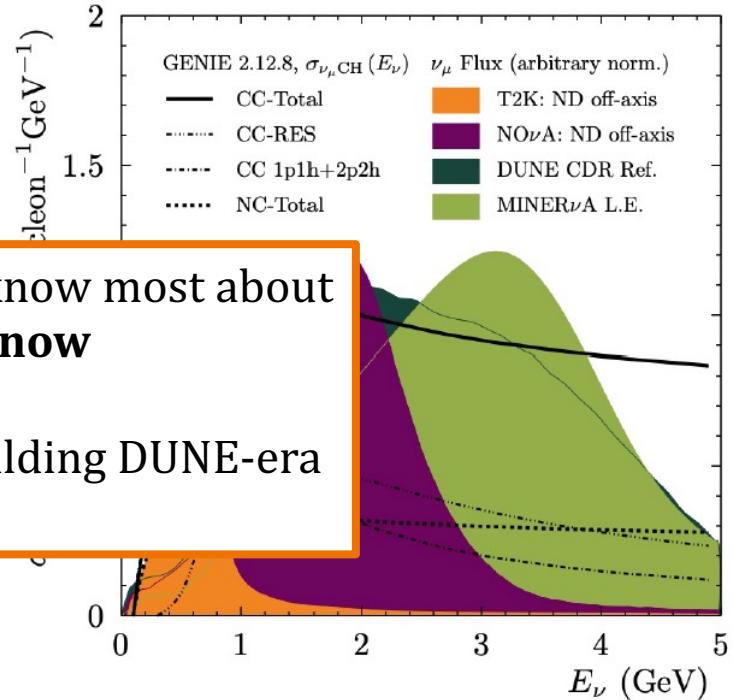
DUNE is not QE dominated
But QE interactions are **those which we can measure best**
Remain important for **precision measurements**

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The QE regime is where we know most about
what we don't know

Perfect starting point for building DUNE-era
analyses

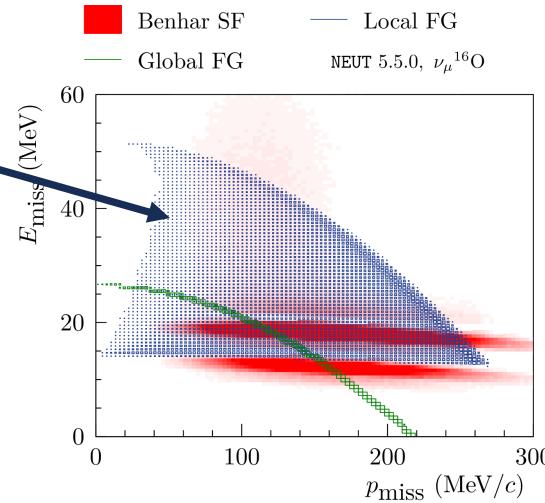


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The base model

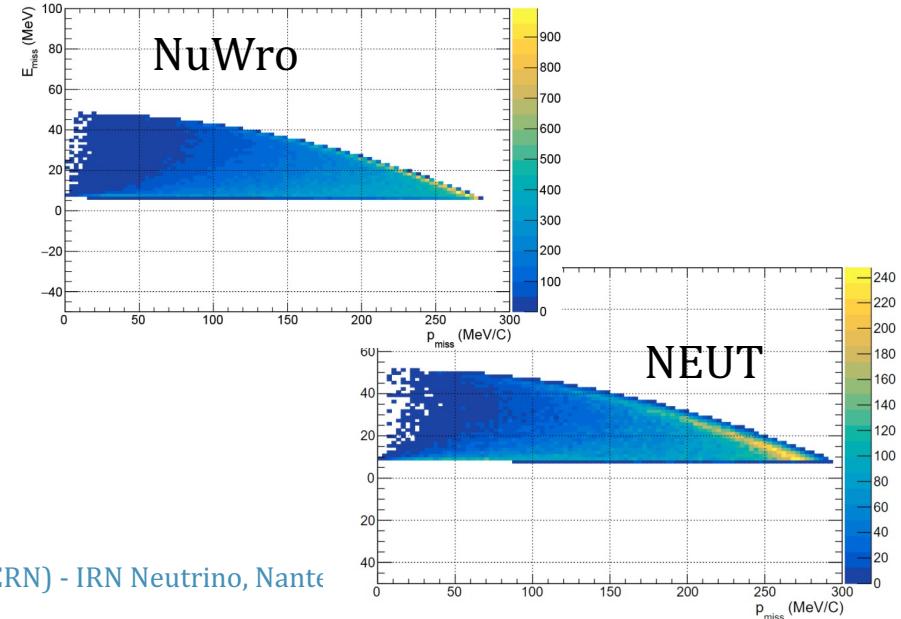
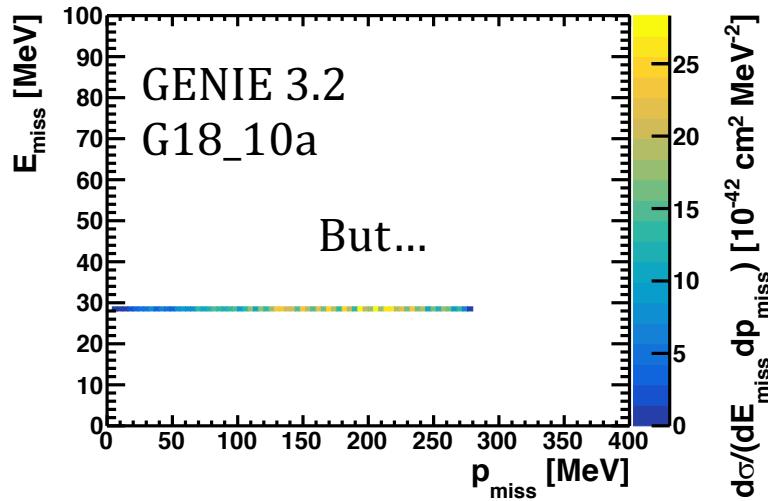
- The DUNE production framework is optimized to use the GENIE event generator
- GENIE does not have a 2D spectral function model yet
- But want to take inspiration from what we've learned from T2K
 - Choose an LFG model!

Can be reweighted to a
spectral function!



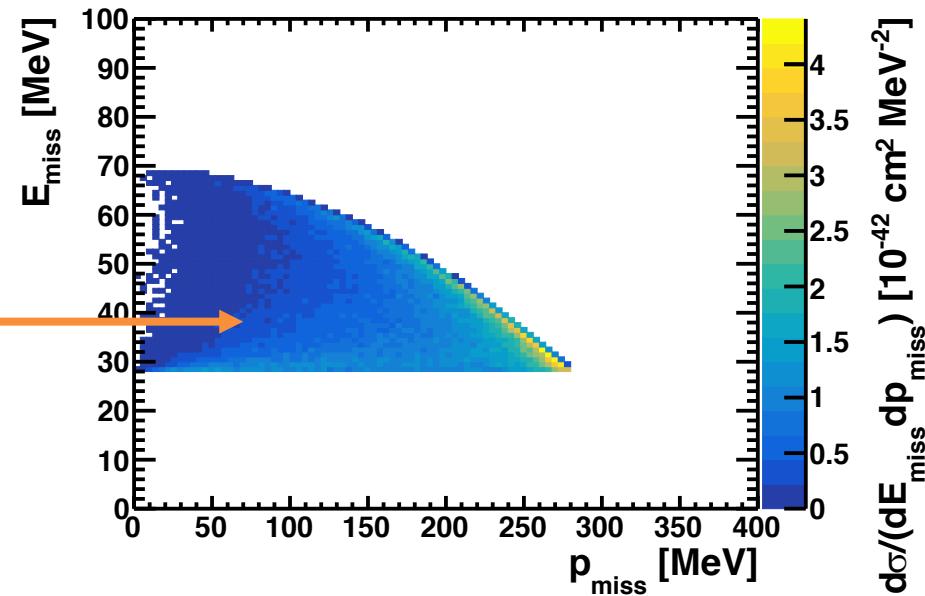
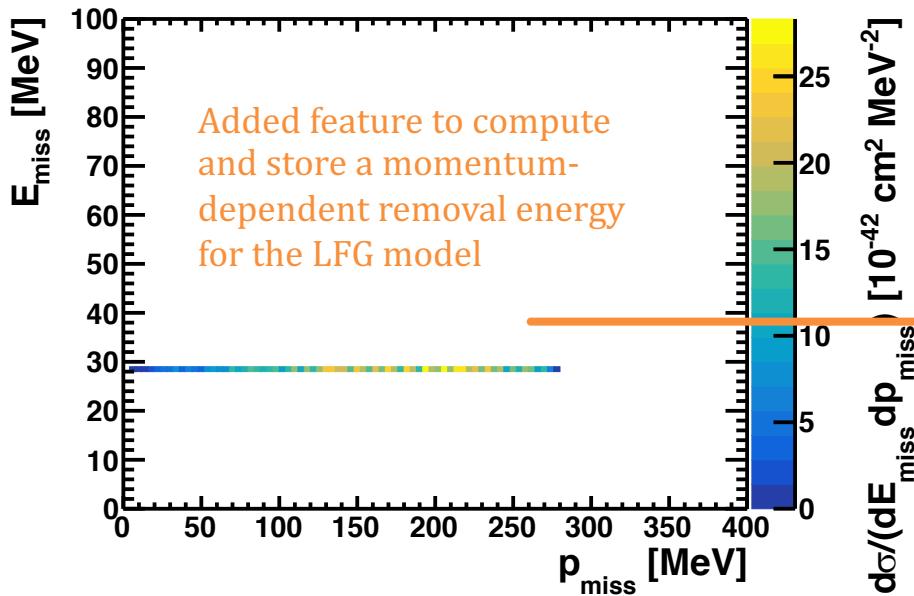
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Expanding the ground state phase space

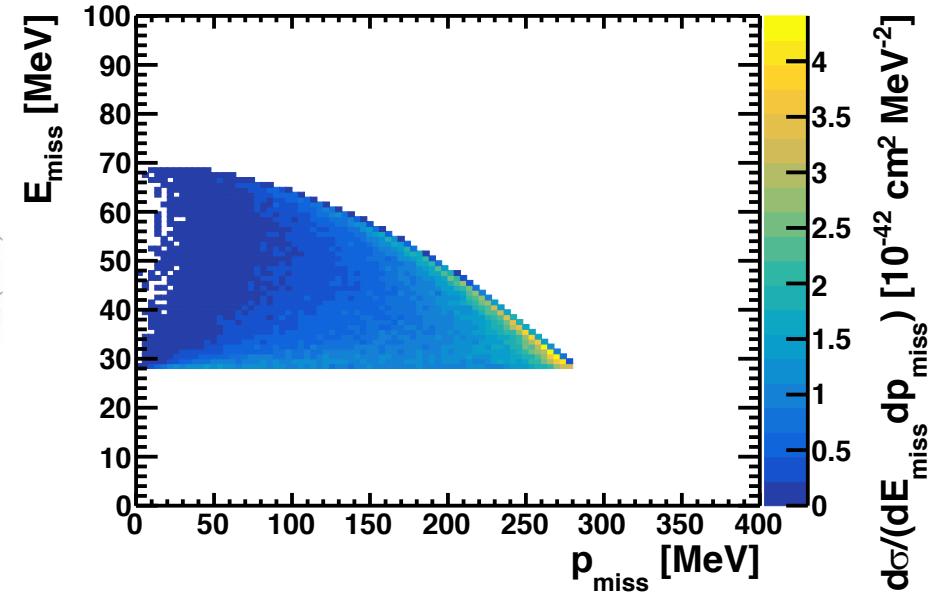
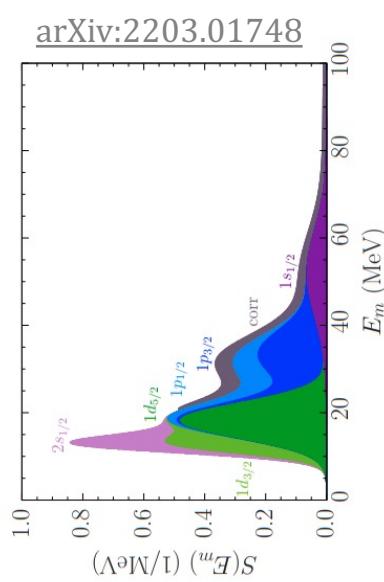
- For an LFG model, the dependence between the removal energy and the nucleon momentum is well-understood (Fermi gas approach)



Expanding the ground state phase space

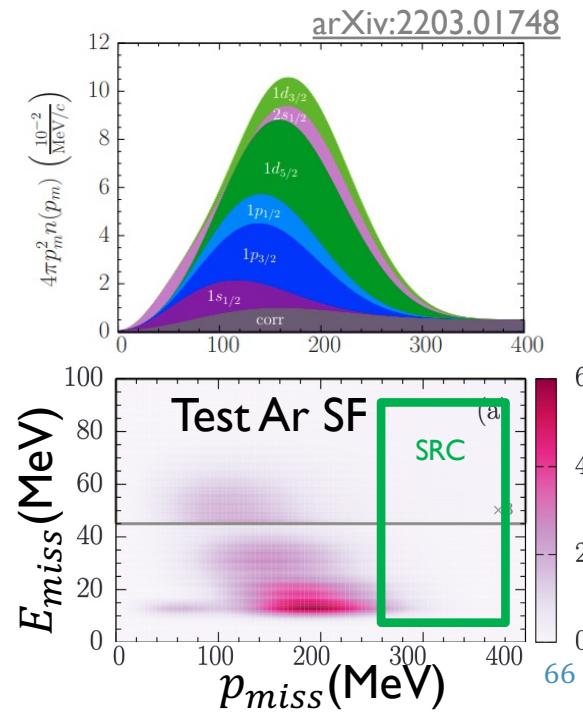
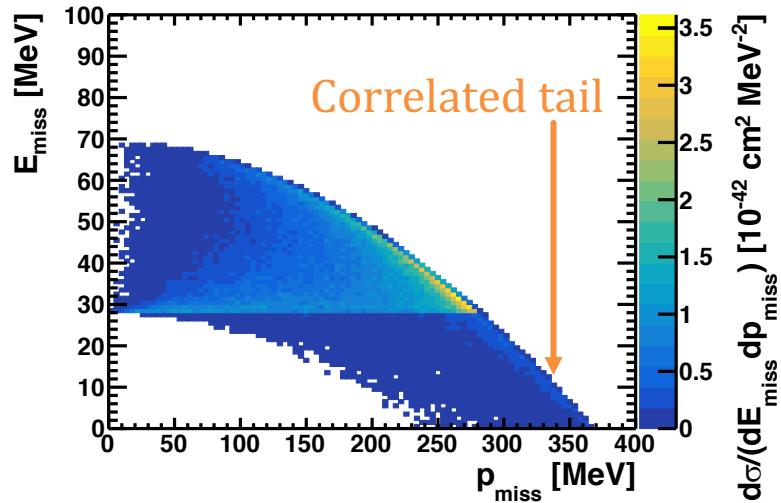
- For an LFG model, the dependence between the removal energy and the nucleon momentum is well-understood (Fermi gas approach)

Covers relevant energies
for Argon nuclear shells!



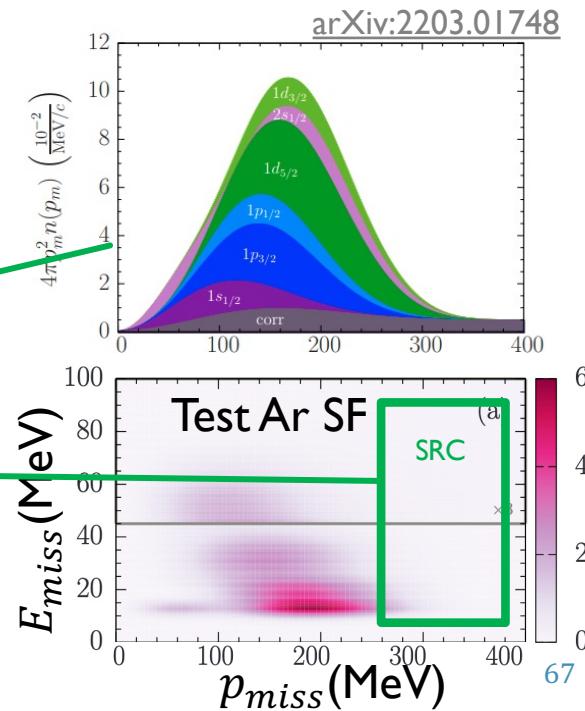
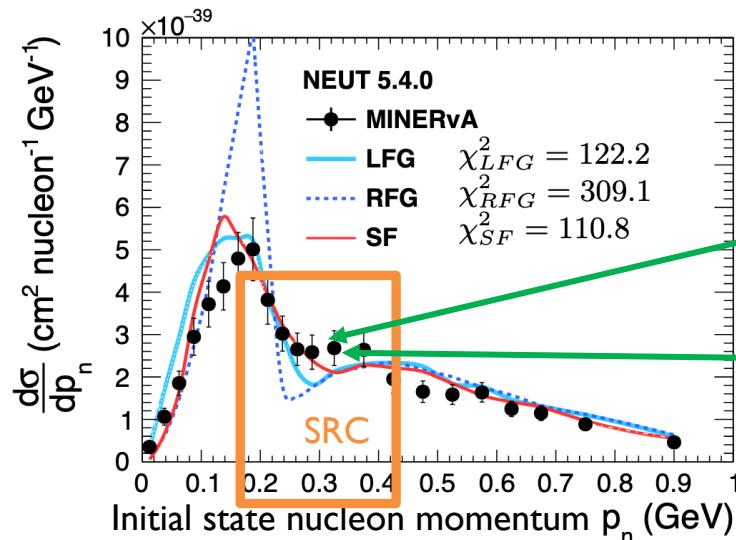
Expanding the ground state phase space

- Will be used in conjunction with “correlated tail” feature



Expanding the ground state phase space

- Will be used in conjunction with “correlated tail” feature
- Probes nucleon correlations – high model disagreement



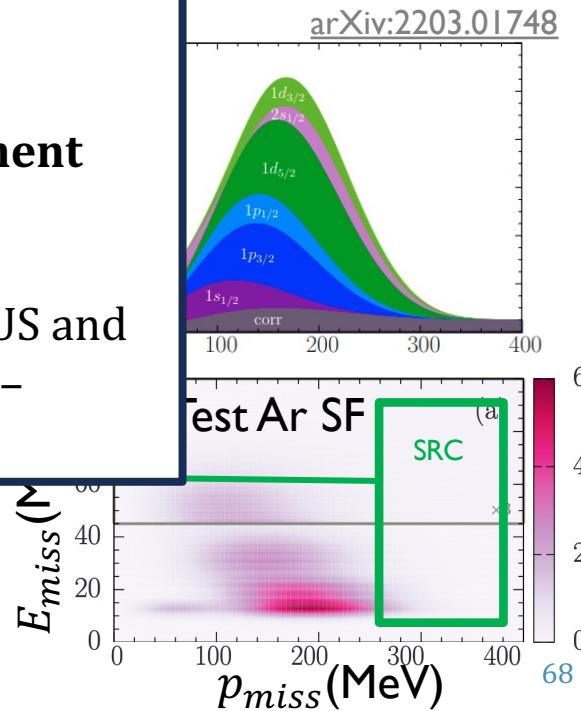
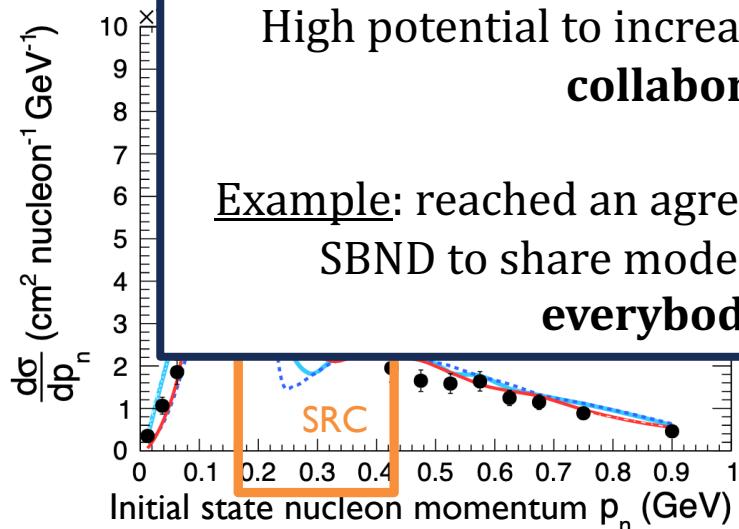
Expanding the ground state phase space

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- Probes nu

This work is as relevant for DUNE as it is for
Hyper-K and SBN experiments

High potential to increase **cross-experiment collaboration!**

Example: reached an agreement with ICARUS and SBND to share model and systematics – **everybody wins!**



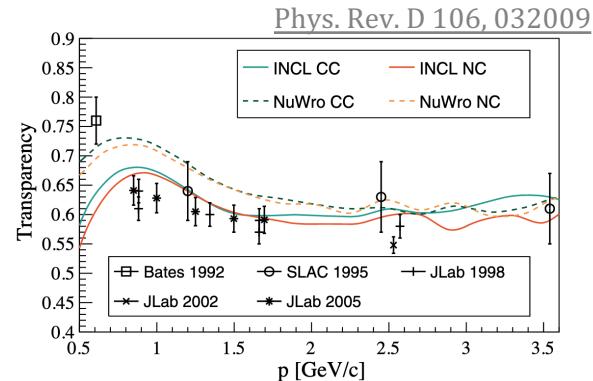
Beyond the nuclear ground state

- CCQE(-like) interactions are easiest to study experimentally
 - Main channel for Hyper-K and future SBN experiments (ICARUS, SBND)
 - And provide a gateway towards understanding more complex processes relevant at DUNE energies
- But the work is just getting started! We need to understand:

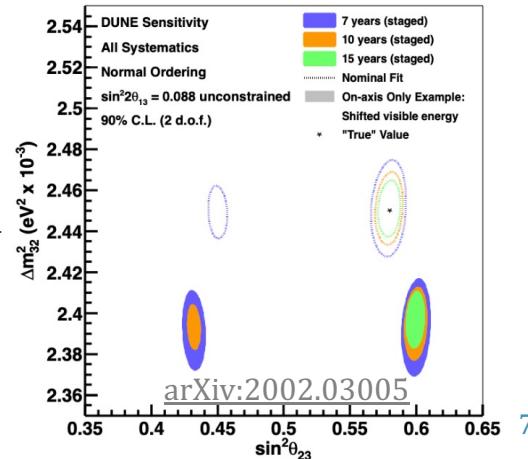
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Bias in oscillation parameters if
20% of proton energy were
carried away by neutrons
FSI-like effect

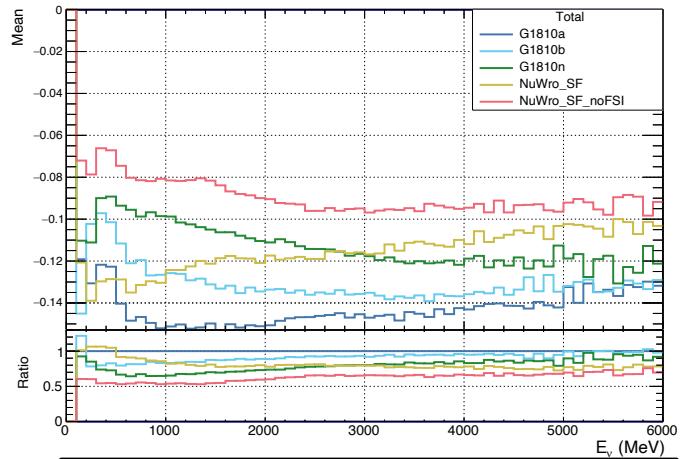


For Carbon: 40% of protons
re-interact inside the nucleus



Beyond the nuclear ground state

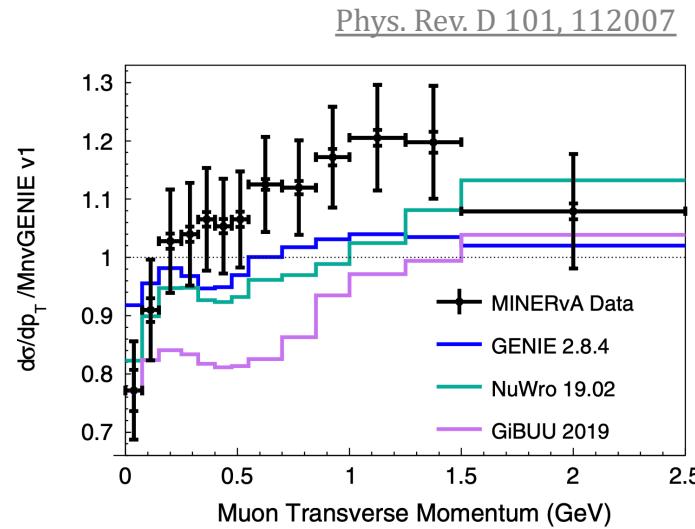
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 - Cross-section energy dependence



Mean of neutrino energy bias as a
function of neutrino energy for DUNE flux
High model disagreement

Beyond the nuclear ground state

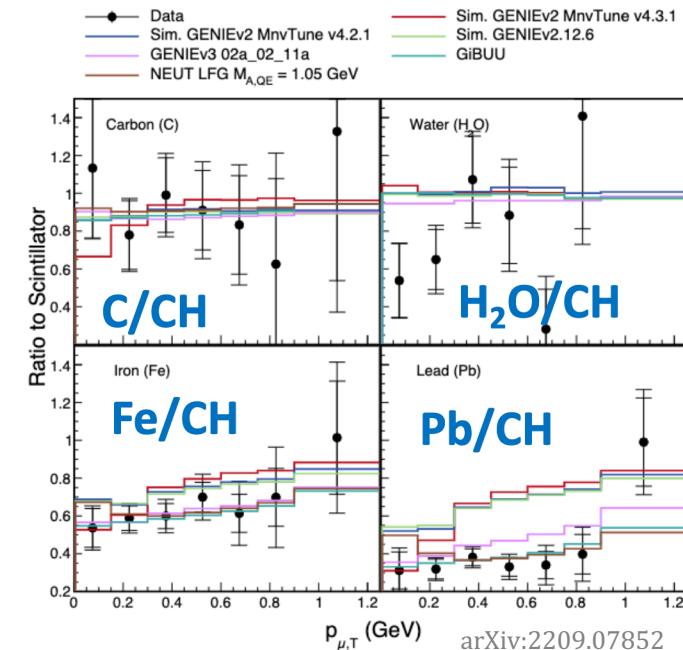
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 - Cross-section energy dependence
 - Resonant interactions and the deep inelastic regime



DIS-enriched measurement from MINERvA
High model disagreement

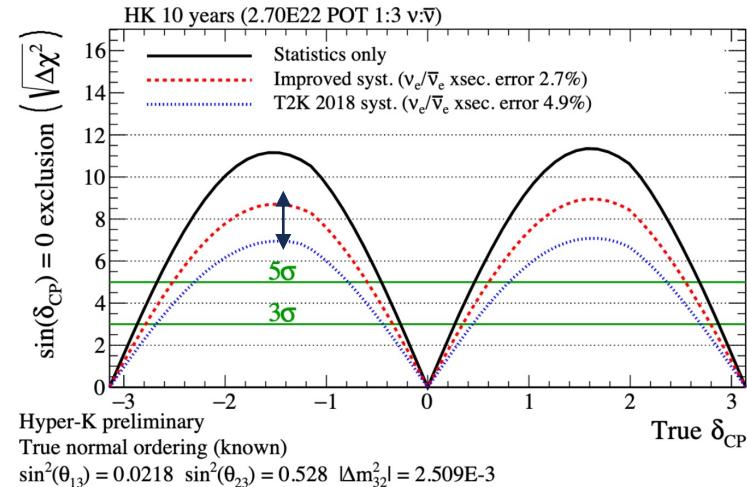
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 - How nuclear effects vary with atomic number



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 - How nuclear effects vary with atomic number
 - Differences in ν_e vs ν_μ cross-sections (**crucial for δ_{CP} !**)*



Majority of the systematic error contribution

*Leading uncertainties from nuclear ground state model!

The T2K ND280 Upgrade project

[ND280 Upgrade TDR](#)

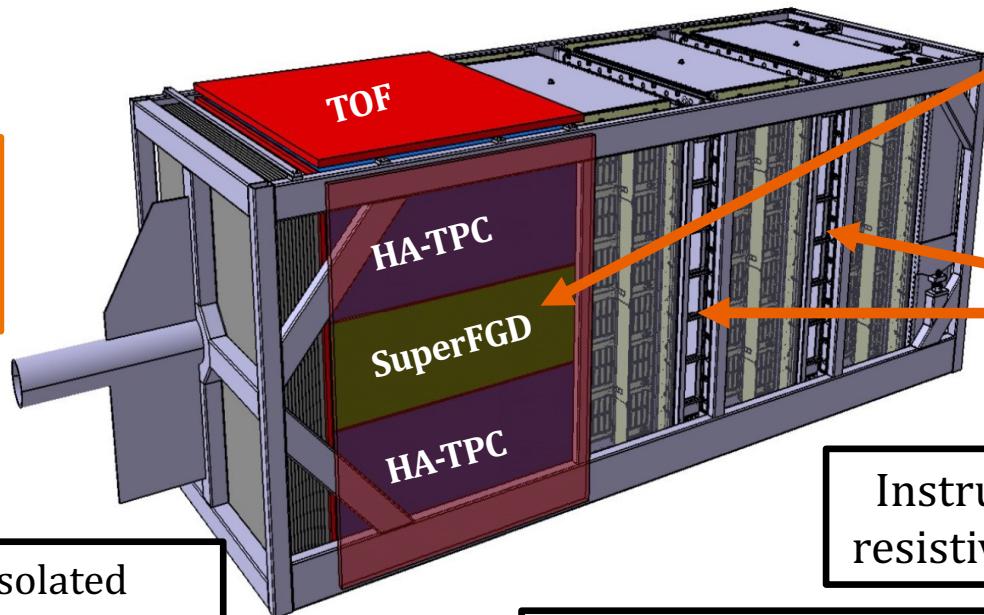
Replace P0D by new suite of detectors

>100 researchers
22 institutes
7 countries



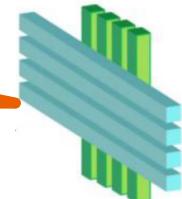
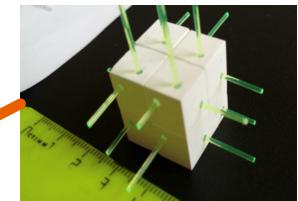
2M 1cm³ optically isolated scintillator cubes

2x fiducial mass of current FGDs



Instrumented with resistive Micromegas

150 ps timing resolution for PID

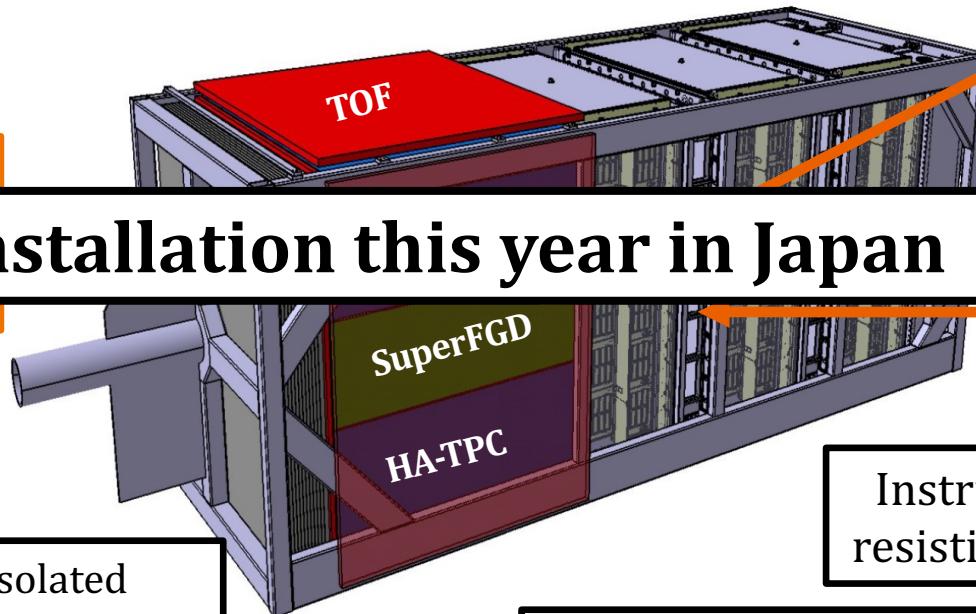


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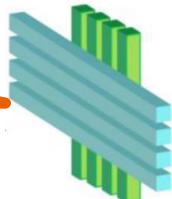
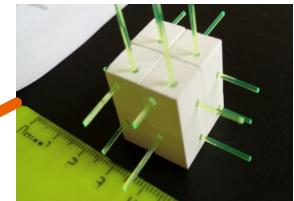


2M 1cm³ optically isolated scintillator cubes

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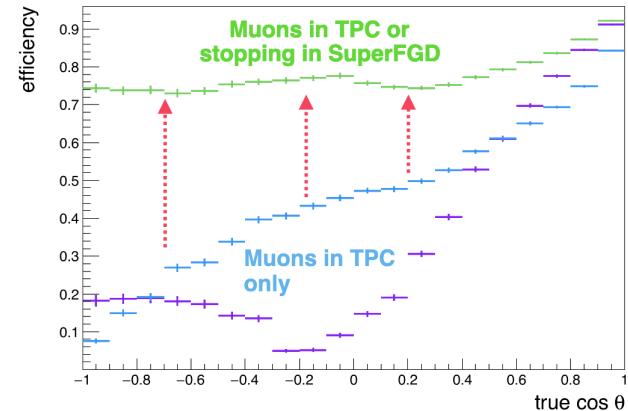
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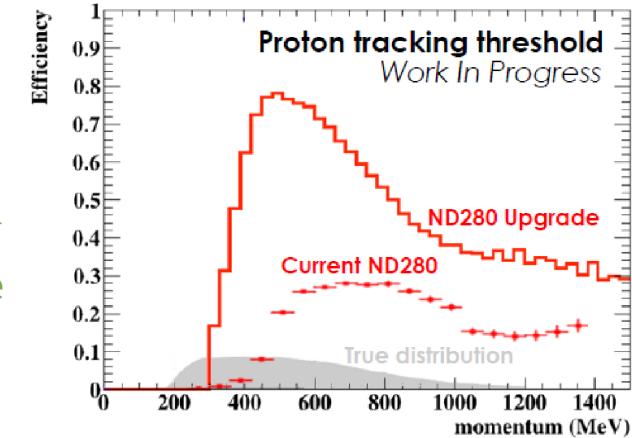
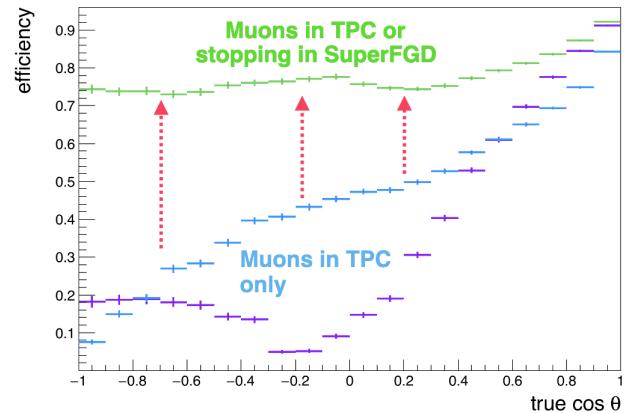
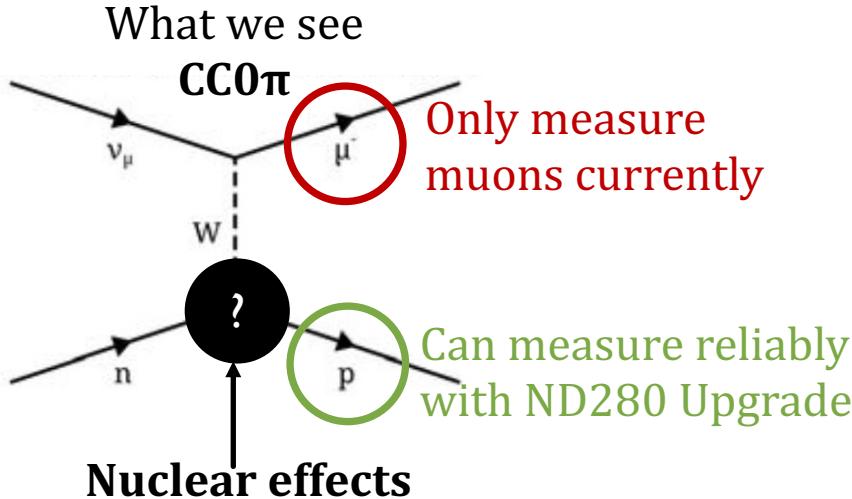
Capabilities of the ND280 Upgrade

- Full 4π angular coverage (same as SK)



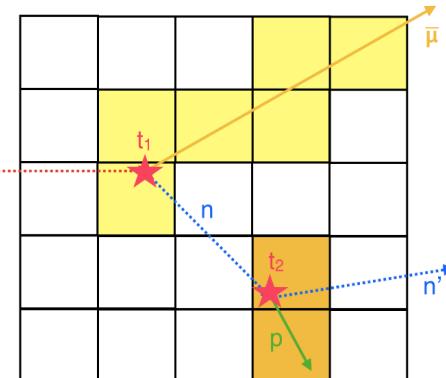
Capabilities of the ND280 Upgrade

- Full 4π angular coverage (same as SK)
- Low momentum thresholds for detecting **protons**
 - Can reliably measure full final state in neutrino interactions!



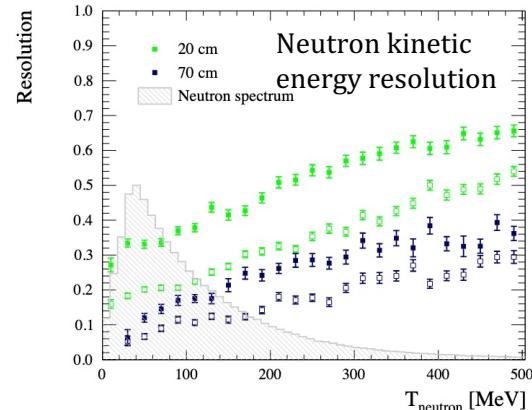
Capabilities of the ND280 Upgrade

- Full 4π angular coverage (same as SK)
- Low momentum thresholds for detecting **protons**
 - Can reliably measure full final state in neutrino interactions!
- Capability to measure (not just tag!) **neutrons**
 - Unmatched by Liquid Argon TPCs

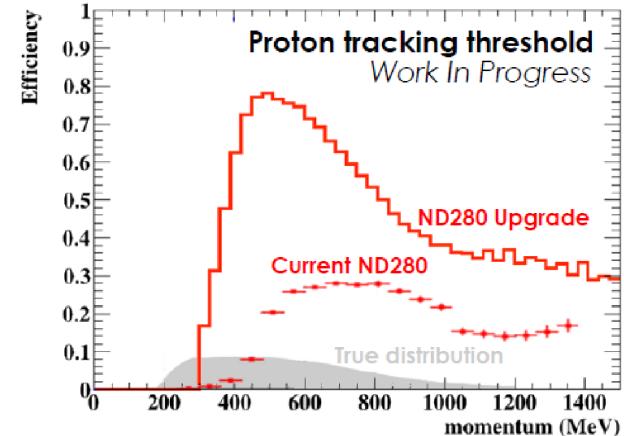
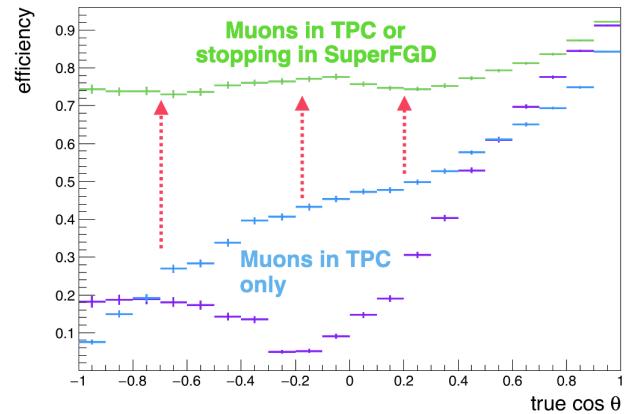


Phys. Rev. D 101, 092003

20.06.2023



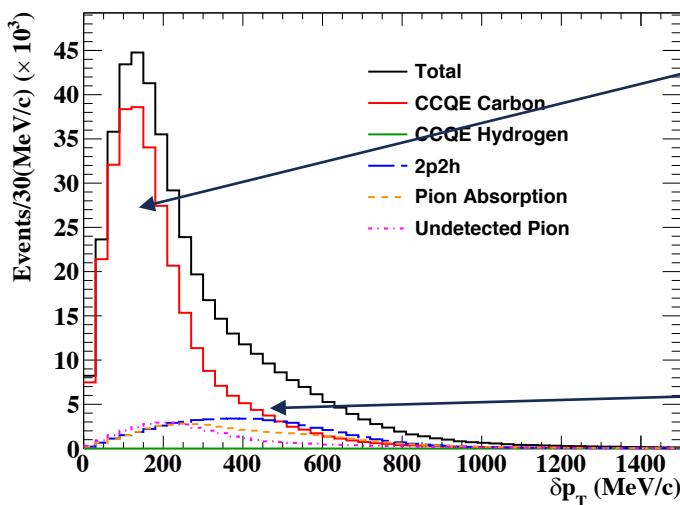
Laura Munteanu (CERN) - IRN Neutrino, Nantes



79

What can the ND280 Upgrade do for us?

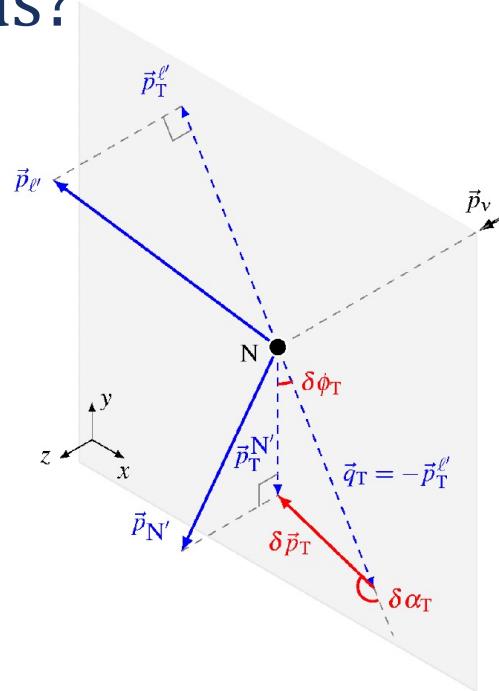
- Access to **exclusive** variables
 - E.g. probe nuclear effects by looking at Transverse Kinematic Imbalance (TKI)



Bulk dominated
by CCQE

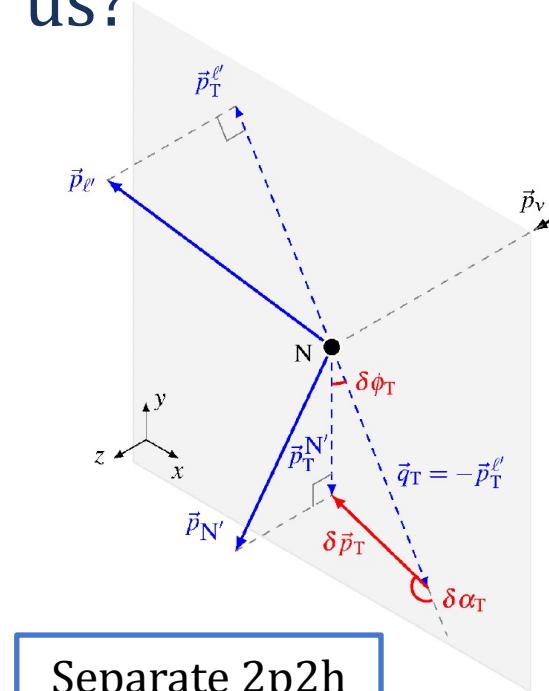
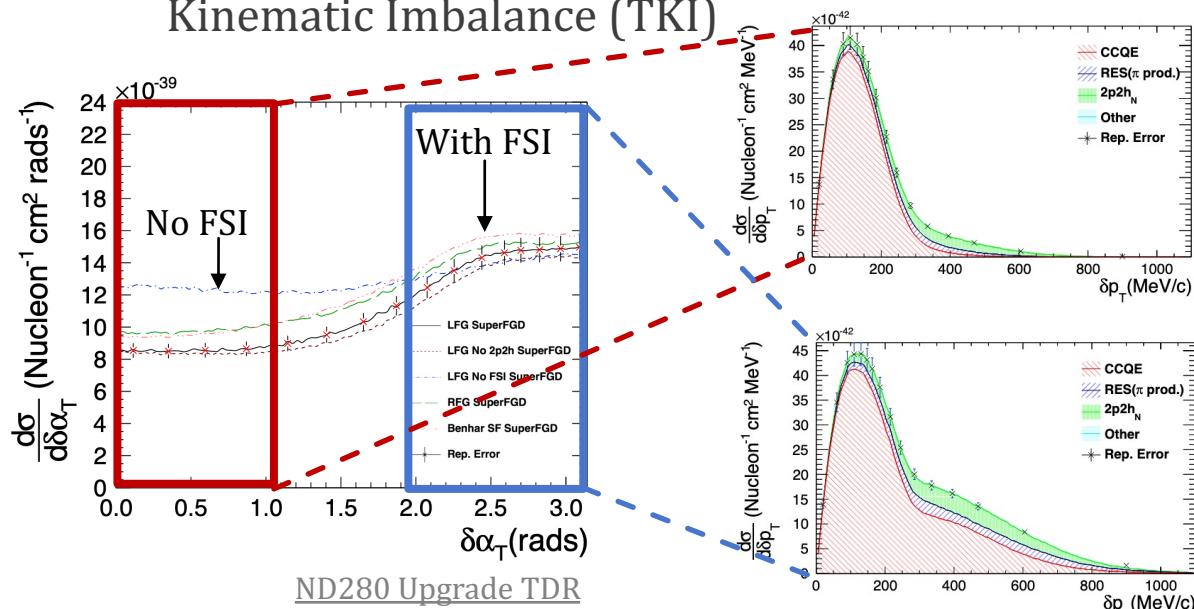
Tails dominated
by FSI+2p2h

Phys. Rev. D 105, 032010



What can the ND280 Upgrade do for us?

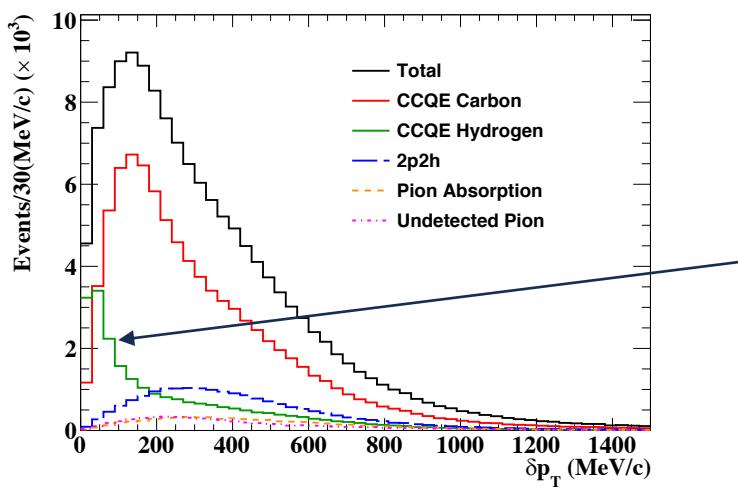
- Access to **exclusive** variables
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Separate 2p2h
and FSI effects!

What can the ND280 Upgrade do for us?

- Access to **exclusive** variables
 - E.g. probe nuclear effects by looking at Transverse Kinematic Imbalance (TKI)



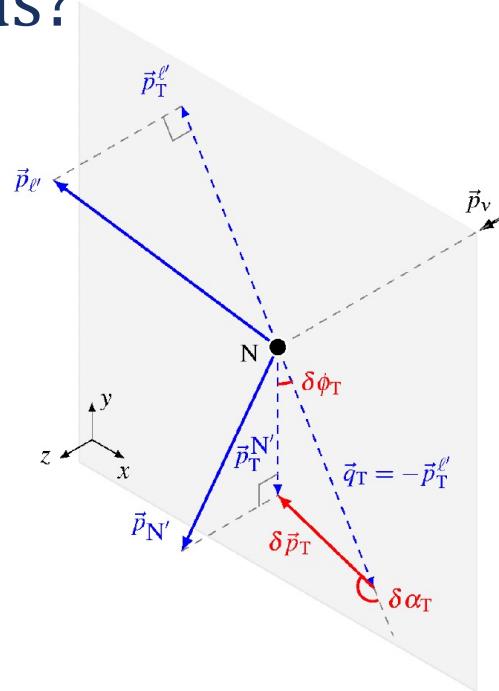
Phys. Rev. D 105, 032010

Phys. Rev. D 101, 092003

Antineutrinos:
Peak from interactions
on hydrogen

No nuclear effects

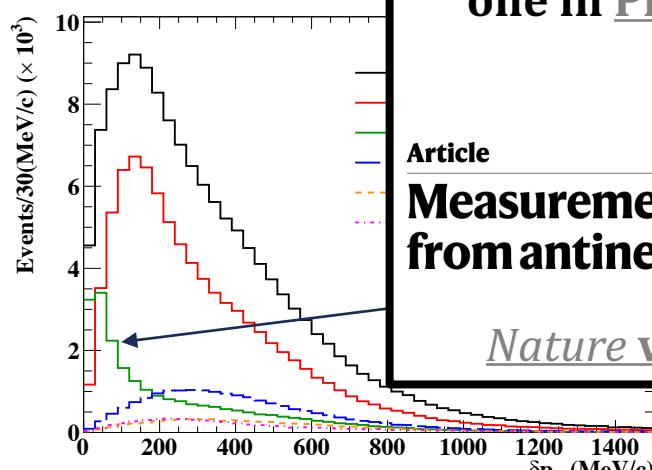
Possible thanks to
neutron detection!



What can the ND280 Upgrade do for us?

- Access to **exclusive** variables

- E.g. probe neutrino Kinematic Invariance



A recent measurement from the MINERVA collaboration using **a method inspired by the one in Phys. Rev. D 101, 092003** has been published in Nature

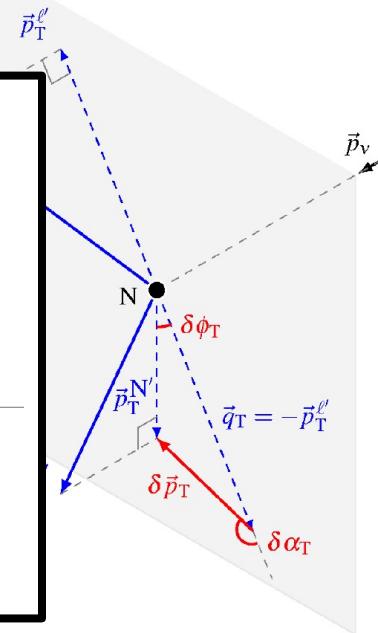
Article
Measurement of the axial vector form factor from antineutrino–proton scattering

Nature volume 614, pages 48–53 (2023)

Possible thanks to
neutron detection!

Phys. Rev. D 105, 032010

Phys. Rev. D 101, 092003

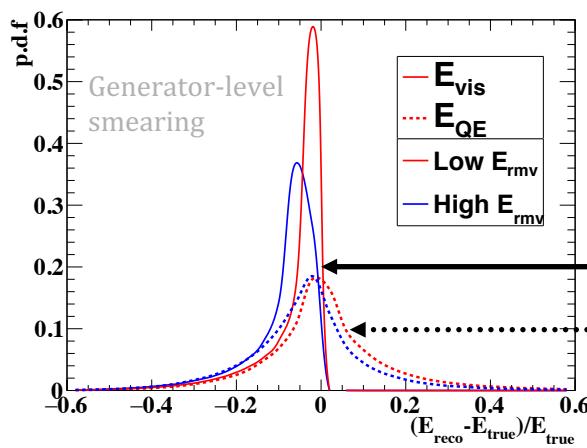


What can the ND280 Upgrade do for us?

- Access to **exclusive** variables
 - E.g. probe nuclear effects by looking at Transverse Kinematic Imbalance (TKI)
 - Using calorimetric estimators for neutrino energy

$$E_{vis} = E_\mu + T_N$$

Method used by NOvA & DUNE

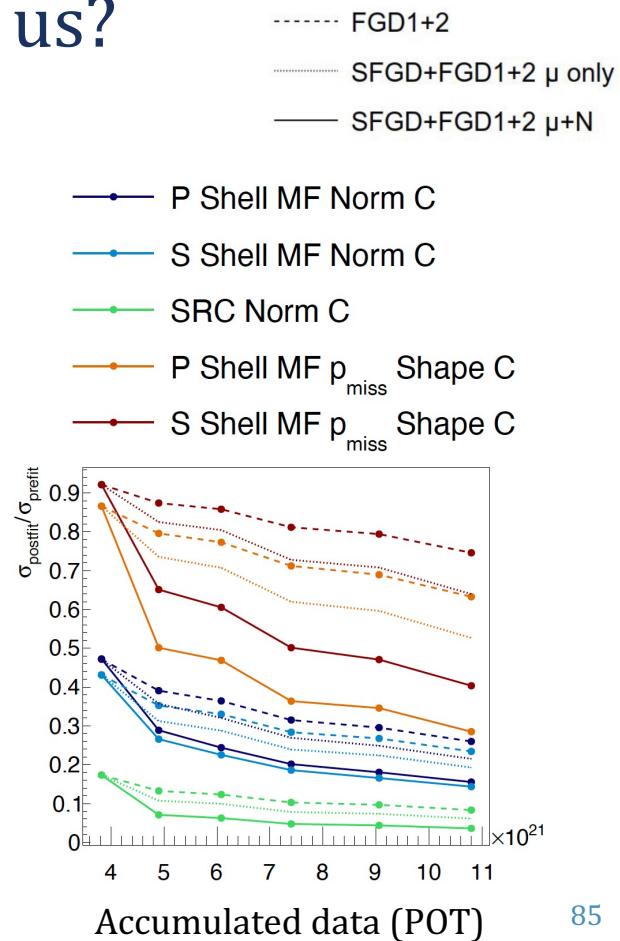


Reduced bias in neutrino energy reconstruction

E_{vis} bias dominated by **detector resolution**
 E_{QE} dominated by **nuclear effects**

What can the ND280 Upgrade do for us?

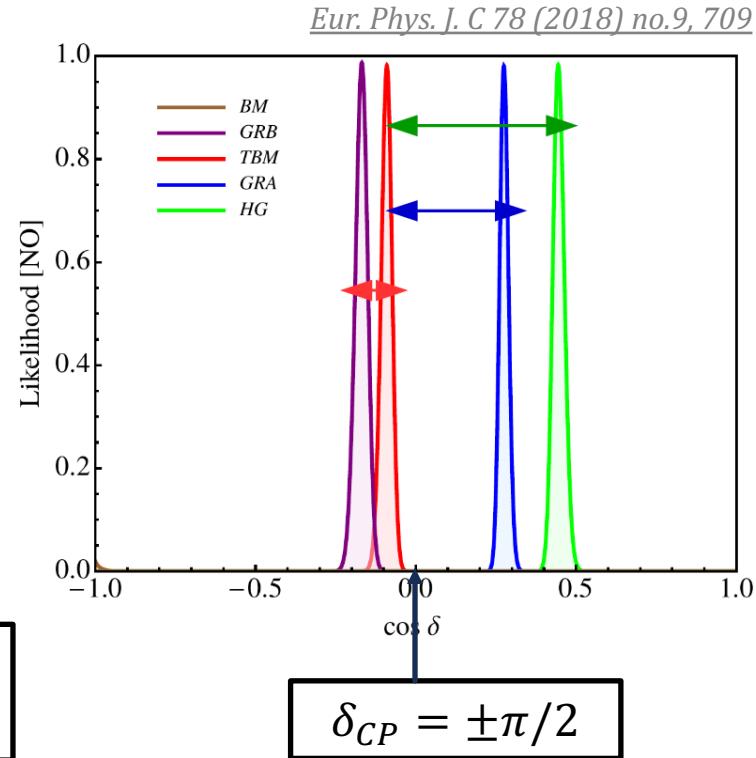
- Access to **exclusive** variables
 - E.g. probe nuclear effects by looking at Transverse Kinematic Imbalance (TKI)
 - Using calorimetric estimators for neutrino energy
- Significantly **improve constraints** on systematic error parameters for oscillation analyses
- Help build more **robust** analyses
- Give us **novel measurements!**



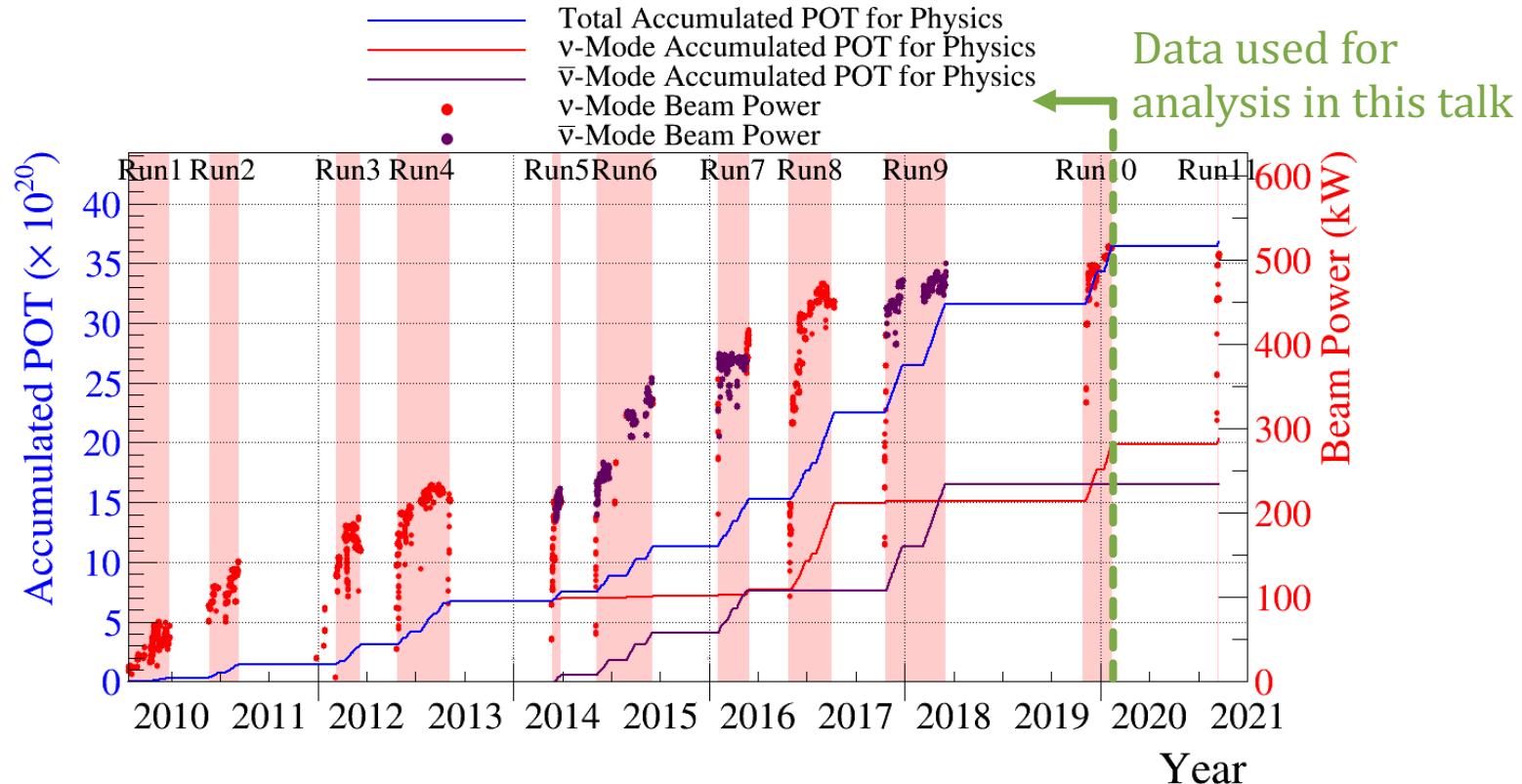
Precision measurements of δ_{CP}

- Measuring CP violation (CPV) \neq measuring δ_{CP}
- If no CPV in neutrino oscillations: exclude matter-antimatter asymmetry explanation through lepton sector CPV
- If CPV is observed: we can **constrain leptogenesis models**
 - But need **precise** measurement of δ_{CP} !
- Model separation power vs δ_{CP} resolution:
 - **Satisfactory:** $\sigma(\delta_{CP}) < 30^\circ$
 - **Good:** $\sigma(\delta_{CP}) < 23^\circ$
 - **Excellent:** $\sigma(\delta_{CP}) < 5^\circ$

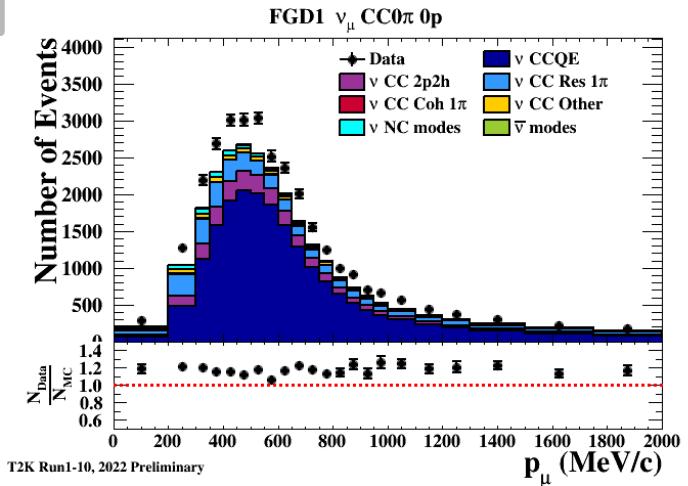
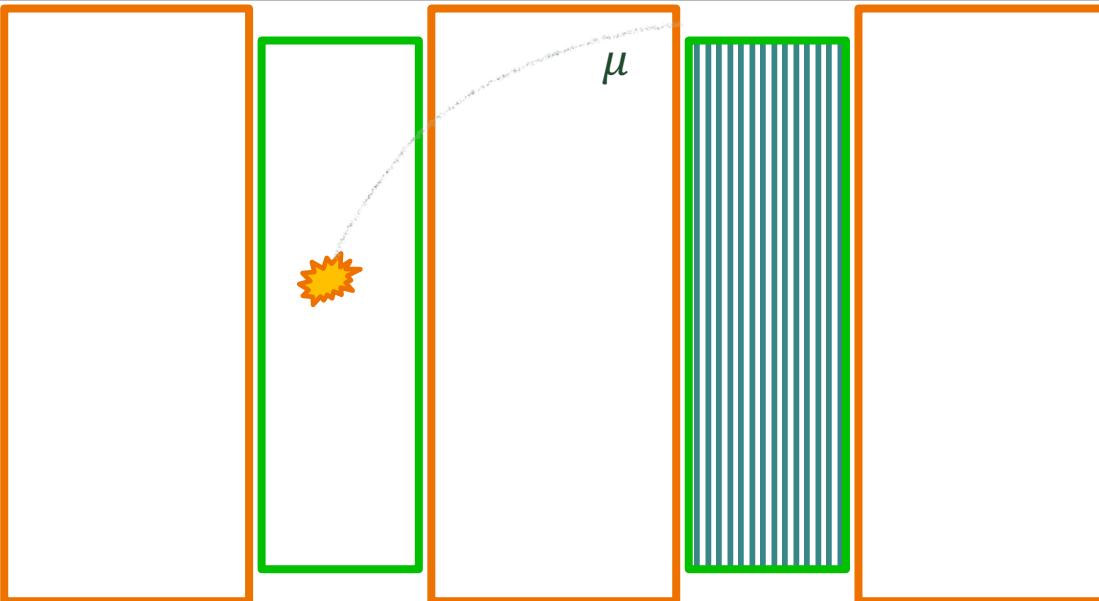
Hyper-K and
DUNE target



Collected T2K data

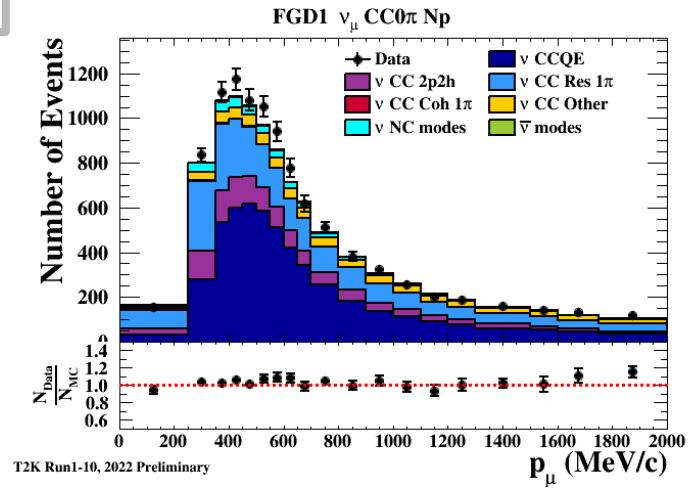
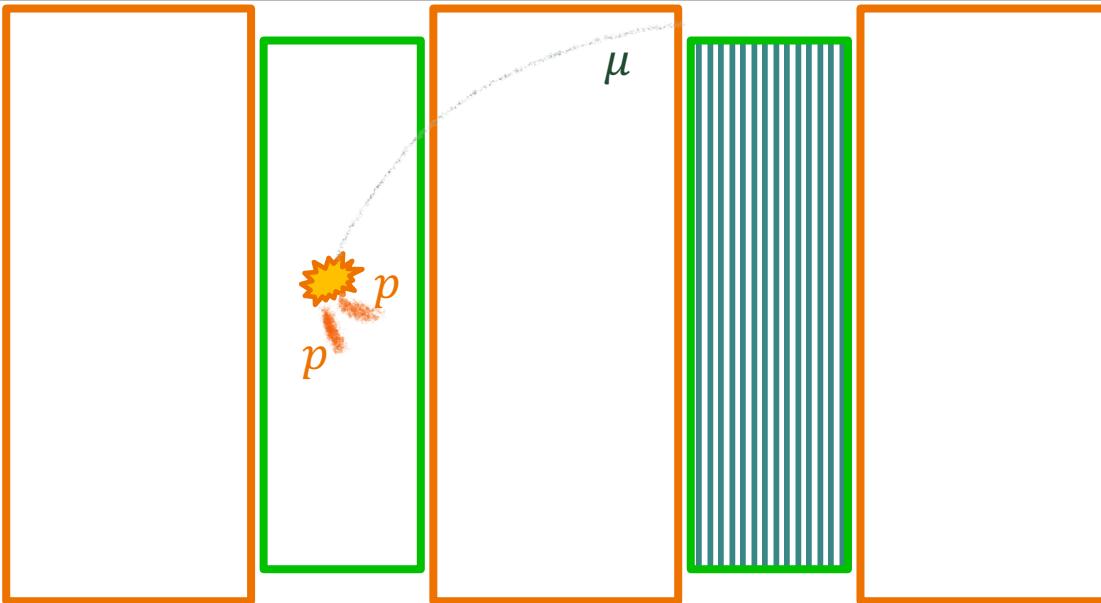


Samples: CC $0\pi 0p$



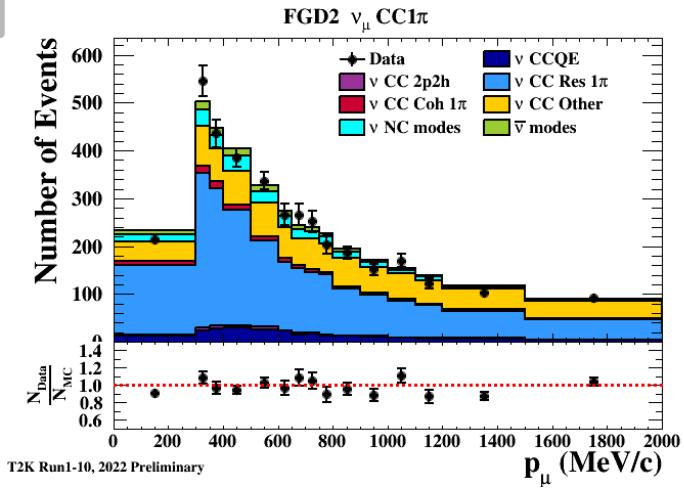
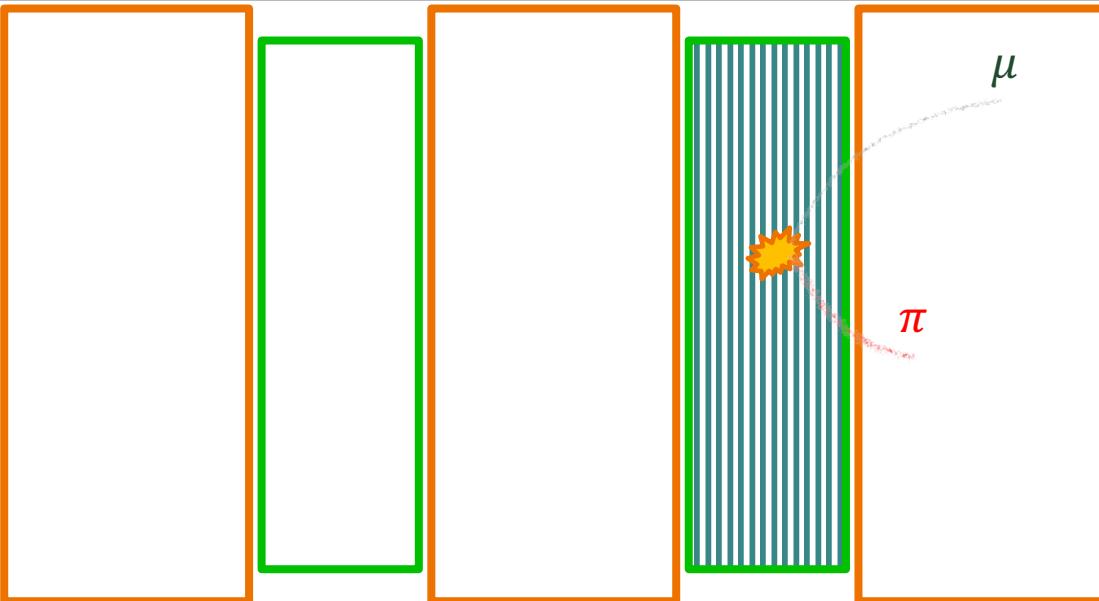
$1\mu, 0\pi, 0p$
Enriched in CCQE interactions

Samples: CC0 π Np



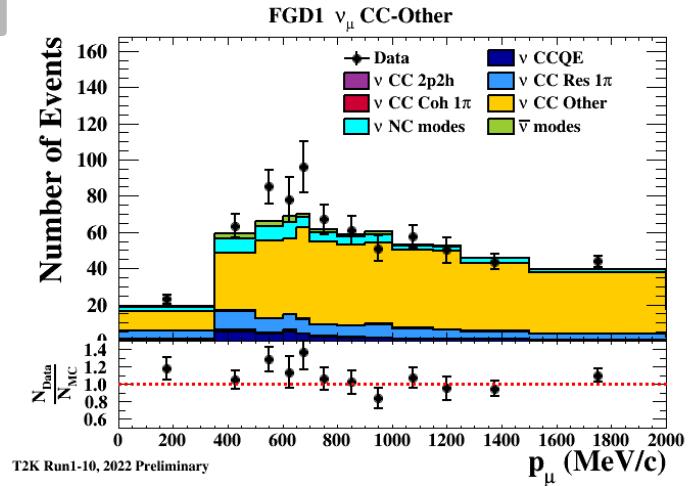
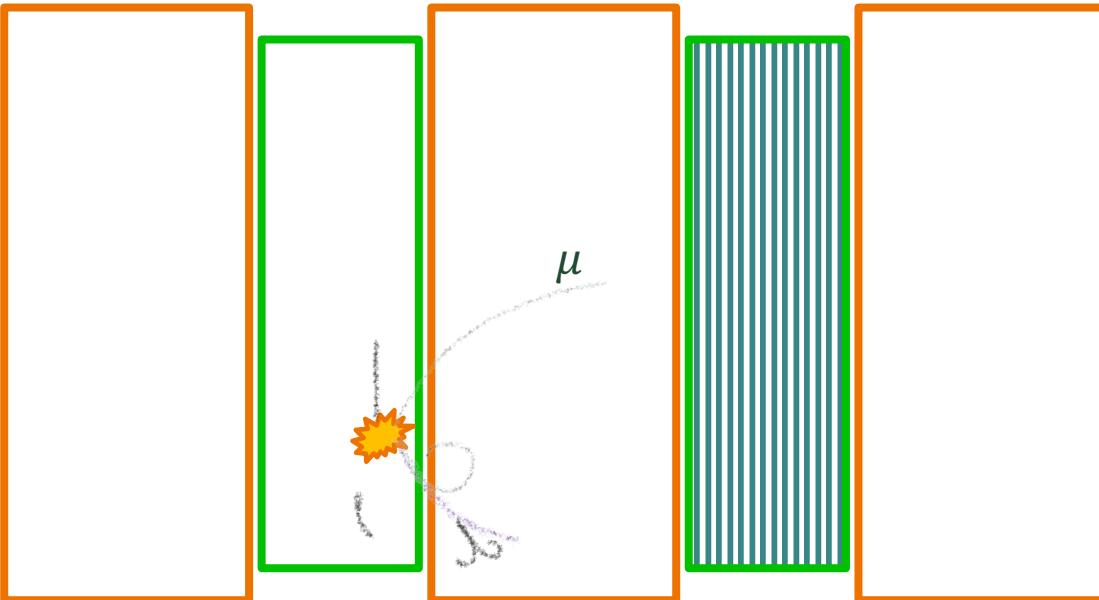
$1\mu, 0\pi, Np$
Constrains FSI-related
background

Samples: CC1 π



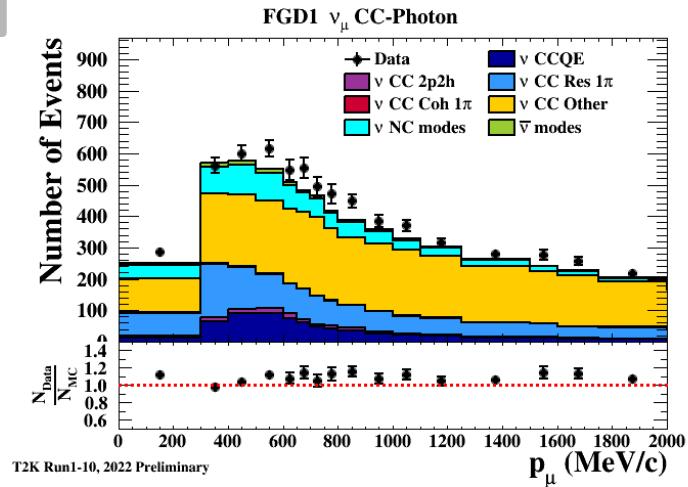
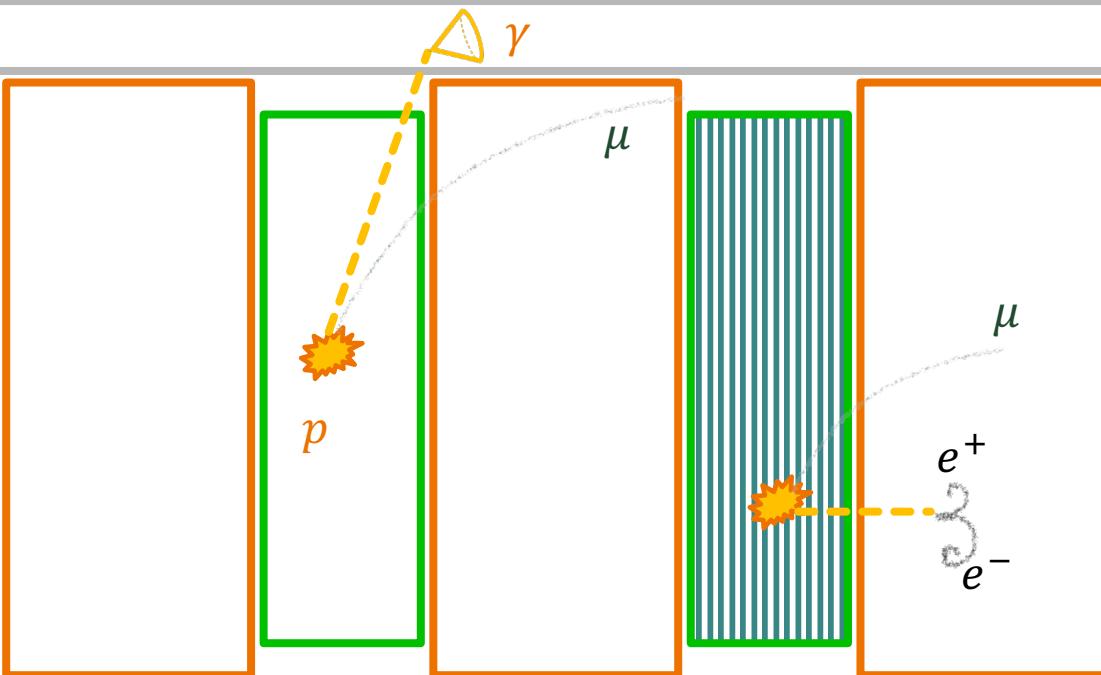
$1\mu, 1\pi$
Enriched in resonant
interactions

Samples: CCOther



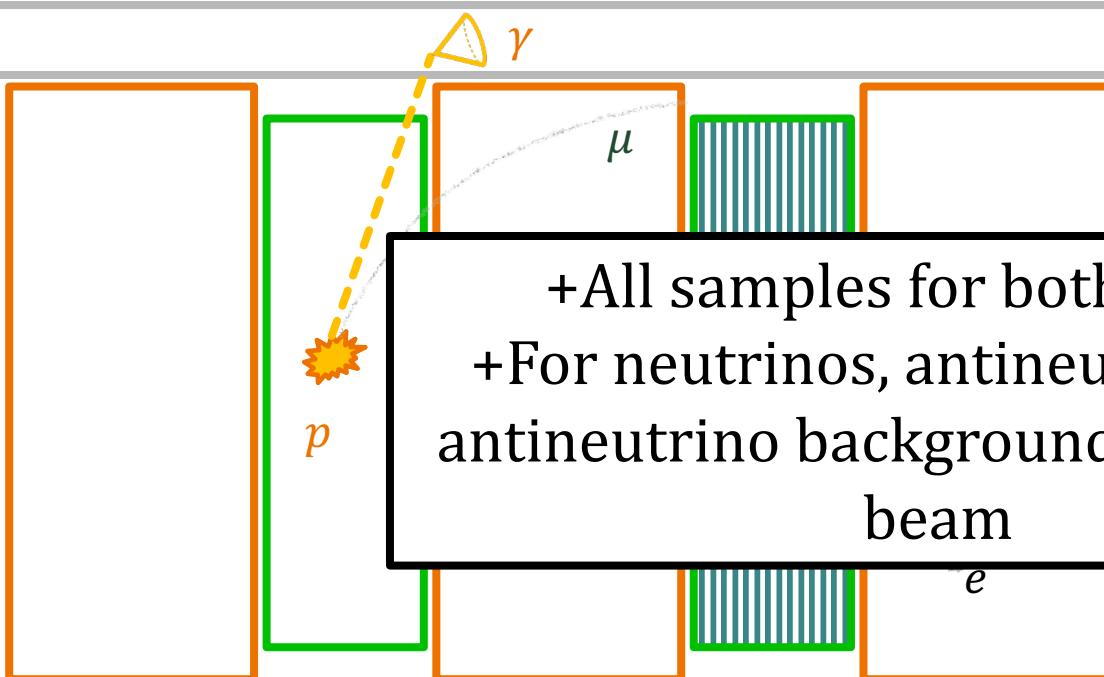
$1\mu + \text{any other tracks}$
Enriched in DIS interactions

Samples: CC γ

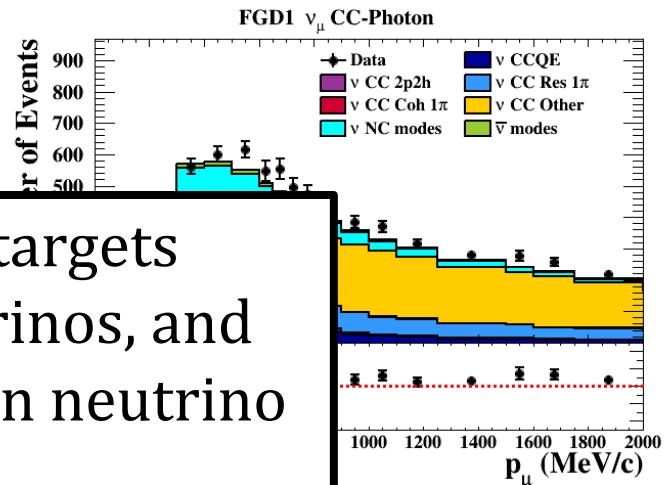


$1\mu + \text{photon}$
Enriched in DIS interactions

Samples: CC γ

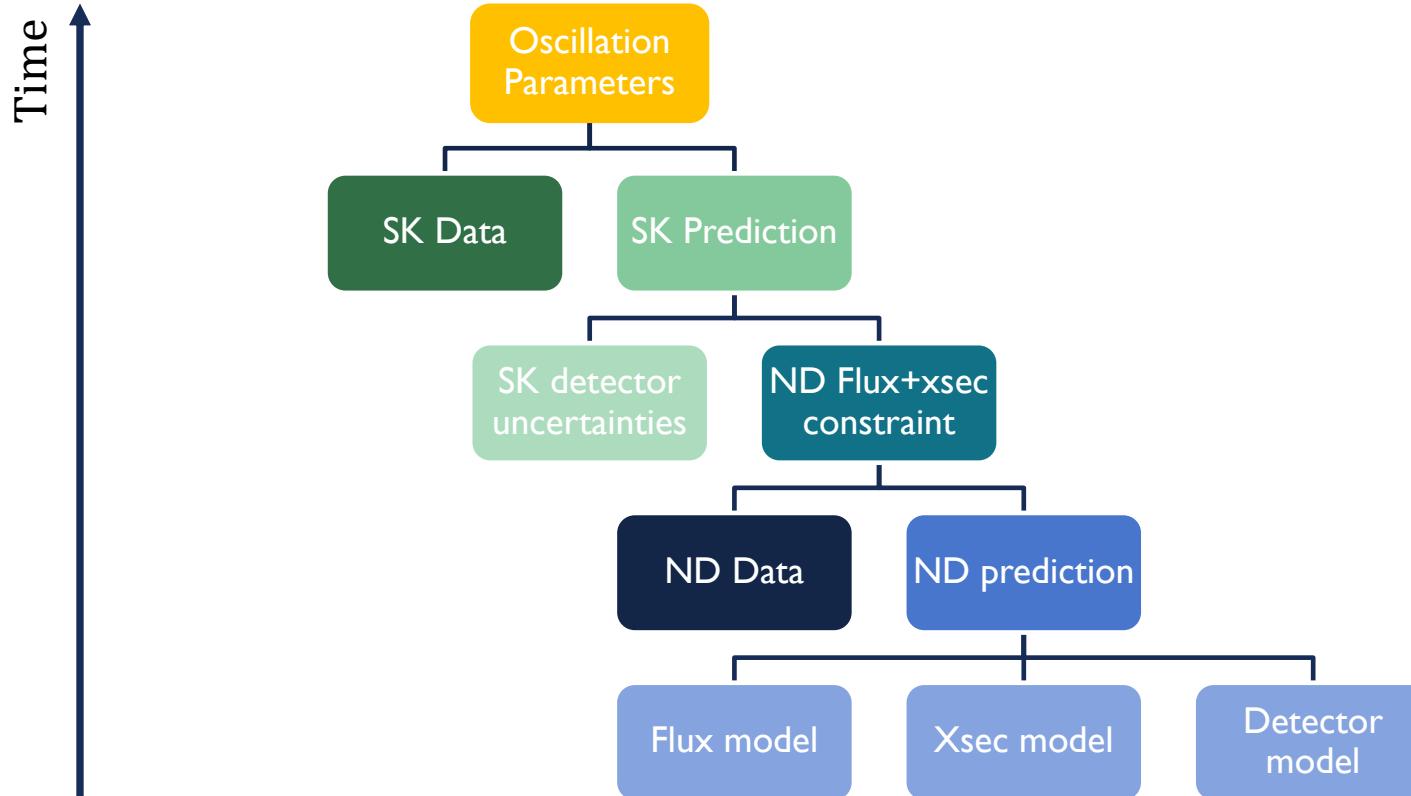


+All samples for both targets
+For neutrinos, antineutrinos, and
antineutrino background in neutrino
beam

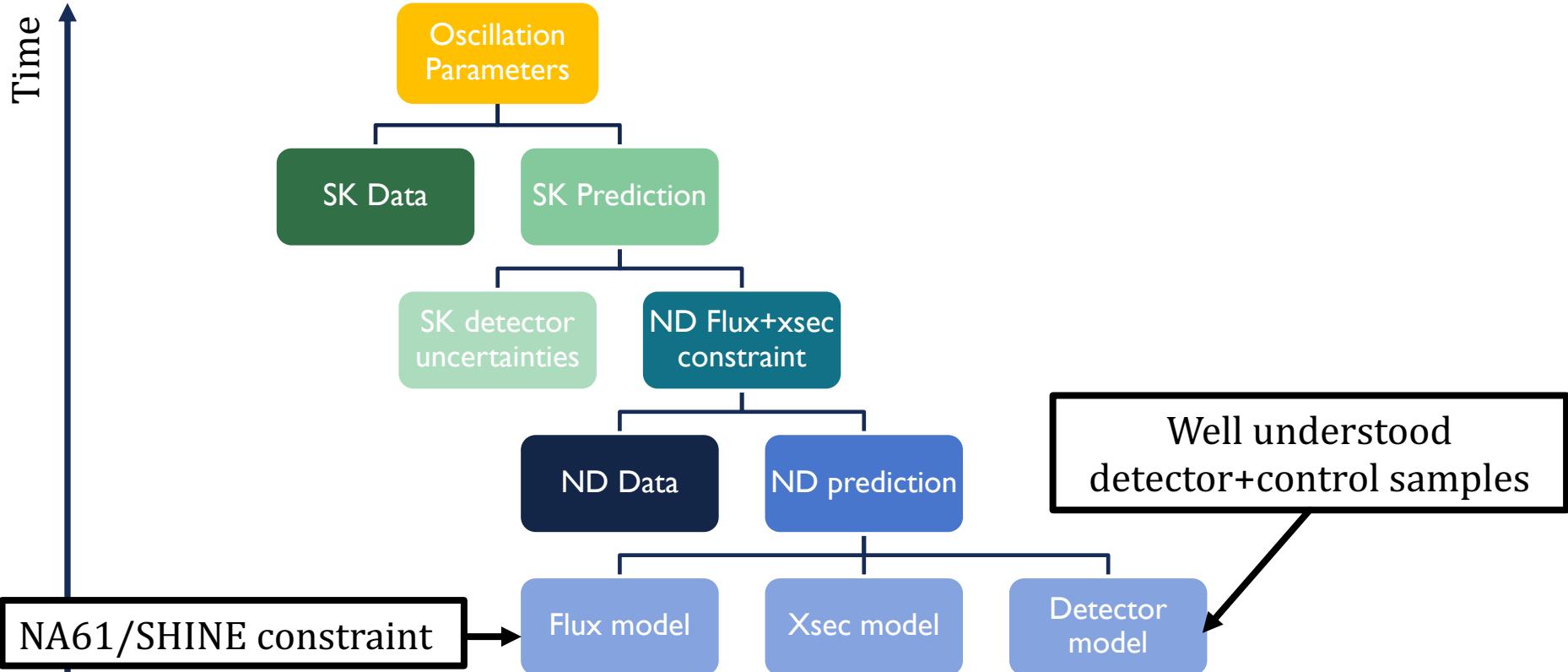


$\tau\mu + \text{photon}$
Enriched in DIS interactions

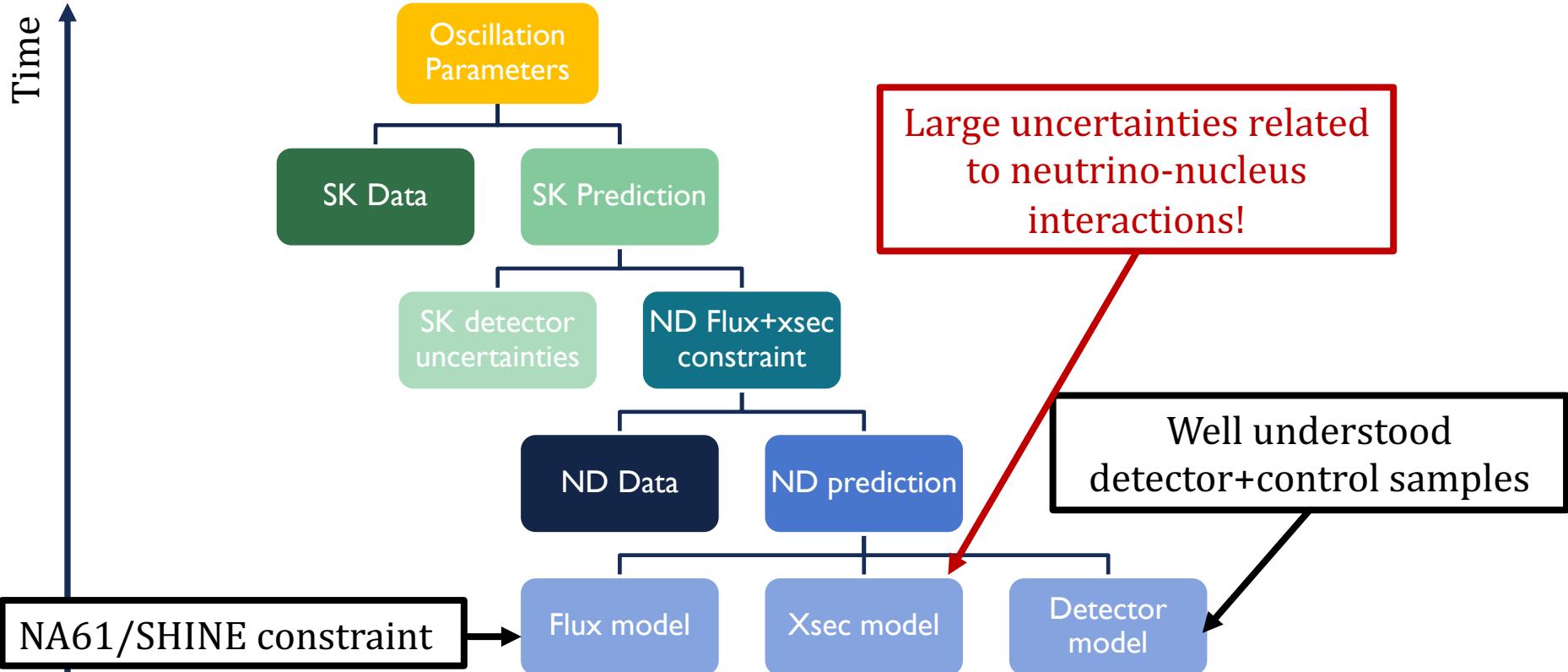
Analysis strategy



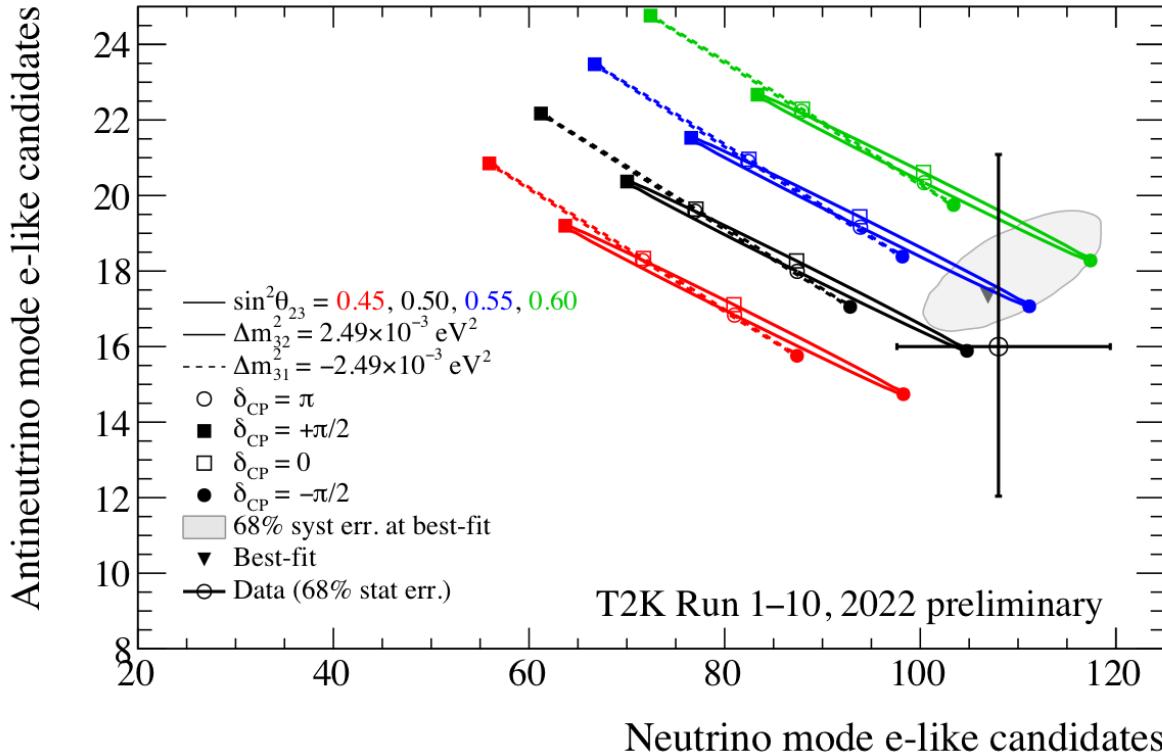
Analysis strategy



Analysis strategy

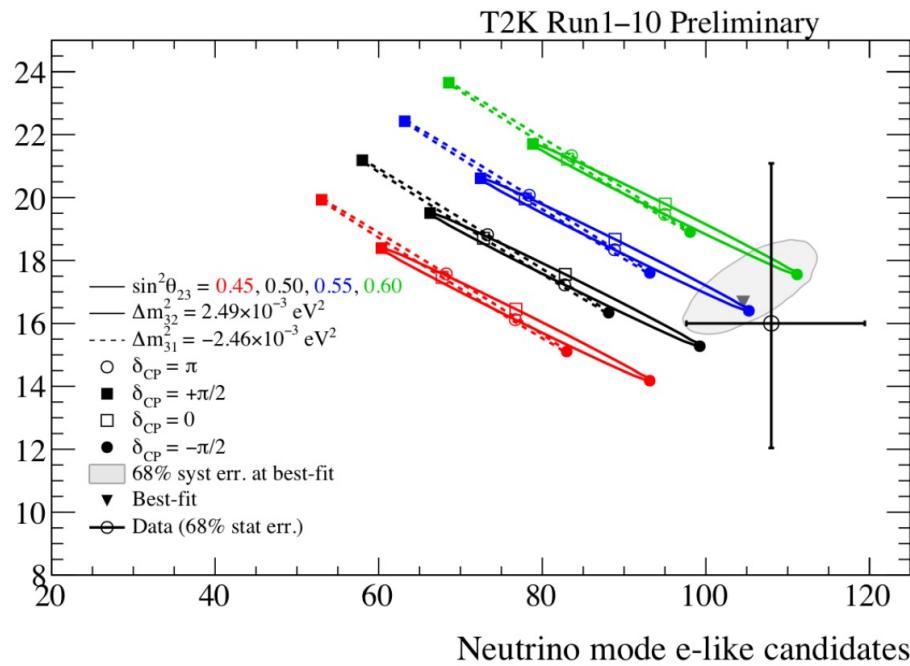


T2K Bi-probability plot



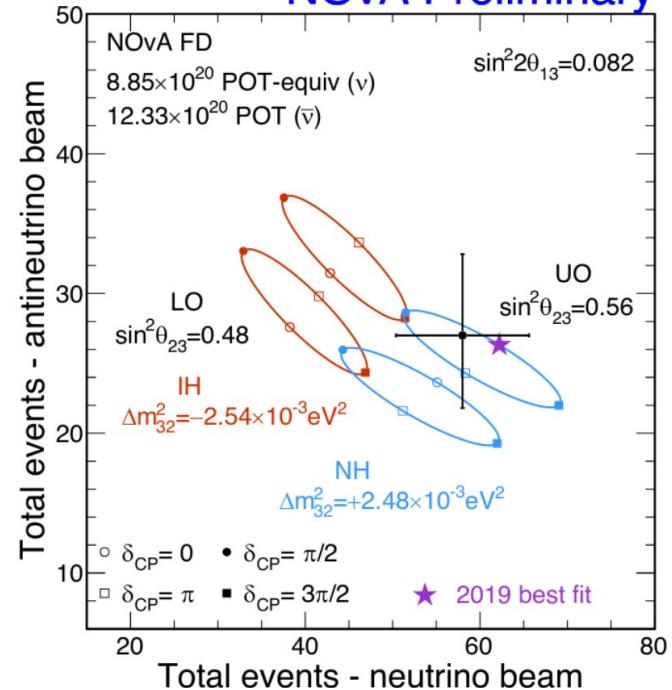
T2K-NOvA joint fit

Antineutrino mode e-like candidates

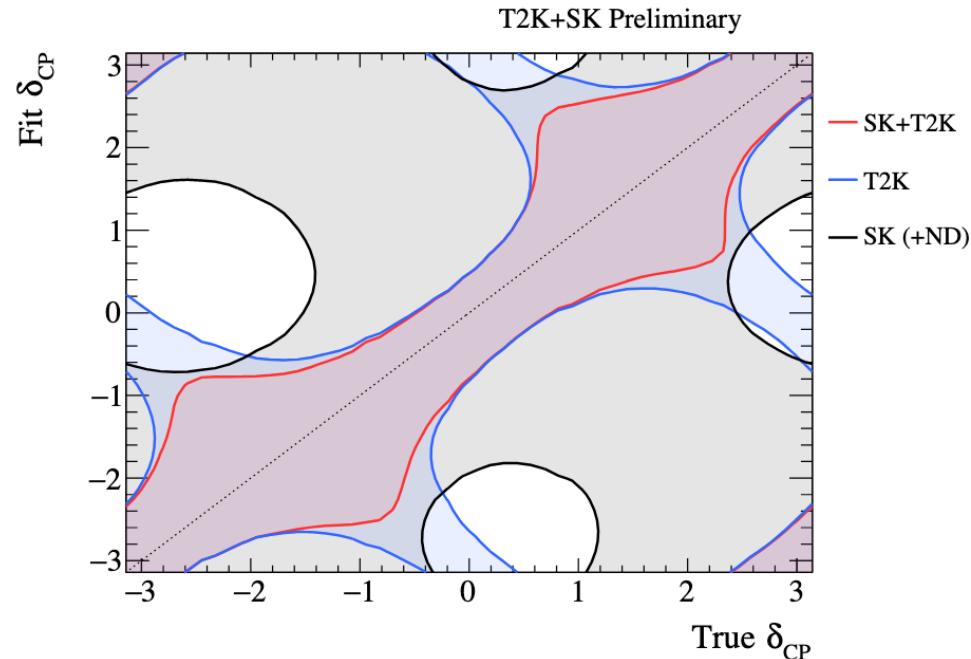
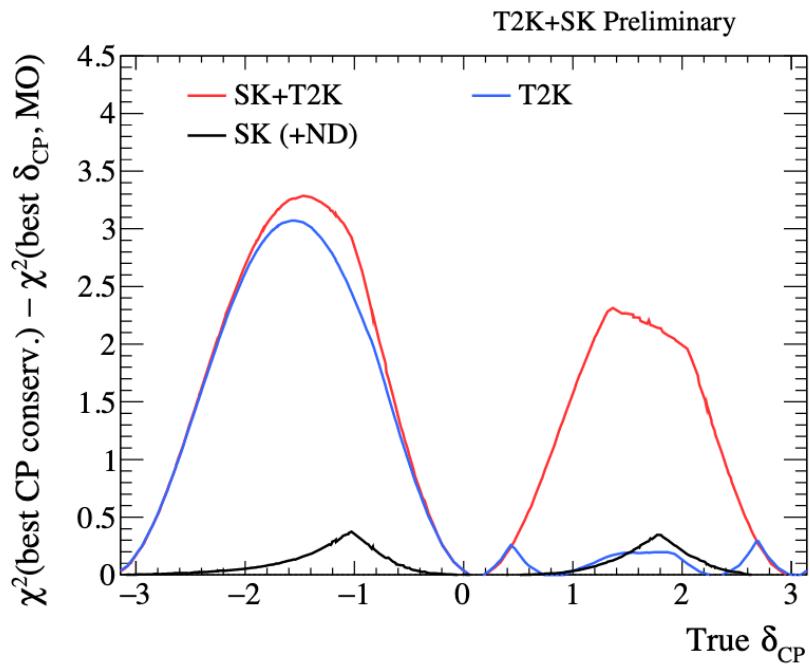


FNAL Users Meeting2019

NOvA Preliminary

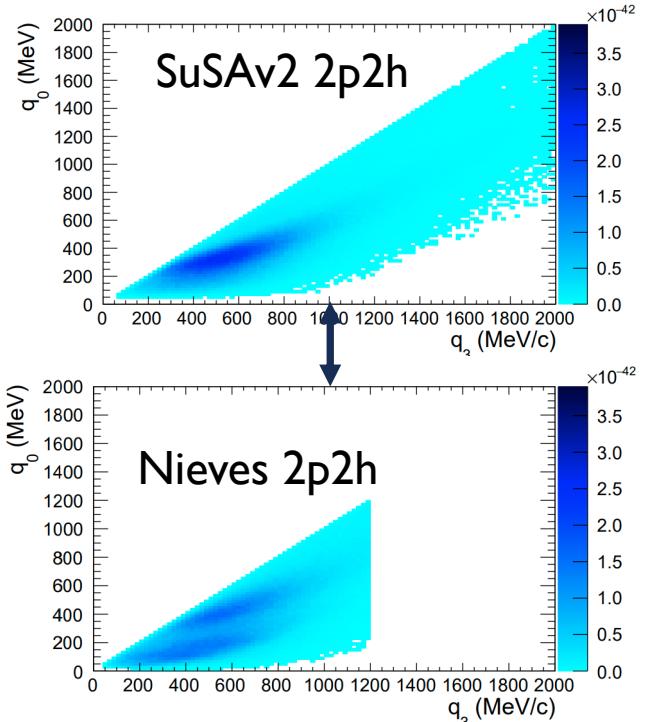


T2K+SK Atospheric Joint Fit



DUNE interaction model

- The expansion of the ground state model is part of a broader effort to increase the flexibility of our analyses
- Interaction model choices:
 - Valencia 1p1h model for CCQE interactions
 - Z-expansion form factor
 - SuSAv2 2p2h
 - Berger-Sehgal for Resonant interactions
 - Bodek-Yang DIS
 - hA2018 for hadronization
- Each of these parts of the model requires the development of systematic uncertainties tailored to DUNE-era statistics



[Phys. Rev. D 101, 033003 \(2020\)](#)